

thanks for  
the logo btw



# Updates on Galapy

a fast API for modelling galaxy SEDs with Bayesian sampling

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**@Spoke3-WP1&2 monthly meeting**

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# Extract (astro-)Physical Info from the SED of galaxies

At different **wavelengths**  
→ different **components** dominate

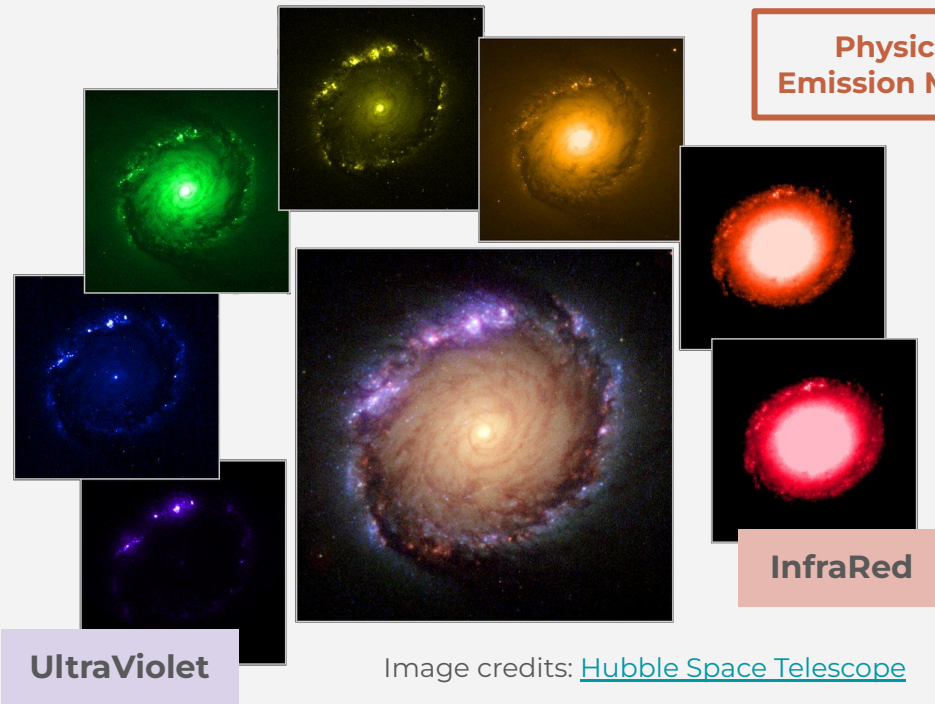
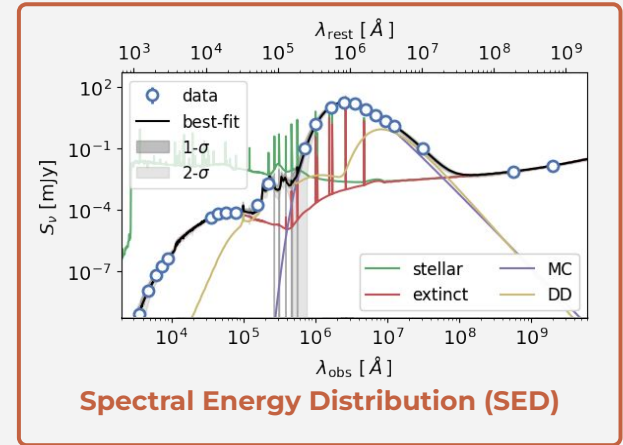
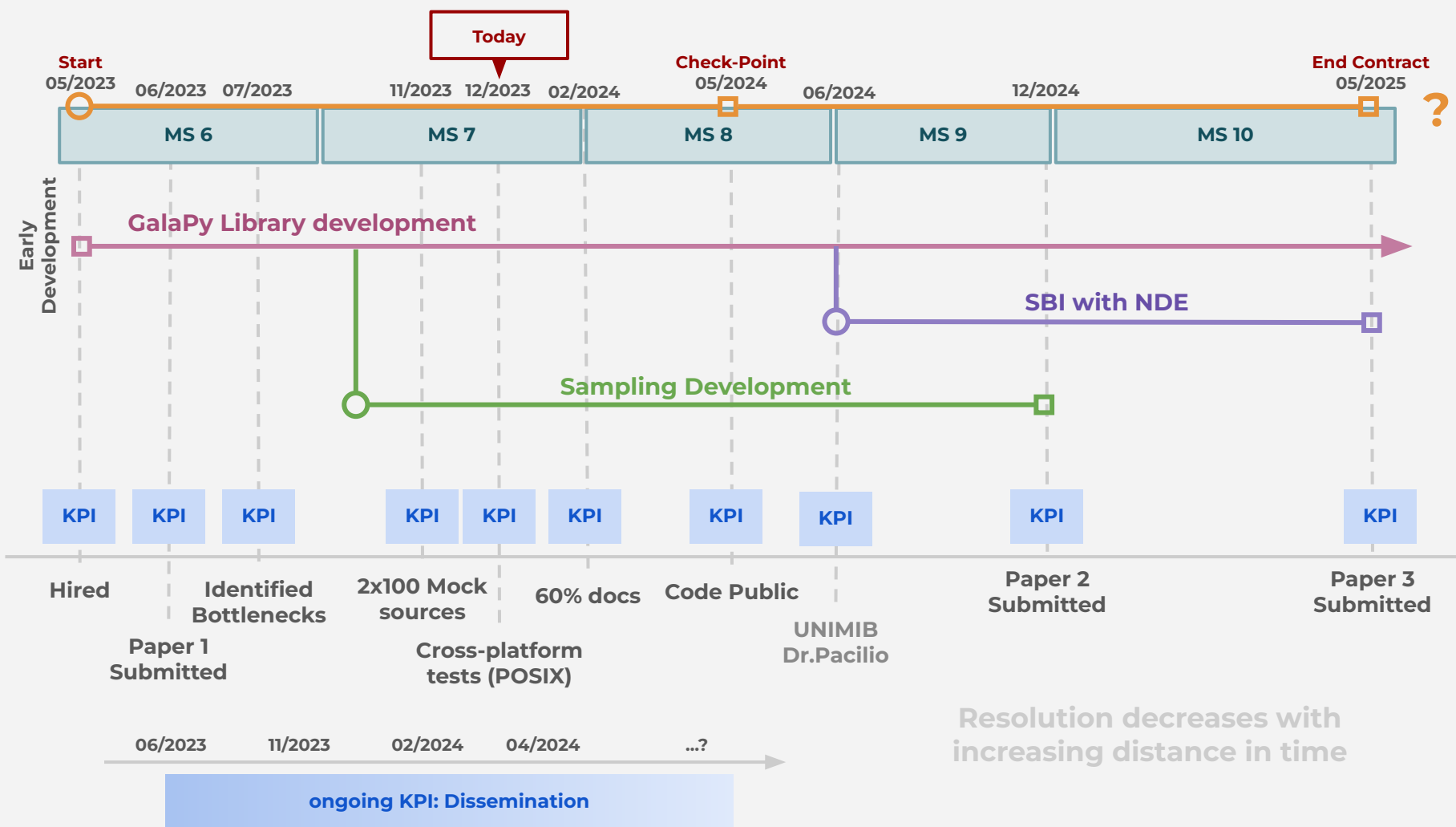


Image credits: [Hubble Space Telescope](#)



→ 1 zettabyte of data per year!



# Key Performance Indicators

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MS6 & MS7



**SUBMITTED!**

Soon public on [GitHub](#) and [PyPI](#)

## New modelling

- **Analytical solutions for Galaxy Evolution:**  
realistic modelling of dust absorption/emission  
Gas, Stars, Metals and Dust  
([Lapi et al., 2018](#); [2020](#), [Pantoni et al., 2019](#))
- **2-components dust model:**  
as seen in GRASIL ([Silva et al., 1998](#))  
**but AGE DEPENDENT ENERGY BALANCE**  
(no radiative transfer + physical temperature)

## Performance

- **Hybrid C++/CPython** implementation (performance + user friendly)
- First release: **shared memory parallelization** from Python

## Bayesian framework

First release:

- **Markov Chain Monte Carlo** with [emcee](#) [Foreman-Mackey et al., 2013](#)
- **Dynamic Nested Sampling** with [dynesty](#) [Speagle, 2019](#)

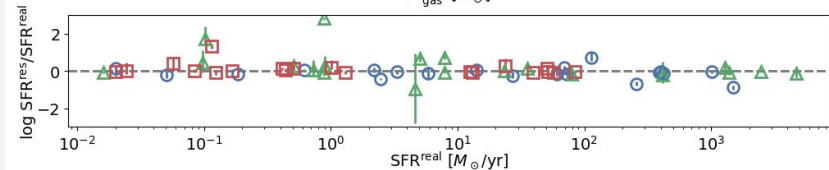
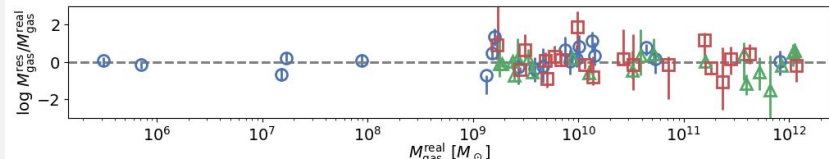
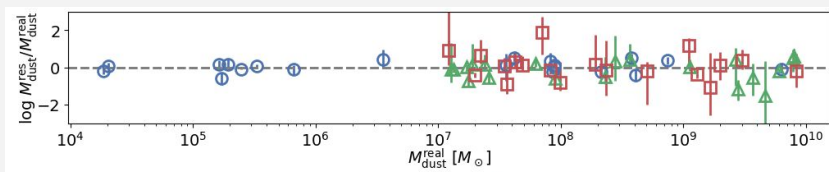
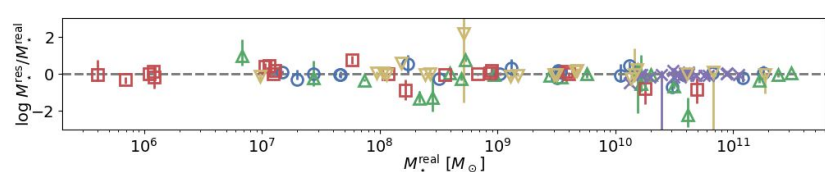
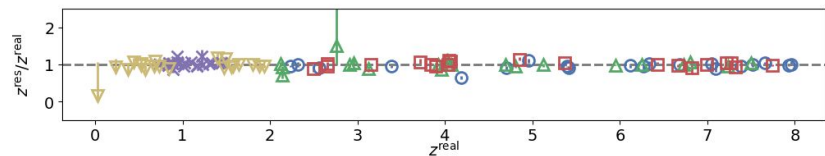
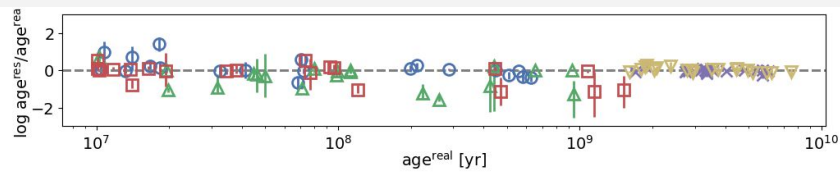
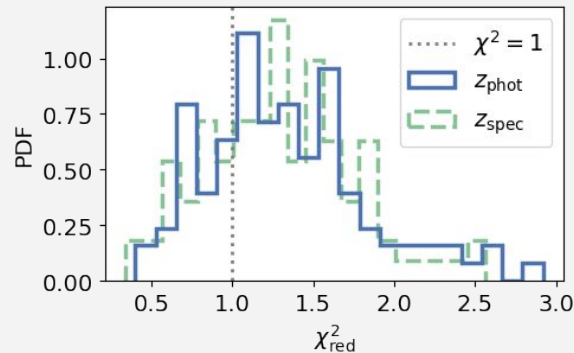
# Validation on 2x100 mock sources

## Different evolution phases

Different Models

	ACTIVE	PASSIVE
In-Situ		
Delayed-Exp.		
Constant		

- + Uninformative priors
- + Different experiments configuration



# 1st Bottleneck = CSP computation

KPI

SSP

- wavelength  $N\lambda > 2000$
- age of SSP  $N\tau > 100$
- metallicity  $NZ \sim 10$

$> 10^6$  double precision



~10 MB



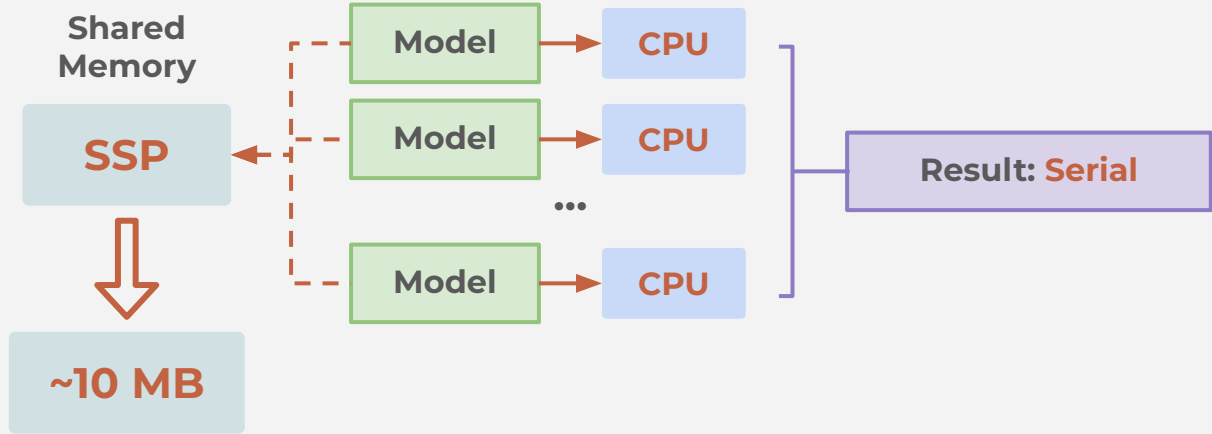
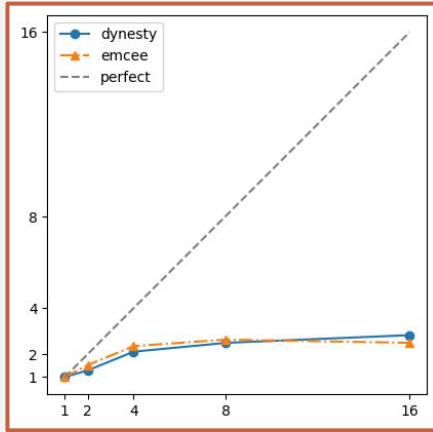
SFH

- at a given AGE

- $SFR(\tau < AGE)$
- $Z_{stars}(\tau < AGE)$
- $idx_Z(\tau < AGE)$

$$L_{CSP}(\lambda_i, \tau_{GXY}) = \sum_{\forall j > 0 \mid \tau_j \lesssim \tau_{GXY}} \frac{\tau_j - \tau_{j-1}}{2} \times$$
$$\times \left\{ \psi(\tau_j) P_{L_{SSP}}^{(1)}[\lambda_i, \tau_j, Z_*(\tau_{GXY} - \tau_j)] + \right.$$
$$\left. + \psi(\tau_{j-1}) P_{L_{SSP}}^{(1)}[\lambda_i, \tau_{j-1}, Z_*(\tau_{GXY} - \tau_{j-1})] \right\}$$

# 2nd “Bottleneck” = scaling, but why?



```
$ lscpu
[...]
Caches (sum of all):
  L1d:    256 KiB (8 instances) -> 32 KiB
  L1i:    256 KiB (8 instances) -> 32 KiB
  L2:     2 MiB (8 instances)   -> 256 KiB
  L3:    16 MiB (1 instance)
[...]
```

- It is still bugging me the way memory is dealt by the processes in particular, **access to the SSP matrix**
- my two cents: since the matrix cannot be loaded all in L1/L2 private caches, **other processes trigger cache misses**
- is it therefore a problem of **cache invalidation?**



# Others on MS7

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## KPI

Cross-Platform testing

**Currently:**

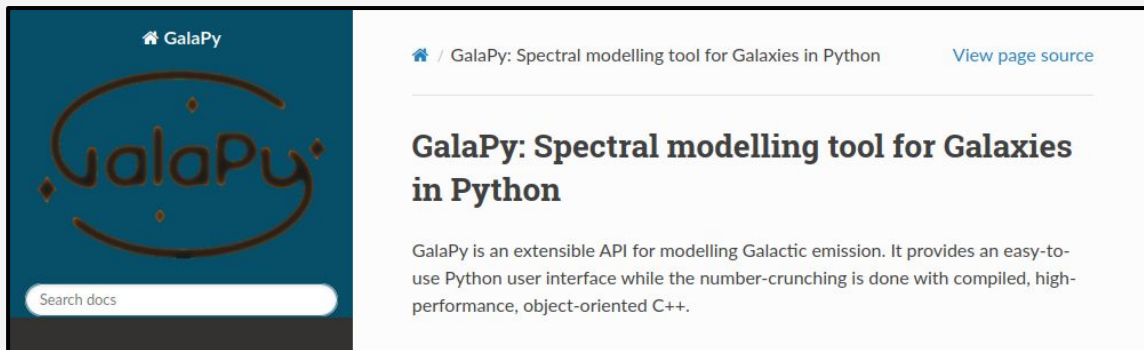
- Linux
- MacOS

**Missing:**

- Windows (..sigh)

## KPI

**~60% Documentation** completed

A screenshot of the GalaPy documentation page. The page has a dark blue header with the GalaPy logo and a search bar. The main content area is white and contains the title "GalaPy: Spectral modelling tool for Galaxies in Python" and a brief description of the tool. The description states: "GalaPy is an extensible API for modelling Galactic emission. It provides an easy-to-use Python user interface while the number-crunching is done with compiled, high-performance, object-oriented C++." The page also includes a "View page source" link in the top right corner.

🏠 GalaPy

🏠 / GalaPy: Spectral modelling tool for Galaxies in Python [View page source](#)

## GalaPy: Spectral modelling tool for Galaxies in Python

GalaPy is an extensible API for modelling Galactic emission. It provides an easy-to-use Python user interface while the number-crunching is done with compiled, high-performance, object-oriented C++.

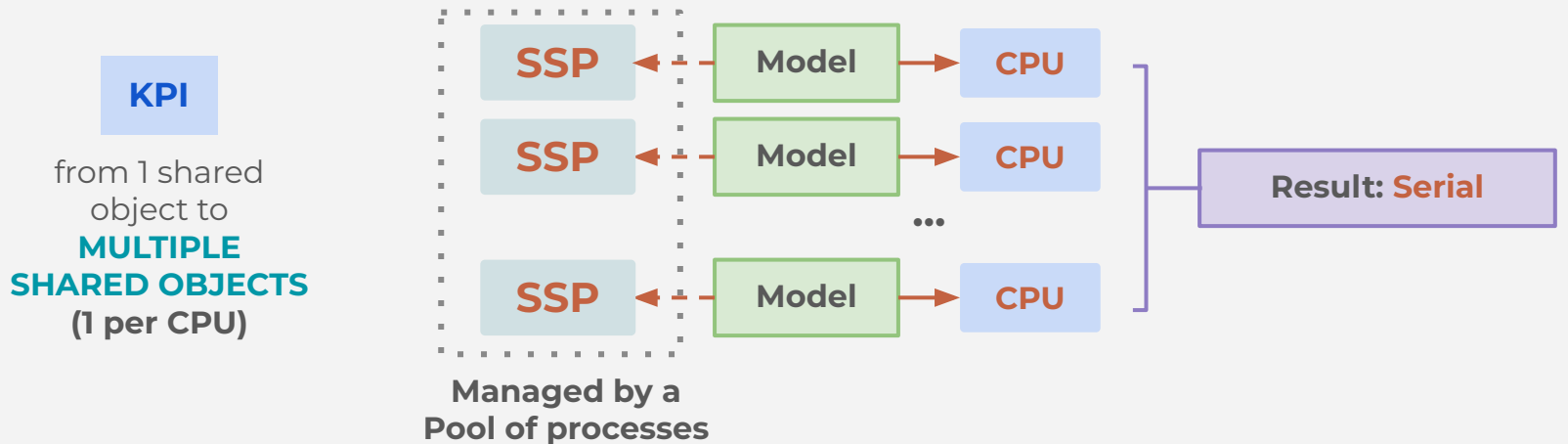
**Currently ~30%**

## **(foreseen) KPIs for MS8**

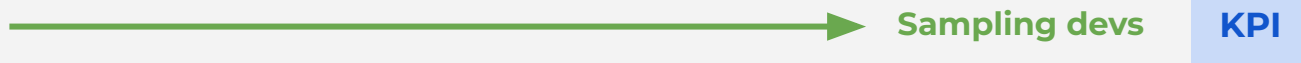
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- **Paper 1 + Code public**
- **Parallelisation development**

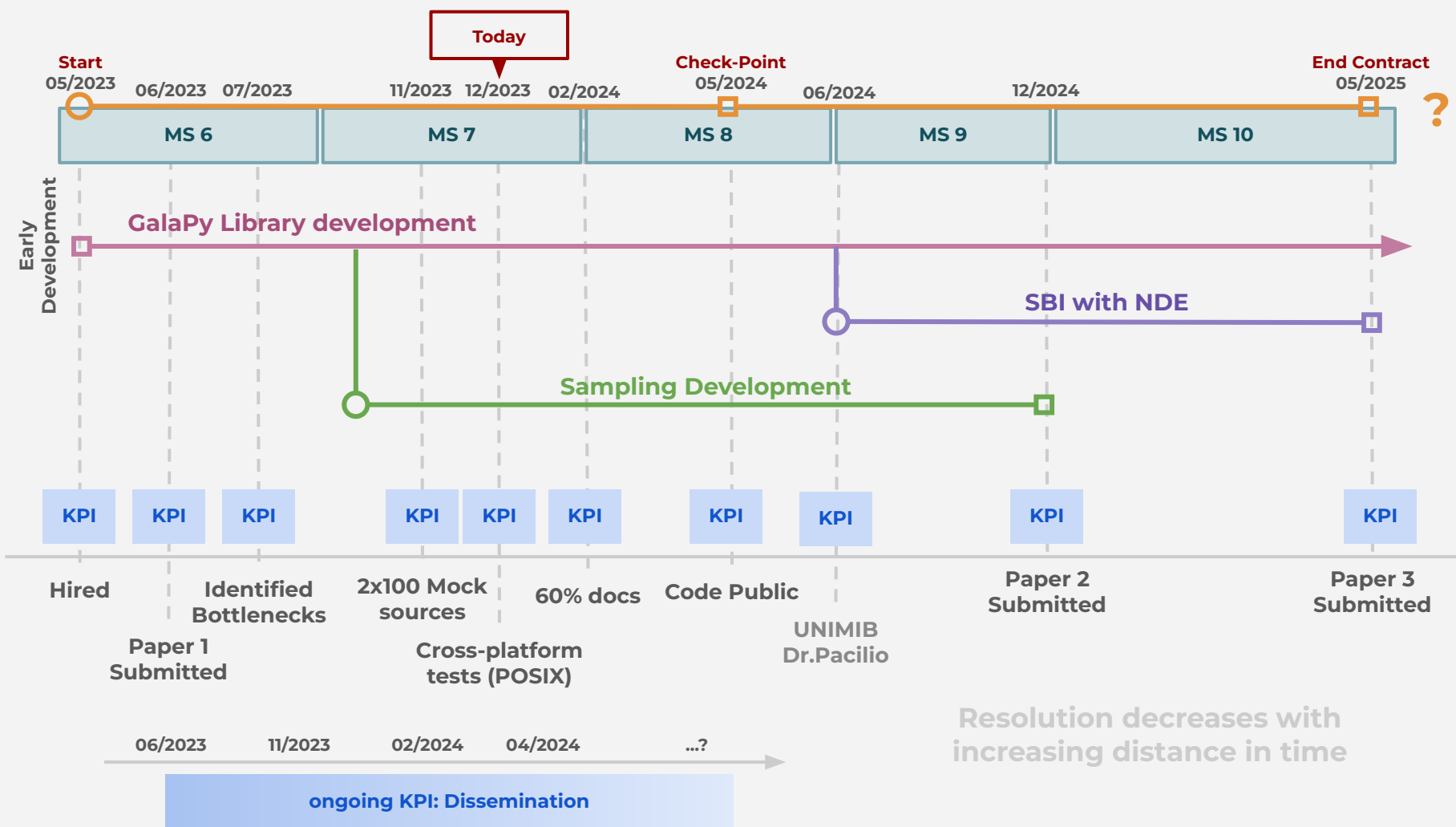
# Foreseen KPIs for MS8



- **1 source: multiple positions** in parameter space
- **Multiple sources: multiple models** sampled serially



(extends up to end of MS9)

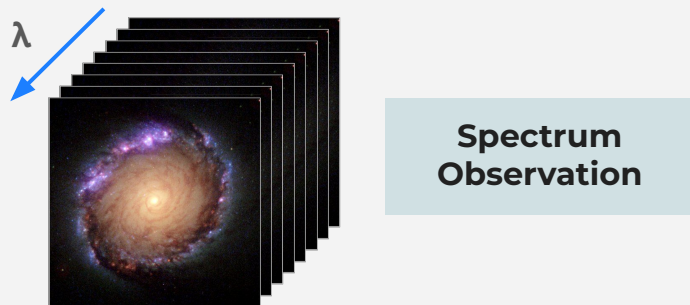




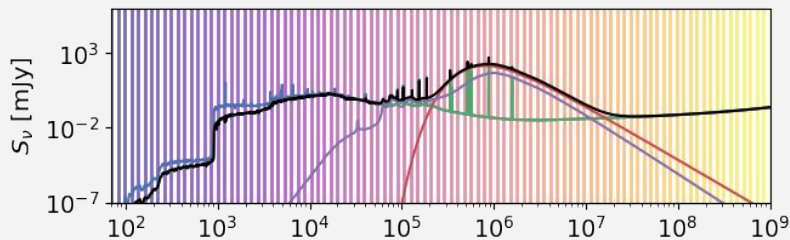
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... and this is my last slide, thanx!

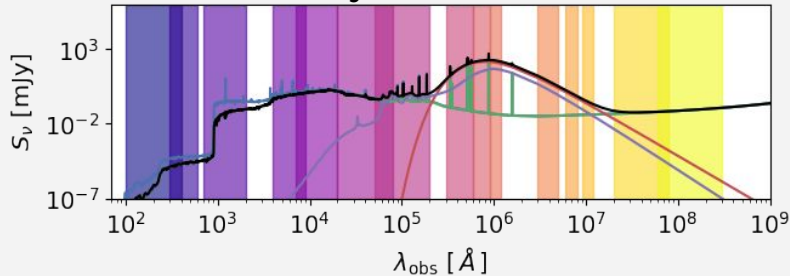
# SED fitting 101



→ Ideally:

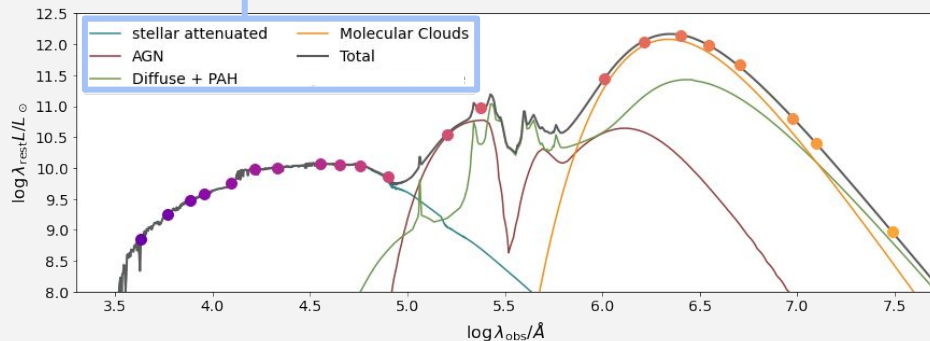


→ Realistically:



Physical Emission Model

← parameterized

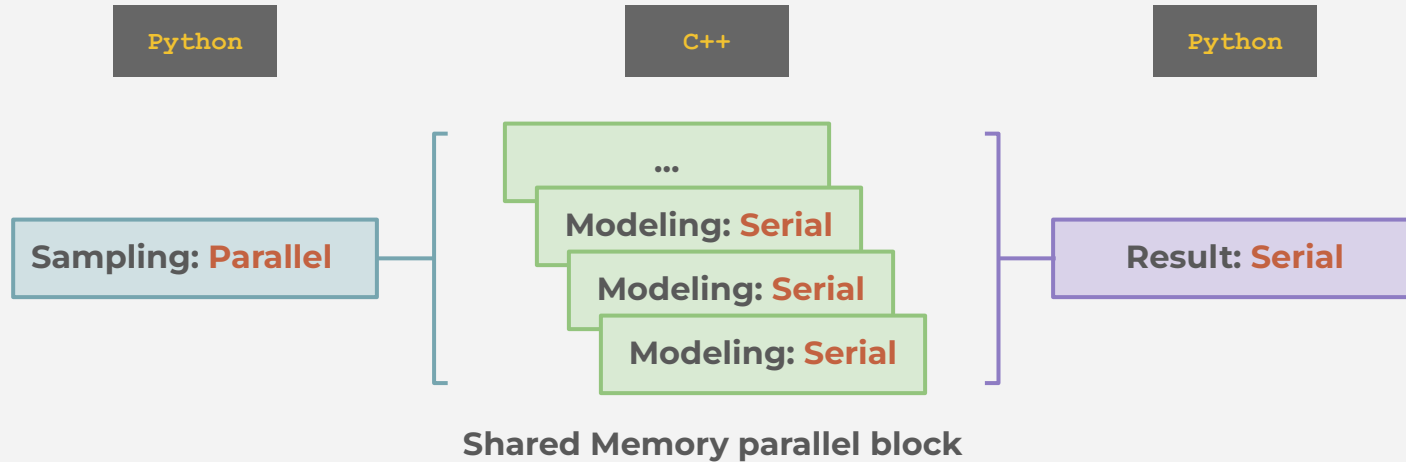


- + observation
- + parameterized model
- + sampling/fitting strategy

→ **optimal model**  
representing your data  
(hopefully)

# Parallelisation scheme and why

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- **User-oriented:** I do want users to be able to tailor parallelization on their system easily, so this has to happen on the higher level interface
- **Future devs in the sampling strategy:** Especially considering a future development with **hierarchical bayesian sampling** I will need to parallelize sampling on multiple sources