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Ministero  
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Italiadomani  
PIANO NAZIONALE  
DI RIPRESA E RESILIENZA



# Machine Learning Based Cosmological Radio Source Detection

(aka...kill two birds with one stone)

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Andrea Botteon (INAF-IRA) and Franco Vazza (UniBo)



USCVIII – General Assembly, Galzignano October 15<sup>th</sup> 2024

# HaMMon (Hazard Mapping and vulnerability Monitoring) : WP 4

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## WP4 objective:

Development of vulnerability curves for flood, windstorm, hail and seismic risk, based on the characteristics of the buildings automatically identifiable by the algorithms.



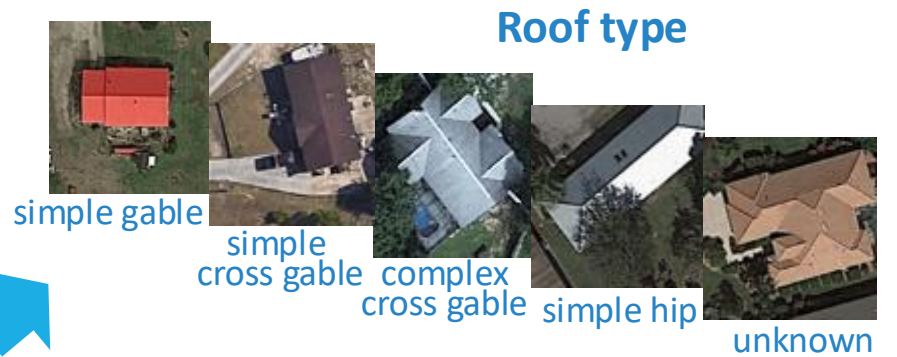
Automatic assessment of Italy's buildings patrimony through Machine Learning

## INAF focus on WP4



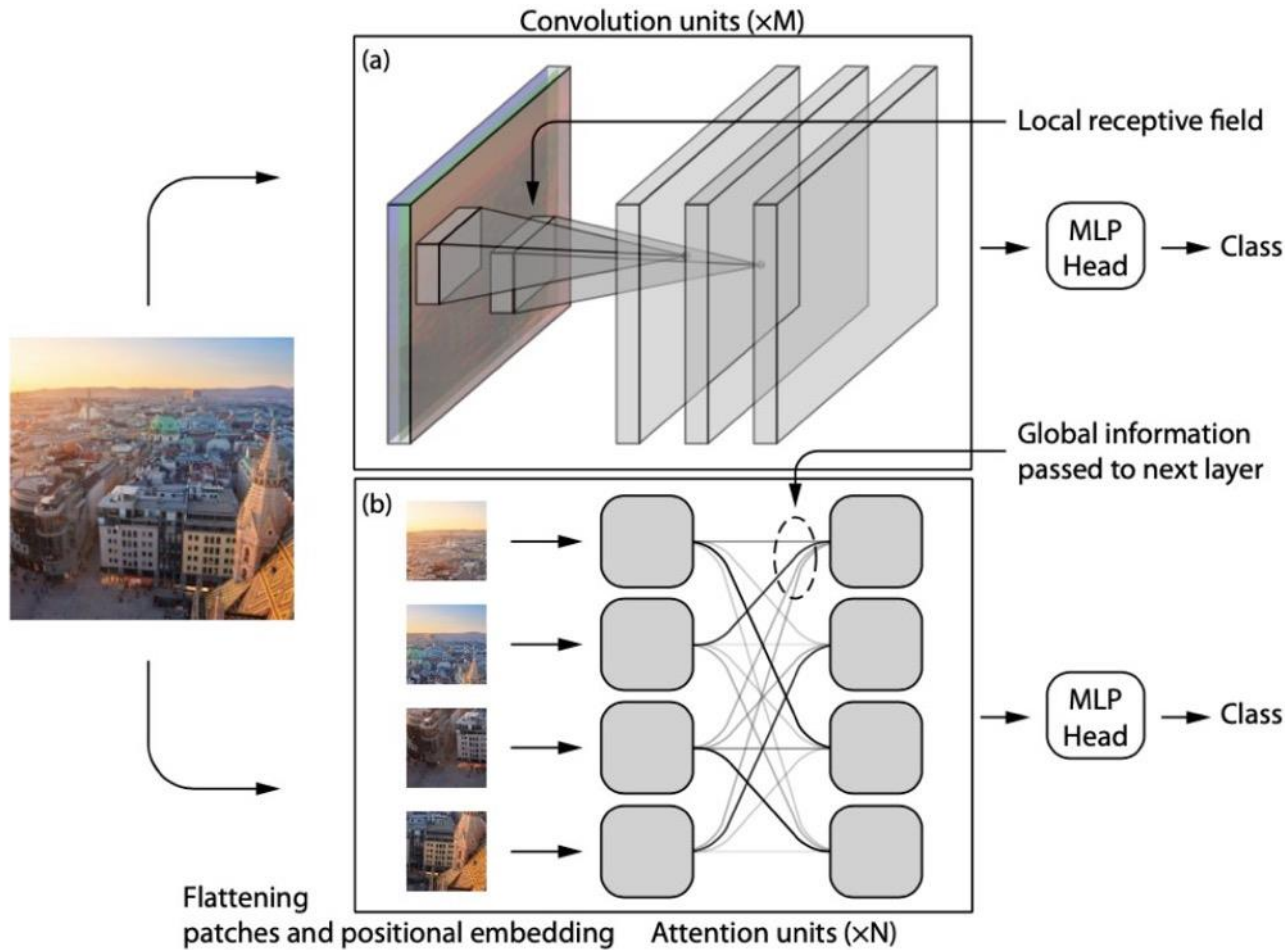
# AI solutions for automatic extraction of building features

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# Approches : CNN and Vision Transformers (ViT) architectures

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**Convolution uses a kernel to apply sliding multiplication and addition to local data points:**

- Strong ability to model local properties
- Less data for training
- Scaling linearly with the datasize  $O(N)$

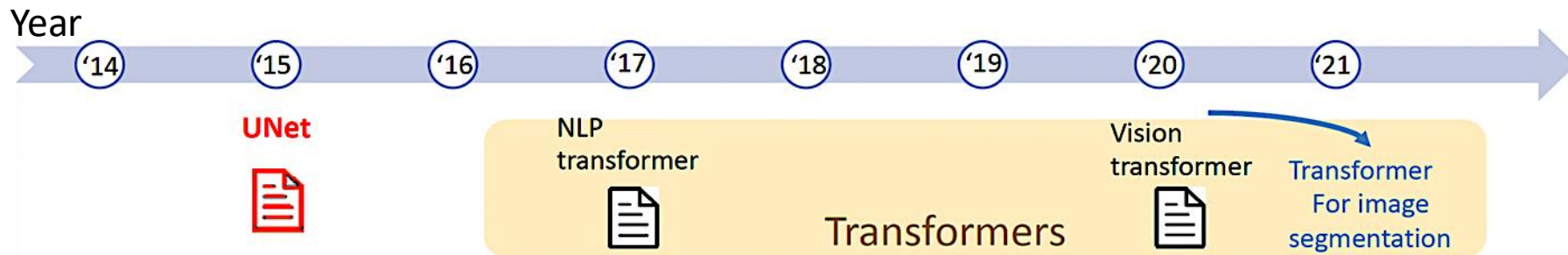
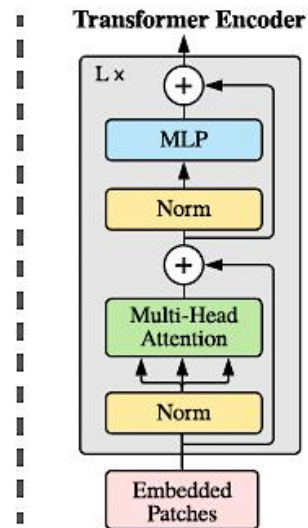
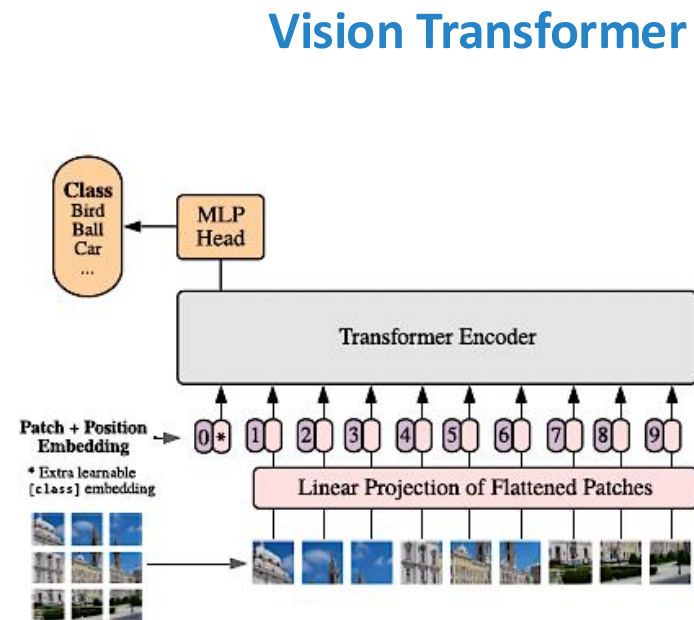
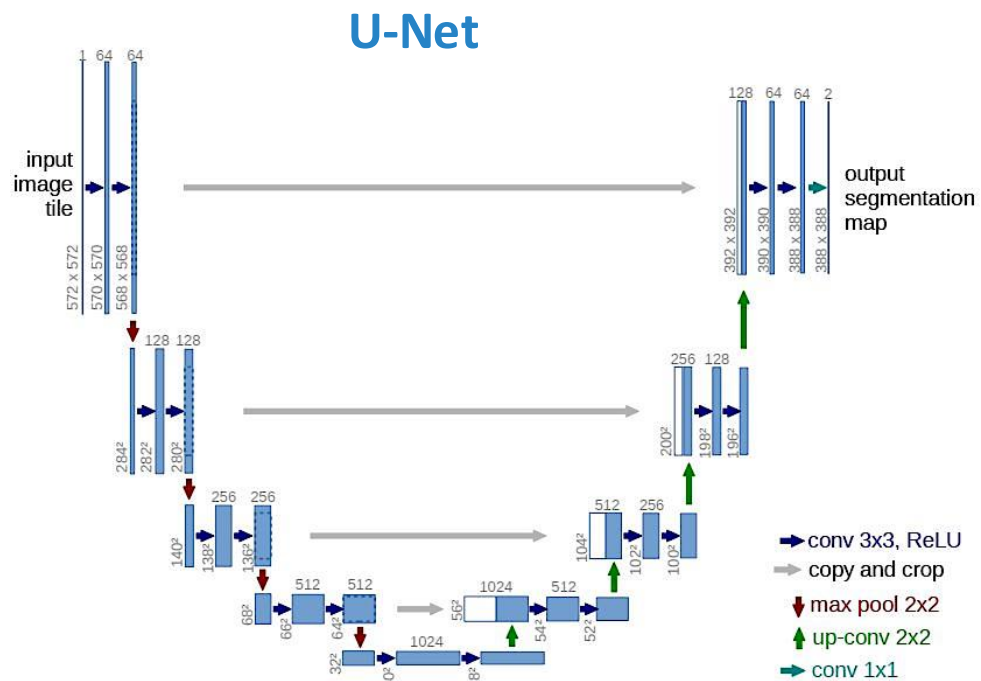
**Vision Transformers treat images as sequences of image patches.** The Transformer encoder is a stack of **Attention layers** that allow the model to learn the relationships between the patches:

- Strong ability to model long-range dependencies
- More data for training
- Scaling quadratically with the datasize  $O(N^2)$

(Tuli, Shikhar et al. 2021 "Are Convolutional Neural Networks or Transformers more like human vision?" *ArXiv* abs/2105.071972021)

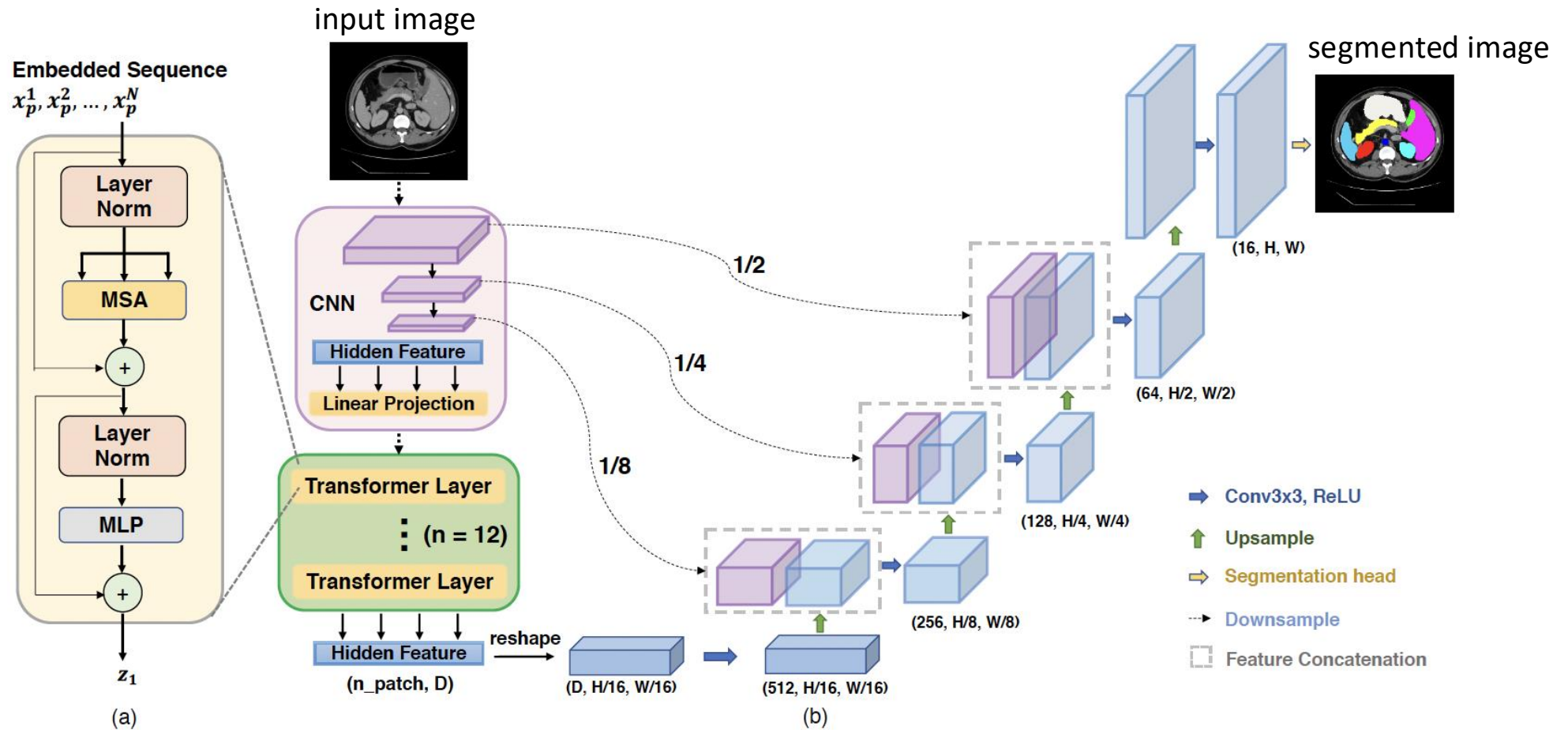


# Hybrid ViTs with Convolutions



# TransUnet

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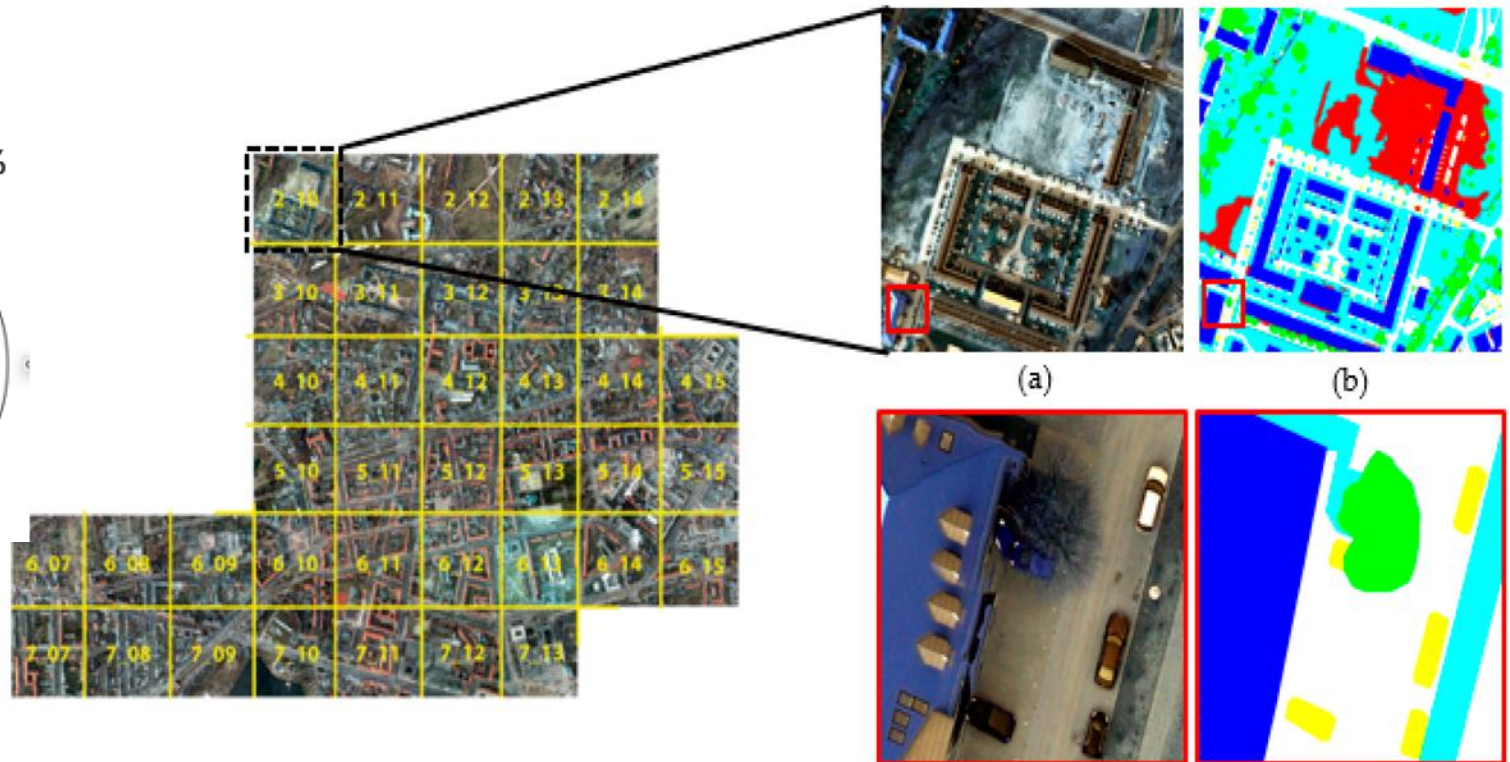
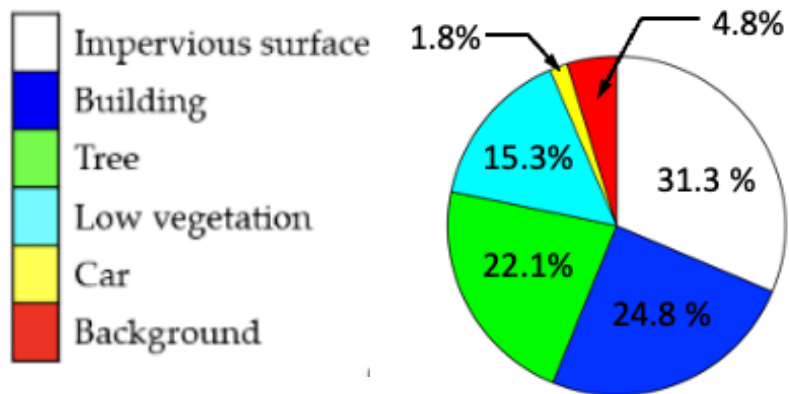


Chen, Jieneng, et al. (2021) "Transunet: Transformers make strong encoders for medical image segmentation." *arXiv preprint arXiv:2102.04306*

# Open benchmark Potsdam/Germany dataset

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- 38 high-resolution airborne images (6000 × 6000 pixels)
- training dataset 24 images, testing dataset 14 images
- spatial resolution of 5 cm.
- Near-infrared (NIR), RGB, orthorectified imagery
- ground-truth comprising 6 classes





# Comparison of predictions across Resolutions

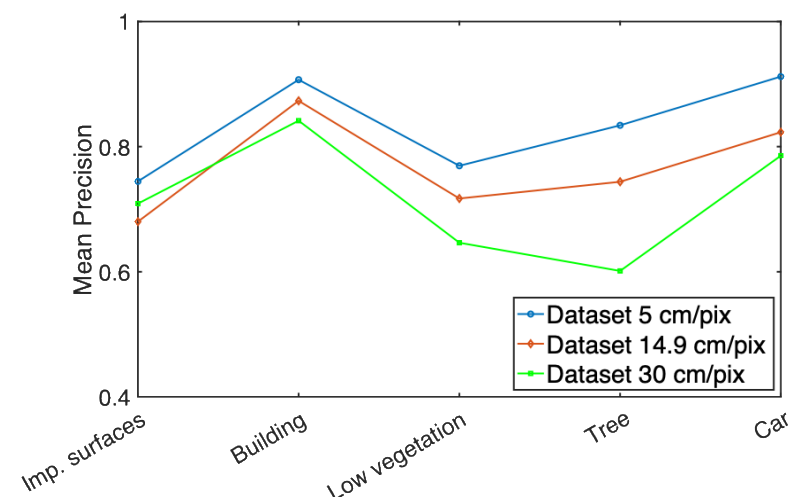
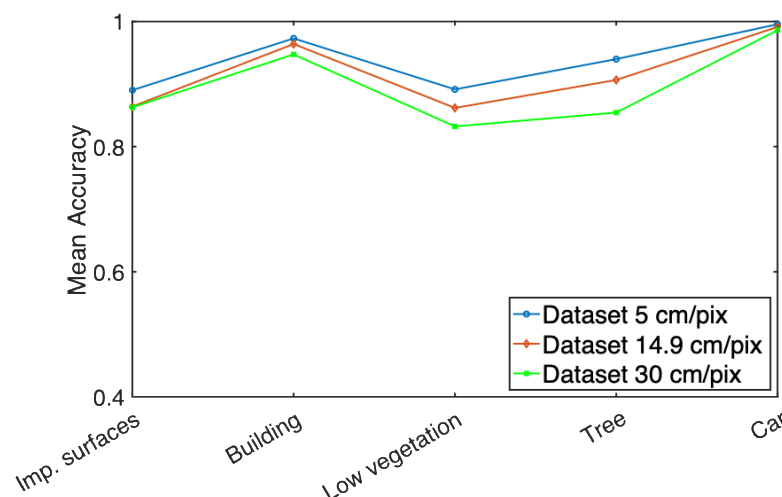
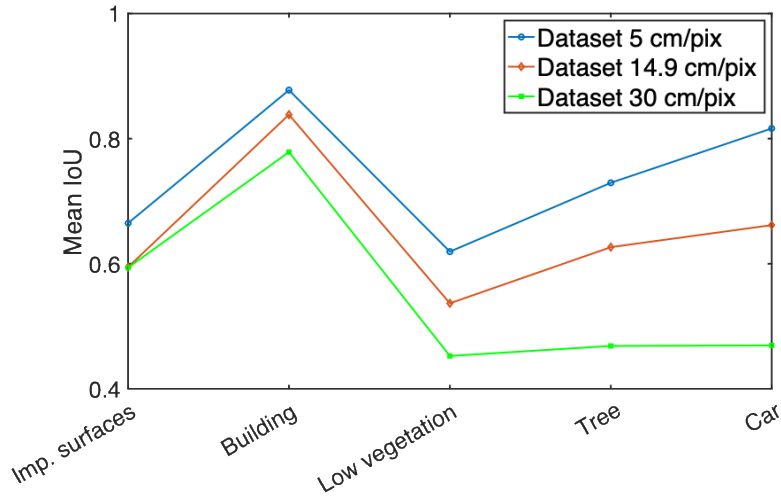
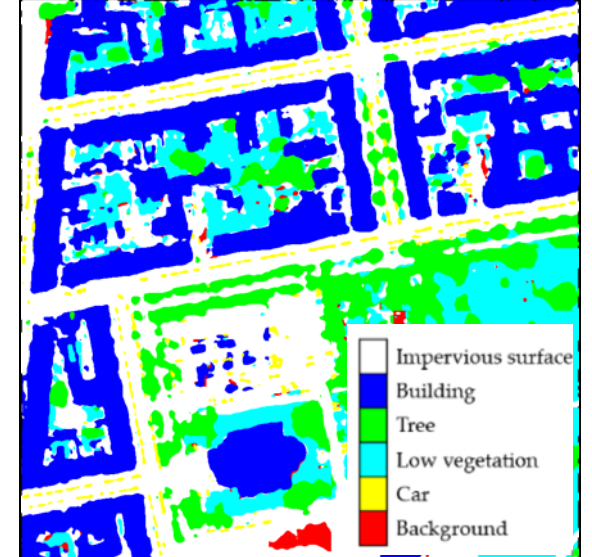
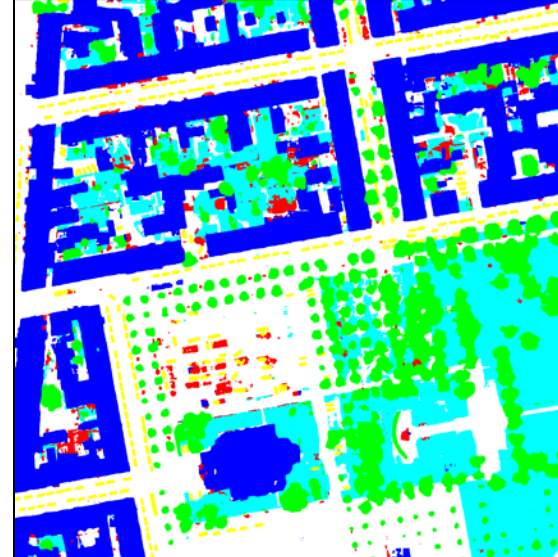
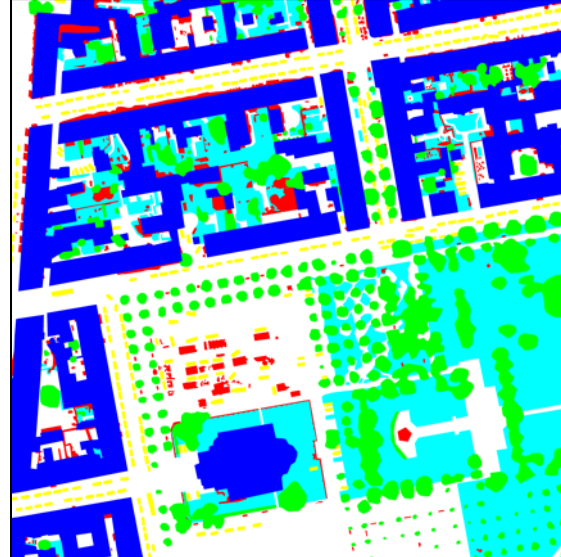
8

TransUnet: lr 0.001, epoch 20,  
batch size 24

ground truth

prediction 5.5 cm/pix

prediction 30 cm/pix





# Radio Source Detection

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Target: detection of radio sources in data collected from radio telescopes (galaxies, cluster of galaxies, cosmic web, etc.).

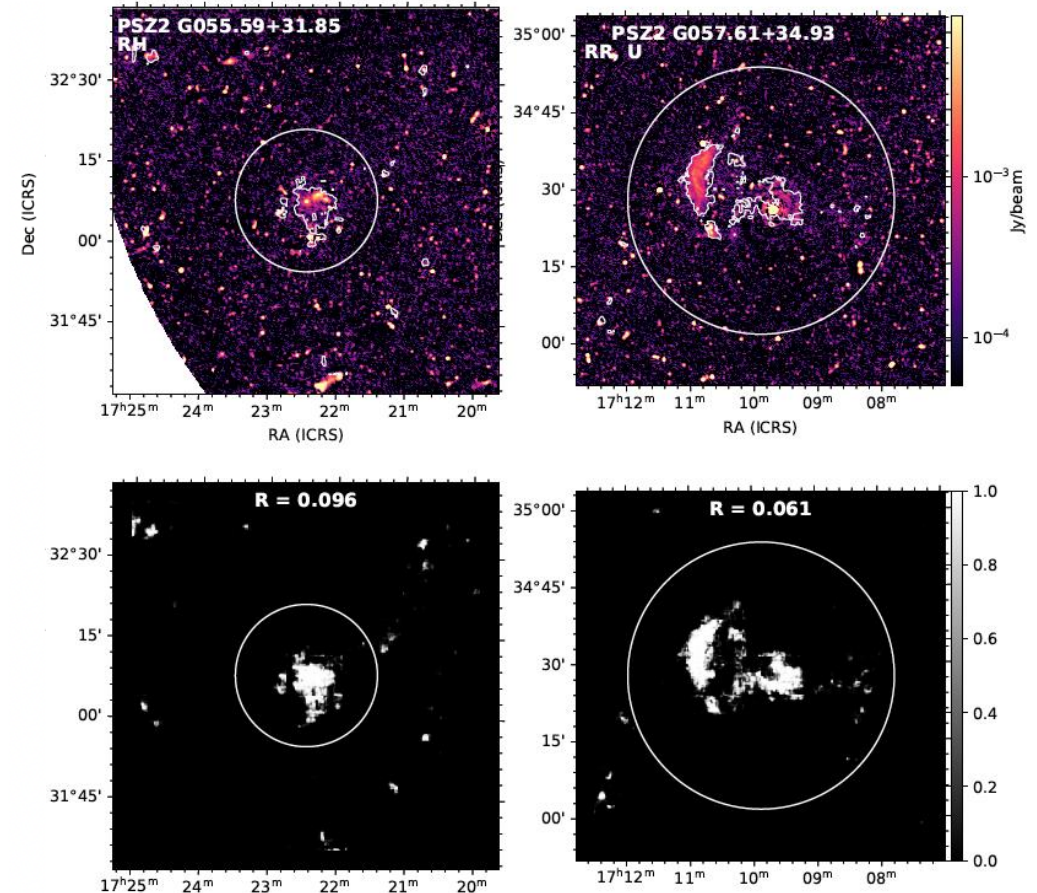
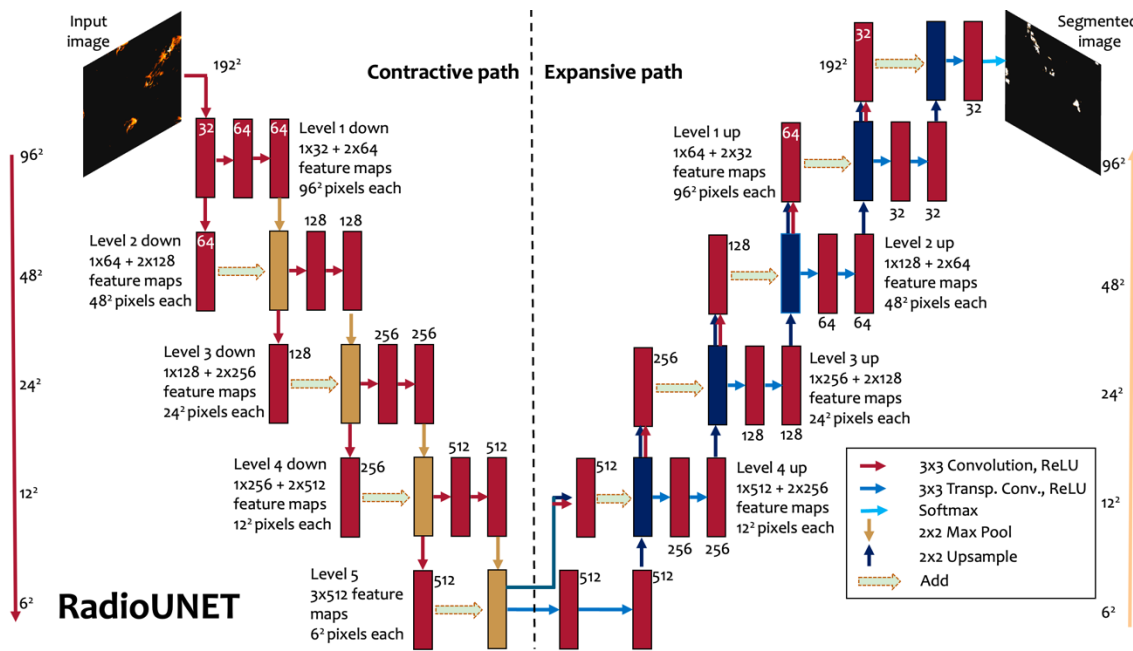
Due to the size and complexity of the data produced by current and upcoming astrophysical observatories (LOFAR, MeerKAT, SKA, etc), its processing needs to be:

- automated
- accurate
- fast



# Deep learning approach: RadioU-Net

C Stuardi, C Gheller, F Vazza, A Botteon, Radio U-Net: a convolutional neural network to detect diffuse radio sources in galaxy clusters and beyond, *Monthly Notices of the Royal Astronomical Society*, Volume 533, Issue 3, September 2024

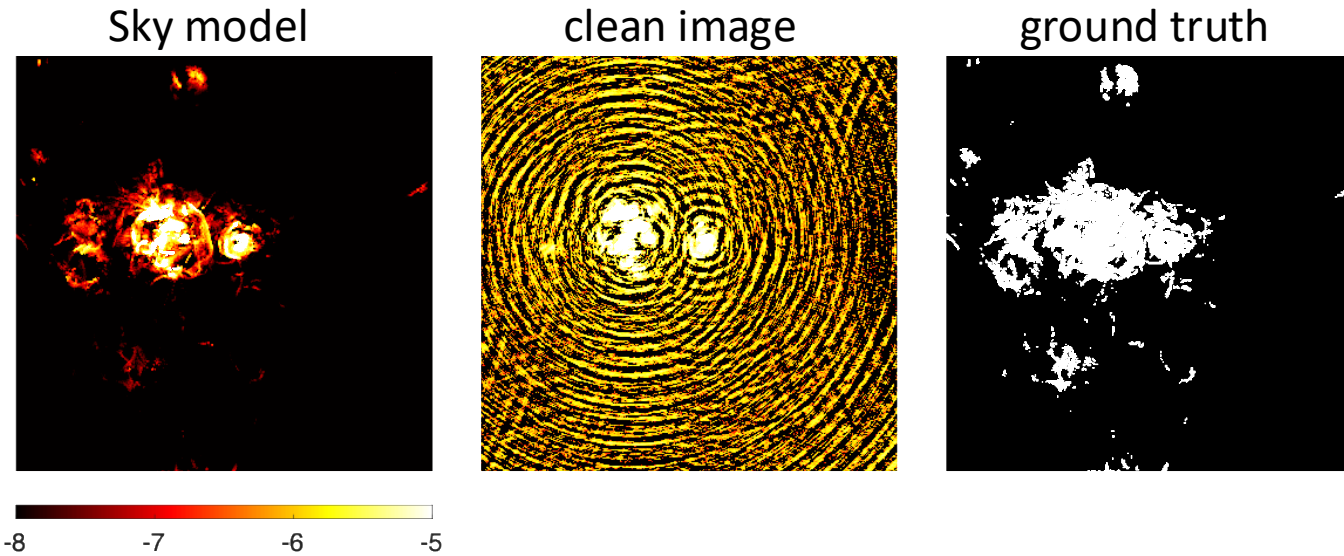


Examples of LOFAR LoTSS-DR2/PS2 galaxy clusters processed by Radio U-Net and successfully classified

# Exploring Vision Transformers for Radio Source Detection

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Comparison of RadioUnet vs TransUnet (TU) based on the simulated dataset



## Mock observation:

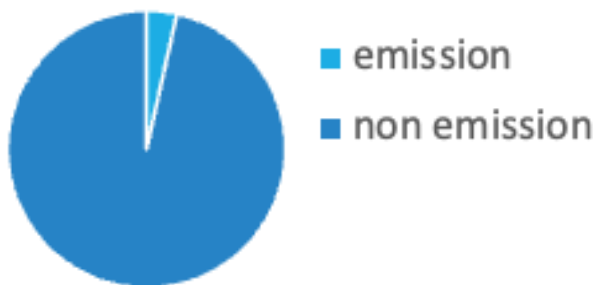
- Light cones from cosmological simulations (sky model)
- Observations simulated with WSClean (<https://gitlab.com/aroffringa/ws-clean>) + custom tools to reproduce noise and artifacts **such as**
  - ✓ **Random instrumental noise** (thermal, electronics...)
  - ✓ **Incomplete sampling of visibility space** (PSF or dirty beam)
- LOFAR HBA – 8 hrs observation



# Comparison of Evaluation Metrics

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- Dataset with 140 mock images (2000 × 2000 px)
- training 100 images, testing 40 images
- nominal spatial resolution 2 arcsec
- single frequency 150 MHz
- ground-truth comprising 2 classes:
  - ✓ class 1: emission (3.5%)
  - ✓ class 0: non emission (96.5%)



## Class 1 Statistics: Emission

RadioUnet: 200 epochs, lr  $10^{-4}$ , batch size 50, tile size 192 px

metric	value	std
Dice	0.4559	$\pm 0.0074$
IoU	0.3054	$\pm 0.0055$
Precision	0.6372	$\pm 0.0096$
Recall	0.3916	$\pm 0.0078$

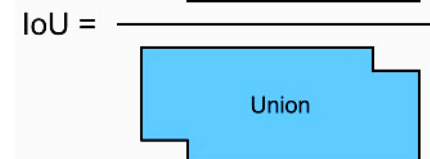
TransUnet: 200 epochs, lr 0.005, batch size 24, tile size 512 px

metric	value	std
Dice	0.5377	$\pm 0.0096$
IoU	0.3766	$\pm 0.0090$
Precision	0.6121	$\pm 0.0184$
Recall	0.4987	$\pm 0.0120$

$$\text{Precision} = \frac{\text{True Positives}}{\text{True Positives} + \text{False Positives}}$$

$$\text{Recall} = \frac{\text{True Positives}}{\text{True Positives} + \text{False Negatives}}$$

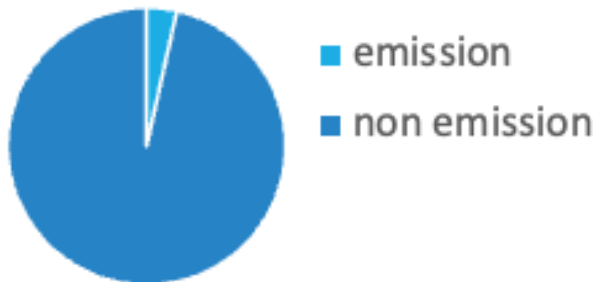
$$\text{Dice} = \frac{2 \cdot \text{IoU}}{1 + \text{IoU}}$$



# Comparison of Evaluation Metrics

13

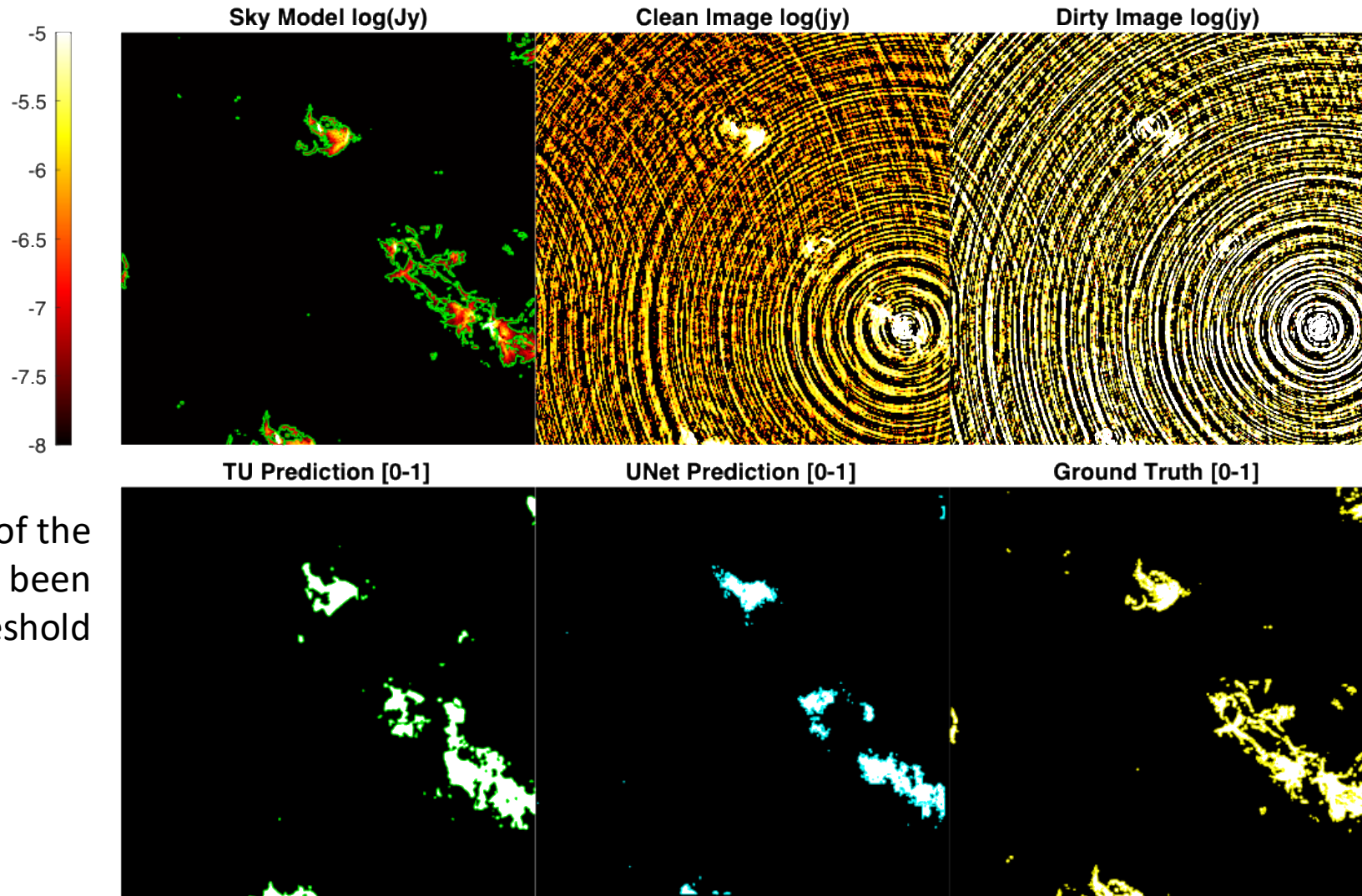
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Computational Performance (NVIDIA A100)	TransUNet	UNet
Inference time (sec/image)	2.9	0.36
Training time (sec/epoch)	88.5	13.5
Image resolution	2000x2000	2000x2000

# Results

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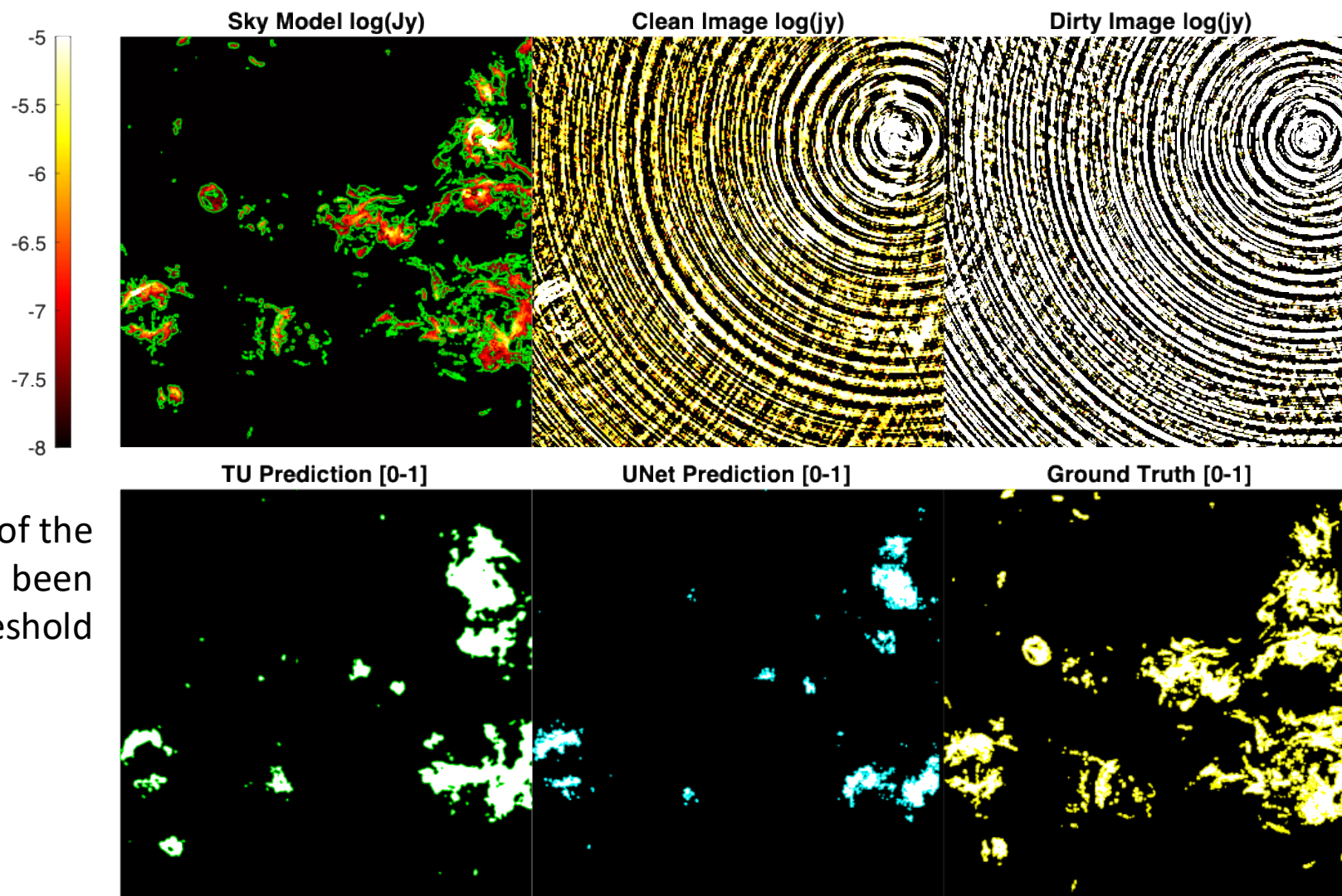


The mask (ground truth) of the network prediction has been set with probability threshold  $> 0.5$



# Results

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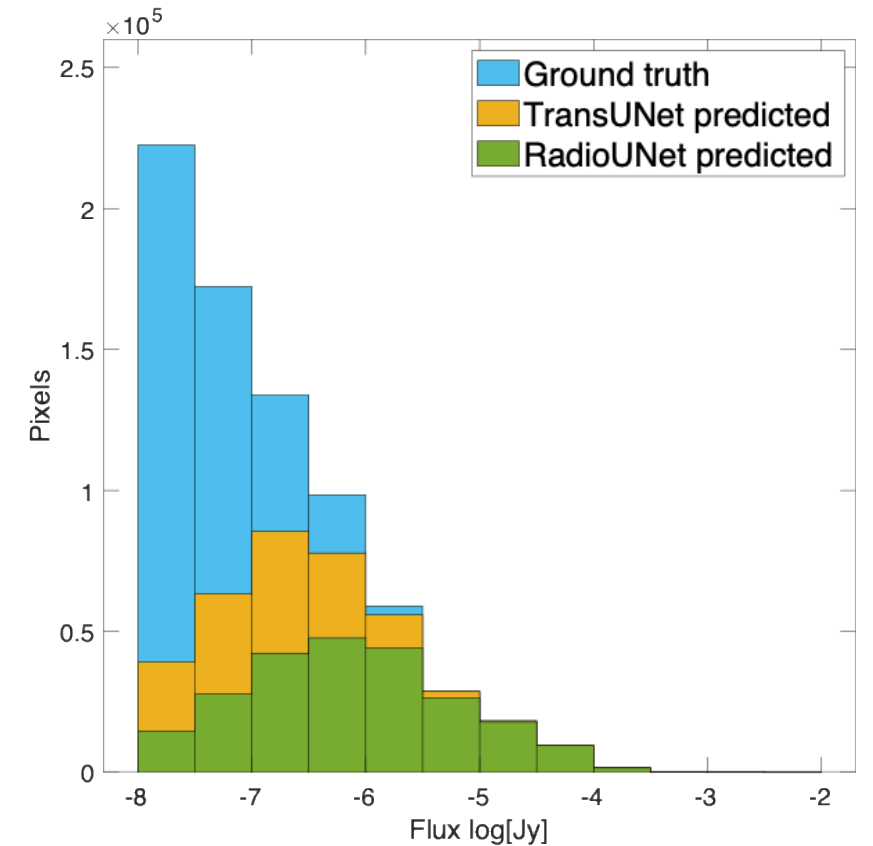
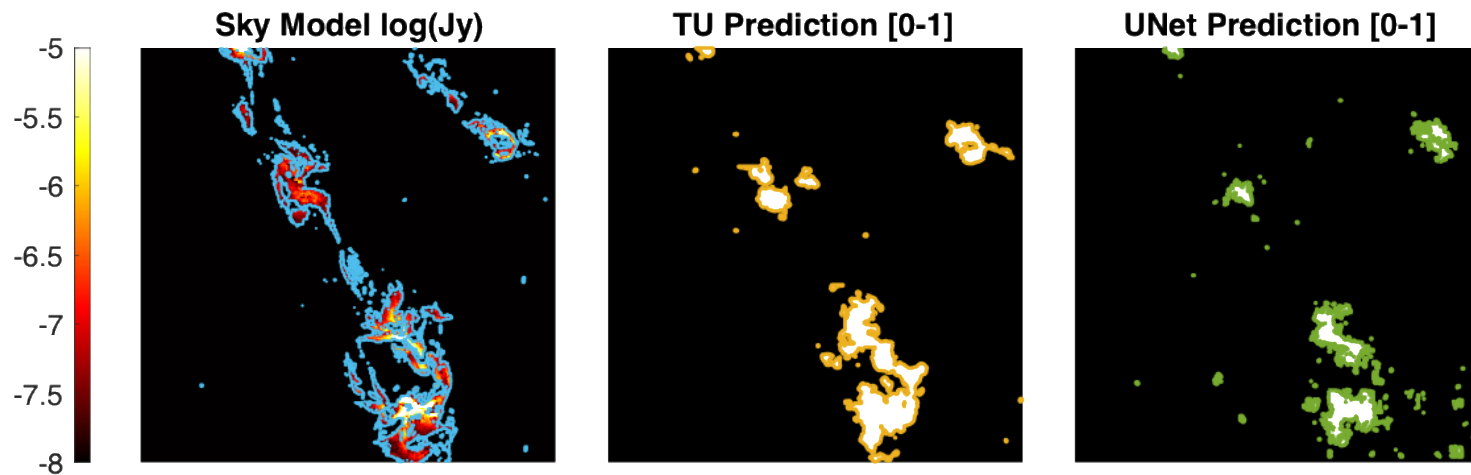
The mask (ground truth) of the network prediction has been set with probability threshold  $> 0.5$

# Results: Flux Distribution of the test dataset

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Flux distribution of:

- all pixels of the sky images with flux  $> 10^{-8}$  Jy, i.e. ground truth
- all pixels of the sky images overlapping the TransUNet predictions
- all pixels of the sky images overlapping the RadioUnet predictions



# Take Home Messages

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- ⑩ **The same network** can be used successfully on **completely different data and fields**
- ⑩ Deep Learning can be **effectively used to process and analyze large and complex astronomical images**
- ⑩ **Vision transformers** can be effective in enhancing segmentation performance of traditional convolutional approach but application on real astronomical images is to be investigated





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**THANK YOU  
FOR YOUR ATTENTION**



USCVIII – General Assembly, Galzignano October 15<sup>th</sup> 2024