

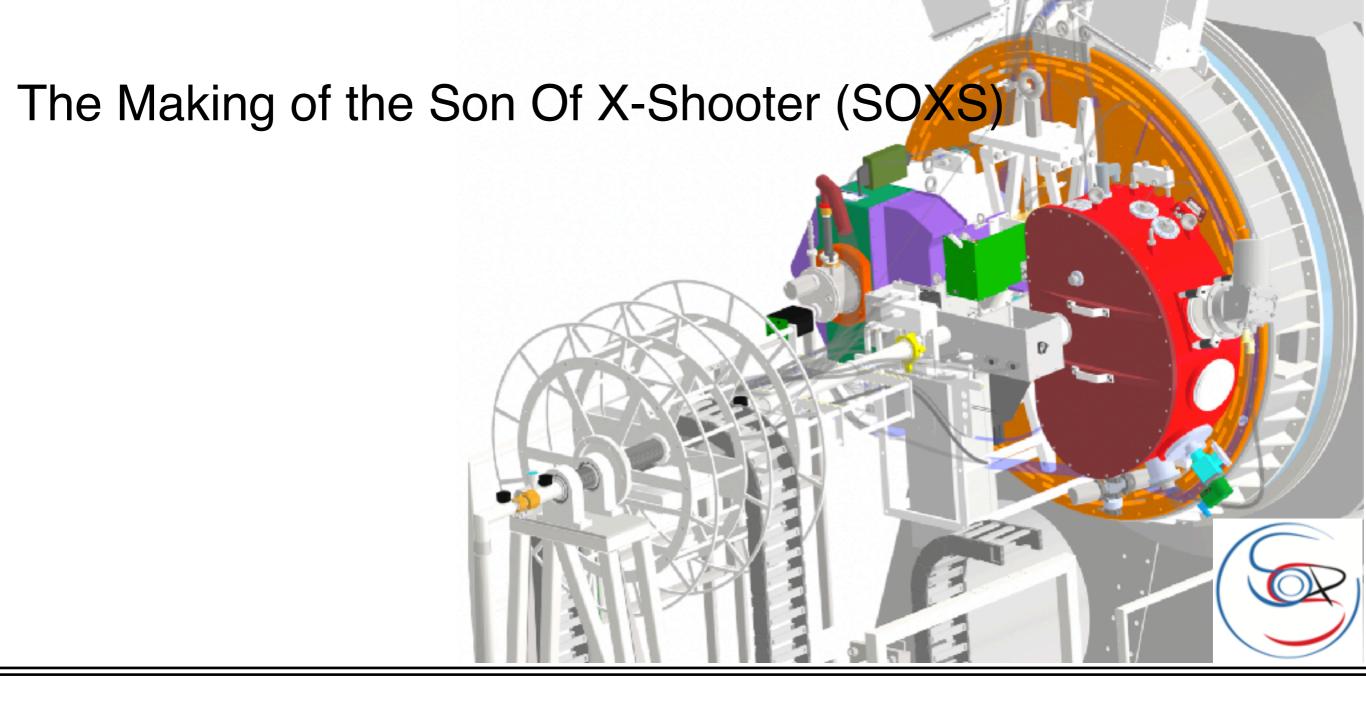
Kalyan Radhakrishnan & AIV Team

SOXS Instrument AIV Lead & Deputy System Engineer, Astronomical Instrumentation and Adaptive Optics Group, INAF-Osservatorio Astronomico di Padova (INAF-OAPD), Padova, Italy

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"SOXS Science Meeting" INAF-Osservatorio Astronomico di Capodimonte 25 June 2024



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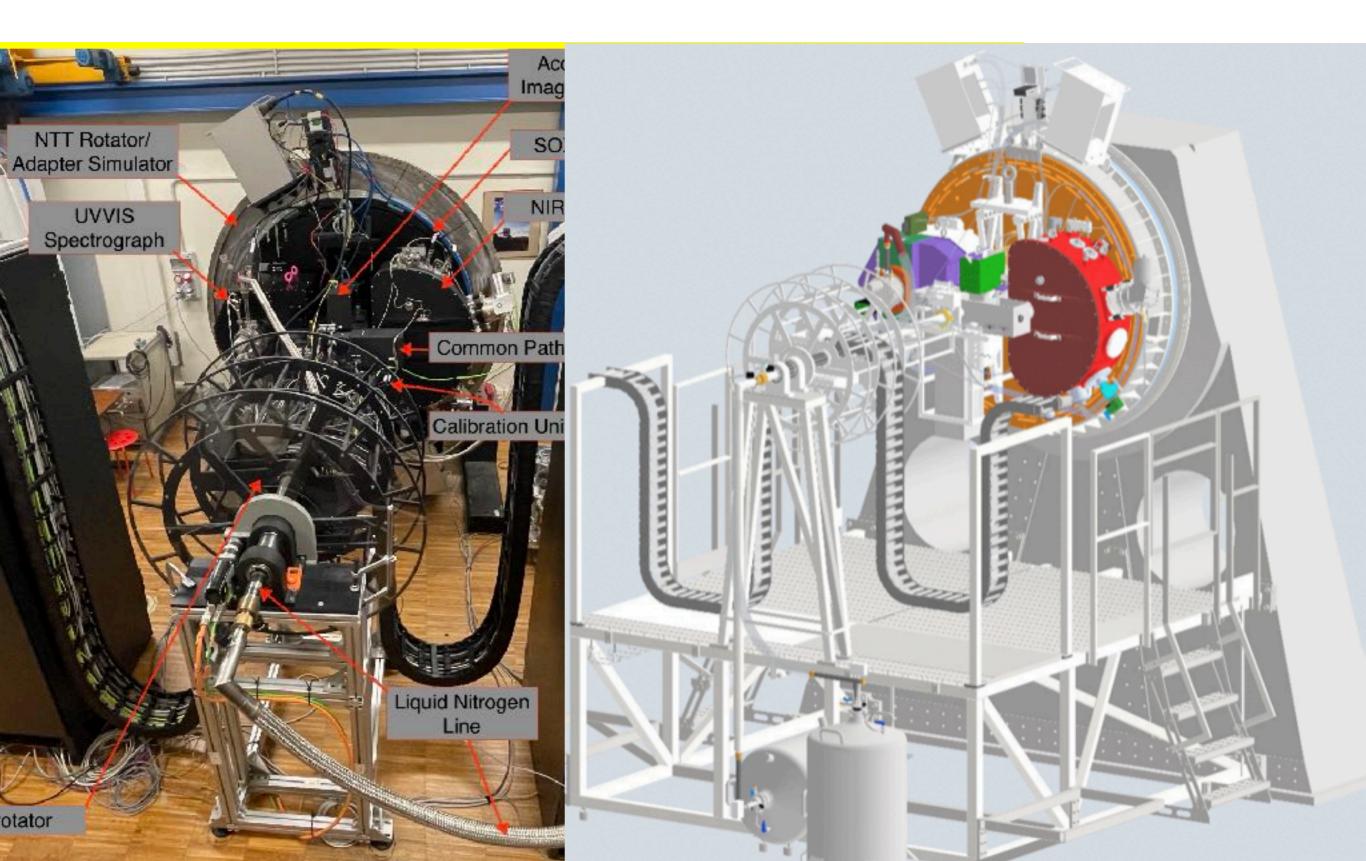


"SOXS Science Meeting" INAF-Osservatorio Astronomico di Capodimonte 25 June 2024

Outline



La Silla vs Padova

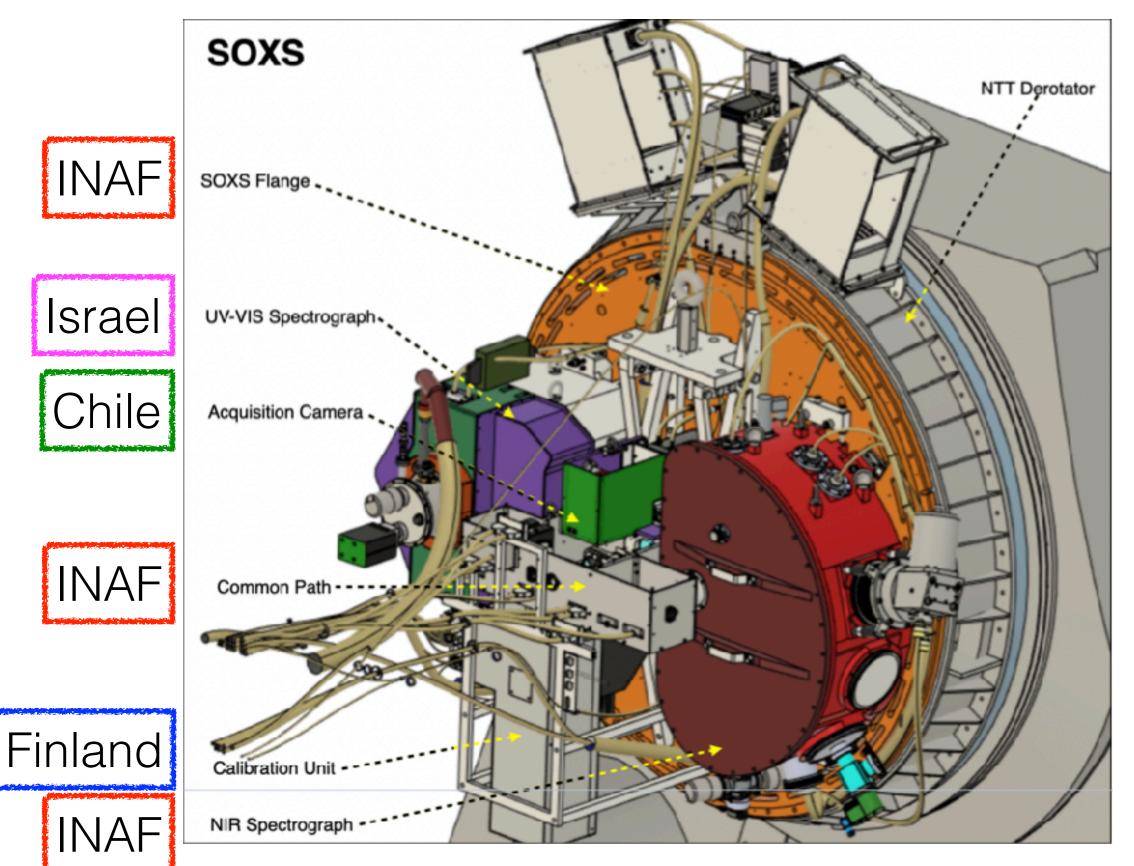


SOXS in numbers

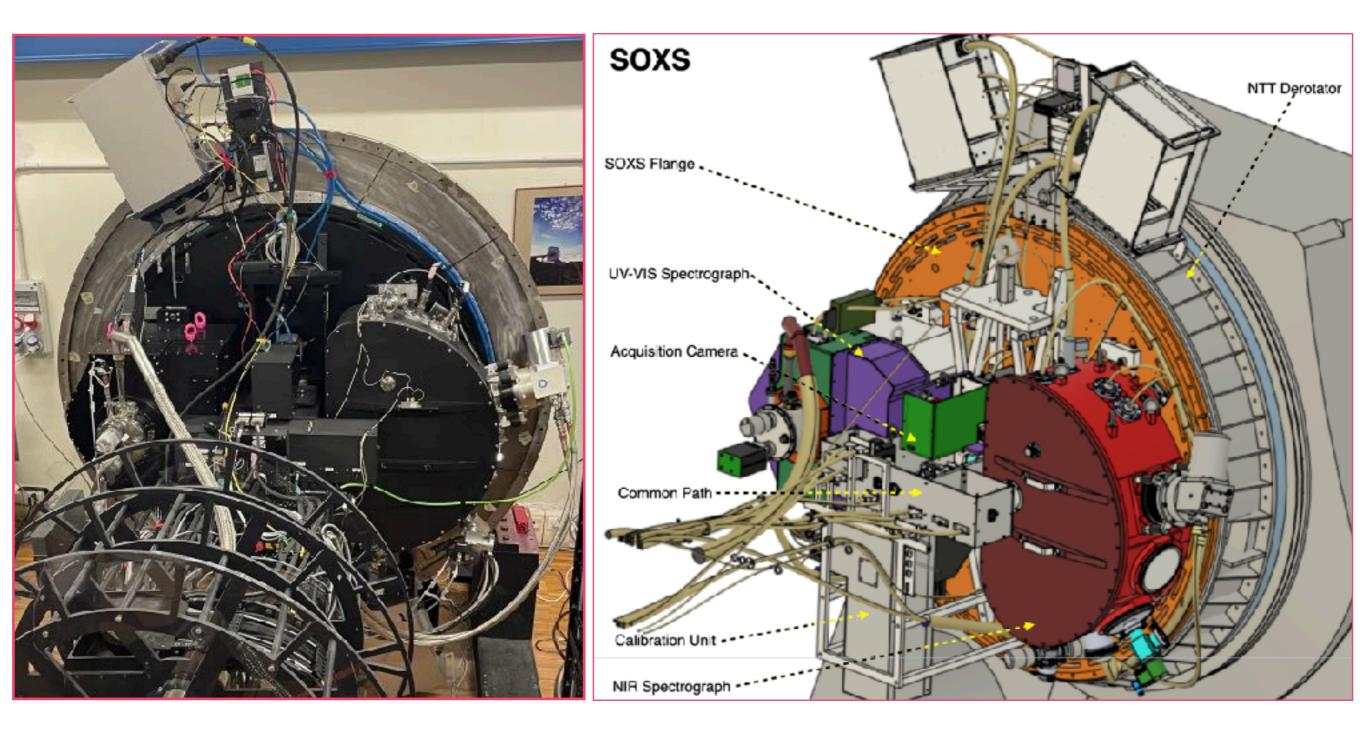
	UV-VIS	NIR
Spectral Range	350 - 850 nm	800 - 2000 nm
Resolution (1" slit)	>3600 (~4500 avg)	5000
Slit Widths	0.5", 1", 1.5", 5"	0.5", 1", 1.5", 5"
Slit height	12"	12"
Detector	e2V CCD44-82 2k x4k	Teledyne H2RG 2k x 2k
Pixel Scale	15um	18 um
Detector Scale	0.28"/pixel	0.25"/pixel

	Acquisition Camera	
Spectral Range	360 - 970 nm	
Filters	SDSS u, g, r, i, z, LSST y, and VIMOS V	
FoV	3.5" x 3.5"	
Detector	Andor iKon M-934 1k x 1k	
Pixel Scale	13um	
Detector Scale	0.205"/pixel	

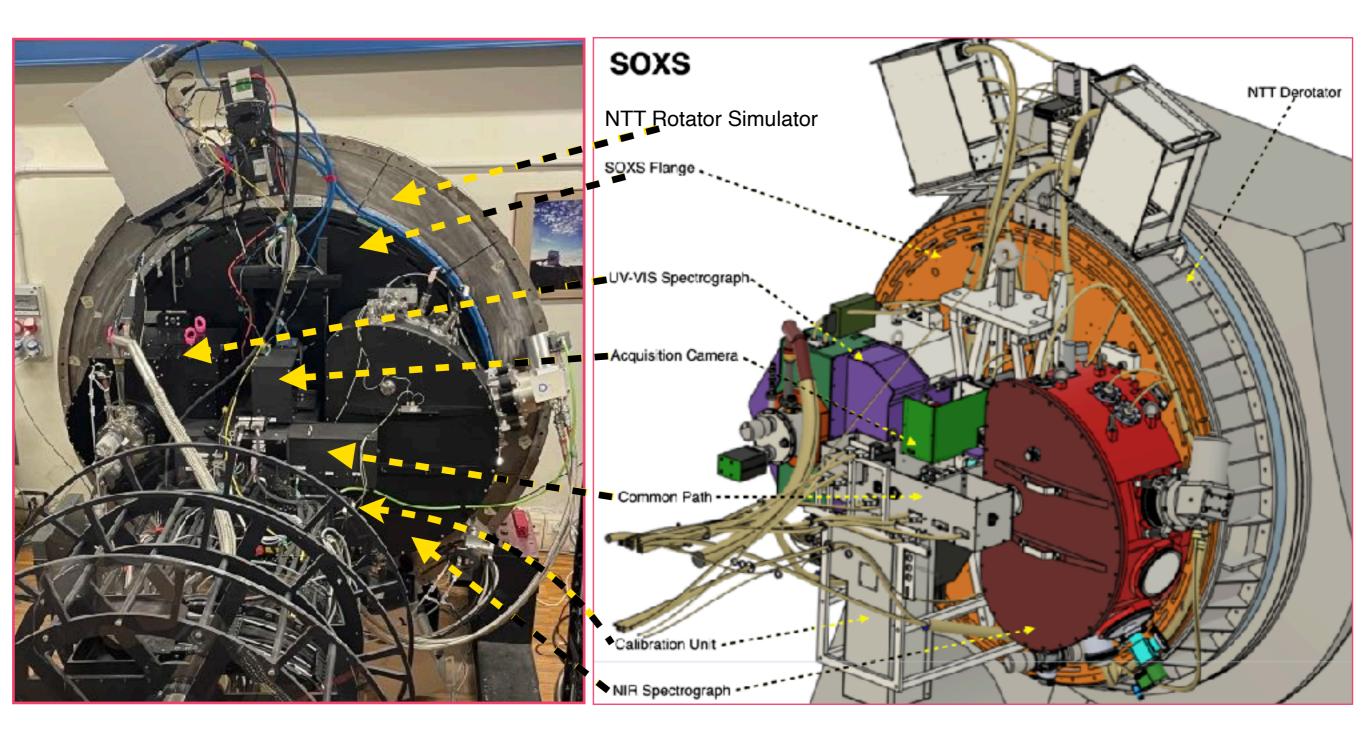
SOXS subsystems



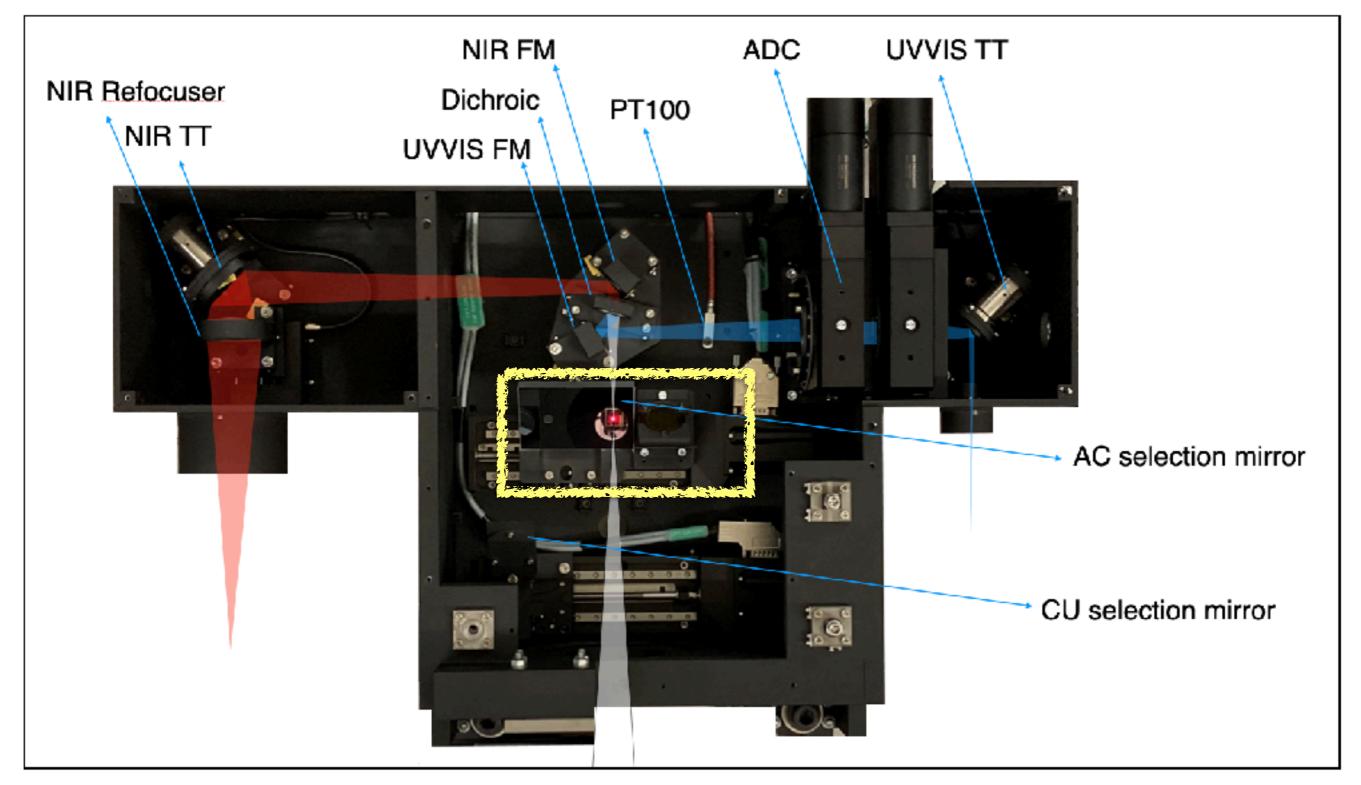
SOXS subsystems



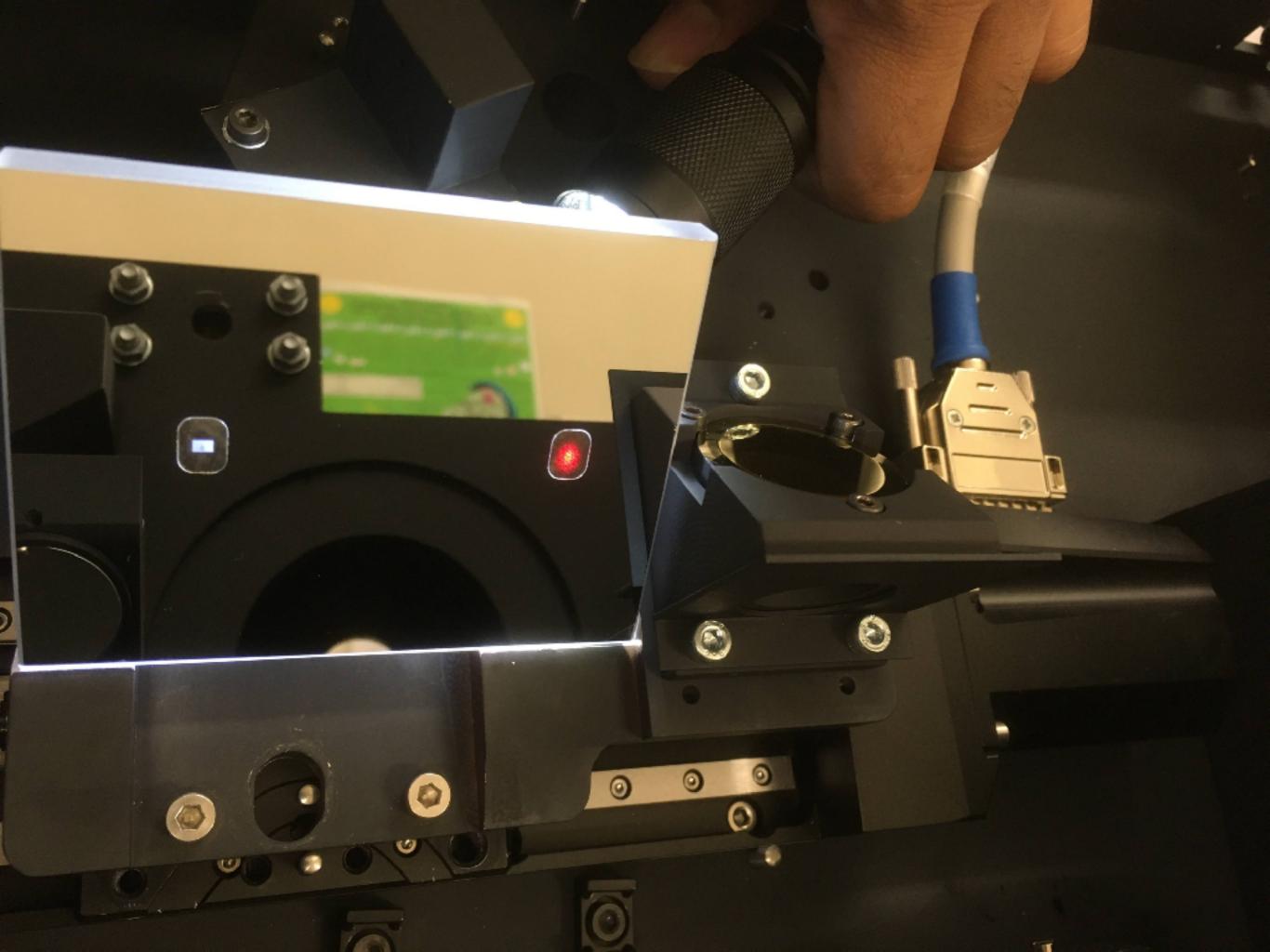
SOXS subsystems



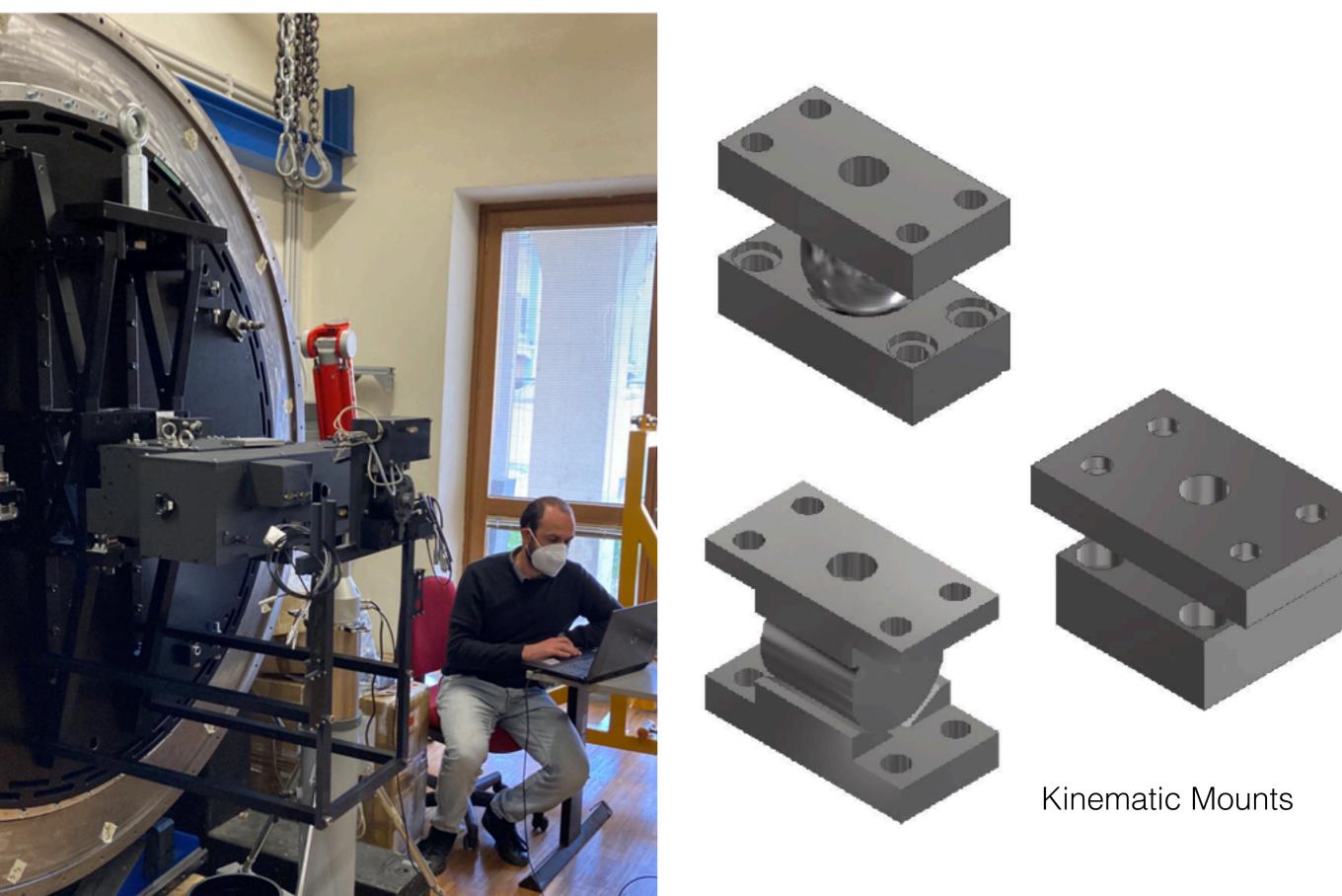
Common Path

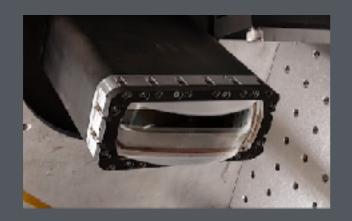


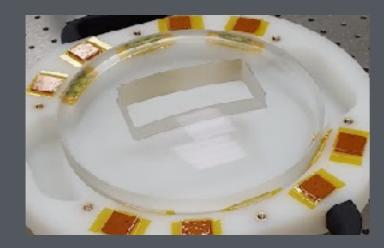
Receives the F/11 beam from the telescope and feeds the spectrographs with an F/6.5 beam

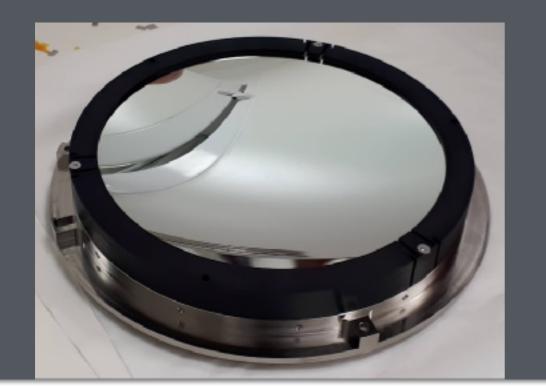


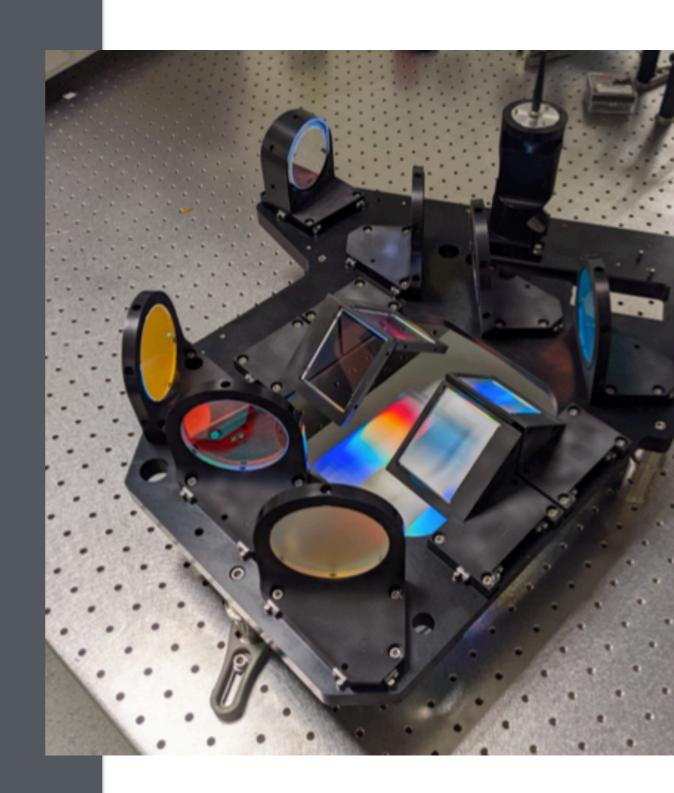
Mounting to the Flange using Kinematic Mounts







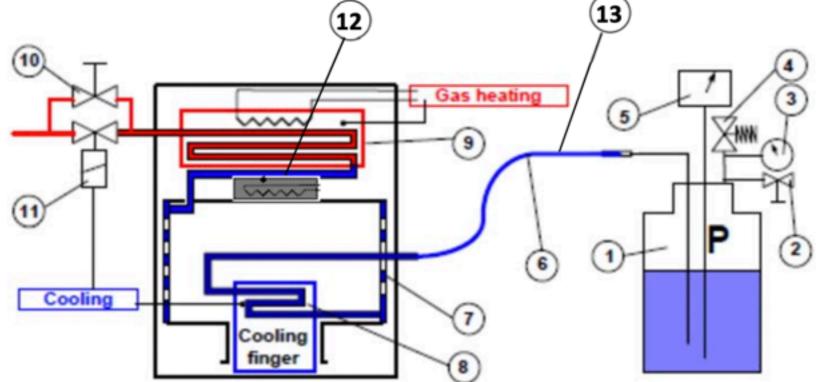




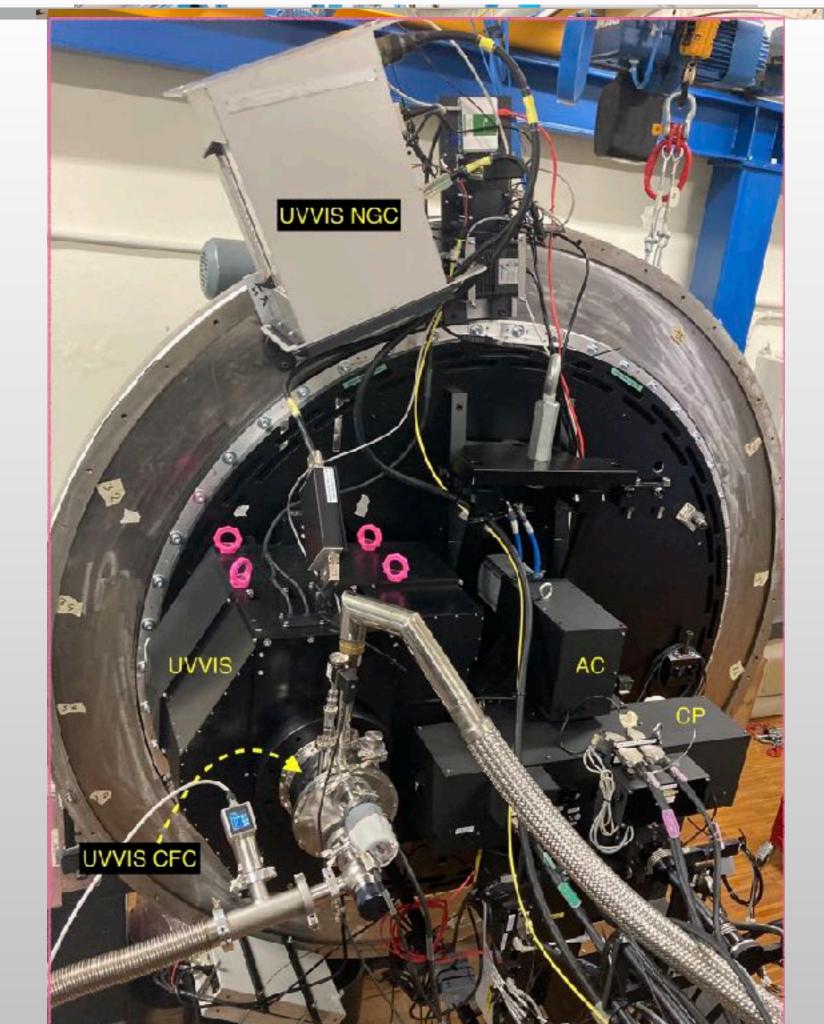
UV-VIS Cryogenic System

- A Continuous Flow Cryostat (CFC) is used to cool down the CCD.
- The CFC together with the detector head makes the UV-VIS cryostat.

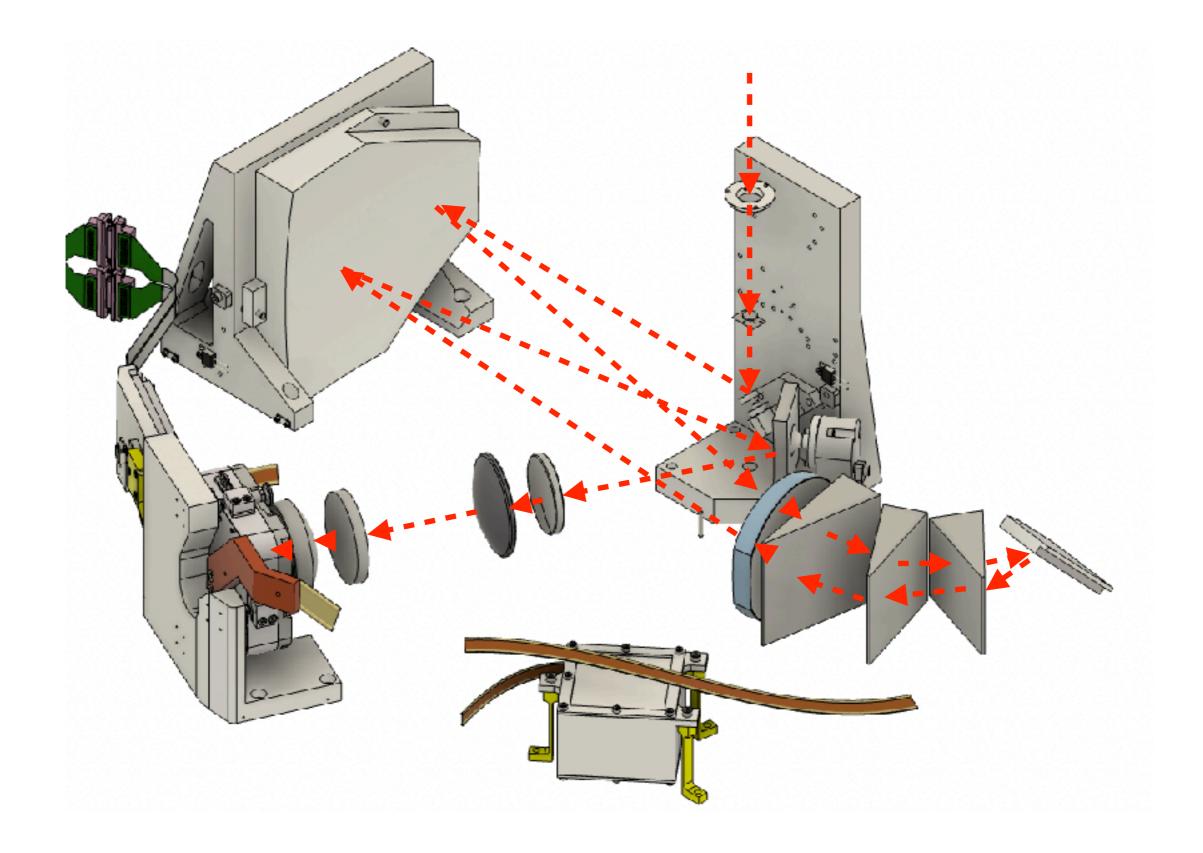




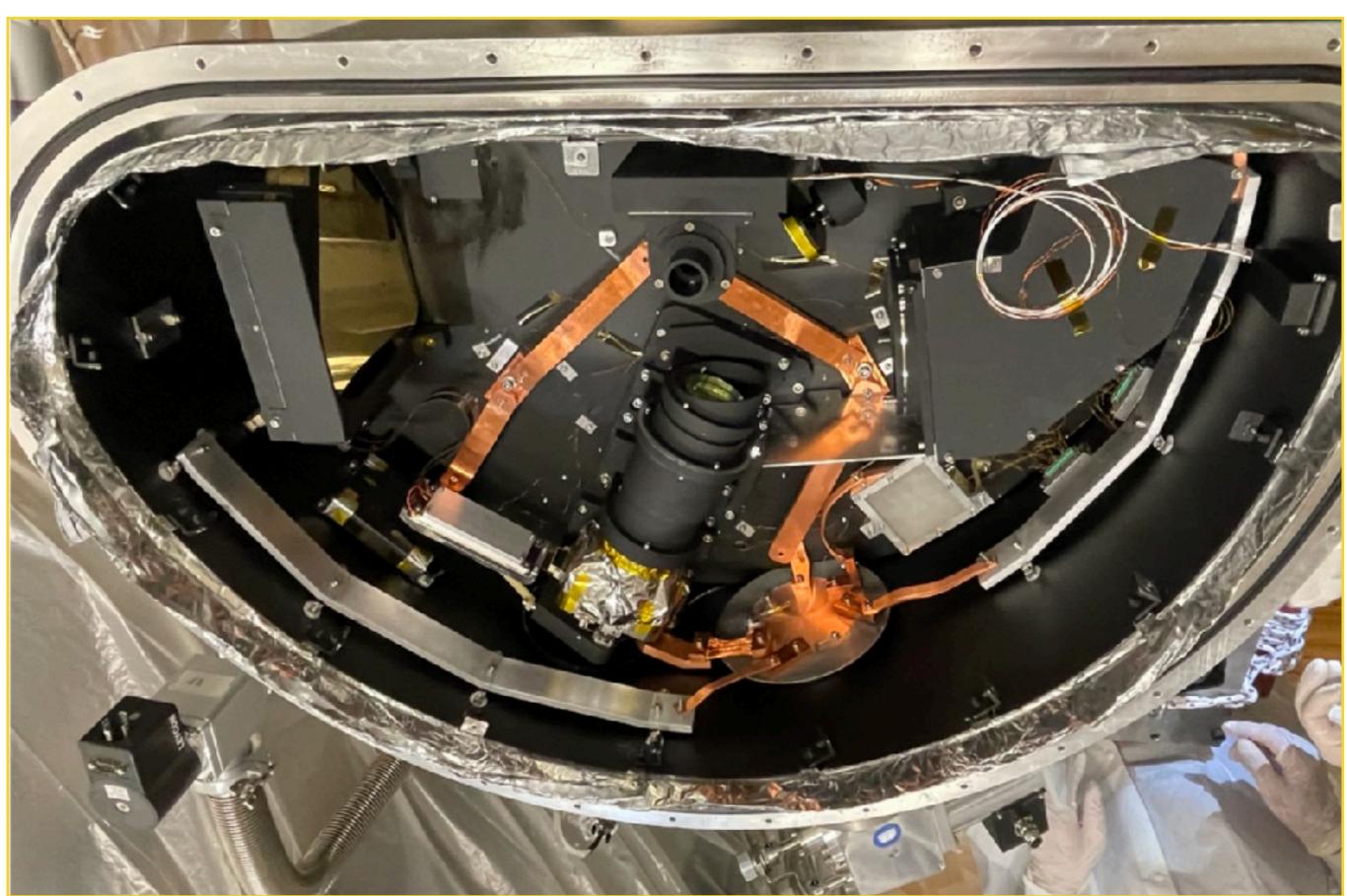
CFC principle: 1. LN2 storage tank, 2. De-pressurization valve, 3. Manometer, 4. Over-pressure valve, 5. Liquid nitrogen level gauge, 6. Vacuum insulated transfer line, 7. Radiation shield heat exchanger, 8. Cooling heat exchanger, 9. Gas heater, 10. Bypass valve, 11. Regulation valve.

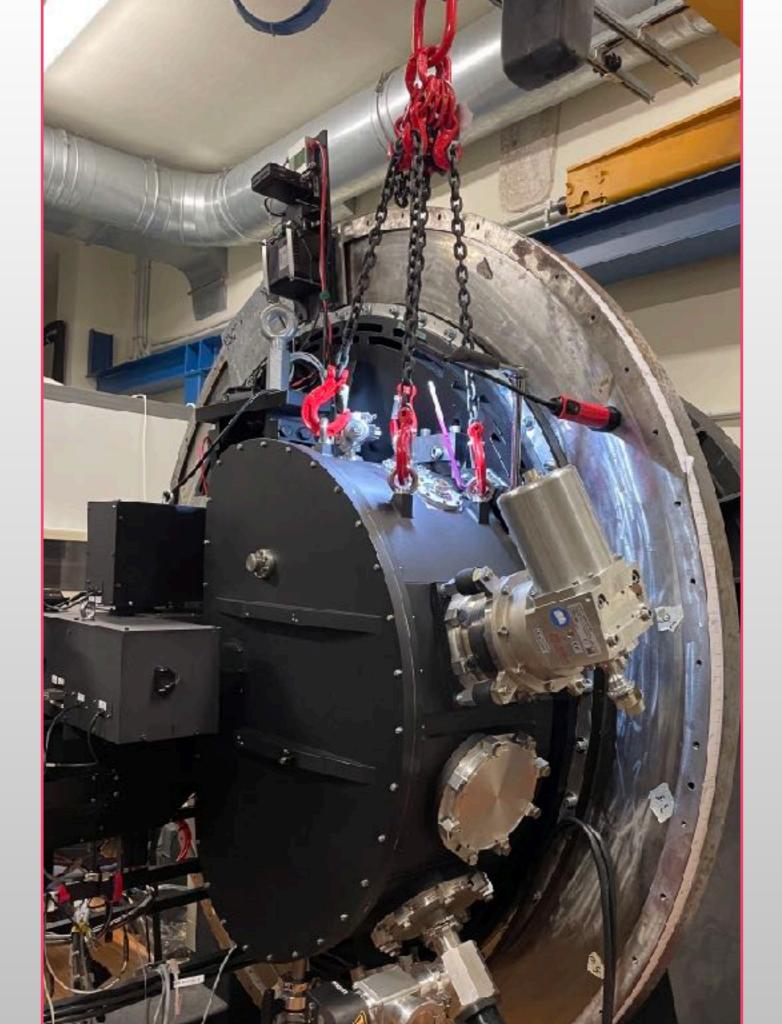


NIR Spectrograph

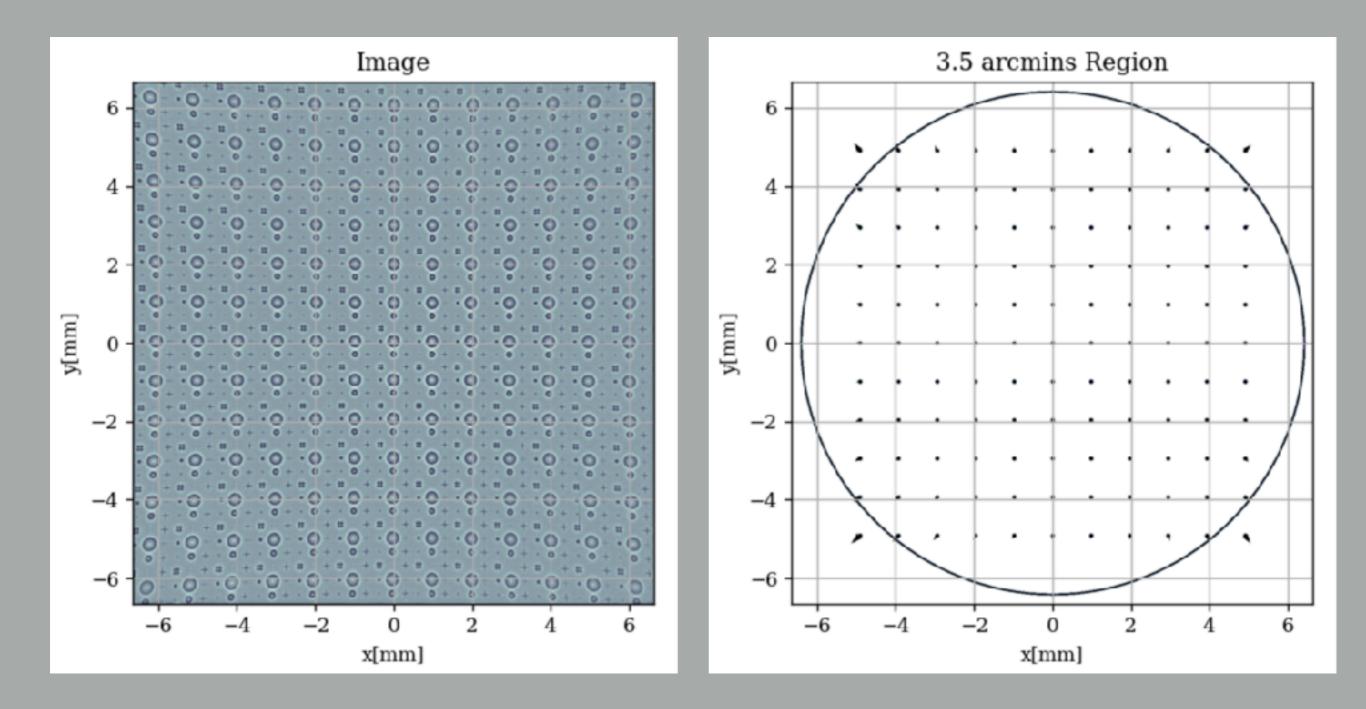


NIR Spectrograph

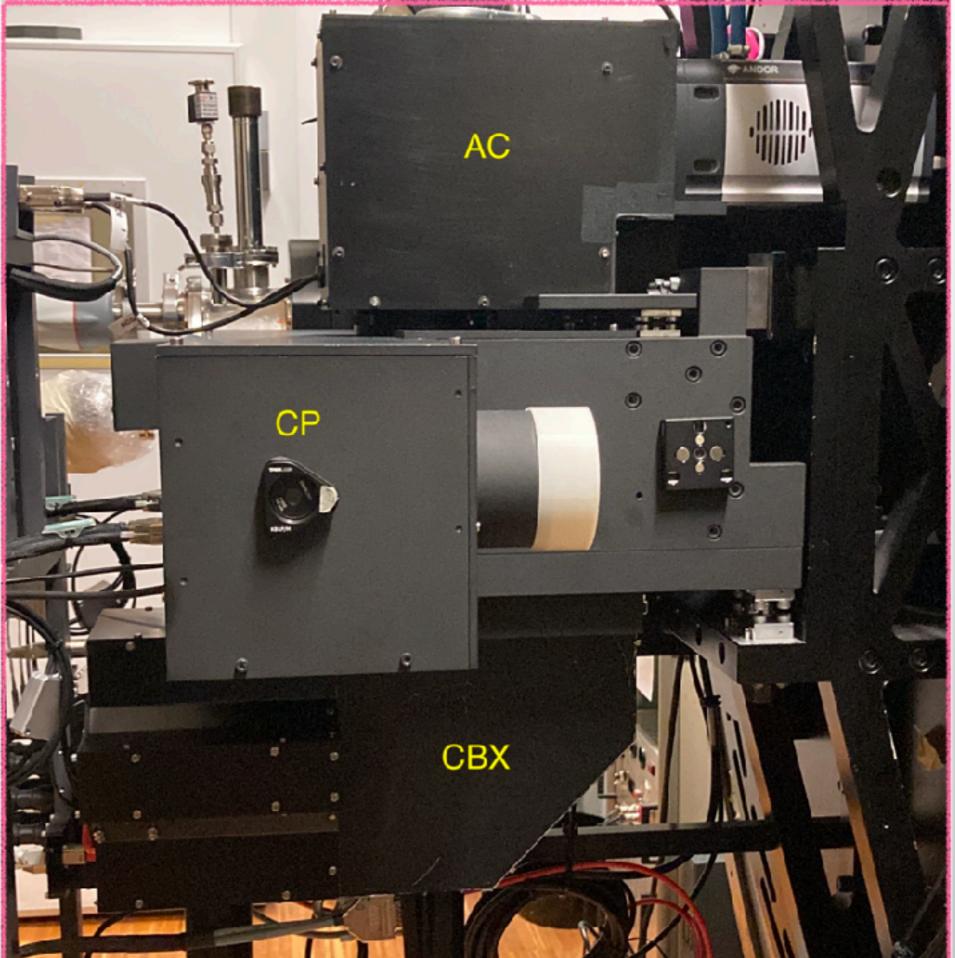




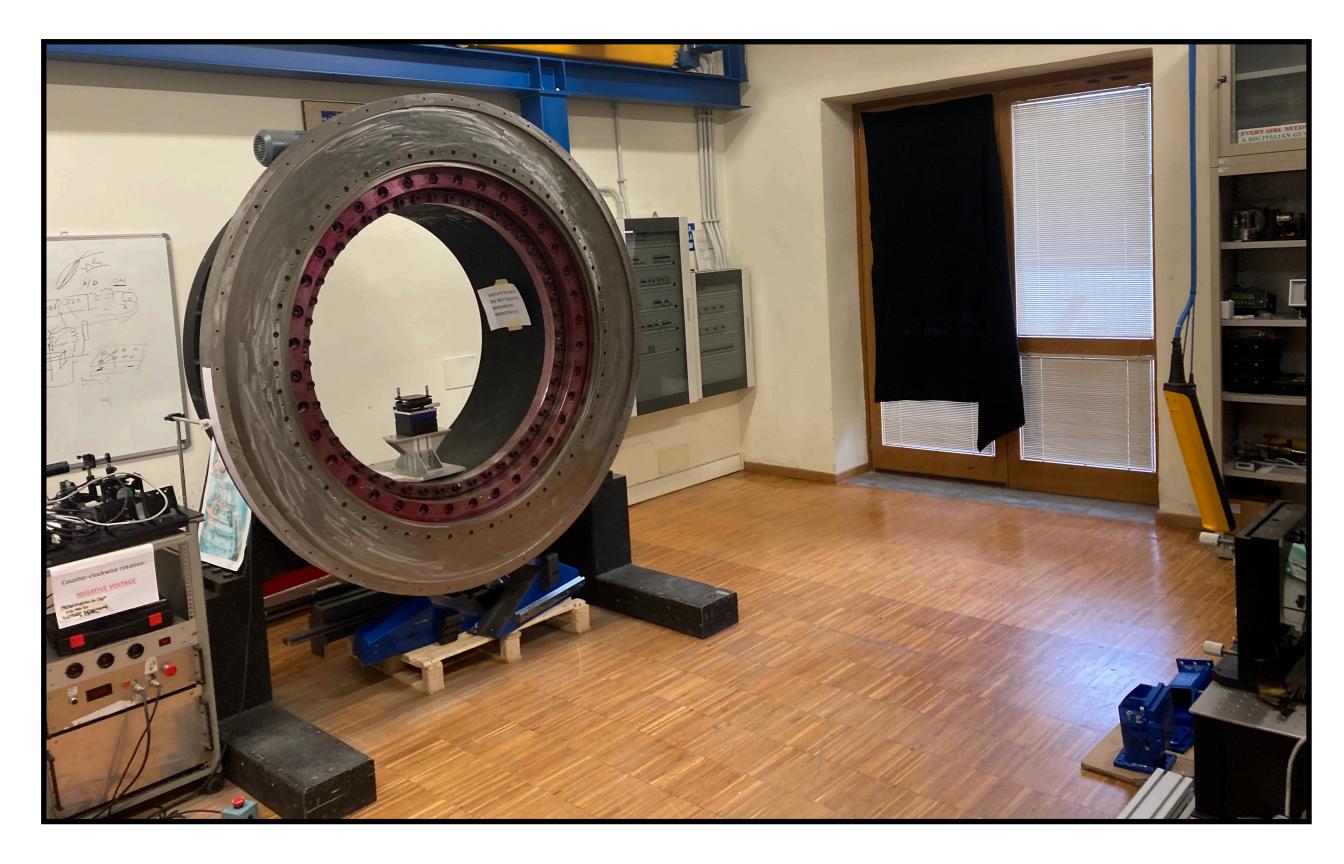
Acquisition Camera



Calibration Unit



Where We Started ...



Mounting the SOXS Flange



Corotator Holding Structure



Common Path Mounting



Mounting the Calibration Unit (old)



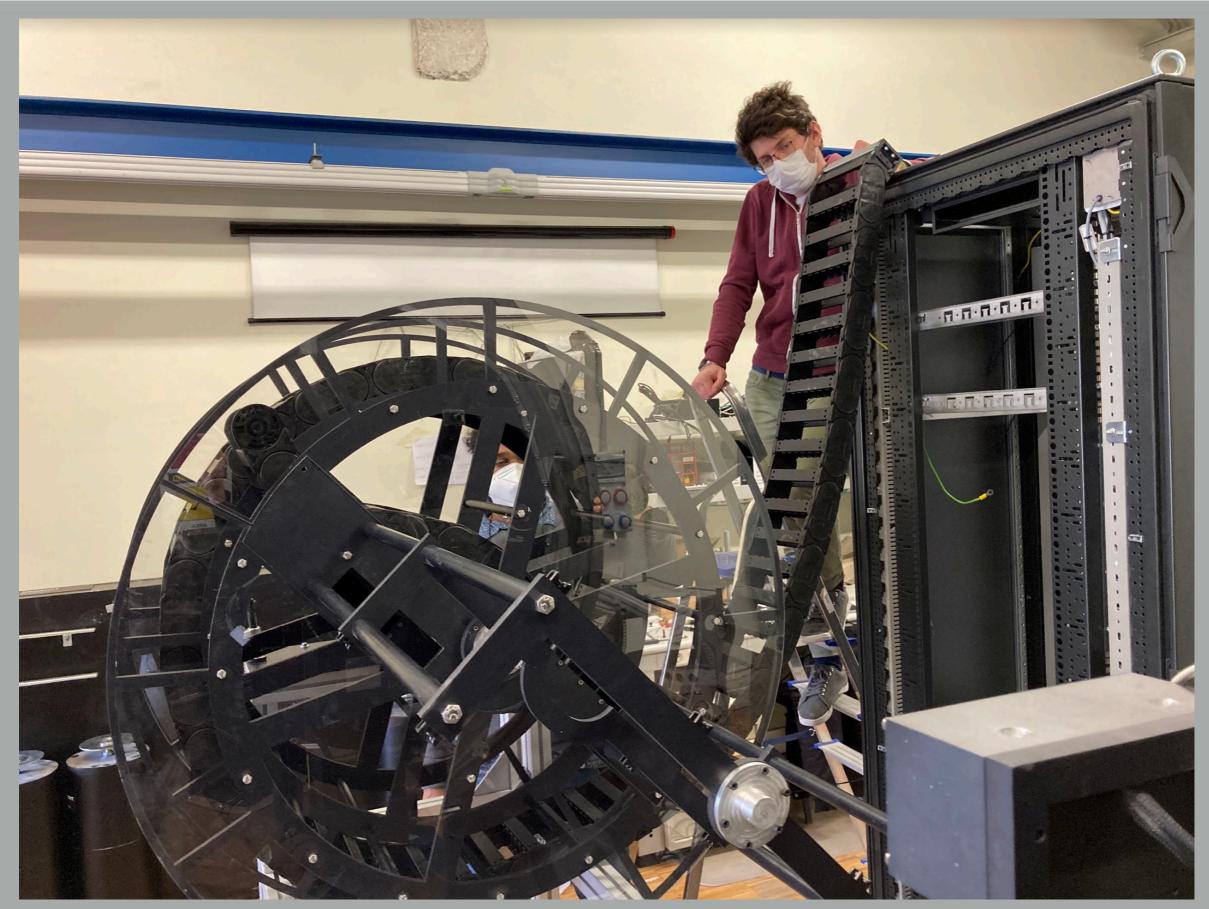




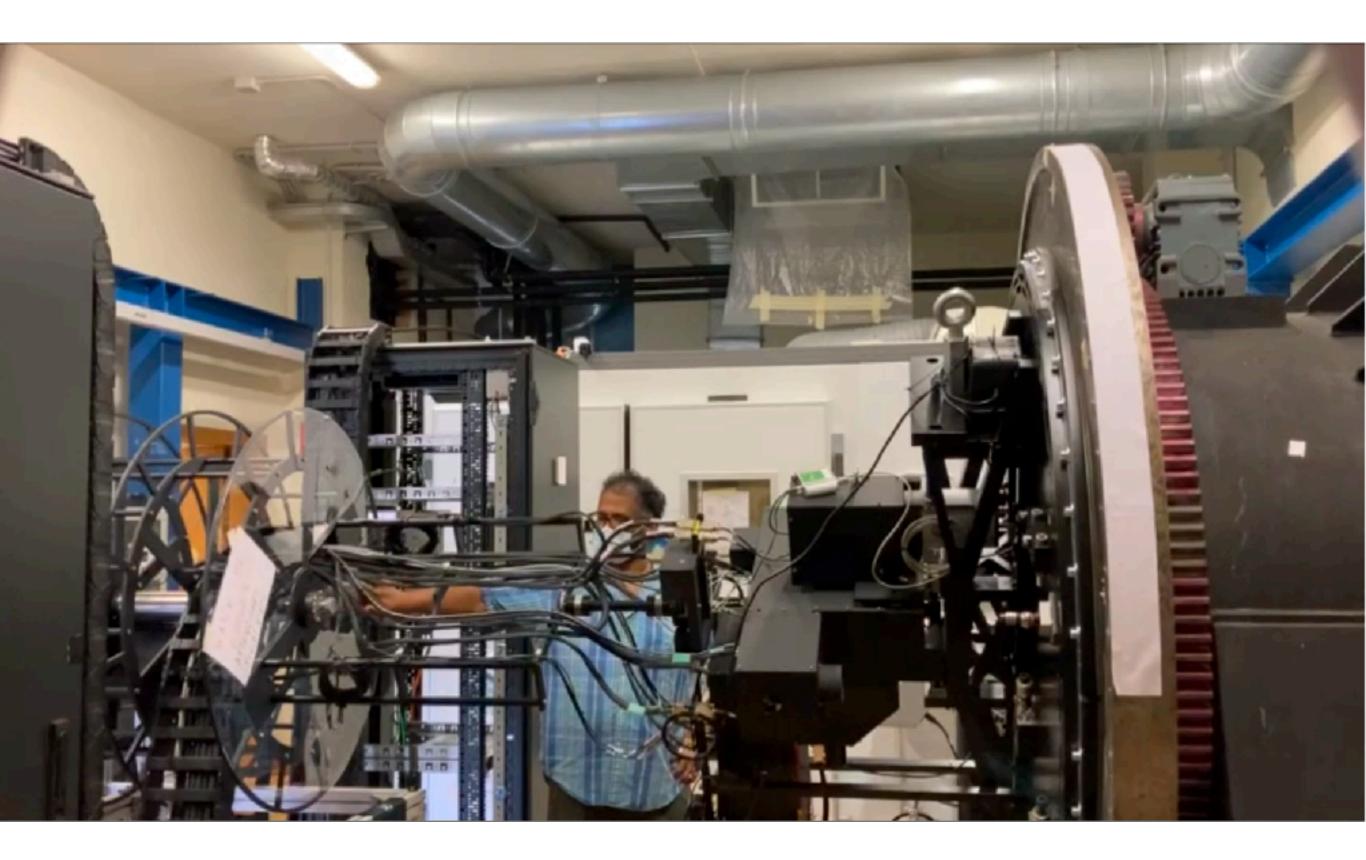
Mounting the Acquisition Camera



Cabinets and Cable Chains



Corotator co-rotating...



Finally everything together...



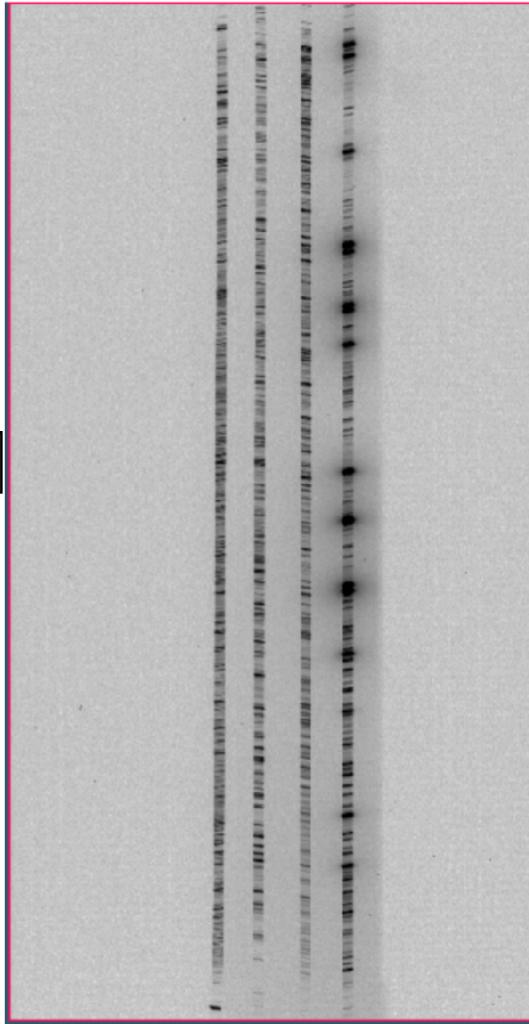
Now, some UVVIS spectra ...

1" slit, ThAr arc lamps with DIT of 300s

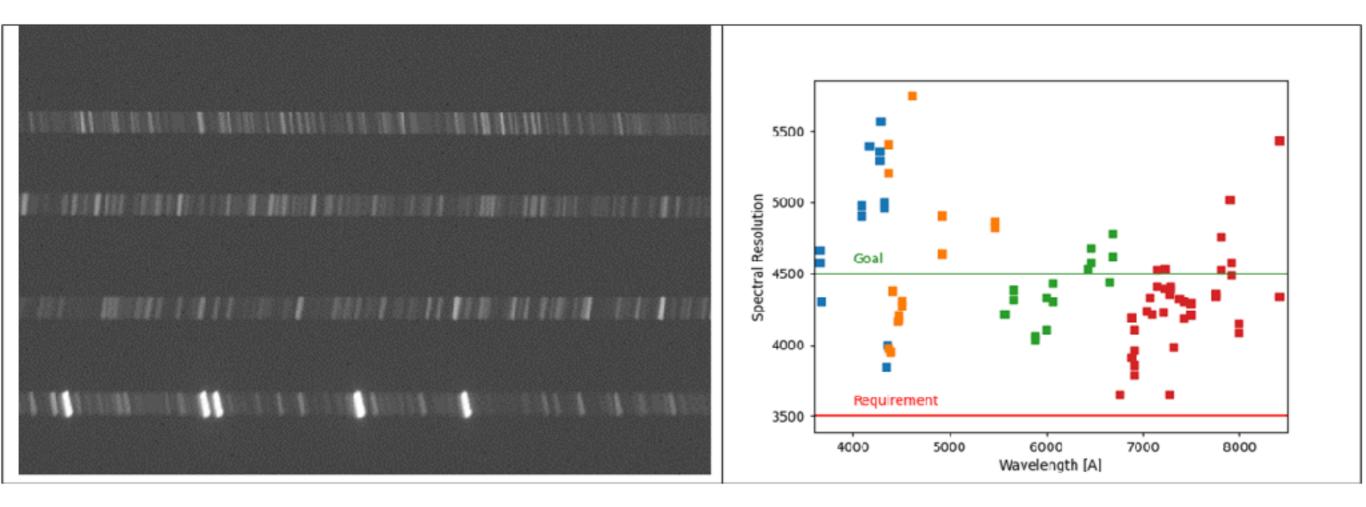
Multi-Pinhole, ThAr arc lamp with DIT of 300s

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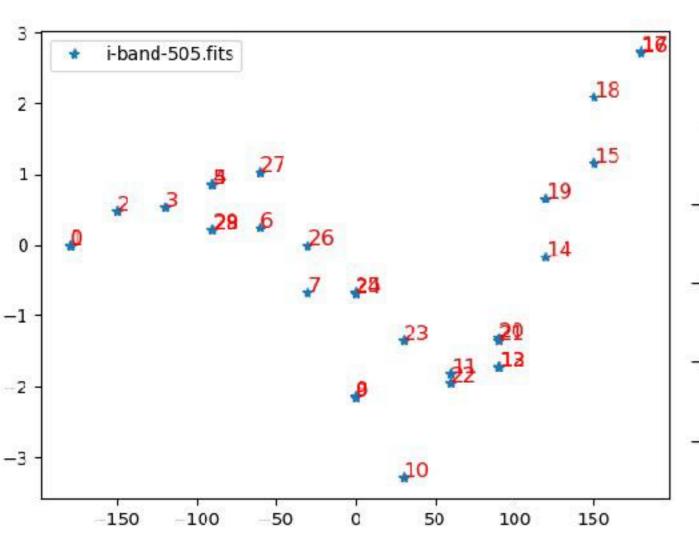
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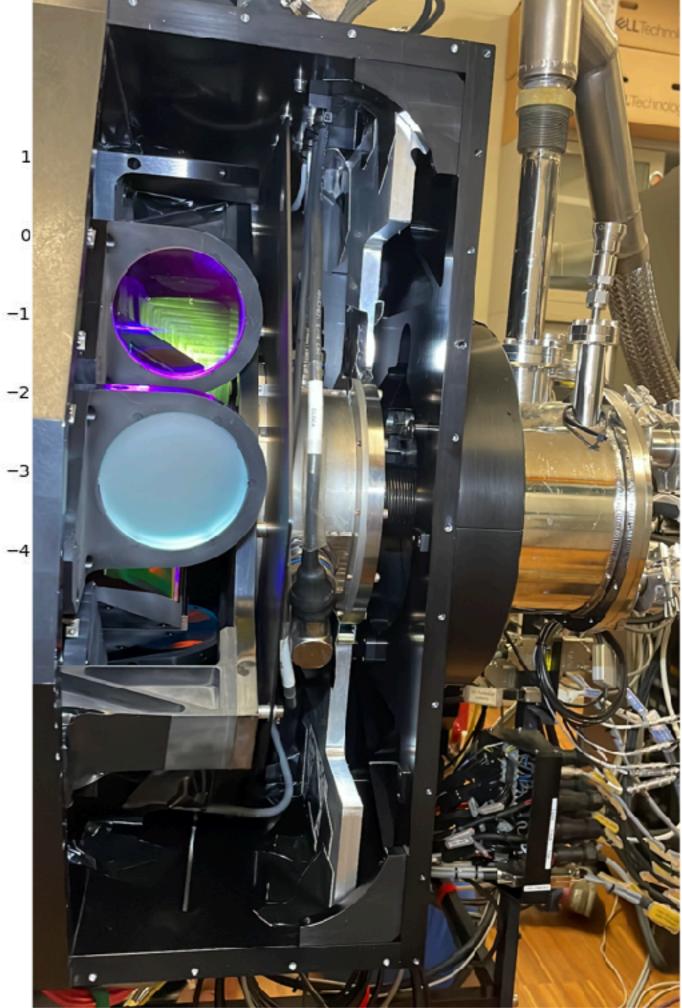


UVVIS Resolution

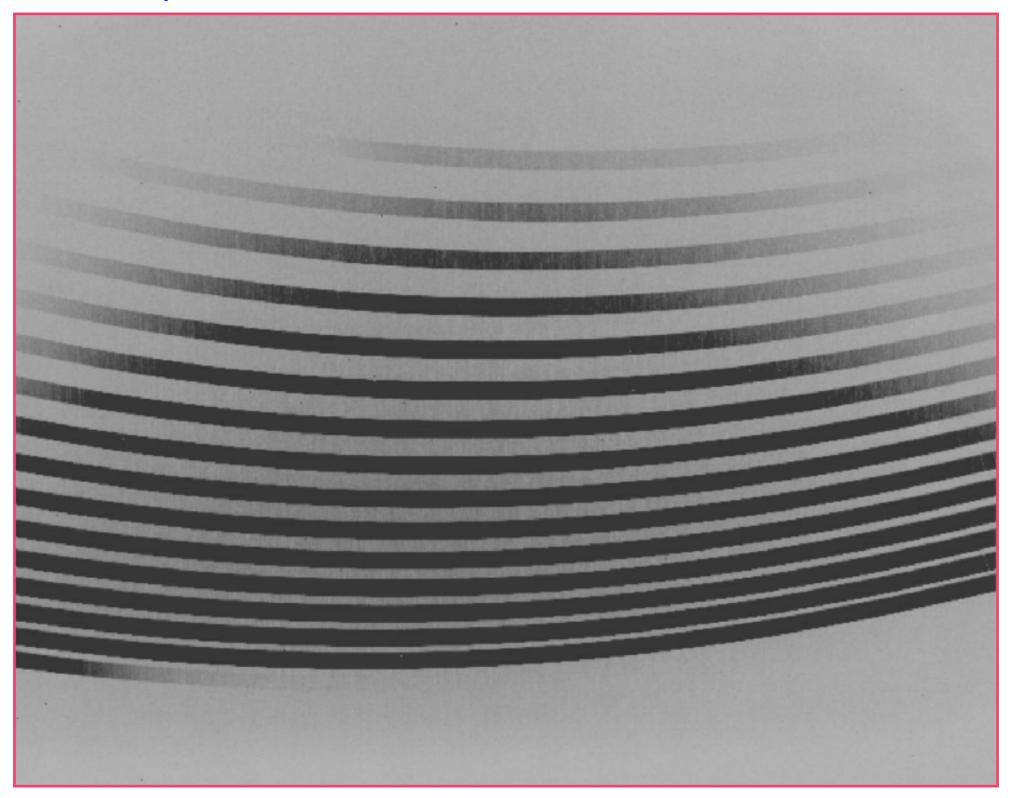


UVVIS Flexure



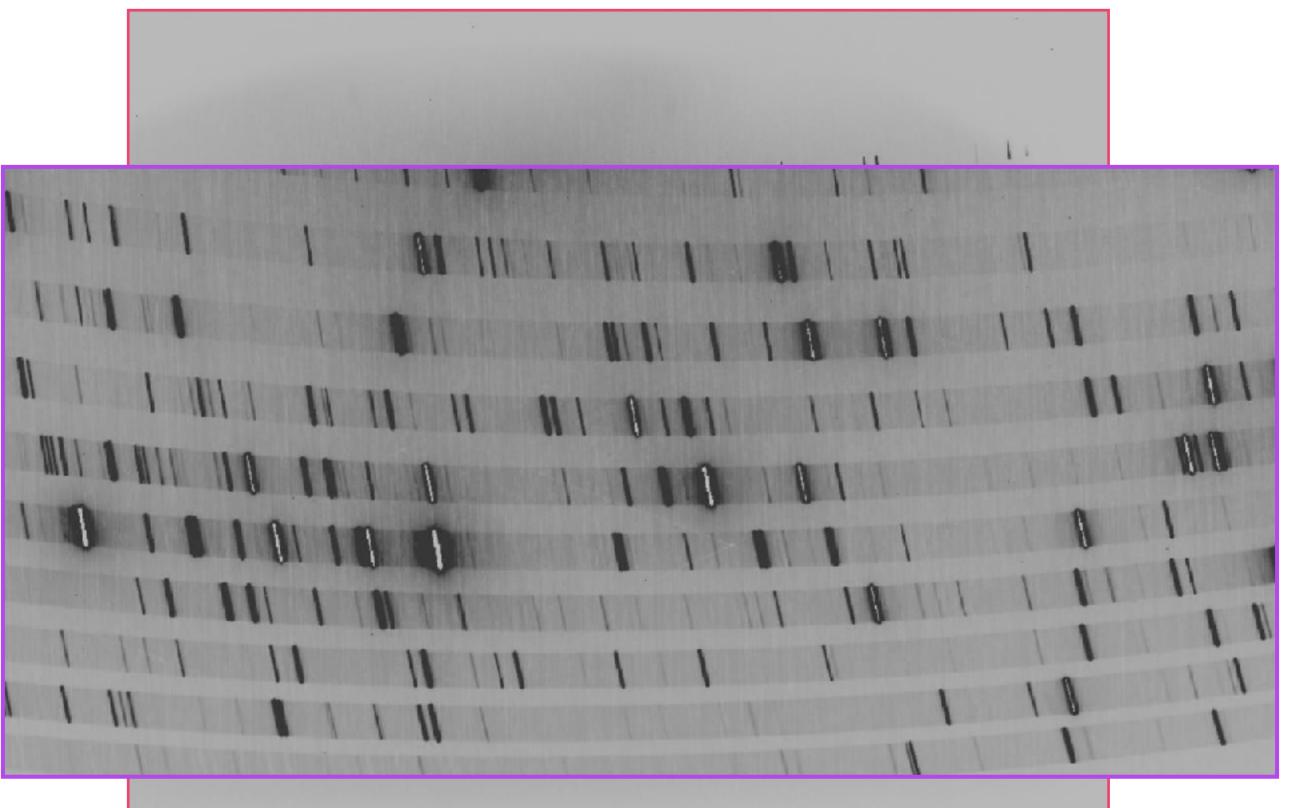


And NIR spectra ...



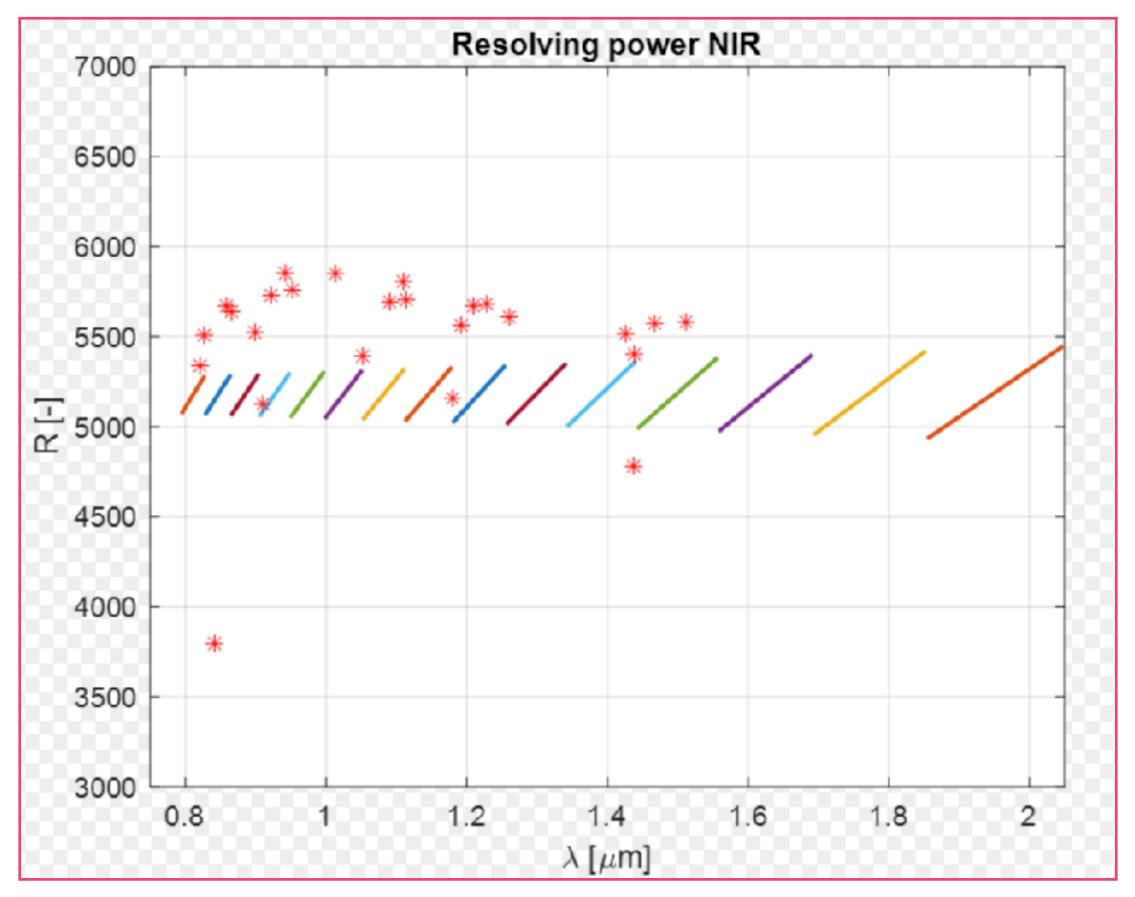
QTH continuum lamp with 0.5" slit and 5s exposure time

And NIR spectra ...

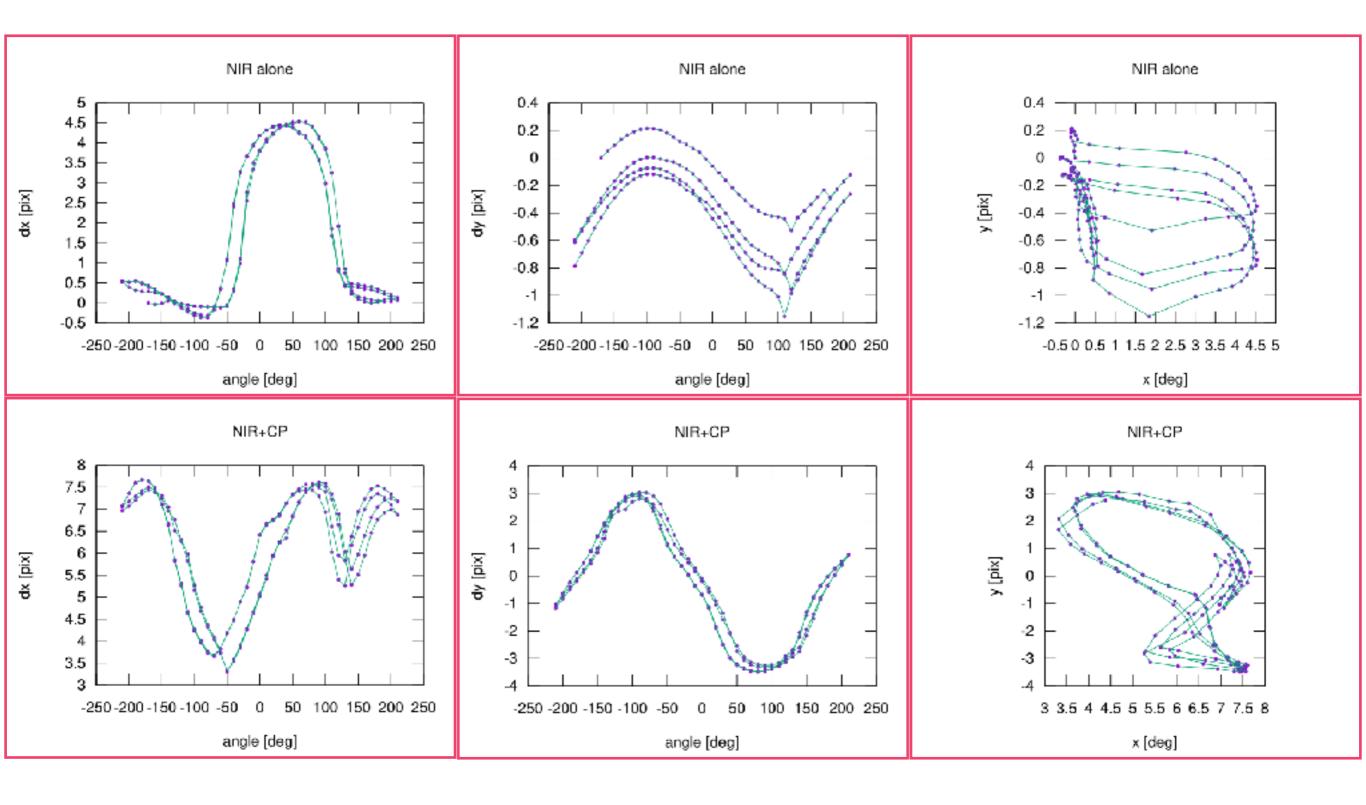


Arc lamps (Ar, Ne, Hg, & Xe) with 0.5" slit and 15s exposure time

NIR Resolution

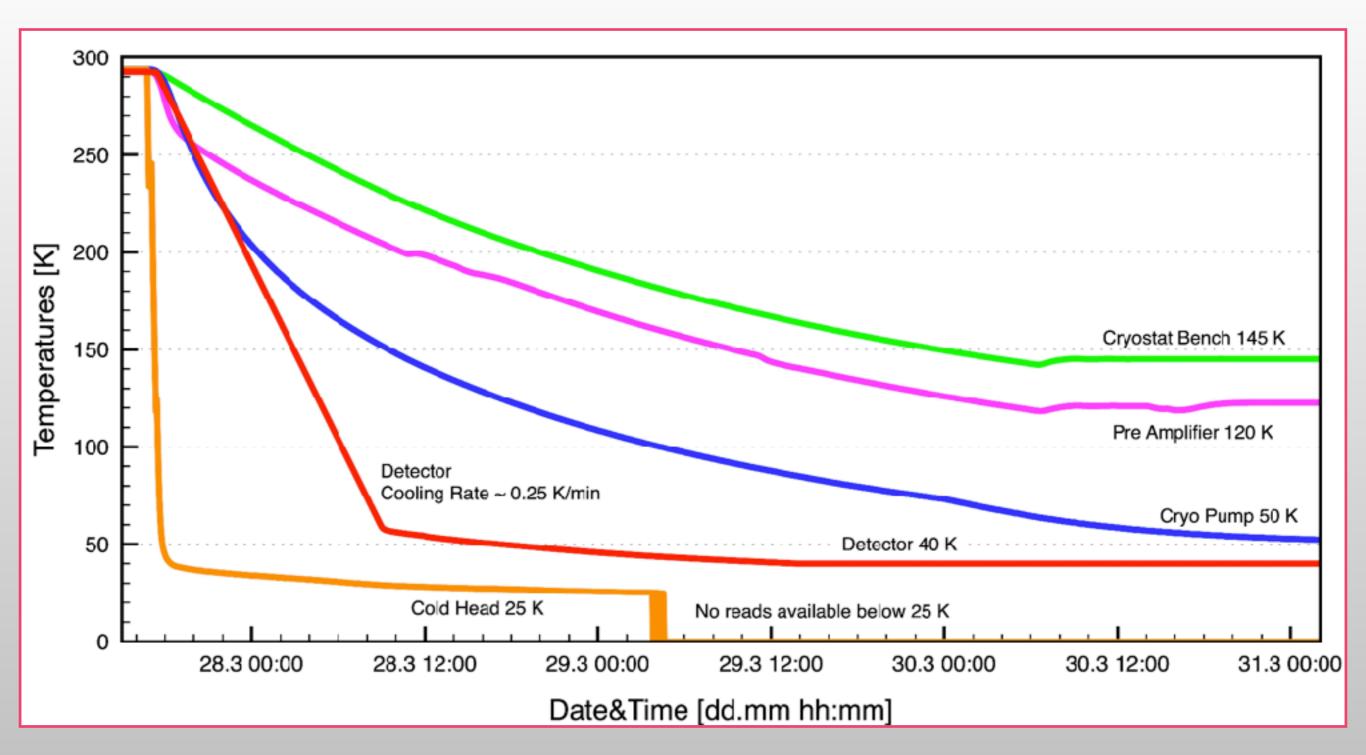


NIR Flexure

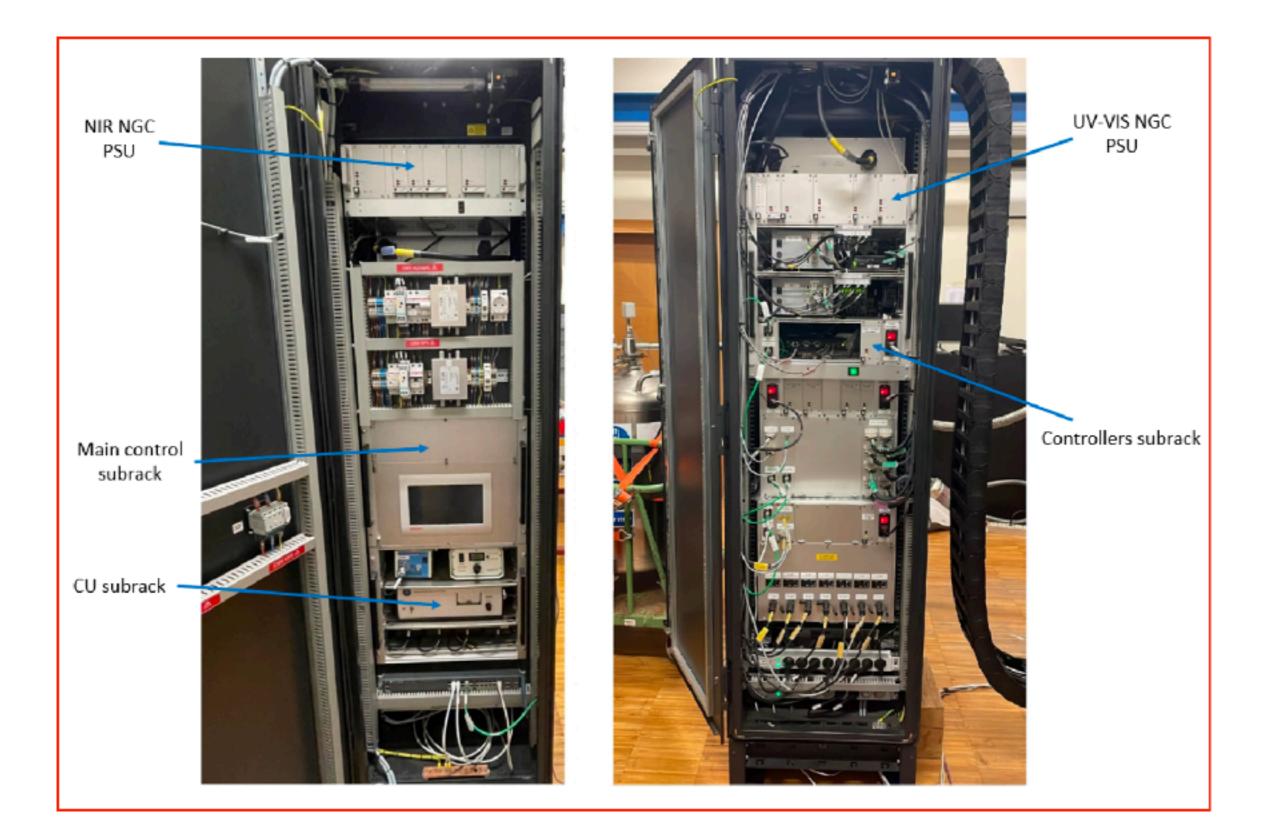


Measured flexure of the NIR alone (top) and NIR+CP (bottom)

Cryo-Vacuum System



Instrument Control Electronics



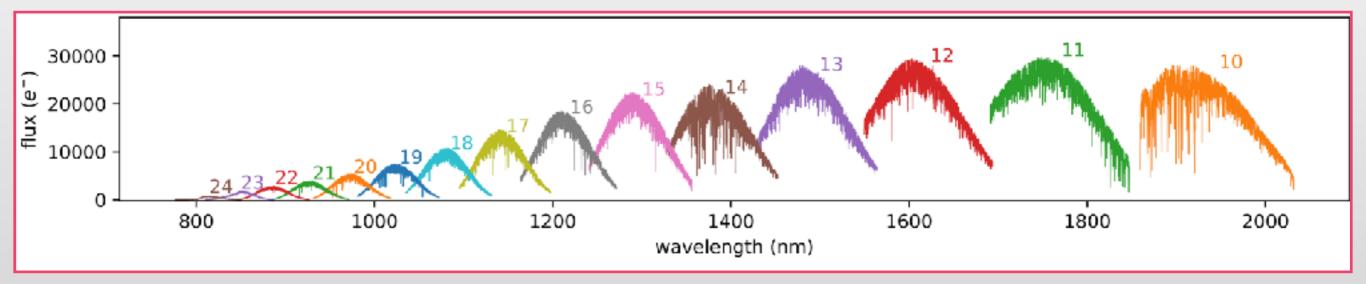
Instrument Control Electronics



Instrument Control Software

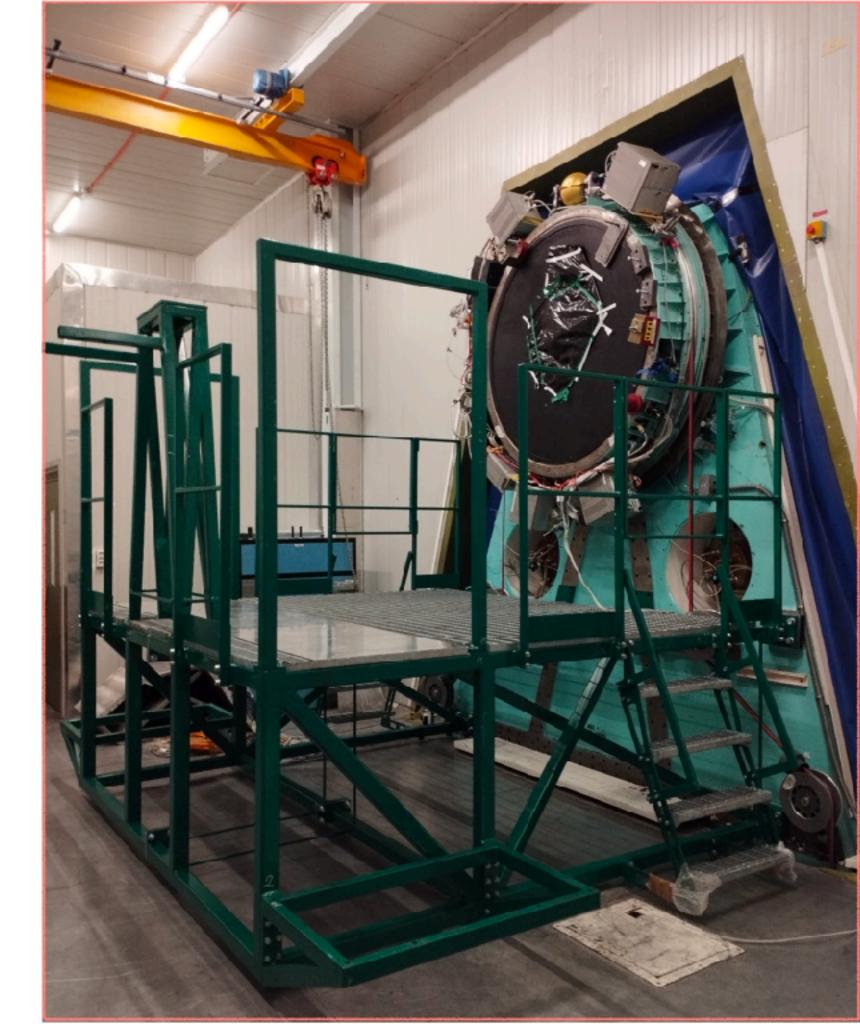
			SOXS Statu	is - @wsx	5				^ _ ×
File Std. Opt	tions								Help
SOXS S	tatus			NIR - N	lear Infrared	Spect	rograph		
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	Substate		setup		Remaining		2	NDIT	1
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ACQS ONLI	INE HW	Spect	roscopy 4.33	AFC1	ONLINE	Н₩		Mode	STAT
		- L			Current X	1000).0 Y	999.9	
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Data Reduction Software



All the spectral orders extracted from the NIR flat QTH calibration spectrum

SOXS Platform



AIV @ La Silla

Activity Name	Duration [hours]		0.2	0.3	0.4	D.5	D.6	D.7	D.8		D 10	D 11	D 12	D 13	0.14	D 15	D 16	D 1	0.11	D 19	D 20	D 21	D 22	D 23	D 24	D 25	D 26	0 27	D 28	0 29	D 30 C	0 31
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Flarge Position Repeatability Venfication	3	x		-	-			-	c	c	-	-	-	-	-		-		10	end				-	-		-	-	-	-	\rightarrow	-
Flange Position Nighttime Verification	4	<u> </u>	×					_	c	c	-	-	-	-	-	-	x	Dav	time A	-	,			-	-		-	-	-	-	-+	-
Electronics Cabinets Installation	8		x	×				_	c	c	-	-	-	-	+	-			nical			ted	-	-	-		-	-	-	-	\rightarrow	-
Corotator structure installation and testing	3	~	<u> </u>	÷.					~	c	-			-	-				tingen			NEU		-							\rightarrow	_
Instrument Workstation Installation	4	Ŷ		^				_	c	c	-	-	-	-	-	-	-	0.011	-	.,,			-	-	-		-	-	-	-	-+	-
INS network setup	4	x	×	-	-			_	c	c	-	-	-	-	+	+	-	-	+		-	<u> </u>	-	-	-		-		-	-	\rightarrow	-
Electronics and SW functional verification	4	<u> </u>	· ·					_	c	c	-	-	-	-	-		-	-	-	-	-	-	-	-	-		-		-	-	\rightarrow	-
Setting up interlock b/w corotator and NTT rotator	4		-	~	×			-	c	c	-	-	-	-	-		-	-	+	-	-	-	-	-	-		-		-	-	\rightarrow	_
Corotator Position Verification and Functioning	4			-	X			_	c	c	-	-	-	-	+	+	-	-	+	+		-	-	-			-		-		\rightarrow	
				-	A.		v	_	_	-	-	-	-		+		-	-	+				-	-			-		-	-	\rightarrow	
Setting up the alarm handling system	4						X	_	C	C	-	-	-		+		-	-			-	-	-	-			-				\rightarrow	
CP Alignment Verification	14	<u> </u>	×	к	<u>×</u>	~			C	C					-	-		-	-	-				-			_		_	_		_
CP Mounting to the Flange	2			-	X	X		_	C	C	-	-	-		+	+	-	-	+	-	-	-	-	-			-				\rightarrow	
CP Functional Verification	3		-	-	x	Y		_	с	с	-	-	-	-			-	-			-		-	-			-			-	\rightarrow	_
CP Nighttime Verification	4	<u> </u>				X			c	c	-										<u> </u>		-	-							\rightarrow	_
AC Alignment Verification	10	<u> </u>		ĸ	×	X		_	с	с	-	<u> </u>	-	<u> </u>			-	-			<u> </u>	<u> </u>	-	-						-	\rightarrow	_
AC Mounting to Flange	1	<u> </u>	-			X	X		c	C	-	<u> </u>	-	-		-	-	-		-			-	-					_	_	\rightarrow	
AC Functional Verification	3	<u> </u>	-			X	Х	X	_	С	_	L		L		<u> </u>	<u> </u>	-		<u> </u>	<u> </u>			L	 						\rightarrow	
CBX Mounting Lamps	4	L	<u> </u>	×	X				c	c					-	<u> </u>	<u> </u>		-	<u> </u>				L	<u> </u>						\rightarrow	
CBX Testing Lamps	3	<u> </u>				X				с	_				-	<u> </u>			-	<u> </u>												
CBX Mounting to Flange	1	<u> </u>					Х	×		c	_				-	<u> </u>			-	<u> </u>											\rightarrow	
CBX Functional Verification	6						Х	<u>x</u>	с	с																						
CP + AC + CBX Nighttime Testing	3								х																							
UVVIS: Integrating CCD chamber to the main unit	6									ж	×							С														
UVVIS: Stand-alone Alignmen: and Punctional Tests	3										×	х						C														
UVVIS: Warm Functional Tests with SOXS electronics	2											х						С														
UVVIS: Cryo-vacuum Functional Tests (without connecting to UVVIS)	6										х	Х						С														
UVVIS: Cryo-pump Regeneration	2											х	X					С														
UVVIS: Cryo-vacuum Functional Tests (Warm)	6											х	X					С														
UVVIS: Vacuum and cool down the spectrograph	4												х	x				С														
UVVIS: Verification of the Image Quality & Spectral Resolution	8	1												×				С														
UVVIS: Mounting to Flange	6	1													×			С														
UVVIS: Alignment to CP	3														×			С														
UVVIS: Functional Verification	2															х		С														
UVVIS: Vacuum & Cool down on the flange, cryo-vacuum tests	4	1														х	x	С														
UVVIS: Verification of Image Quality and Spectral Resolution on the flange	4															×	x	С														
UVVIS: Nighttime Verification	4										-				-			X		-												
NIR: Mourting on Flange	6	1																×	×							с						
NIR: Alignment to C? (using mechanical references)	4	1																	×	X						с					\neg	
NIR Cryo-Vacuum Testing (without connecting to HW)	4	1													-			×	-							c					\rightarrow	-
NIR: Cryo-pump Regeneration	2	1-													-	1			×	x						c					\rightarrow	-
NIR: Warm Functional Tests (including cryo-vacuum)	8	1-									-				-	<u> </u>		\vdash	-	x						c				-	\rightarrow	-
NIR: Cooldown the vessel	54	1-									-			-	-	<u> </u>		\vdash	-	x	x	x	Y			c				-	-+	-
NIR: Optical Alignment Verification	8	1-	-								-		-	-	-	-			-	-	~	Ŷ	Y			c				-	\rightarrow	-
NIR: Verification of the Image Quality and Spectral Resolution	16	1	-							-	+				-	-			+	-		-	Y	x	×.	c				-	\rightarrow	-
NIR: Nighttime testing	4	1									-				-	<u> </u>		\vdash	-	-			-	~	^	v				-	\rightarrow	
Verification of the position of the TT mirrors for various NTT derotator angles	- ·	1	-		-						-	-	-	-	-	-	-	-	-	-		<u> </u>	-	-	-				-	-	\rightarrow	-
compensating the flexure	16								х	x																x	x					
Verification of the UZVIS and NIR image quality and Spectral resolution for various derotator angles using CRX light sources	16																										x	x	x			
Full system Electronics and SW tests (including testing some OBs for repeatability											\vdash				\square				\square												\neg	\neg
and reliability)	40	<u> </u>		-	-						-	×	×	-		×	×				X		-						×			
Training the ESO staff for preventive and corrective maintenance - Part 1	16	<u> </u>	-						ж	×		<u> </u>	-	<u> </u>	-		<u> </u>	×		-				-	-					×	×	
Training the ESO staff for preventive and corrective maintenance - Part 2	16	<u> </u>	-						х	×					-	<u> </u>	<u> </u>	Х		<u> </u>				L	<u> </u>					х	×	
Full System Nighttime testing before the Commissioning	32												1			1				1										X	X	x

SOXS timeline

Date	Activity	
June 2014	ESO Call for New Instruments at NTT	~
February 2015	Proposal Submission	\checkmark
May 2015	SOXS selected by ESO (out of 19 proposals)	\checkmark
October 2016 - July 2017	INAF approval + PDR phase	\checkmark
August 2017 - Sept 2018	FDR Phase	\checkmark
October 2018 - February 2022	Procurement, Sub-system AIVT (delay due to Covid-19)	\checkmark
March 2022	Integration Started	\checkmark
March 2022 - July 2024	Integration & System level tests	\checkmark
July 2024	Preliminary Acceptance Europe start	 Image: A second s
November 2024	PAE complete and start packing	\checkmark
December 2024	Shipping SOXS to La Silla, Chile	\checkmark
January 2025	AIV @ La Silla	\checkmark
March 2025	Commissioning Start	 Image: A second s



The SOXS Instrument @ INAF-OAPD



Name	Initials	Institution	D 1	D 2	n 3	D 4	n 5	D 6	n z	D 8	D 9	D 10	D 11	D 12	0.12	0.14	n 15	D 16	D 17	0.18	n 19	n 20	D 21	D 22	D 22	D 24	n 25	D 26	B 27	n 28	D 29	D 30	D 31	n 22	n 22	D 34 I	0.351	Costal
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Kalyan Radhakrishnan		INAF-Padova	X	X	X	X	X	X	X	X	×	x	x	Х	х	-			<u> </u>															X	X	X	X	17
	FBA	INAF-Padova	X	х	X	X	X	X	X	х	-	<u> </u>	-			<u> </u>		<u> </u>	<u> </u>						Х	Х	х	X	х	X	X	х	X	x	x	X	х	21
	SDF	INAF-Padova	X	X	×	×	X	Х	X	X							-	L													\vdash		\square	\vdash	\vdash	$ \rightarrow $	_	8
	RCL	INAF-Padova									×	X	Х	X	X	X	×	X	Х														\square	\square		$ \rightarrow $		9
	LCA	INAF-Padova														x	×	X	X	X	x	×	x	X	X									\square		$ \longrightarrow $		10
Davide Ricci	DRI	INAF-Padova	X	X	X	X	X	X	X	X	X	X	X										X	X	X	х	X	X	X	X	X	X	X					22
SoftWare Person	SWP	INAF-Padova												X	X	x	X	X	X	X	x	X												х	X	X	X	13
Matteo Aliverti	MAL	INAF-Brera	X	X	X	×	X	X	X							X	X	X	X	X	X	×	X	X														16
Matteo Genoni	MGE	INAF-Brera																		X	x	×	X	Х	X	x	x	×	X	X	X	х	X					14
Sergio Campana	SCA	INAF-Brera																																x	×	X	x	4
Paolo D'Avanzo	PDA	INAF-Brera																																х	X	×	х	4
Pietro Schipani	PSC	INAF-Napoli	X	X	X	×	X	X	X	×	×	X													X	X	X	x	Х	X	X	х	X	X	X	X	х	23
Mirko Colapietro	MCO	INAF-Napoli	Х	X	х	×	х	X	X																													7
Sergio D'Orsi	SDO	INAE-Napoli	X	×	X	×	X	х	X																													7
Salvatore Savarese	SSA	INAE-Napoli																							X	X	X	X	X	Х	X	X	X	x	x	X	X	13
Guilio Capasso	GCA	INAF-Napoli																							X	х	x	x	X	Х	X	х	X	x	X	X	X	13
Fabrizio Vitali	FVI	INAF-Rome														×	×	X	х	X	x	×	x	Х														9
Francesco D'Alessio	FDA	INAF-Rome														x	X	x	Х	X	х	x	х	X														9
Salvatore Scuderi	SSC	INAF-Milan											X	X	X	X	X	X	X	X	×	×	X	X														12
Antonio Miccichè	AMI	INAF-Catania														x	×	X	х	X	x	×	x	х														9
Sagi Ben-Ami	SBA	Weizmann								x	X	X	х	Х	x																							6
Adam Rubin	ARU	Weizmann								x	×	X	X	X	X																							6
Ofir Hershko	OHE	Weizmann								х	x	X	х	X	X																							6
Rosario Cosentino	RCO	TNG									×	X	х	Х	X																							5
Rachael Bruch	RBR	Weizmann																								x	X	×	X	Х	X	х	X	x	X	×	X	12
Guiliano Pignata	GPI	Universidad An																										×	X	Х	X	X	X					6
		Total/day	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	7	7	7	8	8	8	8	8	8	9	9	9	9	
																														Total	num	ber d	ays fe	or all	people	e on-s	ite	281



- SOXS is a Single object spectrograph offering simultaneous spectral coverage in UV-VIS and NIR, with imaging capabilities in the visible.
- It will be a precious facility for the spectroscopic followup of transient sources.
- SOXS installation, commissioning, and science verification will be in the year of 2024.

Mounting the UVVIS Spectrograph on the SOXS Flange



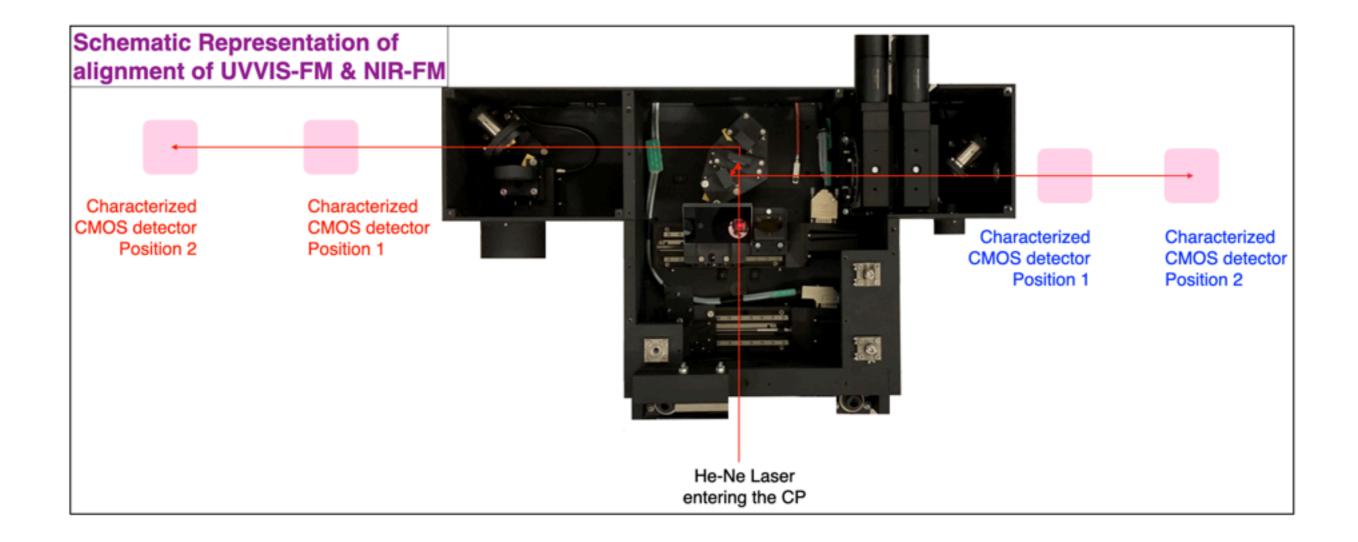
Thanks

SOXS Consortium

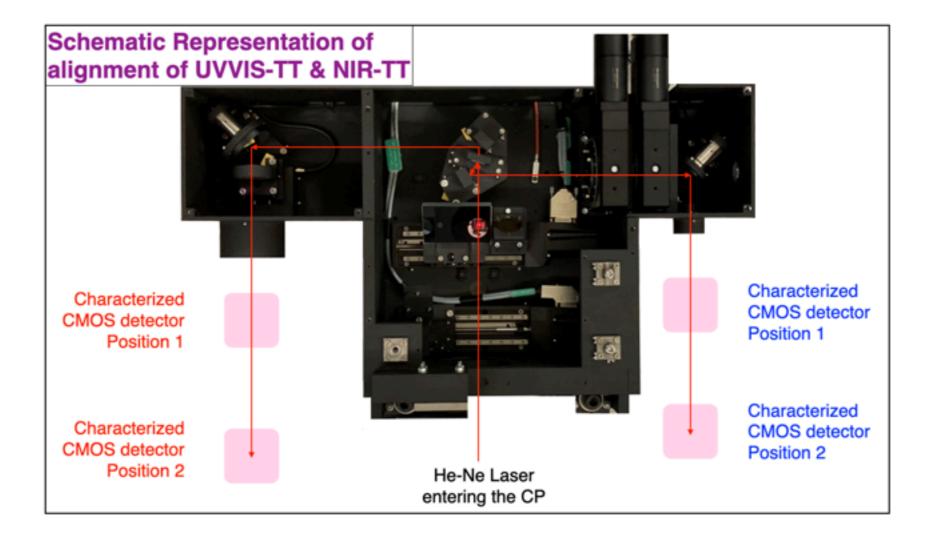
- INAF Common Path subsystem, NIR-spectrograph, control software and electronics, vacuum and cryogenics, detector control.
- Weizmann Institute of Science -UV-VIS spectrograph
- Universidad Andres Bello & Instituto Milenio de Astrofisica -Acquisition Camera sub-system
- Turku University Calibration Unit sub-system
- Queen's University -Data reduction



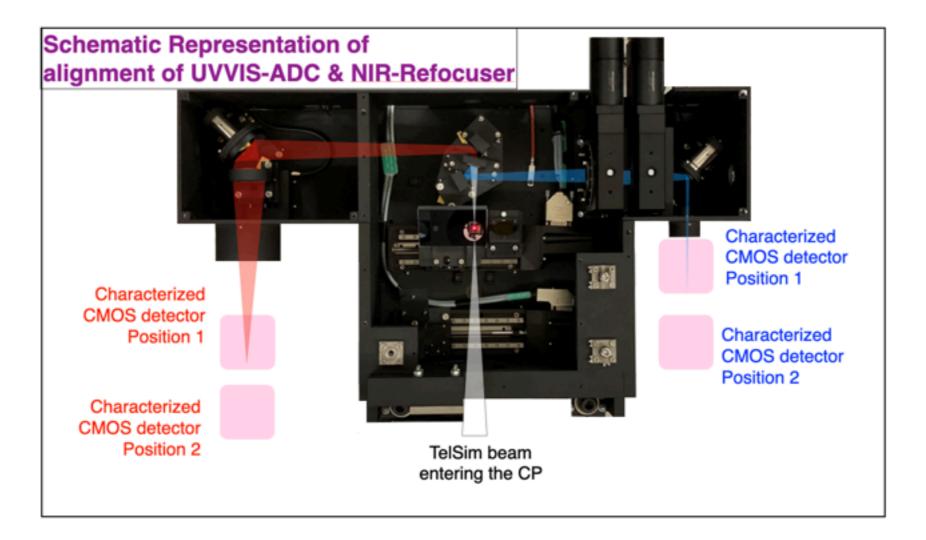
Common Path Alignment



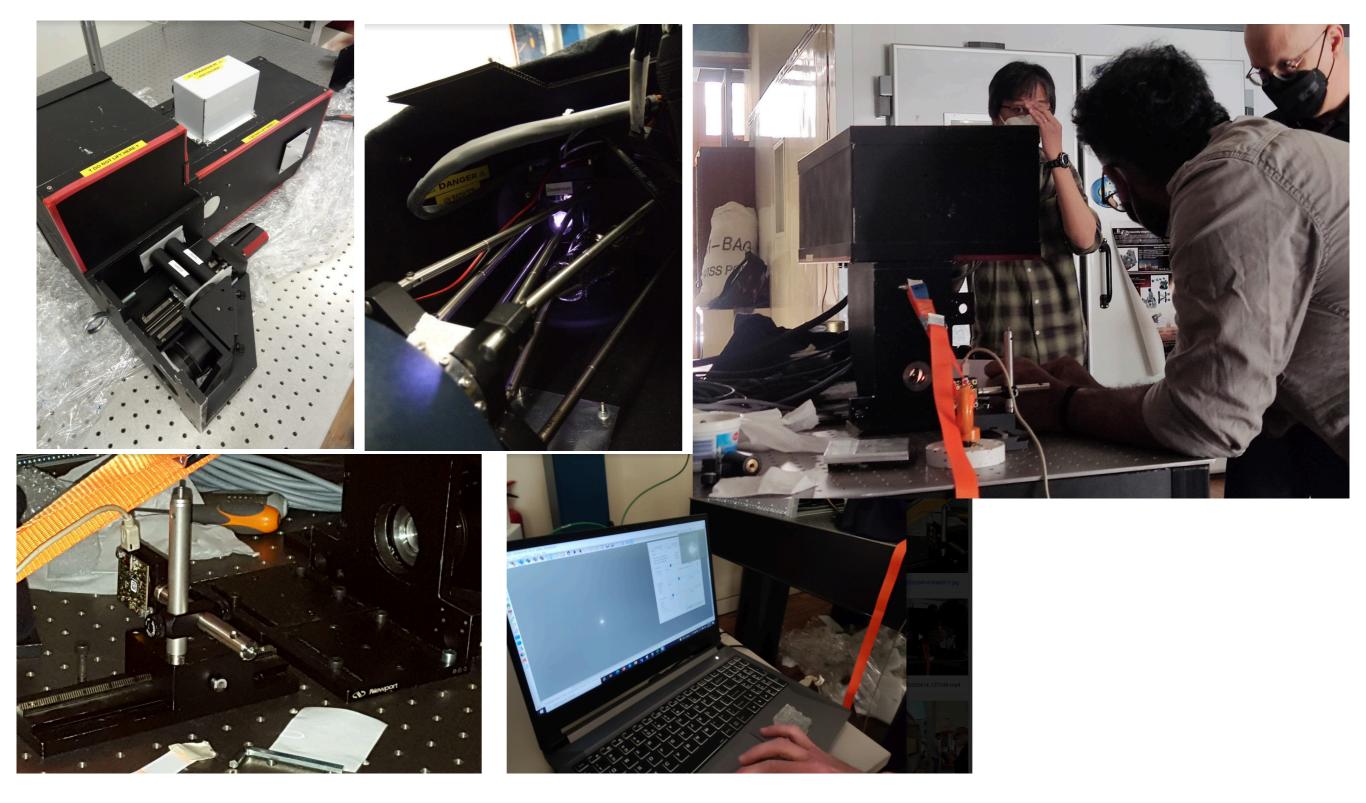
Common Path Alignment



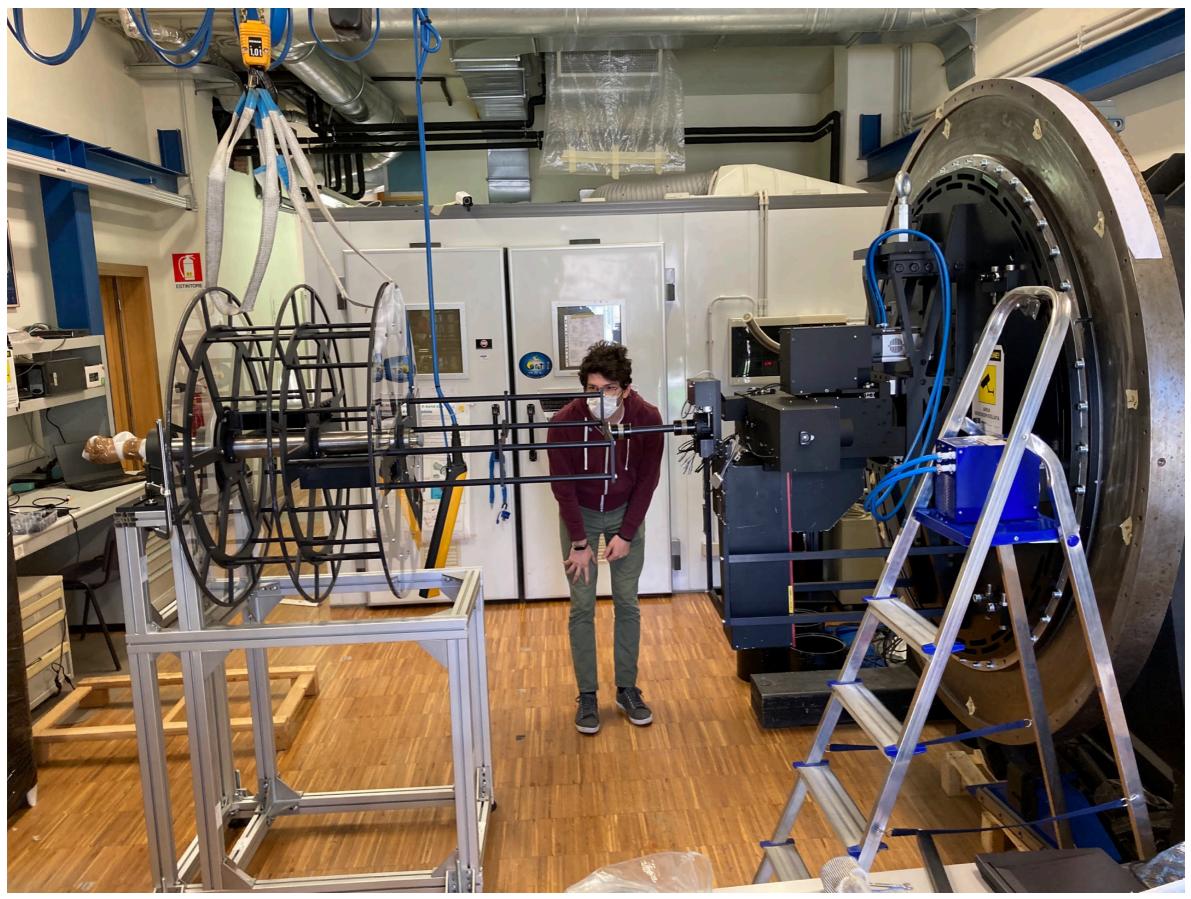
Common Path Alignment











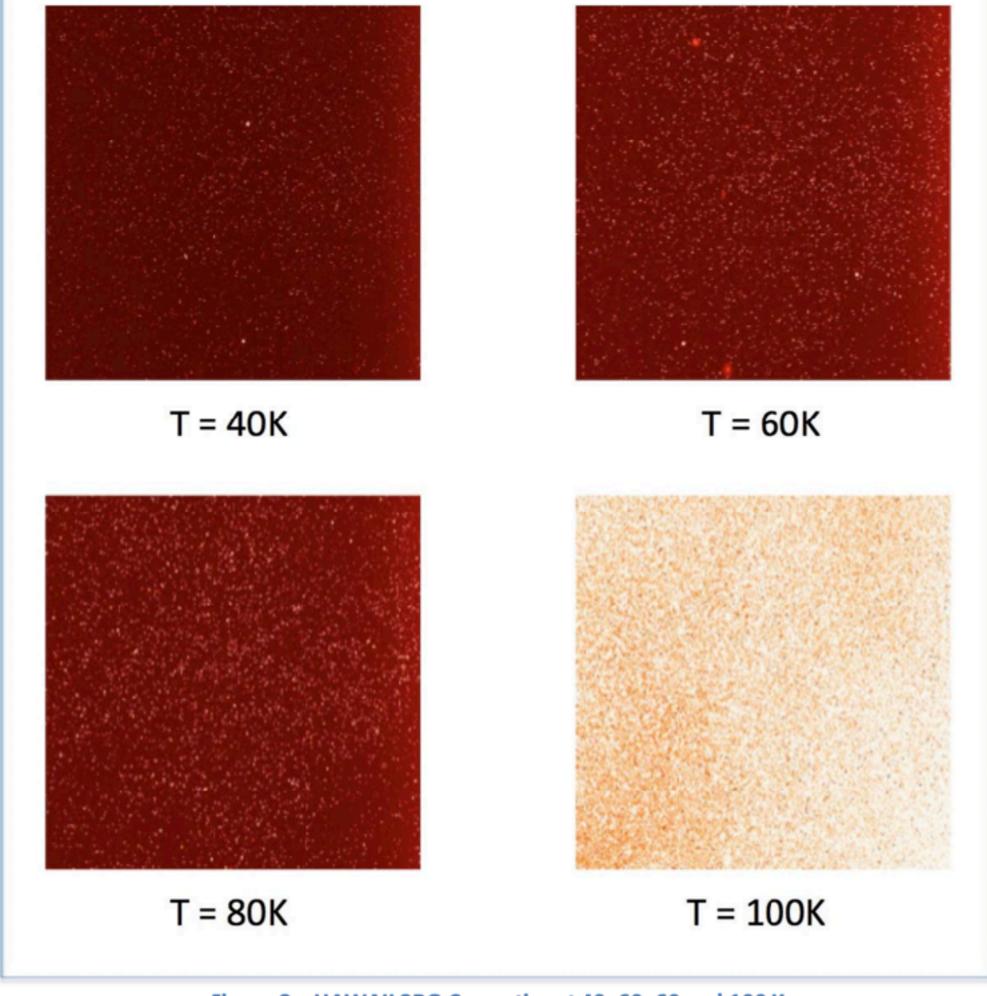


Figure 2 – HAWAII 2RG Cosmetics at 40, 60, 60 and 100 K

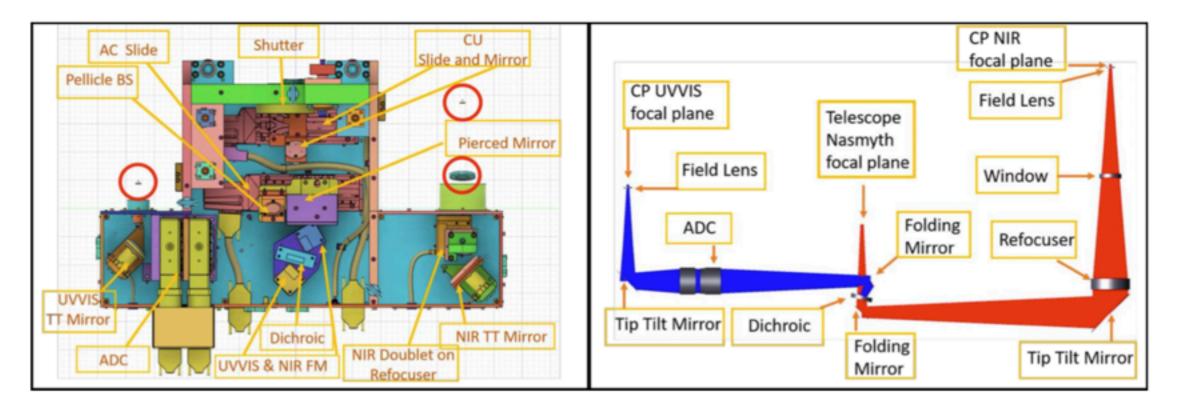


Figure 1. Left panel: Common Path CAD image displaying its components. Right panel: The CP light path.

The optical components UVVIS field lens, NIR window, and NIR field lens (marked within red circles in Figure 1 are formally a part of the CP, but physically present within the spectrographs. Without these components, the CP produces an F/6.91 beam at the UVVIS CP exit and an F/6.8 beam at the NIR CP exit.

History (more recent)



ESO call for new instruments at NTT (06/2014)

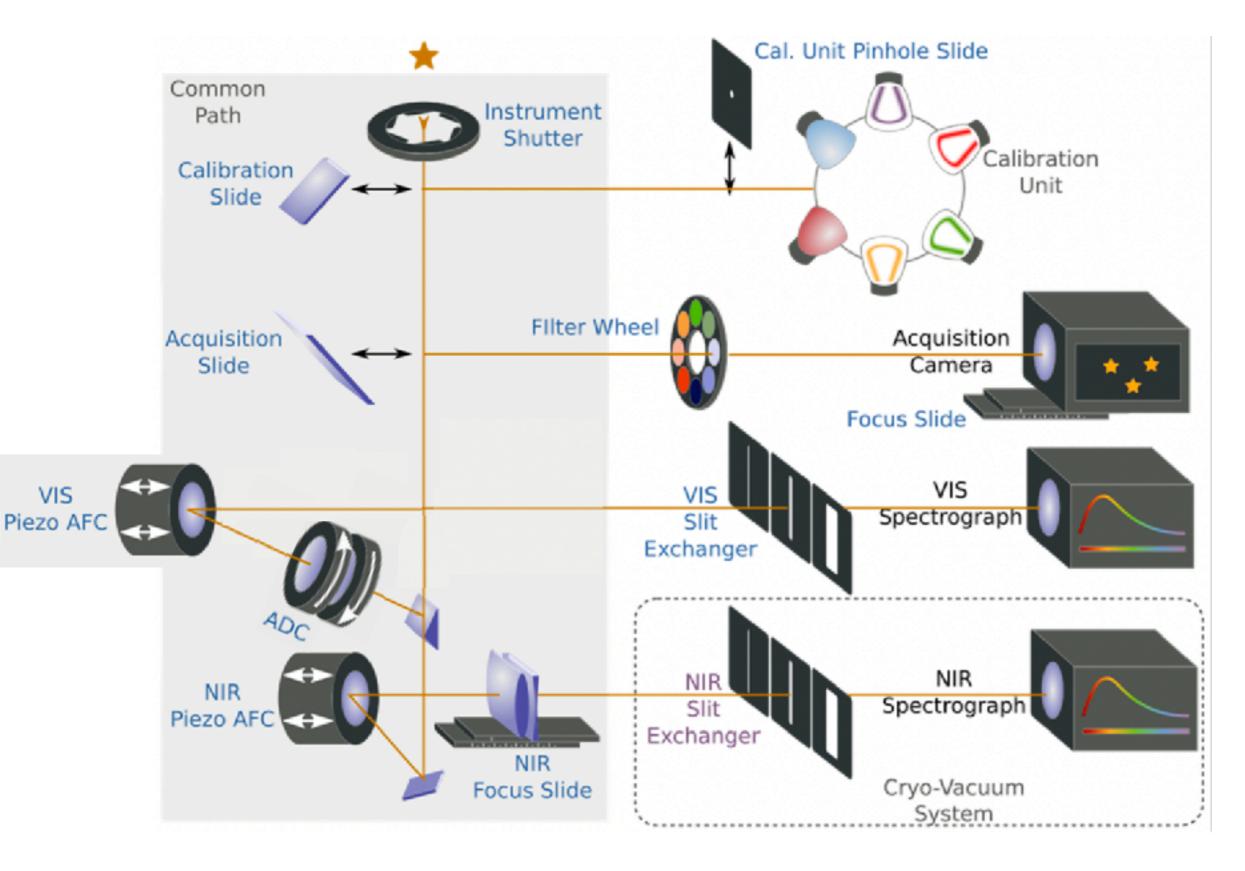
Proposal submission (02/2015)

SOXS selected by ESO (05/2015) out of 19

Signed MoU INAF-ESO Signed MoU INAF-Partners

Project Phase	Start	End	Duration
Preliminary Design	08/2016	07/2017	12 months
Final Design	08/2017	10/2018	14 months
MAIT	11/2018	11/2022	48 months+COVID
PAE	12/2022	02/2023	3 months
Commissioning & SV & PAC	03/2023	09/2023	6 months
Operations & GTO	2023	2028	

SOXS block diagram



4.1 Mounting to the NTT Derotator

The following table describes what all components go on the NTT derotator, their weights, and where they are attached.

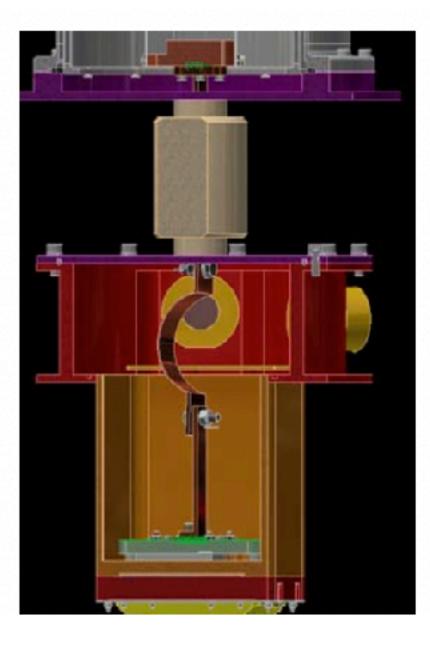
#	What	Weight measured at INAF-PD [kg]	Where is it attached
1	SOXS Flange + Support Structure	~ 355	NTT Derotator
2	Common Path	46	SOXS Flange
3	Calibration Unit	9	SOXS CP
4	Acquisition Camera	11	SOXS CP
5	UVVIS Spectrograph	140	SOXS Flange
6	NIR Spectrograph	~ 217	SOXS Flange
7	NGC - UVVIS + Support Structure	17	NTT Derotator
8	NGC - NIR + Support Structure	~ 17	NTT Derotator
9	Power Supply Support Structure	~ 6	NTT Derotator

 Table 2. The table shows the weight of the components that will be mounted to the flange (highlighted in green) and onto the NTT rotator adaptor (highlighted in blue).

UV-VIS Cryogenic System

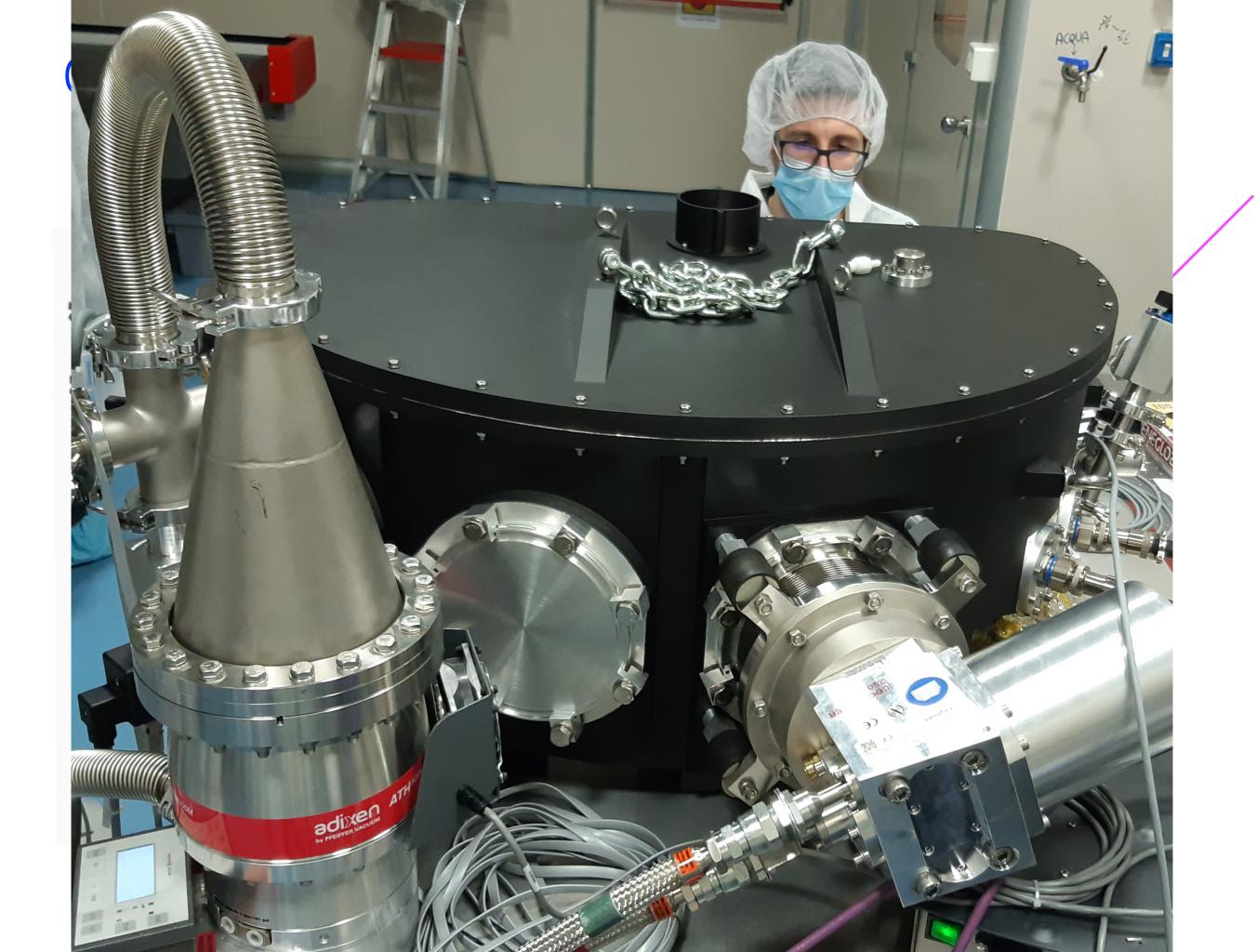




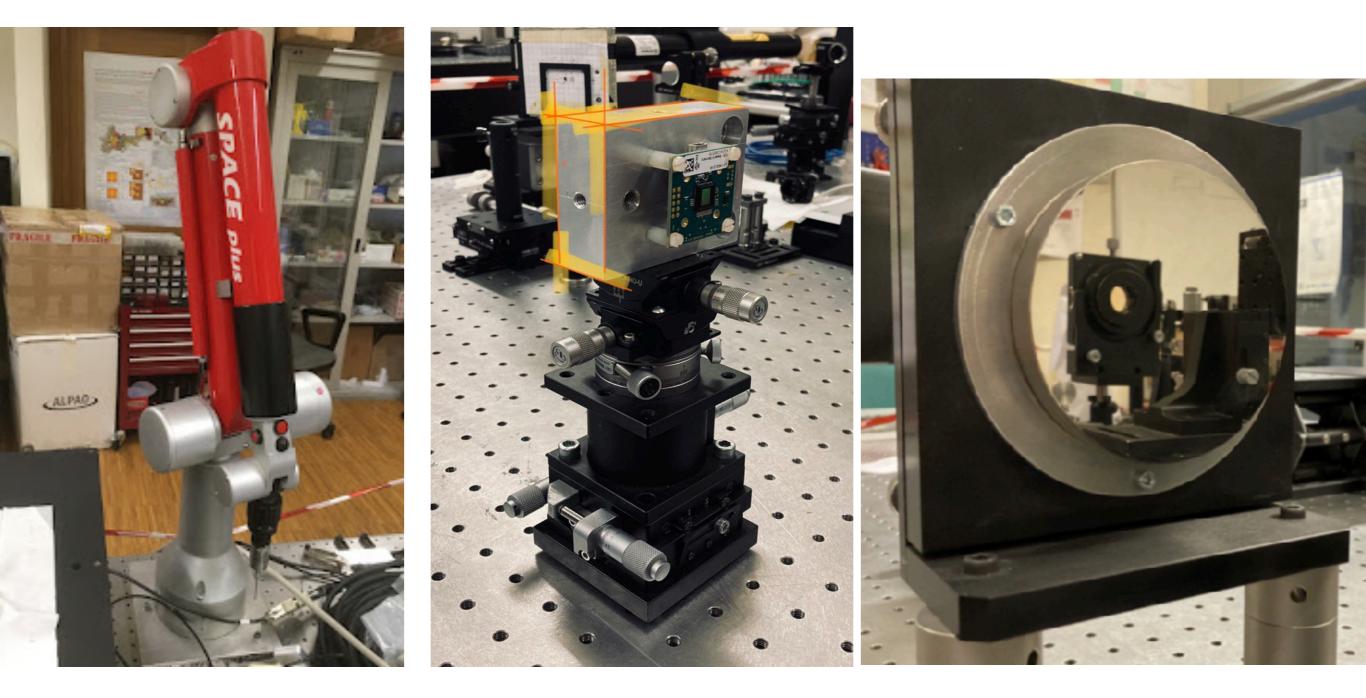


The cold finger of CFC.

The thermal connection from CFC to CCD. Section of the UV-VIS cryostat showing the thermal connection.



Common Path Alignment Tools



Portable Coordinate Measuring Machine

Characterized CMOS detector

10cm diameter (aluminum) mirror

What is SOXS?

- Single object spectrograph offering simultaneous spectral coverage in UV-VIS (350-850nm) and NIR (800-2000nm) with an average R ~ 4500 for an 1" slit.
- Can perform photometry in 360-970nm (ugVrizY) (3.5' x 3.5', 0.2"/pixel).
- Final destination: Nasmyth platform of the 3.58m ESO New Technology Telescope (NTT) at the La Silla Observatory in the Southern part of the Chilean Atacama Desert.
- Designed to observe all kinds of transients and variable sources.
- SOXS consists of 5 sub-systems CommonPath (CP), Calibration unit (CU), Acquisition Camera (AC), UV-VIS spectrograph, and the NIR spectrograph.

Science with SOXS

- Classification
- Super Novae
- Gravitational Waves EM counter parts
- Tidal Disruptive Events & Nuclear transients
- Gamma Ray Bursts and Fast Radio Bursts
- Blazars & AGN
- X-ray binaries & magnetars
- YSOs & Stars
- Novae, Cataclysmic Variables, & White Dwarfs

Astroids & Comets

Solar System

Galactic

• Unknown

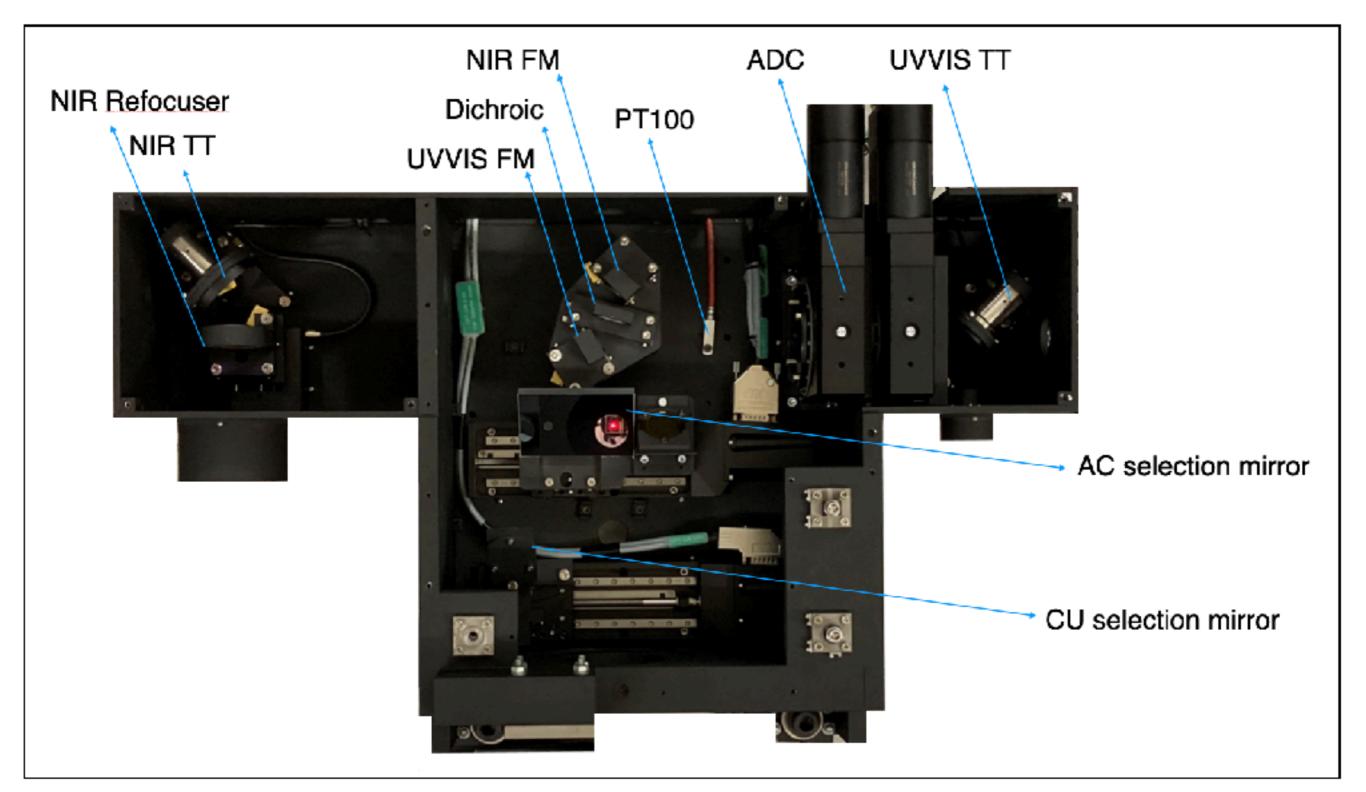
ExtraGalactic

SOXS Consortium

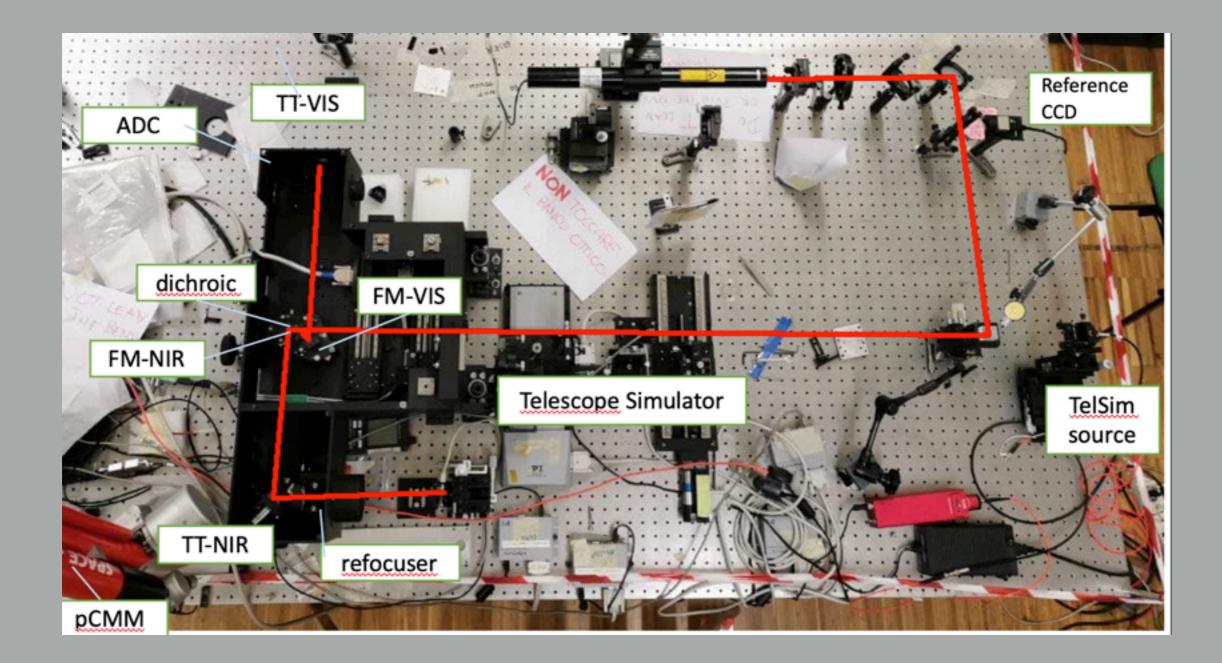
- Istituto Nationale di AstroFisica (INAF) - Italy
- Weizmann Institute of Science -Israel
- Universidad Andres Bello & Instituto Milenio de Astrofisica, Chile
- Turku University, Finland
- Queen's University, UK
- Tel Aviv University, Israel
- Niels Bohr & Aarhus University, Denmark



Common Path

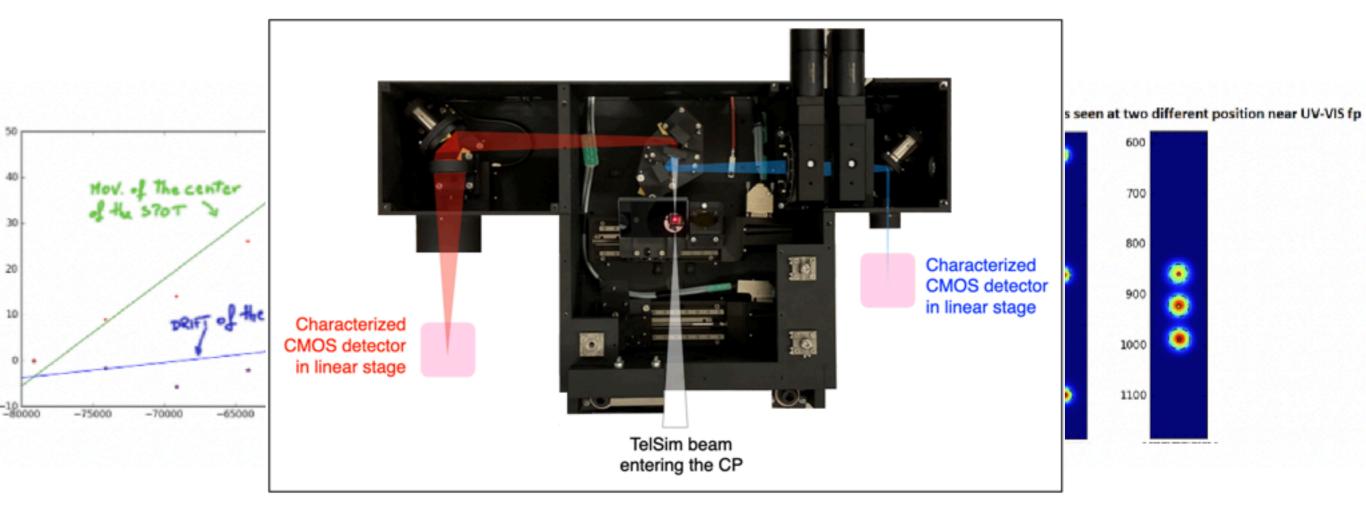


Receives the F/11 beam from the telescope and feeds the spectrographs with an F/6.5 beam



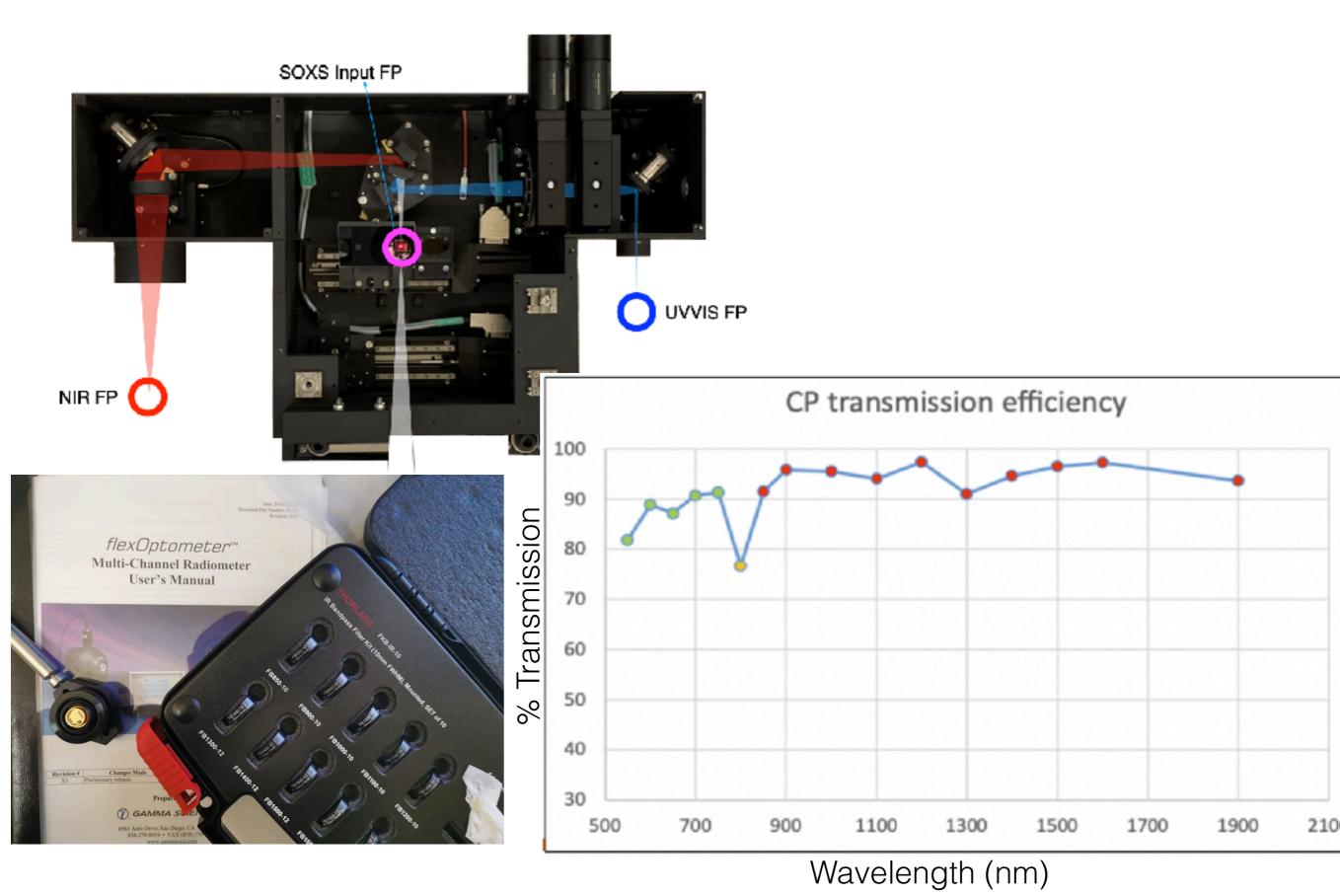
Common Path Alignment Verification

• CP exit beam position, tip-tilt, and focal number

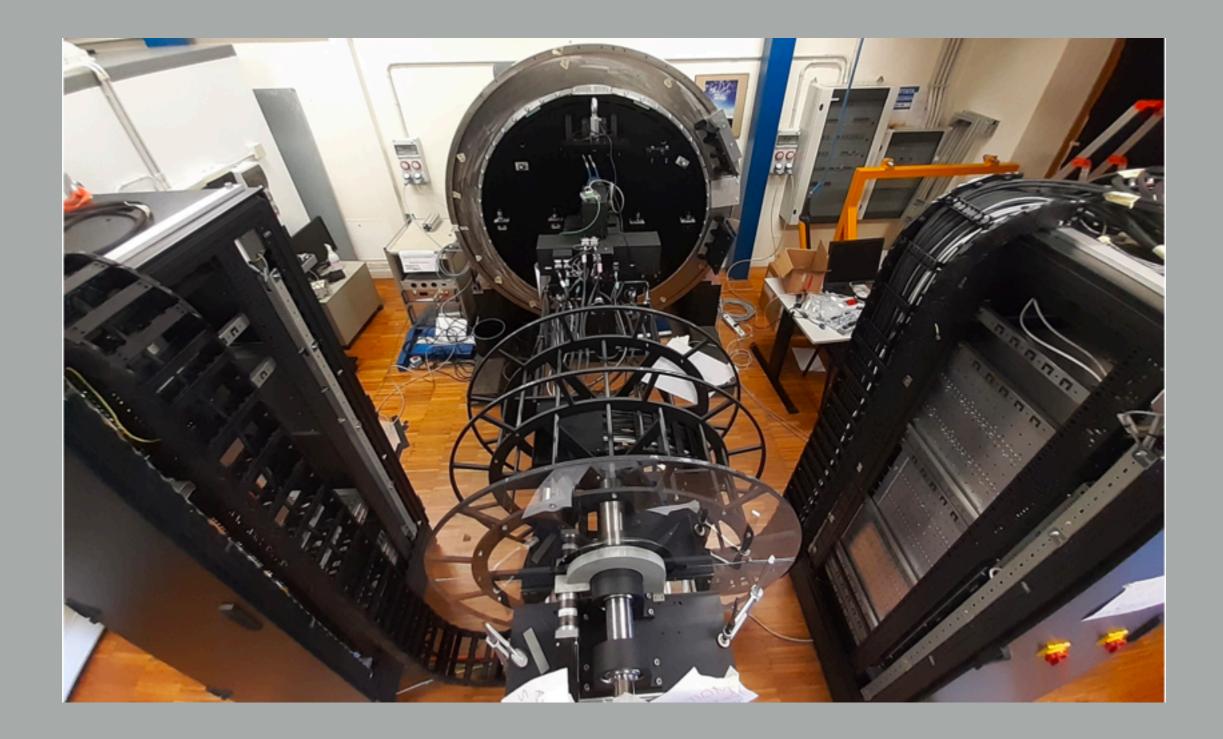


	Decenter–X (μm)	Decenter-Y (μm)	Decenter–Z (μm)	Tilt-X (")	Tilt-Y (")	F/#
UVVIS exit beam	230 ± 60	-145 ± 60	$<30\pm60$	-413 ± 100	37 ± 100	6.93
NIR exit beam	45 ± 60	3 ± 60	$<30\pm60$	-250 ± 32	-255 ± 32	6.94

Common Path Alignment Verification









Continuum source spectra

ThAr Arc lamp spectra