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ICSC Centro Nazionale di Ricerca in HPC, Big Data and Quantum Computing

Retrieval of exoplanetary atmosheres with GUIBRUSH®: status and prospects

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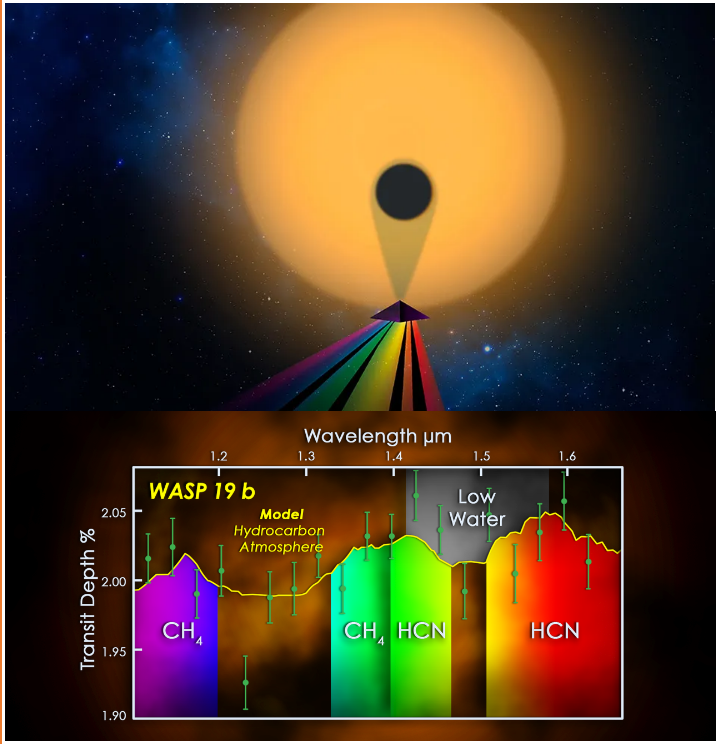
Scientific Rationale

EXOPLANETARY ATMOSPHERES

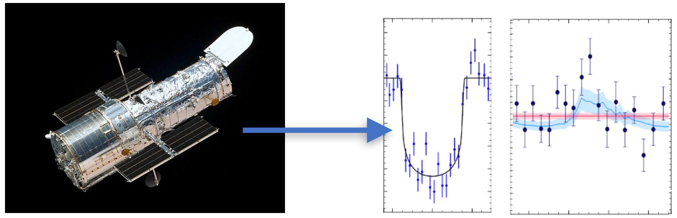
- Encoded within a planet spectrum there is information about the *formation and migration history*.

Transiting planets represent a gold booty to perform atmospheric studies.



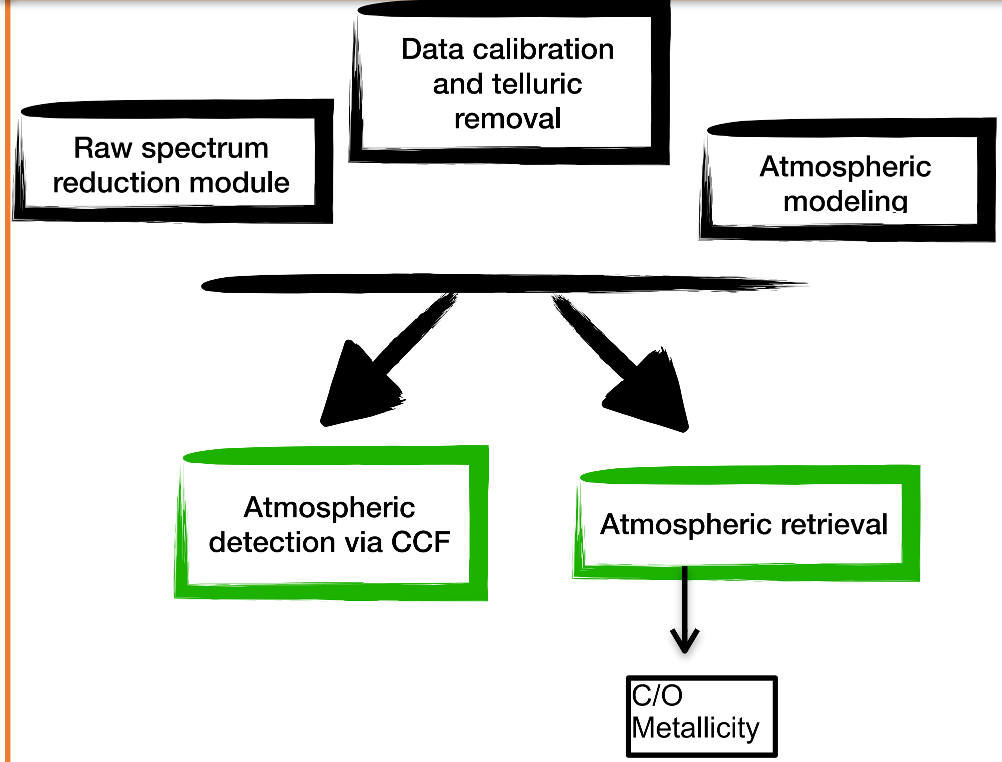
How do we probe exoplanetary atmospheres?

Space-borne Low-Resolution Spectroscopy



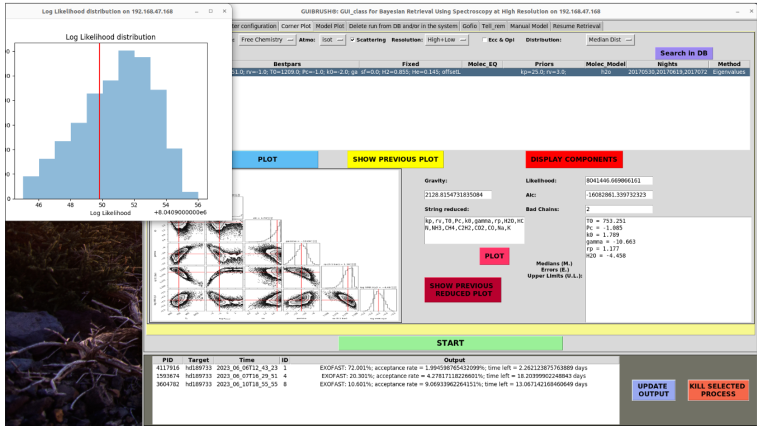
Ground-based High-Resolution Spectroscopy (R>20 000)

GUIBRUSH®:
Is a user friendly workspace to study and characterize exoplanetary atmospheres



Technical Objectives, Methodologies and Solutions

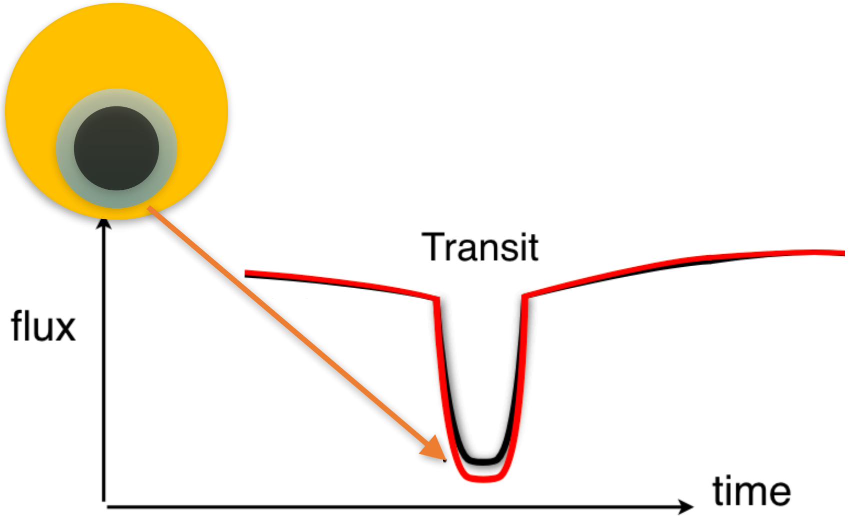
- **The Code: GUIBRUSH®** is coded in Python > 3.8 and makes use of the Bayesian differential evolution Markov chain Monte Carlo (DE-MÇMC) technique to explore the parameter space and derive the posterior distributions of the free parameters such as the abundances of the probed chemical species.



- GUIBRUSH® is currently run on the HPE Proliant DL560 Gen10 Server at INAF-Osservatorio Astrofisico di Torino, which was purchased with the PRIN-INAF 2019 project "HOT-ATMOS" (PI: A. S. Bonomo) and currently has 2 processors, 48 2.3Ghz cores and 256GB R AM;
- We are buying another server with CPU + GPU Nvidia (16 k€) thanks to an INAF minigrant (PI: Giacobbe) + overhead CN-HPC Spoke 3 + overhead PRIN MUR 2022 ESPLORA (PI: A. S. Bonomo).
  - The procurement procedure was delayed by more than 5 months due to administrative problems: the procedure was started in October 2023;
  - This server is needed to test the porting of the radiative transport code to GPU;

Technical Objectives, Methodologies and Solutions

- The main <u>bottleneck</u> is due to the slowness of the radiative transfer code, which takes ~10 s to produce a single atmospheric model at each step of each DE-MCMC chain to be compared with the observed spectrum and thus compute the likelihood function. The currently employed radiative transfer code is the publicly available, open-source tool **petitRADTRANS** (Mollière 2019).



$$\text{Transit depth} = \frac{\Delta F_\lambda}{F_\lambda} = \frac{F_\lambda^{\text{out}} - F_\lambda^{\text{in}}}{F_\lambda^{\text{out}}}$$

RADIATIVE TRANSFER: $I_\lambda(\tau) = I_0 e^{-\tau_\lambda}$

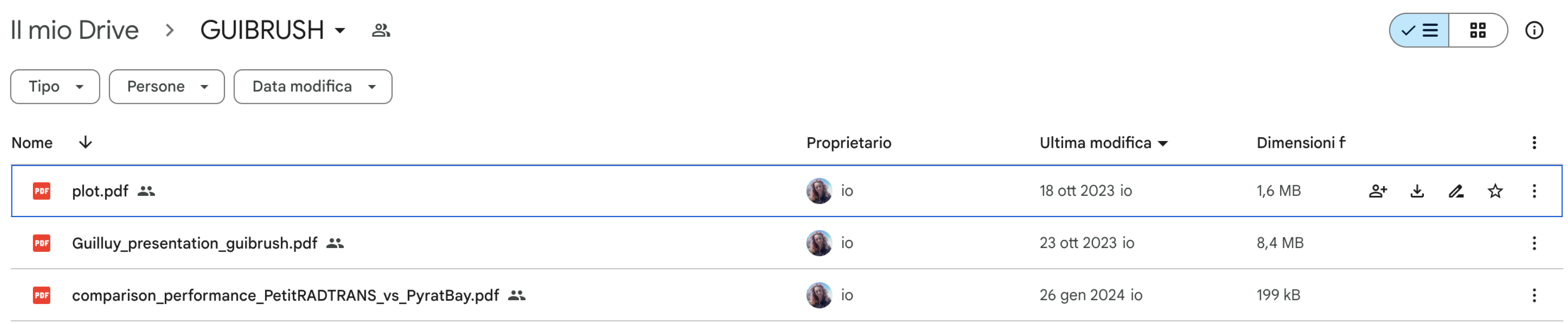
$$\text{Transit depth} = \frac{1}{R_*^2}\left[R_{\text{top}}^2 - 2\int_0^{R_{\text{top}}} e^{-\tau_\lambda} b\,db\right]$$

- **Goal**: reduce the computation time of a single atmospheric model by at least a factor of 10, that is about or less than 1 s.

Timescale, Milestones and KPIs

Milestone	Target	KPIs	Date
M6	Translation of the differential evolution Markov chain Monte Carlo (DE-MCMC) Bayesian code from IDL to Python > 3.8 and parallelization of the DE-MCMC code with the Multiprocessing Python library (process class)	code available at https://www.ict.inaf.it/gitlab/paolo.giacobbe/giano-b (private gitlab repository); plot of the DE-MCMC posterior distributions of the free parameters obtained after a GUIBRUSH-R analysis for a typical hot Jupiter exoplanet (plot.pdf at https://drive.google.com/drive/folders/1uPRI3zgJqUNDUYHafKRmCd3HLH26HePQ?usp=sharing); presentation made by G. Guilluy at the Spoke 3 Technical Meeting in Trieste, 9-11 October 2023, (Guilluy_presentation_guibrush.pdf at https://drive.google.com/drive/folders/1uPRI3zgJqUNDUYHafKRmCd3HLH26HePQ?usp=sharing)	August 2023

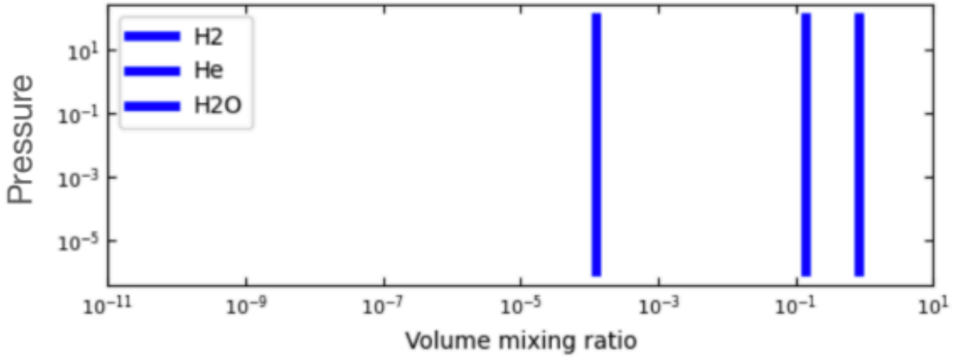
Decrease in computing time by a factor of 15



Timescale, Milestones and KPIs

Milestone	Target	KPIs	Date
M7	Decision on which radiative transfer code to use	internal report available at https://drive.google.com/drive/folders/1uPRI3zgJqUNDUYHafKRmCd3HLH26HePQ?usp=sharing (file comparison_performance_PetitRADTRANS_vs_PyratBay.pdf)	Dec 2023

- We generated a transmission spectrum for an atmosphere composed of H2O, H2 and He with both PYRATBAY and PetitRADTRANS. The investigated atoms and molecule have a uniform distribution at different pressures (atmospheric layers) with fixed volume mixing ratio. We simulated 100 atmospheric layers with a pressure between $10^{-6} - 10^{+2}$ bar



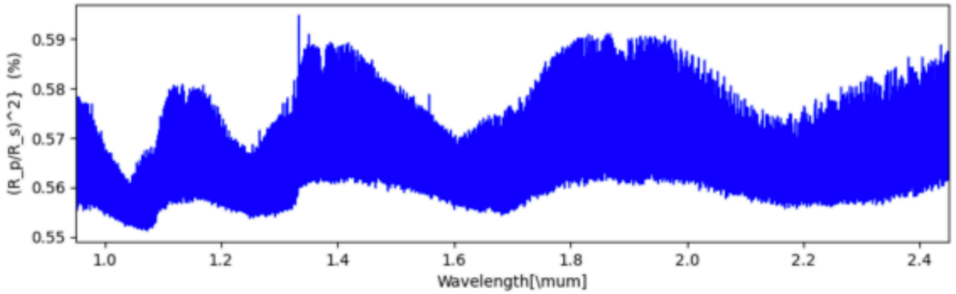


Figure 2, Transmission spectrum as a function of wavelength

- We computed the opacity for H2O from the HITEMP linelist at a resolution of R=250,000. We used a wavelength range [0.95-2.45] µm, which is the same range of the GIANO-B spectrograph at the Telescopio Nazionale Galileo (La Palma island).

FINAL RESULT=> PetitRADTRANS takes: 2.630 s, while PYRATBAY takes: 2.405 s.

We thus decided to use the RADIATIVE TRANSFER CODE implemented in PYRATBAY, as it is a bit faster and the developer of the code is a close collaborator.

Timescale, Milestones and KPIs

Milestone	Target	KPIs
M8	Coding and tests for preliminary porting of the radiative transfer code from CPU to GPU	comparison of code execution times with and without GPUs as model complexity changes (first just H2O, then two molecules, then four, eight, and so on...)

- We have written a preliminary code for porting to GPU (with an initial help of Valentina Cesare, INAF-OACt),

—> We are currently using PyOpenCL;

—> We tested the new code to generate a model with only water for the GIANO-B range, we recorded a slight increase in speed compared to the original code. <u>However, the code is still much slower than desired;</u>

—> We updated our new code in a <u>private repository on GitHub</u> and shared it with Giuseppe Puglisi (INAF-OACt) who is helping us to review it;

—> However, we could not test it well yet due to lack of resources. We are waiting for the new server with the Nvidia GPU

Accomplished Work, Results

- We **translated the DE-MCMC routines from IDL to Python and parallelized them with the Multiprocessing Python library** (process class), which has decreased the computation time approximately by a factor of 15;

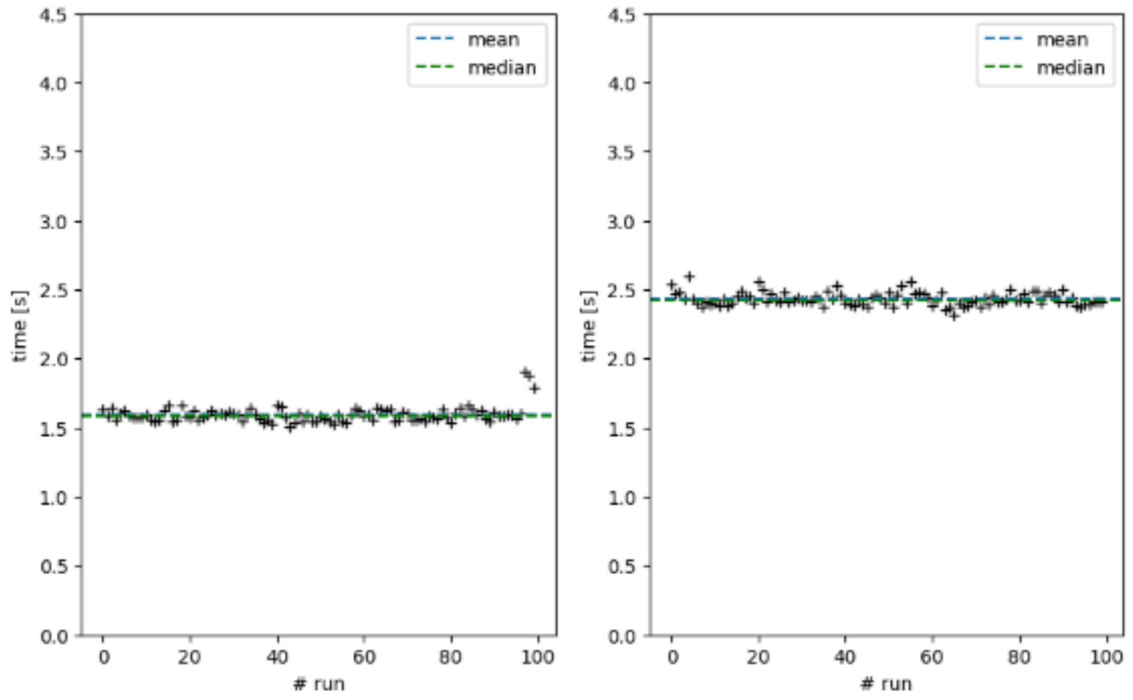
- We studied and got familiar with **radiative transfer codes PYRAT BAY** (Cubillos & Blecic 2021) and **petitRADTRANS** (Mollière 2019);

- We compared the performance of PYRAT BAY with that of petitRADTRANS;

- We are working on the **porting of PYRAT BAY to GPU.** We will also try to do this for PetitRADTRANS "time permitting"

Next Steps and Expected Results

- Ongoing collaboration with UniCT (Dr. G. Puglisi)
- Try to decrease the computation time
- Finalize the GPU porting of the radiative transfer code
- Implementation of the GPU ported radiative transfer code within GUIBRUSH®
- Final tests of GUIBRUSH® on real low-resolution (HST and JWST) and high-resolution (GIANO-B@TNG) exoplanet transmission spectra



mean
median
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0
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40
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80
100