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MACHINE LEARNING-BASED PHOTOMETRIC CLASSIFICATION OF GALAXIES, QUASARS AND STARS

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ML4astro2

Introduction to the data

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Data from Sloan Digital Sky Survey (SDSS)

Most recent: Data Release 18 (DR18) https://www.sdss.org/dr18/

Contains100,000 patterns described by 43 features and related to three classes: galaxies, stars, and quasars.



Correlation Matrix



We analyzed the data using the correlation matrix, it was observed that one feature, 'rerun', is constant.

-0.8

0.6

-0.3

-0.4

 Additionally, we have seen that several features are highly correlated (correlation coefficient > 0.95). This allowed a reduction in dimensionality (36 features from the initial 43)

Modeling Workflow



 To get an idea of the complexity of the separability problem between classes, a plot 3D with Principal Component Analysis (PCA) is shown.
n components=3

```
overlap of the data
```

- □ The implemented model is therefore composed of the following steps:
- 1. Data Standardization
- 2. Correlation Matrix and Elimination of Features with Correlation > 0.95

3. Feature Selection through a Feature Importance Procedure: <u>led to a</u> <u>further dimensionality reduction, resulting in seventeen final features</u>

4. Classification

Classifier Analysis and Parameter Optimization

Various classifiers were analyzed:

- Random Forest (RF);
- XGBoost (XGB);
- Support Vector Machine (SVM);
- k-Nearest Neighbors (KNN)

For each of them, the relevant parameters were evaluated using grid-search procedures within a cross-validation framework

The SVM classifier, being a binary classifier, was evaluated using both the one-vs-one and one-vs-all strategies.

Results

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□ The best result was obtained with **Random Forest** classifier



- Mean Class Accuracy is affected by the sensitivity on the class (with fewer examples) quasars
- □ the parameter configuration that allowed it is as follows

```
- class_weights ={'STAR': 1.3, 'GALAXY': 1.0, 'QSO': 2.5}
```

n_estimators=301, max_samples=0.75

Conclusions and Future Perspectives

- The <u>dimensionality reduction</u>, besides allowing a more computationally intensive analysis, led to an increase the accuracy of about 8%.
- improving overall performance means enhancing the model's performance for the quasars class
- We plan to use Ensemble methods for further performance improvement.
- □ Thank you for your attention