

Finanziato dall'Unione europea NextGenerationEU







Accelerated radio astronomy with RICK De Rubeis Emanuele, Claudio Gheller, Giovanni Lacopo, Giuliano Taffoni, Luca Tornatore

Spoke 3 General Meeting, Elba 5-9 / 05, 2024

ICSC Italian Research Center on High-Performance Computing, Big Data and Quantum Computing

Missione 4 • Istruzione e Ricerca







SKA1-low

SKA1-mid



Scientific Rationale

Why HPC for radio astronomy?

Current and upcoming radio-interferometers are expected to produce **volumes of data of increasing size**. This means that current **state-of-the-art software needs to be re-designed** to handle such unprecedented **data challenge**.

Imaging in radio astronomy represents one of the most **computational demanding** steps of the processing pipeline, both in terms of memory request and in terms of computing time.

For example, this over 7 billions pixels image can take ~250,000 core hours!



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RICK (Radio Imaging Code Kernels) is a code that addresses the *w*-stacking algorithm (Offringa+14) for imaging, combining parallel and accelerated solutions.

- C, C++, CUDA, HIP
- MPI & OpenMP for CPU parallelization
- The code is now capable of running full on GPUs, using CUDA, HIP or OpenMP for offloading
- An optimized version of the reduce has been developed on both CPU (combining MPI+OpenMP) and GPU (using NCCL or RCCL, for Nvidia and AMD respectively); the FFT is done through the cuFFTMp library for Nvidia
- Currently under benchmarking on **Leonardo** (CINECA, No.4 Top500 June 23)











Why do we need multiple GPU?

Modern and future radio interferometers will produce a huge amount of data, that hardly fit into the memory of a single GPU (not even a single node)



The solution is to distribute the problem among multiple GPUs and multiple nodes









NVIDIA Collective Communication Library (NCCL)

NCCL is a library of multi-GPU collective communication used to support the *Reduce* operation.

- Provides fast collectives over multiple GPUs both intra- and inter-node.
- Supports a variety of **interconnection technologies** (e.g. NVLink, PCIe).
- NCCL closely follows the popular collectives API defined by **MPI**, so can be very "natural" to use.











NVIDIA Collective Communication Library (NCCL)

NCCL implements the *Reduce* operation as an intra-node **ring**, and an inter-node **ring**, when GPUs assigned to the main tasks communicate with RDMA with GPUs in different nodes **without passing through the CPUs**.

ncclUniqueId id; ncclComm_t comm;



PCIe / QPI : 1 unidirectional ring

DGX-1: 4 unidirectional rings

ncclCommInitRank(&comm, size, id, rank);

ncclReduce(send_g, recv_g, size_g, ncclDouble, ncclSum, taget_rank, comm, stream)









NVIDIA Collective Communication Library (NCCL)

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ncclUniqueId id; ncclComm t comm; The requirement of a dedicated stream for the *Reduce* comes from the presence of asynchronous memory copies that collided with the ones within a previous function call

ncclCommInitRank(&comm, size, id, rank);

ncclReduce(send_g, recv_g, size_g, ncclDouble, ncclSum, taget_rank, comm, stream











cuFFTMp

Fast Fourier Transform (FFT) is a **critical operation** in radio astonomy, because it determines the relationship between the "observed" and the "desired" data (the final image).

For the FFT step, RICK now implements the cuFFTMp library, that enables the **distribution of the FFT problem using NVSHMEM**.

NVSHMEM uses asynchronous, GPU-initiated, data transfers, **avoiding synchronization overheads** between the CPU and the GPU.



NVSHMEM











cuFFTMp

Data are distributed among multiple GPUs and inverse-transformed.











cuFFTMp

- We may need to to this FFT process even 100s-1000s times, and each time we need to create and destroy a *descriptor*, which is the ad-hoc data structure used by the cuFFTMp library for the FFT.
- This was critical for the performance, but we overcame this problem using CUDA kernels to write the *to-be-transformed* data for each loop, and then writing them directly inside the descriptor.

- The joint usage of NVSHMEM and NCCL can cause severe runtime errors during the FFT.
- There is the possibility to switch off NCCL support for NVSHMEM at runtime by setting NVSHMEM_DISABLE_NCCL=1.









Accomplished Work, Results

Currently, we are testing RICK on Leonardo (CINECA) using NVIDIA HPC-SDK (v 23.11). The testing dataset are **real, LOFAR-VLBI data**, the closest facility to SKA in terms of overall data size.

Comparing the code on GPUs, with respect to the one on CPUs, we obtained:

- for a small input dataset (4GB) we reached a speed-up factor up to **~x27** for both Reduce and FFT
- for a large input dataset (530GB) we reached a speed-up factor of **~x175** for the Reduce and **~x32** for the FFT









Accomplished Work, Results



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Timescale, Milestones and KPIs

| KPI | Target | Deadline |
|-----------------------|-------------------------------------------------------------|---------------------|
| Release of the v2.0 | Release of RICK v2.0 | December 31st, 2023 |
| | New Reduce and FFT on GPUs available, code fully on GPUs | December 31st, 2023 |
| Paper to be submitted | Paper on the v2.0 release of the code | M8 |
| Release of the v.2.1 | Weighting and uv-tapering, parallel and accelerated version | M9 |
| | Optimization of memory occupation on GPUs | M9 |









Next Steps and Expected Results

- With RICK we test the benefits that the exploitation of HPC resources can bring to future generation of radio interferometers, in our case for imaging data
- Preliminary results suggest that the improvement could possibly be huge, meaning that this can represent a promising approach to the SKA (and precursors) data volumes
- The code is now capable of working fully on NVIDIA GPUs with CUDA.
 Full AMD GPUs support is not yet available because of the lack of a proprietary distributed library for FFT (such as cuFFTMp for NVIDIA)
- The large data size requires distributing the workload among multiple nodes: in return, we obtain that the **comunication** becomes our main bottleneck. **The key is the right choice of computational resources**.