

# Watching the Universe's Acceleration Era with the SKAO



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**Φ IN THE  
SKY**

# Mind Your Cosmological Priors

If  $\Omega = -1$  is a good fit to all the data. If  $w = -1$ , then flat  $\Lambda$ CDM is a good fit to all the data. If  $w = 1$ , then  $\Omega = -1$  is a good fit to all the data.

# The Redshift Drift

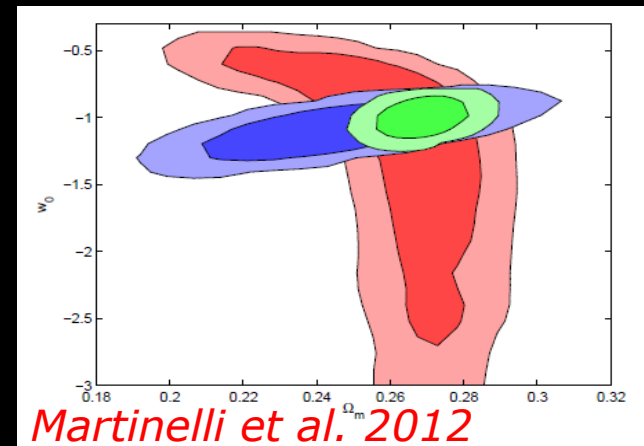
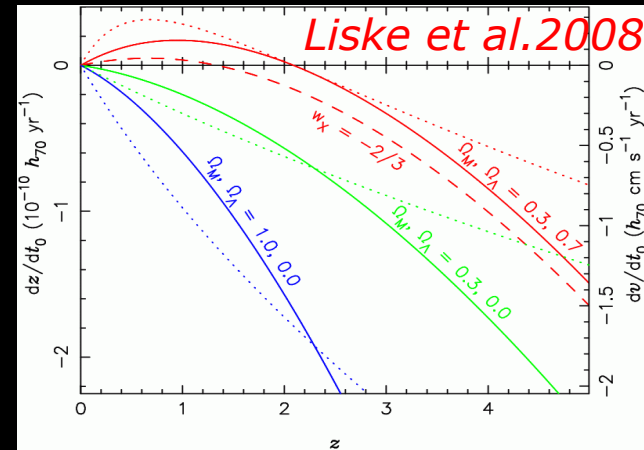
A direct non-geometric model-independent measurement of the universe's expansion history [Sandage 1962]

- Watching the universe expand in real time!
- Independent of gravity, geometry or clustering
- Not mapping (present-day) past light-cone, but directly comparing different past light-cones

$$\dot{z} \equiv \frac{dz}{dt_{\text{obs}}}(t_0) = (1+z)H_0 - H(z)$$

ELT flagship science driver (for  $z > 2$ ) [Liske et al. 2008], unique tool to close consistency loop and break parameter degeneracies

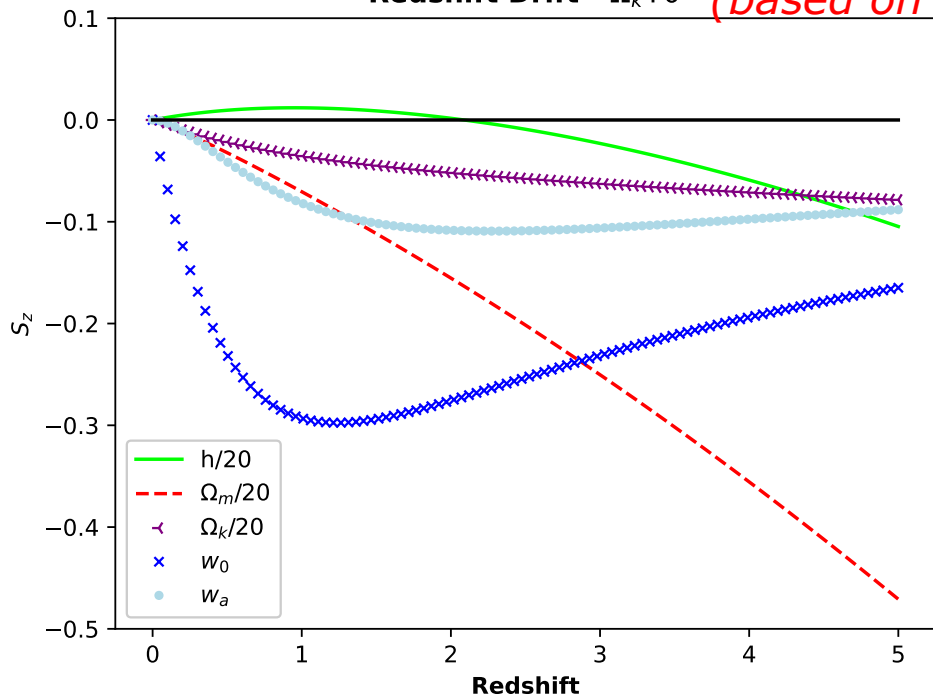
- SKAO can cover  $z < 1$  [Klockner et al. 2015, ...]
- In practice one measures a spectroscopic velocity



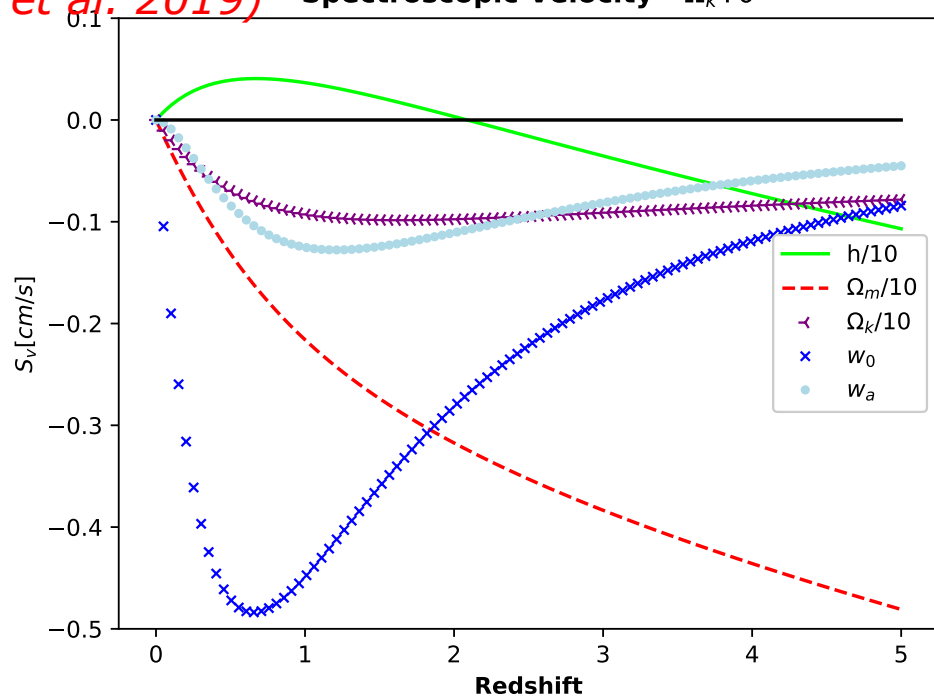
# Cosmological Parameter Sensitivity (for a CPL fiducial model)

*Melo e Sousa et al. 2023*  
(based on Alves et al. 2019)

Redshift Drift -  $\Omega_k : 0$



Spectroscopic Velocity -  $\Omega_k : 0$



# Real-time Cosmography

First and second redshift derivatives are powerful test of the  $\Lambda$ CDM paradigm; cosmographic approach useful here [Martins et al. 2016]

$$Z_1(z) = \frac{1}{H_0} \frac{dz}{dt_0} = 1 + z - E(z)$$

$$Z_2(z) = \frac{1}{H_0^2} \frac{d^2z}{dt_0^2} = \frac{1+q(z)}{1+z} E^2(z) - E(z) - q_0(1+z)$$

$$\frac{dZ_1(z)}{dz} = 1 - E(z)'$$

- To linear order,

$$Z_1 = -q_0 z + O(z^2)$$

$$Z_2 = j_0 z + O(z^2)$$

$$\frac{dZ_1(t_0, z)}{dz} = -q_0 + (q_0^2 - j_0)z + O(z^2)$$

Assuming specs discussed in [Klockner et al. 2015], SKAO redshift drift measurements can reach  $\sigma_{q_0} \sim 0.006$  and  $\sigma_{j_0} \sim 0.13$  [Martins et al. 2016]

- Optimal way to measure  $q_0$  with both accuracy and precision, which is not possible with traditional distance indicators [Neben & Turner 2013]
- A key consistency test:  $j(z)=1$  at all redshifts for a flat  $\Lambda$ CDM universe
- More broadly, a positive drift implies SEC violation, hence dark energy

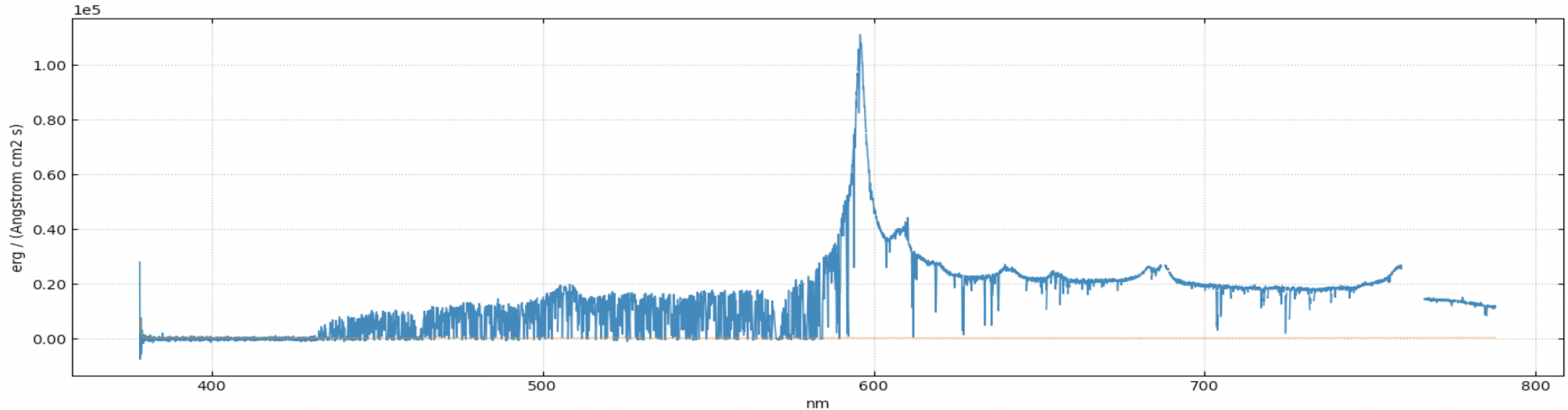
# The ESPRESSO Redshift Drift Experiment

Current limits 1000x larger than expected signal, and systematics-dominated

- [Darling 2012] in the radio at  $z < 0.7$ , [Cooke 2020] in the optical at  $z > 2$

ESPRESSO can improve this by a factor  $\sim 10$  with an experiment time of 1 year and an observation time of 40h for 2 QUBRICS 'superbright' QSOs

- Test and optimise methodology with real data, test instrument stability
- Two independent experiments, also 'zeroth epoch' for ANDES (calibration permitting)



# SKA Forecast Assumptions

Only available estimate of SKA redshift drift sensitivity is [Klockner et al. 2015]

- Observe HI signal of ca.  $10^7$  galaxies up to  $z \sim 1$  at 2+ different epochs
- Observation time ca. 0.5 years, experiment time ca. 12 years – expect  $\Delta\nu \sim 0.1$  Hz
- Only feasible with full SKA (as it was in 2014)
- Sensitivity, number counts, hardware, systematics (e.g. observatory motion) etc.

These forecasts are simplistic and out of date, it would be good to improve them to ELT level.

- Can someone do simulations?
- Good topic for someone's thesis?
- For now, proceed with these...

What upper limits can we get now?  
Existing ones can be improved!

- Relevant Phi in the Sky expertise



PROCEEDINGS  
OF SCIENCE

## Real time cosmology - A direct measure of the expansion rate of the Universe with the SKA

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# Available SKA Forecasts

Under previous assumptions,  
competitive probe of low- $z$   
accelerating universe

- MCMC in [Rocha & Martins 2022]

- FMA in [Marques et al. 2024]

- Both cosmography and specific  
cosmological models

- Latter code publicly available

Can think of measuring the  
drift of the drift, look at  
spatial variations, etc

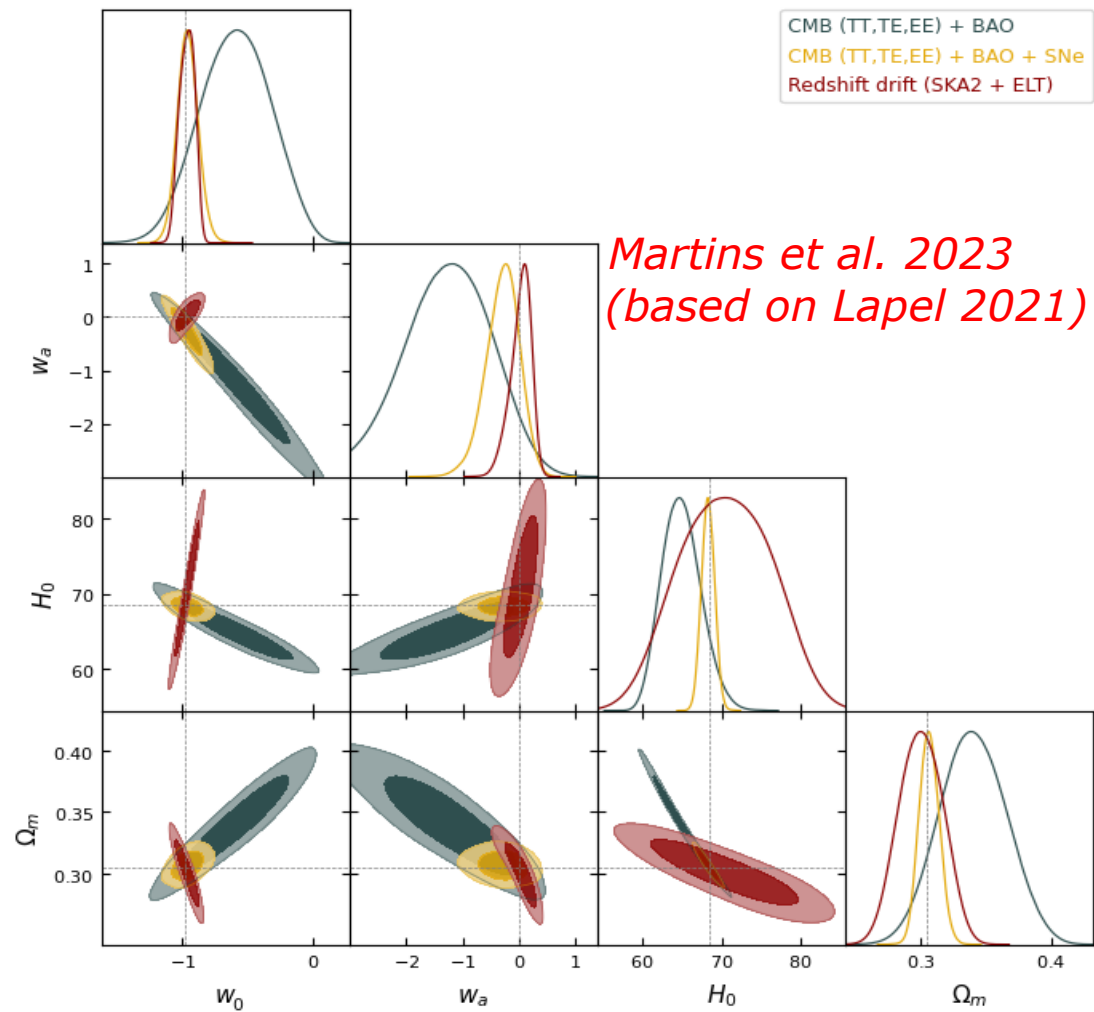
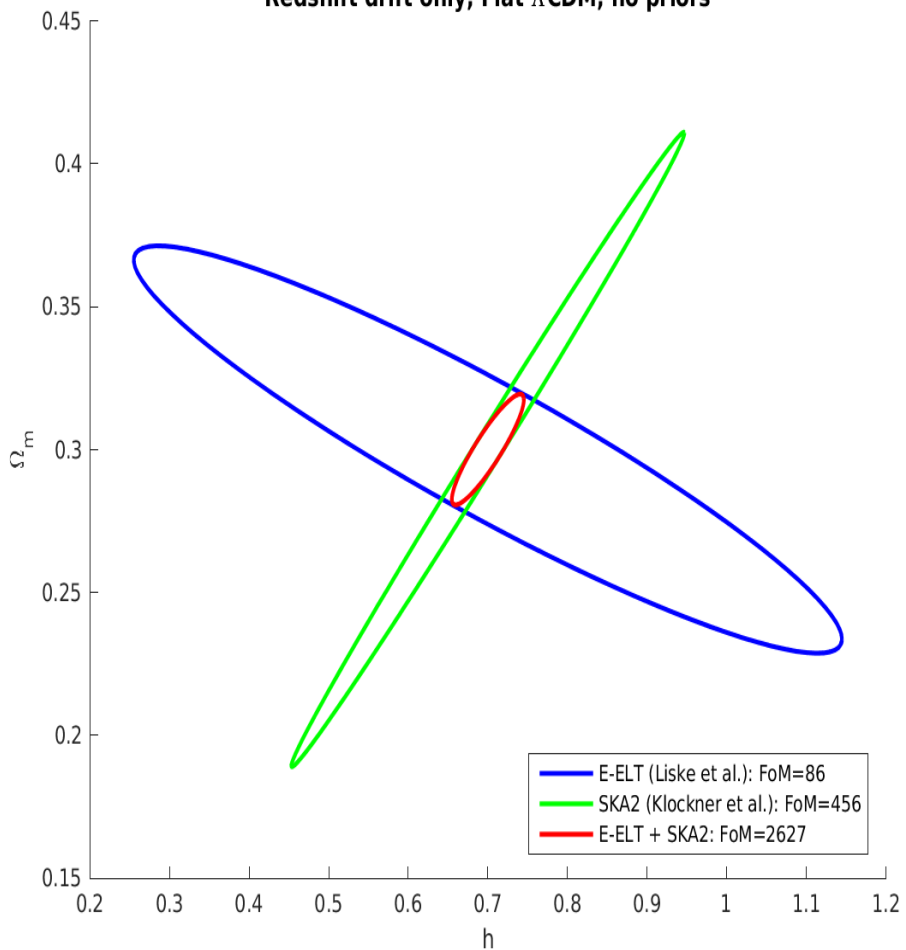
	Parameter	$\sigma_h = 0$			$\sigma_h = 0.05$			$\sigma_h = 0.1$		
		$z \leq 0.3$	$z \leq 0.5$	$z \leq 1.0$	$z \leq 0.3$	$z \leq 0.5$	$z \leq 1.0$	$z \leq 0.3$	$z \leq 0.5$	$z \leq 1.0$
$\Lambda$ CDM	FoM( $w_0, w_a$ )	95	139	252	48	62	110	44	50	77
	$\rho(w_0, w_a)$	-0.624	-0.803	-0.905	-0.226	-0.433	-0.761	-0.156	-0.287	-0.609
	$\sigma(w_0)$	0.038	0.034	0.024	0.054	0.051	0.040	0.062	0.062	0.056
	$\sigma(w_a)$	0.230	0.214	0.207	0.352	0.309	0.285	0.376	0.349	0.316
Freezing	FoM( $w_0, w_a$ )	118	183	352	50	71	139	44	53	91
	$\rho(w_0, w_a)$	-0.614	-0.805	-0.934	-0.221	-0.432	-0.798	-0.154	-0.288	-0.647
	$\sigma(w_0)$	0.035	0.033	0.027	0.048	0.045	0.035	0.055	0.055	0.048
	$\sigma(w_a)$	0.172	0.151	0.147	0.325	0.254	0.205	0.365	0.317	0.245

	Parameter	$\sigma_h = 0$			$\sigma_h = 0.05$			$\sigma_h = 0.1$		
		$z \leq 0.3$	$z \leq 0.5$	$z \leq 1.0$	$z \leq 0.3$	$z \leq 0.5$	$z \leq 1.0$	$z \leq 0.3$	$z \leq 0.5$	$z \leq 1.0$
n=5	FoM( $q_0, j_0$ )	54	90	117	4	7	11	2	3	4
	$\rho(q_0, j_0)$	-0.988	-0.979	-0.969	-0.944	-0.954	-0.940	-0.891	-0.935	-0.935
	$\sigma(q_0)$	0.042	0.029	0.024	0.112	0.098	0.079	0.154	0.152	0.132
	$\sigma(j_0)$	1.226	0.801	0.641	2.736	2.016	1.445	3.196	2.900	2.182
n=4	FoM( $q_0, j_0$ )	37	54	245	4	5	18	2	2	6
	$\rho(q_0, j_0)$	-0.983	-0.986	-0.956	-0.940	-0.936	-0.942	-0.890	-0.908	-0.938
	$\sigma(q_0)$	0.045	0.039	0.016	0.112	0.102	0.068	0.156	0.151	0.124
	$\sigma(j_0)$	1.422	1.203	0.368	2.736	2.365	1.063	3.200	2.969	1.822
n=3	FoM( $q_0, j_0$ )	83	259	845	4	15	81	2	4	25
	$\rho(q_0, j_0)$	-0.986	-0.959	-0.895	-0.959	-0.948	-0.933	-0.925	-0.943	-0.939
	$\sigma(q_0)$	0.034	0.017	0.009	0.123	0.076	0.041	0.177	0.141	0.077
	$\sigma(j_0)$	0.917	0.355	0.129	2.797	1.207	0.365	3.706	2.171	0.644
n=2	FoM( $q_0, j_0$ )	788	1508	4050	42	128	652	11	37	236
	$\rho(q_0, j_0)$	-0.917	-0.865	-0.805	-0.943	-0.937	-0.950	-0.944	-0.942	-0.963
	$\sigma(q_0)$	0.010	0.007	0.005	0.056	0.038	0.023	0.108	0.072	0.043
	$\sigma(j_0)$	0.138	0.078	0.033	0.562	0.259	0.092	1.081	0.483	0.160



# Synergies: ELT + SKAO

Redshift drift only, Flat  $\Lambda$ CDM, no priors





***Let's do it!***