

Finanziato dall'Unione europea NextGenerationEU







Make Cosmological Analysis Accessible: A Unified Pipeline for IV Spectroscopic Survey BAO Analysis E. Sarpa, M.Viel

Spoke 3 General Meeting, Elba 5-9 / 05, 2024

ICSC Italian Research Center on High-Performance Computing, Big Data and Quantum Computing

Missione 4 • Istruzione e Ricerca









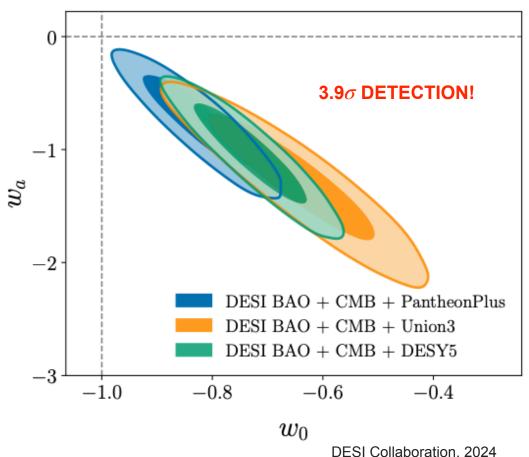
Cosmological Revolution: Rethinking Gravity

Galaxy Surveys

- IV Generation Spectroscopic surveys
- 30 milion galaxies, Full sky, 10Gyr

DESI Results

- Evidence of Dinamical Dark Energy
- Need for New Gravitational model



DESI YEAR 1 DARK ENERGY CONSTRAINTS

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The Importance of Reproducible Results

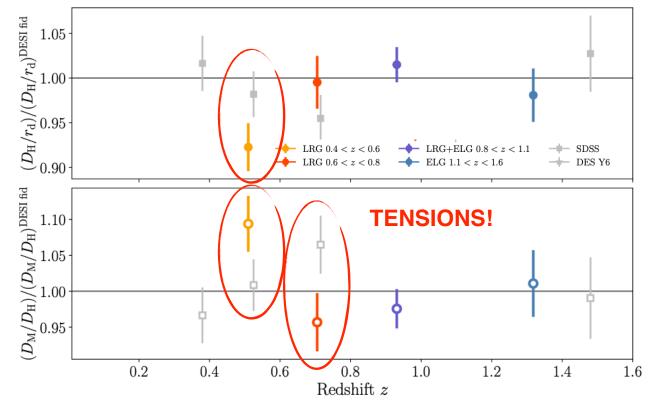
DESI RESULTS

- Comparison with SDSS results
- Tensions on cosmological distances

ACCESSIBLE SCIENCE

- Results should be reproducible
- Analysis piplines should be public

STATE OF THE ART BAO SCALE MEASUREMENTS



DESI Collaboration. 2024

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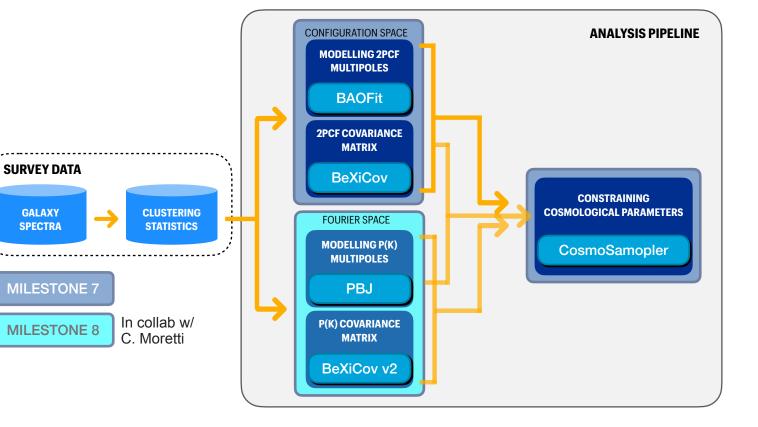
A Comprehensive Analysis Pipline: From measurements to cosmological constraints

Features

- Models for the non-linear two-point clustering signal of mass tracers
- Semi-analytical covariance matrixes for the two-point statistics
- Cosmological inference

Accessibility

- Publicly available
- Accompained by documentation and examples
- Results described in Euclid Consortium papers











BAOFit: modelling the 2PCF multipoles

Models

- Physical description of the signal: linar prediction + non-linear corrections
- Parameterinc modelling of systematics

Sampler

- Parallelized sampling of the Likelihood: emcee
- CPU time: 250k steps, 1 cpuh

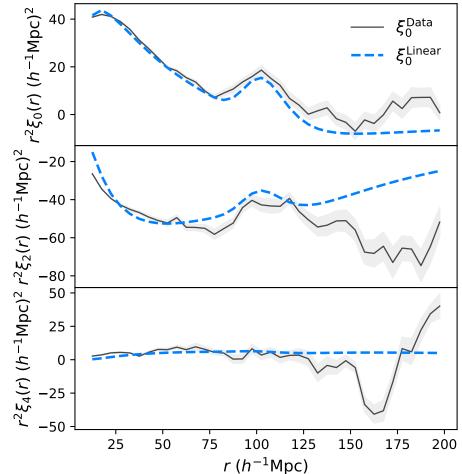
KPI

- Pulication of the code: <u>https://gitlab.com/esarpa1/BAOFit</u>
- Used in Euclid BAO reconstruction forecast paper
 & Observatyional systematics paper

Next steps

- Emulators for the 2PCF multipoles estimated gain: 1/20 CPU time (ongoing within Euclid, BORA emulator by M. Bonici)
- Nested samplers

DATA VS LINEAR MODEL











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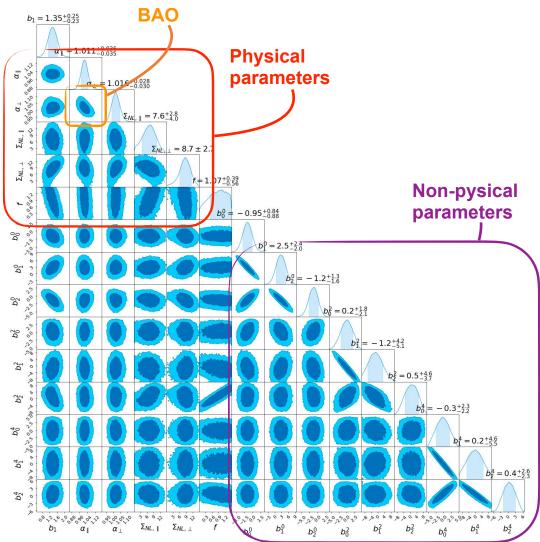
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BAOXiCov: 2PCF multipoles covariance matrix

Models

- Gaussian covariance matrix
- No Window
- Paramtric description of measured clustering signal estimated via iterative fit

Performances

• 0.1 CPUh for 2 iterations of 3 multipoles

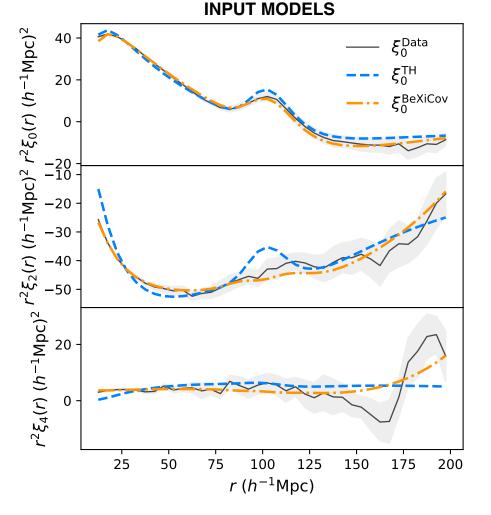
KPI

- Pulication of the code: <u>https://gitlab.com/esarpa1/BeXiCov</u>
- Used in Euclid BAO reconstruction forecast paper

Next steps

- Emulators for the 2PCF multipoles within iterative fit
- Inclusion of the window

(ongoing merging with WinCov code by A. Veropalumbo)











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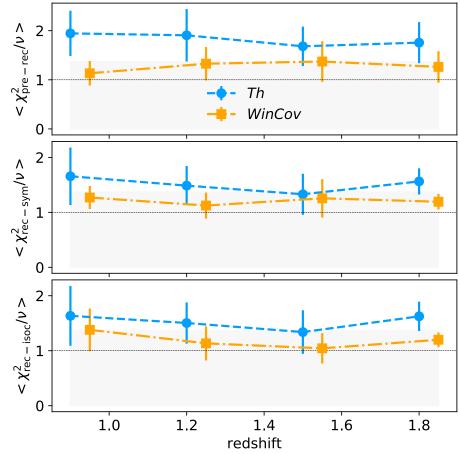
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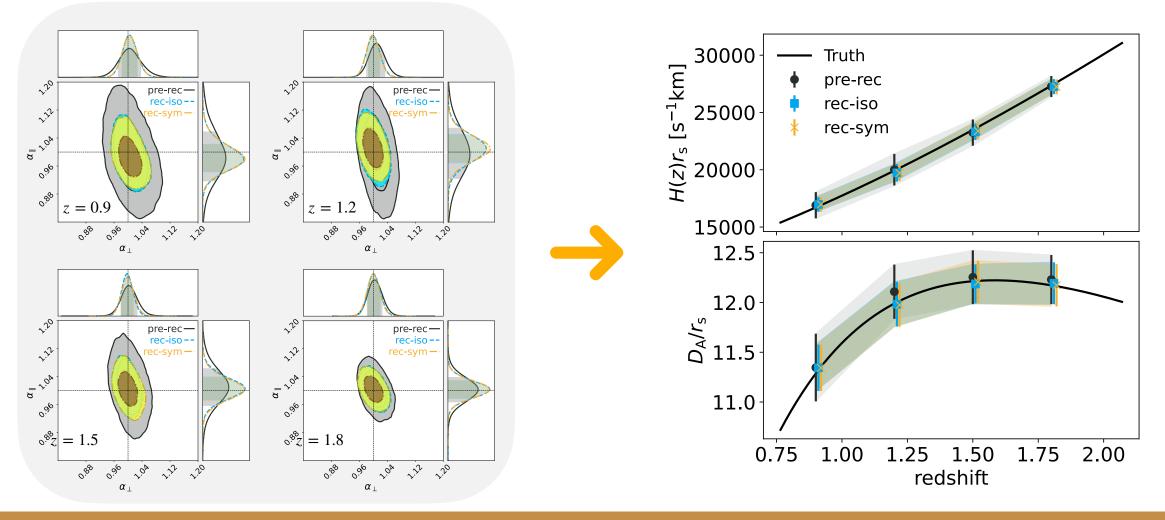








BAOFit + BAOXiCov: lightcone results



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CosmoSampler: from BAO constraints to Dark Energy

Models

- Cosmologies: Flat LCDM, waw0 LCDM (dynamical DE)
- Possibility to include CMB constraints (Planck)

Sampler

- Parallelized sampling of the Likelihood: emcee
- CPU time: 500k steps, 1 cpuh

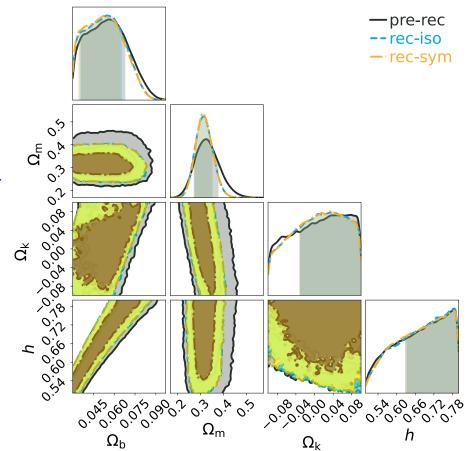
KPI

- Pulication of the code: <u>https://gitlab.com/esarpa1/cosmosampler</u>
- Used in Euclid BAO reconstruction forecast paper
 & Observatyional systematics paper

Next steps

- Beyond BAO analysis
- Inclusion of neutrino cosmologies





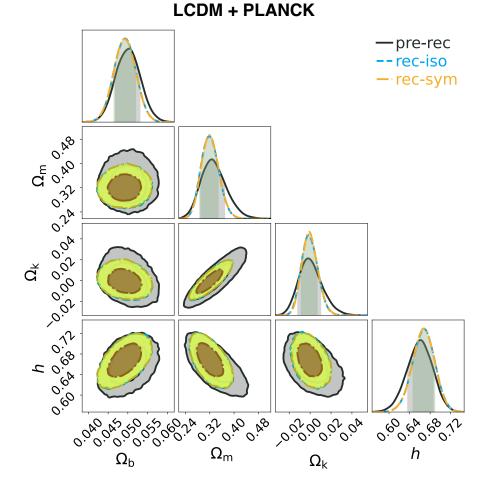


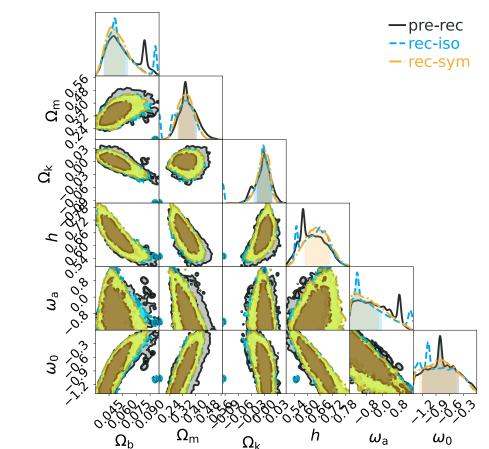






CosmoSampler: Exploring different scenarios





W0WaCDM + PLANCK

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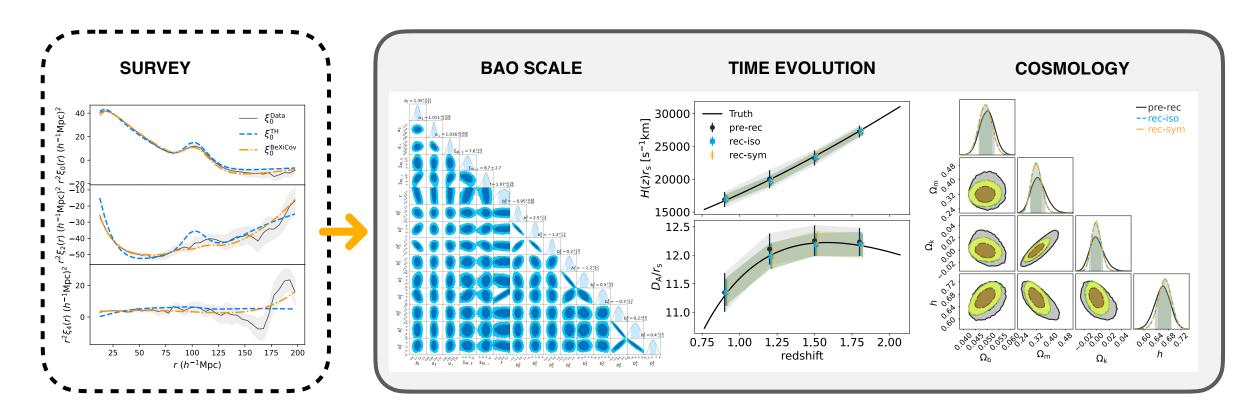








Summary: Full Pipline Run



READY FOR EUCLID & DESI DATA!









Next Steps:

Improving performances with Emulators for the clustering signal

Statuts: ongoing within Eulcid Consortium Milestone: 8 KPI: publication of Euclid BAO paper

Implementing analysis in Fourier-space

Statuts: ongoing in collaboration w/ C. Moretti (PBJ) Milestone: 8 KPI: inclusion in PBJ code

• Bayond BAO analysis: BAO + full-shape joint fit

Statuts: ongoing in collaboration w/ C. Moretti (PBJ) Milestone: 9 KPI: inclusion in PBJ code

