

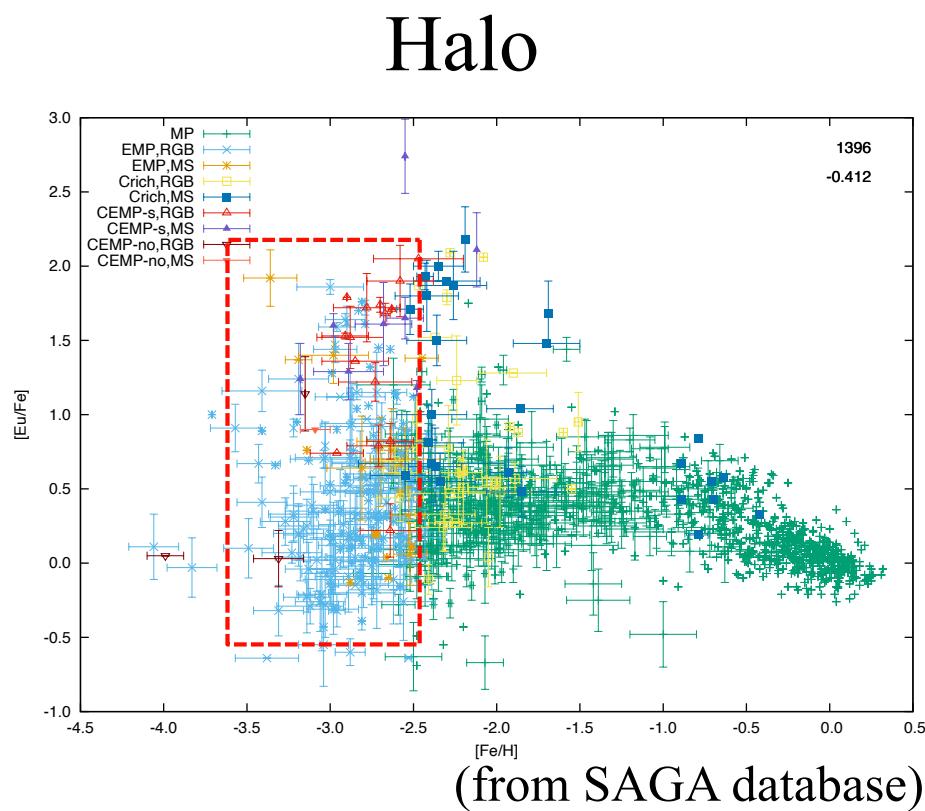
Evidence for r -process production by rare supernovae in the low-metallicity environment

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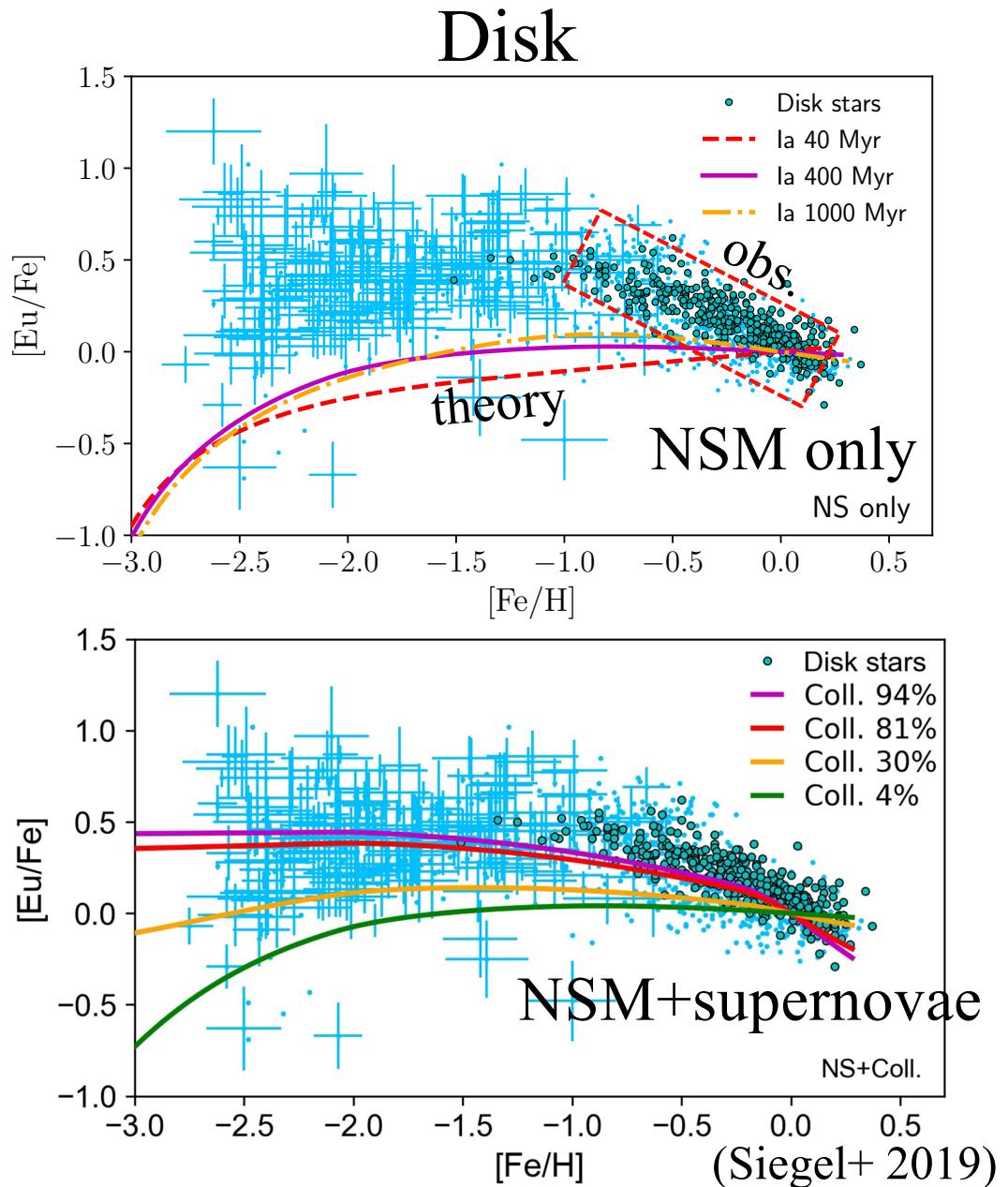
National Astronomical Observatory of Japan

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*Galactic chemical evolution suggests
supernovae must contribute to the *r*-process enrichment*



Extremely metal-poor stars are enriched by *r*-process elements



Candidates of r -process core-collapse SNe

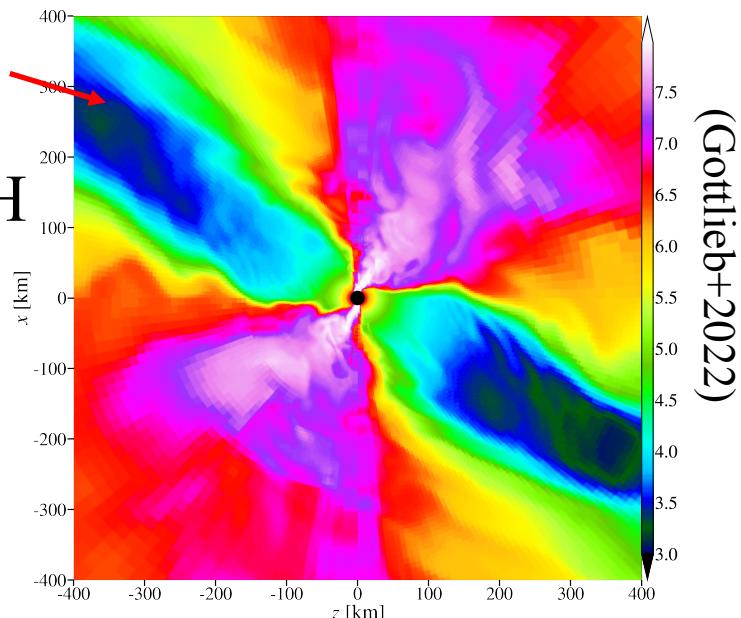
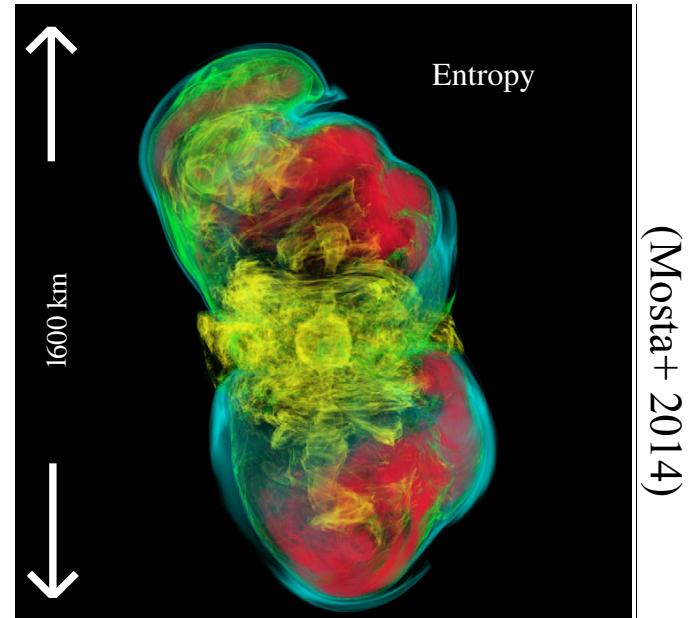
1. Magnetorotational SNe

- ✓ An explosion triggered by fast rotation and high magnetic fields
(e.g., Takiwaki+09, Kuroda+20)

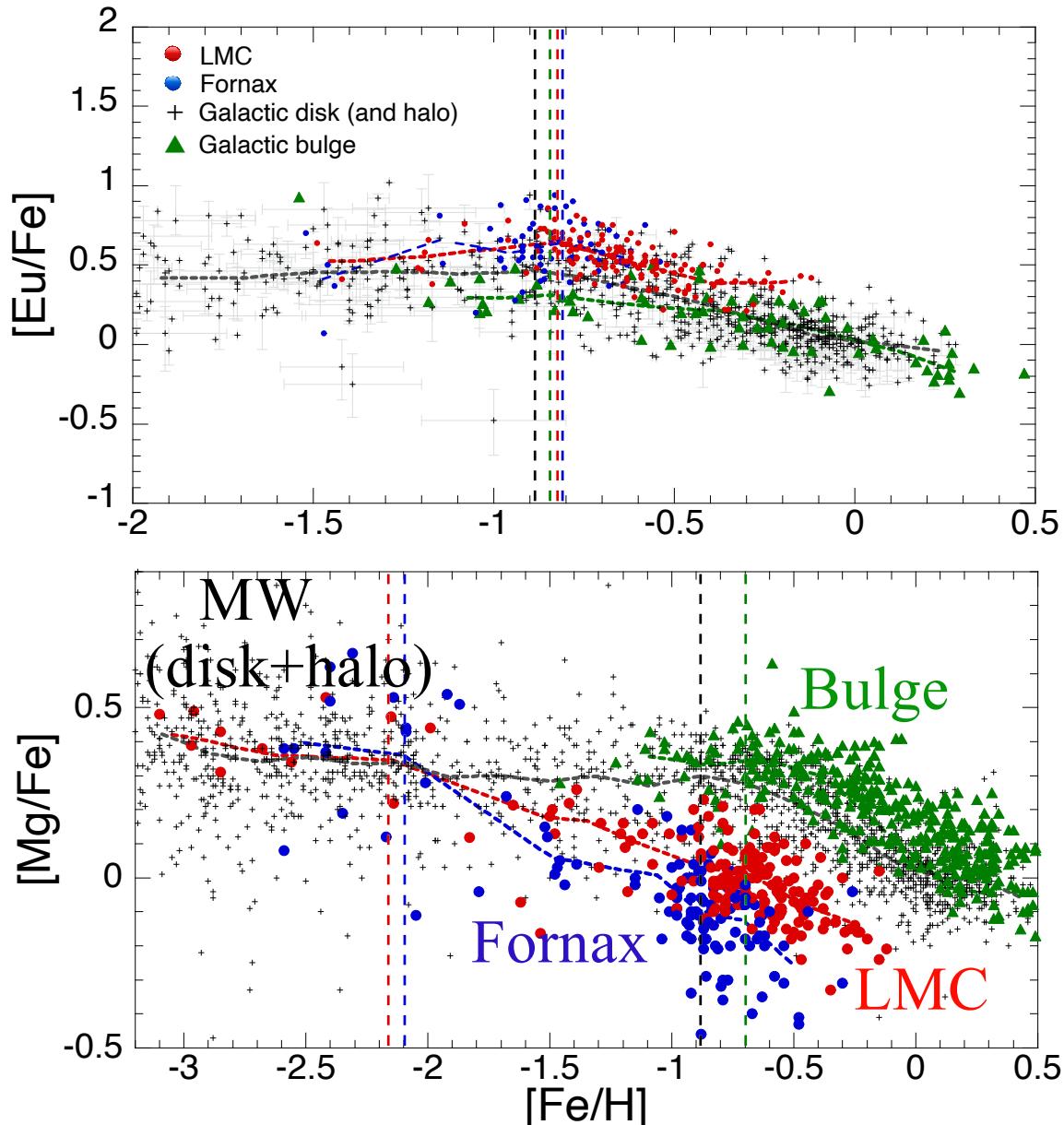
- ✓ r -process nucleosynthesis
(e.g., Winteler+12, Nishimura+15)
- ✓ associated with superluminous SNe (?)

2. Collapsars

- torus around BH
- ✓ Powered by energy from the rotating BH
(MacFadyen & Woosley 99)
- ✓ r -process nucleosynthesis
(Siegel+ 18)
- ✓ associated with long GRBs (?)



What causes the [Eu/Fe]-knee feature??



The [Eu/Fe]-knee locus
solely depends on the
metallicity ($[\text{Fe}/\text{H}] \sim -0.7$)



The [Eu/Fe] knee was not
caused by an onset of the
Fe release from SNe Ia

Instead,

“A metallicity threshold
beyond which r -process
SNe cease to emerge”
is implied

If r-process SNe = collapsars,

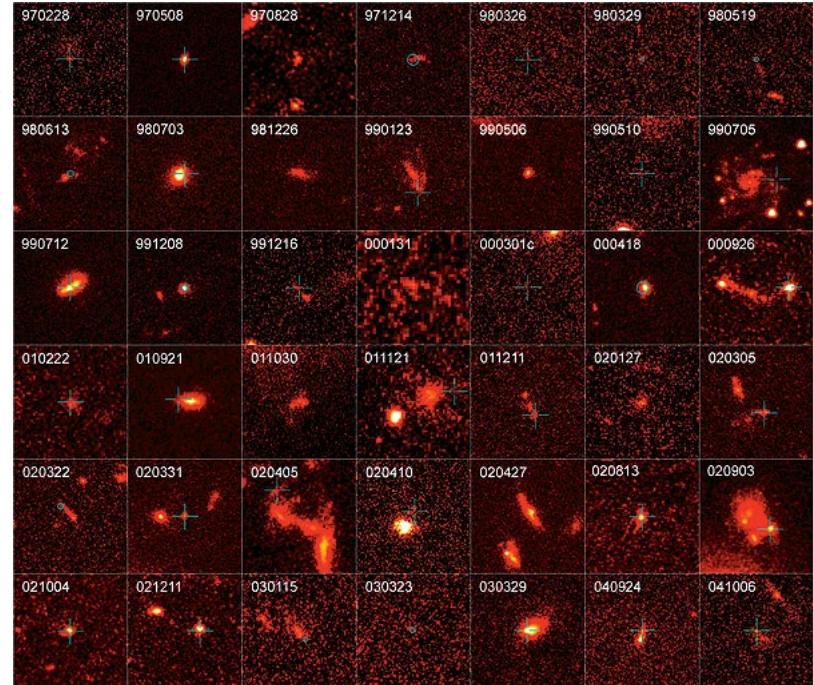
✓ Long GRBs (\approx Collapsars) favor a low-metallicity environment

host galaxies = faint, irregular galaxies
“low-metallicity galaxies”

$$Z < 1/3 Z_{\odot}$$

✓ Theoretically, a metallicity threshold should exist that retains enough angular momentum as $Z < 0.3 Z_{\odot}$

(Woosley & Heger 2006)

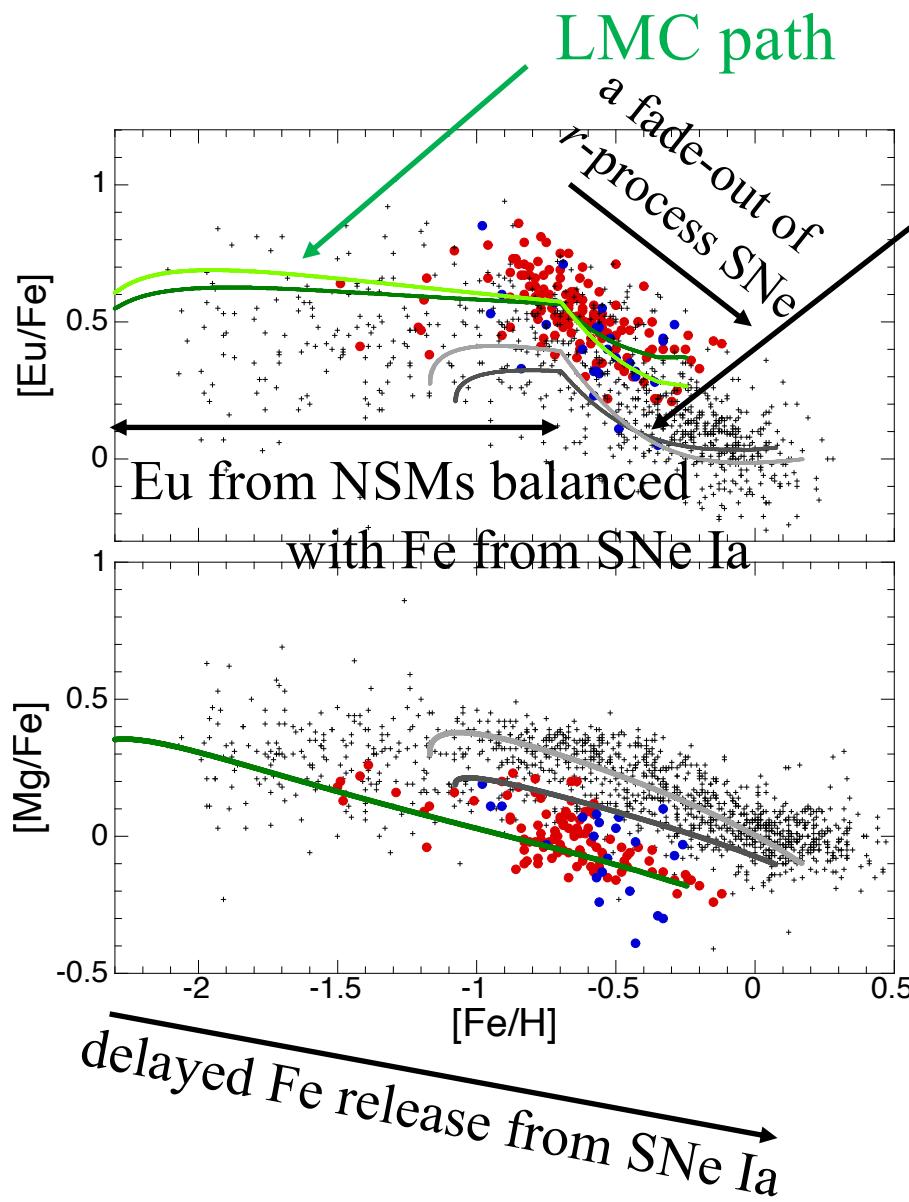


(Fruchter+ 2006)

If r-process SNe = magneto-rotational SNe,

Superluminous SNe, possibly identified with magnetorotational SNe, emerge in low-metallicity galaxies (e.g., Lunnan+ 2014).

Modeling of chemical evolution for MW and LMC



MW path

$[\text{Eu}/\text{Mg}]$

(TT 2024)

$[\text{Fe}/\text{H}]$

delayed r -process enrichment by NSMs

a fade-out of r -process SNe

LMC path

a fade-out of
 r -process SNe

Eu from NSMs balanced
with Fe from SNe Ia

delayed Fe release from SNe Ia