

$^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ at the INFN Bellotti Ion Beam Facility

D. RAPAGNANI, UNIVERSITY OF NAPLES "FEDERICO II"

on behalf of the LUNA collaboration

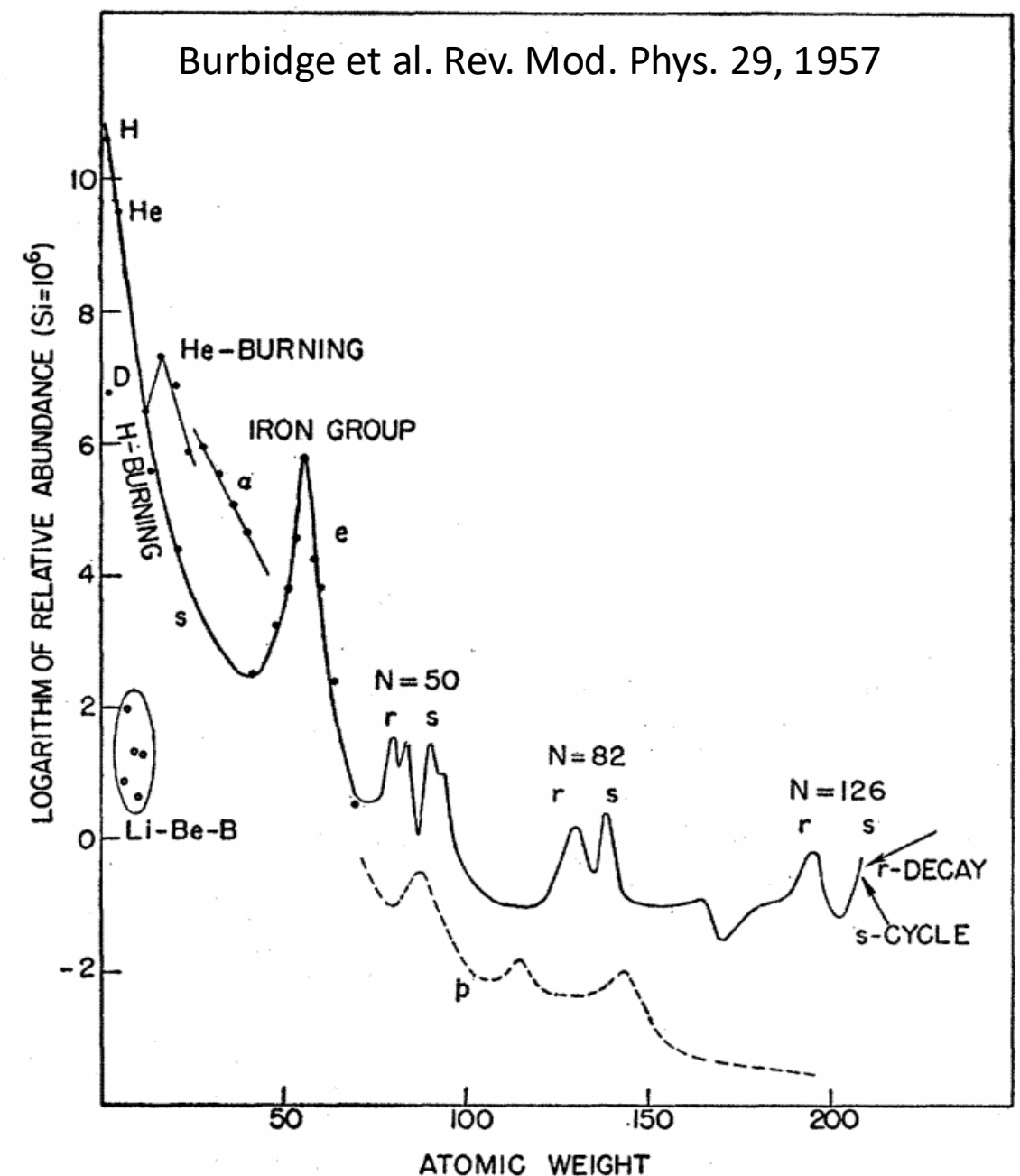
david.rapagnani@unina.it

Slow neutron capture sources

- Elements heavier than Li produced in stars
 - $A < \text{Fe}$ produced through charged-particle capture
 - $A > \text{Fe}$ neutron capture (s-, r-, i-process)

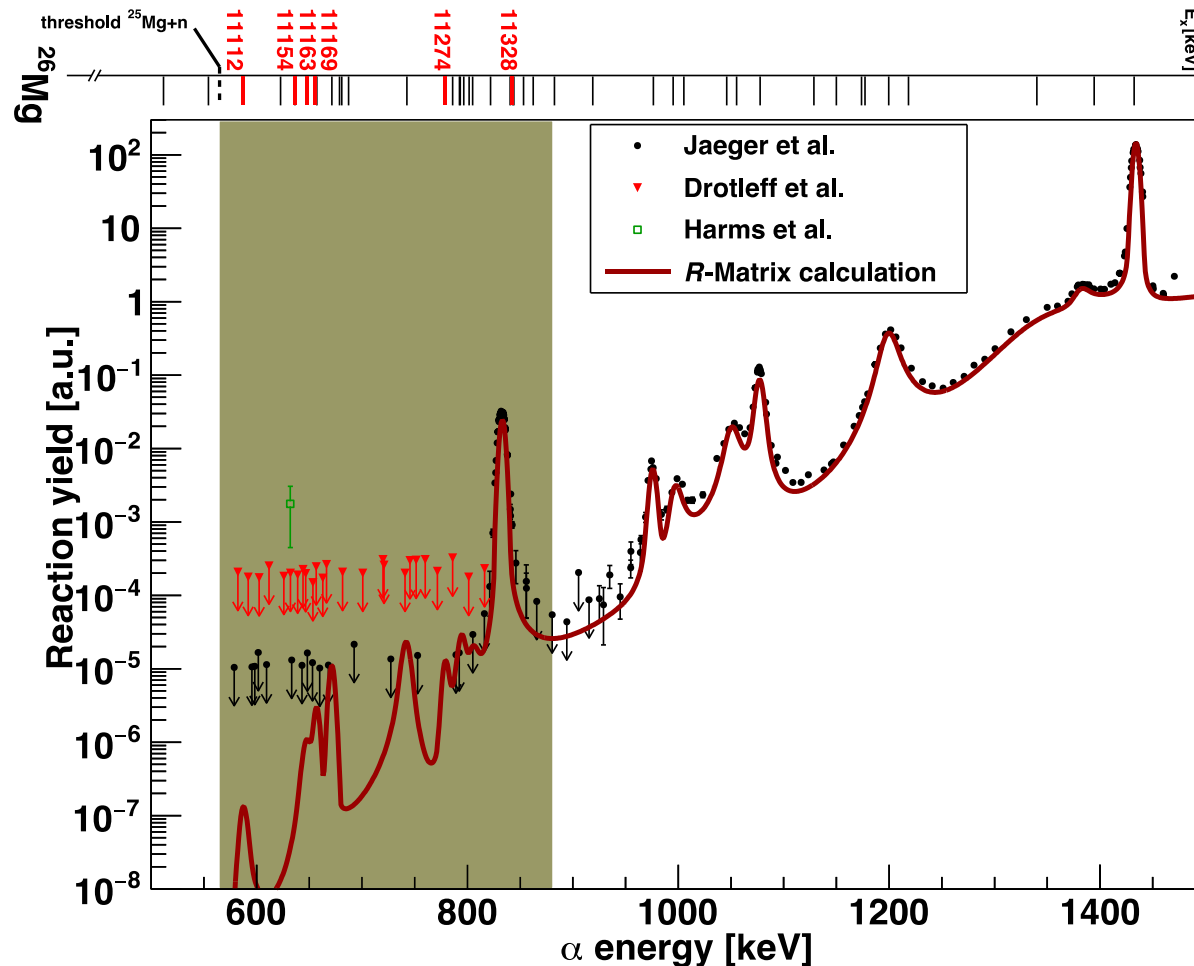
➤ $^{13}\text{C}(\alpha;n)^{16}\text{O}$ and $^{22}\text{Ne}(\alpha;n)^{25}\text{Mg}$ main neutron sources for s-process

- low mass TP-AGB stars and higher mass AGB stars
- Temperature regimes $\sim 90 \text{ MK} / 250 \text{ MK}$ (and higher)
- Astrophysical observables
 - ^{25}Mg and ^{26}Mg abundance (pre-solar grains)
 - $^{25}\text{Mg}/^{24}\text{Mg}$ and $^{26}\text{Mg}/^{24}\text{Mg}$ ratios from stellar spectra
 - ^{60}Fe in type II supernovae
 - ^{56}Ni in type Ia supernovae



Status of the art

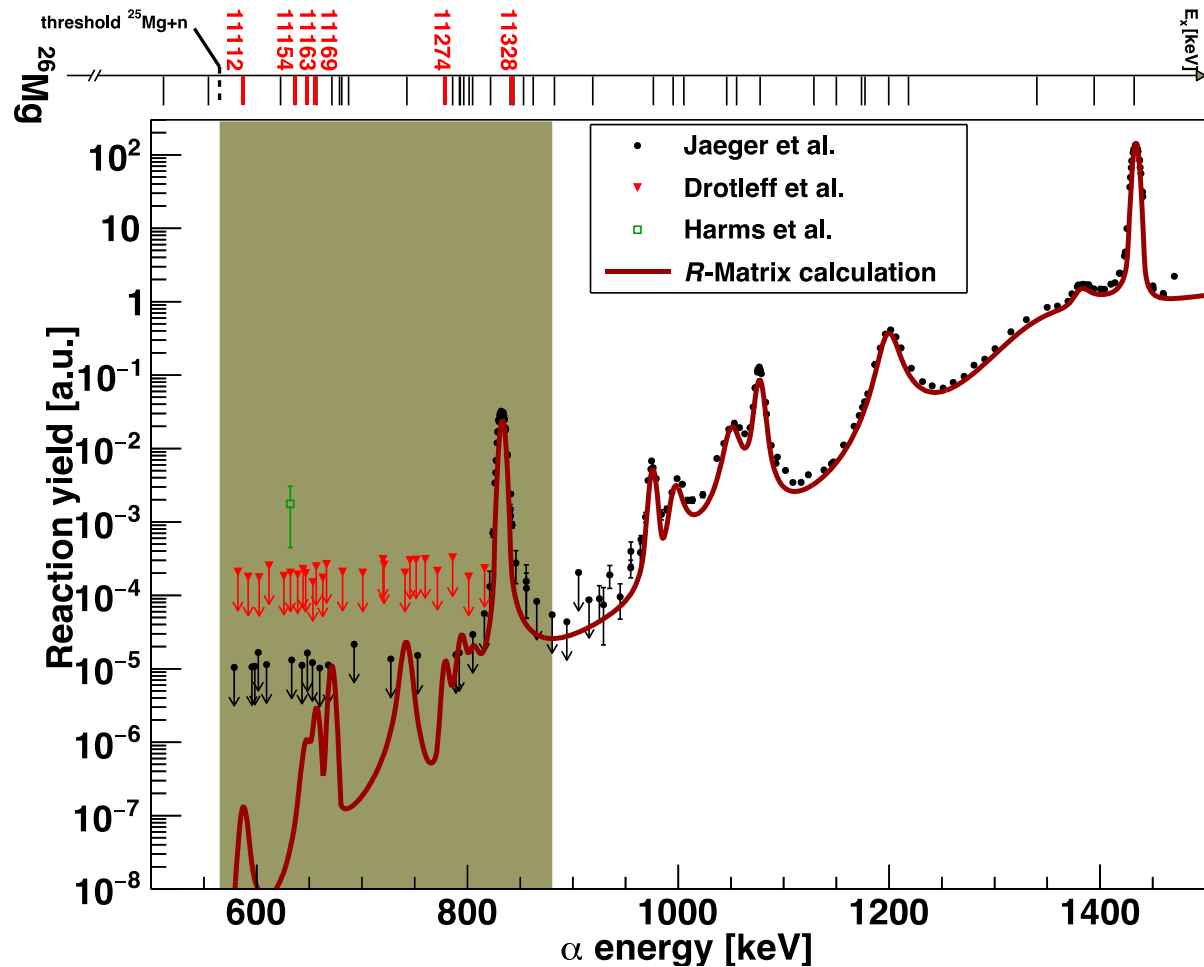
R matrix calculation courtesy of James deBoer



- $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ $Q = -478.296(89)$ keV
- only upper limits at the lower energies for direct experiments
- Γ_n and Γ_γ partial widths from indirect studies
 $\omega\gamma \sim \Gamma_\alpha$, at the moment inferred from α -transfer reactions
- Low event rate and high background main limit for direct data so far

➤ high beam current, reduce background

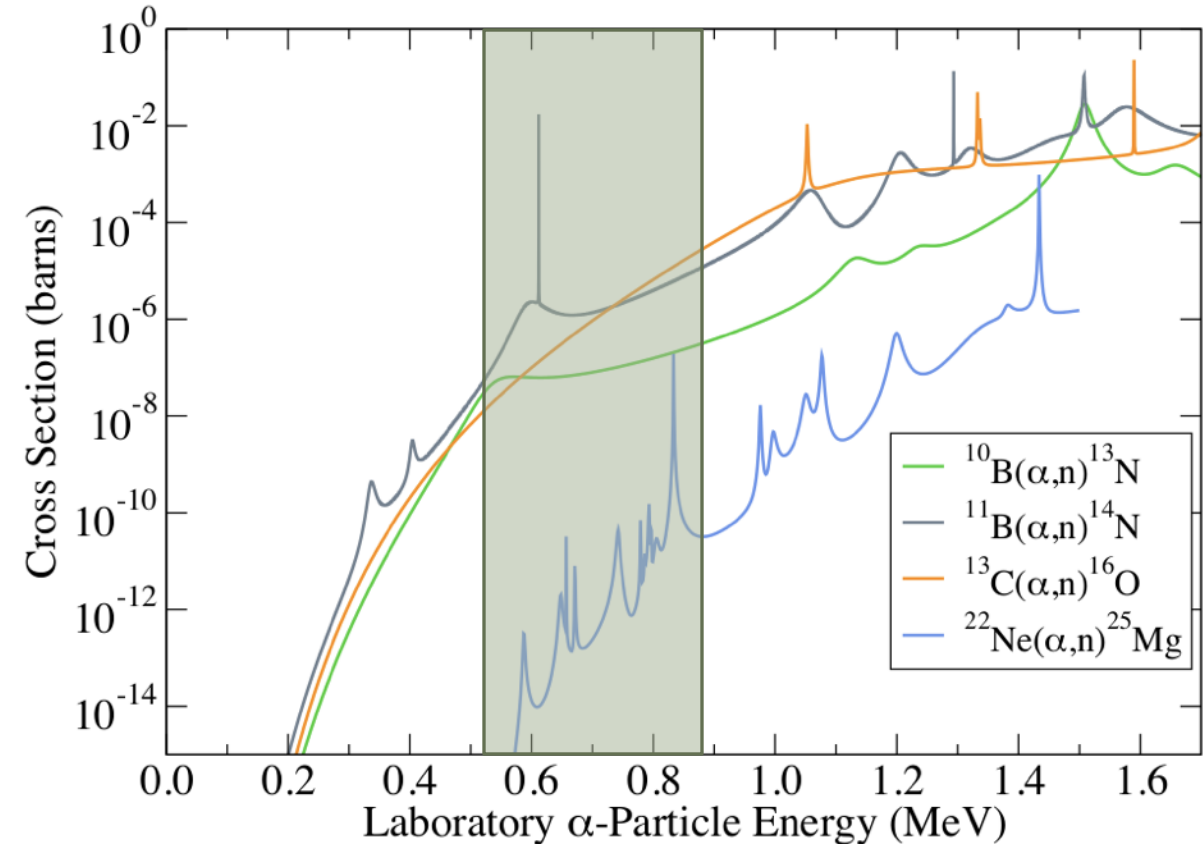
Beam induced background



low relevant/BIB cross section ratio

“clean” setup -> gas target design

neutron spectroscopy -> hybrid detection array

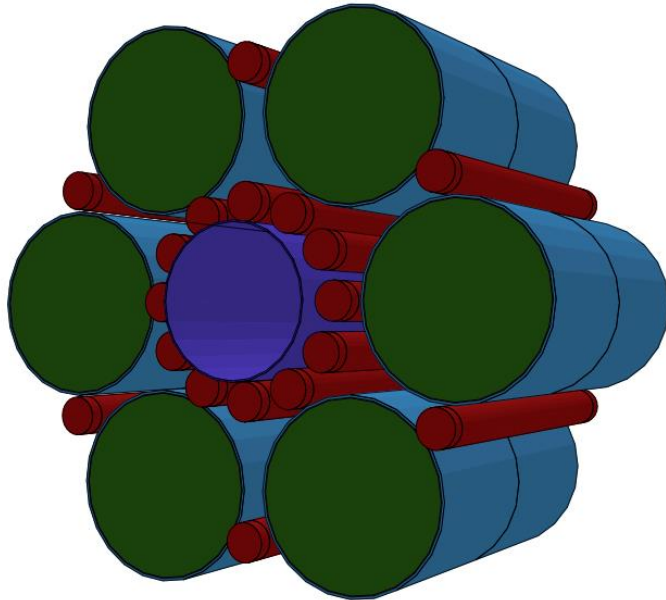


SHADES

Scintillator- ^3He Array for Deep-underground Experiments on the S-process

PI: prof. Andreas Best

Core working group: D. Rapagnani, D. Mercogliano, T. Chillery, A. Best



European Research Council funded project to measure $^{22}\text{Ne}(\alpha, n)^{25}\text{Mg}$ down to neutron threshold



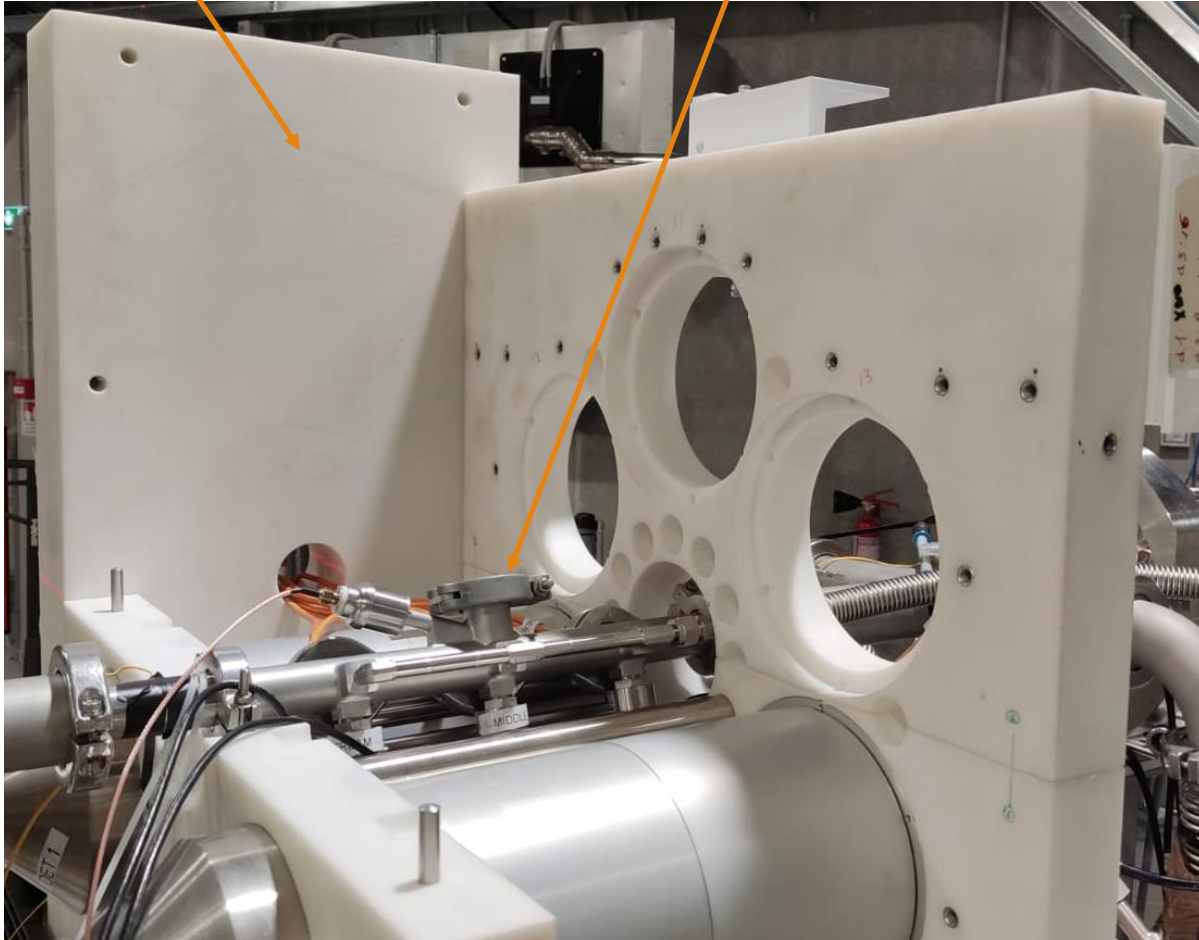
European Research Council

Established by the European Commission

Detection array

PE array holder

reaction chamber

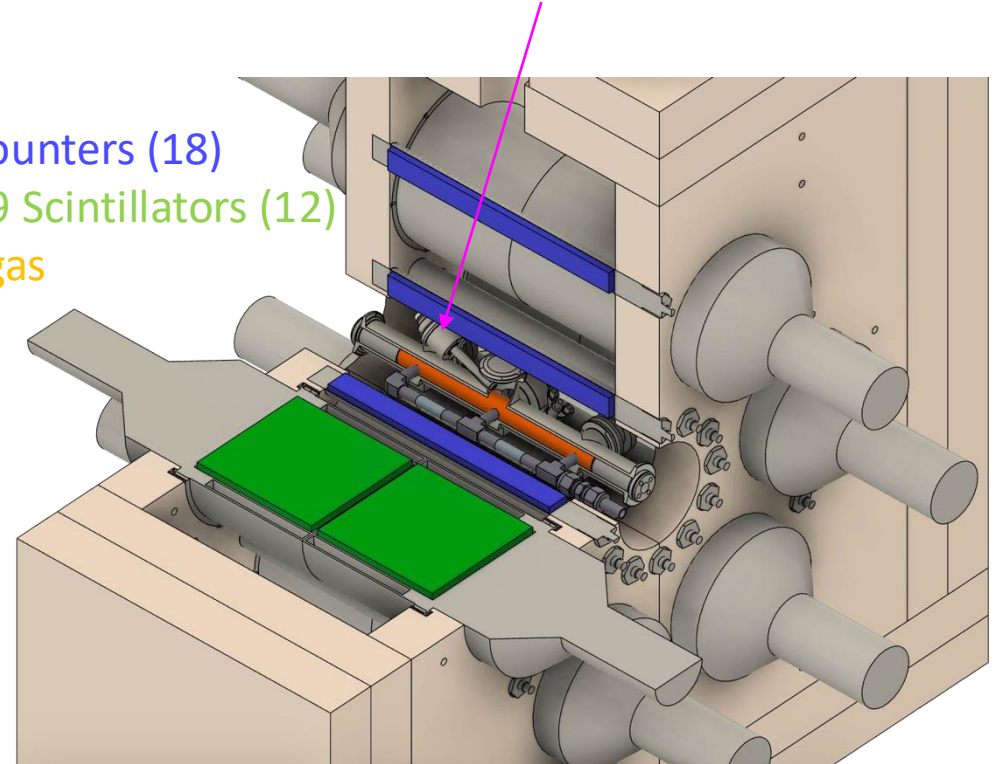


^3He counters (18)

EJ-309 Scintillators (12)

^{22}Ne gas

RBS Si detector

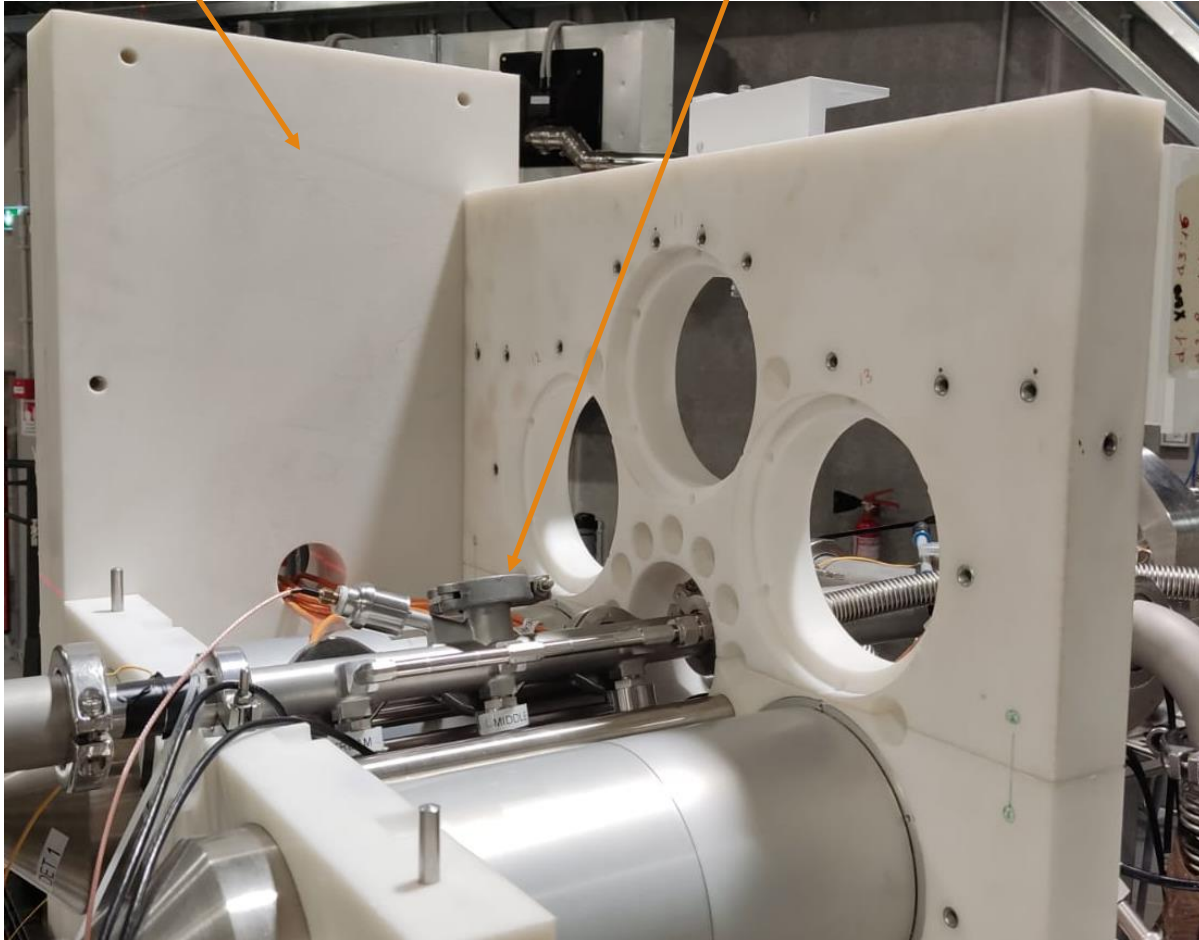


how it works?

Detection array

PE array holder

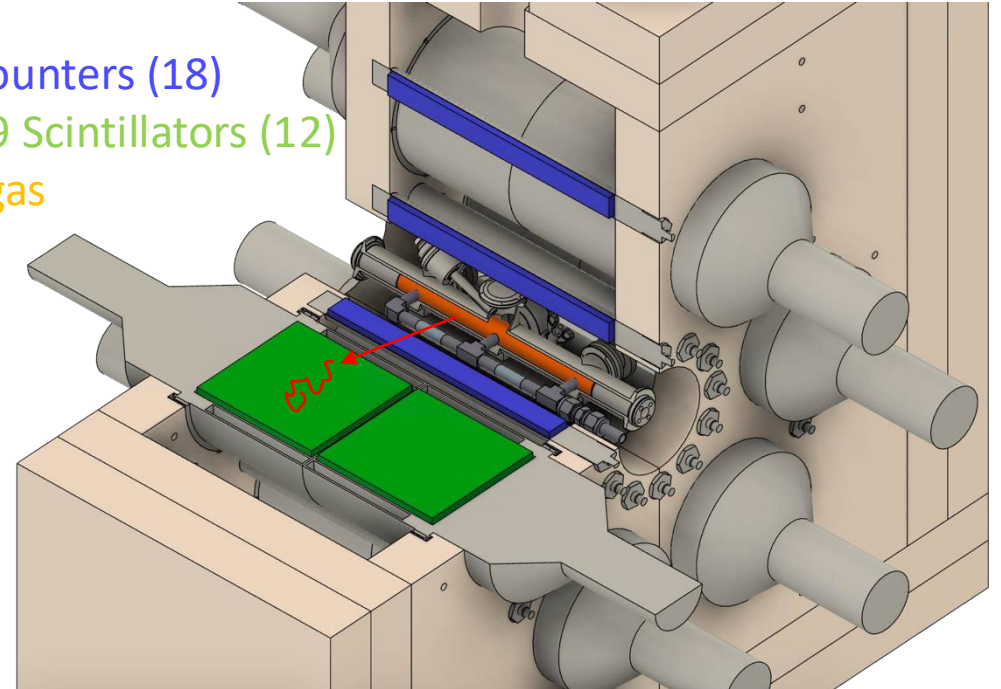
reaction chamber



^3He counters (18)

EJ-309 Scintillators (12)

^{22}Ne gas



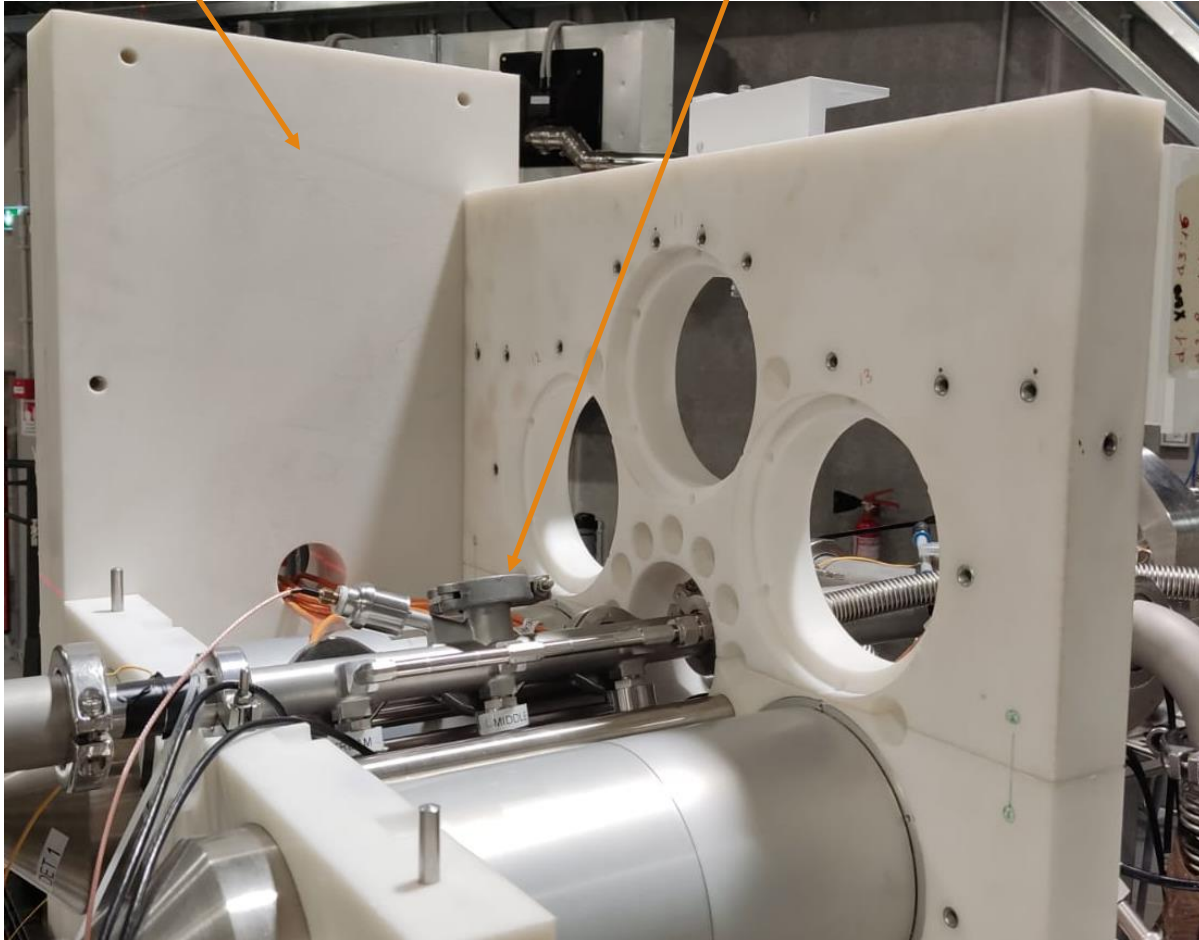
how it works?

- 1) a neutron from the reaction are moderated in an EJ-309 liquid scintillator

Detection array

PE array holder

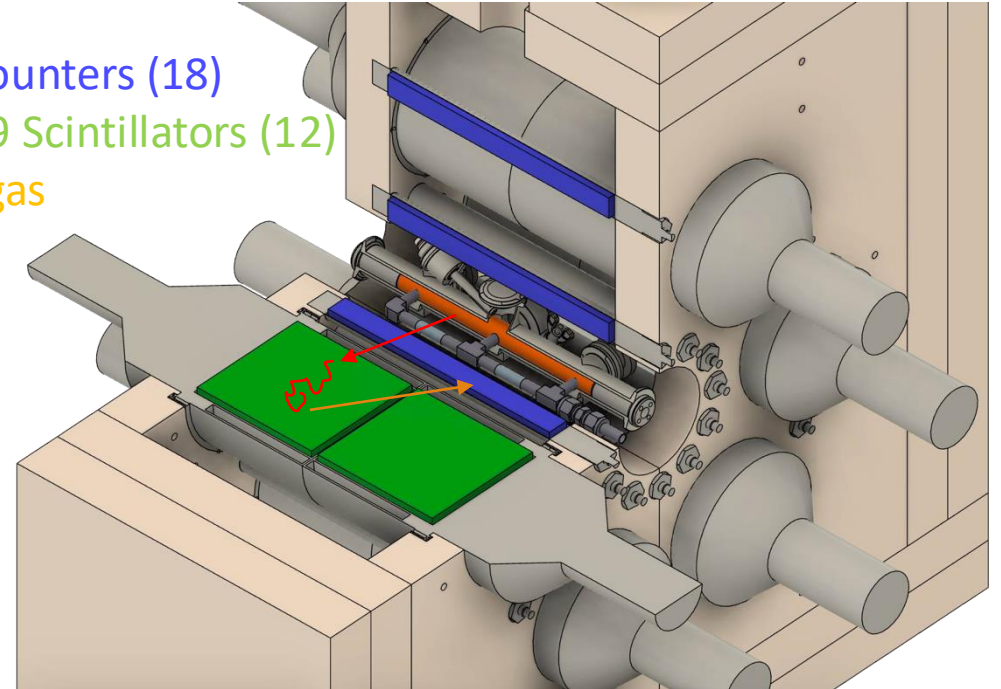
reaction chamber



^3He counters (18)

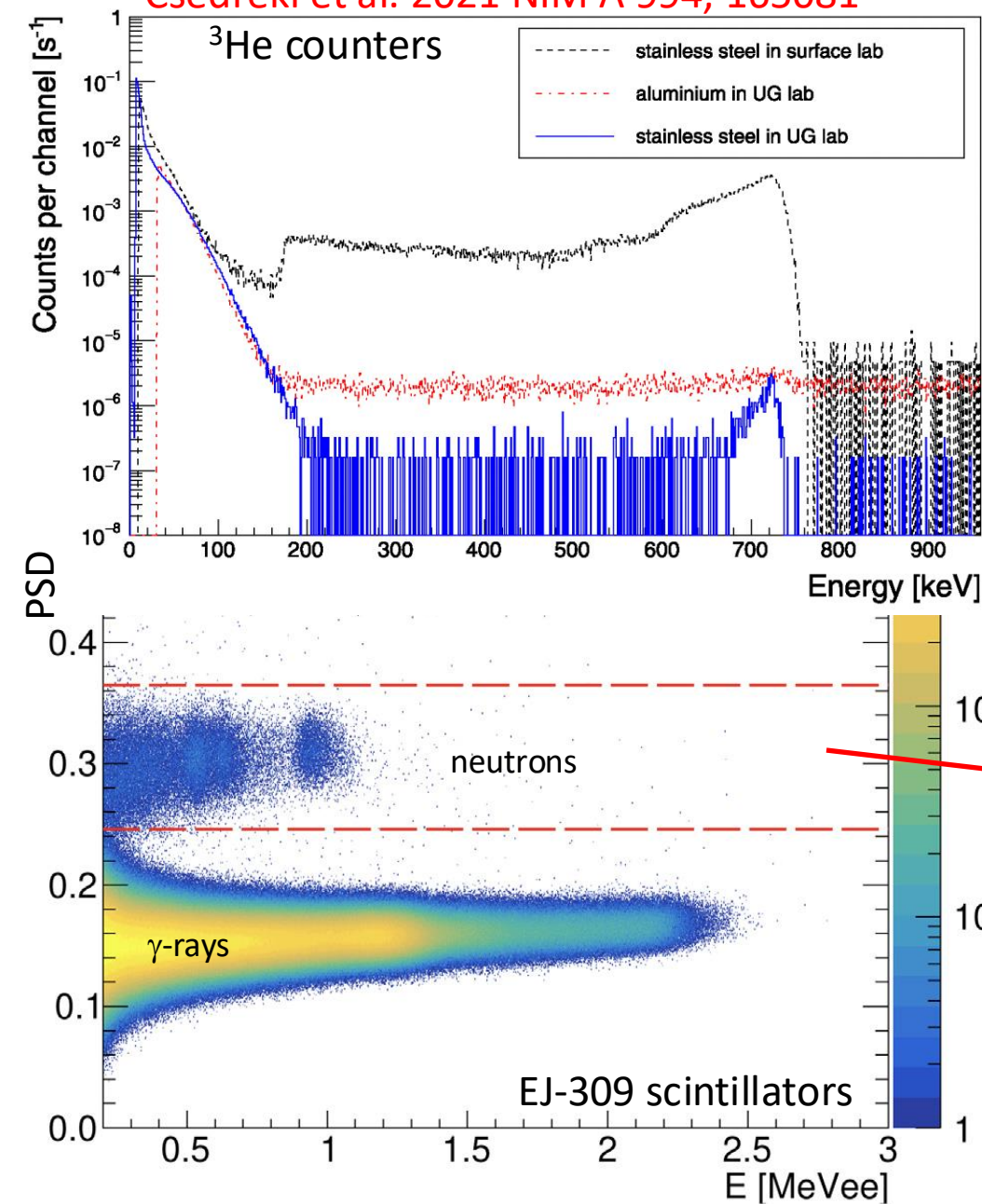
EJ-309 Scintillators (12)

^{22}Ne gas



how it works?

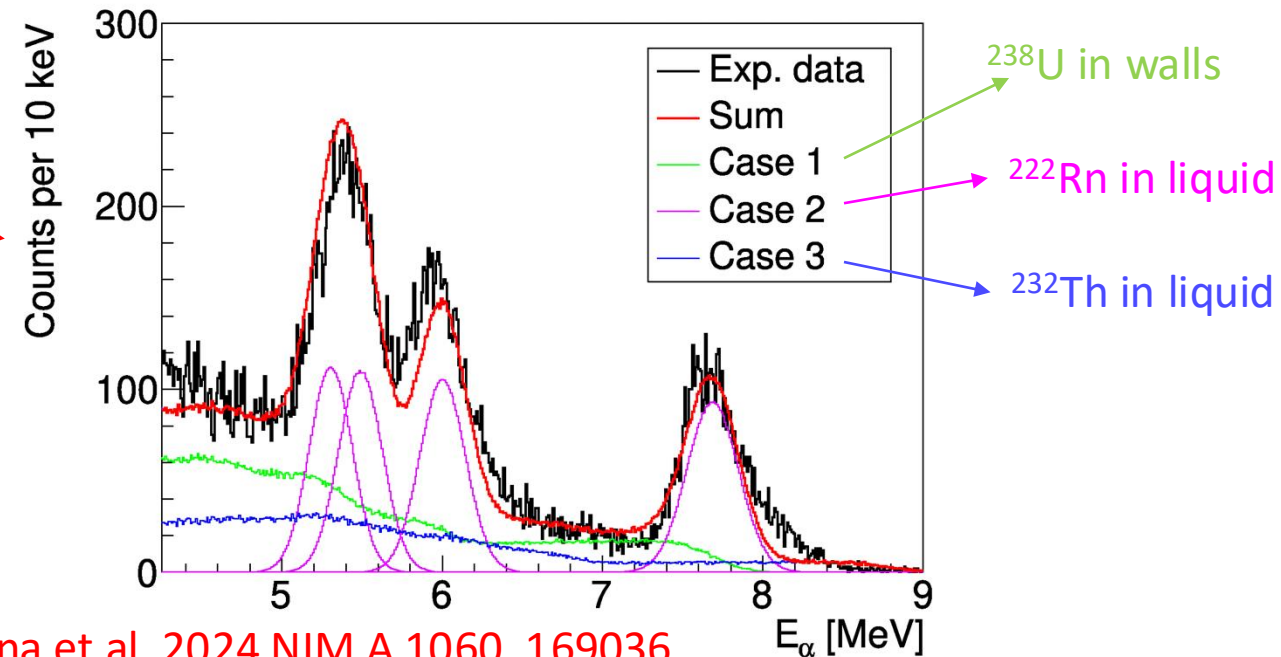
- 1) a neutron from the reaction are moderated in an EJ-309 liquid scintillator
- 2) the thermalized neutron is then captured by a ^3He counter



Detectors intrinsic background at LNGS

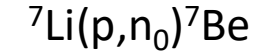
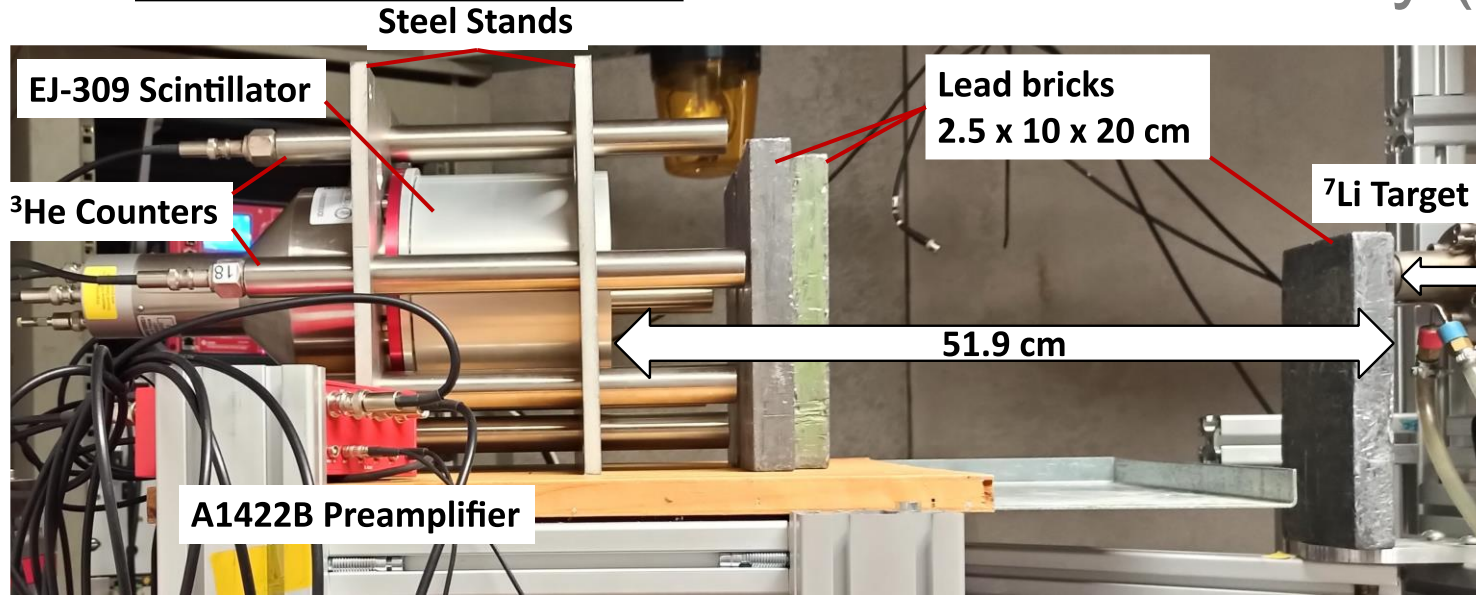
- LNGS environmental background makes intrinsic one relevant
- actinides contamination in both the case and liquid
- 1-month long background measurement in LNGS (LUNA400 hall)
- $8.4(1.8)_{\text{sta}}(1.4)_{\text{sys}} \times 10^{-2}$ ppm for ^{238}U
- $1.62(0.57)_{\text{sta}}(0.03)_{\text{sys}} \times 10^{-1}$ ppm for ^{232}Th
- 64 counts/hour, eliminable with ^3He prop. counter coincidence

➤ intrinsic background from $^{238}\text{U}/^{232}\text{Th}$ identified



Ananna et al. 2024 NIM A 1060, 169036

Calibration at FRANZ at Goethe University (Frankfurt)



$E_p = 1900 - 2450 \text{ keV}$, 50 keV steps

3.1 μm thick ^7Li on copper backings

$E_n = 50 - 720 \text{ keV}$

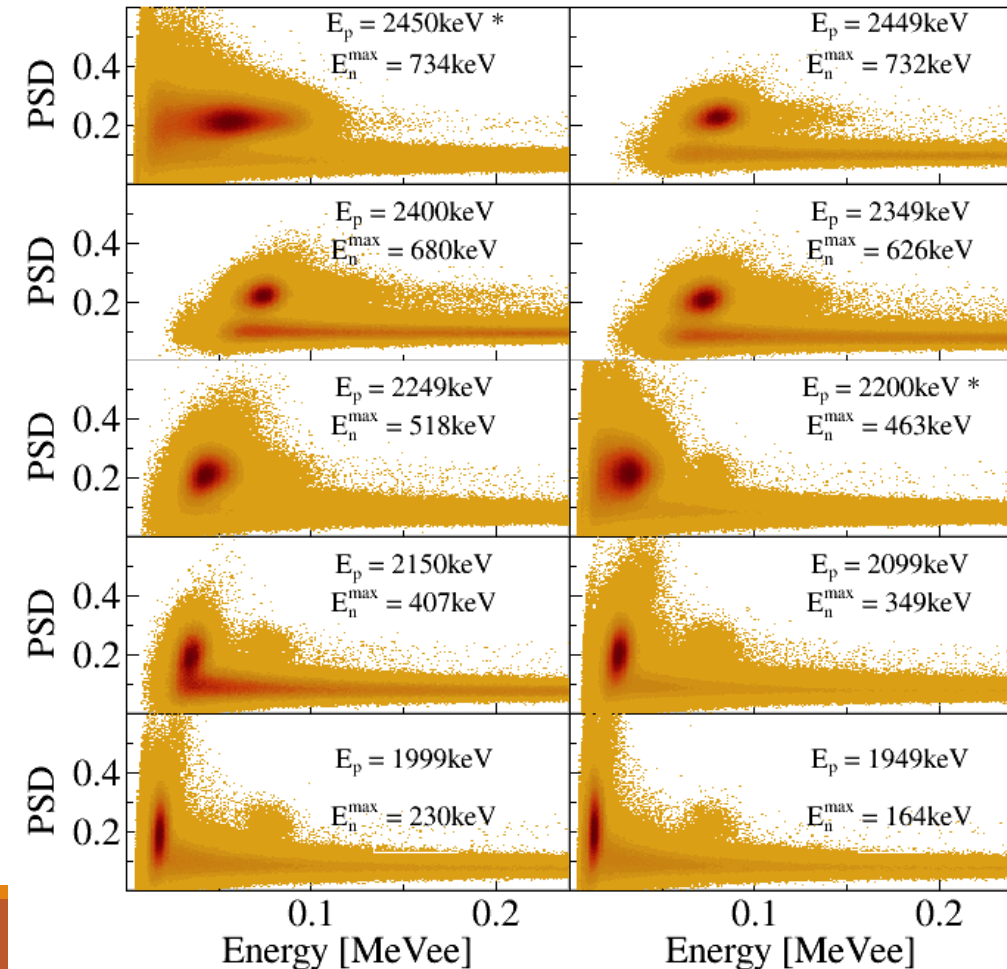
- prototype version, same working principle
- determination of scintillators neutron quenching
- testing on neutron/ γ -rays discrimination techniques (PSD and neural networks)
- counters-scintillators coincidence check

➤ performance assessment (low E limit, resolution, n/ γ separation)

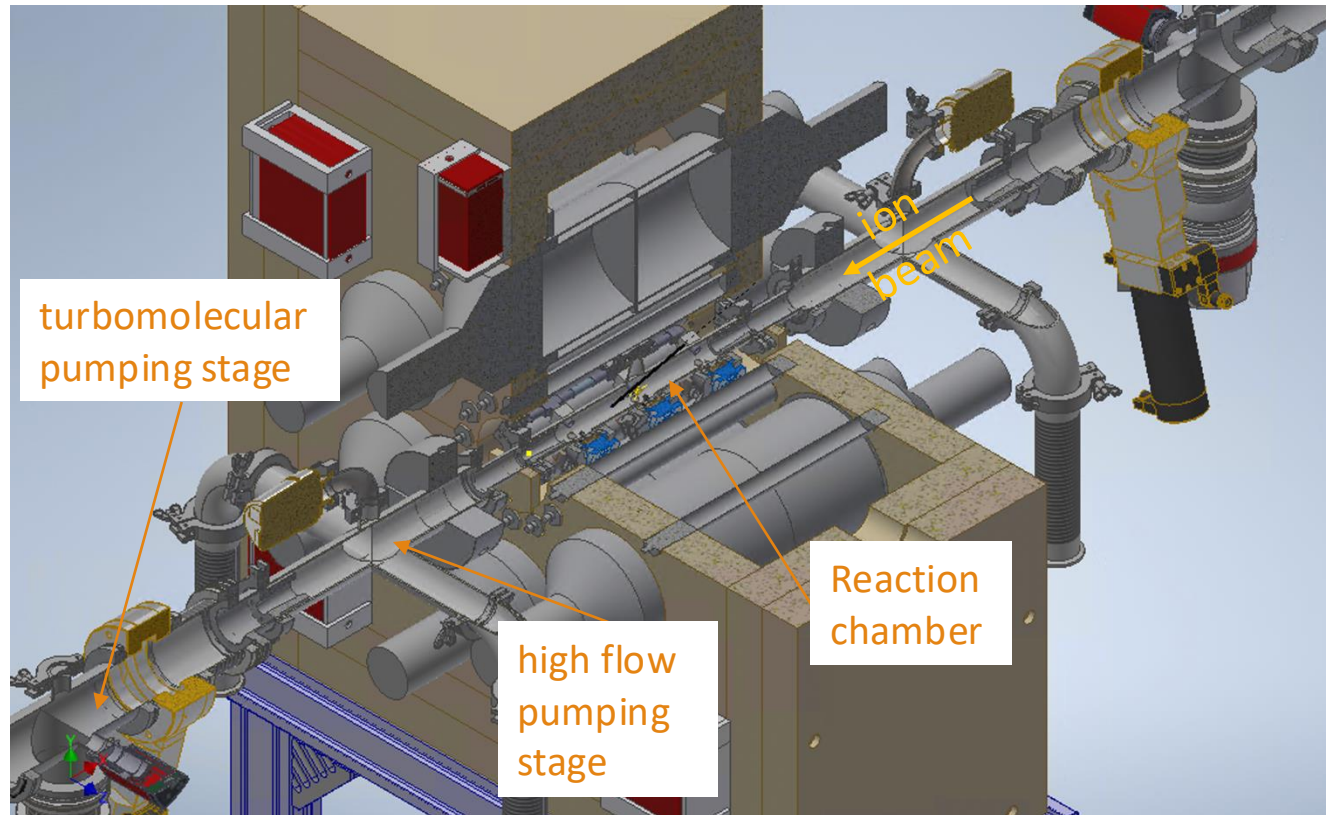
T. Chillery et al. 2025 under Review J. Phys. G

T. Chillery poster

Proton Beam



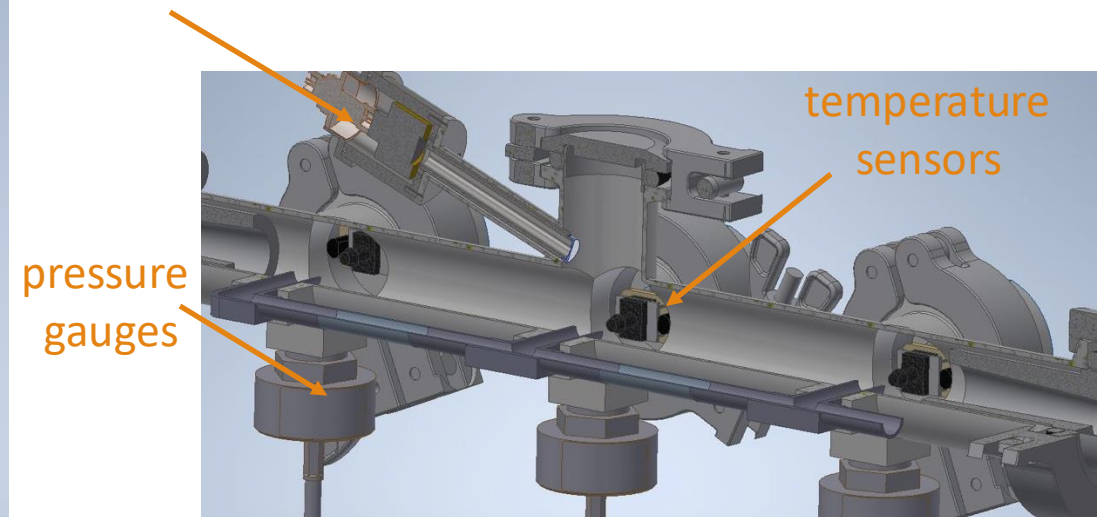
Gas target



- recirculated extended gas target with enriched ^{22}Ne
- compact design to maximize detection efficiency
- hydrocarbons purifier
- Si-industry pumps

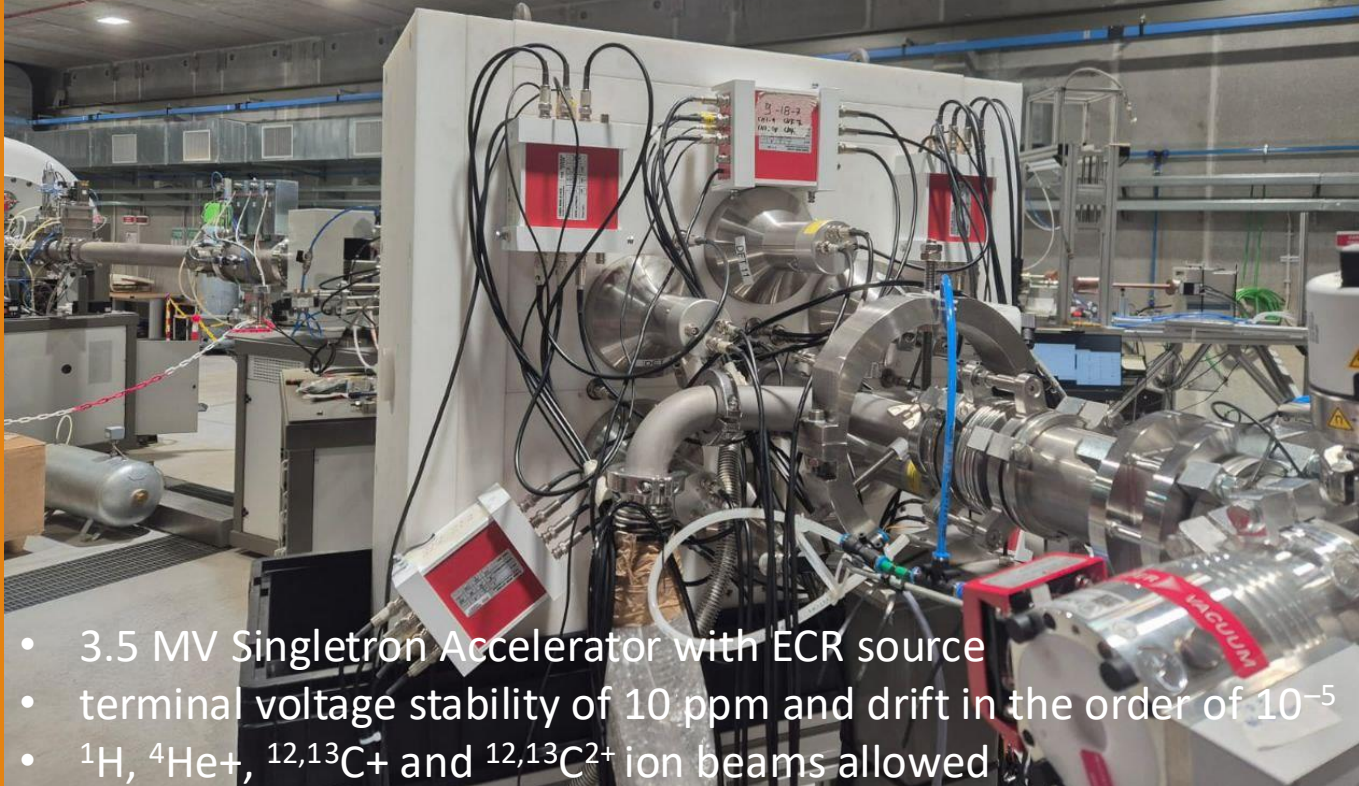
- suppressed BIB events rate
- ion beam current and target thickness monitor

RBS monitor



- Aluminum sealings – low leak rate and no carbon degassing
- OFHC copper coatings of exposed surfaces
- online monitors of pressure and temperatures
- Rutherford Backscattering Spectroscopy for ion beam / target thickness monitor

Bellotti IBF @ LNGS



- 3.5 MV Singletron Accelerator with ECR source
- terminal voltage stability of 10 ppm and drift in the order of 10^{-5}
- ^1H , $^4\text{He}^+$, $^{12,13}\text{C}^+$ and $^{12,13}\text{C}^{2+}$ ion beams allowed
- 500 μA of high energy resolution and stability ^4He beams

➤ enhanced S/N ratio over natural background

Junker et al. 2023 Front. Phys. 11:1291113

completed measurements

$^{14}\text{N}(\text{p},\gamma)^{15}\text{O}$

ongoing measurements

$^{22}\text{Ne}(\alpha,\text{n})^{25}\text{Mg}$

$^{12}\text{C}+^{12}\text{C}$ (γ -channel)

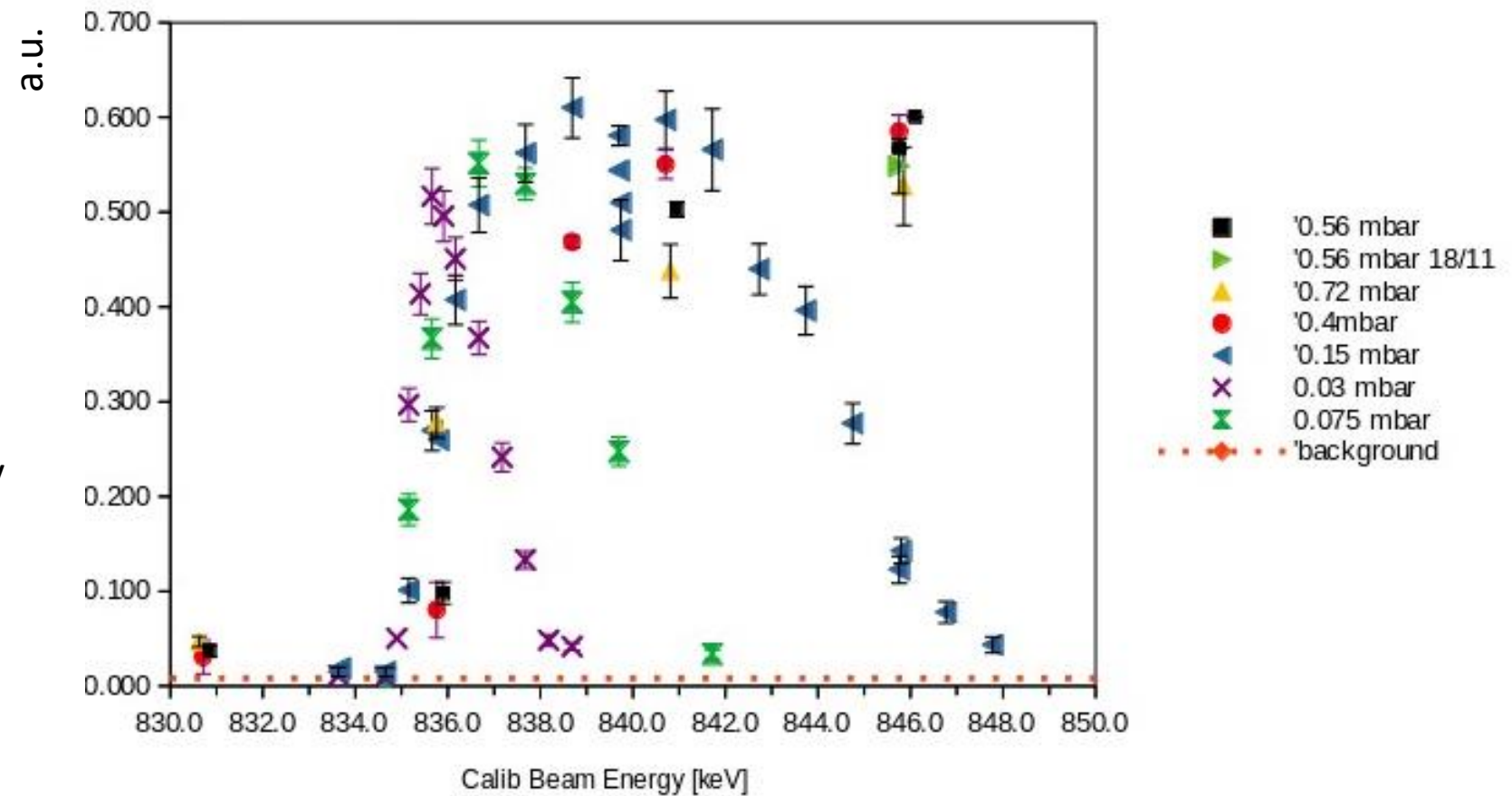
planned measurements

$^{22}\text{Ne}(\alpha,\gamma)^{26}\text{Mg}$ – Daniela Mercogliano talk



Preliminary results & outlooks

- 835 keV resonance investigated with several target thickness
- off-resonance measurements ongoing, complete within 2025
- data analysis and publication by mid-2026





Laboratori Nazionali del Gran Sasso, INFN, ASSERGI, Italy/*GSSI, L'AQUILA, Italy

T. Chillery, F. Ferraro, *R. Gesuè, M. Junker, D. Basak

Università degli Studi di Bari and INFN, BARI, Italy

G.F. Ciani

Konkoly Observatory, Hungarian Academy of Sciences, BUDAPEST, Hungary

M. Lugaro

Institute of Nuclear Research (ATOMKI), DEBRECEN, Hungary

L. Csedreki, Z. Elekes, Zs. Fülöp, Gy. Gyürky, T. Szücs

Helmholtz-Zentrum Dresden-Rossendorf, DRESDEN, Germany

D. Bemmerer, A. Boeltzig, E. Masha

University of Edinburgh, EDINBURGH, United Kingdom

M. Aliotta, L. Barbieri, C.G. Bruno, T. Davinson, J. Jones, J. Marsh, D. Robb, R. Bonnell, A. Compagnucci

Università degli Studi di Genova and INFN, GENOVA, Italy

P. Corvisiero, P. Prati, M. Rossi, S. Zavatarelli

INFN Laboratori Nazionali di Legnaro, LEGNARO, Italy

M. Campostrini, V. Rigato

Università degli Studi di Milano and INFN, MILANO, Italy

R. Depalo, G. Gosta, A. Guglielmetti

Università degli Studi di Napoli "Federico II" and INFN, NAPOLI, Italy

A. Best, D. Dell'Aquila, A. Di Leva, G. Imbriani, D. Mercogliano, D. Rapagnani

Università degli Studi di Padova and INFN, PADOVA, Italy

R. Biasissi, C. Brogгинi, A. Caciolli, R. Menegazzo, D. Piatti, J. Skowronski, S. Turkat

INFN Roma, ROMA, Italy

A. Formicola, C. Gustavino, M. Vagnoni

Osservatorio Astronomico di Collurania, TERAMO and INFN LNGS, Italy

O. Straniero, U. Battino

Università di Torino and INFN, TORINO, Italy

F. Cavanna, P. Colombetti, G. Gervino, R. Sariyal

**THANK YOU FOR
YOUR
ATTENTION**