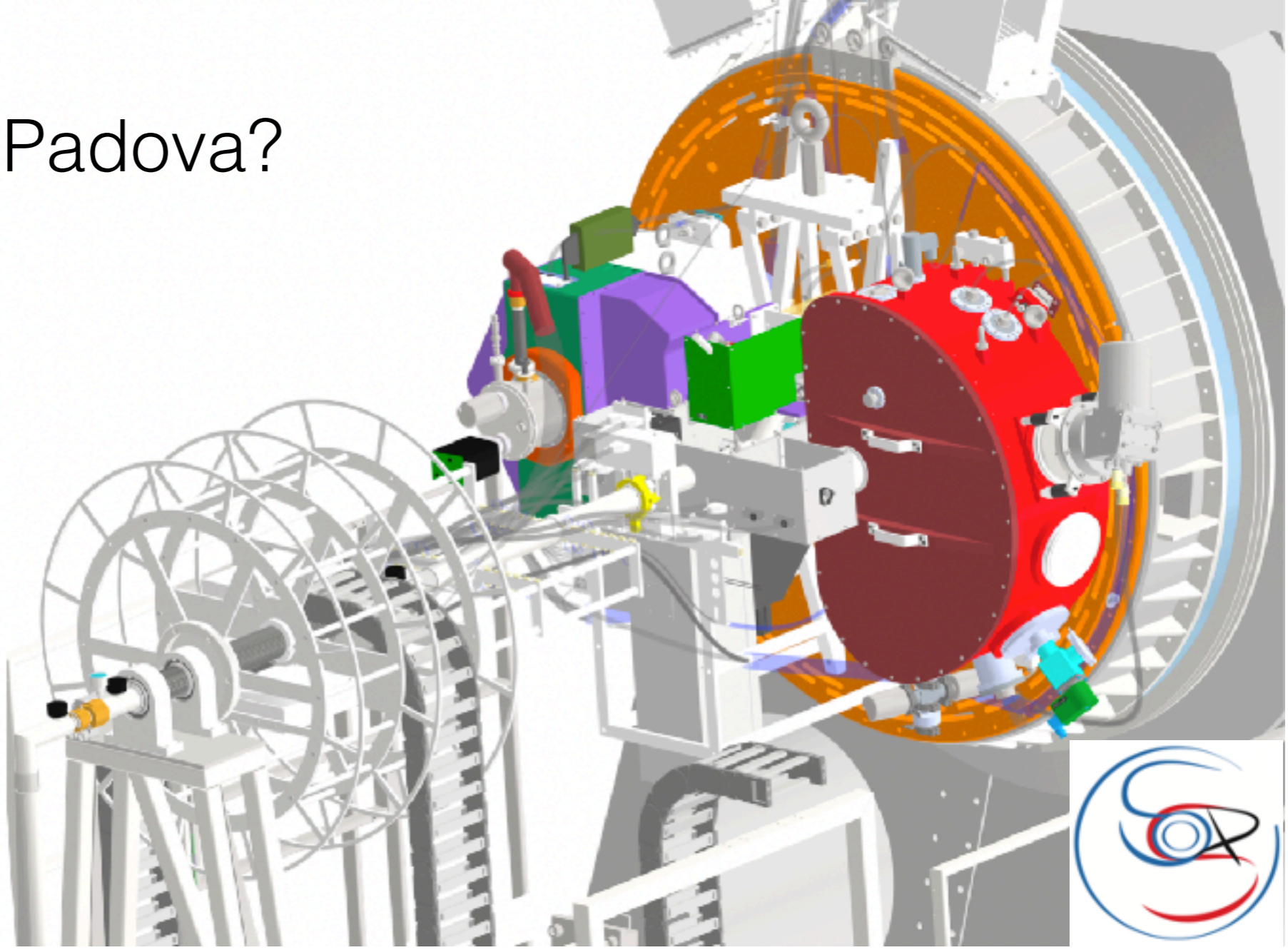


What is going on in Padova?



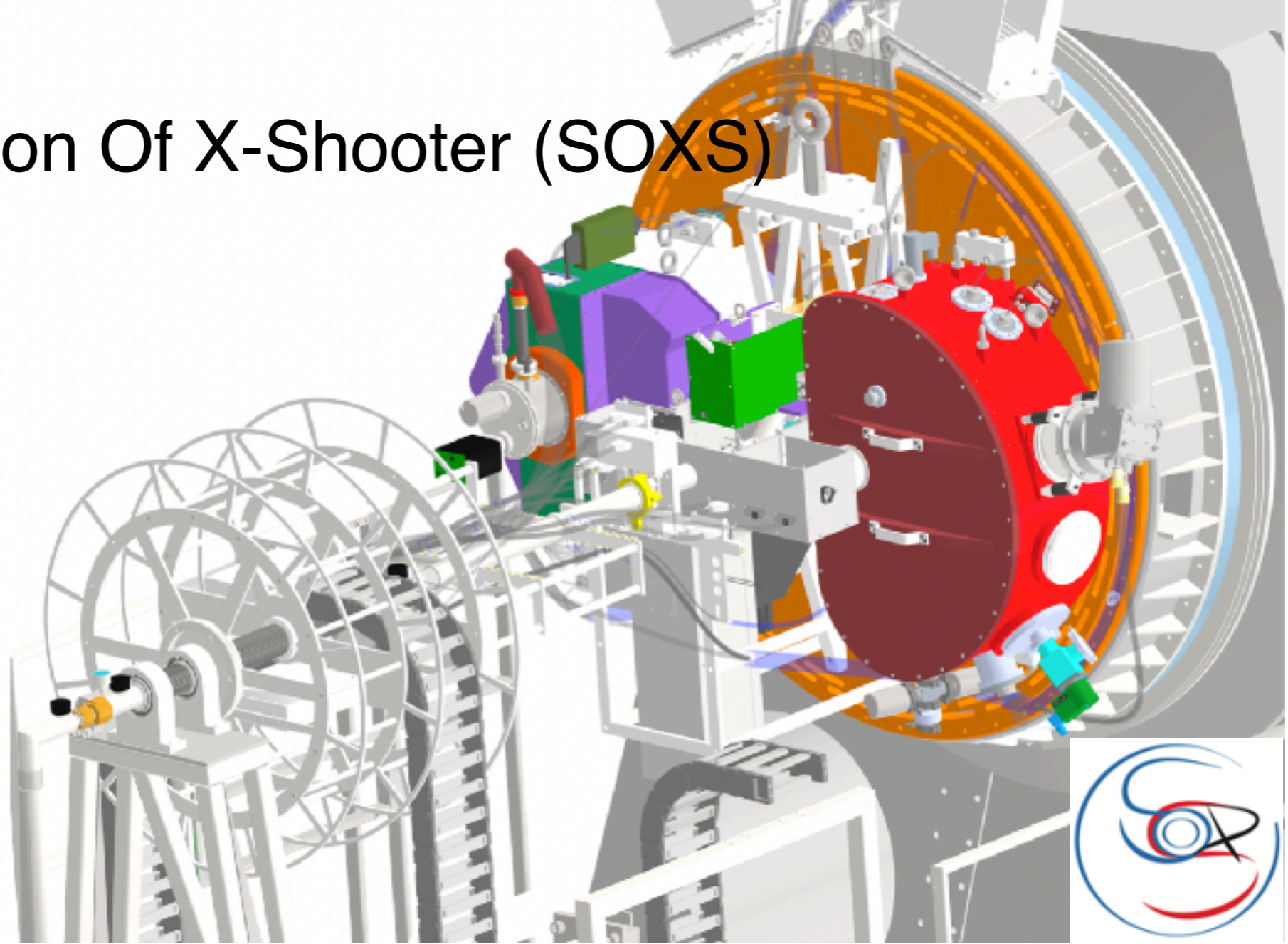
Kalyan Radhakrishnan & AIV Team

SOXS Instrument AIV Lead & Deputy System Engineer,
Astronomical Instrumentation and Adaptive Optics Group,
INAF-Osservatorio Astronomico di Padova (INAF-OAPD),
Padova, Italy



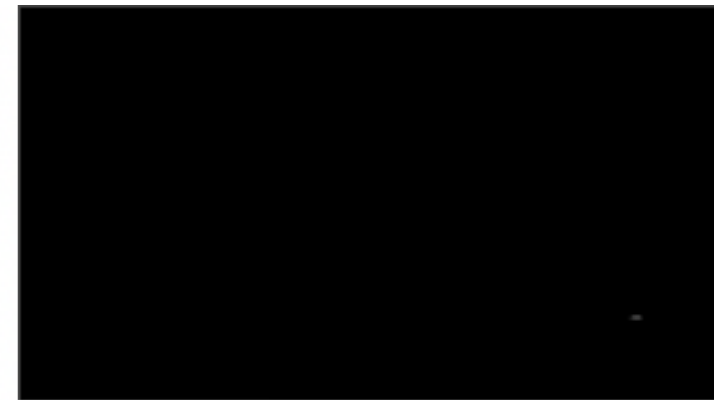
“SOXS Science Meeting”
INAF-Osservatorio Astronomico di Capodimonte
25 June 2024

The Making of the Son Of X-Shooter (SOXS)



Kalyan Radhakrishnan & AIV Team

SOXS Instrument AIV Lead & Deputy System Engineer,
Astronomical Instrumentation and Adaptive Optics Group,
INAF-Osservatorio Astronomico di Padova (INAF-OAPD),
Padova, Italy

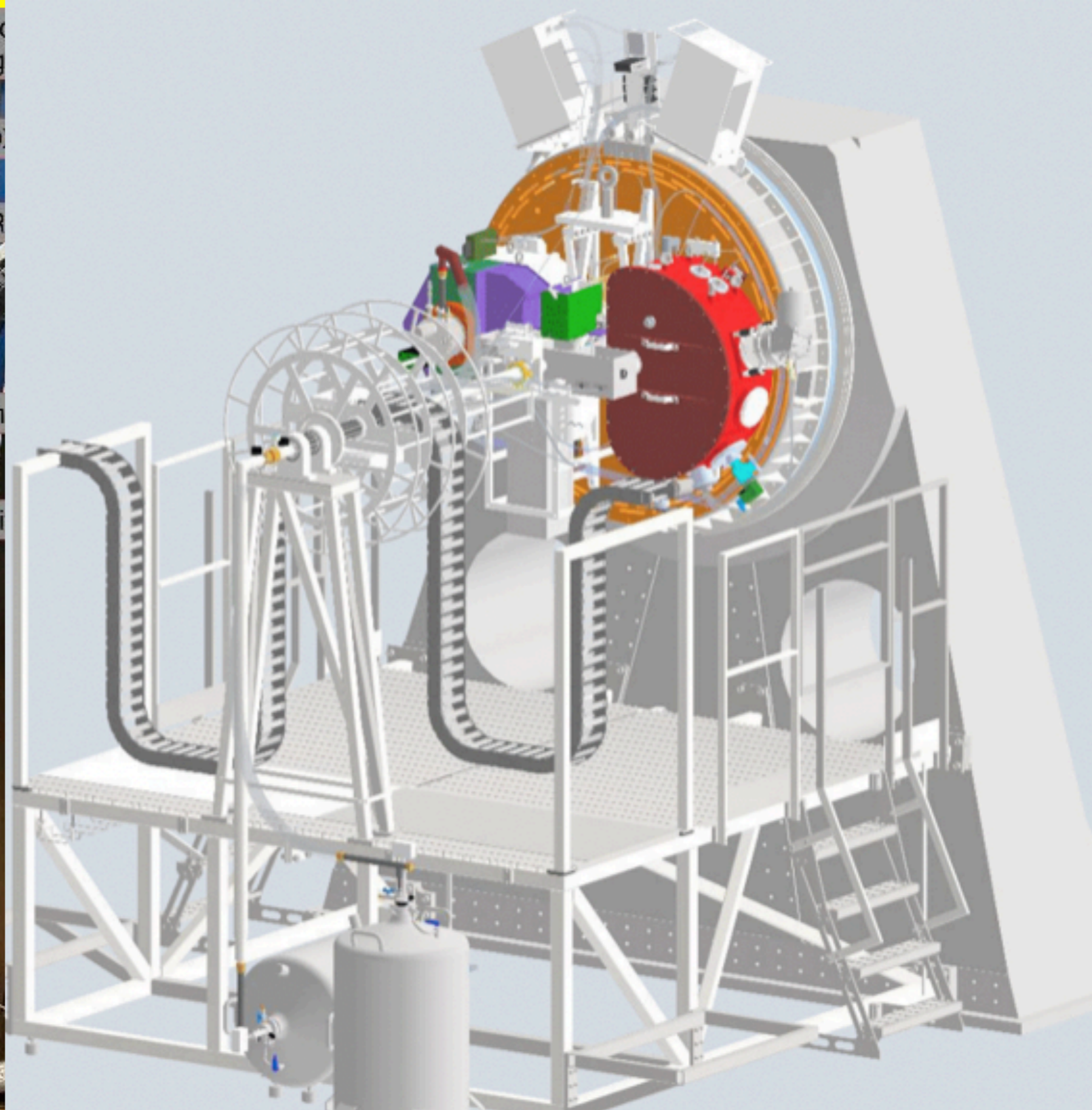
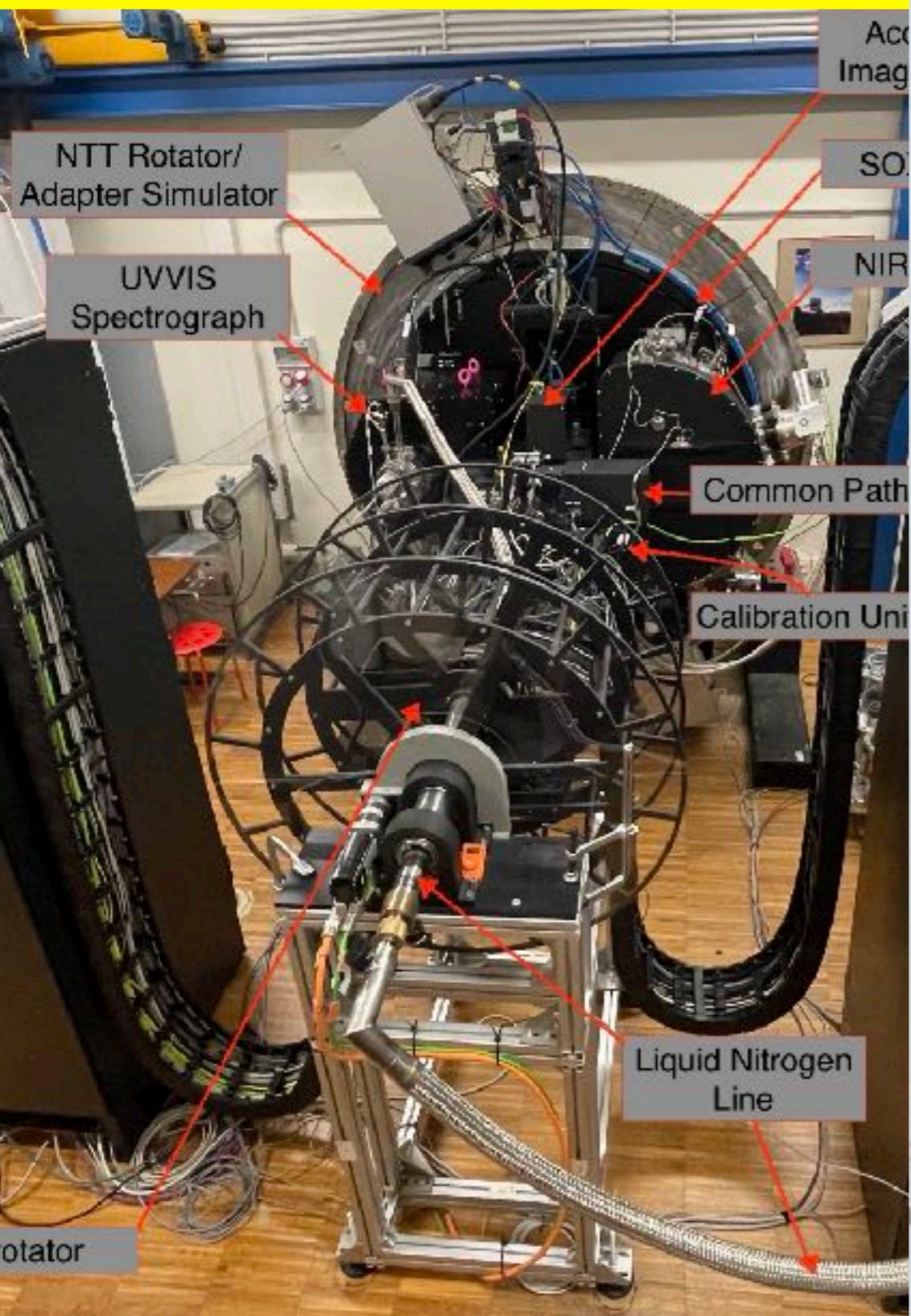


“SOXS Science Meeting”
INAF-Osservatorio Astronomico di Capodimonte
25 June 2024

Outline



La Silla vs Padova



SOXS in numbers

	UV-VIS	NIR
Spectral Range	350 - 850 nm	800 - 2000 nm
Resolution (1" slit)	>3600 (~4500 avg)	5000
Slit Widths	0.5", 1", 1.5", 5"	0.5", 1", 1.5", 5"
Slit height	12"	12"
Detector	e2V CCD44-82 2k x4k	Teledyne H2RG 2k x 2k
Pixel Scale	15um	18 um
Detector Scale	0.28"/pixel	0.25"/pixel

	Acquisition Camera
Spectral Range	360 - 970 nm
Filters	SDSS u, g, r, i, z, LSST y, and VIMOS V
FoV	3.5" x 3.5"
Detector	Andor iKon M-934 1k x 1k
Pixel Scale	13um
Detector Scale	0.205"/pixel

SOXS subsystems

INAF

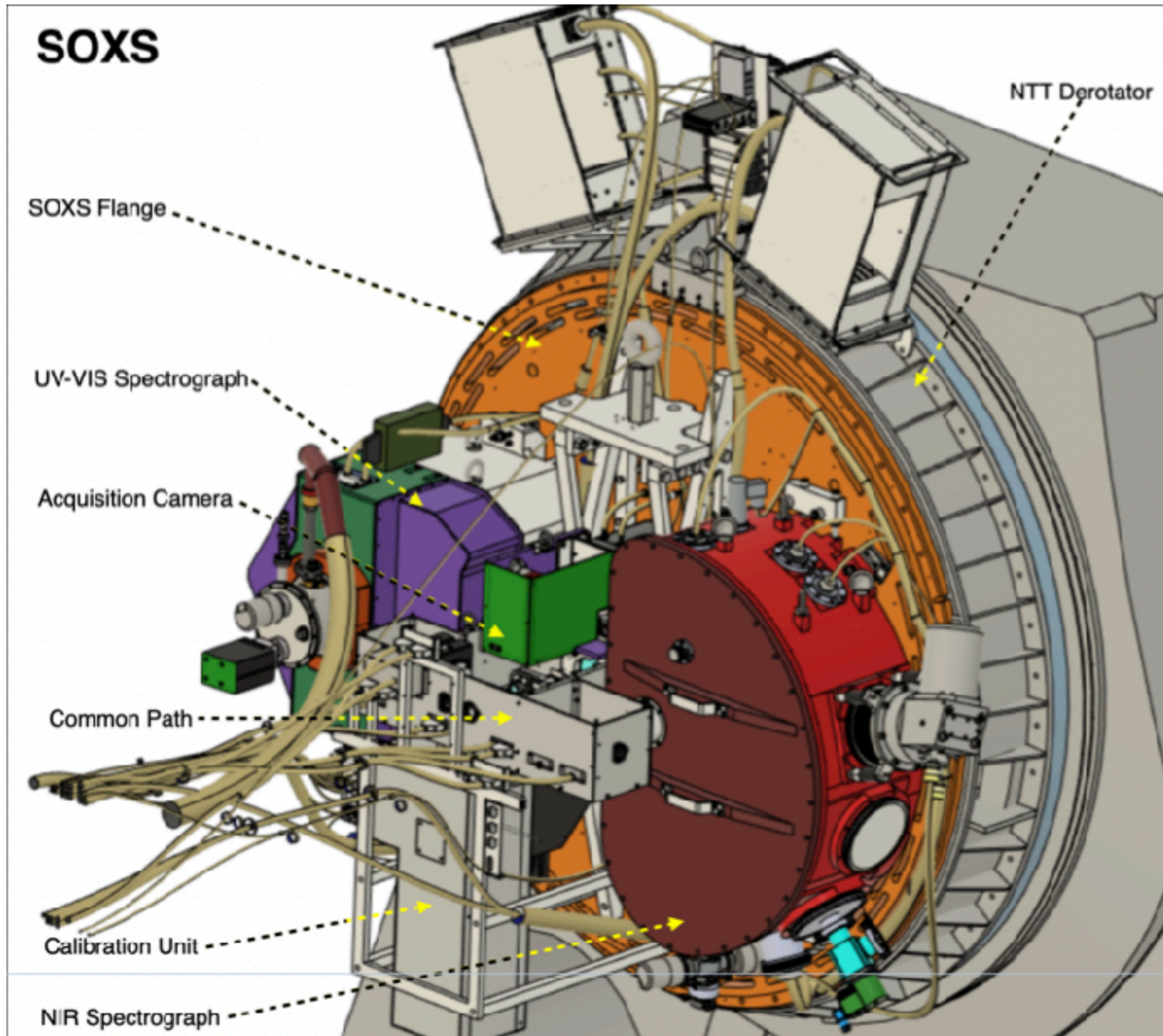
Israel

Chile

INAF

Finland

INAF



SOXS subsystems



SOXS

SOXS Flange

UV-VIS Spectrograph

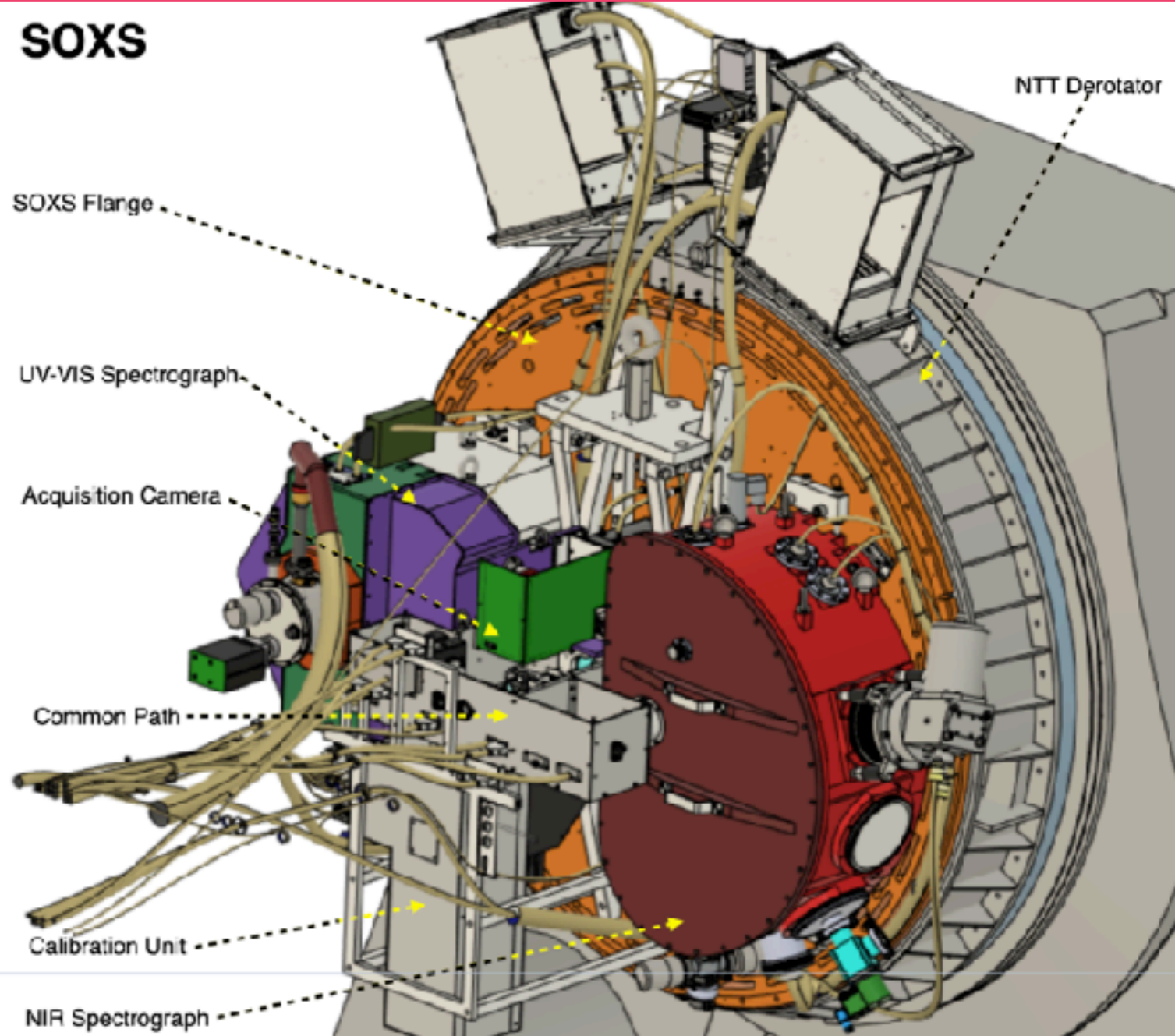
Acquisition Camera

Common Path

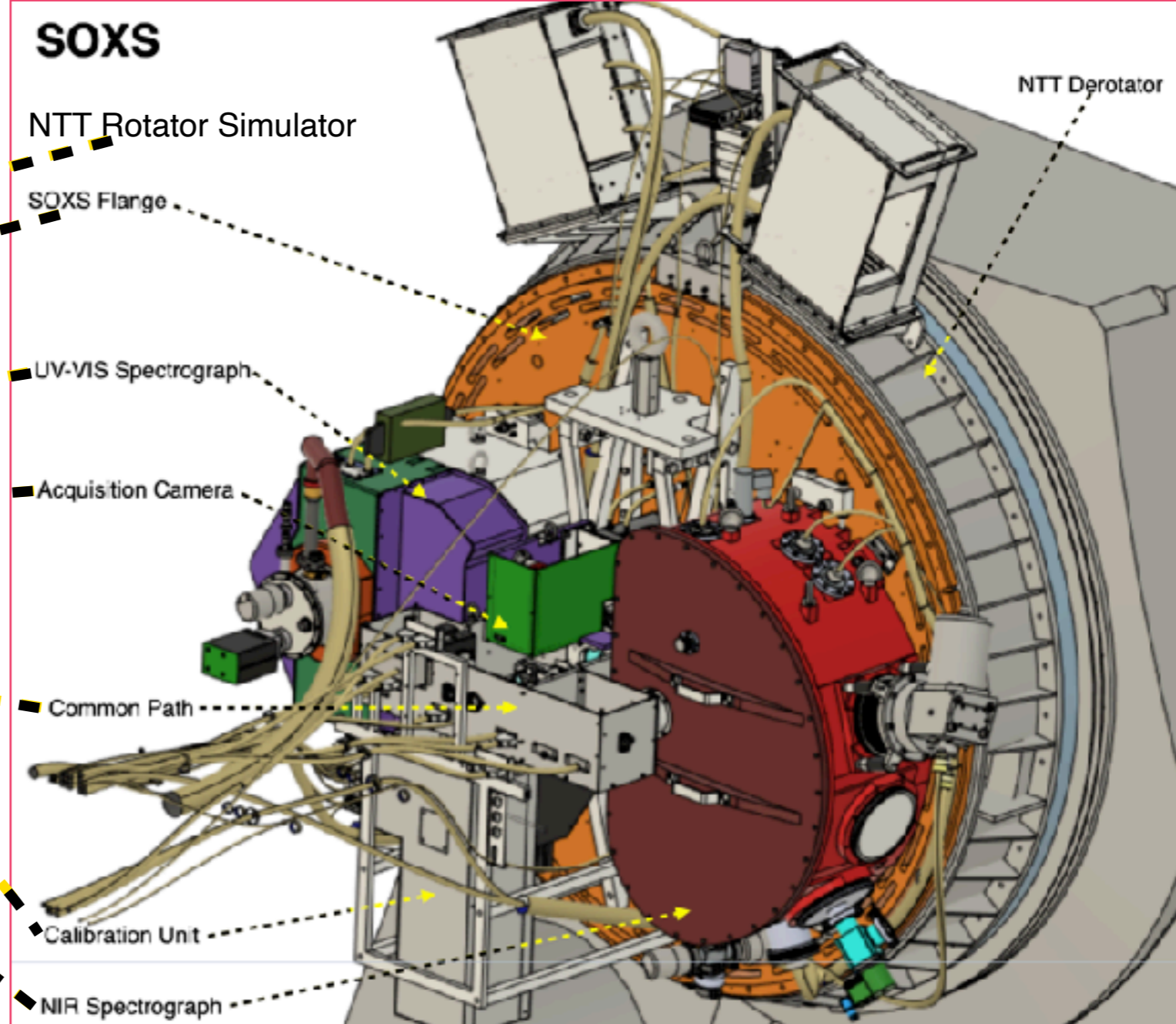
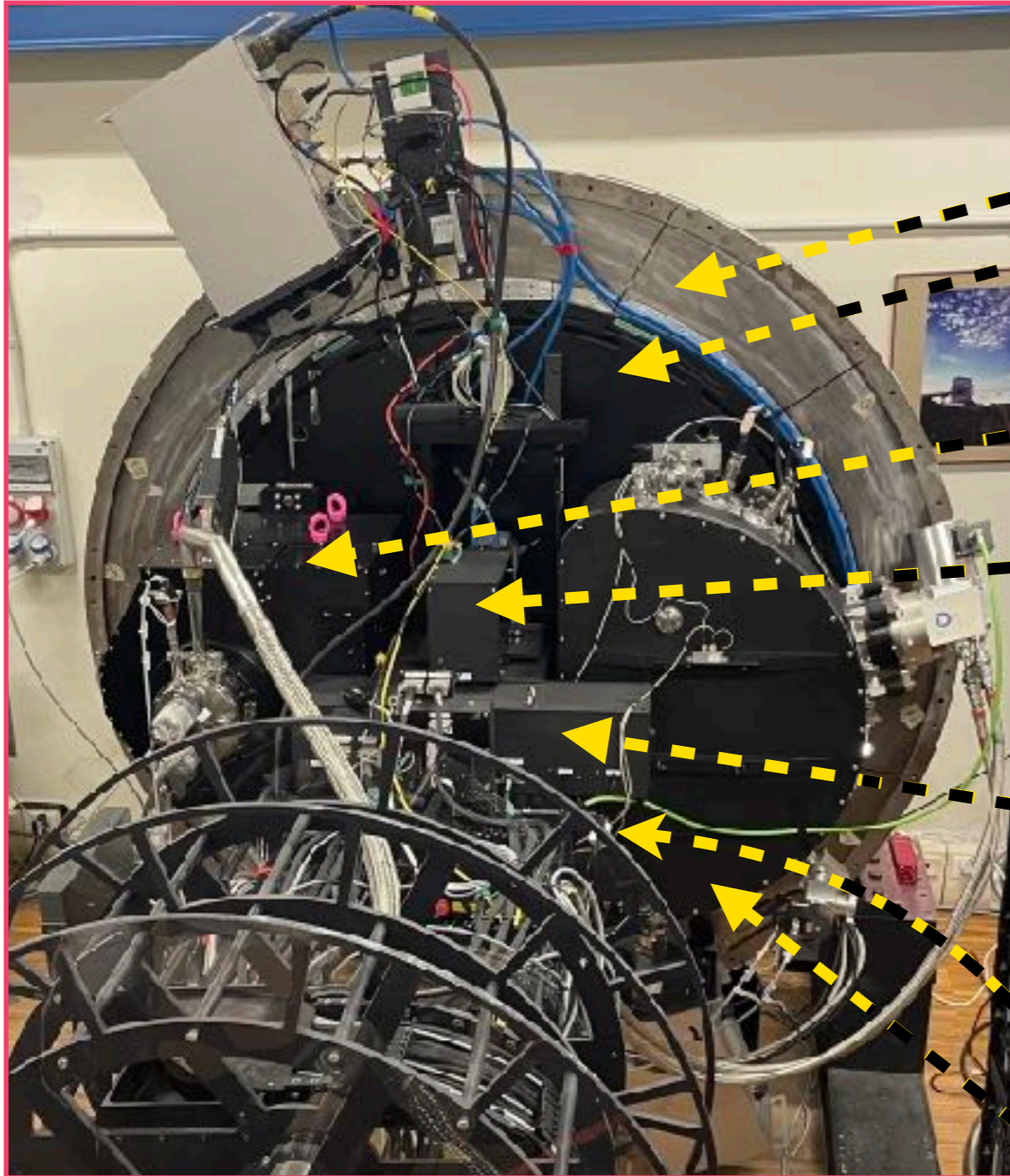
Calibration Unit

NIR Spectrograph

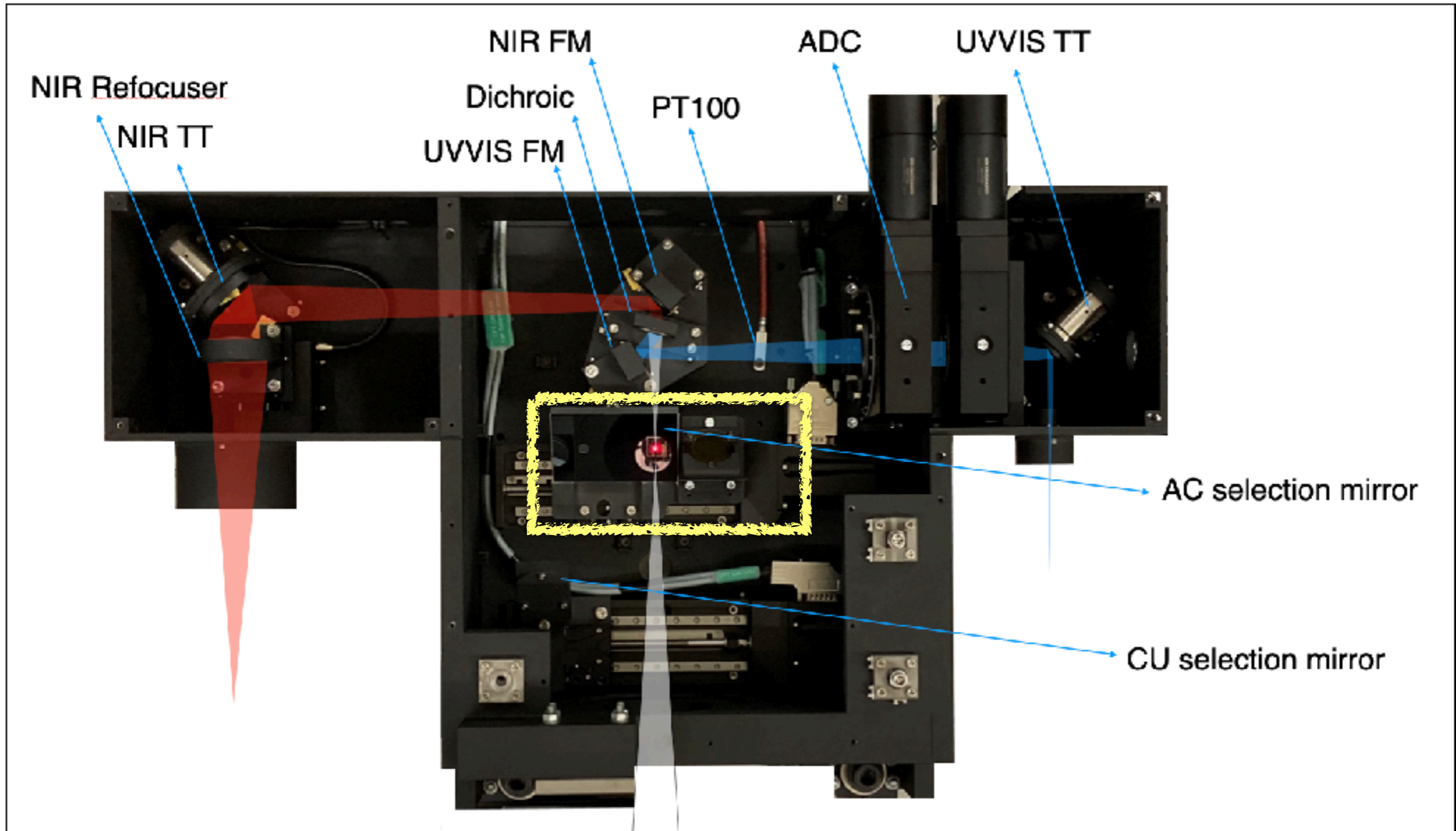
NTT Derotator



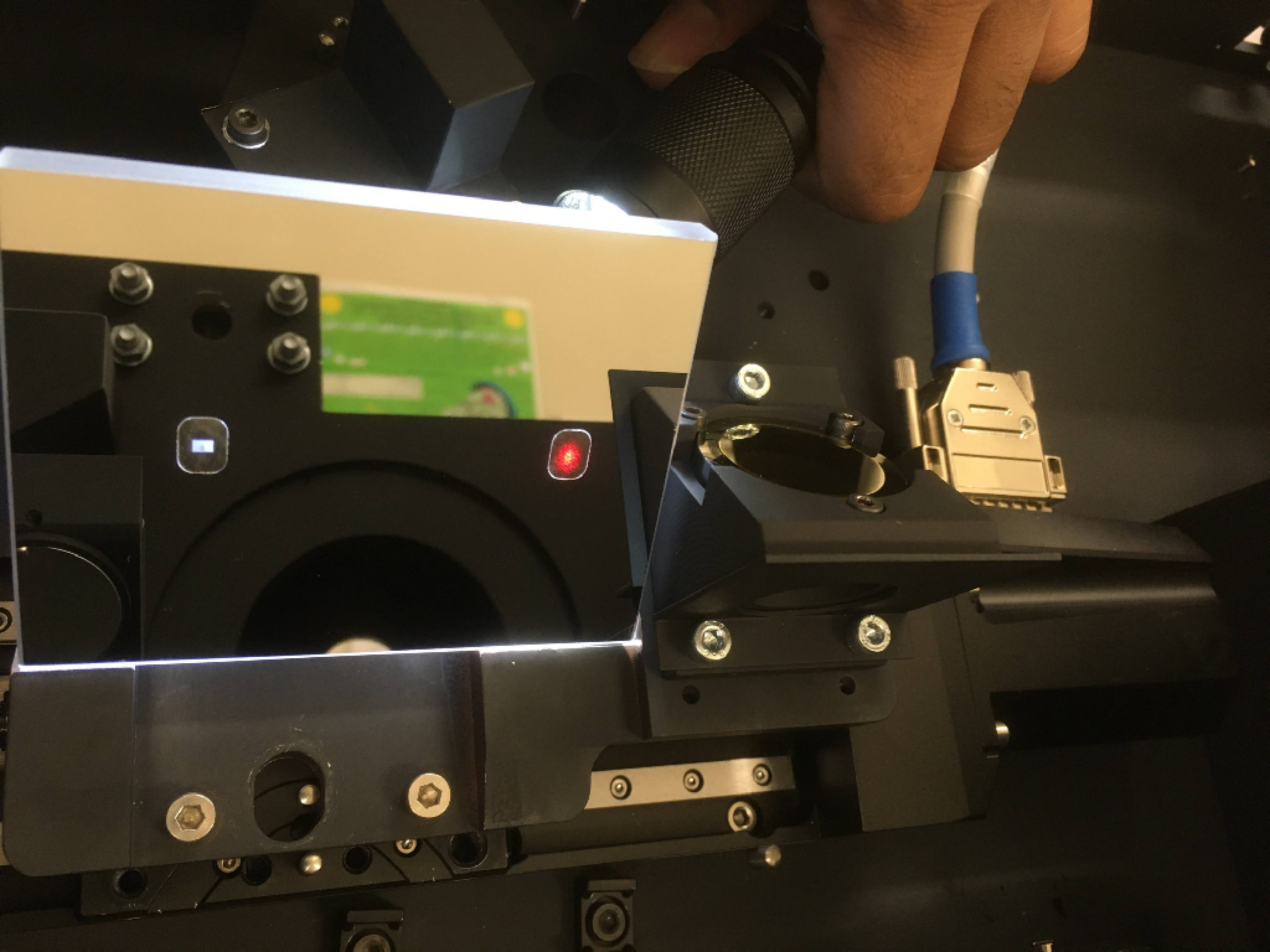
SOXS subsystems



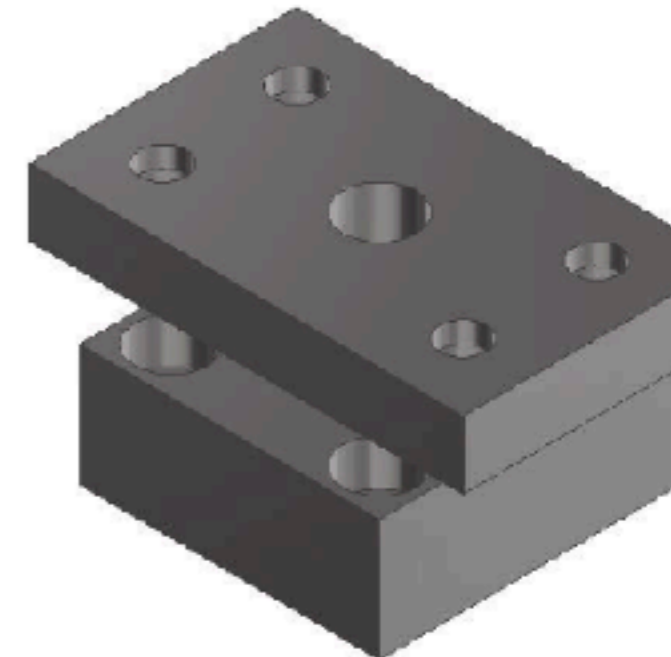
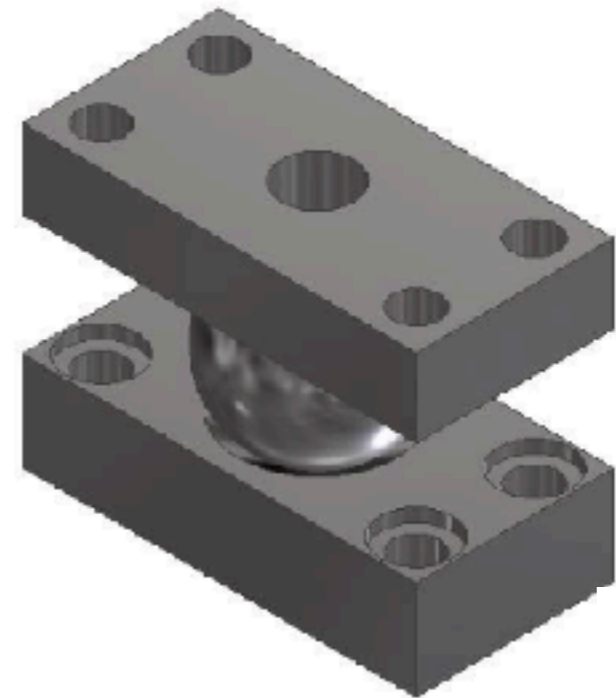
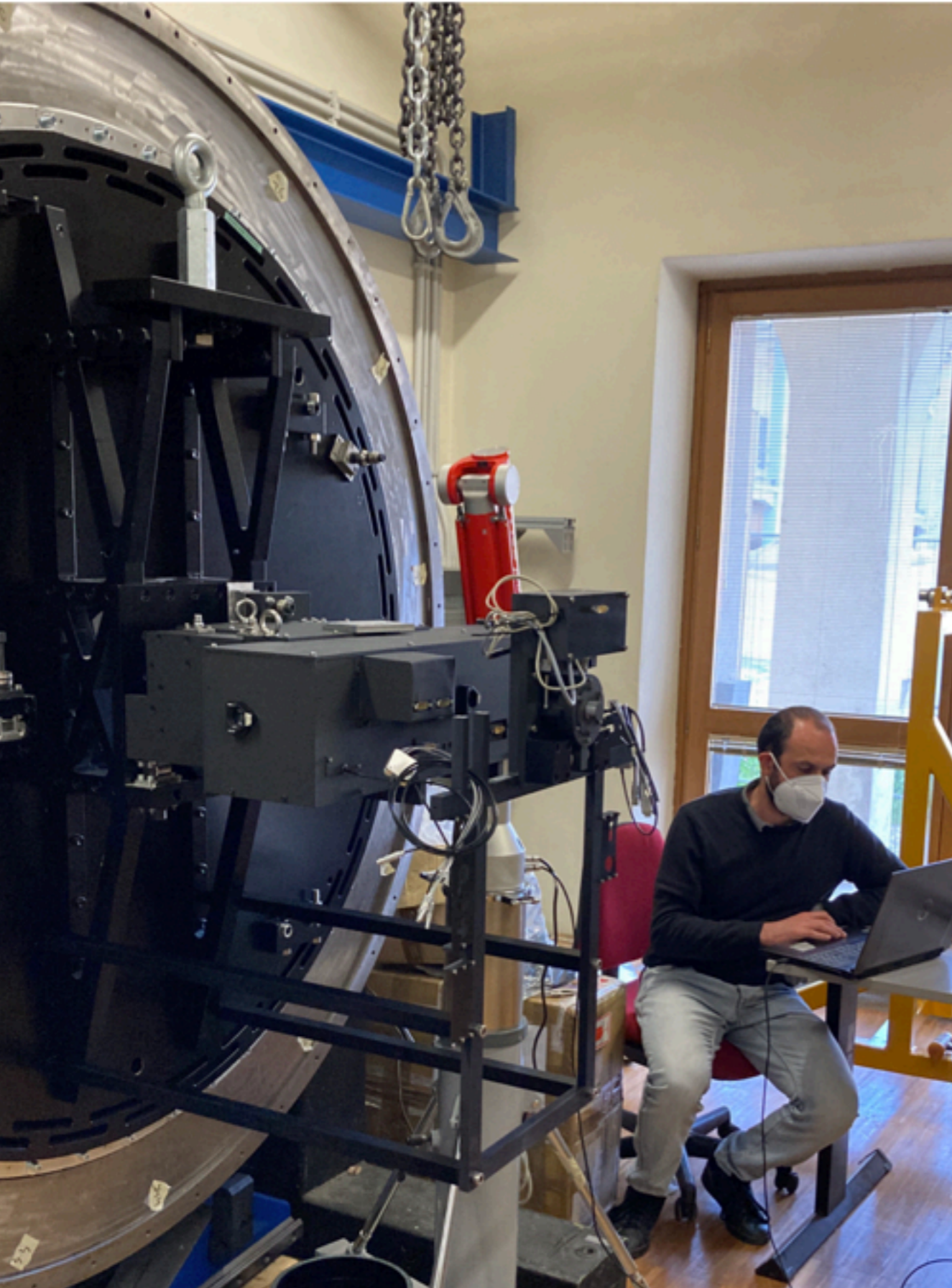
Common Path



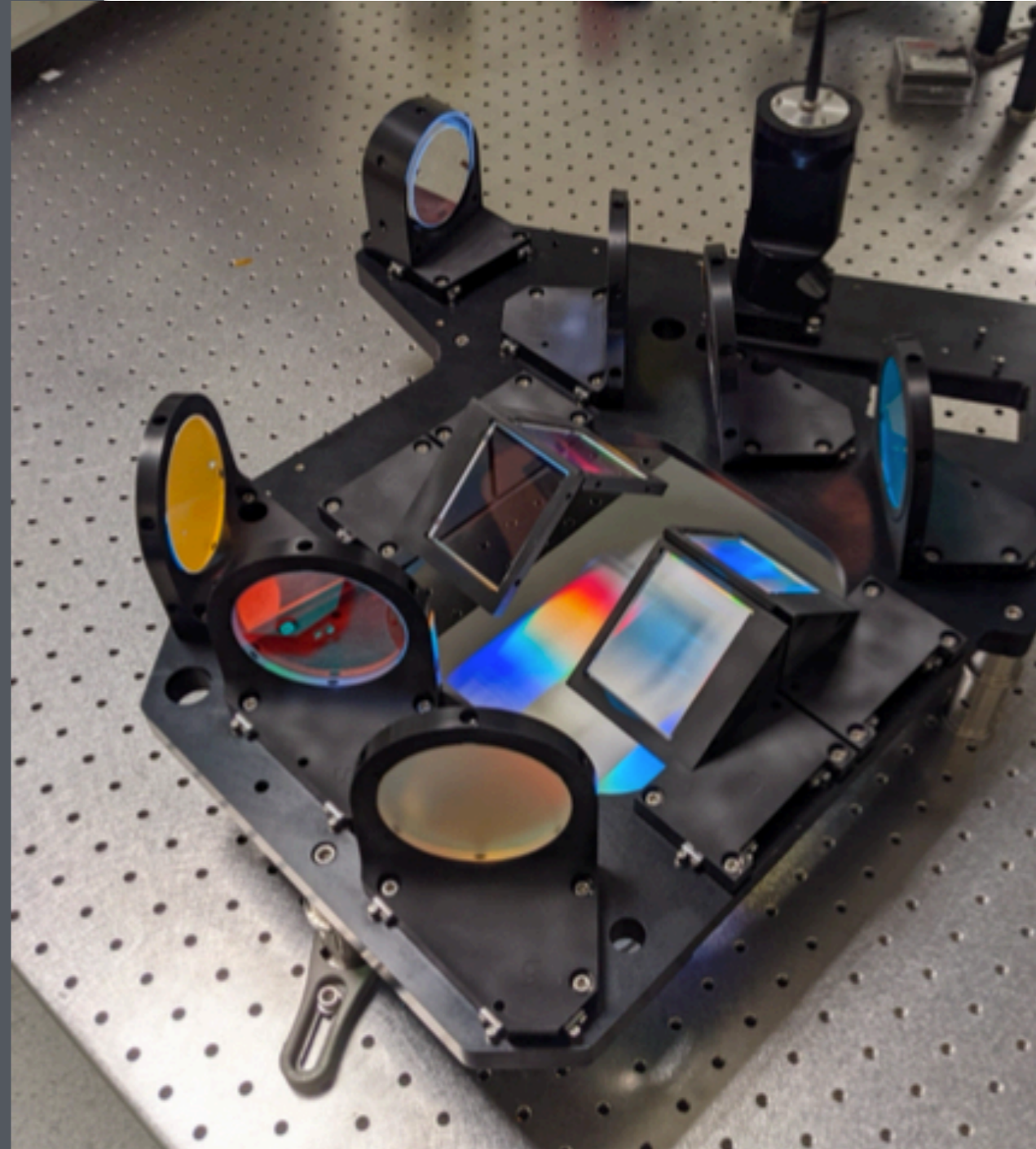
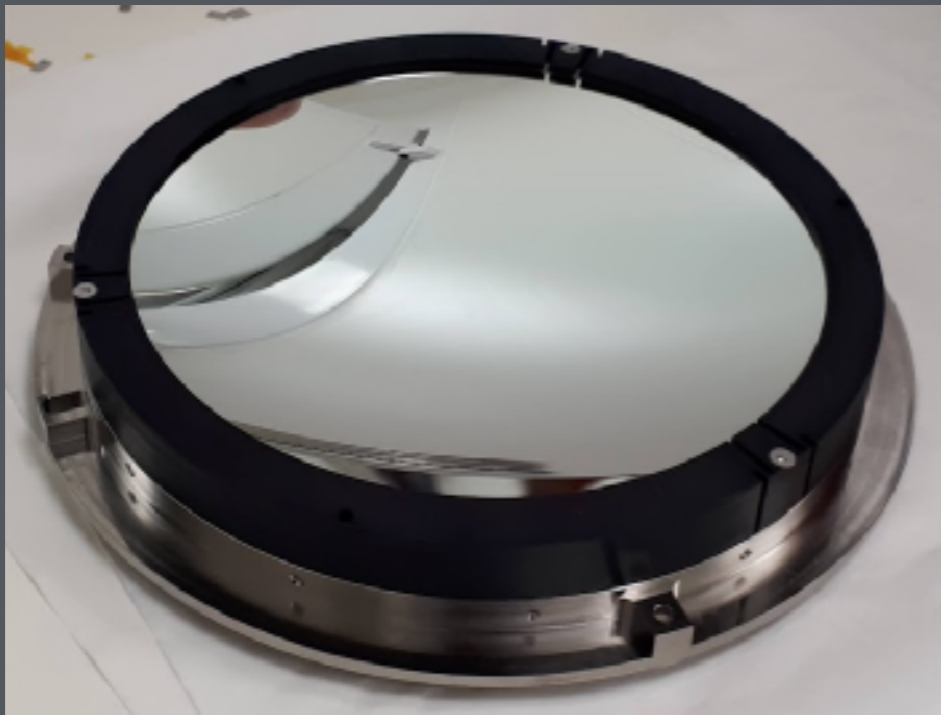
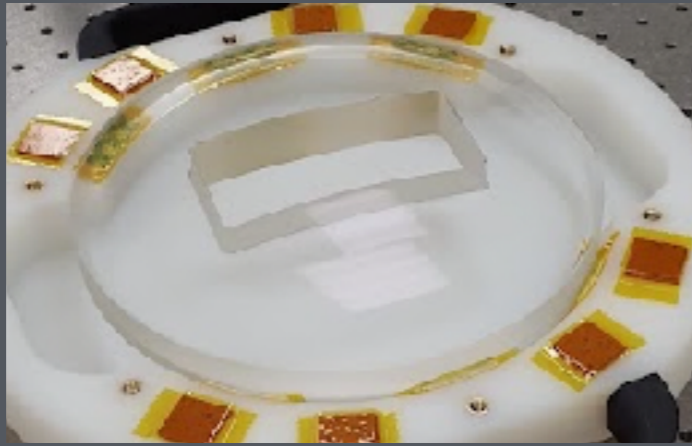
- Receives the F/11 beam from the telescope and feeds the spectrographs with an F/6.5 beam



Mounting to the Flange using Kinematic Mounts

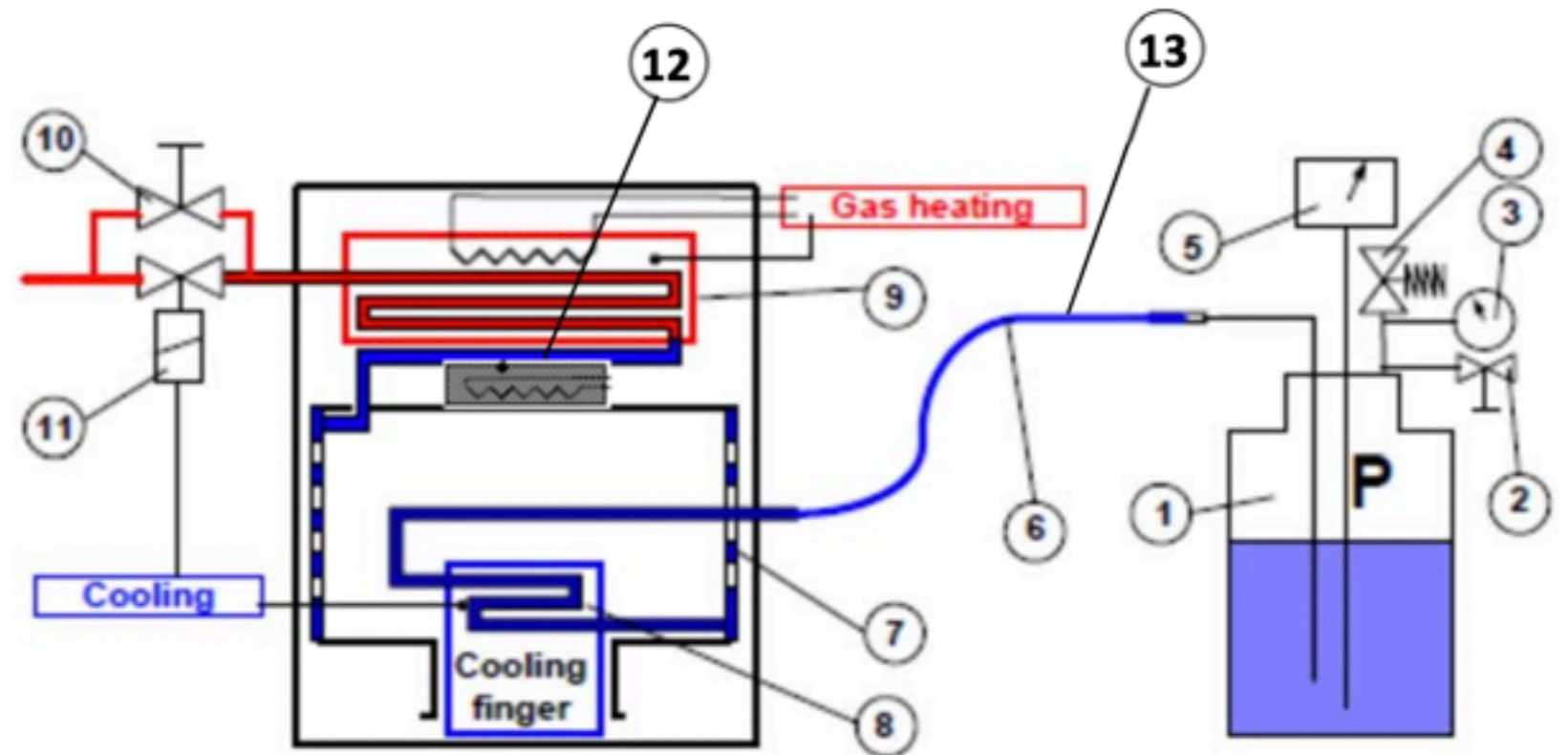


Kinematic Mounts

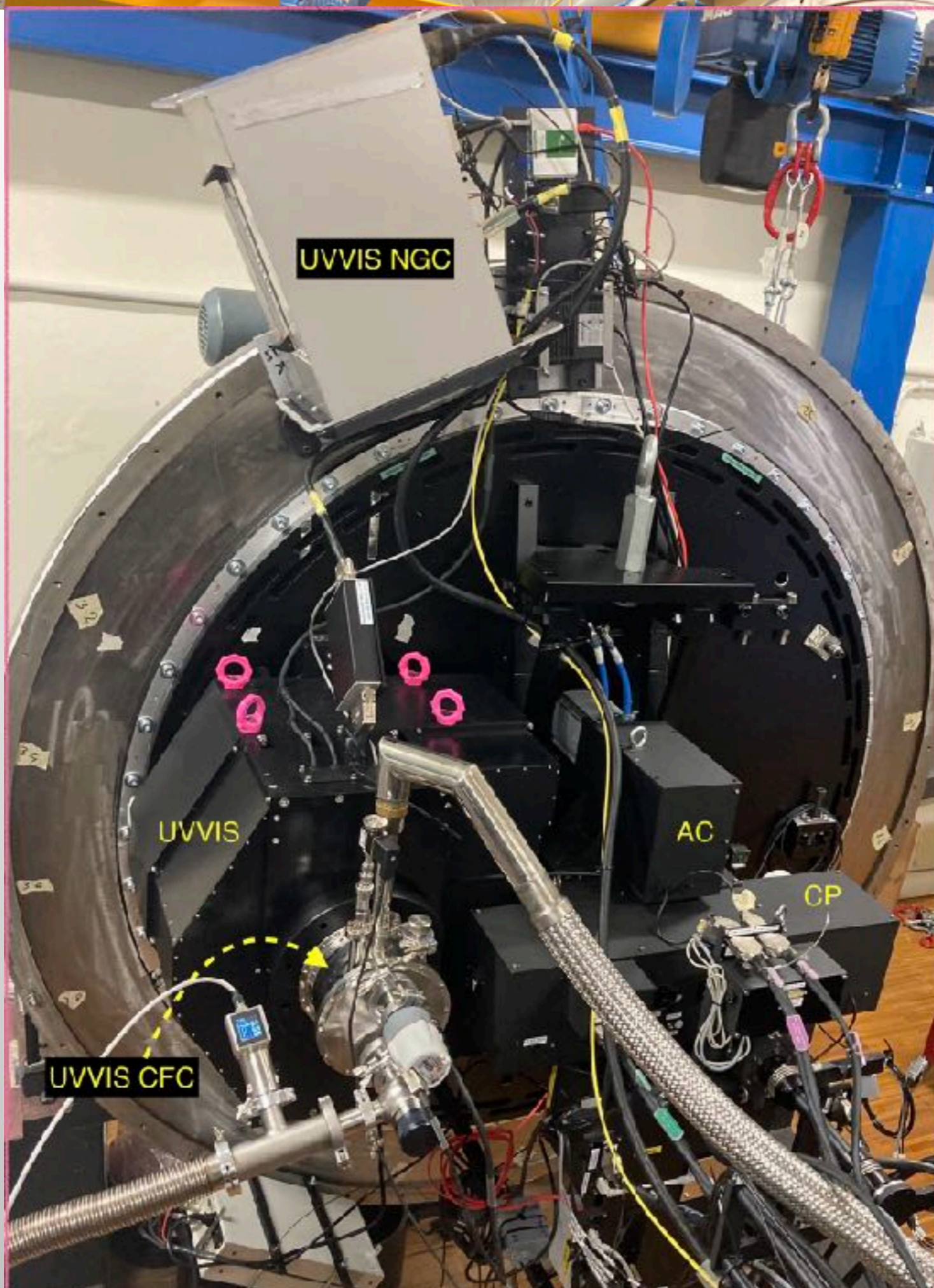


UV-VIS Cryogenic System

- A Continuous Flow Cryostat (CFC) is used to cool down the CCD.
- The CFC together with the detector head makes the UV-VIS cryostat.



CFC principle: 1. LN2 storage tank, 2. De-pressurization valve, 3. Manometer, 4. Over-pressure valve, 5. Liquid nitrogen level gauge, 6. Vacuum insulated transfer line, 7. Radiation shield heat exchanger, 8. Cooling heat exchanger, 9. Gas heater, 10. Bypass valve, 11. Regulation valve.



UVVIS NGC

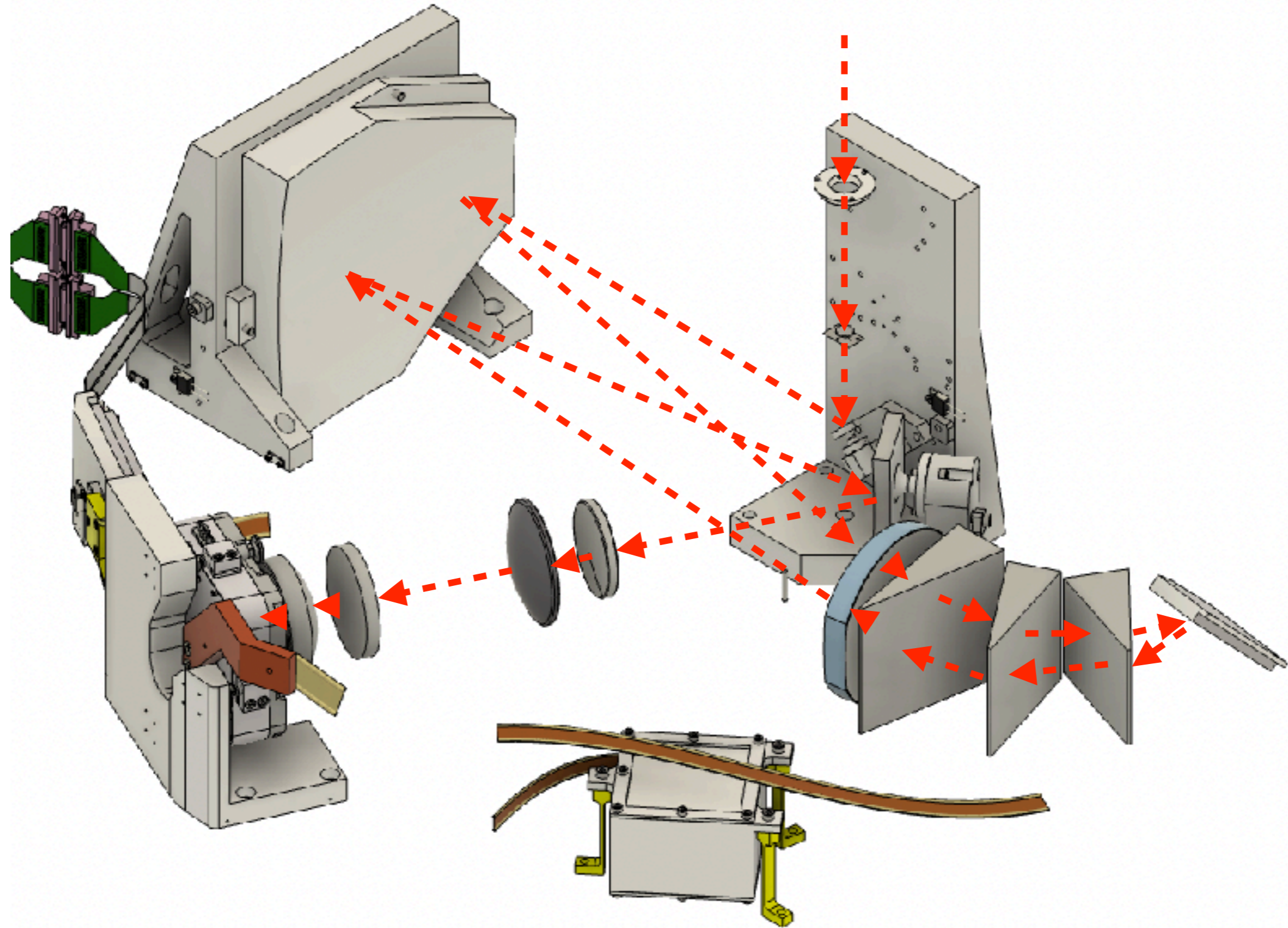
UVVIS

UVVIS CFC

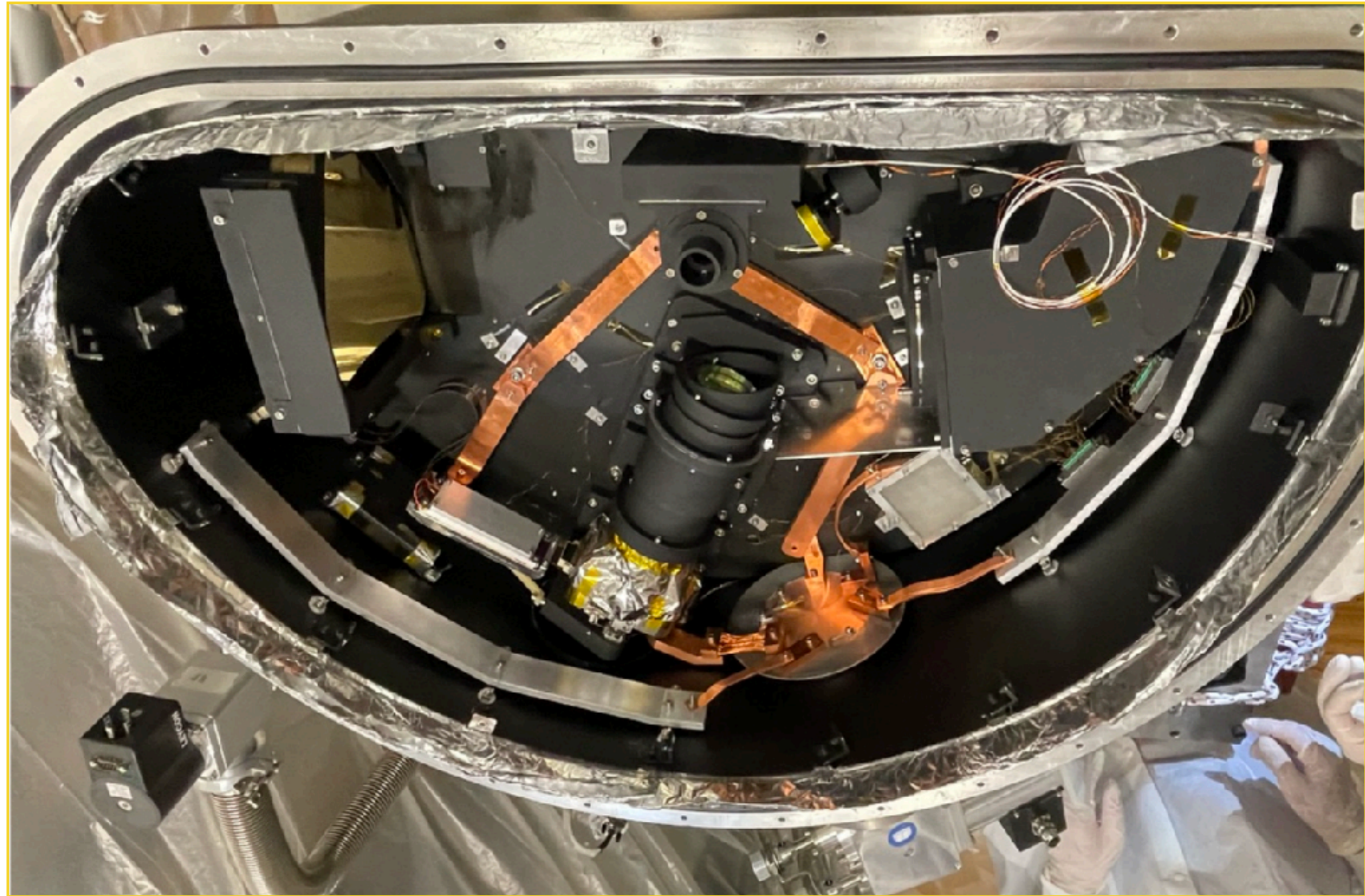
AC

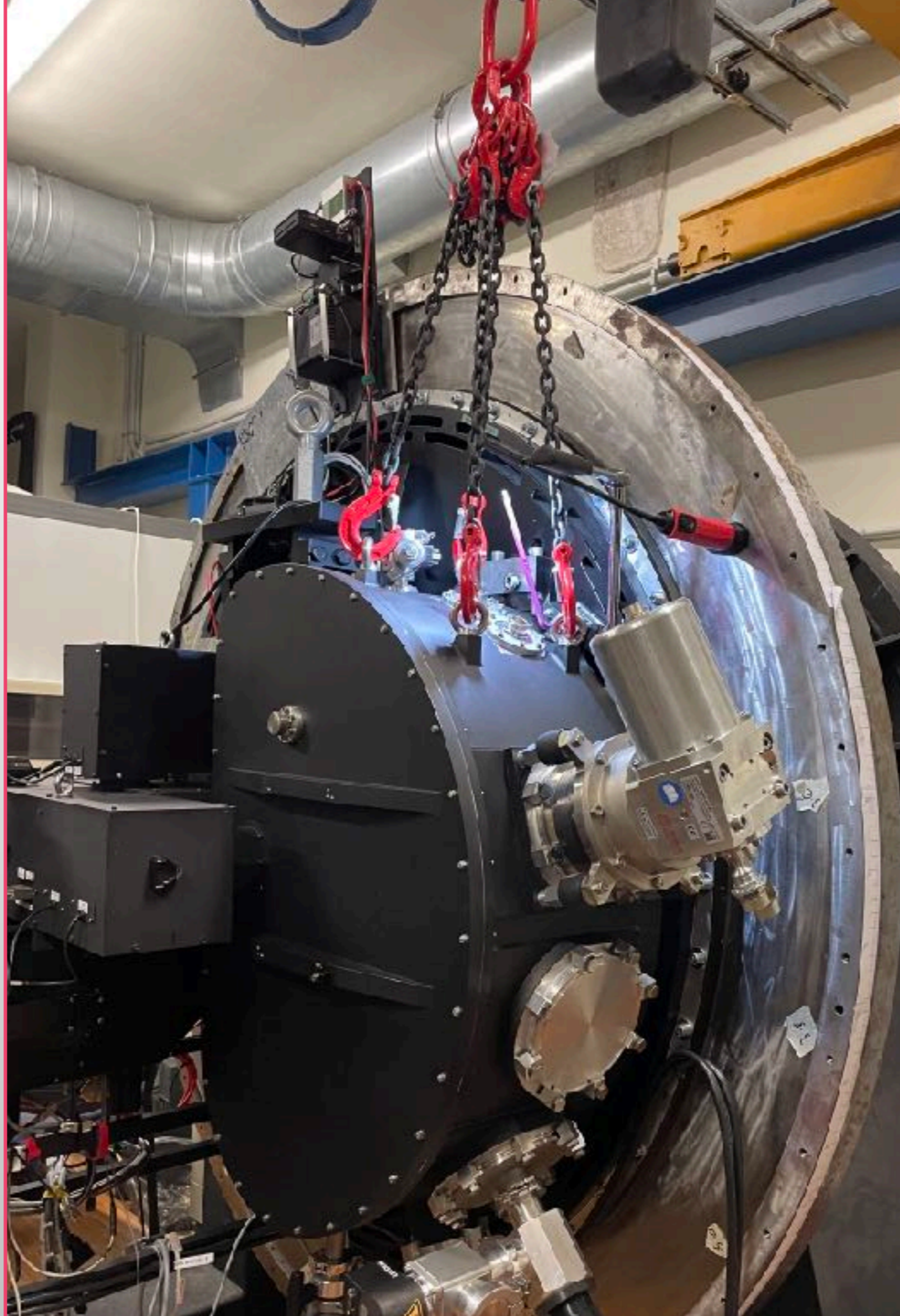
CP

NIR Spectrograph

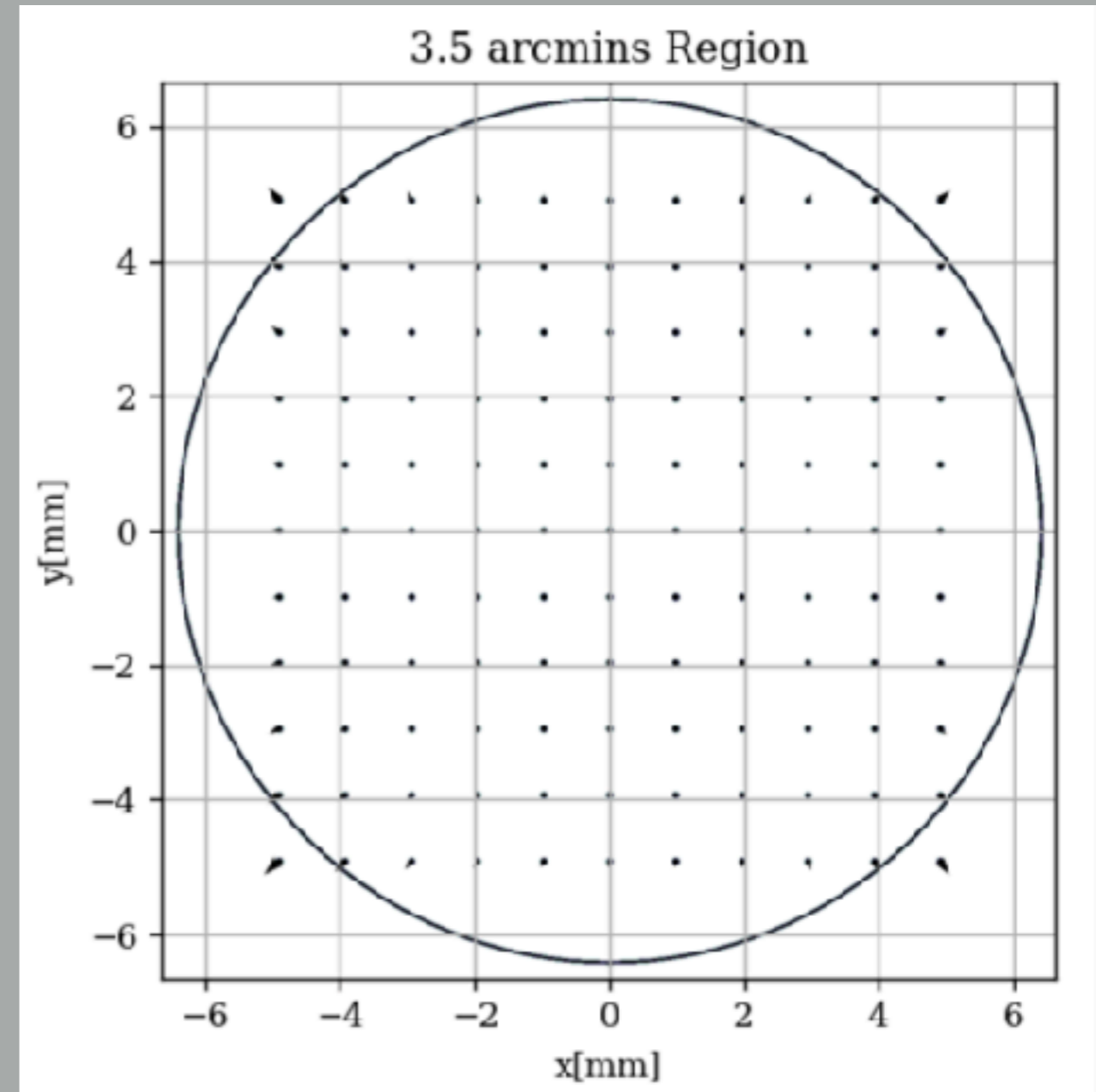
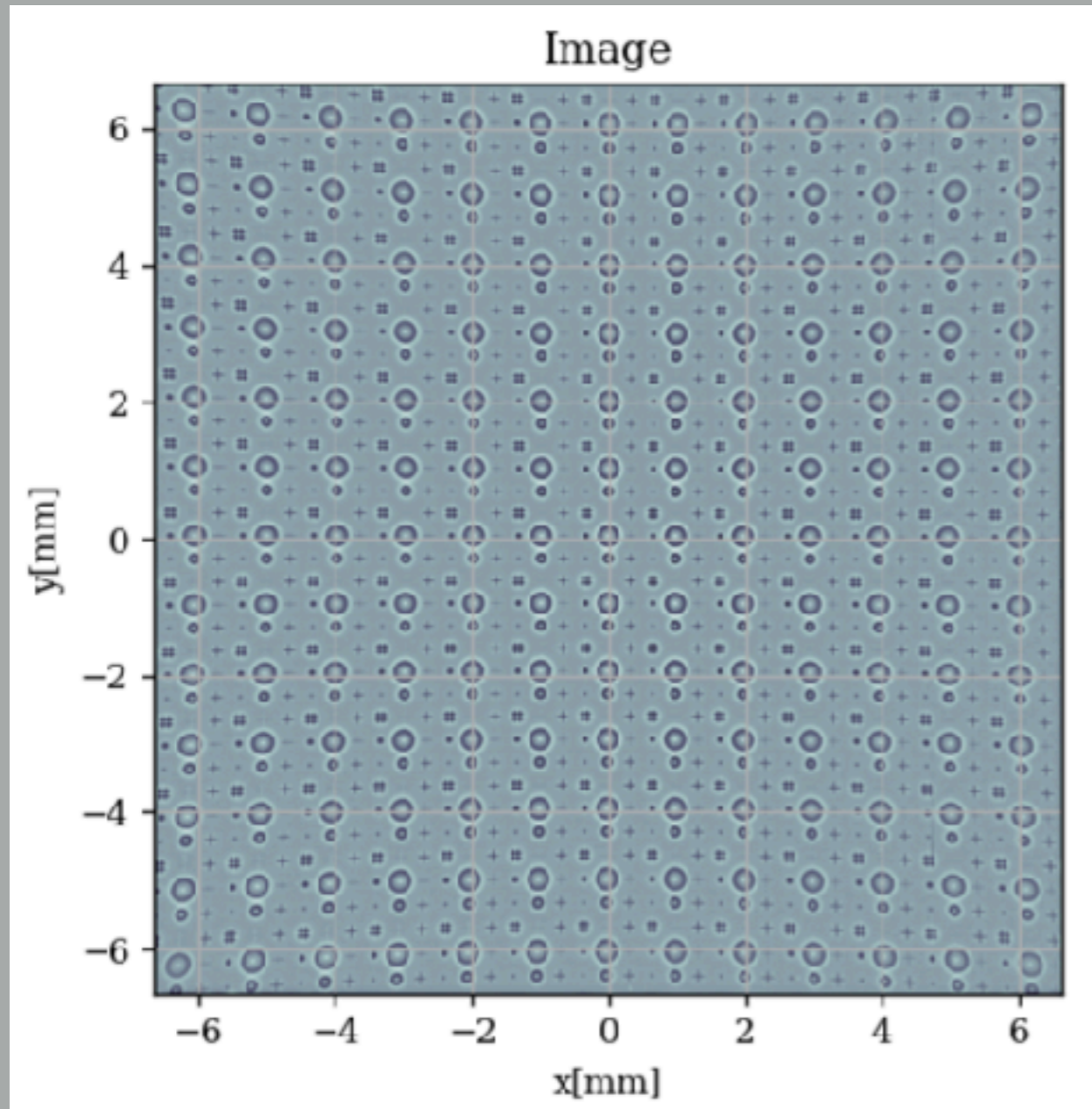


NIR Spectrograph

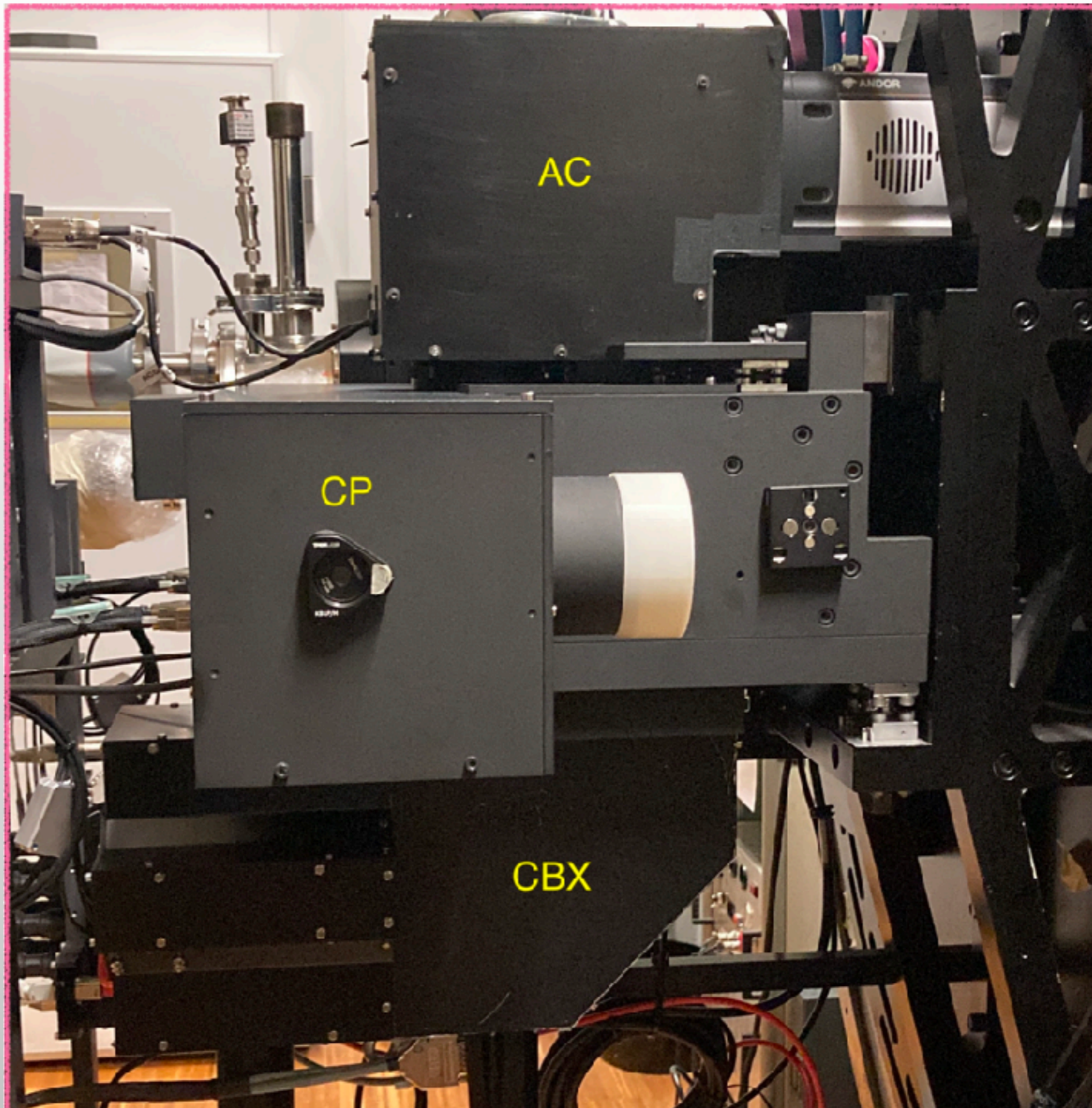




Acquisition Camera



Calibration Unit



Where We Started ...



Mounting the SOXS Flange



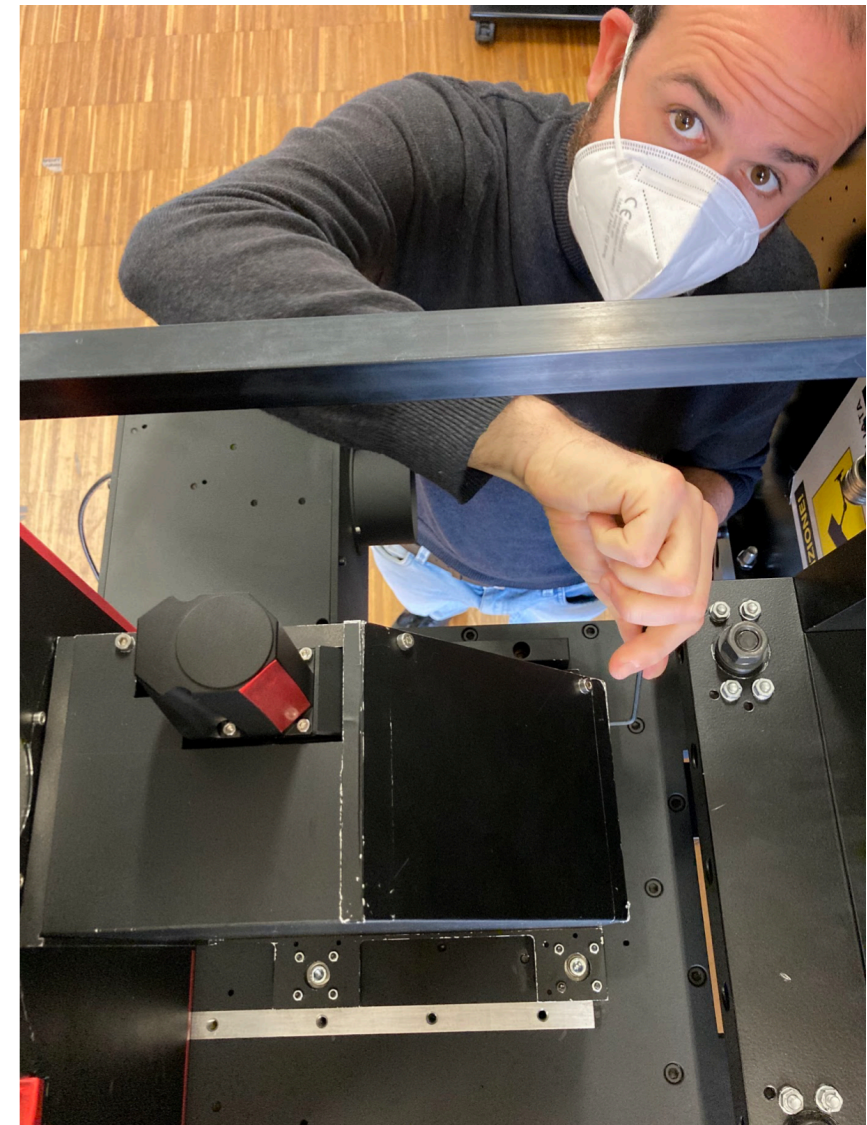
Corotator Holding Structure



Common Path Mounting



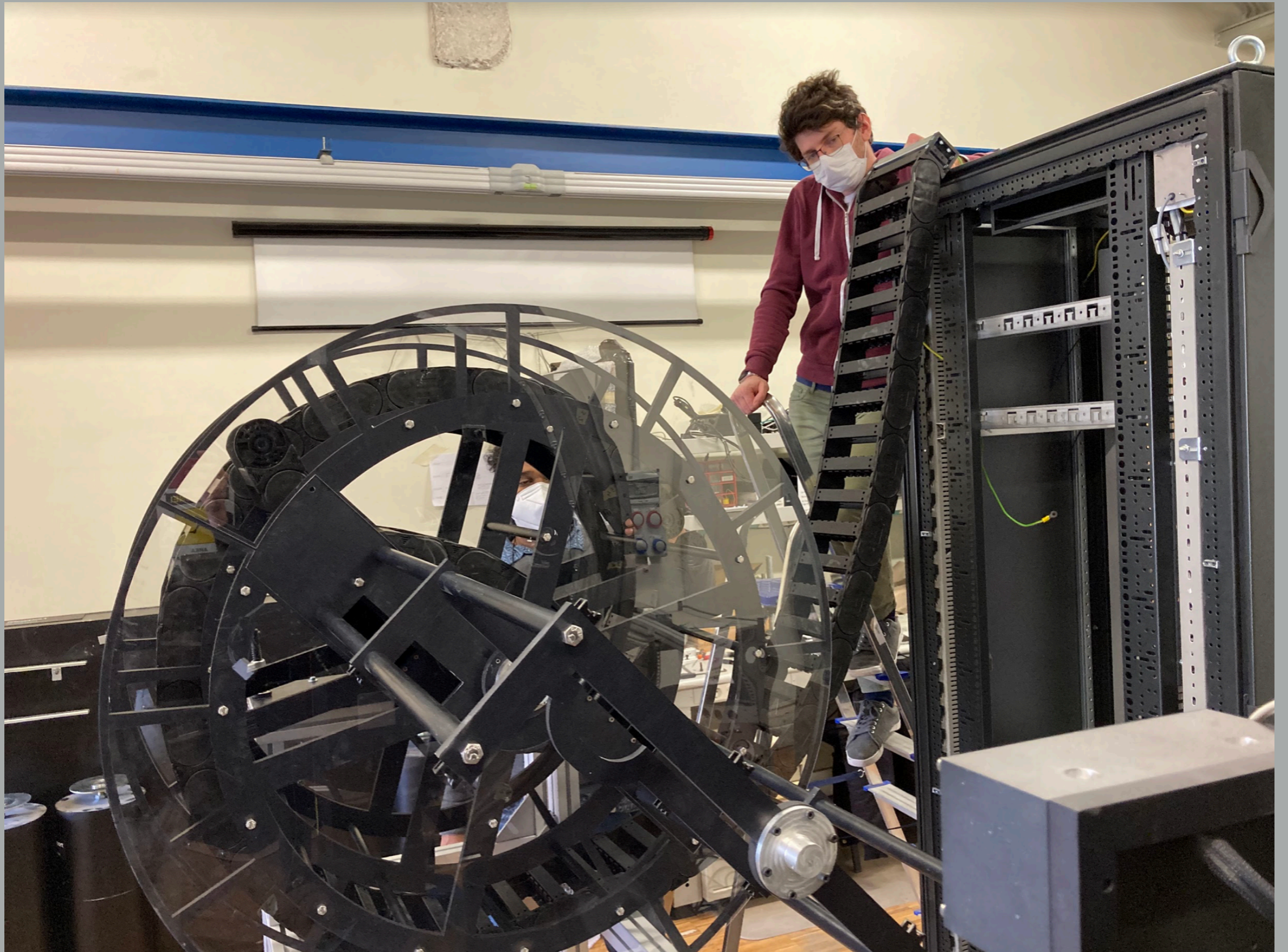
Mounting the Calibration Unit (old)



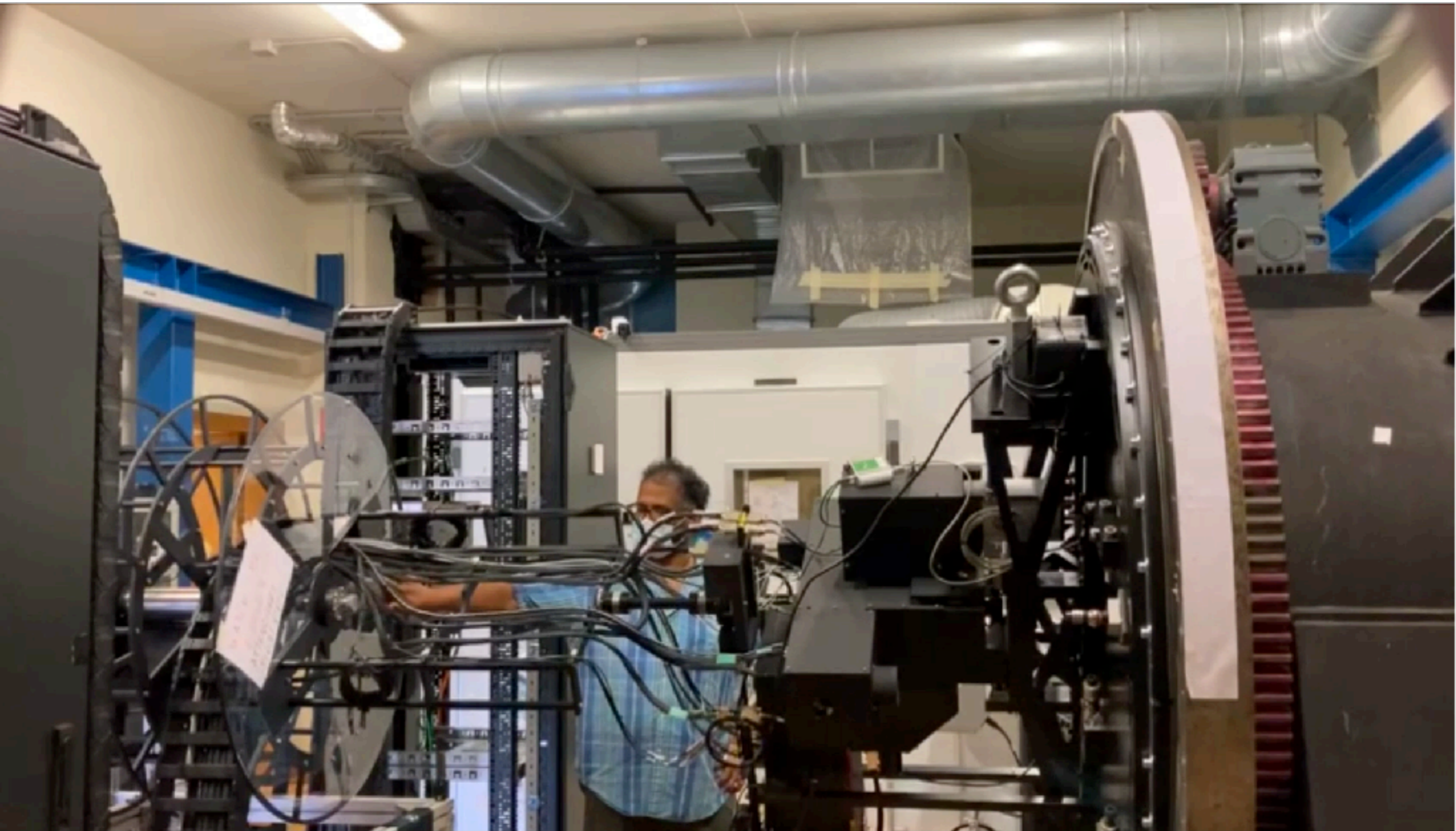
Mounting the Acquisition Camera



Cabinets and Cable Chains



Corotator co-rotating...



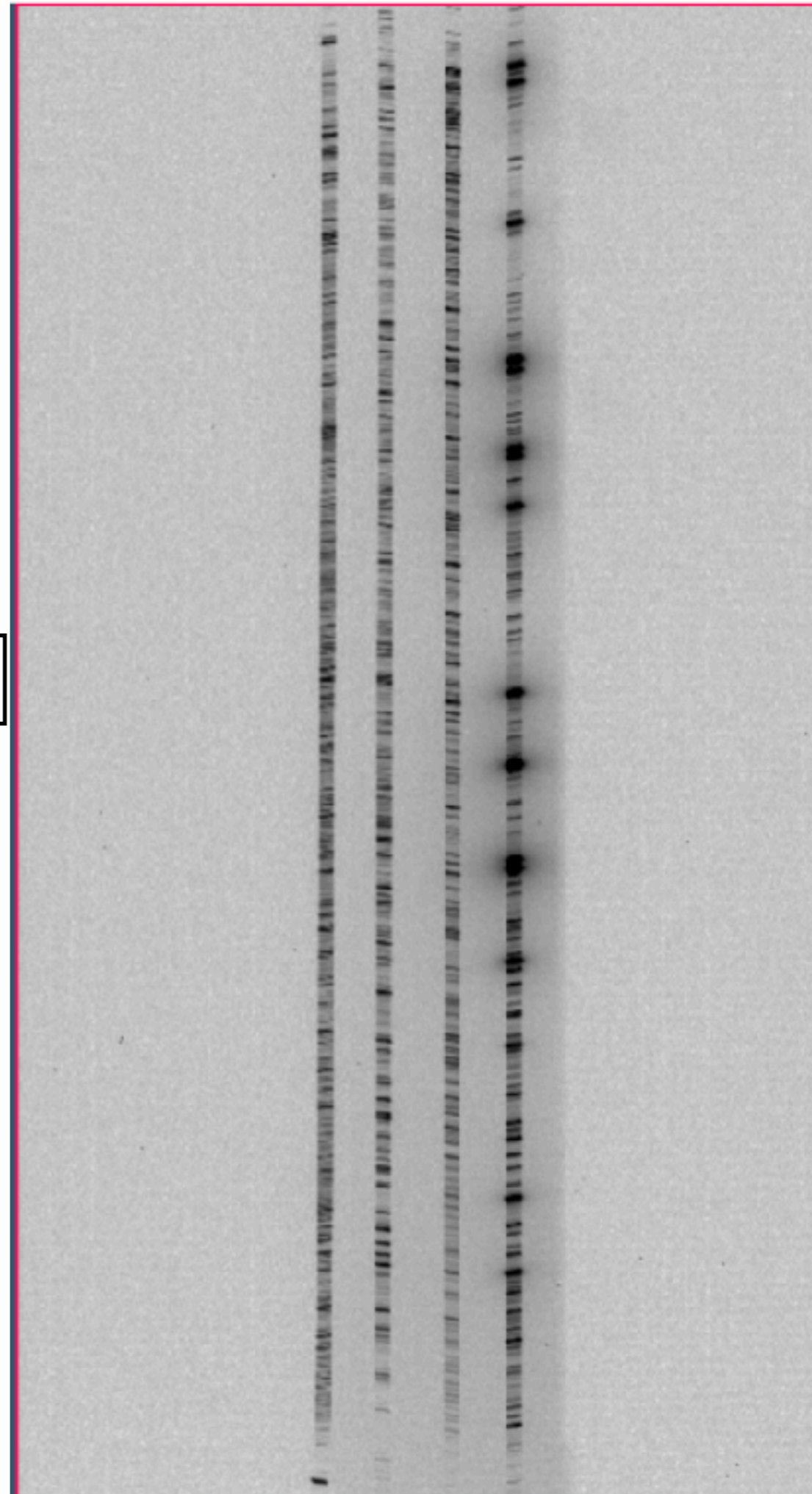
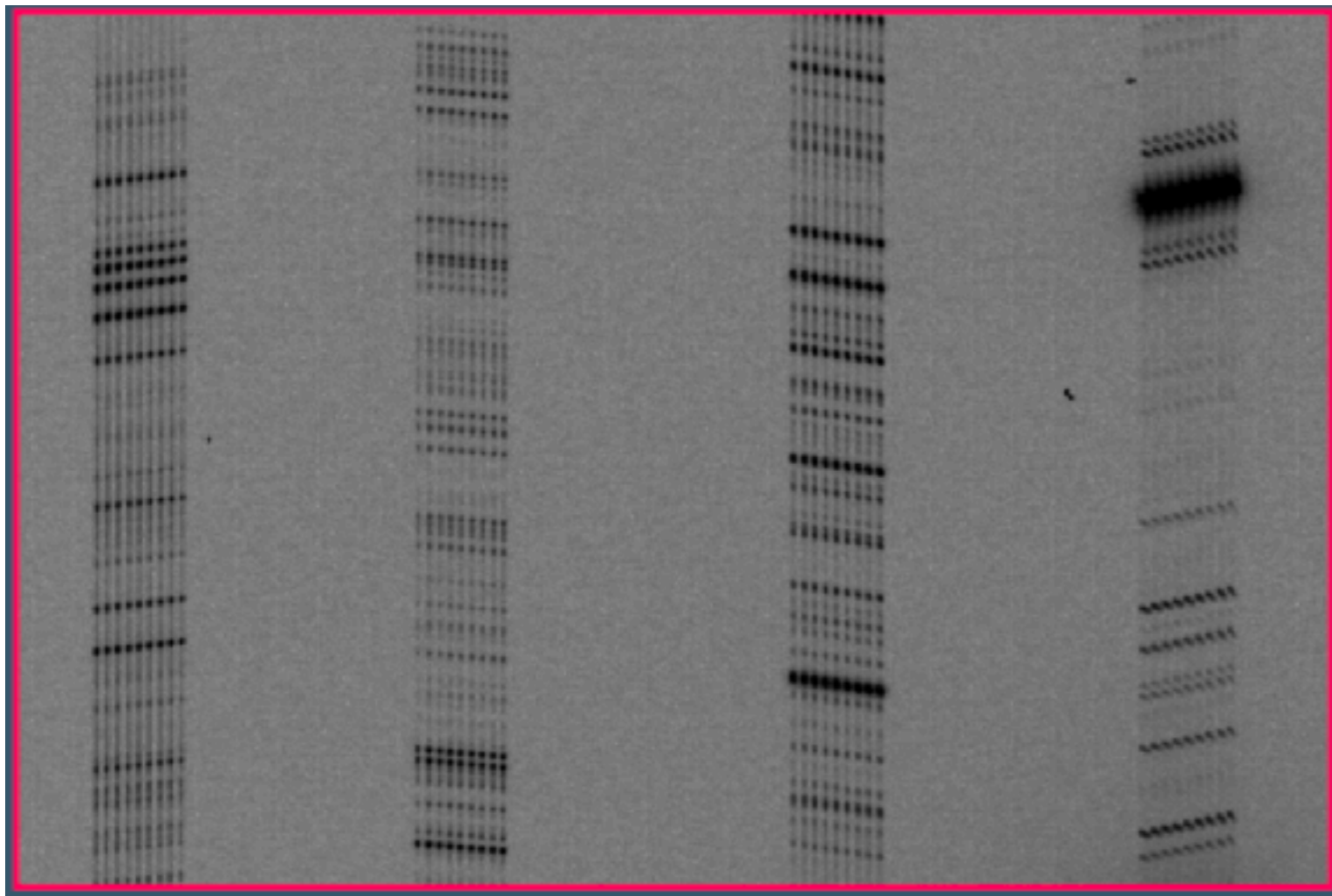
Finally everything together...



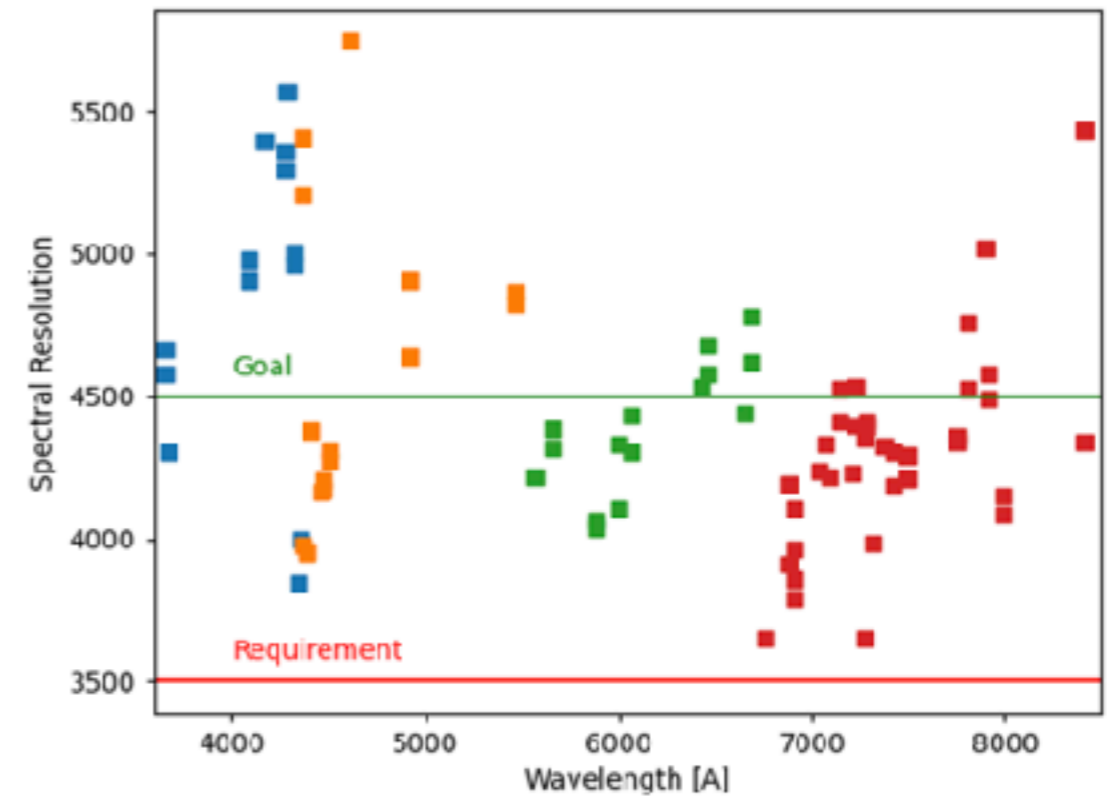
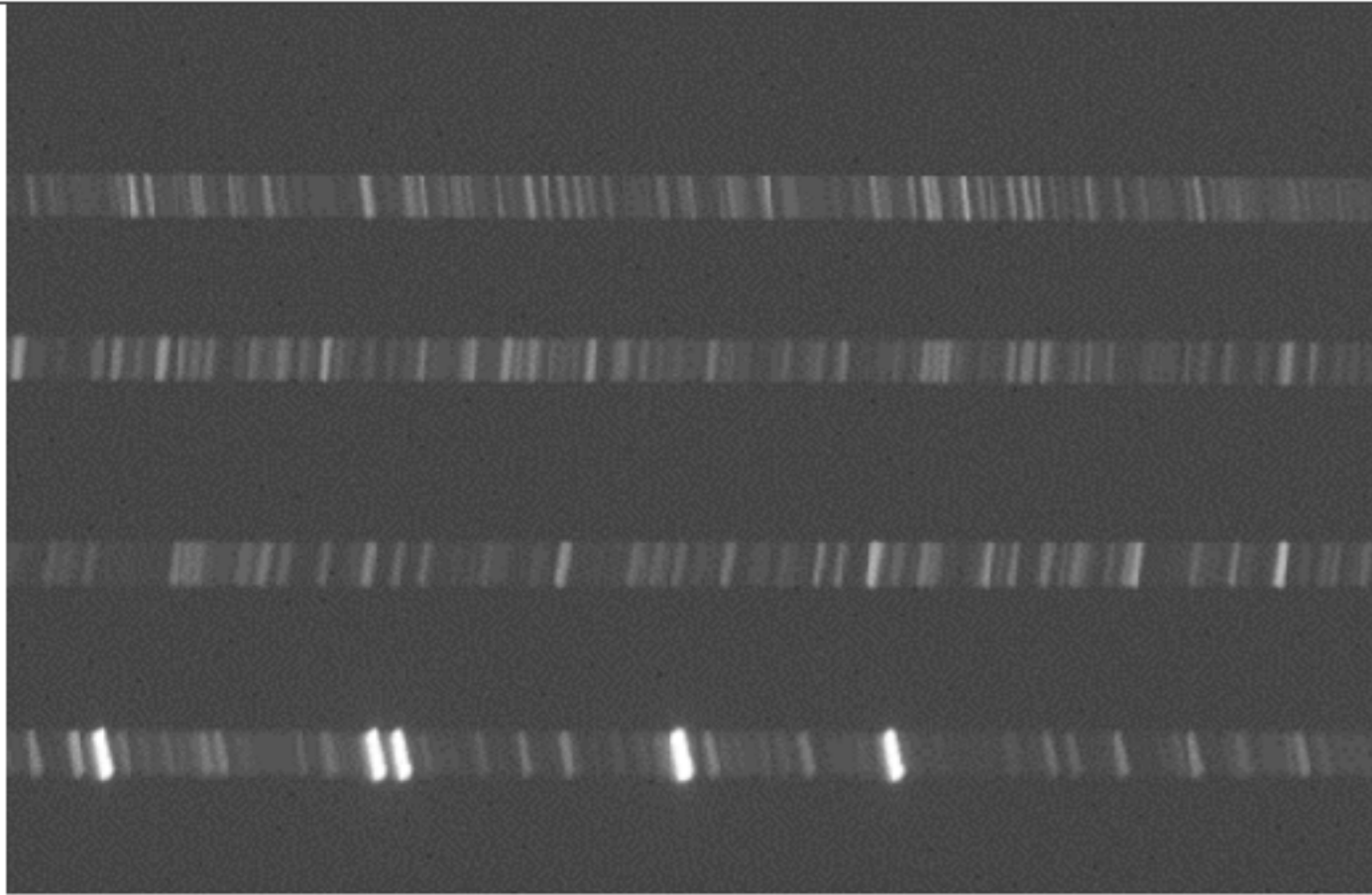
Now, some UVVIS spectra ...

1" slit, ThAr arc lamps with DIT of 300s

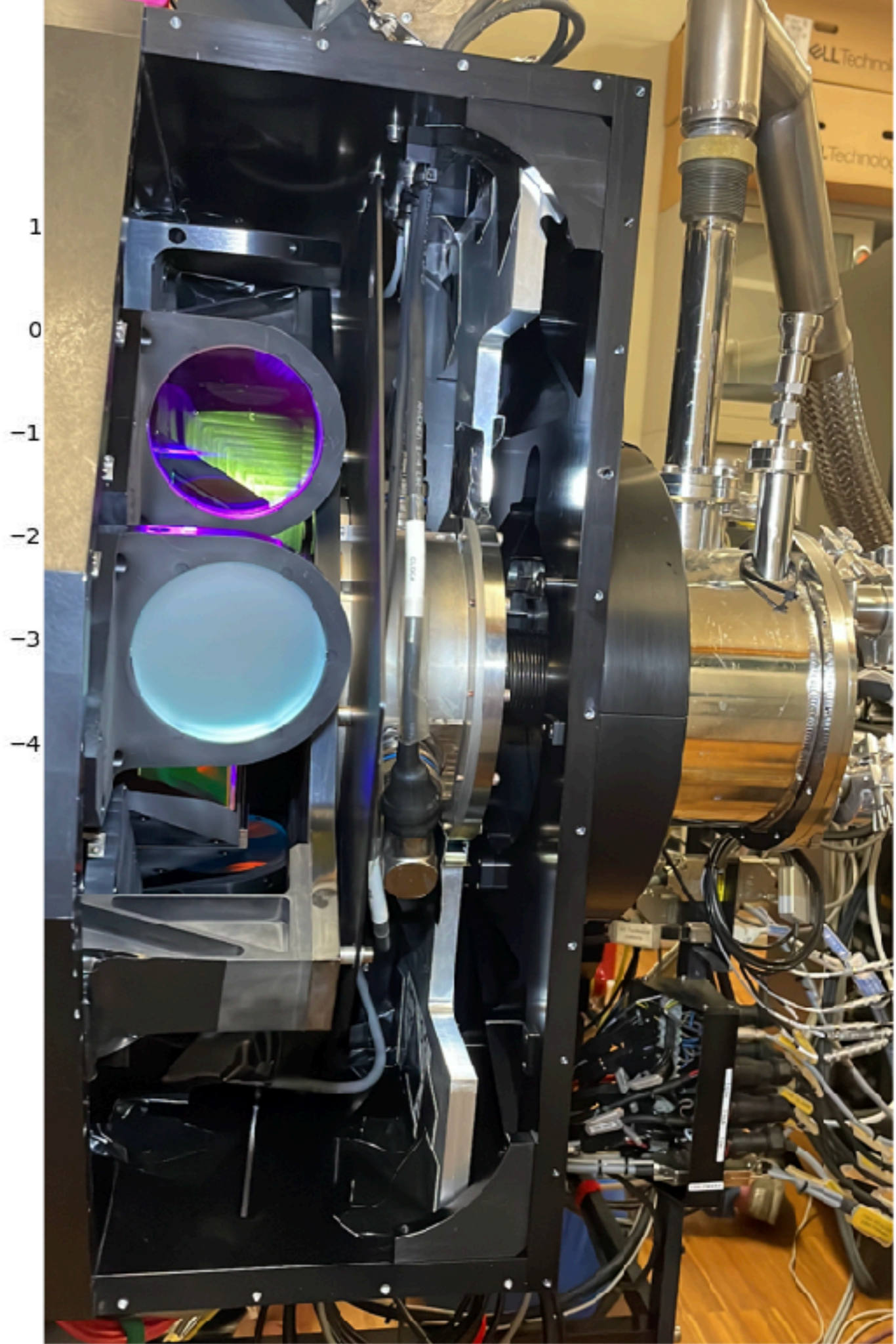
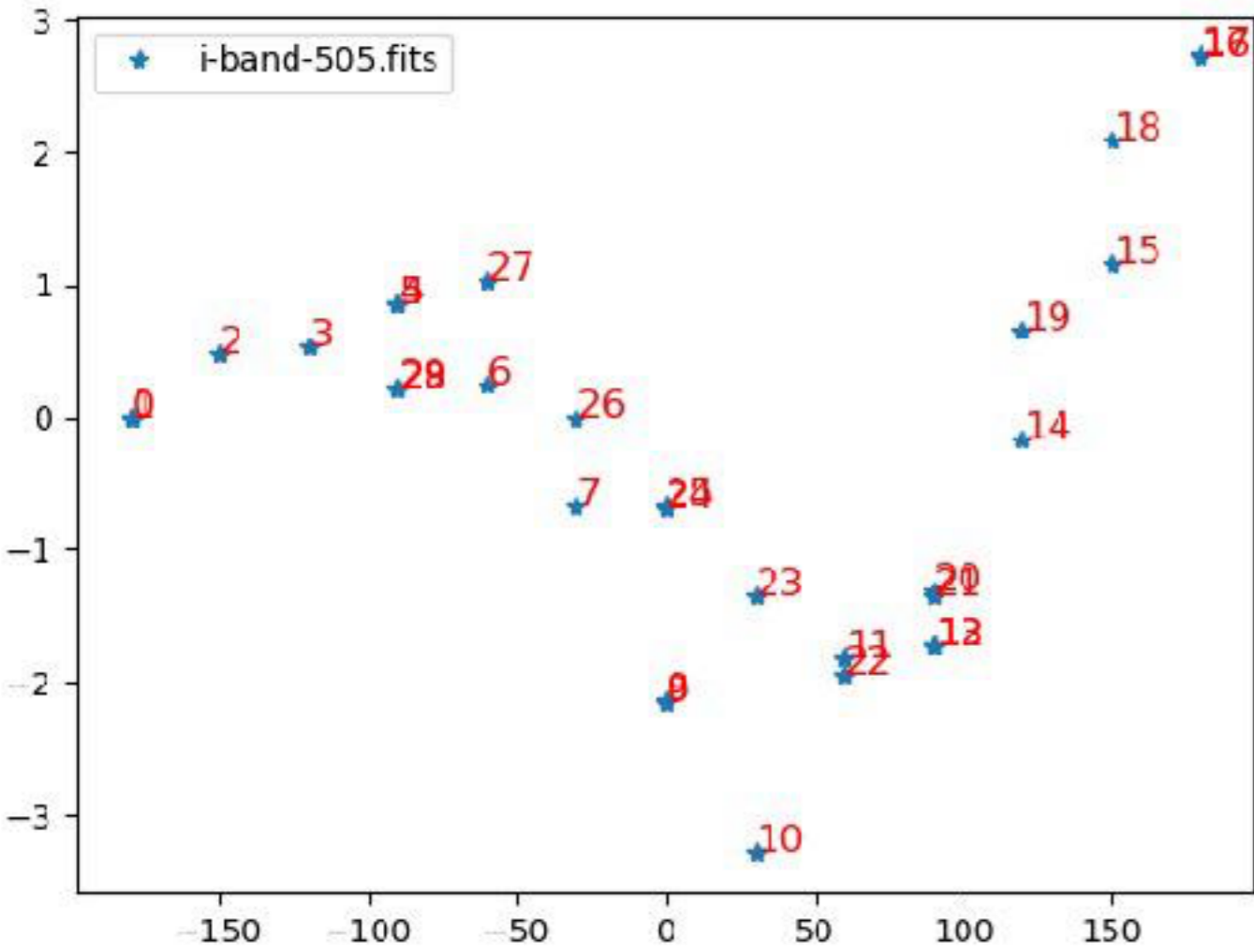
Multi-Pinhole, ThAr arc lamp with DIT of 300s



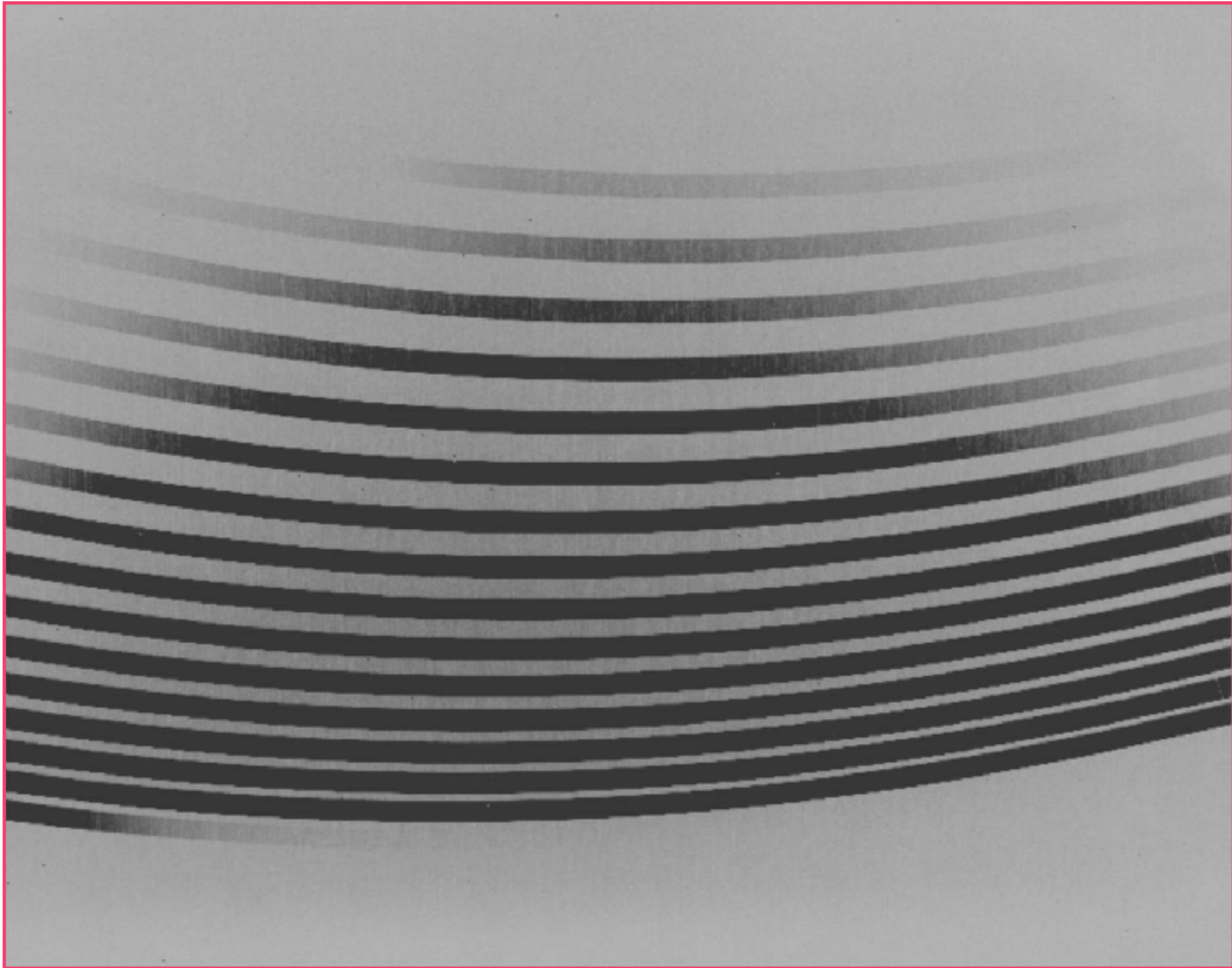
UVVIS Resolution



UVVIS Flexure

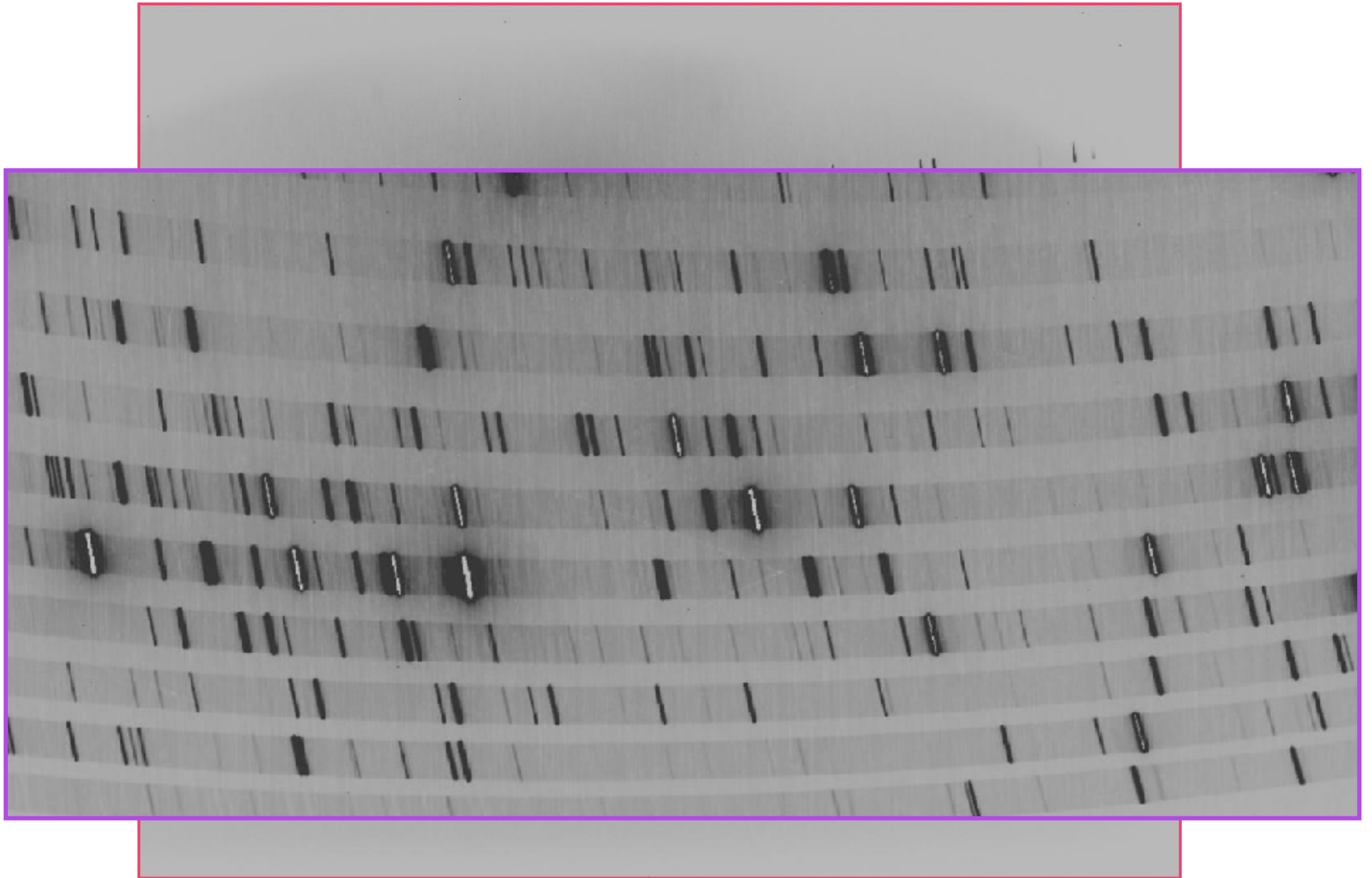


And NIR spectra ...



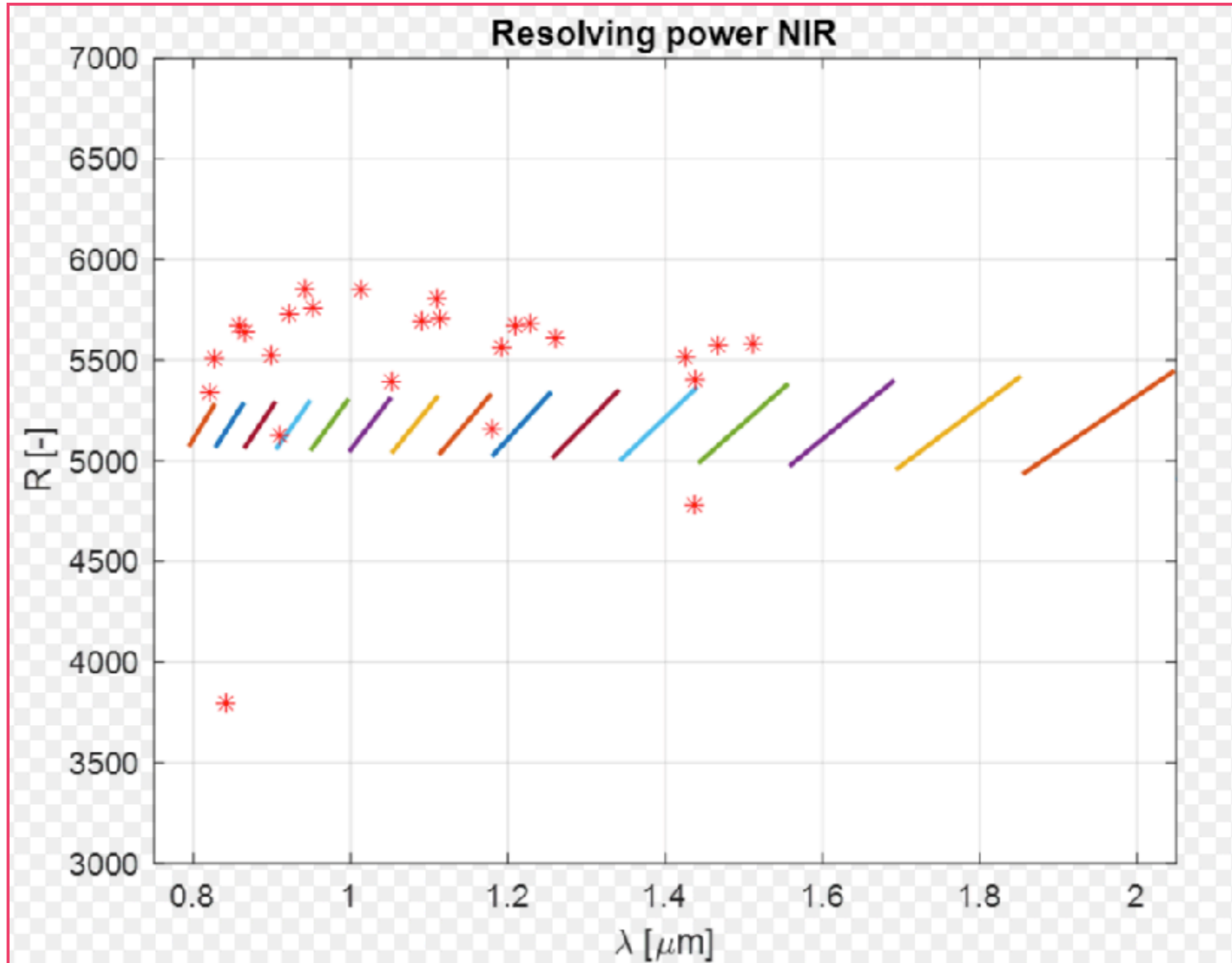
QTH continuum lamp with 0.5" slit and 5s exposure time

And NIR spectra ...

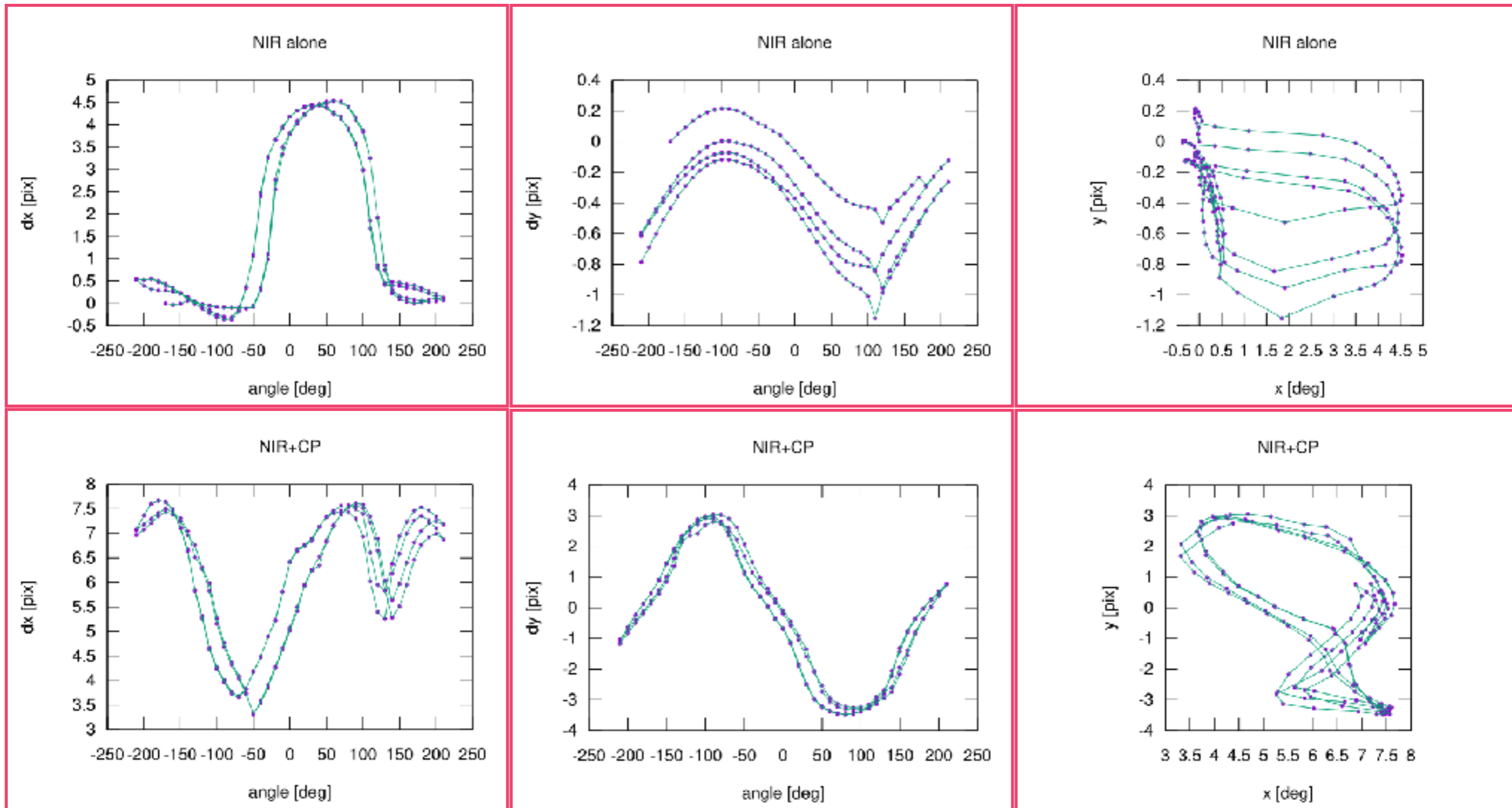


Arc lamps (Ar, Ne, Hg, & Xe) with 0.5" slit and 15s exposure time

NIR Resolution

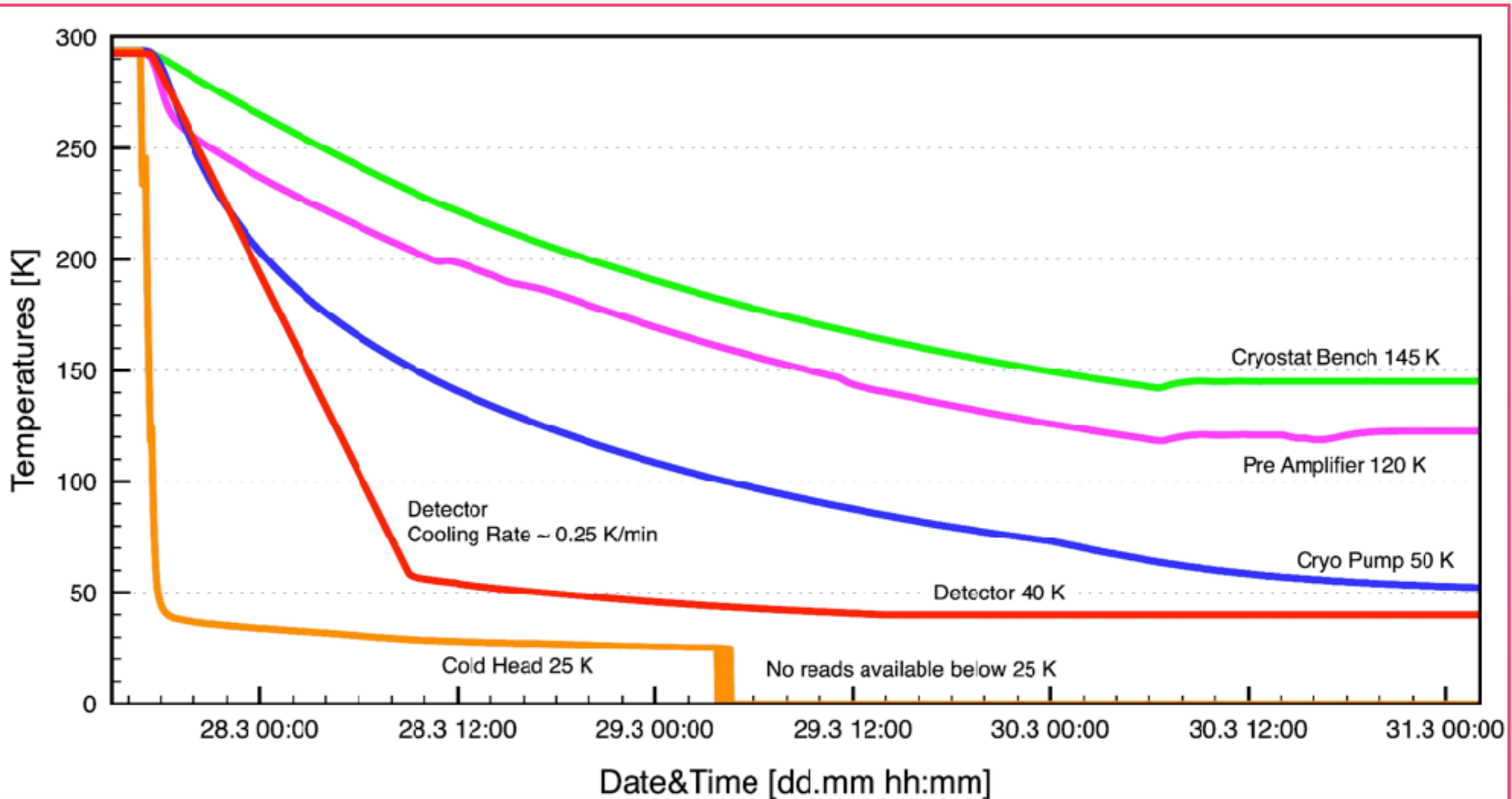


NIR Flexure

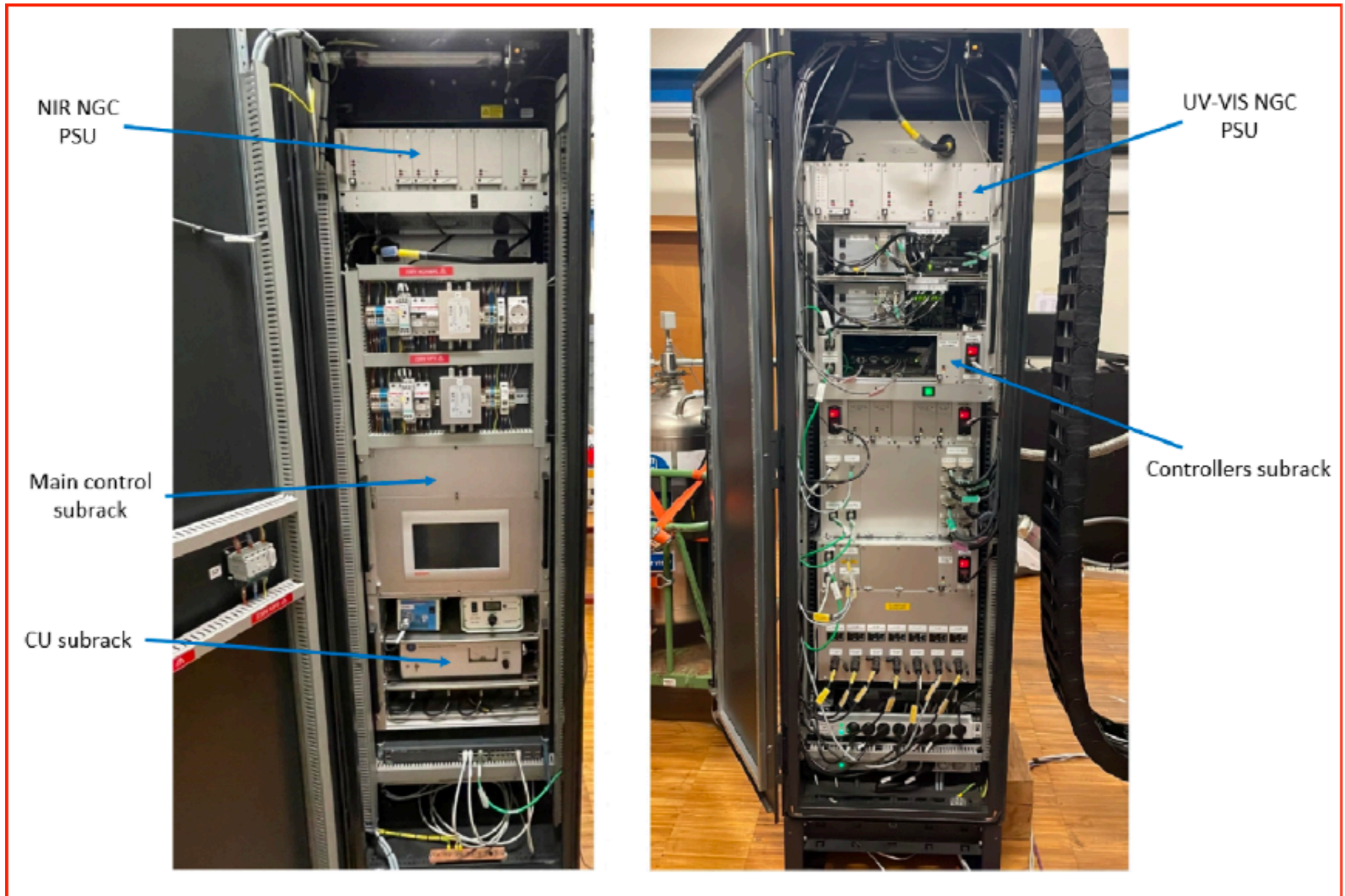


Measured flexure of the NIR alone (top) and NIR+CP (bottom)

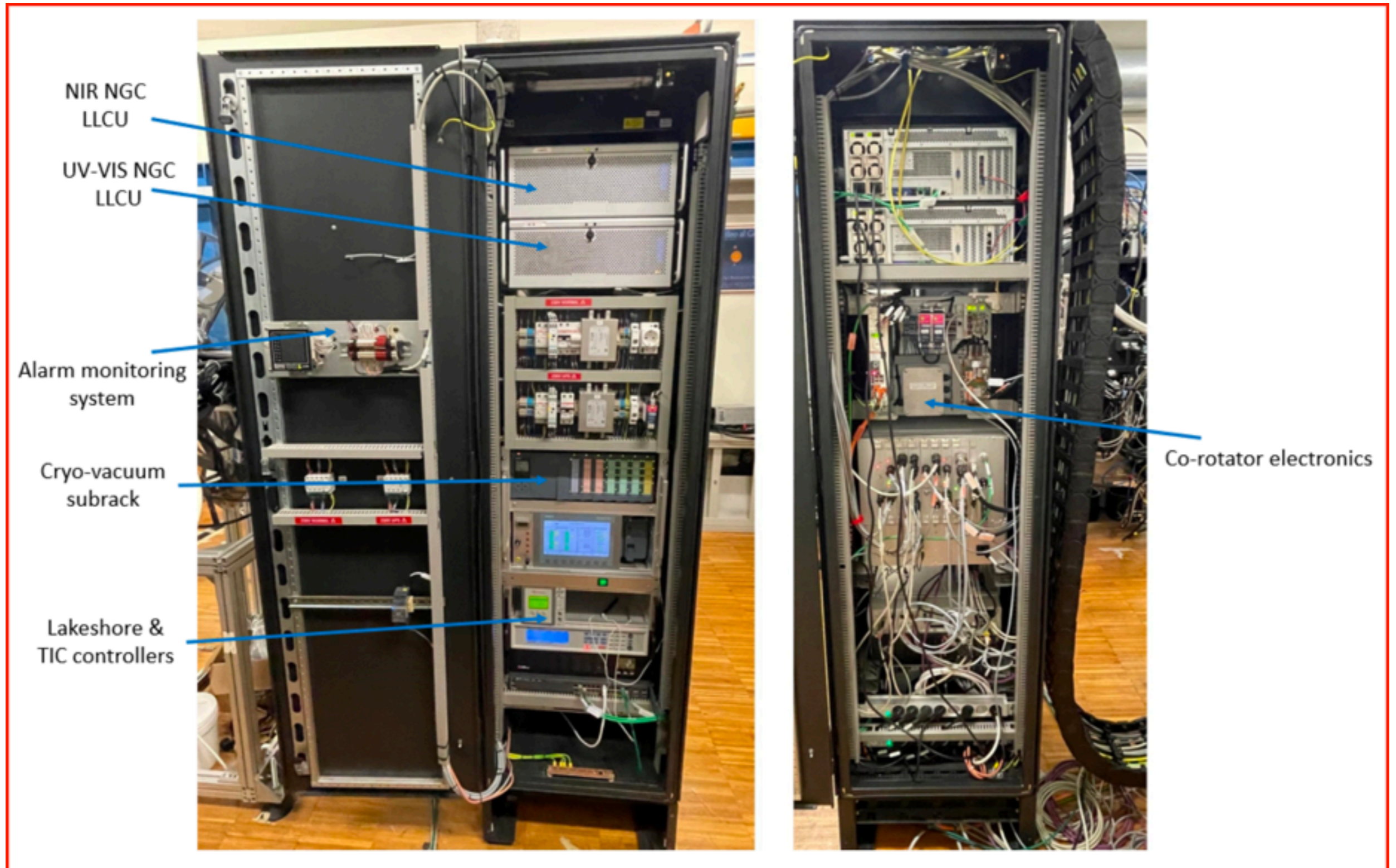
Cryo-Vacuum System



Instrument Control Electronics



Instrument Control Electronics




Instrument Control Software

SOXS Status - @wsxs

File Std. Options Help

SOXS Status



Ins Mode **SPECTROSCOPY**
 State **ONLINE**
 Substate **setup**

TCS

State **ONLINE** Tracking **ENABLED**
 Substate **IDLE** RA **115.531227**
 Access **NORMAL** DEC **-795132.131**

Shutter and slides

INSH	ONLINE	HW	CLOSED	
CUPS	ONLINE	HW	Free	35.00
CALS	ONLINE	HW	Calib	3.05
ACQS	ONLINE	HW	Spectroscopy	4.33

Calibration Lamps

QTH	ONLINE	HW	OFF
DEUT	ONLINE	HW	OFF
NEON	ONLINE	HW	OFF
ARGO	ONLINE	HW	OFF
XENO	ONLINE	HW	OFF
MERC	ONLINE	HW	OFF
THAR	ONLINE	HW	OFF

Other

CROT **ONLINE**
 Active
 Touched
 Moving
 Fault
 CPTS **31.0**

NIR - Near Infrared Spectrograph

NIR	ONLINE	NORMAL	Sub.	idle
Exposure	success	DIT	15.00	
Remaining	2	NDIT	1	
AFC2	ONLINE	HW	Mode	STAT
Current X	1013.0	Y	1014.5	
NFCC	ONLINE	HW		2.80
NISE	ONLINE	HW	Multi_Pinhole	-12.10

VIS - Visible Spectrograph

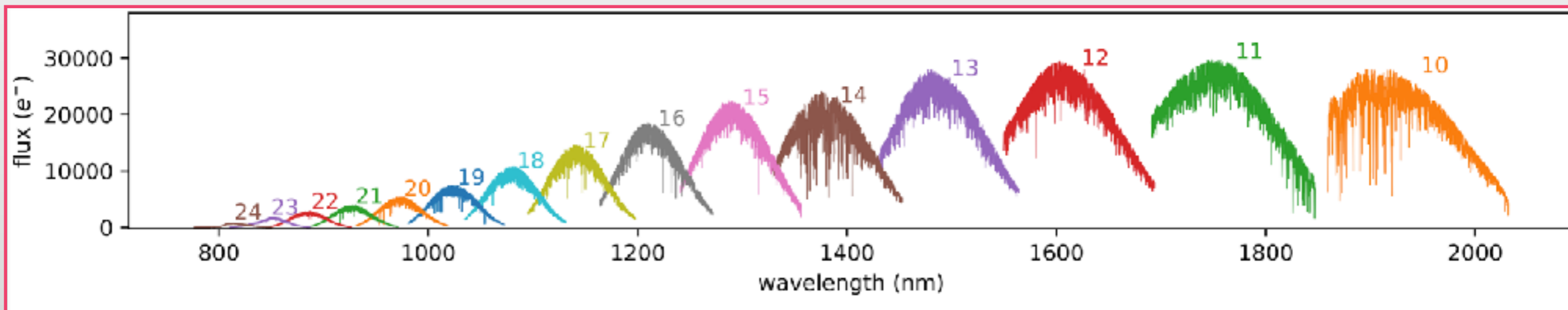
VIS	ONLINE	NORMAL	Sub.	idle
Exposure	128	UIT		
Remaining	0			
AFC1	ONLINE	HW	Mode	STAT
Current X	1000.0	Y	999.9	
ADC1	ONLINE	HW	STANDING	OFF
ADC2	ONLINE	HW	STANDING	OFF
WISE	ONLINE	HW	slit_0.5	11.33
TECH	ONLINE	HW	IDLE	

ACQ - Acquisition Camera

ACQ	ONLINE	HW	IDLE	
ACFW	ONLINE	HW	SDSS-u	205.00
AFOC	ONLINE	HW		9.70

DB

Data Reduction Software



All the spectral orders extracted from the NIR flat QTH calibration spectrum

SOXS Platform



SOXS timeline

Date	Activity	
June 2014	ESO Call for New Instruments at NTT	✓
February 2015	Proposal Submission	✓
May 2015	SOXS selected by ESO (out of 19 proposals)	✓
October 2016 - July 2017	INAF approval + PDR phase	✓
August 2017 - Sept 2018	FDR Phase	✓
October 2018 - February 2022	Procurement, Sub-system AIVT (delay due to Covid-19)	✓
March 2022	Integration Started	✓
March 2022 - July 2024	Integration & System level tests	✓
July 2024	Preliminary Acceptance Europe start	✓
November 2024	PAE complete and start packing	✓
December 2024	Shipping SOXS to La Silla, Chile	✓
January 2025	AIV @ La Silla	✓
March 2025	Commissioning Start	✓



The SOXS Instrument @ INAF-OAPD

Extra Slides

Name	Initials	Institution	D 1	D 2	D 3	D 4	D 5	D 6	D 7	D 8	D 9	D 10	D 11	D 12	D 13	D 14	D 15	D 16	D 17	D 18	D 19	D 20	D 21	D 22	D 23	D 24	D 25	D 26	D 27	D 28	D 29	D 30	D 31	D 32	D 33	D 34	D 35	Total		
Kalyan Radhakrishnan	KRA	INAF-Padova	X	X	X	X	X	X	X	X	X	X	X	X																					X	X	X	X	17	
Federico Battaini	FBA	INAF-Padova	X	X	X	X	X	X	X	X															X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	21
Simone Di Filippo	SDF	INAF-Padova	X	X	X	X	X	X	X	X																													8	
Riccardo Claudi	RCL	INAF-Padova									X	X	X	X	X	X	X	X	X																				9	
Lorenzo Cabona	LCA	INAF-Padova													X	X	X	X	X	X	X	X	X	X															10	
Davide Ricci	DRI	INAF-Padova	X	X	X	X	X	X	X	X	X	X	X										X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	22	
Software Person	SWP	INAF-Padova												X	X	X	X	X	X	X	X	X	X												X	X	X	X	13	
Matteo Aliverti	MAL	INAF-Brera	X	X	X	X	X	X	X						X	X	X	X	X	X	X	X	X																16	
Matteo Genoni	MGE	INAF-Brera																		X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	X	14		
Sergio Campana	SCA	INAF-Brera																																	X	X	X	X	4	
Paolo D'Avanzo	PDA	INAF-Brera																																	X	X	X	X	4	
Pietro Schipani	PSC	INAF-Napoli	X	X	X	X	X	X	X	X	X	X													X	X	X	X	X	X	X	X	X	X	X	X	X	X	23	
Mirko Colapietro	MCO	INAF-Napoli	X	X	X	X	X	X	X																														7	
Sergio D'Orsi	SDO	INAF-Napoli	X	X	X	X	X	X	X																														7	
Salvatore Savarese	SSA	INAF-Napoli																							X	X	X	X	X	X	X	X	X	X	X	X	X	X	13	
Gullia Capasso	GCA	INAF-Napoli																							X	X	X	X	X	X	X	X	X	X	X	X	X	X	13	
Fabrizio Vitelli	FVI	INAF-Rome													X	X	X	X	X	X	X	X	X																9	
Francesco D'Alessio	FDA	INAF-Rome													X	X	X	X	X	X	X	X	X																9	
Salvatore Scuderi	SSC	INAF-Milan											X	X	X	X	X	X	X	X	X	X	X																12	
Antonio Micciché	AMI	INAF-Catania													X	X	X	X	X	X	X	X	X																9	
Sagi Ben-Ami	SBA	Weizmann								X	X	X	X	X	X																								6	
Adam Rubin	ARU	Weizmann								X	X	X	X	X	X																								6	
Ofir Hershko	OHE	Weizmann								X	X	X	X	X	X																								6	
Rosario Cosentino	RCO	TNG									X	X	X	X	X																								5	
Rachael Bruch	RBR	Weizmann																								X	X	X	X	X	X	X	X	X	X	X	X	X	12	
Gulliano Pignata	GPI	Universidad An																								X	X	X	X	X	X	X	X	X	X	X	X	X	6	
		Total/day	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	8	7	7	7	8	8	8	8	8	8	8	9	9	9	9		
Total number days for all people on-site																																							281	

Summary

- SOXS is a Single object spectrograph offering simultaneous spectral coverage in UV-VIS and NIR, with imaging capabilities in the visible.
- It will be a precious facility for the spectroscopic follow-up of transient sources.
- SOXS installation, commissioning, and science verification will be in the year of 2024.

Mounting the UVVIS Spectrograph on the SOXS Flange



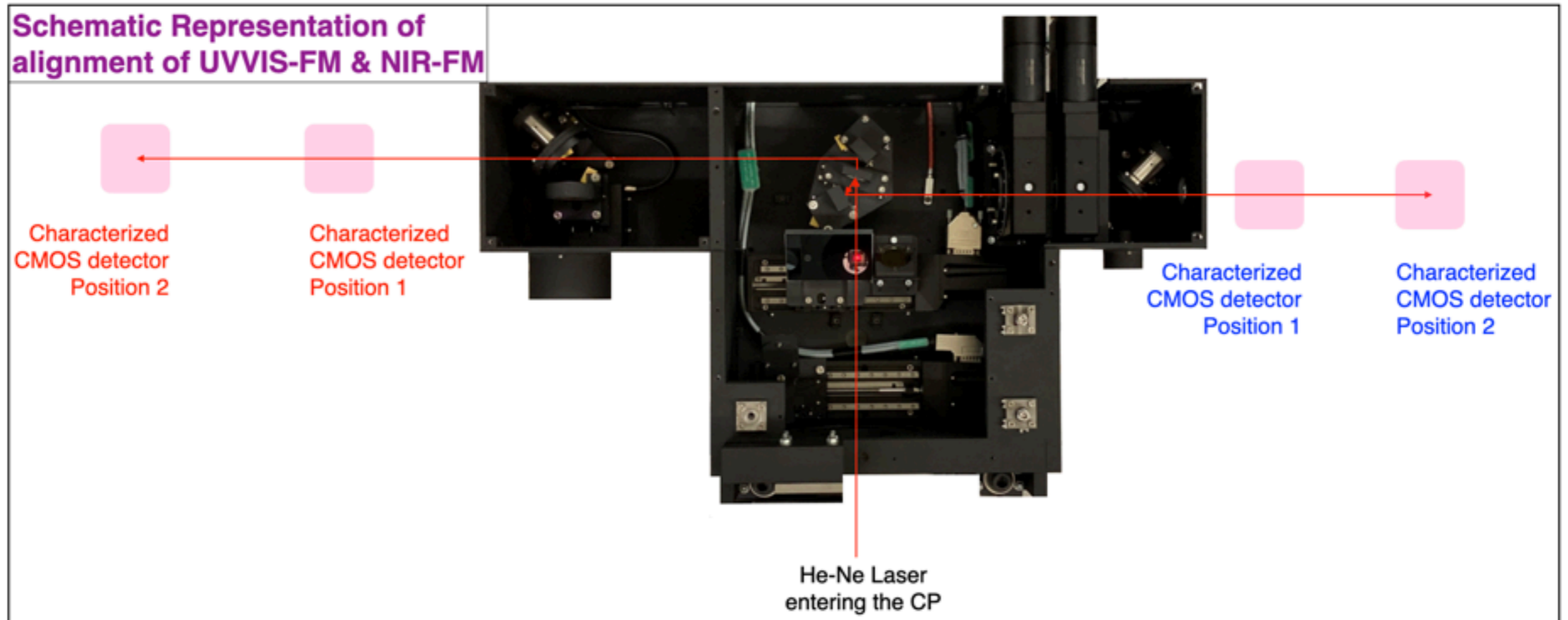
Thanks

SOXS Consortium

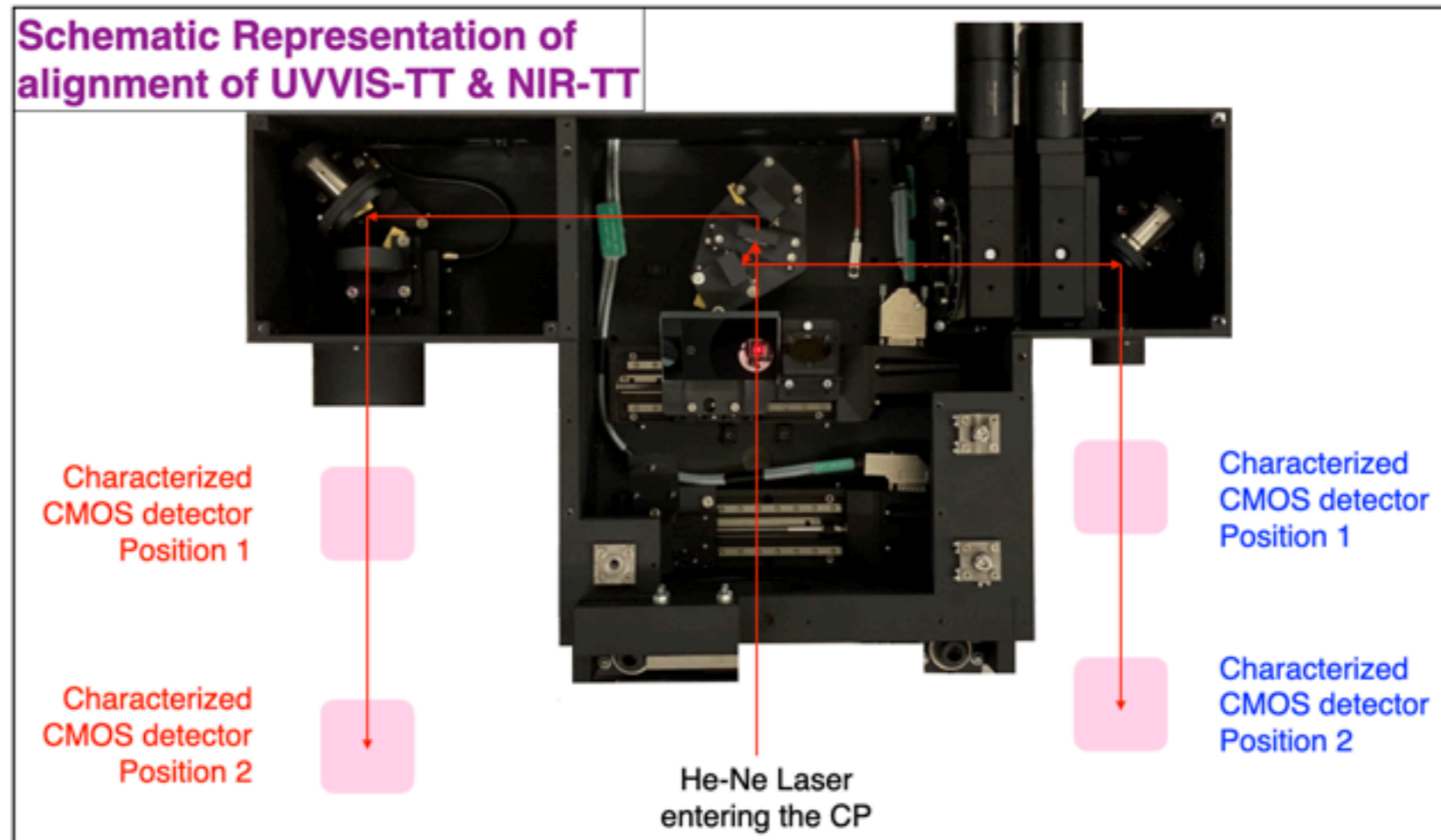
- INAF - Common Path sub-system, NIR-spectrograph, control software and electronics, vacuum and cryogenics, detector control.
- Weizmann Institute of Science - UV-VIS spectrograph
- Universidad Andres Bello & Instituto Milenio de Astrofisica - Acquisition Camera sub-system
- Turku University - Calibration Unit sub-system
- Queen's University - Data reduction



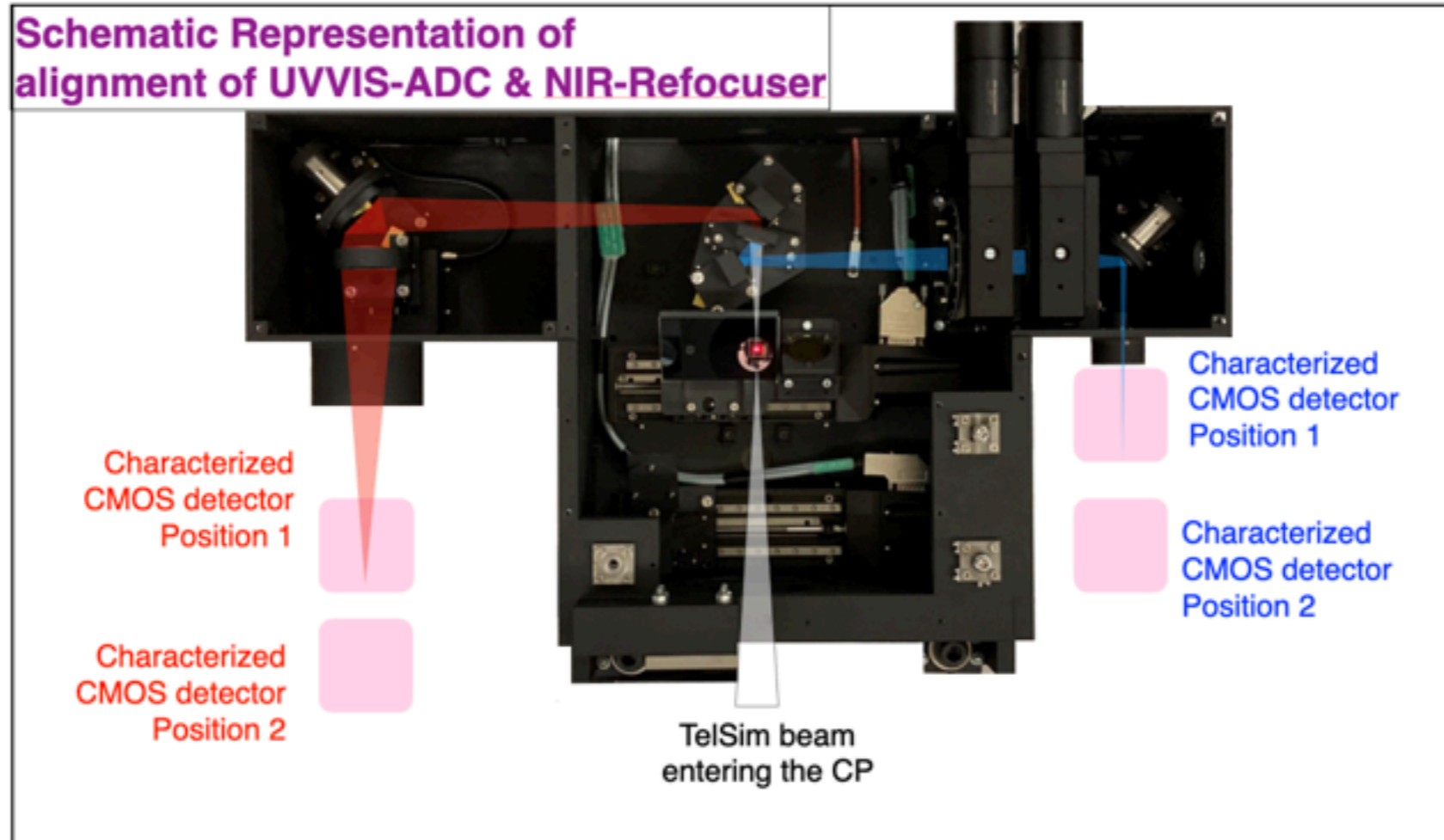
Common Path Alignment



Common Path Alignment



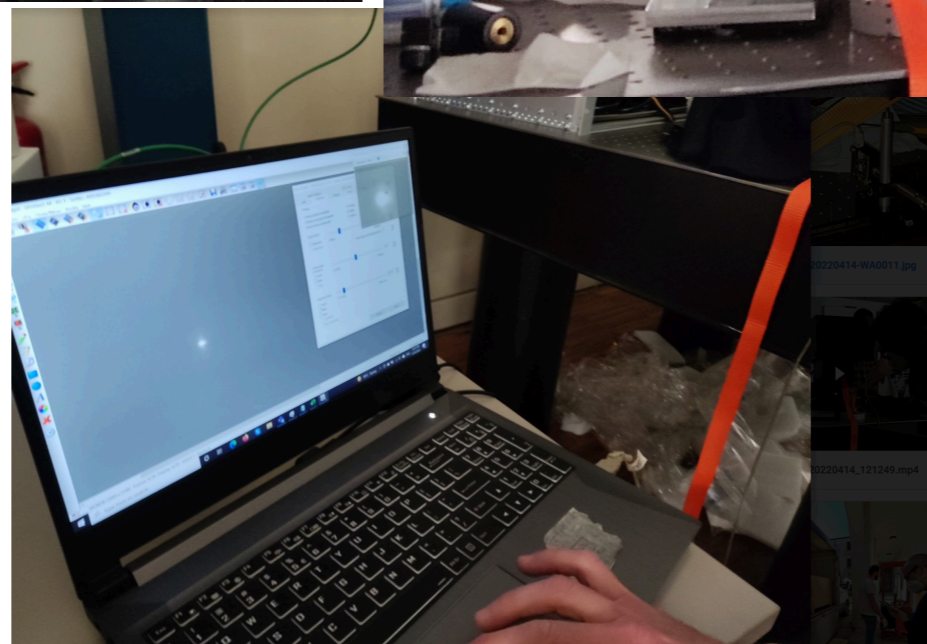
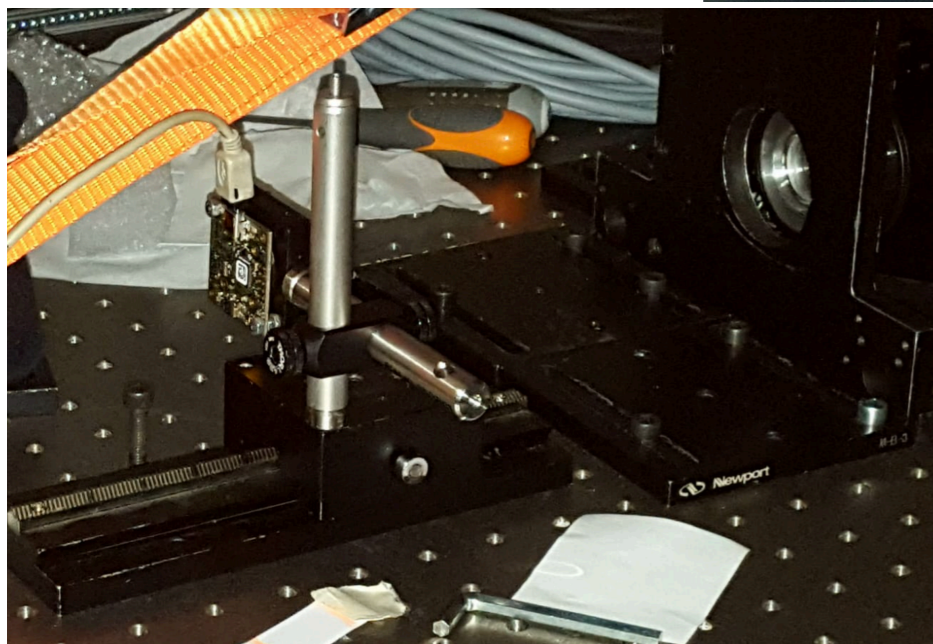
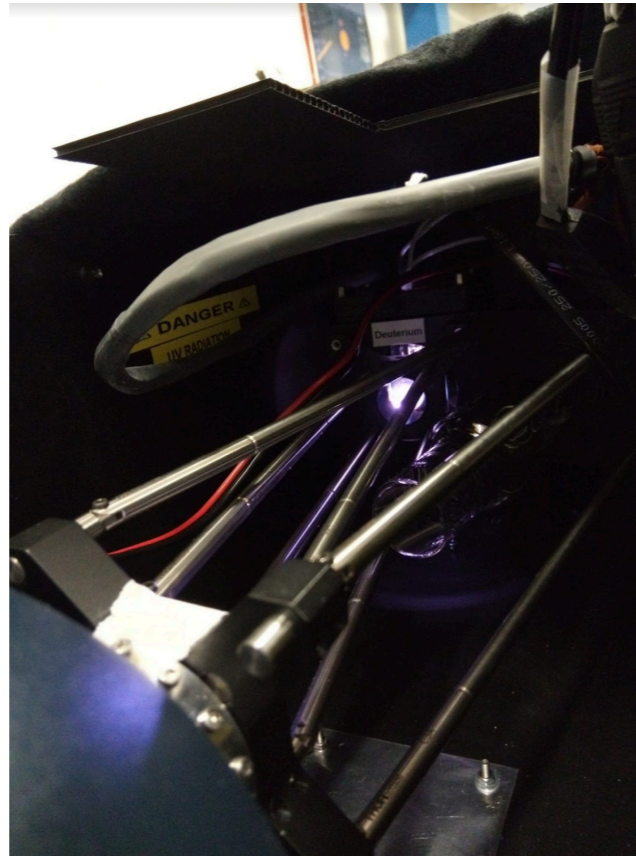
Common Path Alignment



Integration and Testing



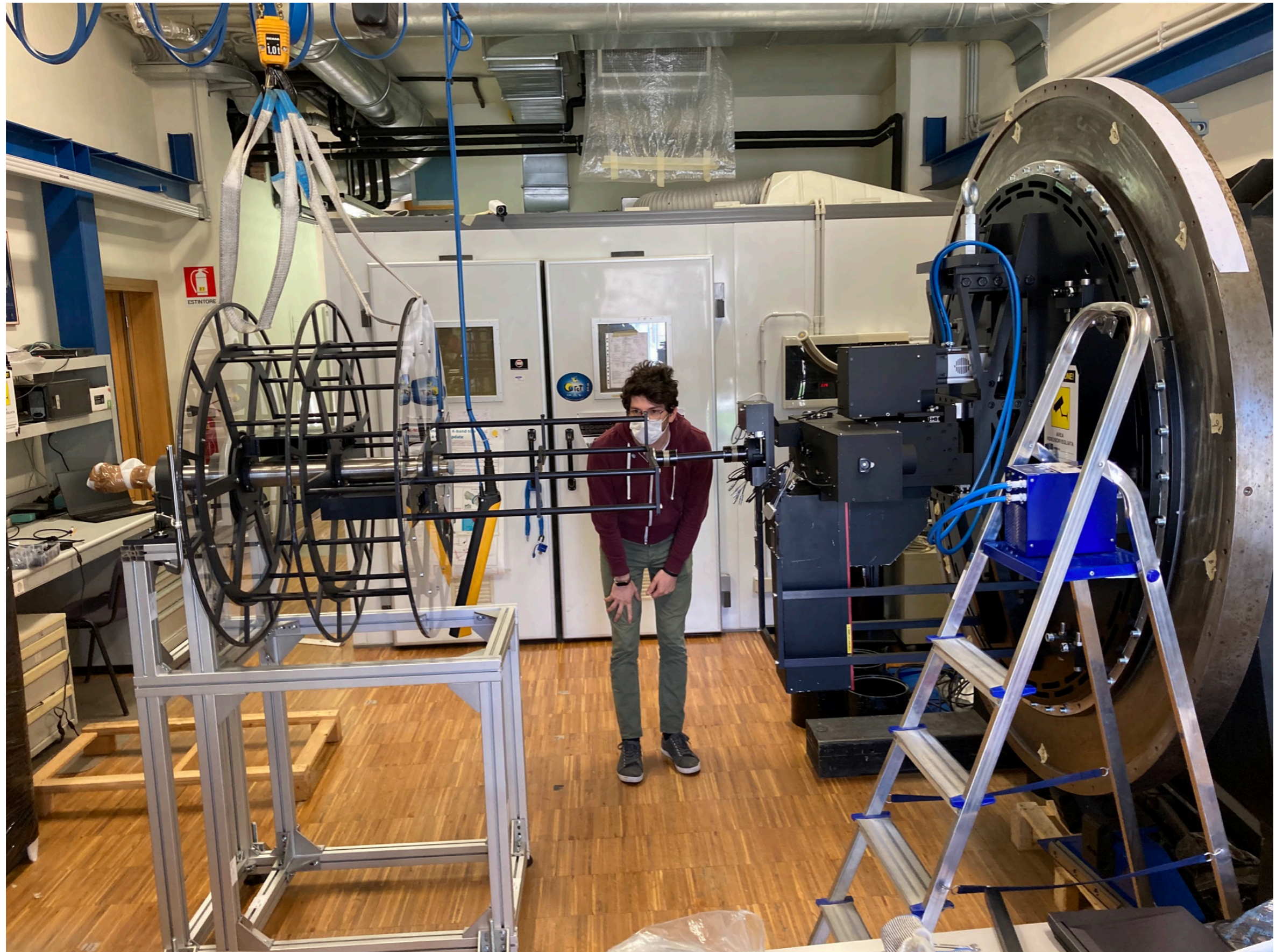
Integration and Testing

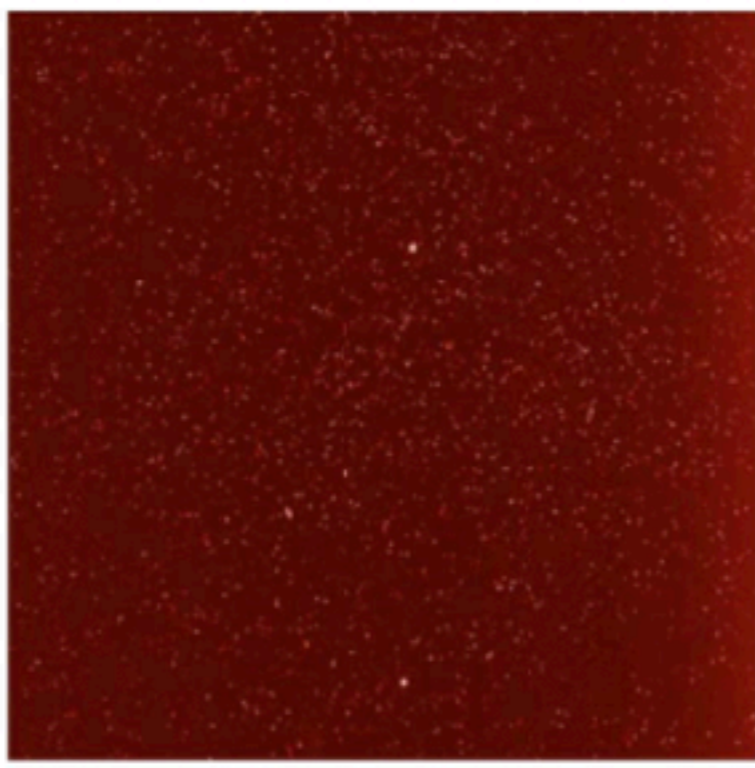


Integration and Testing

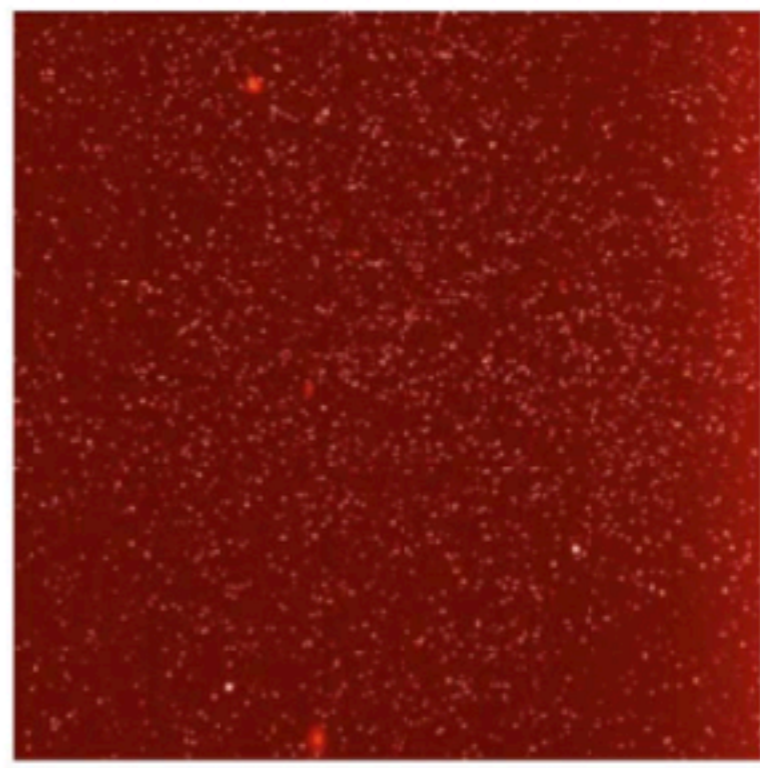


Integration and Testing

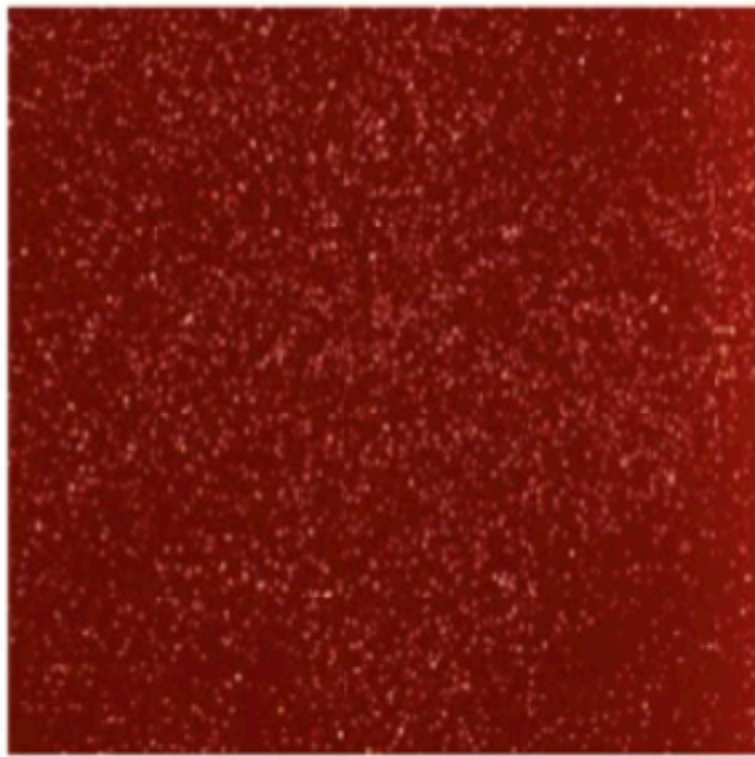




T = 40K



T = 60K



T = 80K



T = 100K

Figure 2 – HAWAII 2RG Cosmetics at 40, 60, 60 and 100 K

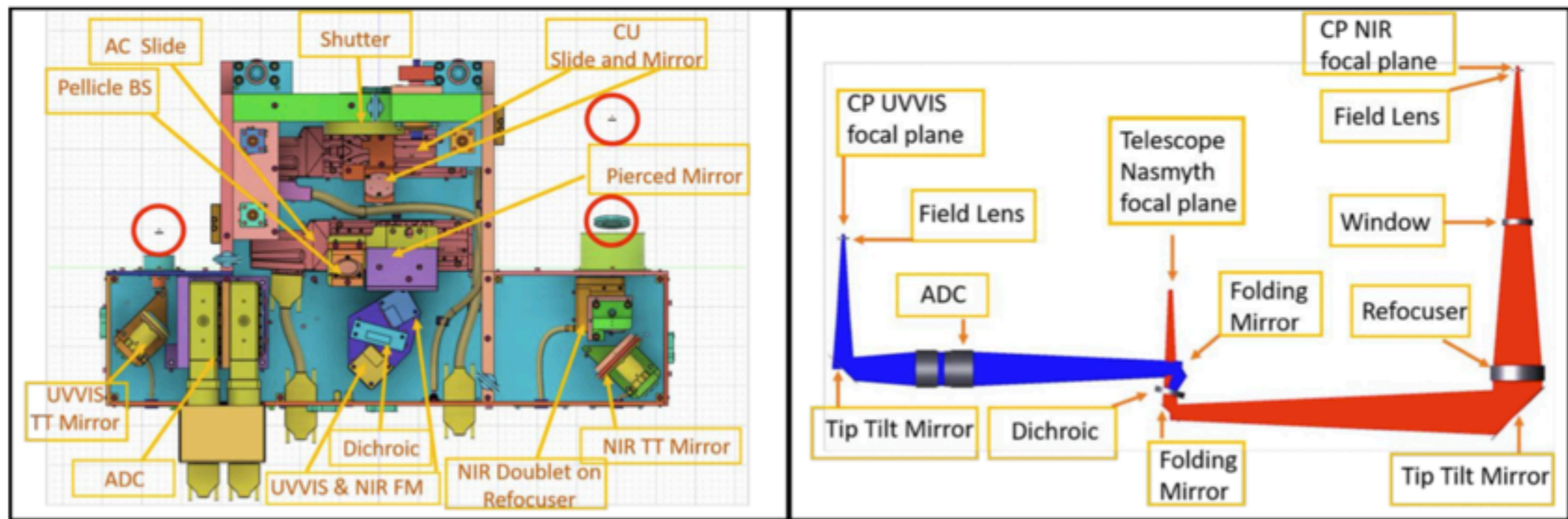


Figure 1. *Left panel:* Common Path CAD image displaying its components. *Right panel:* The CP light path.

The optical components UVVIS field lens, NIR window, and NIR field lens (marked within red circles in Figure 1) are formally a part of the CP, but physically present within the spectrographs. Without these components, the CP produces an $F/6.91$ beam at the UVVIS CP exit and an $F/6.8$ beam at the NIR CP exit.

History (more recent)



ESO call for new instruments at NTT (06/2014)

Proposal submission (02/2015)

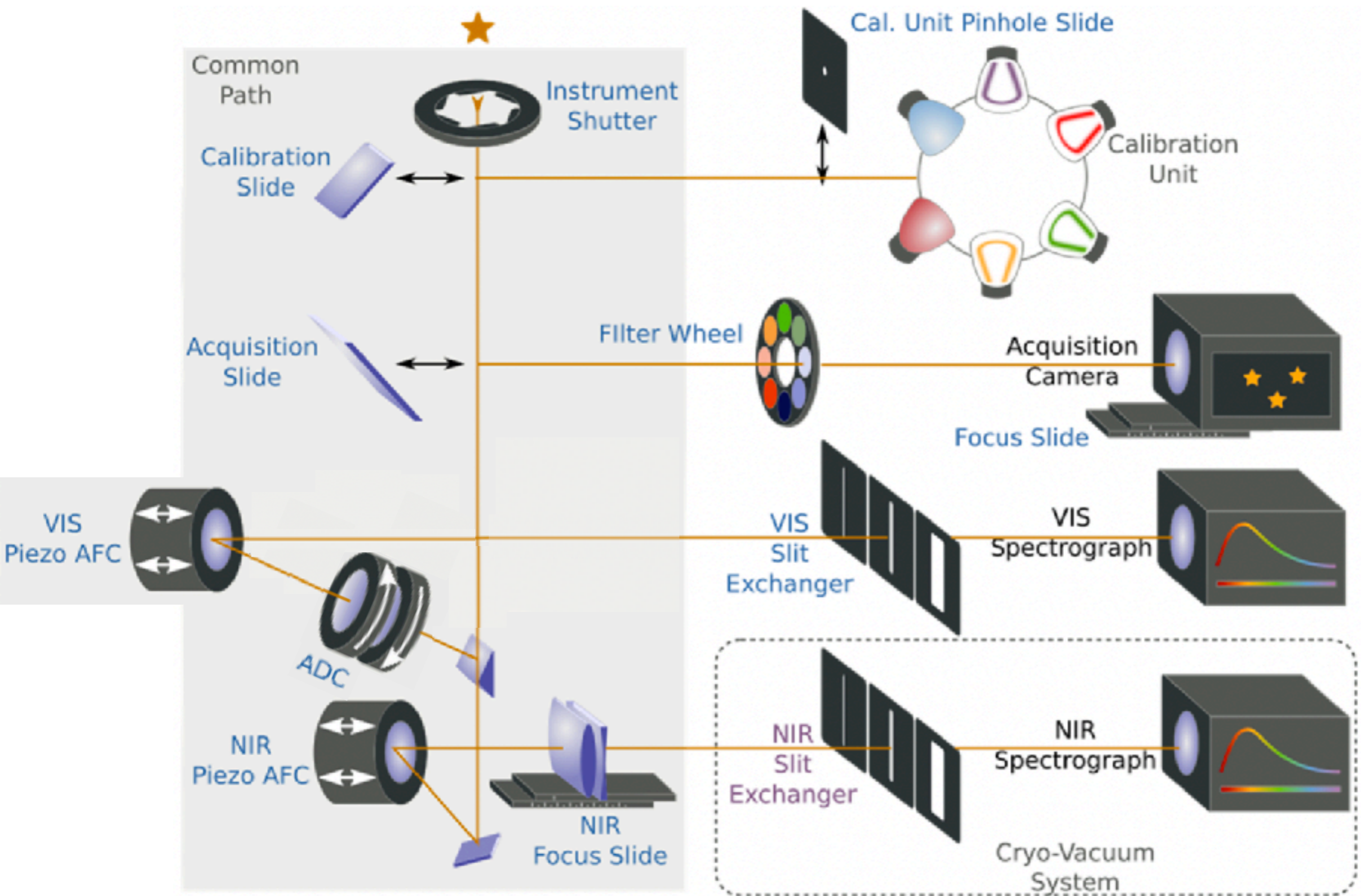
SOXS selected by ESO (05/2015) out of 19

Signed MoU INAF-ESO
Signed MoU INAF-Partners

Project Phase	Start	End	Duration
Preliminary Design	08/2016	07/2017	12 months
Final Design	08/2017	10/2018	14 months
MAIT	11/2018	11/2022	48 months+COVID
PAE	12/2022	02/2023	3 months
Commissioning & SV & PAC	03/2023	09/2023	6 months
Operations & GTO	2023	2028	



SOXS block diagram



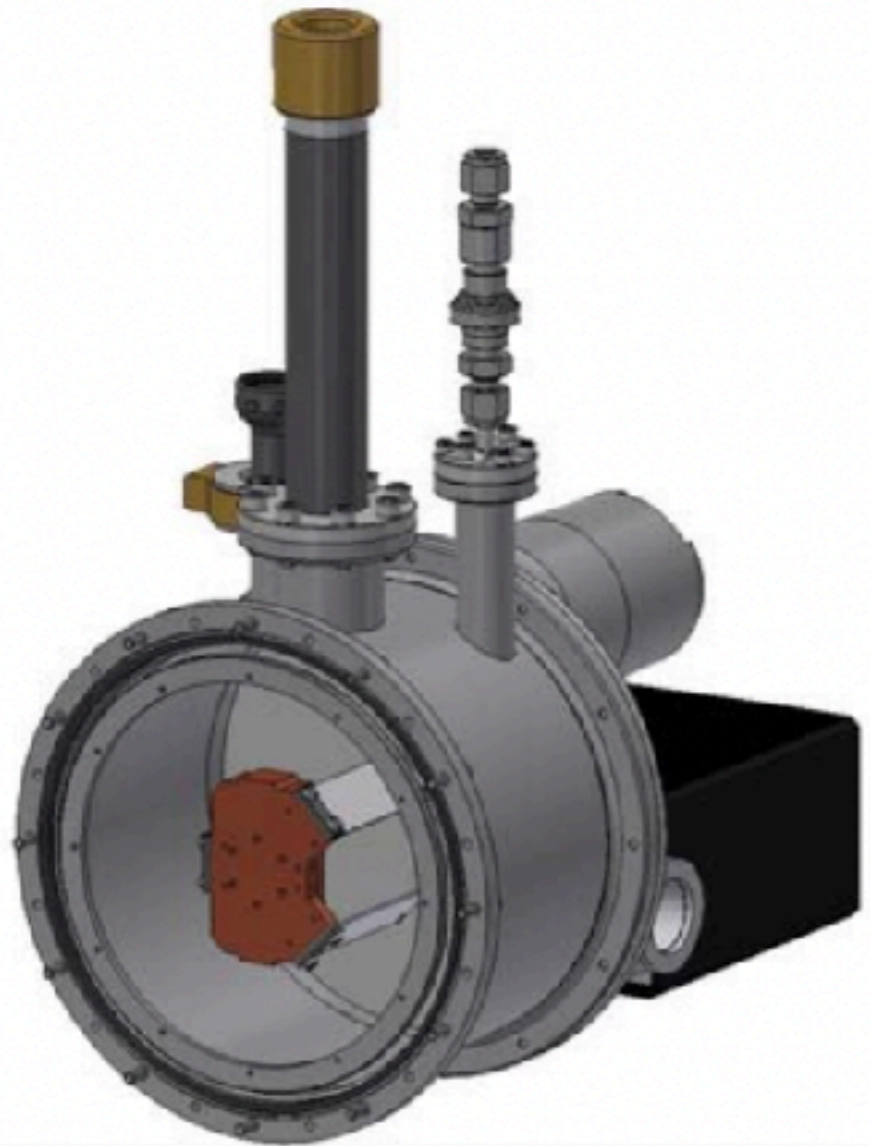
4.1 Mounting to the NTT Derotator

The following table describes what all components go on the NTT derotator, their weights, and where they are attached.

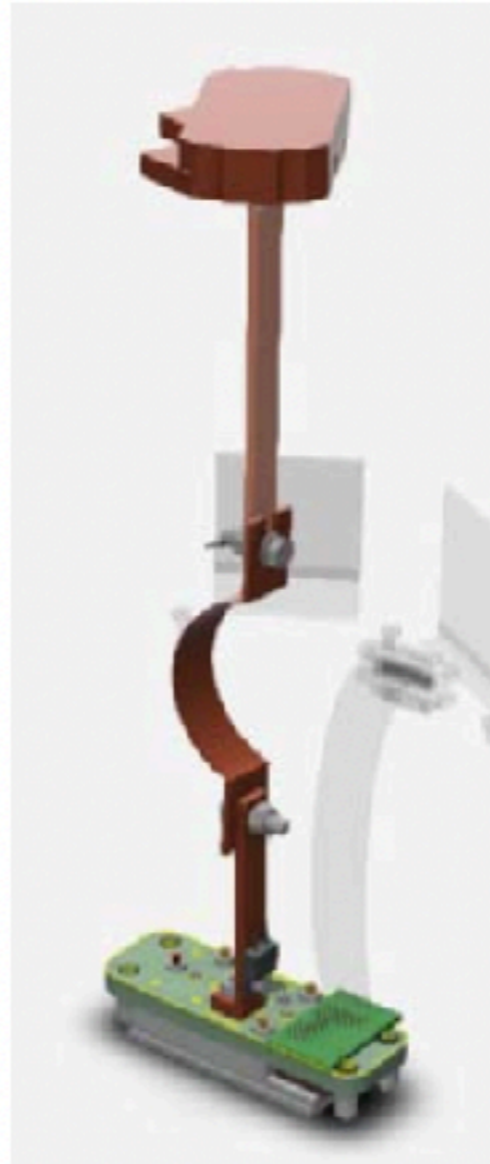
#	What	Weight measured at INAF-PD [kg]	Where is it attached
1	SOXS Flange + Support Structure	~ 355	NTT Derotator
2	Common Path	46	SOXS Flange
3	Calibration Unit	9	SOXS CP
4	Acquisition Camera	11	SOXS CP
5	UVVIS Spectrograph	140	SOXS Flange
6	NIR Spectrograph	~ 217	SOXS Flange
7	NGC - UVVIS + Support Structure	17	NTT Derotator
8	NGC - NIR + Support Structure	~ 17	NTT Derotator
9	Power Supply Support Structure	~ 6	NTT Derotator

Table 2. The table shows the weight of the components that will be mounted to the flange (highlighted in green) and onto the NTT rotator adaptor (highlighted in blue).

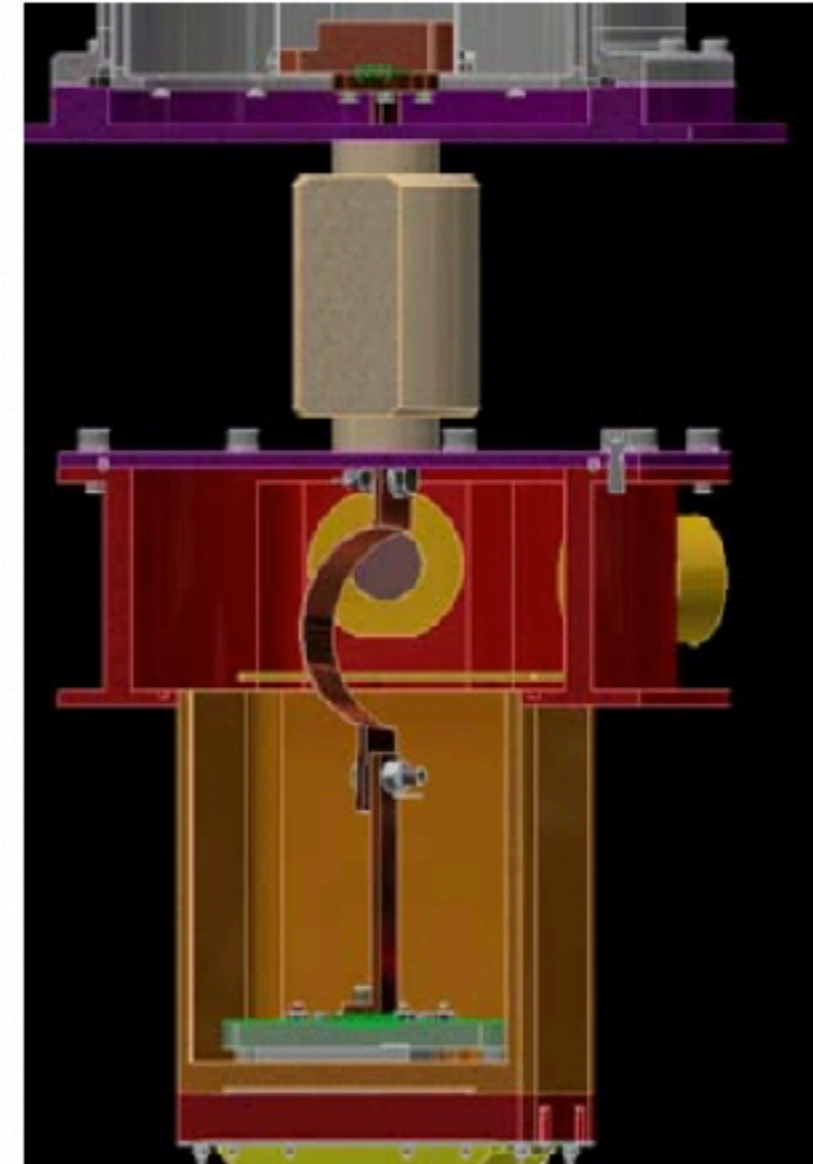
UV-VIS Cryogenic System



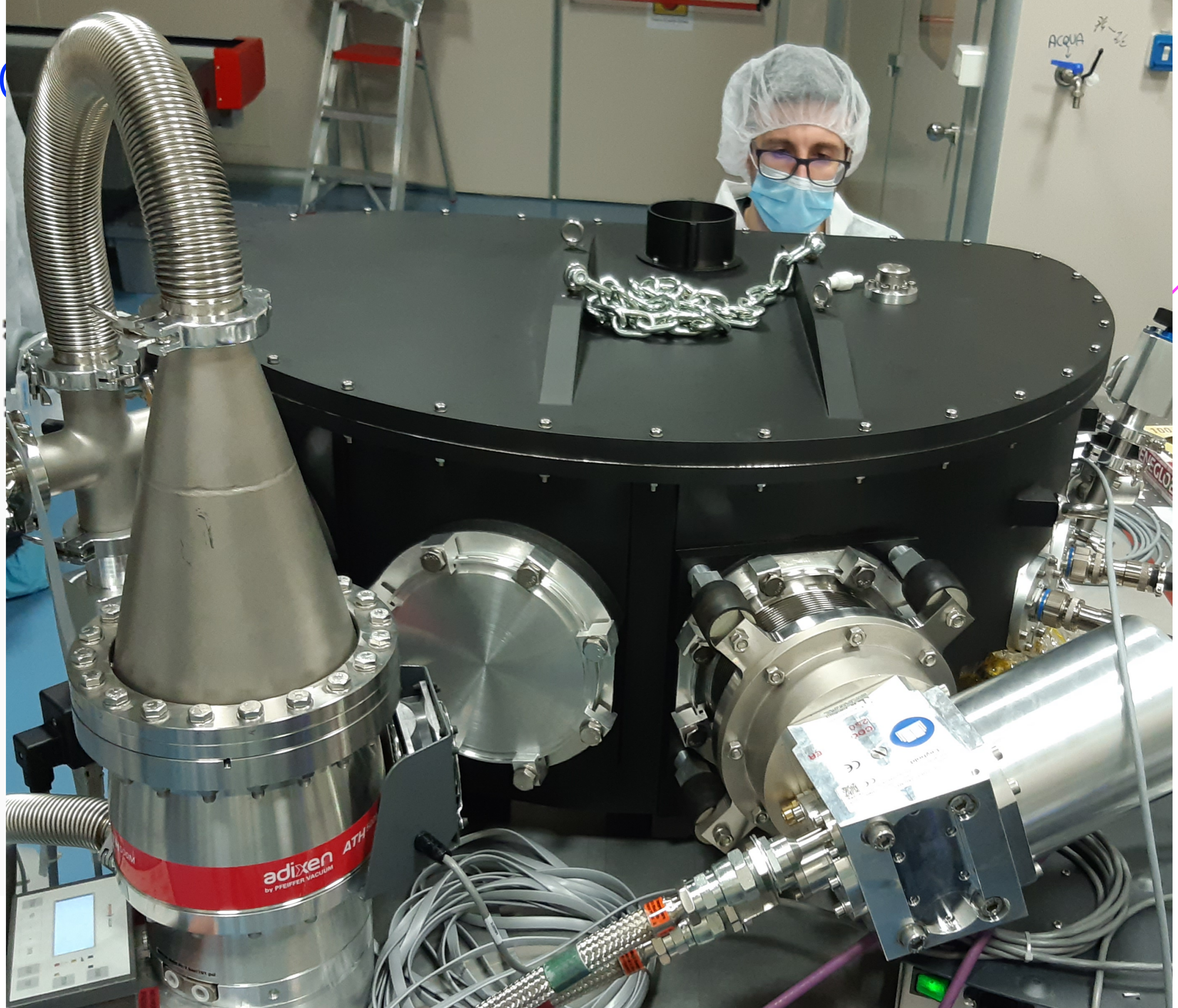
The cold finger of CFC.



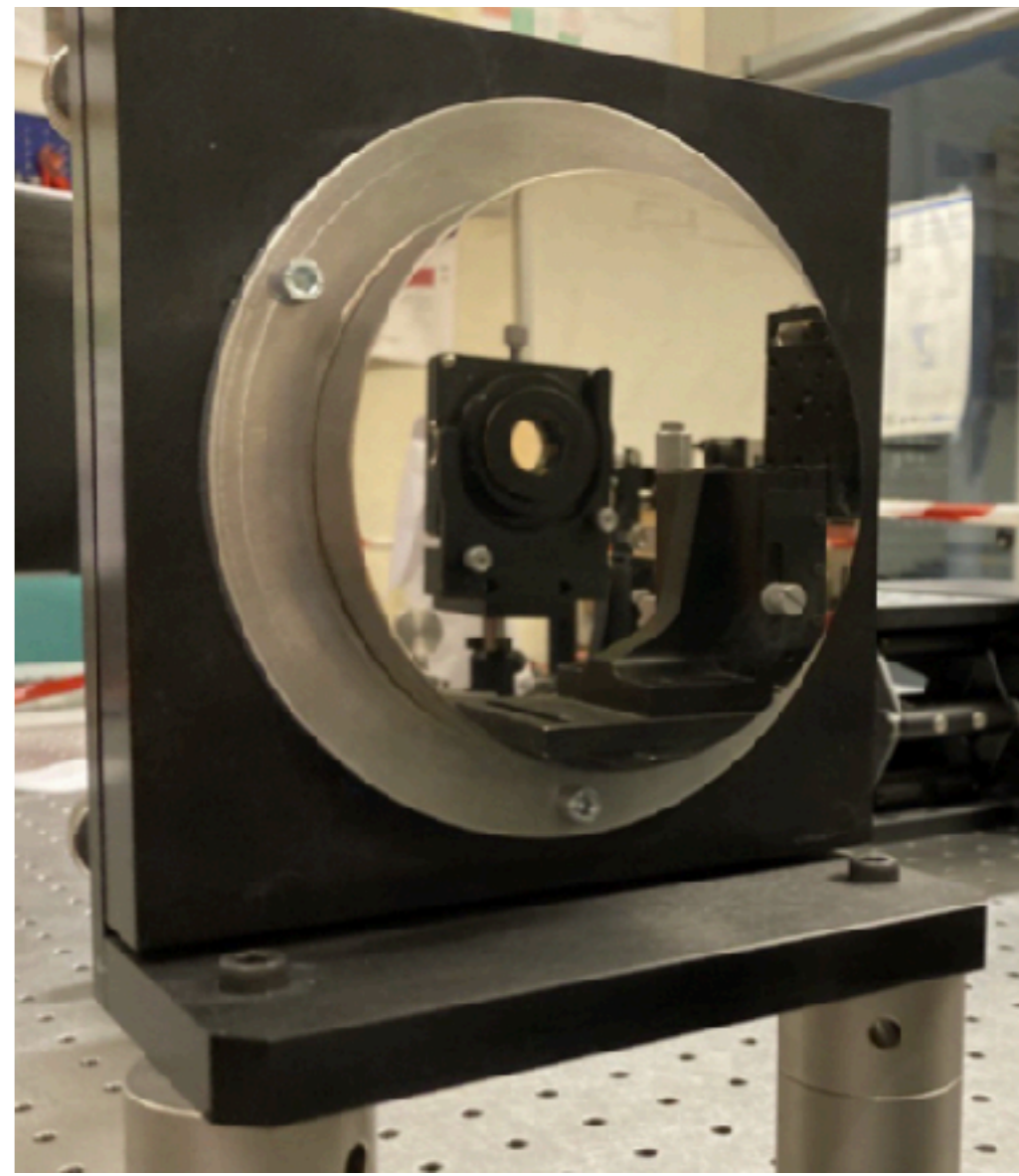
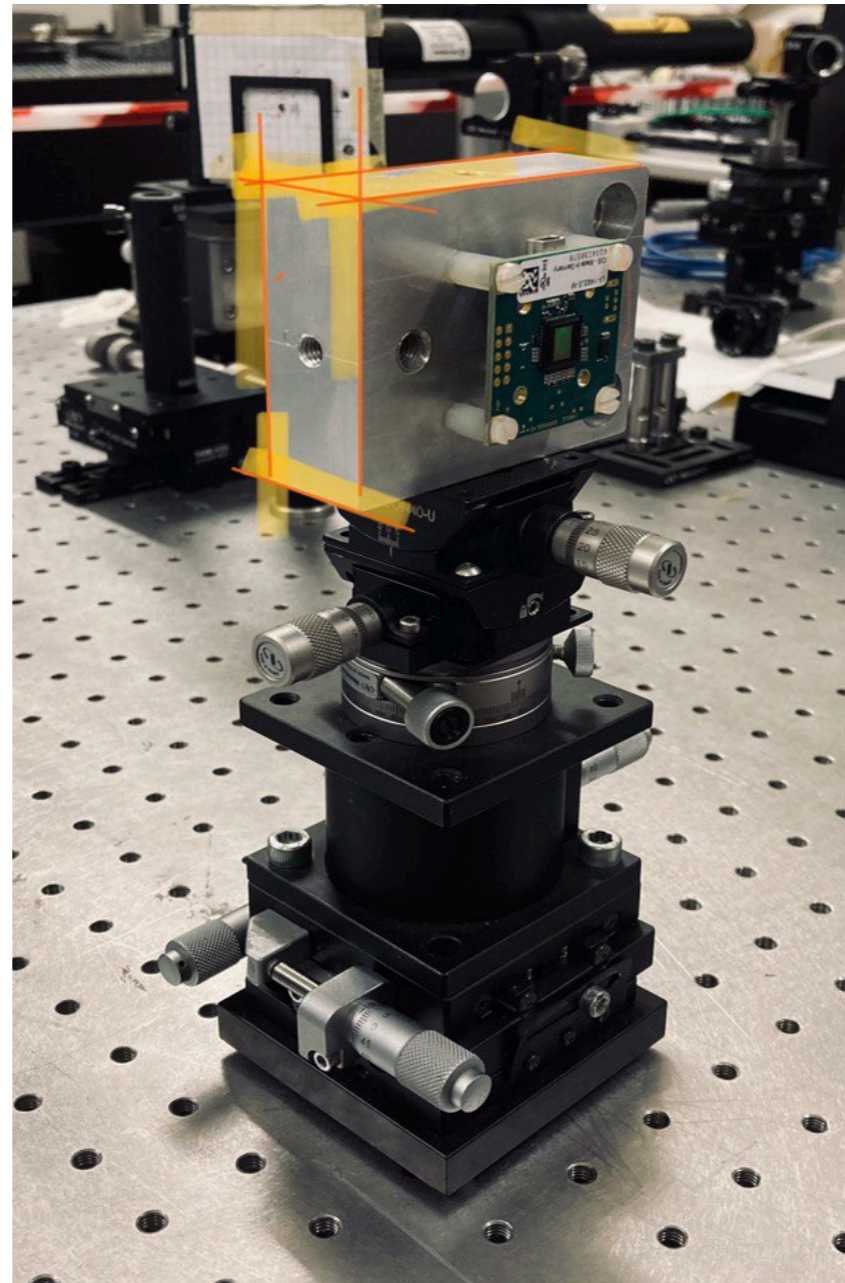
The thermal connection from CFC to CCD.



Section of the UV-VIS cryostat showing the thermal connection.



Common Path Alignment Tools



Portable Coordinate Measuring Machine

Characterized CMOS detector

10cm diameter (aluminum) mirror

What is SOXS?

- Single object spectrograph offering simultaneous spectral coverage in UV-VIS (350-850nm) and NIR (800-2000nm) with an average $R \sim 4500$ for an 1" slit.
- Can perform photometry in 360-970nm (ugVrizY) (3.5' x 3.5', 0.2"/pixel).
- Final destination: Nasmyth platform of the 3.58m ESO New Technology Telescope (NTT) at the La Silla Observatory in the Southern part of the Chilean Atacama Desert.
- Designed to observe all kinds of transients and variable sources.
- SOXS consists of 5 sub-systems - CommonPath (CP), Calibration unit (CU), Acquisition Camera (AC), UV-VIS spectrograph, and the NIR spectrograph.

Science with SOXS

- Classification

- Super Novae

- Gravitational Waves EM counter parts

- Tidal Disruptive Events & Nuclear transients

- Gamma Ray Bursts and Fast Radio Bursts

- Blazars & AGN

ExtraGalactic

- X-ray binaries & magnetars

- YSOs & Stars

- Novae, Cataclysmic Variables, & White Dwarfs

Galactic

- Astroids & Comets

Solar System

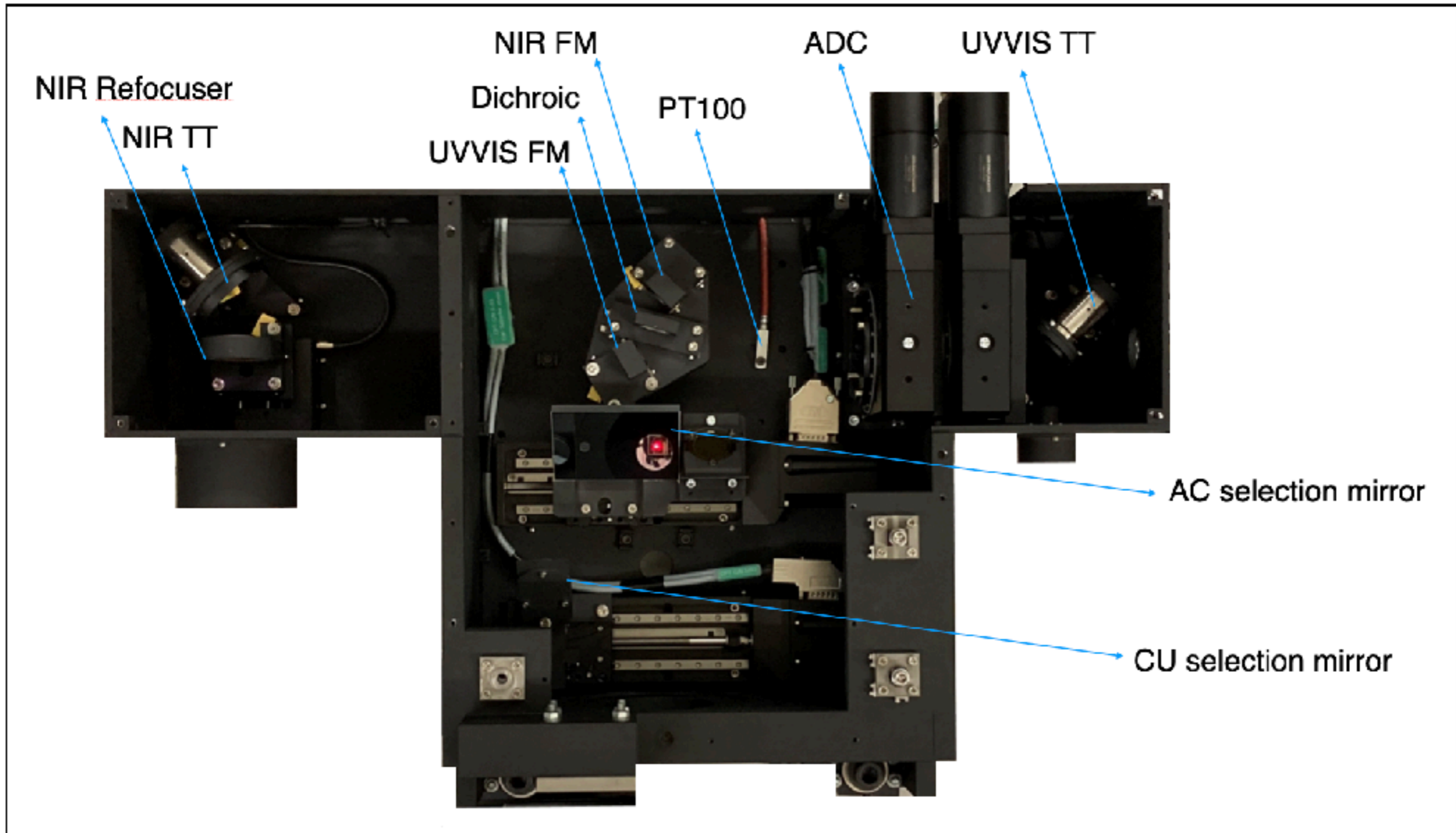
- Unknown

SOXS Consortium

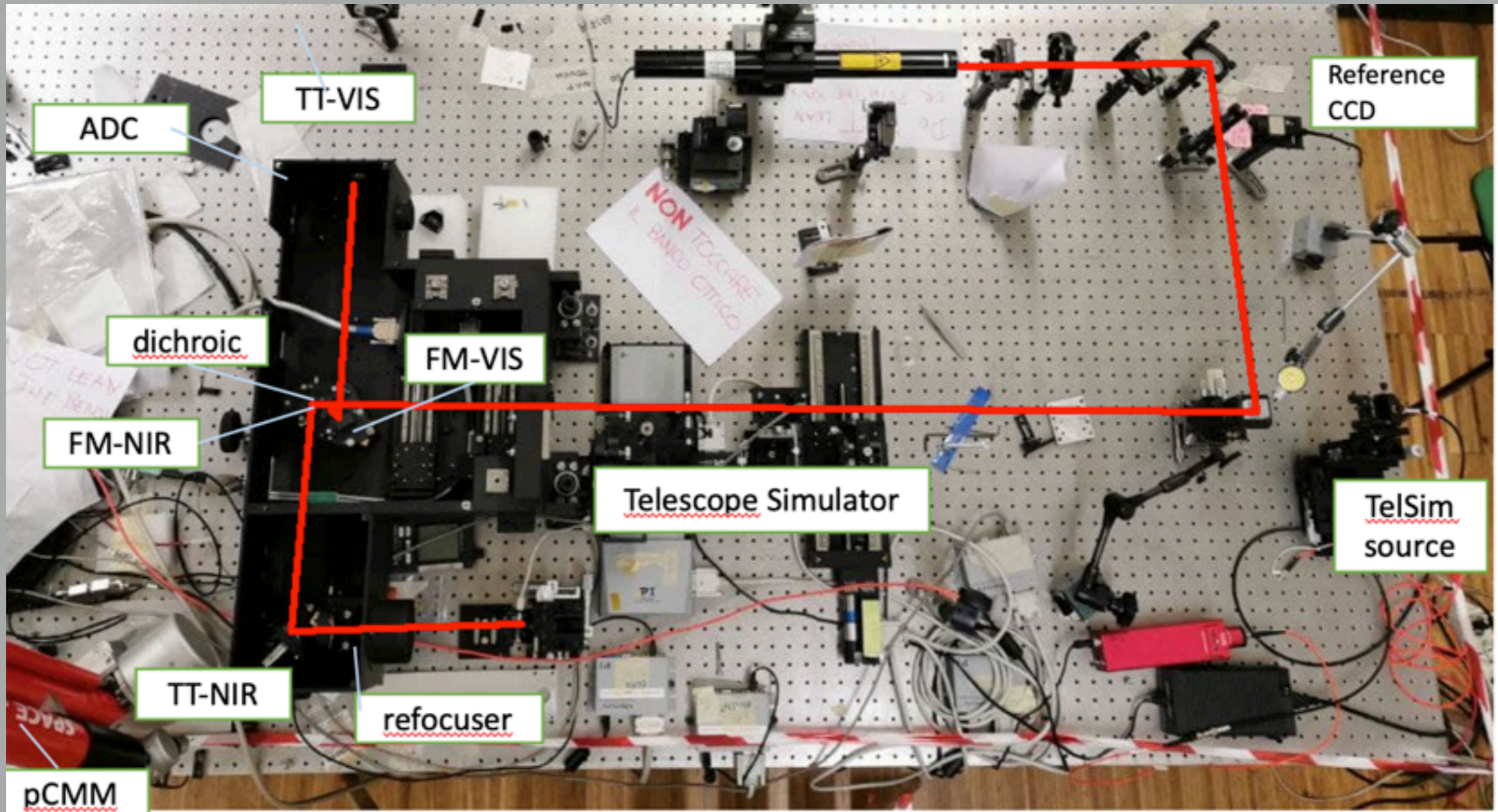
- Istituto Nazionale di AstroFisica (INAF) - Italy
- Weizmann Institute of Science - Israel
- Universidad Andres Bello & Instituto Milenio de Astrofisica, Chile
- Turku University, Finland
- Queen's University, UK
- Tel Aviv University, Israel
- Niels Bohr & Aarhus University, Denmark



Common Path

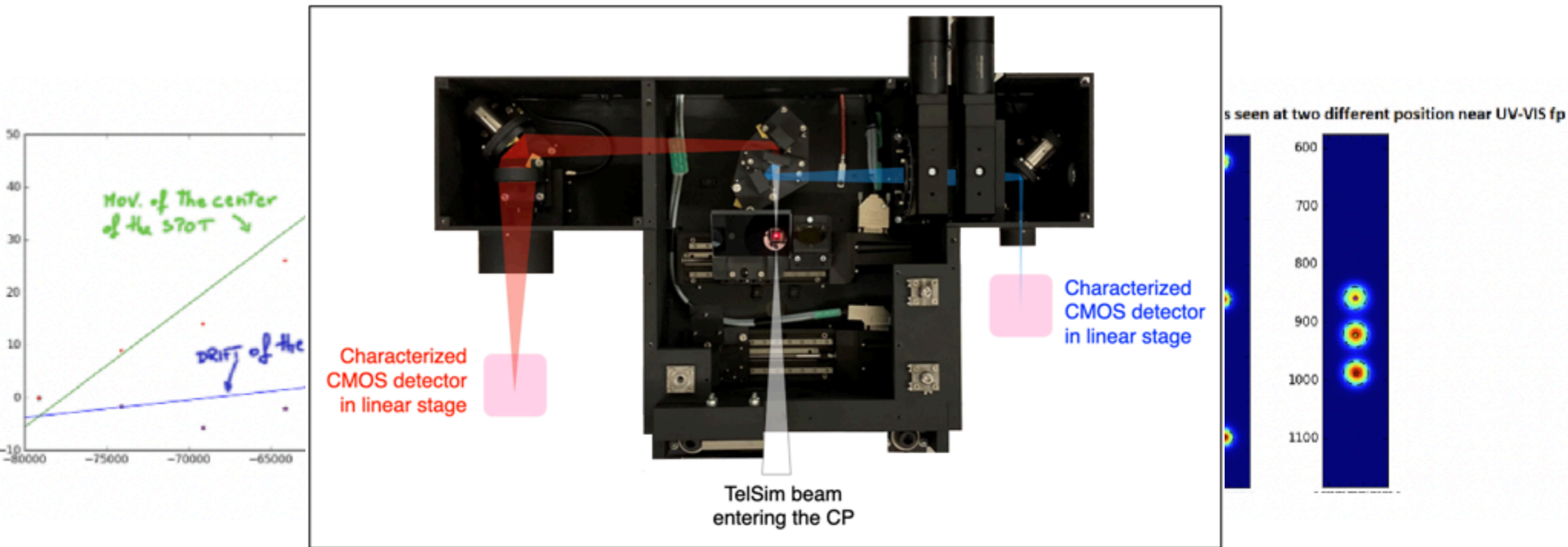


- Receives the F/11 beam from the telescope and feeds the spectrographs with an F/6.5 beam



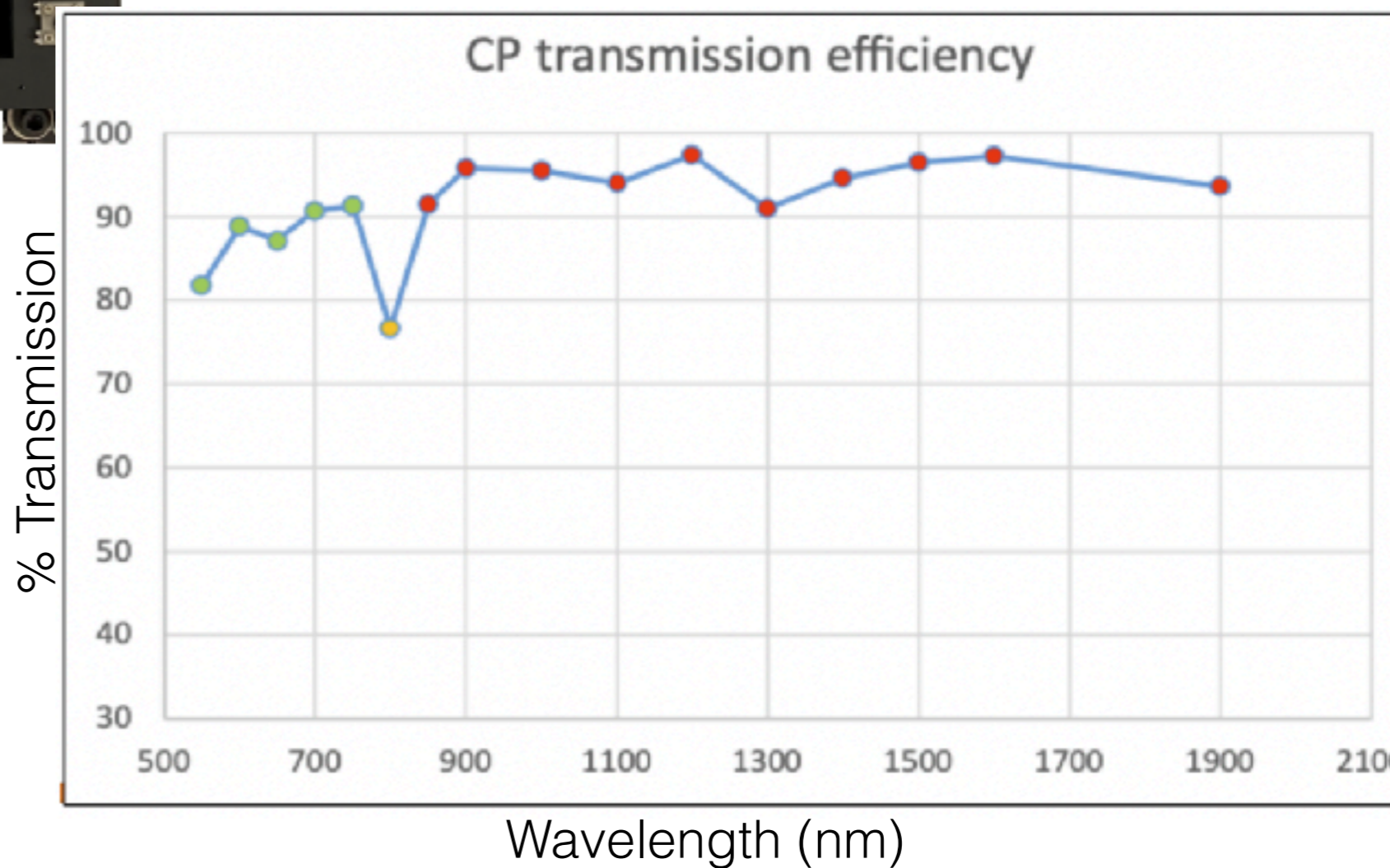
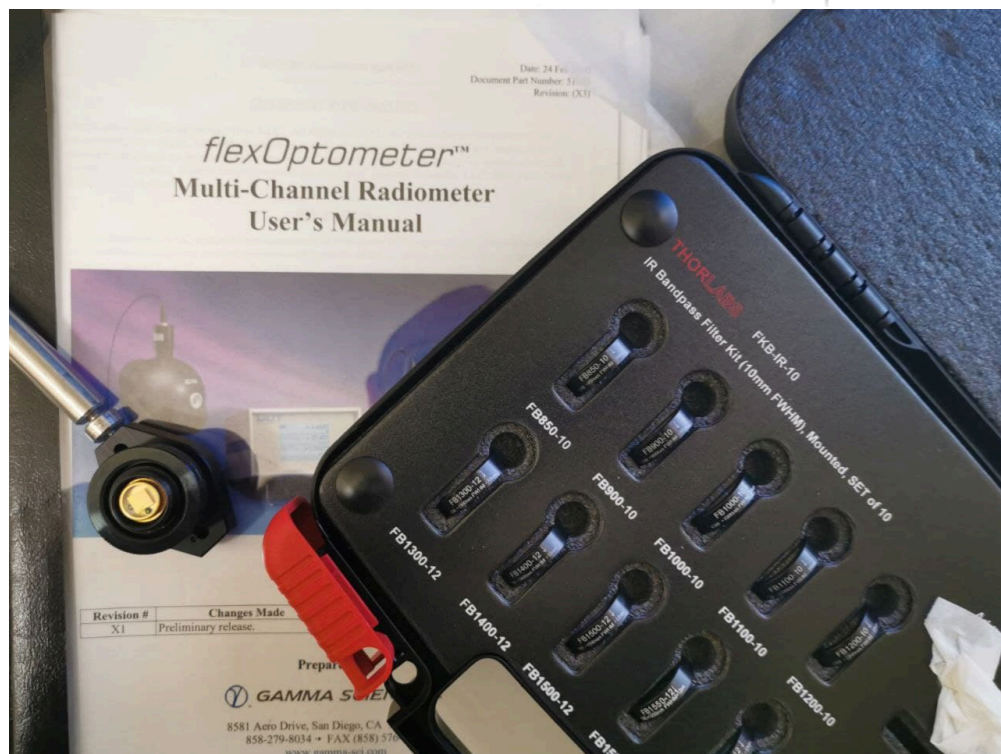
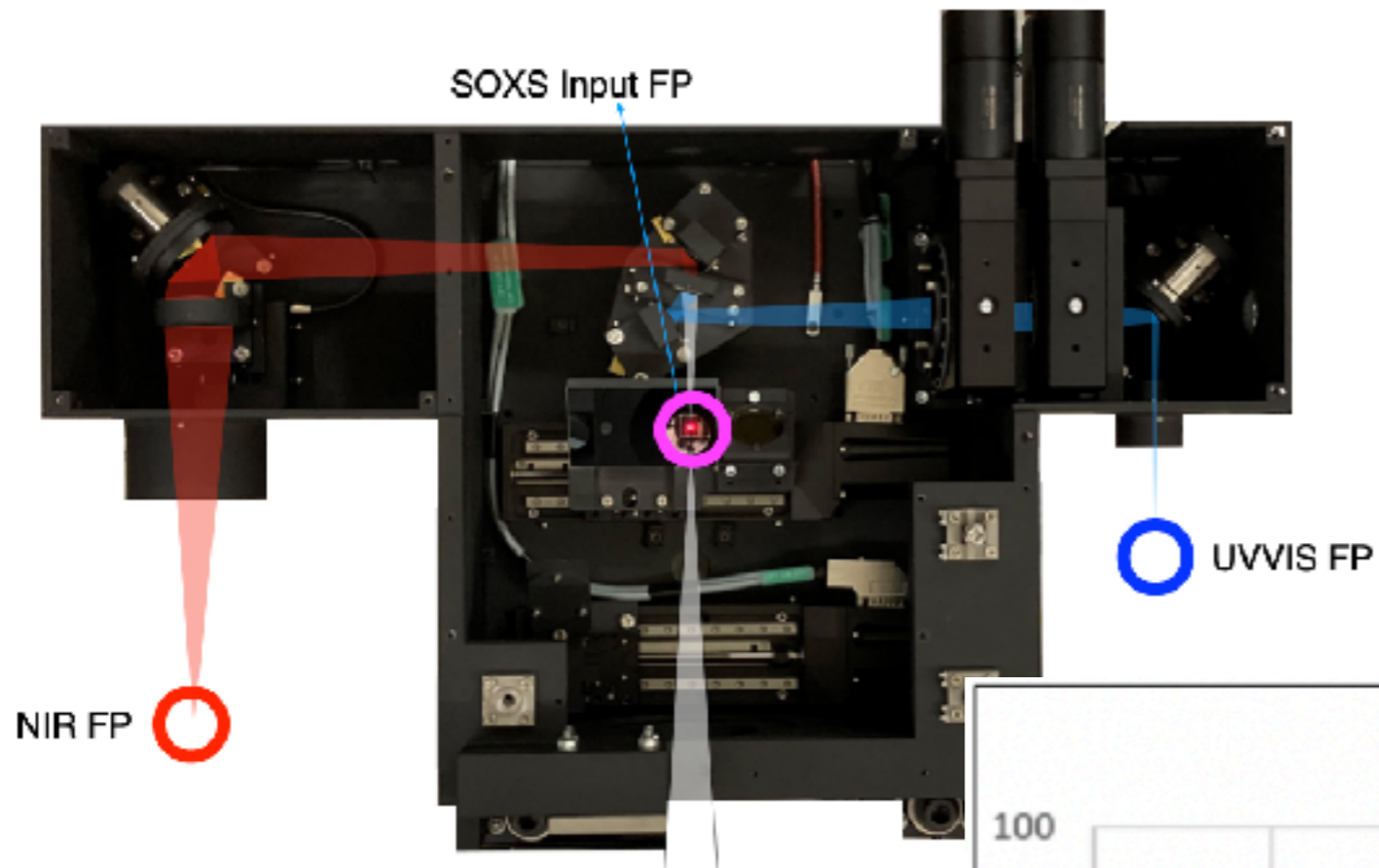
Common Path Alignment Verification

- CP exit beam position, tip-tilt, and focal number



	Decenter-X (μm)	Decenter-Y (μm)	Decenter-Z (μm)	Tilt-X (")	Tilt-Y (")	$F/\#$
UVVIS exit beam	230 ± 60	-145 ± 60	$<30 \pm 60$	-413 ± 100	37 ± 100	6.93
NIR exit beam	45 ± 60	3 ± 60	$<30 \pm 60$	-250 ± 32	-255 ± 32	6.94

Common Path Alignment Verification



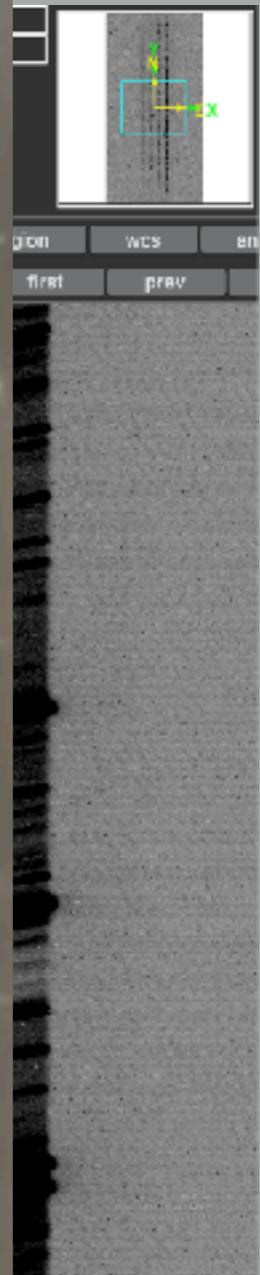
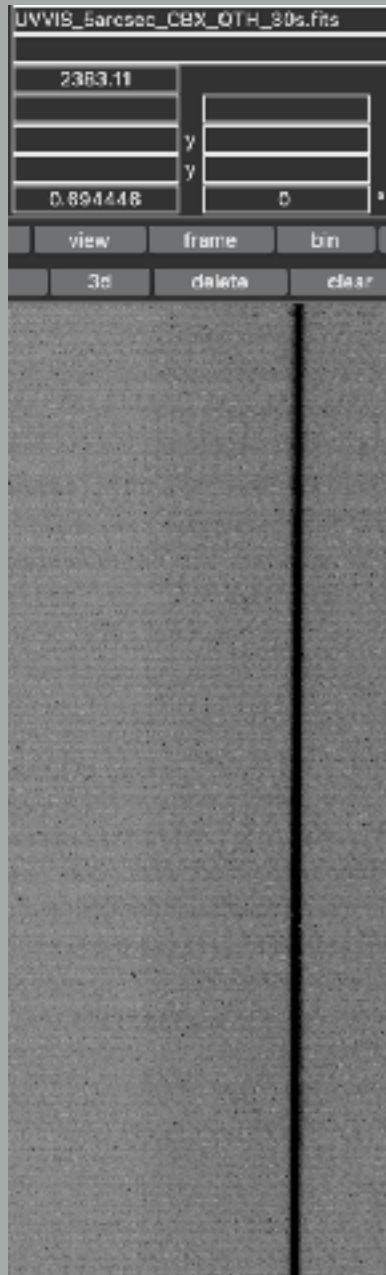
Integration and Testing



Integration and Testing



Integration and Testing



Continuum source spectra

ThAr Arc lamp spectra