Matter-Dominated Gravitation Wave Sources

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Key Points

- GW dHz astronomy will make history
- Largest number of science cases
- Transient sky
- Research needed





Seeing with GWs

- New eyes for astronomy
- Unseen objects, e.g. BBHs Abbot+16
- Hidden regions, e.g. DNSs Abbot+17
- Unique signal profile
- All directions
- ET, CE will see all BBH mergers Branchesi+23

GW:
$$h_{\rm GW} \propto \frac{1}{d} \longrightarrow N_{\rm objects} \propto d^3 \propto h_{\rm GW}^{-3}$$

EM: $F_{\rm EM} \propto \frac{1}{d^2} \longrightarrow N_{\rm objects} \propto d^3 \propto F_{\rm EM}^{-3/2}$



dHz Band: Interesting binaries

 Binaries about to do something interesting:

LIGO:
$$\tau_{\rm GW} \propto 2.9 \sec \left(\frac{M_{\rm chirp}}{M_{\odot}}\right)^{-5/3} \left(\frac{f_{\rm GW}}{100 \,{\rm Hz}}\right)^{-8/3}$$

LGWA: $\tau_{\rm GW} \propto 9.3 \,{\rm yr} \left(\frac{M_{\rm chirp}}{M_{\odot}}\right)^{-5/3} \left(\frac{f_{\rm GW}}{0.1 \,{\rm Hz}}\right)^{-8/3}$
LISA: $\tau_{\rm GW} \propto 2.0 \,{\rm Myr} \left(\frac{M_{\rm chirp}}{M_{\odot}}\right)^{-5/3} \left(\frac{f_{\rm GW}}{0.001 \,{\rm Hz}}\right)^{-8/3}$



Double White Dwarfs

- Detached population extragalactic
- Detection horizon 30-100Mpc
- About 100DWDs, few Galactic
- Interaction phase short AB+17
- Formation of interacting DWDs

$$\tau_{\rm GW} \propto 9.3 \, {\rm yr} \left(\frac{M_{\rm chirp}}{M_\odot}\right)^{-5/3} \left(\frac{f_{\rm GW}}{0.1 \, {\rm Hz}}\right)^{-8/3}$$

Ajith+24

Credit: Caltech/IPAC

$$\frac{\tau_{\rm GW,LGWA}}{\tau_{\rm GW,LISA}} \propto \frac{N_{\rm DWD,LGWA}}{N_{\rm DWD,LISA}} \propto \left(\frac{f_{\rm GW,LGWA}}{f_{\rm GW,LISA}}\right)^{-8/3} \approx 5 \cdot 10^{-6} \approx \frac{\rm several}{10^6}$$

Talk by J.Morán-Fraile today Talks by S.Toonen, T.Kupfer today



WD-NS Binaries

- About 10x fewer than DWDs Toonen+19
- All extragalactic, 10 systems
- Short-lived XRBs AB+17
- May witness formation of ULXs
- LGWA may identify transients



Credit: M.Garlick



Talk by J.Morán-Fraile today

dHz Band: Shocks

- Shocks
- GW memory
 - Changes in v, a, phi, ... permanent changes in the metric tensor

LIGO:
$$L_{\rm acc} \propto 100 \,\mathrm{km} \left(\frac{v}{10000 \,\mathrm{km/s}}\right) \left(\frac{f}{100 \,\mathrm{Hz}}\right)^{-1}$$

LGWA: $L_{\rm acc} \propto 0.14 \,R_{\odot} \left(\frac{v}{10000 \,\mathrm{km/s}}\right) \left(\frac{f}{0.1 \,\mathrm{Hz}}\right)^{-1}$
LISA: $L_{\rm acc} \propto 14 \,R_{\odot} \left(\frac{v}{10000 \,\mathrm{km/s}}\right) \left(\frac{f}{0.001 \,\mathrm{Hz}}\right)^{-1}$



DWD Explosions



• Supernovae Type la

- Details of detonations
- Other transients



Core-Collapse SNe



Credit: Choi+24

• Informative, multimessenger



(Choked) Gamma-Ray Bursts



Credit: Senno+16

• Local Universe, may be hidden



Tidal Disruption Events



Credit: Jankovič+23

• Local Universe, low-mass MBHs

Talk by Elisabeth Kepler today



He Flashes



Credit: CSIRO

- Asymmetric ignition
- < 1min timescale
 Paxton+11
- 1 per decade Ajith+24



The dHz Band

$$f_{\rm GW} \propto 4 \cdot 10^{-5} \,\mathrm{Hz} \, \sqrt{\frac{\langle \rho_{\rm matter} \rangle}{g/\mathrm{cm}^3}}$$

LIGO: $\langle \rho_{\text{matter,LIGO}} \rangle \approx 10^{13} \,\text{g/cm}^3$ LGWA: $\langle \rho_{\text{matter,LGWA}} \rangle \approx 10^7 \,\text{g/cm}^3$ LISA: $\langle \rho_{\text{matter,LISA}} \rangle \approx 10^3 \,\text{g/cm}^3$

- Other types of sources:
 - Stellar cores?
 - Accretion discs?
 - Stellar jets?



Combined Signal

- Dominated by extragalactic binaries (BBHs, DWDs, WDNSs)
- Few very bright sources
- Diverse burst signals expected
- Importance of EM



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Possible Discussion Points

- Science value relative to kHz/mHz GWs
- Need to identify dominant signals
- Landscape is new new discoveries possible
- Possible that 90s-00s science could not appreciate dHz astronomy
- Synergies with neutrino astronomy
- Synergies with EM
- Many of the white paper topics may be developed into separate papers
- Science with Soundcheck
- Other science cases

