



INAF
leads Spoke 3
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member of Spoke 1
member of Spoke 10

INAF

OA—Trieste
OA—Bologna
I. Radioastronomy (BO)
OA—Roma
OA—Catania

STAFF — 33 Months

2 Full
2 Associate
5 Researcher

**P
L
A
N**

WP 1

Numerical simulations
Nbody - fluids

WP 2

Multiparameter
Optimisation

WP 3

Machine Learning
catalogs
features

WP 4

raw data level
(radio, pixel)
objects,
components

**1 PhD
UniTS**

45% south

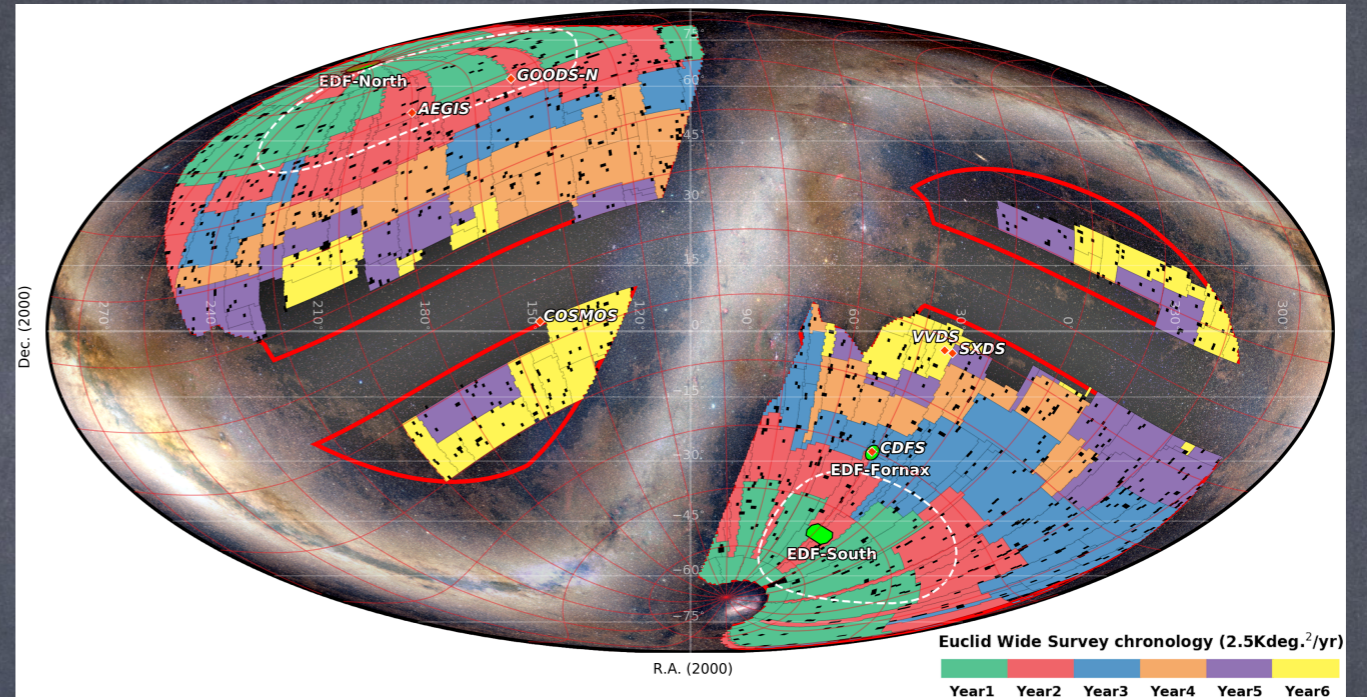
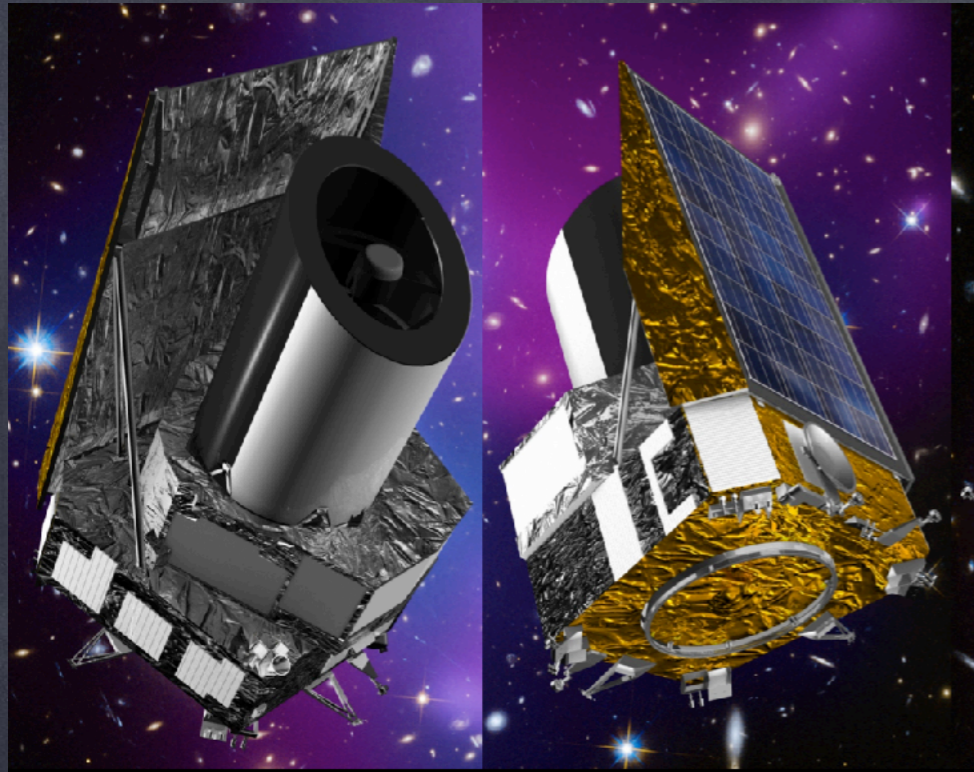
**1 TD (2y)
OA Napoli**

**1 TD (2y)
OA Catania**

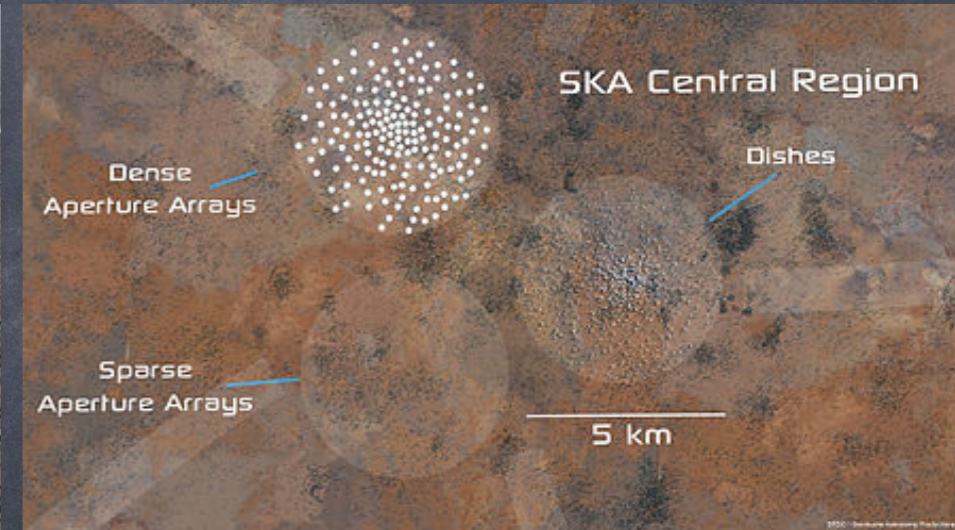
**1 TD (2y)
OA Catania**

Plenty of HUGE data and problems. Examples: Euclid, SKA

R. Scaramella et al.: The *Euclid* Wide Survey



LOFAR



Goals:

- find if some of actual methods/problems can benefit from QC (select some toy models; examine & discard; test most promising with emulators)
- think/investigate NEW ways of approaching old problems

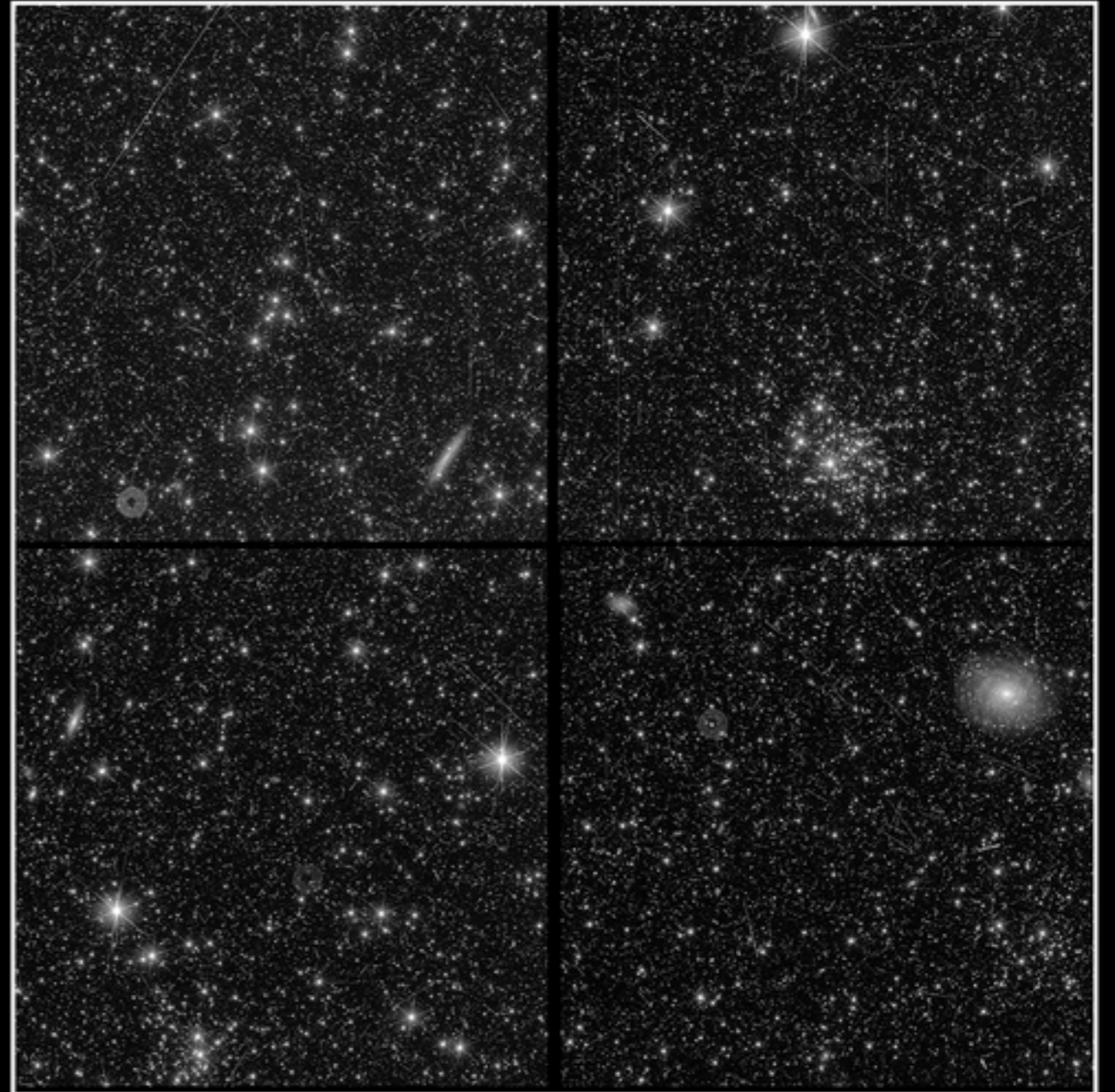
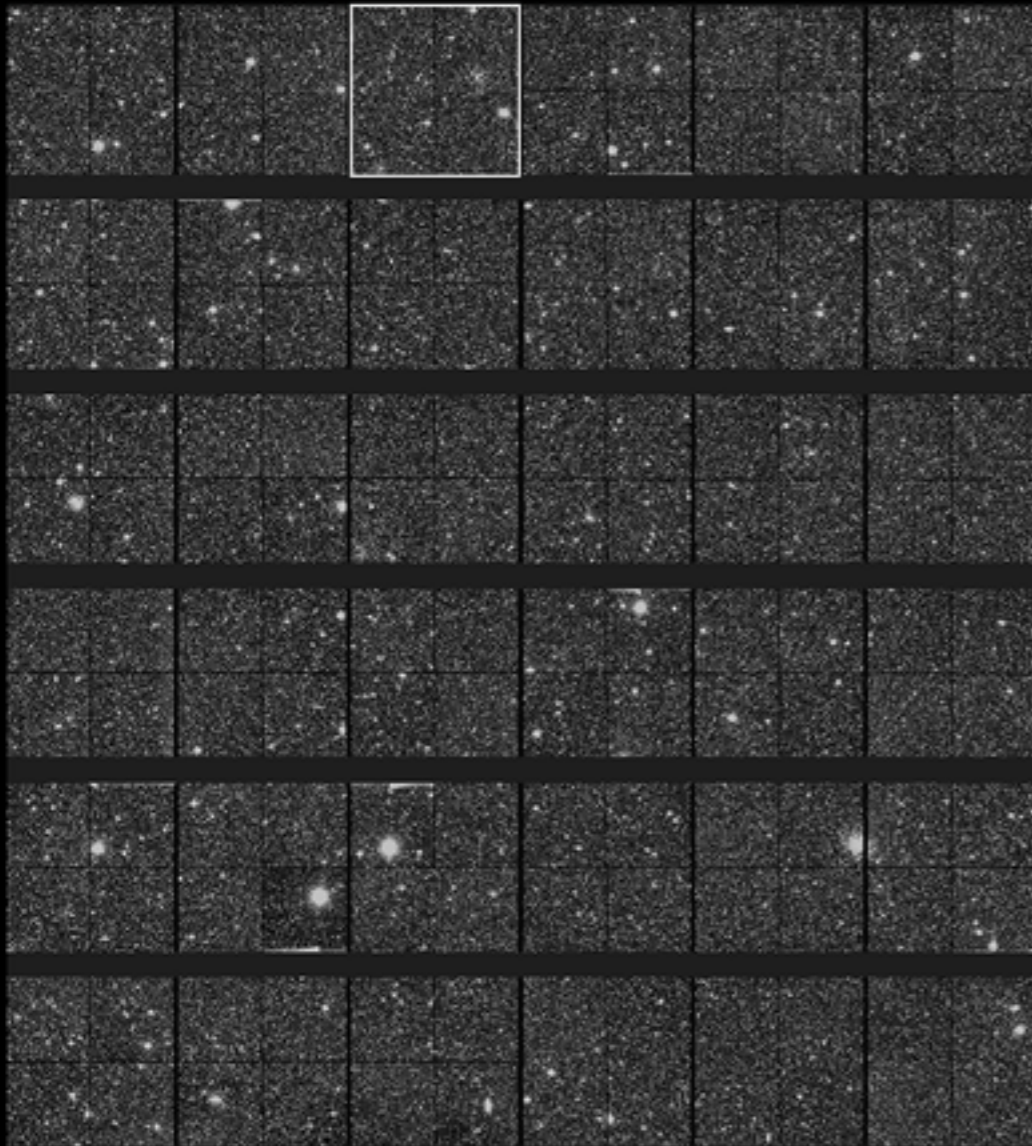


 esa

Euclid Mission (Dark Energy, Dark Matter and billions of galaxies: satellite launched July 1st 2023

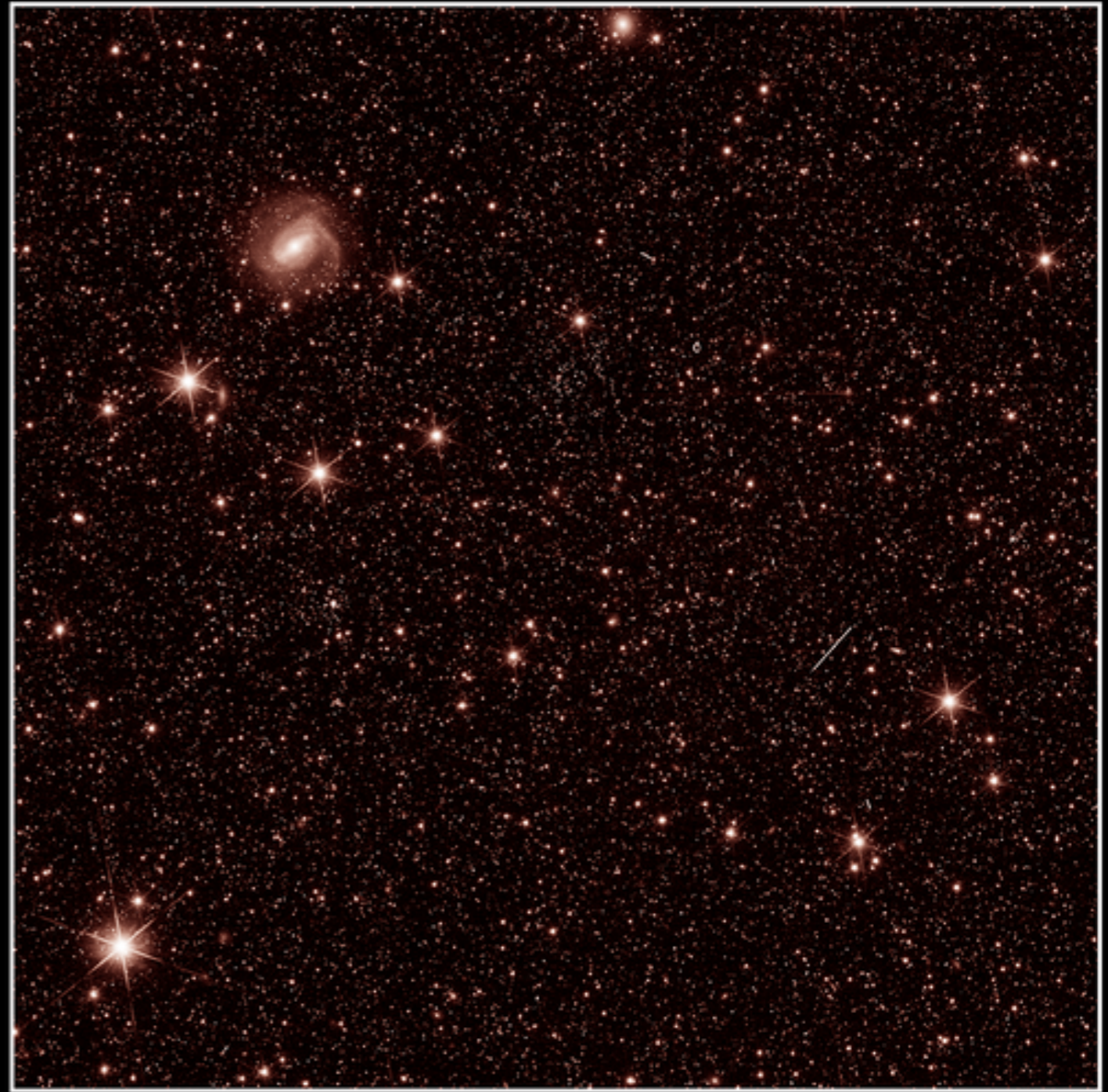
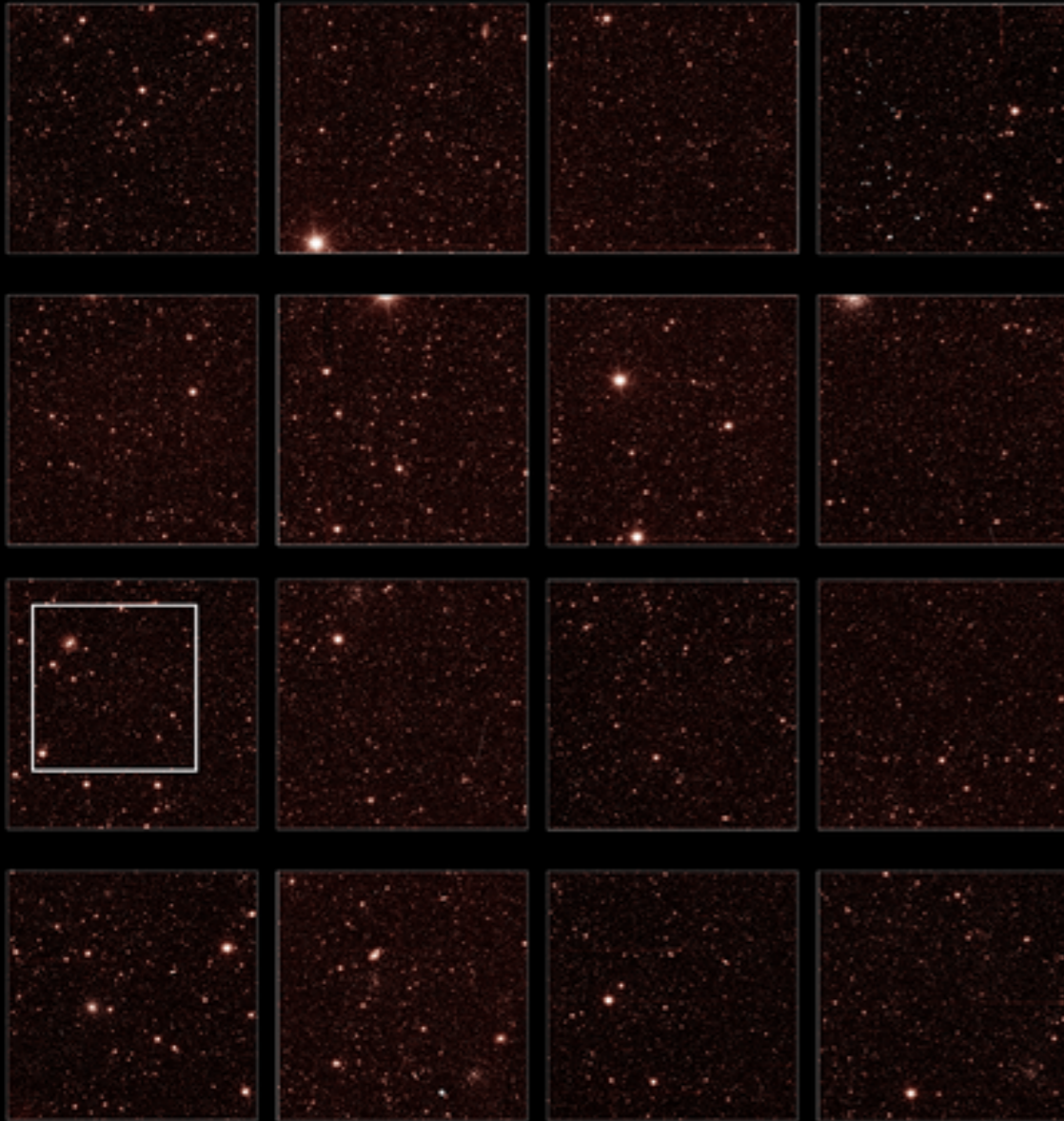
Cost ~1.5 G€, ~150 Science Institutes, ~1500 scientists

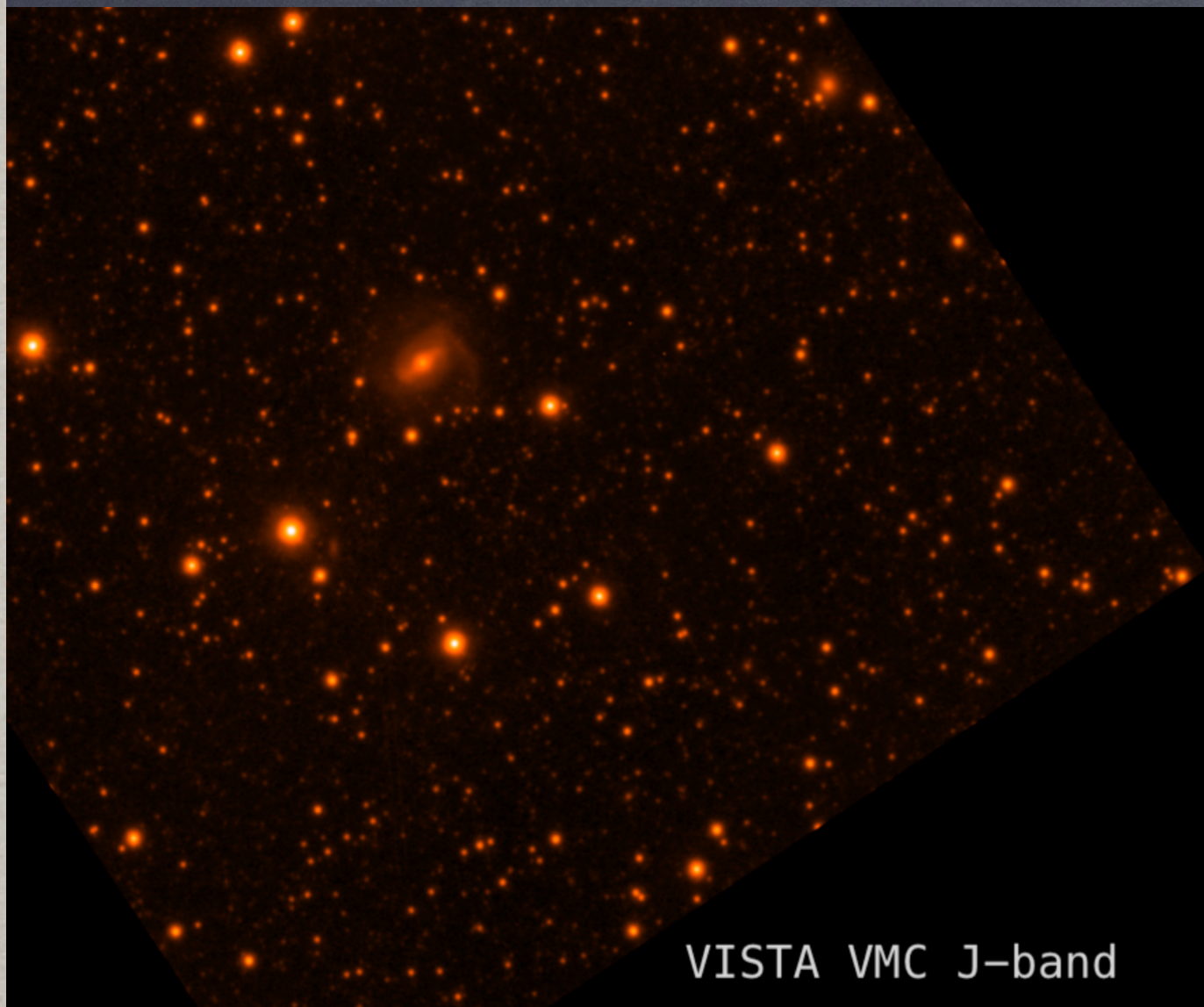
EARLY COMMISSIONING TEST IMAGE, VIS INSTRUMENT



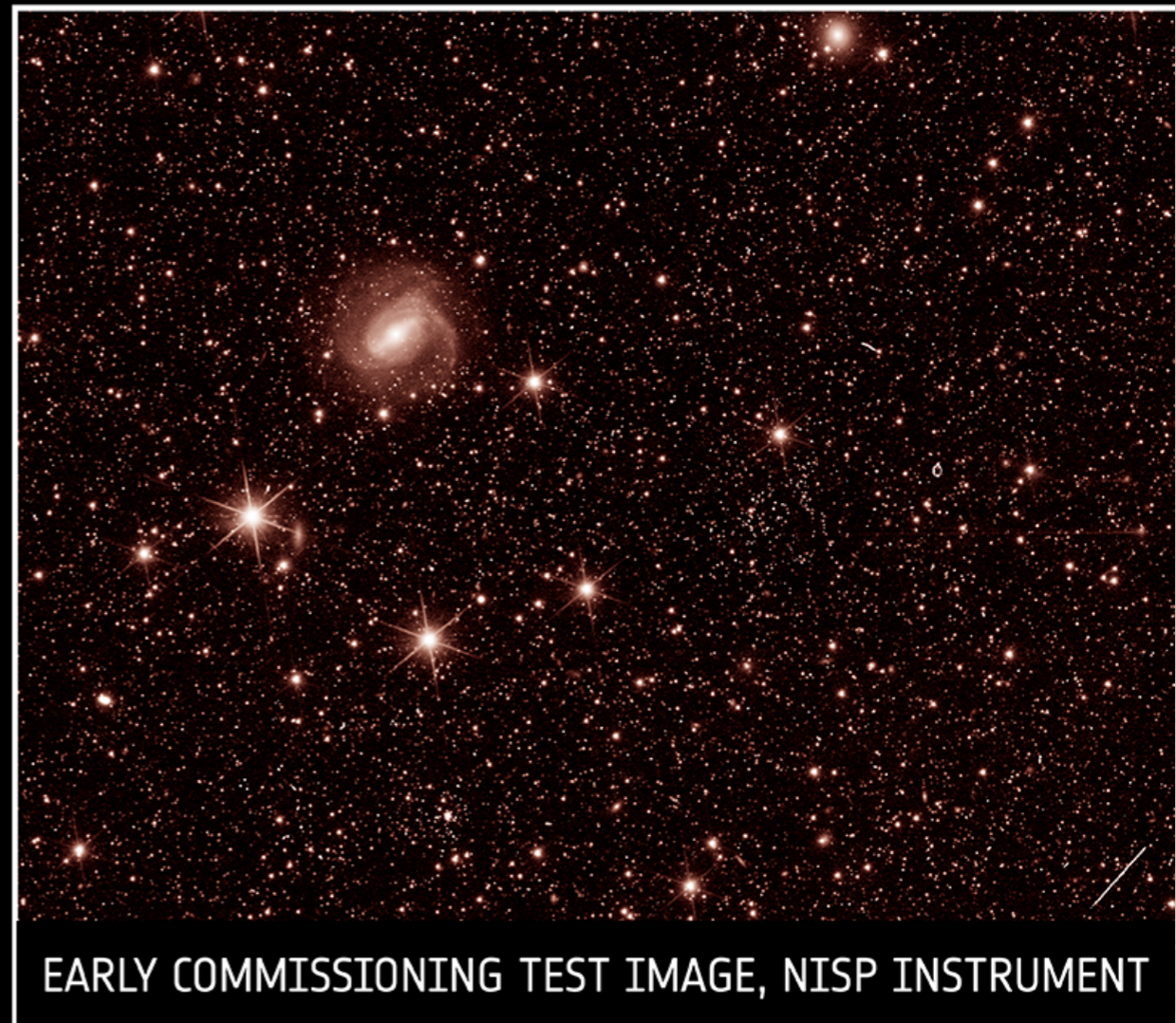
Euclid Mission: wide survey started in 2024, end 2030

EARLY COMMISSIONING TEST IMAGE, NISP INSTRUMENT





VISTA VMC J-band



EARLY COMMISSIONING TEST IMAGE, NISP INSTRUMENT

- ~ 1/3 of the sky at resolution 0.1"/pix**
- ~ 6E6 large images from two instruments**
- Several billions objects to study**

Food for thought

Superposition

Bell inequalities

Hybrid algorithms

Probability

Quantum entanglement

Einstein-[Podolski]-Rosen

Non locality

THE MAP OF QUANTUM COMPUTING

CLASSICAL COMPUTERS

1 STATE AT A TIME

BITS ARE INDEPENDENT OF EACH OTHER

CLASSICAL VS. QUANTUM

QUANTUM COMPUTERS

SUPERPOSITION
ENTANGLEMENT
INTERFERENCE

MANY STATES AT A TIME

QUBITS ARE IN A COMBINED STATE TOGETHER

SUPERPOSITION

MEASUREMENT

50% 0
50% 1

ENTANGLEMENT

PROBABILITY DISTRIBUTION

| | | | |
|-----|-----|-----|-----|
| 00 | 01 | 10 | 11 |
| 25% | 10% | 40% | 25% |

ONE MAIN MODEL OF CLASSICAL COMPUTING

NOT GATE: 0 → 1, 1 → 0

NOR GATES: A, B

GATE MODEL

OR CIRCUIT MODEL QUANTUM COMPUTING

QUBITS ENTANGLED WITH EACH OTHER

ALGORITHM IS A SERIES OF GATE OPERATIONS

SINGLE QUBIT GATES: H, X, Z

TWO-QUBIT GATES: CNOT

MEASUREMENT

WHAT TO BUILD QUBITS FROM?

FUNDAMENTAL PARTICLES?

ATOMS? ELECTRONS? PHOTONS?

NEED A 2-LEVEL QUANTUM SYSTEM

TO ENCODE THE 2 BINARY STATES

BULK QUANTUM SYSTEMS?

MEASUREMENT BASED (OR ONE-WAY) QUANTUM COMPUTING

MODELS OF QUANTUM COMPUTING

ADIABATIC QUANTUM COMPUTING

MINIMUM ENERGY STATE IS THE ANSWER TO YOUR PROBLEM

ENERGY LANDSCAPE

CORRECT ANSWER

EQUIVALENT TO GATE MODEL

TOPOLOGICAL QUANTUM COMPUTING

MOST THEORETICAL MODEL

MAJORANA ZERO-MODE QUASI-PARTICLE

NON-ABELIAN ANYON

QUASI-PARTICLE EXAMPLE: ELECTRON HOLE

A HOLE IN A SEA OF ELECTRONS HAS PARTICLE-LIKE PROPERTIES

QUANTUM ANNEALING

ALSO ENERGY MINIMISATION BUT NOT UNIVERSAL

OBSTACLES

DECOHERENCE: WANT THEM TO ENTANGLE TO EACH OTHER BUT NOT TO THE ENVIRONMENT

SCALABILITY: DOES YOUR DESIGN SCALE TO LARGE NUMBERS OF QUBITS? MASSIVE ENGINEERING CHALLENGE!

NOISE: COSMIC RAYS, RADIATION, HEAT, PARTICLES

CONTROL WIRE, CONNECTING WIRES (TO OTHER QUBITS), READOUT WIRE

QUANTUM ERROR CORRECTION

FAULT TOLERANT QUANTUM COMPUTERS

MANY NOISY QUBITS → 'PERFECT' QUBIT

PHYSICAL QUBITS → LOGICAL QUBIT

HOW MANY YOU NEED DEPENDS ON QUALITY OF QUBITS

(ESTIMATE) ~100 TO 1000 QUBITS FOR 1 LOGICAL QUBIT

OBSTACLES

DECOHERENCE

WANT THEM TO ENTANGLE TO EACH OTHER BUT NOT TO THE ENVIRONMENT

SCALABILITY

CONTROL WIRE, CONNECTING WIRES (TO OTHER QUBITS), READOUT WIRE

DOES YOUR DESIGN SCALE TO LARGE NUMBERS OF QUBITS? MASSIVE ENGINEERING CHALLENGE!

INTERFERENCE

QUBIT REALLY DESCRIBED BY QUANTUM WAVEFUNCTION

CONSTRUCTIVE INTERFERENCE

DESTRUCTIVE INTERFERENCE

QUBITS → QUANTUM WAVEFUNCTIONS → OVERALL WAVEFUNCTION → PROBABILITY DISTRIBUTION

ON MEASUREMENT GET ONE ANSWER OUT

| NUMBER OF QUBITS | NUMBER OF STATES |
|------------------|------------------|
| 1 | 2 |
| 2 | 4 |
| 3 | 8 |
| 4 | 16 |
| 5 | 32 |
| ⋮ | ⋮ |
| N | 2 ^N |

POTENTIAL APPLICATIONS OF QUANTUM COMPUTERS

QUANTUM SIMULATION

RAPIDLY PROTOTYPE MANY DIFFERENT MATERIALS

IS FASTER THAN PHYSICALLY MAKING AND TESTING THEM

CHEMICAL REACTIONS

WANT TO SIMULATE LARGE QUANTUM SYSTEMS ON A QUANTUM COMPUTER

ELECTRONIC PROPERTIES

IMPROVING SOLAR PANELS

SIMULATING AS FEW AS 30 PARTICLES ON A SUPERCOMPUTER IS DIFFICULT

BETTER CATALYST FOR FERTILIZER PRODUCTION

FeMoCo CATALYST

Fe, Mo, S, C

IMPROVING BATTERIES

CURRENTLY 2% OF GLOBAL CO2 EMISSIONS

DRUG DEVELOPMENT

MATERIALS FOR AEROSPACE

NEW CHEMICALS

PHYSICAL REALISATIONS

SUPERCONDUCTING QUANTUM COMPUTERS

TRANSMON QUBIT

2-LEVEL SYSTEM

FREQUENCY OF CHARGE OSCILLATION

COOPER PAIR

FLUX QUBIT

OR MAGNETIC FLUX OR SUPERCONDUCTING PHASE

QUANTUM DOT QUANTUM COMPUTERS

ALSO SILICON SPIN QUANTUM COMPUTERS

QUANTUM DOT ELECTRONS

2-LEVEL SYSTEM

SPIN OR CHARGE

CONTROL WITH MICROWAVES OR VOLTAGES OR MAGNETIC FIELDS

MADE OF: SILICON, GALLIUM ARSENIDE, SILICON CARBIDE OR DIAMOND

LINEAR OPTICAL QUANTUM COMPUTERS

LINEAR OPTICAL ELEMENTS: MIRRORS, WAVEPLATES, INTERFEROMETERS

INTEGRATED PHOTONICS CHIPS

CONTROLLED WITH VOLTAGES

2-LEVEL SYSTEM

OR

1 0

NUMBER OF PHOTONS

TRAPPED ION QUANTUM COMPUTERS

CONTROL WITH MICROWAVES OR LASERS

IONISED ATOMS TRAPPED IN MAGNETIC FIELDS

2-LEVEL SYSTEM

1 0

ENERGY LEVELS OF THE ATOM

COLOUR CENTRE QUANTUM COMPUTERS

2-LEVEL SYSTEM

SPIN OF THE ATOM

CONTROL WITH MICROWAVES OR LASERS

NEUTRAL ATOMS IN OPTICAL LATTICES

CONTROL WITH LASERS

TRAPPED ATOMS

COOLED TO MILLIONTHS OF A KELVIN

OPTICAL LATTICE

CAN ALSO BE USED TO MAKE PURE QUANTUM SIMULATORS

A 10 THOUSAND ATOM QUANTUM SIMULATOR HAS BEEN MADE

OTHER APPROACHES

ELECTRON-ON-HELIUM QUBIT

CAVITY QUANTUM ELECTRODYNAMICS

MAGNETIC MOLECULE

NUCLEAR MAGNETIC RESONANCE

MOLECULAR SPINS

QUANTUM ALGORITHMS

MULTIPLICATION: 7177 x 3001 → 21538177 (EASY! EFFICIENT CLASSICAL ALGORITHM)

FACTORISATION: 21538177 → 7177 x 3001 (HARD! NO EFFICIENT CLASSICAL ALGORITHM)

USED FOR ENCRYPTION

SHOR'S ALGORITHM: 21538177 → 7177 3001 (EFFICIENT QUANTUM ALGORITHM)

COMPLEXITY THEORY

HOW MUCH HARDER IS IT TO SOLVE THE PROBLEM AS THE PROBLEM GETS LARGER?

NP-COMplete: TRAVELING SALESMAN, MAP COLOURING

NP: GRAPH ISOMORPHISM

BQP: INTEGER FACTORIZATION, DISCRETE LOGARITHM

P: TESTING IF PRIME, MULTIPLICATION

EFFICIENT FOR A QUANTUM COMPUTER

EFFICIENT FOR A CLASSICAL COMPUTER

CLASSICAL COMPUTERS

ARE VERY VERSATILE DEVICES

PERSHAPS AN EFFICIENT CLASSICAL ALGORITHM WILL BE FOUND UNLIKELY BUT NOT RULED OUT

NON-COMPUTABILITY

COMPUTATIONALLY EQUIVALENT

CAN SIMULATE A QUANTUM COMPUTER ON A CLASSICAL COMPUTER

NEITHER CLASSICAL OR QUANTUM COMPUTERS CAN SOLVE NON-COMPUTABLE PROBLEMS E.G. THE HALTING PROBLEM

COMPUTATIONALLY EQUIVALENT

BUT EXPONENTIALLY DIFFICULT!

GROVER'S ALGORITHM

PROBLEM: FIND THE NUMBER 42

02 16 99 28 42 73 01

QUADRATIC SPEEDUP OVER CLASSICAL

DON'T HAVE TO WORRY ABOUT YOUR BANK ACCOUNT YET

BUT NEED UPWARDS OF 1 MILLION QUBITS!

1,000,000

NOT THERE YET OUR COMPUTERS ARE SAFE FOR NOW

QUANTUM COMPLEXITY THEORY

PROBLEM: FACTORISE A NUMBER WITH N DIGITS

N=8

21538177

SHOR'S ALGORITHM IS POLYNOMIAL log(N)

HOW MUCH HARDER IS IT TO FACTORISE A NUMBER WHERE N=9?

ANYTHING WITH THE N IN THE EXPONENT IS HARD

SCALING OF FACTORISATION IS = 2^{N/2}

BEST CLASSICAL ALGORITHM IS EXPONENTIAL (IF YOU HAVE A WORKING QUANTUM COMPUTER)

OPTIMIZATION PROBLEMS

MACHINE LEARNING AND AI

CLIMATE CHANGE

WEATHER FORECASTING

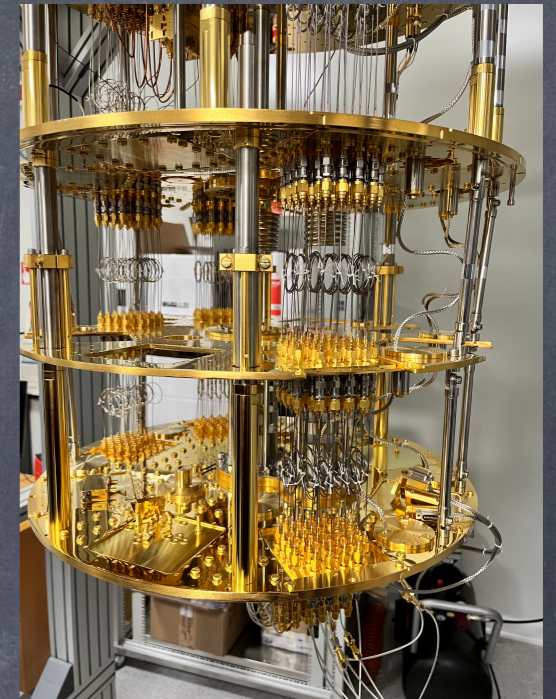
FINANCIAL MODELING

CYBERSECURITY

Quantum computing is..

NOT a general purpose super duper computer

Undergoing **FAST** hardware evolution
(but bottlenecks, errors)



In principle allows a **QUANTUM LEAP** in selected problems (optimisation, factorisation etc etc)

For the time being is **FUN!**

Cosmological numerical simulations (Schrödinger -Poisson & Vlasov-Poisson equation) with Quantum Computers

Schrödinger-Poisson:

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + mV\psi,$$

$$\rho = |\psi|^2$$

$$\nabla^2 V = 4\pi G(\rho - \bar{\rho}),$$

Nonlocal quantum pressure:

$$p_Q = -\left(\frac{\hbar}{2m}\right)^2 \rho \nabla \otimes \nabla \ln \rho.$$

Becomes Euler-Poisson
when m tends to infinity

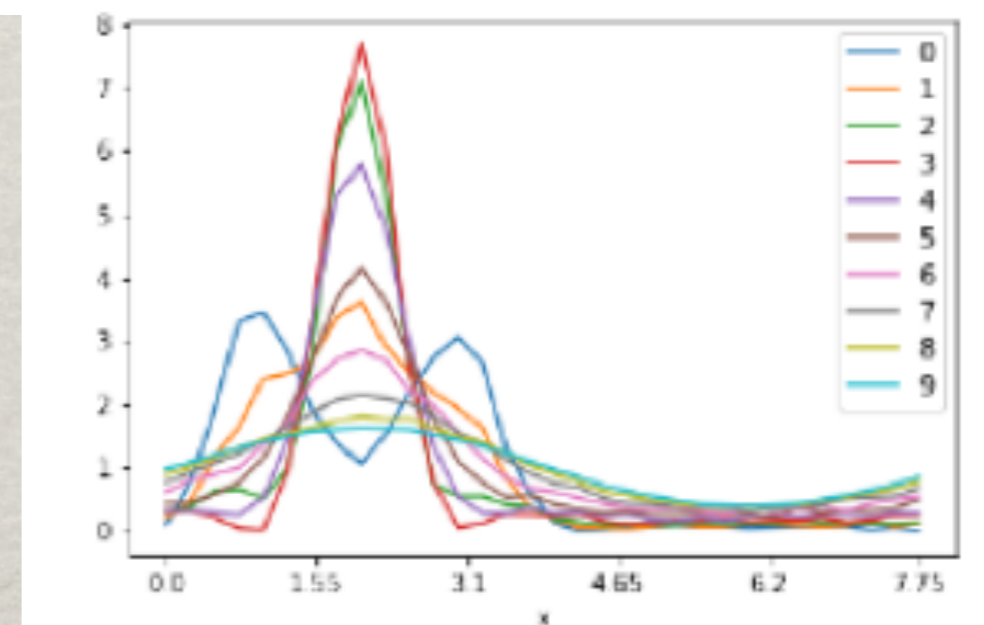
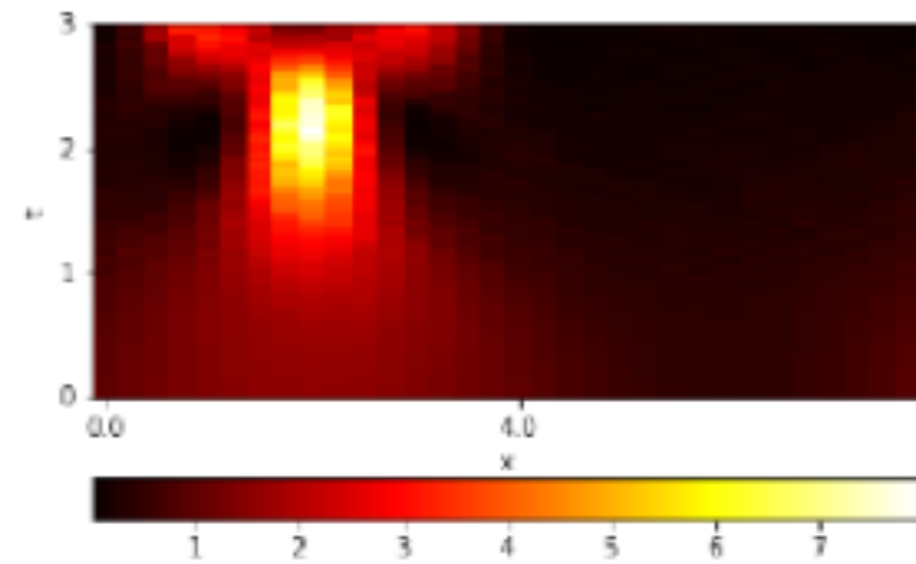


Figure 6.9: Exact potential simulation with a total number of ansatz parameters $M = 60$. $N_t = 2 \cdot 10^4$, $r_c = 10^{-7}$, No regularization.

- **One dimensional**
- Following an idea by Mocz & Szasz, 2021, ApJ, 910, 29
- **Variational algorithm completely rewritten**

**Luca Cappelli won the PhD position at Trieste University
IBM Zurich collaborates with the PhD project**

Research lines for WP1

Quantum Technicality

- Study scaling properties of the QC algorithm
- Reduce circuit depth of the variational algorithm
- Implement the algorithm on a real quantum device
- Study the feasibility of 3D simulations
- Study the quantum advantage when m is large

Physics

- Application: fuzzy dark matter
- Comparison of Schroedinger-Poisson with Vlasov-Poisson when the field mass m varies
- Study the possibility to use SP as a proxy for VP, m acting as *the softening* in a N-body simulation
- Study if a similar variational algorithm can be applied to hydrodynamics

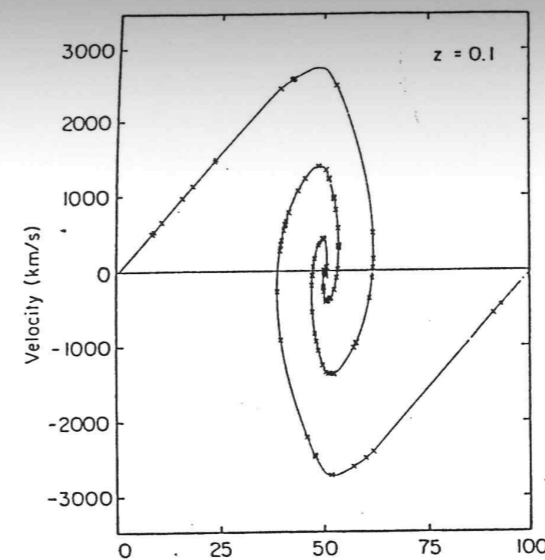


Fig. (4.13a) $z=0.1, a=0.9$

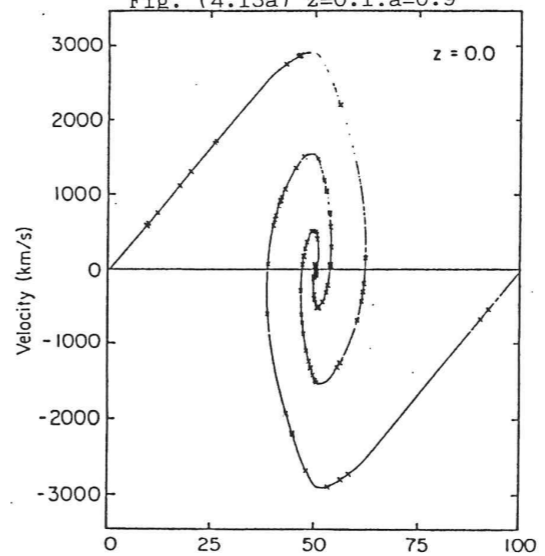


Fig. (4.14a) $z=0.0, a=1.0$

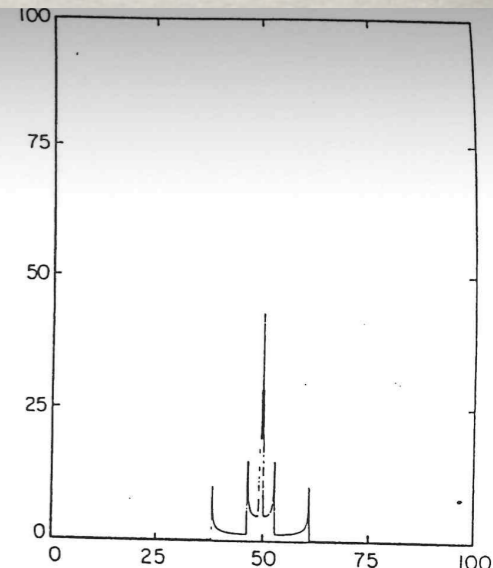


Fig. (4.13b) $z=0.1, a=0.9$

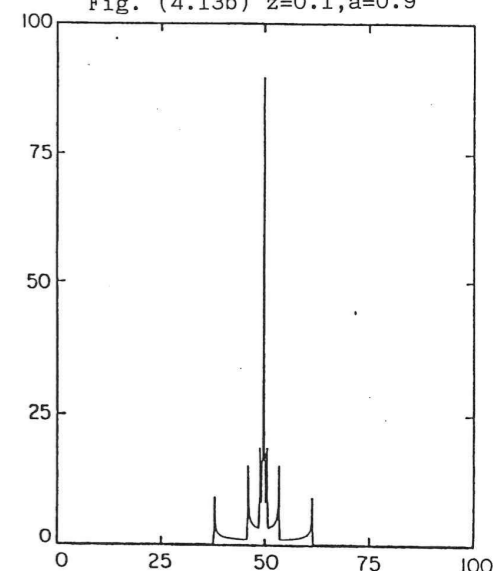


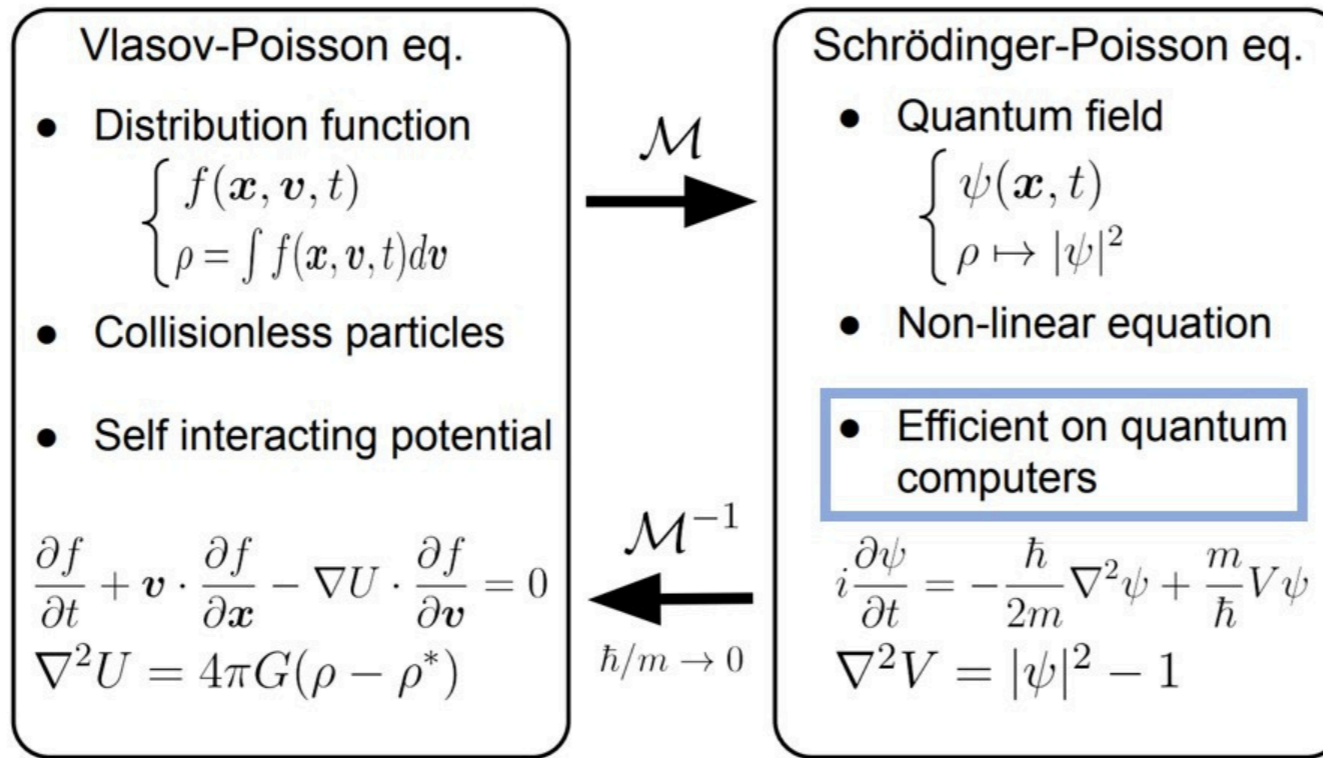
Fig. (4.14b) $z=0.1, a=1.0$

One-dimensional VP simulation. Left panel: phase space, v vs x . Right panel: density vs x

WP1

Schrodinger vs Vlasov - Poisson

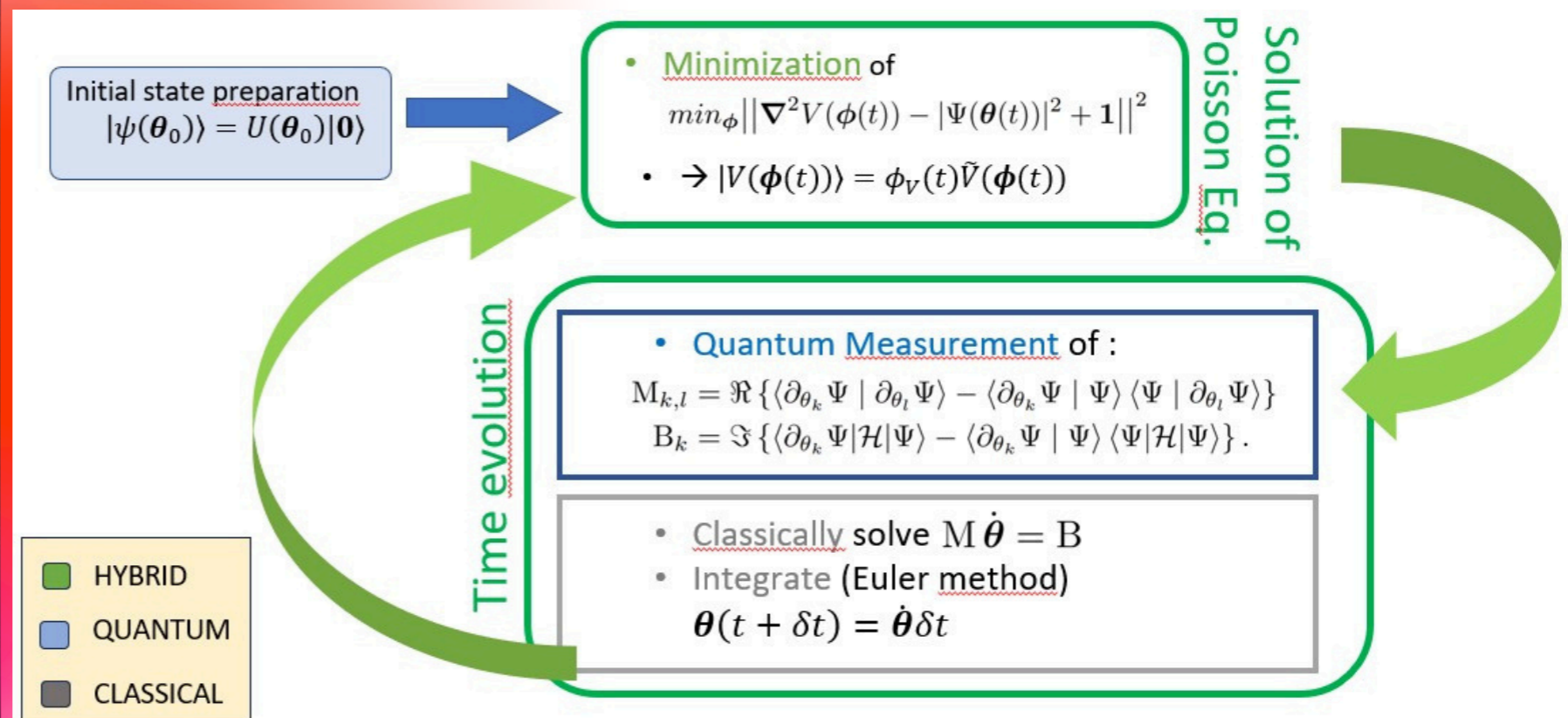
- Collisionless dark matter
- Vlasov - Poisson equation
- hard to solve numerically
- N - body simulations
- Schrodinger equation
- goes to VP for m to infinity



VARIATIONAL ALGORITHM FOR SOLVING SCHROEDINGER EQUATION

CAPPELLI ET AL., 2024, PHYSICAL REVIEW RESEARCH, 6, 013282

Hybrid - Quantum Variational Algorithm



Combinatorial Optimisation Problems

input models from a set S and m conditions defining a cost function to be minimised or maximised

examples:

- maximise a cosmological likelihood
- minimise the chisq of a lens model
- choose the specifics maximizing FoM

Quantum Approximation Optimization Algorithm

Quantum Annealing

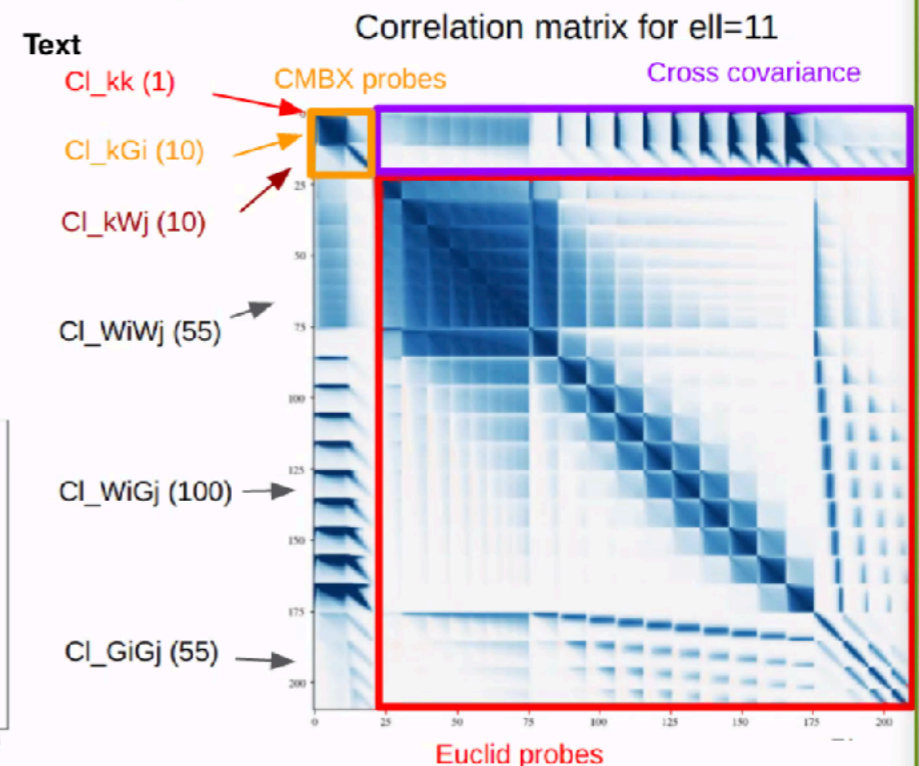
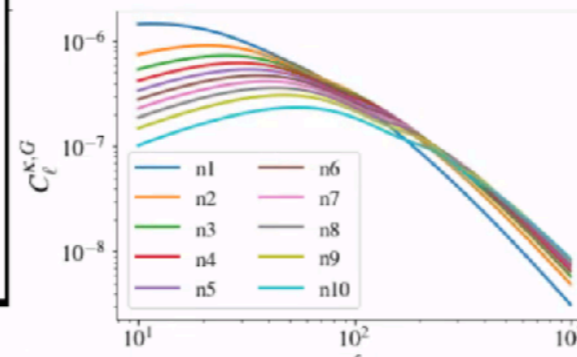
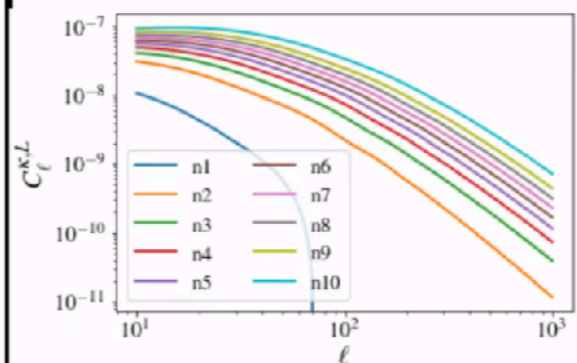
Quantum Evolutionary Algorithm

Joint analysis of Euclid photometric probes and CMB lensing

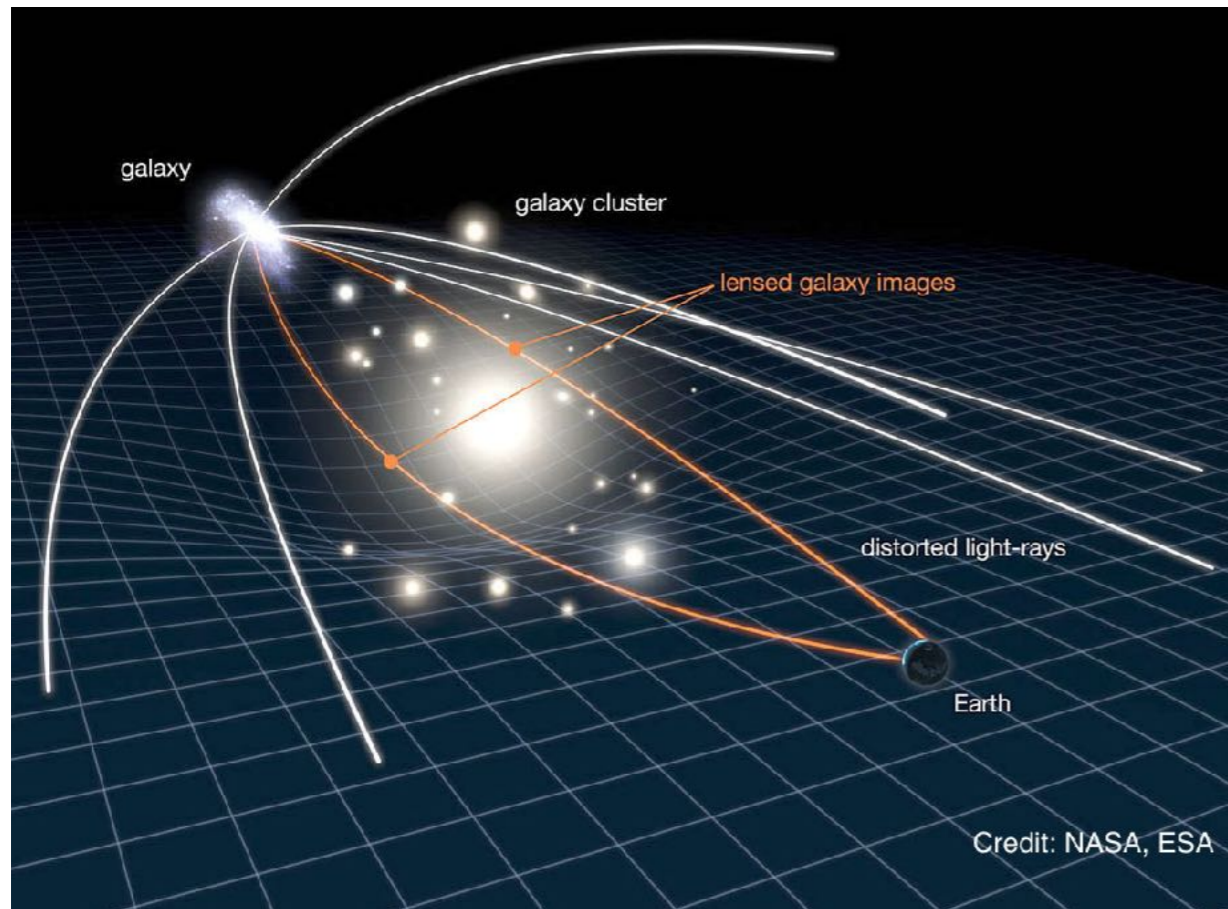
$$\vec{\Theta}_{XC}(\ell) = \{ \underset{\substack{\uparrow \\ \text{CMB lensing}}}{C_{\ell}^{\kappa_{\text{CMB}}, \kappa_{\text{CMB}}}}, C_{\ell}^{\kappa_{\text{CMB}}, \text{GCph}_i}, C_{\ell}^{\kappa_{\text{CMB}}, \text{WL}_i}, C_{\ell}^{\text{GCph}_i, \text{GCph}_j}, C_{\ell}^{\text{WL}_i, \text{WL}_j}, C_{\ell}^{\text{WL}_i, \text{GCph}_j} \}$$

Xcorr
Euclid 3x2pt data vector

- 6x2 points data vector: CMB-L, Euclid WL and GC phot, cross correlations
- Interface with CLOE (mirror git)
- Limber and Gaussian likelihood (following IST:L)
- Code development progressing well
- Next step is to sample the joint likelihood with mock data



GRAVITATIONAL LENSING



Gravitational lensing: mass-energy curving space time.

Deflection angle

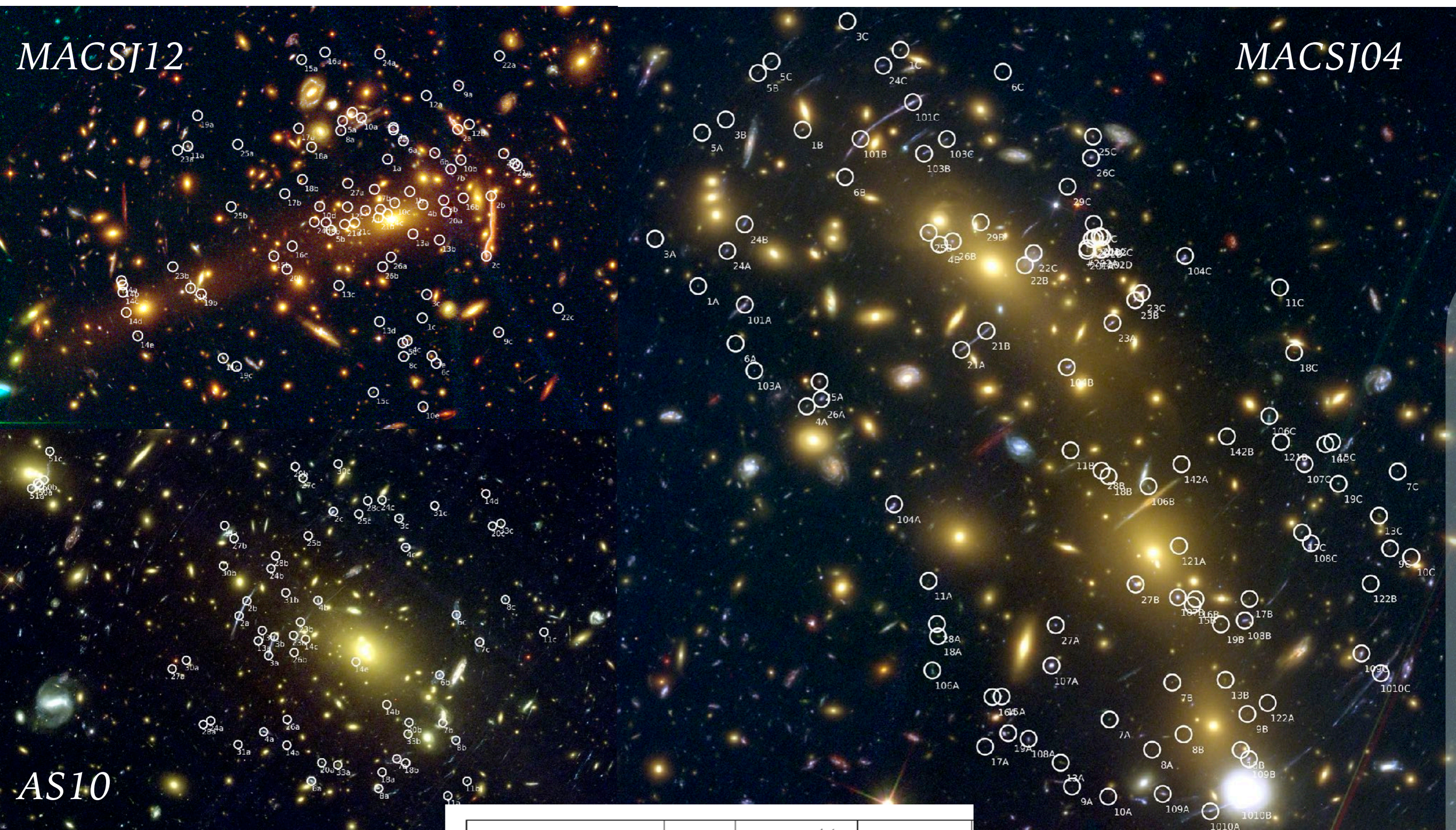
$$\vec{\alpha}(\vec{\theta}) \propto \nabla \int \Phi(\vec{\theta}, z) dz$$

Gravitational potential

Consequences:

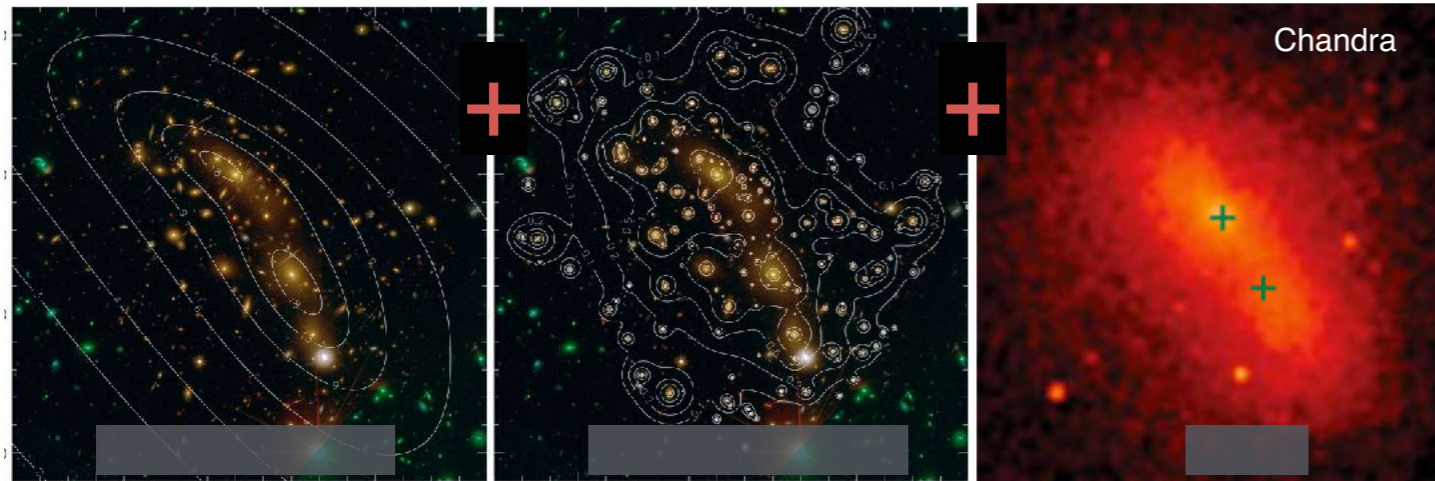
- *Multiple images* (strong lensing — SL)
- *Distortions:* gravitational arcs (SL), induced ellipticities (WL)
- *Magnifications*

All these effects can be used to recover the mass distribution of the lens (dark matter, gas, stars)



| Cluster | z | N_m^{meas} (N_m^{tot}) | N_{im} (N_{fam}) |
|-------------------|-------|------------------------------|------------------------|
| MACS J1206.2-0847 | 0.439 | 58 (258) | 82 (27) |
| MACS J0416.1-0403 | 0.396 | 49 (193) | 102 (37) |
| Abell S1063 | 0.348 | 37 (222) | 55 (20) |

EXAMPLE: SL MODEL OPTIMISATION



Total gravitational potential is the sum of each component:

$$\phi_{tot}(\vec{\xi}) = \sum_{i=1}^{N_h} \phi_i^{halo}(\vec{\xi}_{halo}) + \sum_{k=1}^{N_{gal}} \phi_k^{gal}(\vec{\xi}_{gal}) + \phi_{shear}(\vec{\xi}_{shear}) + \phi_{gas}$$

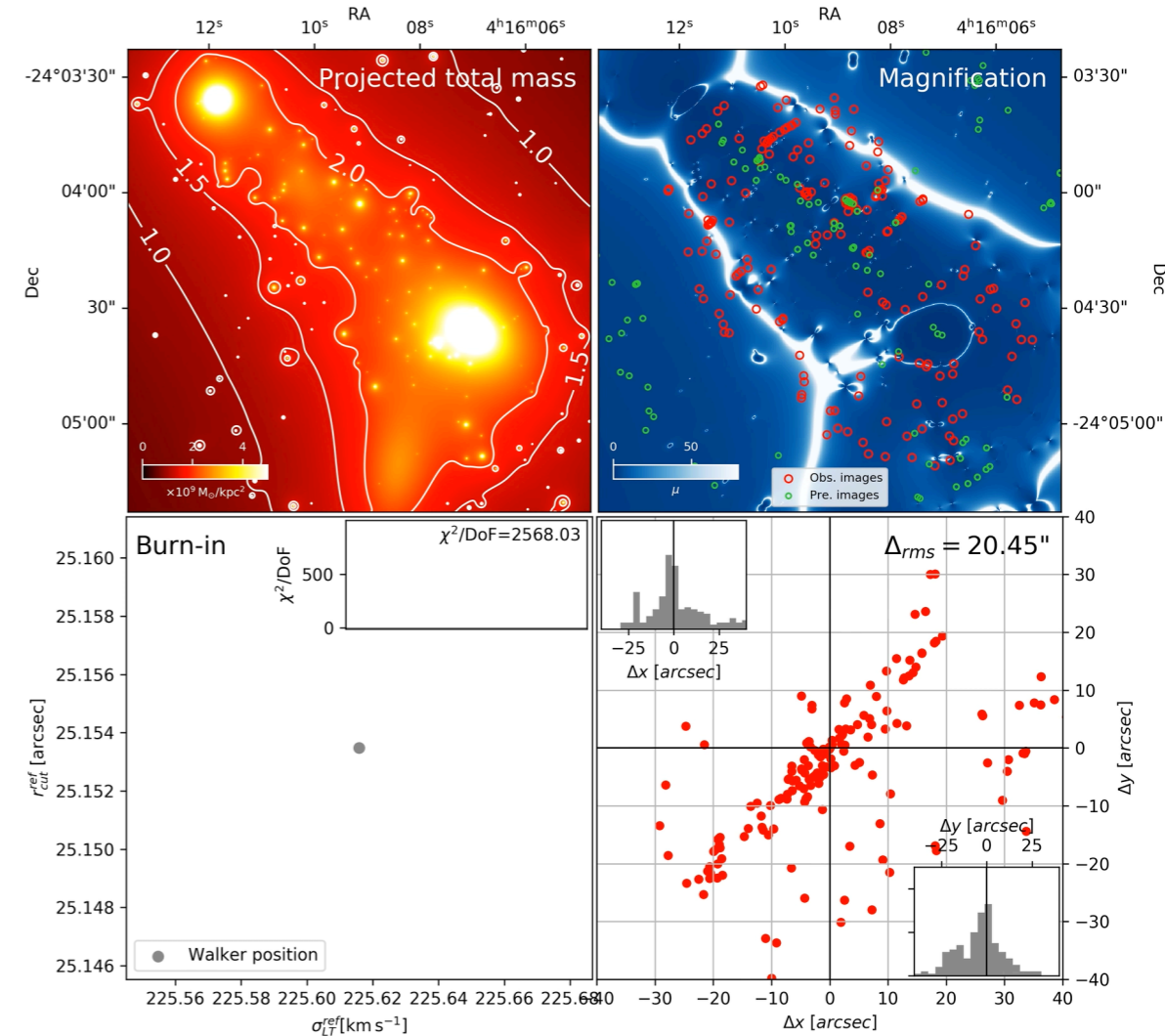
Each mass component is parametrised:

Pseudo (non-singular) Isothermal Ellipsoids:

$$\rho_{PIEMD}(r) = \frac{\sigma_0}{2\pi G(1 + r^2/r_{core}^2)}$$

Truncated-Isothermal Spheres:

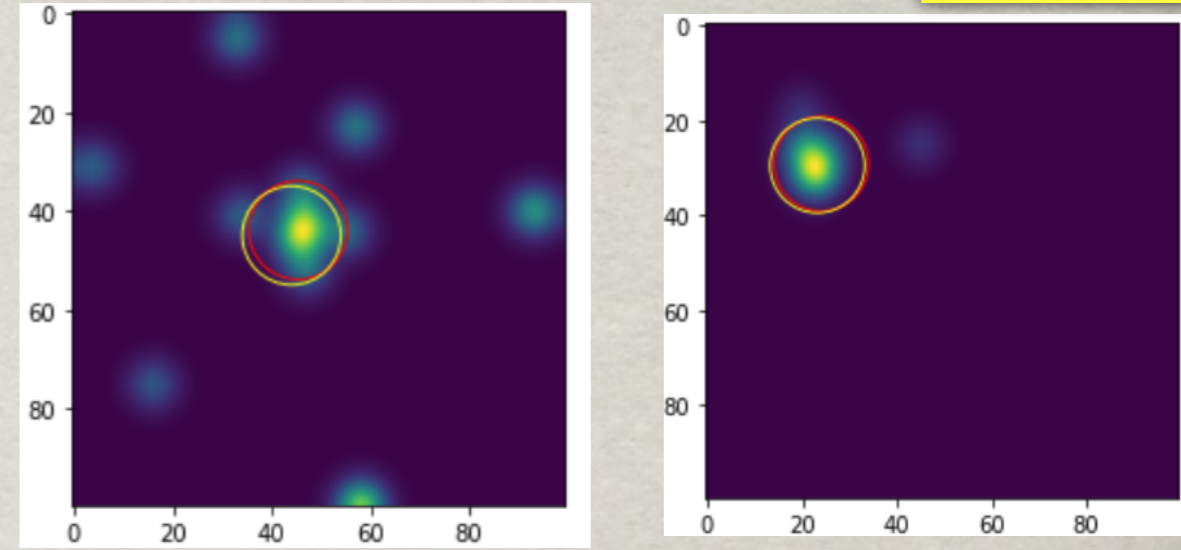
$$\rho_{sub-halo}(r) = \frac{\rho_0}{(1 + r^2/r_{core}^2)(1 + r^2/r_{cut}^2)}$$



Model optimisation

GRB detection localization in AGILE/GRID data

- We developed a new method for detecting and localizing GRB in the AGILE/GRID sky maps as a reaction to external science alerts.
- The science alerts can have error regions with different sizes depending on the instruments that detected the transient event. For this reason, we trained this method to detect GRBs in the AGILE sky maps located in a radius of 20 degrees from the map center; this radius is larger than 99.5 % of the error region present in the GRBWeb catalog.
- The method comprises two Deep Learning models implemented with two Convolutional Neural Networks. The first model detects if the sky map contains a GRB, and the second model localizes the GRB in the sky maps filtered from the first model.
- We trained and tested the models using simulated sky maps and GRBs. The detection model achieves an accuracy of 95.7 %, and the localization model has a mean error lower than 0.8 degrees.



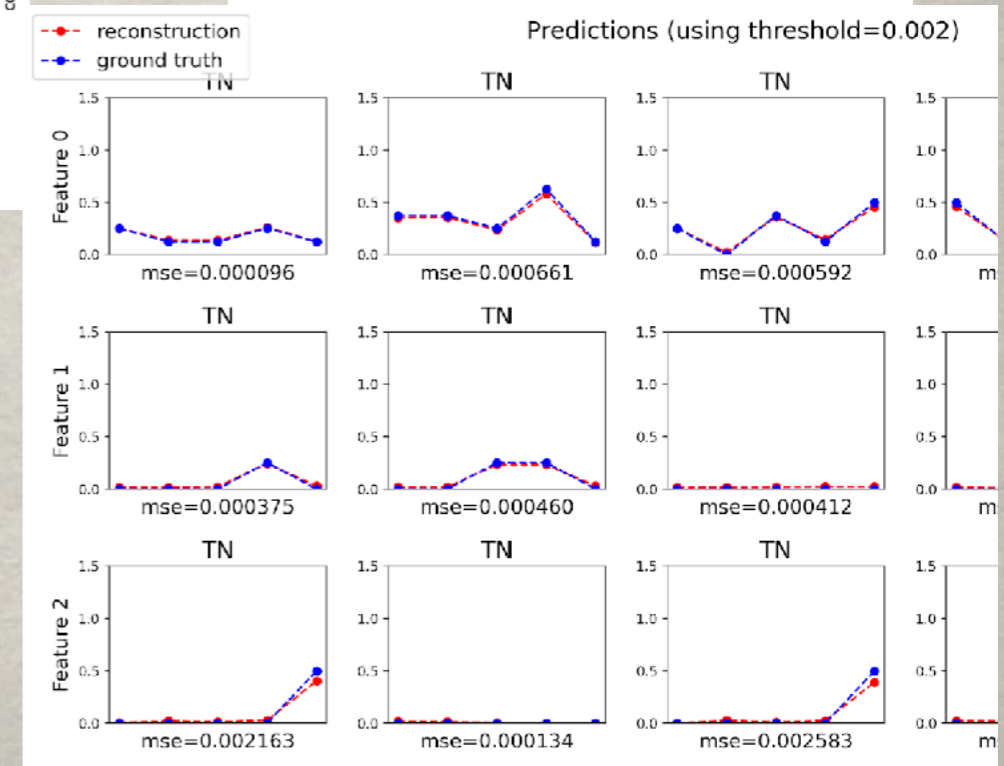
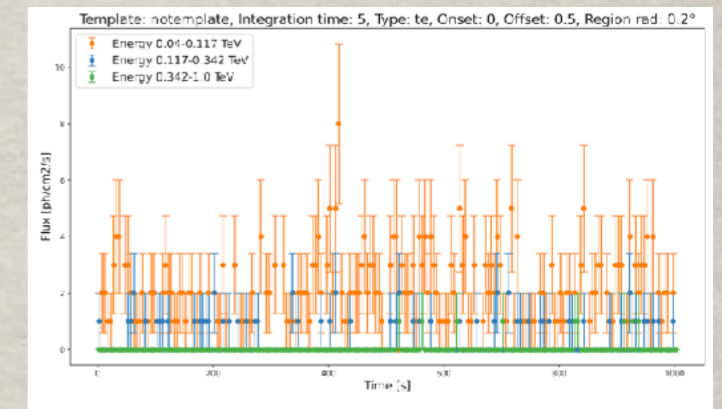
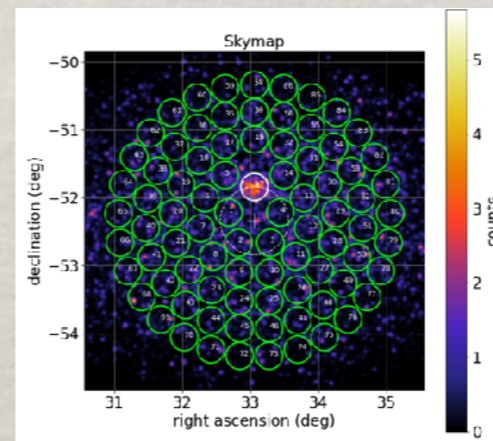
Anomaly detection for GRB search in light curves

This method performs source detection with a statistical gaussian significance $\geq 5\sigma$.

- No assumptions on the source position.
- No assumptions on the source γ -rays emission / background models.

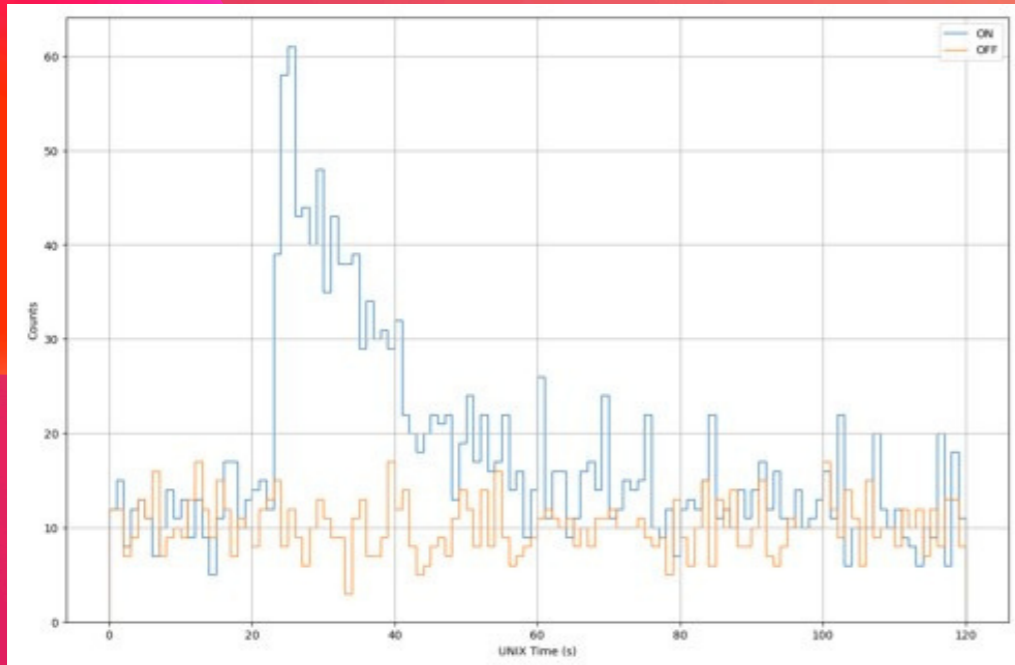
Details:

- the input data is composed by multivariate time series.
- the chosen anomaly detection techniques is based on deep learning (CNN/RNN autoencoders).
- the AE is trained offline with normal samples only (semi-supervised approach) and it learns to reconstruct the input, minimizing the reconstruction error.
- Then, the AE is fed with online data:
 - the reconstruction error for anomalous time series will be higher;
 - a threshold guides the classification.



WP3

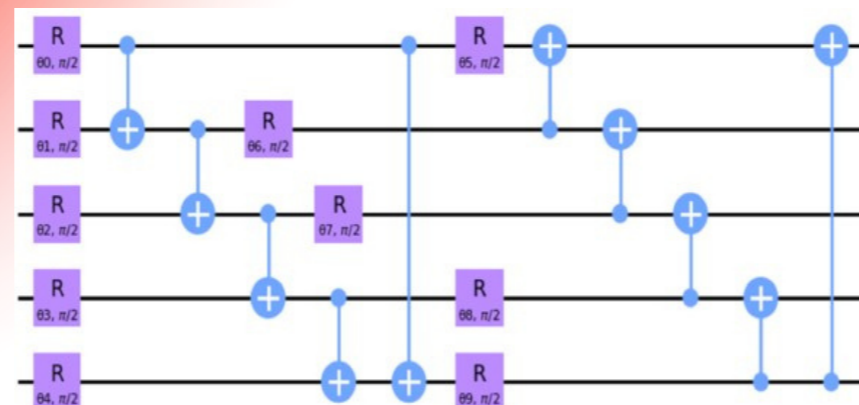
Quantum Convolutional Neural Network for GRB detection in CTA data



Classical Data

- GRB lightcurves vs noise
- 250 time series for training
- 150 time series for testing
- data mimicking CTA data
- see Farsian, F. et al. (in prep)

Quantum Convolutional Neural Network



QUANTUM CIRCUIT

parameterized in Qiskit

QUANTUM ENCODING

data reuploading method

OPTIMIZATION

performed using the COBYLA optimizer

LOSS FUNCTION

binary cross entropy

Results: 1st implementation (smaller number of parameters in QCNN)

| NN model | Num of Qubits | Num of parameters | Accuracy on Train set | Accuracy on Test set | Time |
|---------------|---------------|-------------------|-----------------------|----------------------|------|
| Classical CNN | - | 56 | 99.7% | 97.35% | 21s |
| Fully Quantum | 6 | 12 | 99.38% | 97.5% | 62s |

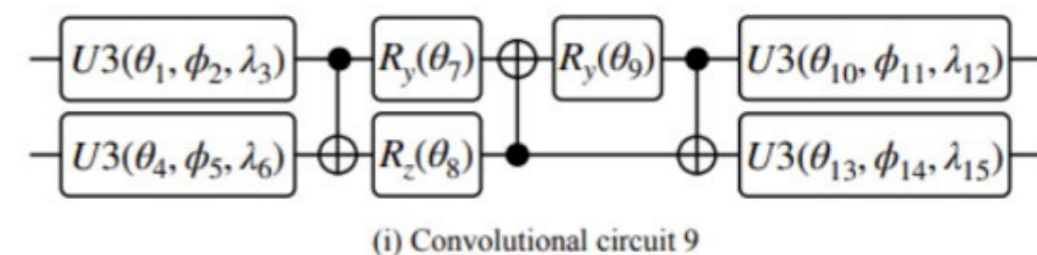
Results: 2nd implementation (only 20 light curves in the training sample)

| NN model | Num of Qubits | Num of parameters | Accuracy on Train set | Accuracy on Test set | Time |
|---------------|---------------|-------------------|-----------------------|----------------------|------------|
| Classical CNN | - | 56 | 54% | 52% | 100s |
| Fully Quantum | 6 | 24 | 99.7% | 98.35% | 23s |

Quantum Autoencoders for GRB detection in AGILE

Classical Algorithm

- classical convolutional autoencoder
- convolutional variational autoencoder
- encoders composed of 1D convolutions
- decoders composed of 1D transpose convolutions
- works on simulated data and ready to be tested on AGILE ones



Quantum Algorithm

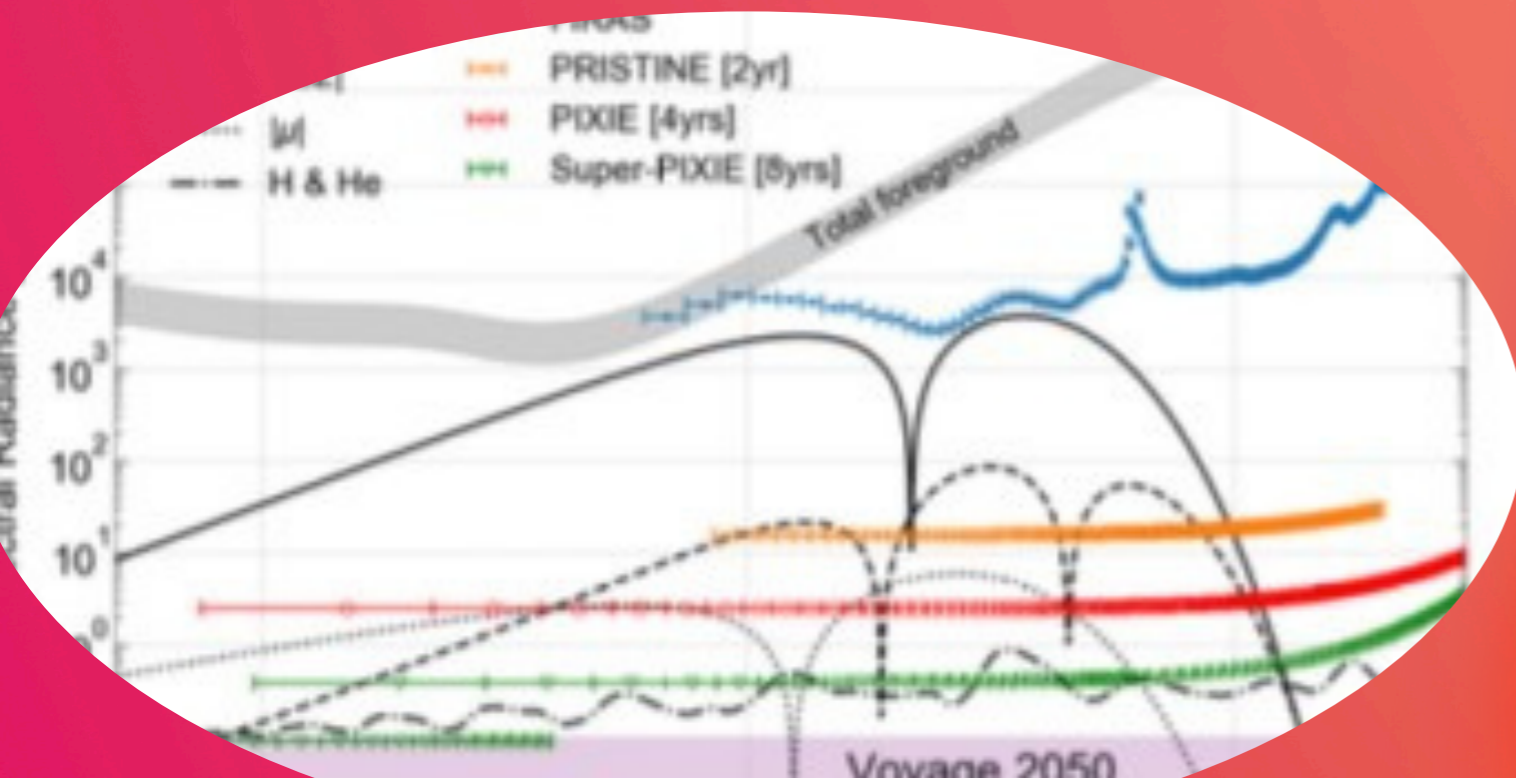
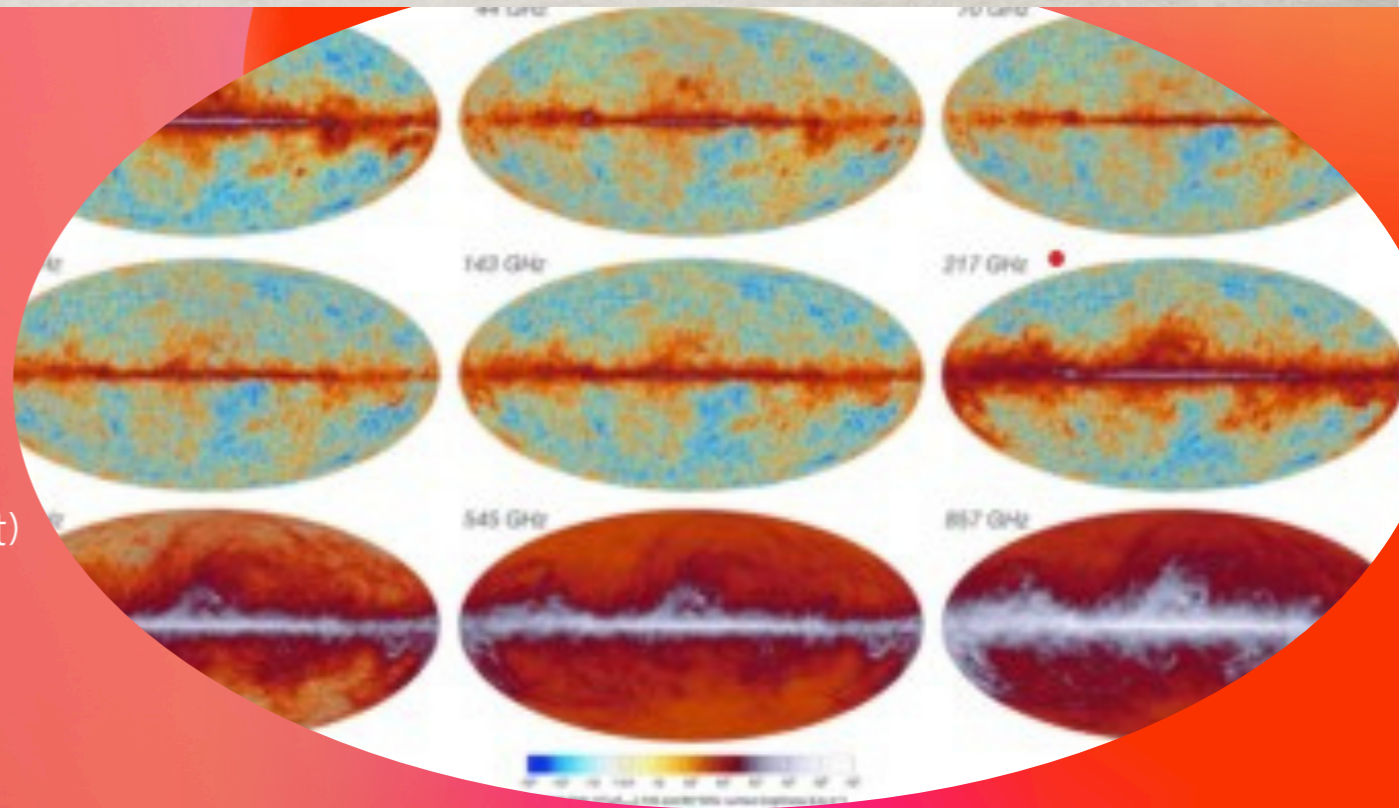
- quantum encoder + classical latent space + classical decoder
- quantum encoder + quantum latent space + classical decoder
- data reuploading and amplitude embedding as encoding techniques
- 8 qubits and 16 qubits
- interesting preliminary results but work still in progress (A. Rizzo, et al. in prep)

**From 1E5 parameters
→ 51 parameters!!**

WP4

CMB components separation

- CMB maps as sum of different contributions
 - cosmological signal (anisotropy power spectrum)
 - galactic foregrounds (e.g., synchrotron and thermal dust)
 - Cosmic Infrared Background)
 - extragalactic radio and far IR sources
- difficult to separate and time consuming



Research Plan

- different spectral features
- already available methods, e.g.
 - template fitting
 - internal linear combination
 - PCA decomposition
- develop quantum counterparts of classical methods
- use Quantum Machine Learning techniques

WP2 Multiparameter Optimisation

COSMOLOGY

Constraining a small set of cosmological parameters in a sea of nuisance ones

STRONG LENSING

Constraining halo dark matter profile marginalising over single galaxies ones

SAMPLING

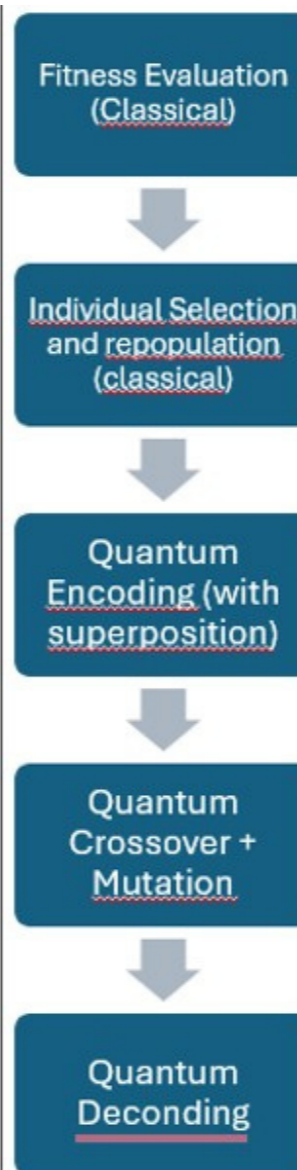
Reconstructing the posterior density in many dimensional spaces

First tests

Quantum Genetic Algorithm

FINDING BEST FIT

- COSMOLOGICAL
- PARAMETERS FROM THE
- FIT TO SNEIA, BAO AND
- CMB DATA



FITNESS EVALUATION

quantify the agreement between model and data

QUANTUM ENCODING

encode model DNA through amplitude encoding

QUANTUM CROSSOVER AND MUTATION

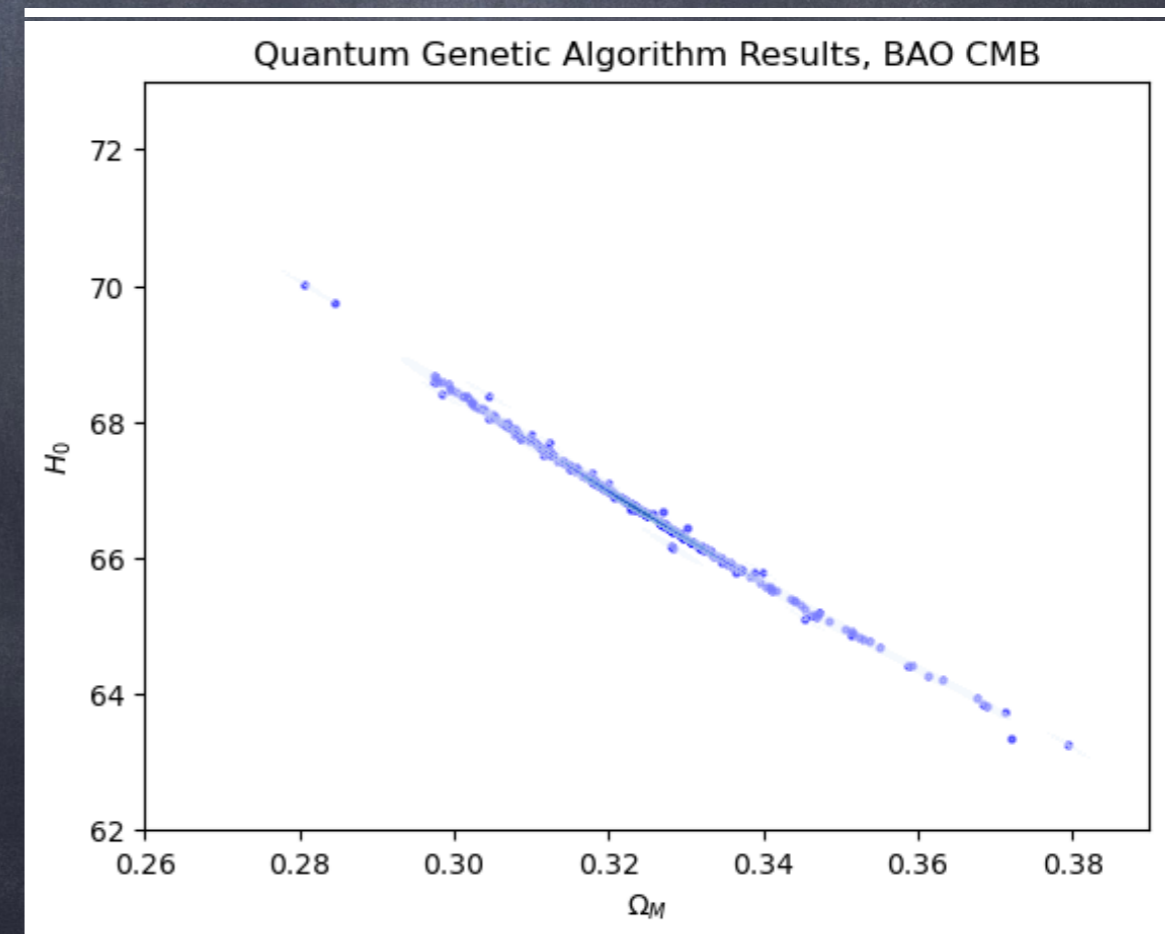
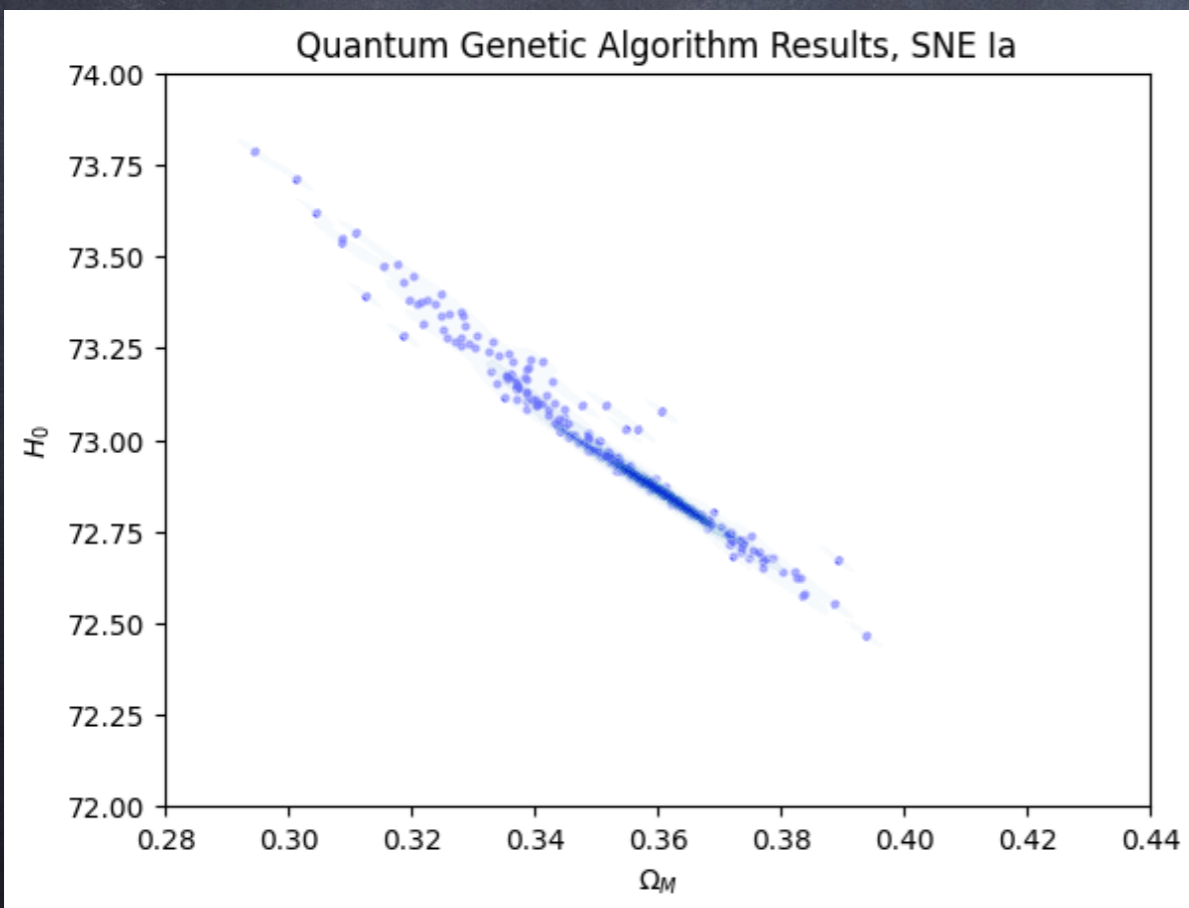
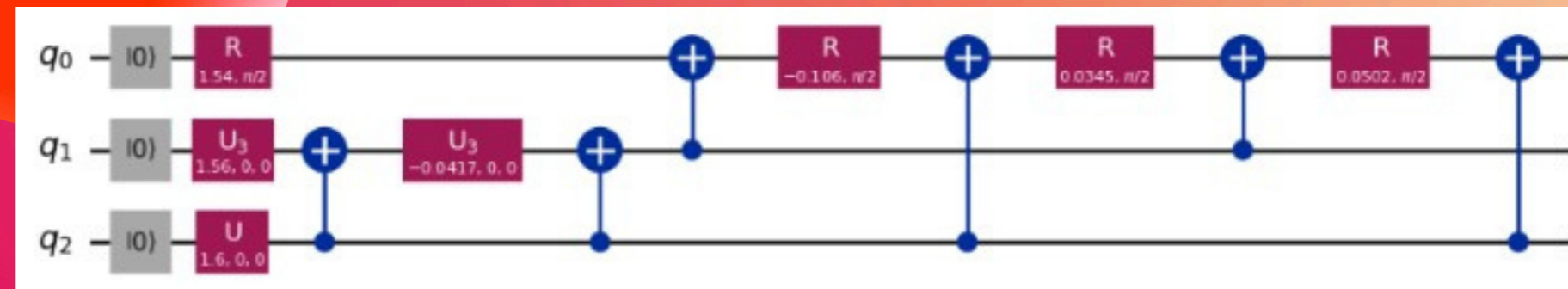
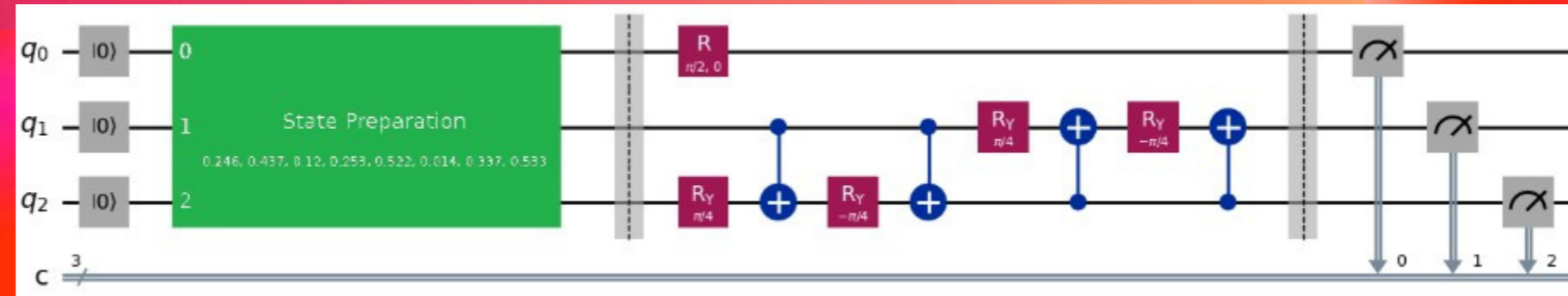
use quantum operations for genetic operations

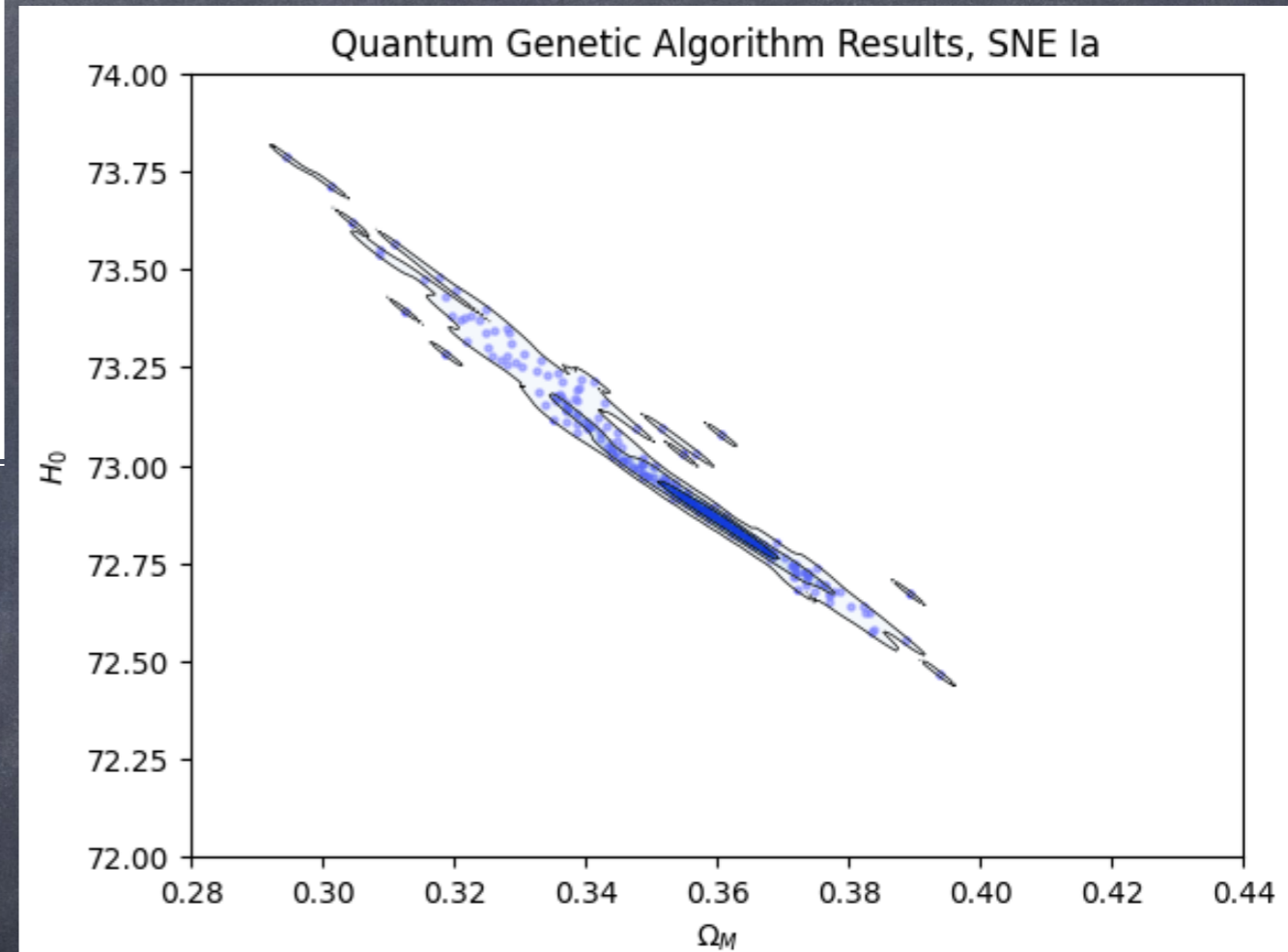
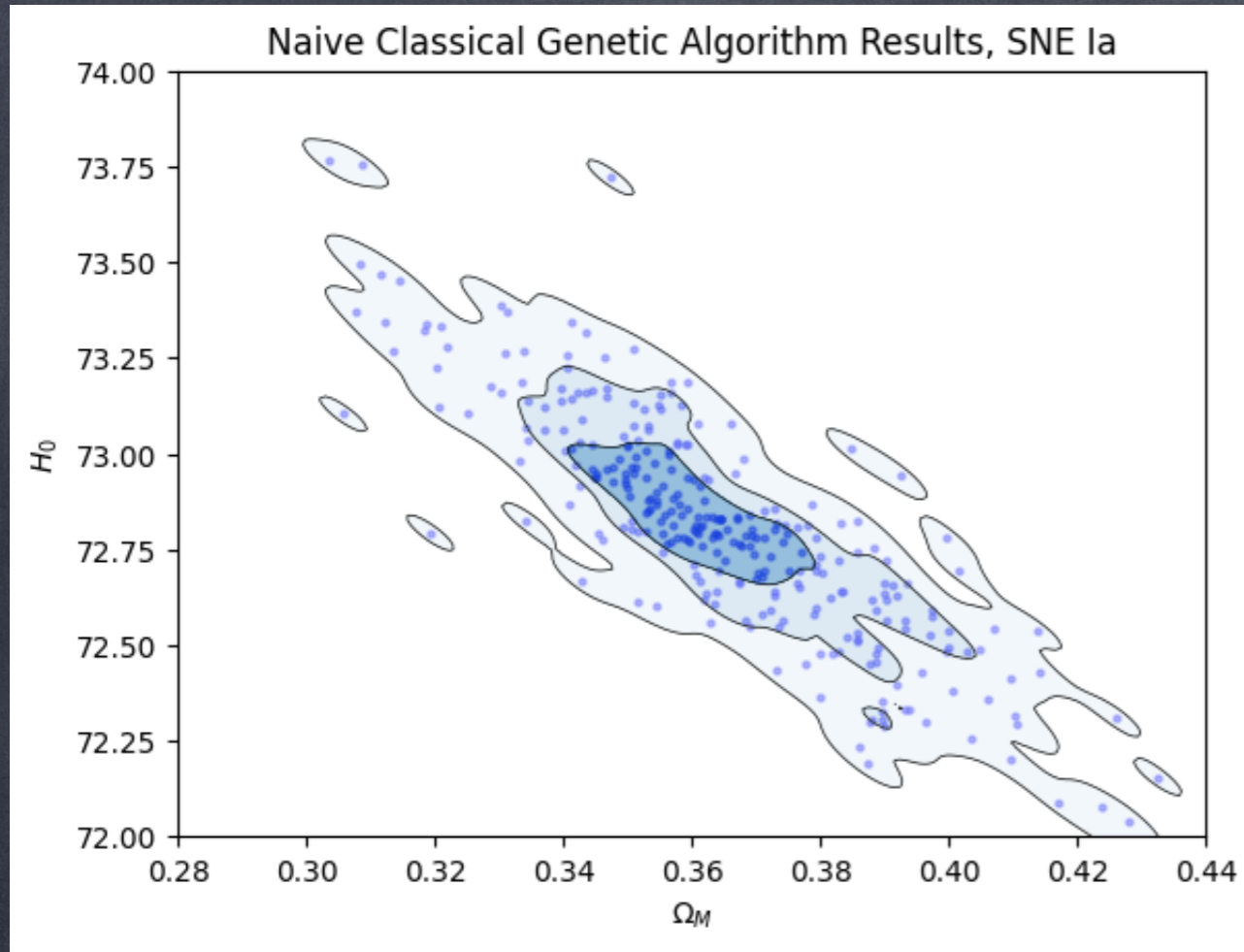
QUANTUM DECODING

decoding back to classical algorithm and iterate

Quantum Genetic Algorithm Circuit

Find best Ω_0 and H_0 from SN and CMB





Quite encouraging!!

Stay tuned ...