

Tracing ongoing quenching in jellyfish galaxies

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& the GASP collaboration



OSSERVATORIO
ASTRONOMICO DI PADOVA

A long, long time ago...

Abell 2744 with JWST + HST by Peter Watson



Ram-pressure stripping and galaxy evolution

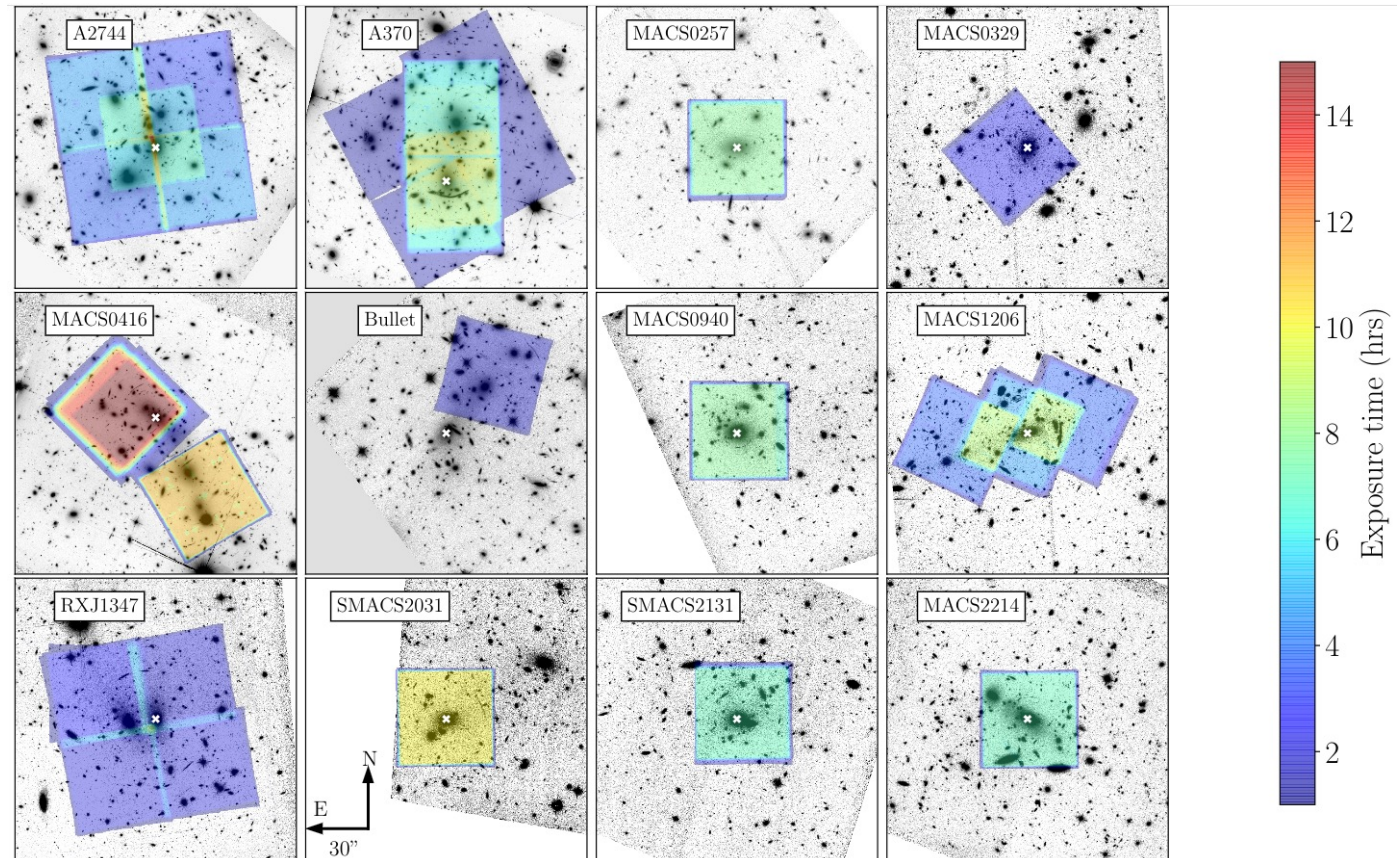
Jellyfish galaxies showcase **long tails of gas** that is being stripped by ram-pressure.

This process can lead to **fast quenching**, turning star-forming galaxies into **post-starbursts**.



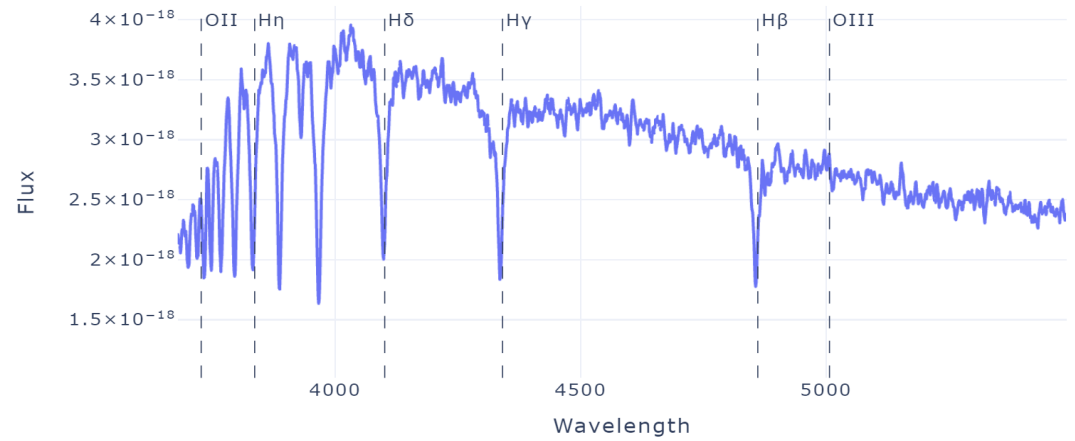
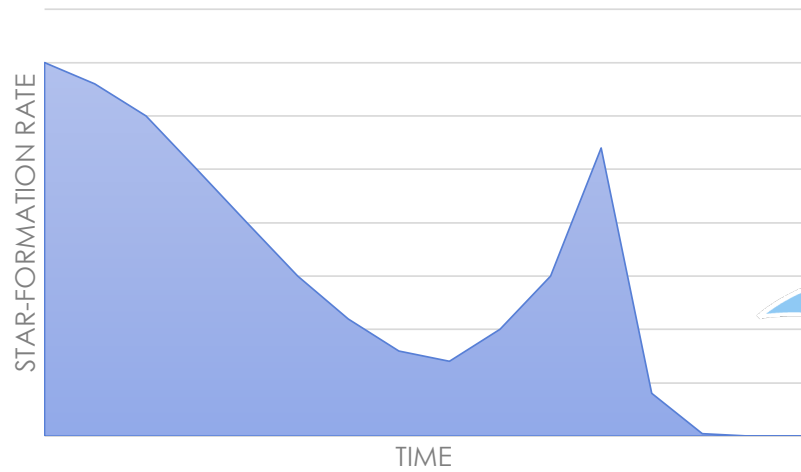
Gullieuszik et al. 2023, see also Giunchi et al. 2024ab and Werle et al. 2024

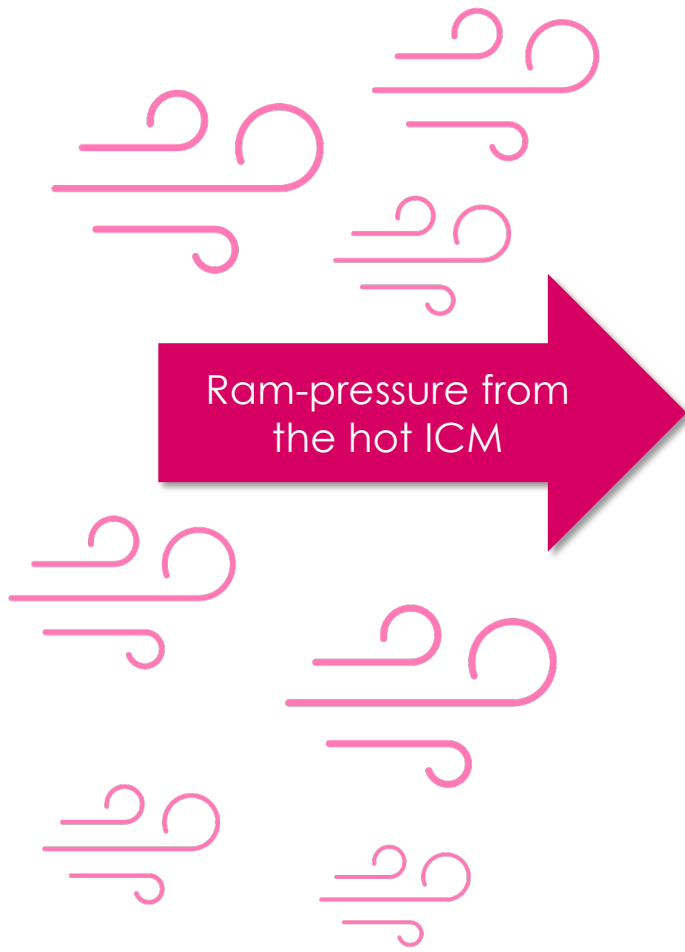
MUSE GTO clusters



Post-starburst galaxies

Star-formation history



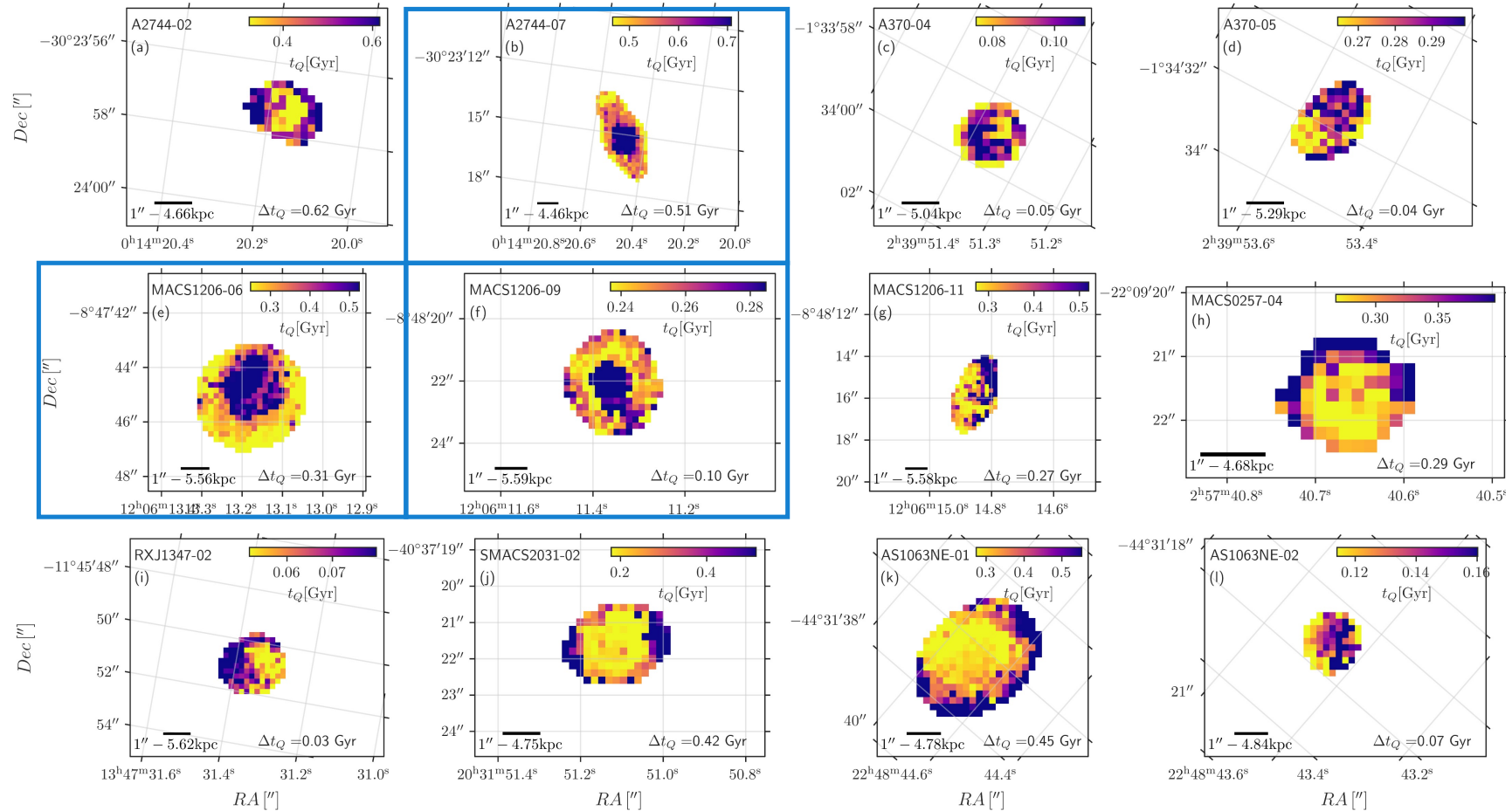


Side-to-side quenching



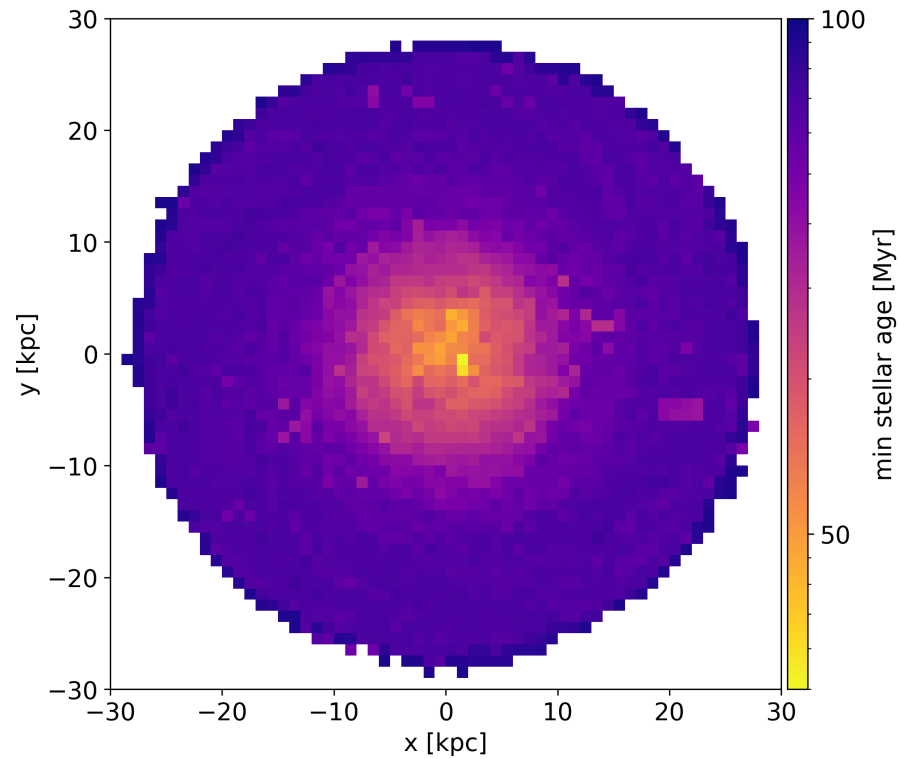
Outside-in quenching

Quenching directions in post-starbursts

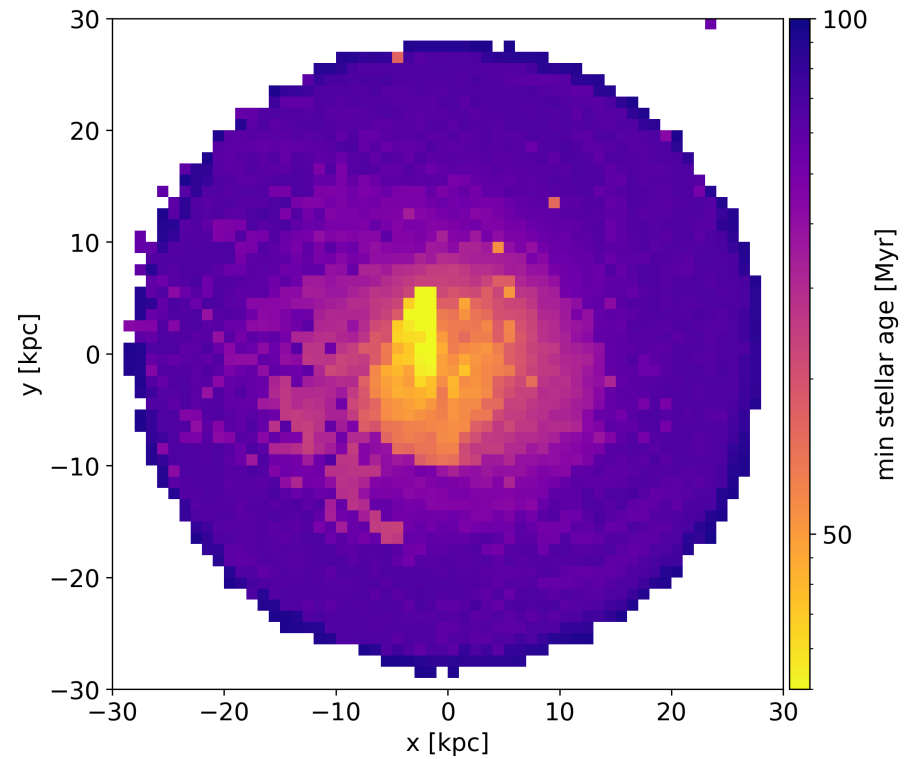


Simulations

Face-on wind



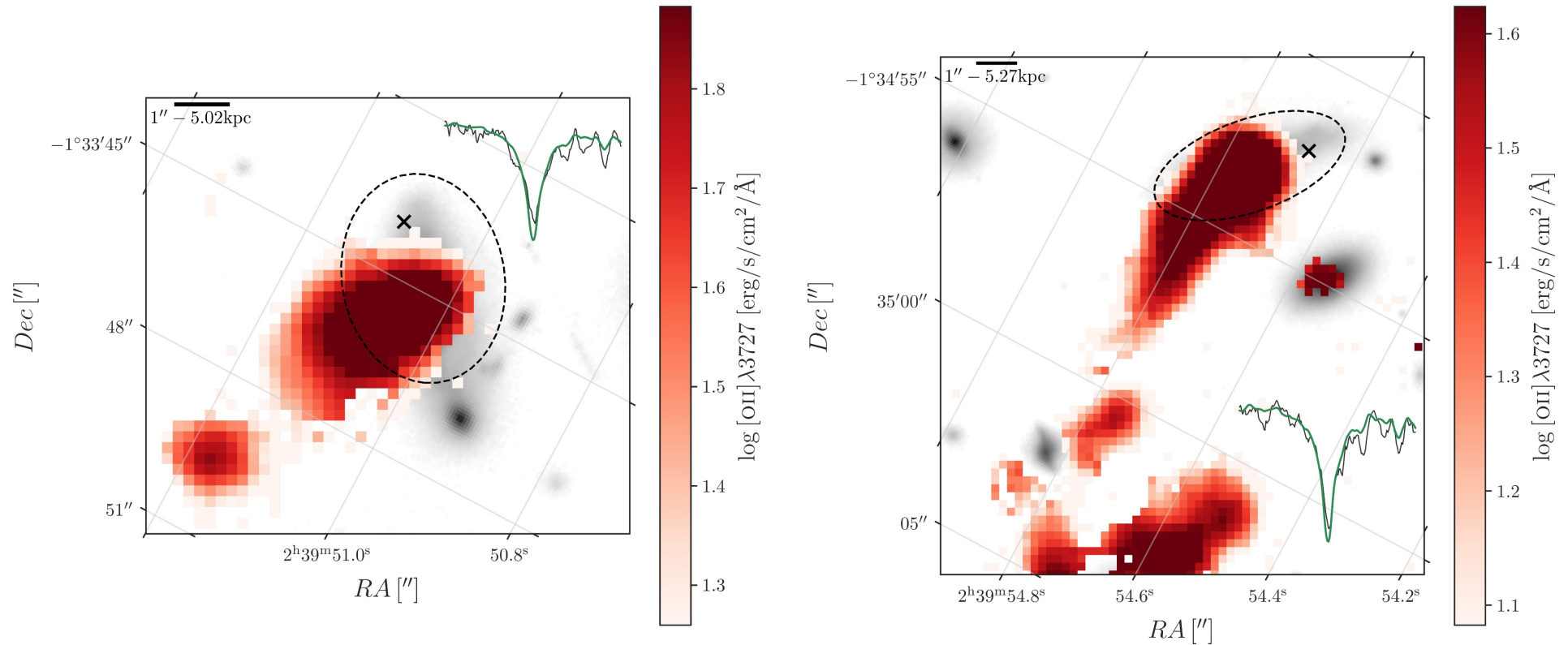
45° wind



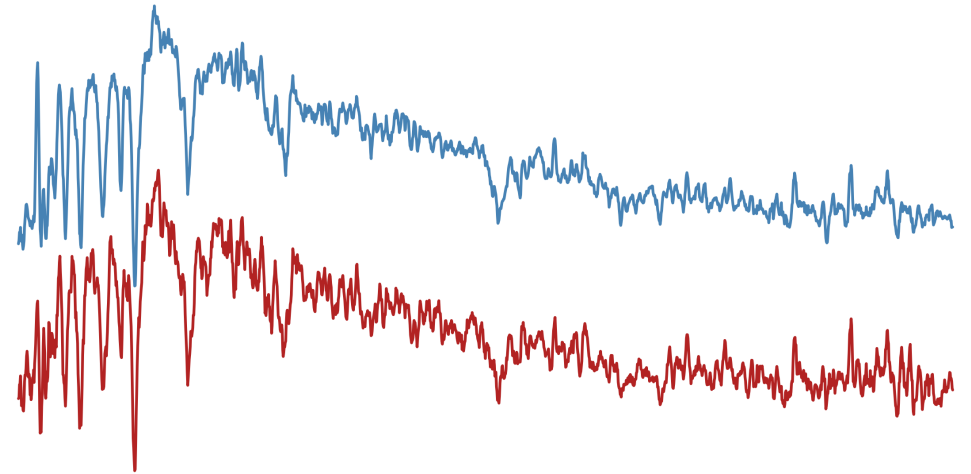
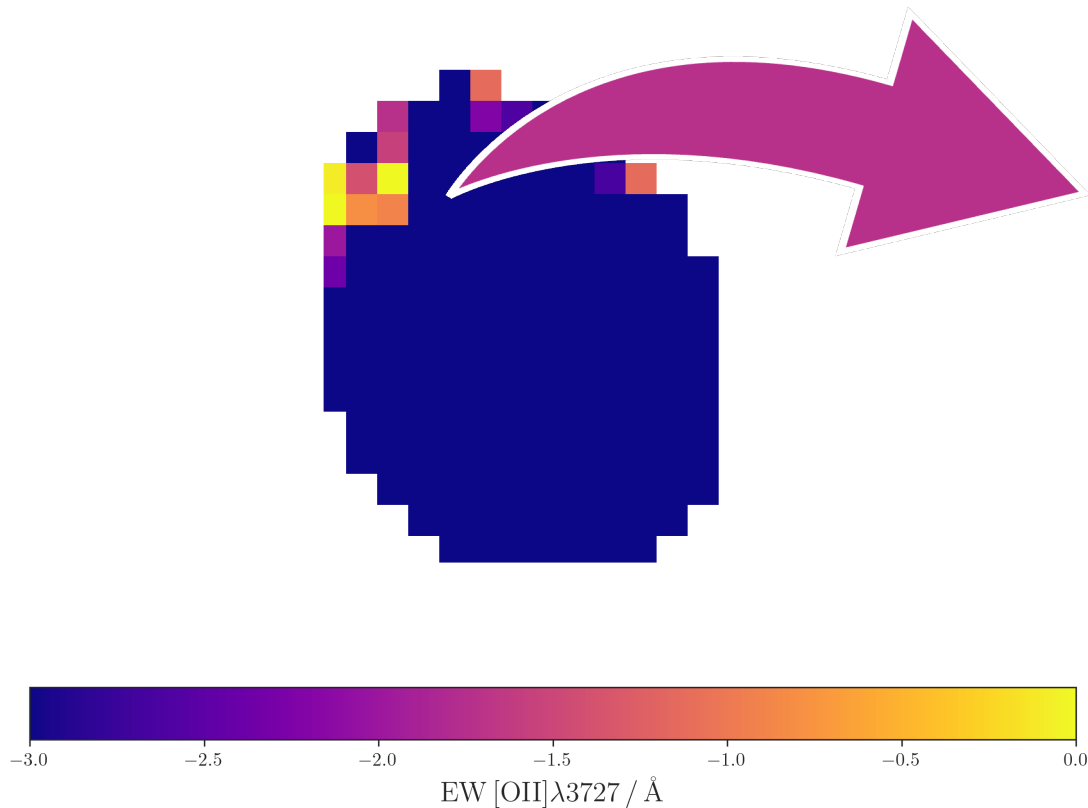
Edge-on case doesn't
quench!

Simulations by Nina
Akerman

Post-starburst features in jellyfish galaxies



The problem with traditional classification



Too much importance given to one spectral feature!

Method should consider the **entire spectrum** and be **probabilistic**!

Our proposed solution

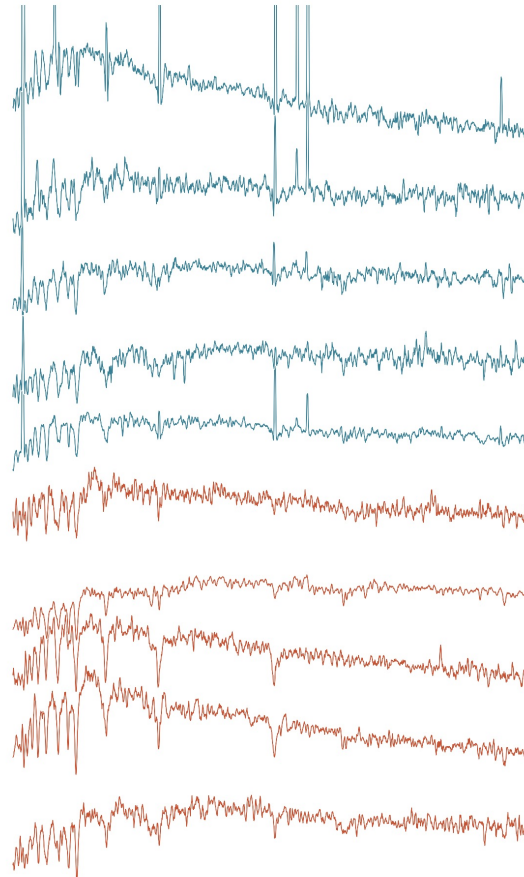


There are loads of **star-forming** and **quiescent** galaxies in our observations. We can use spaxel-by-spaxel spectra of these objects to build a training set to classify regions of jellyfish galaxies!

Our proposed solution

Training set obtained from **star-forming** and **quiescent** galaxies in the same MUSE data cubes (individual spaxels).

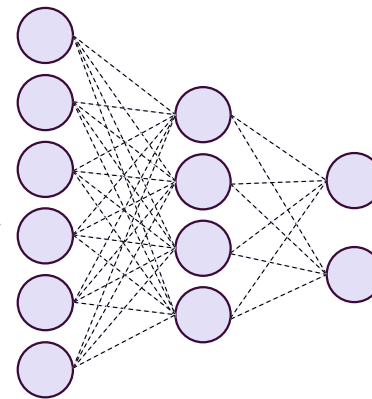
Normalized and **resframed** using the redshift of the galaxy!



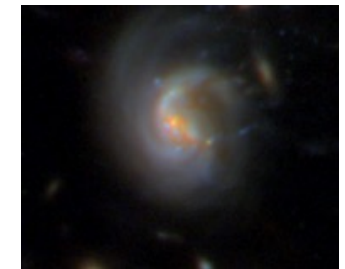
Star-Forming

Quiescent

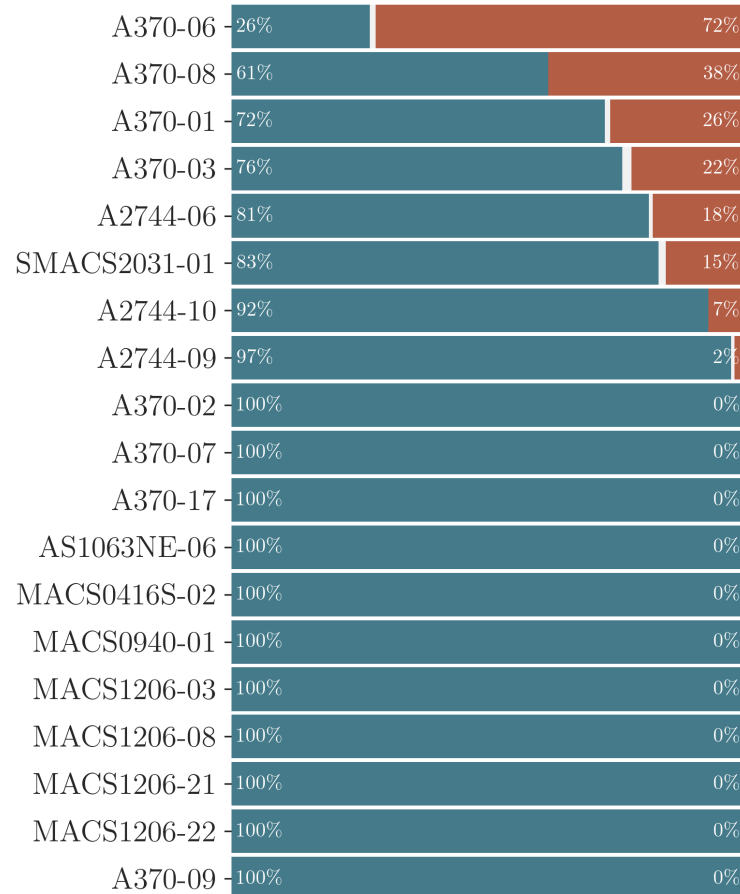
1D CNN classifier



Classify spaxels in a sample of 19 Jellyfish galaxies.



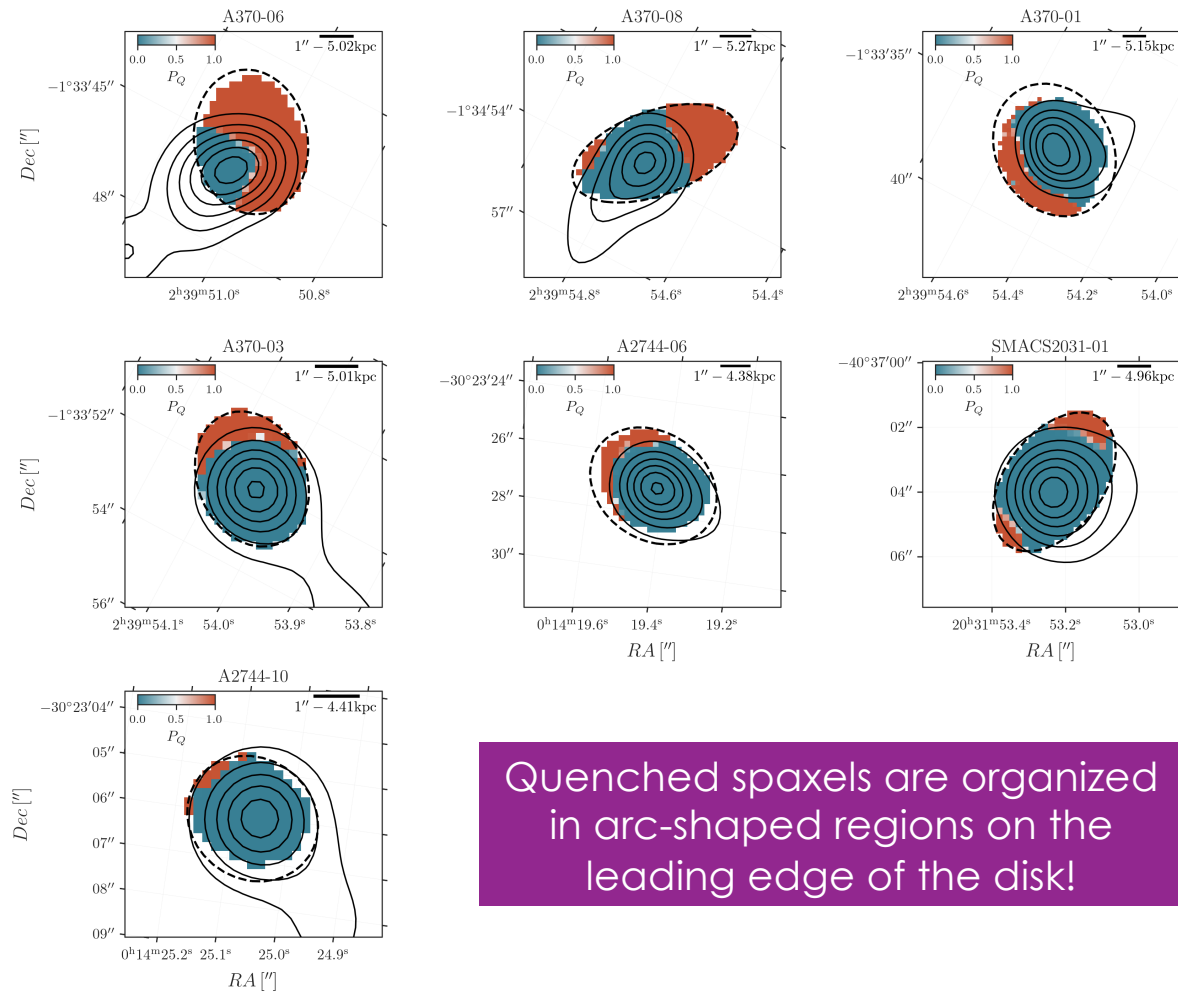
How much quenching do we find?



Seven galaxies with substantial quenching fraction! (Arguably)

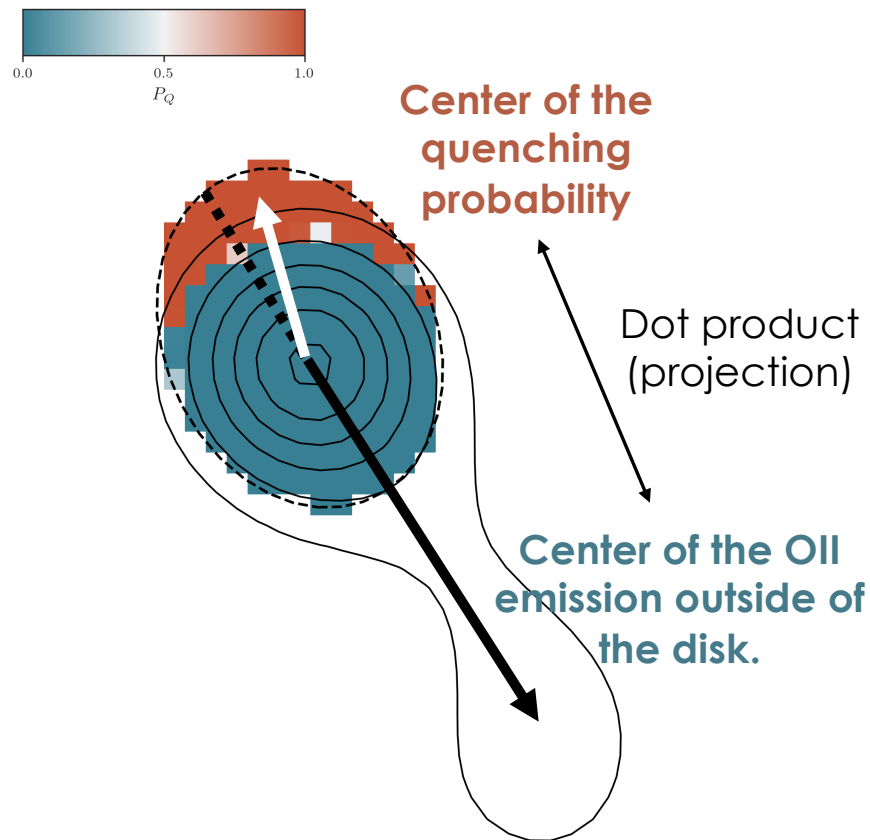
Eleven galaxies with no quenching signatures!

Morphology of quenched regions

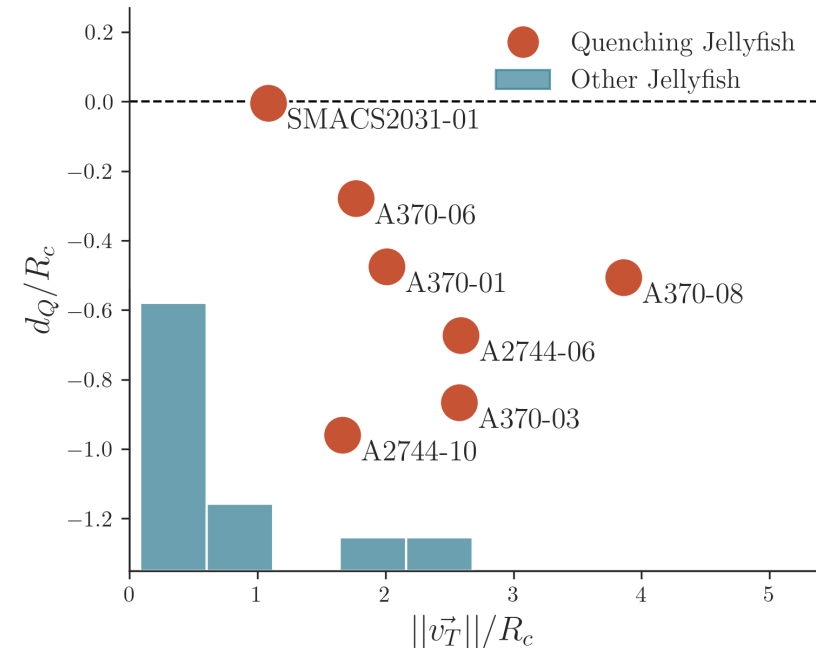


Quenched spaxels are organized in arc-shaped regions on the leading edge of the disk!

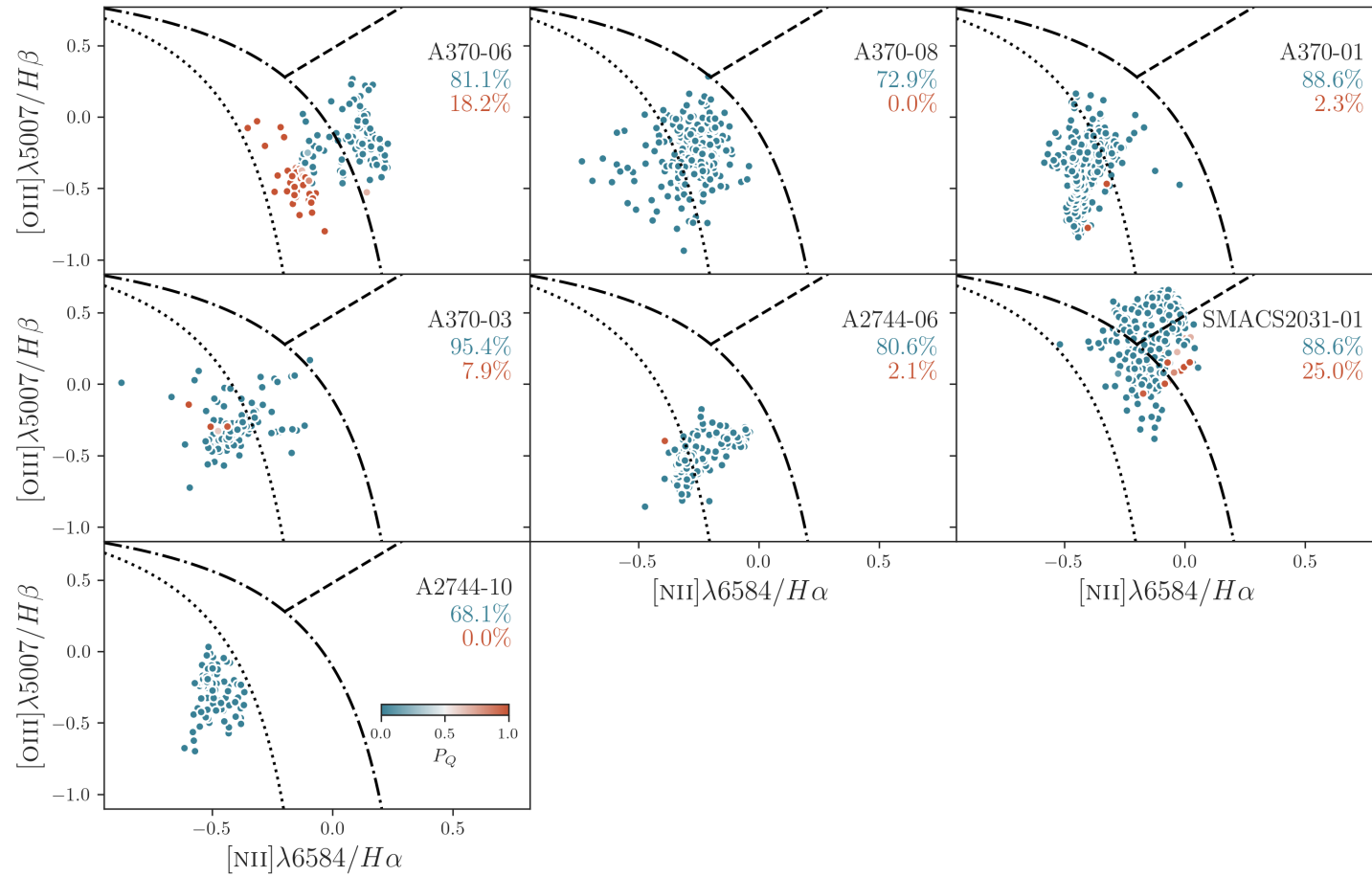
Morphology of quenched regions



Almost all galaxies with long tails ($>1.5 R_e$) are undergoing quenching.

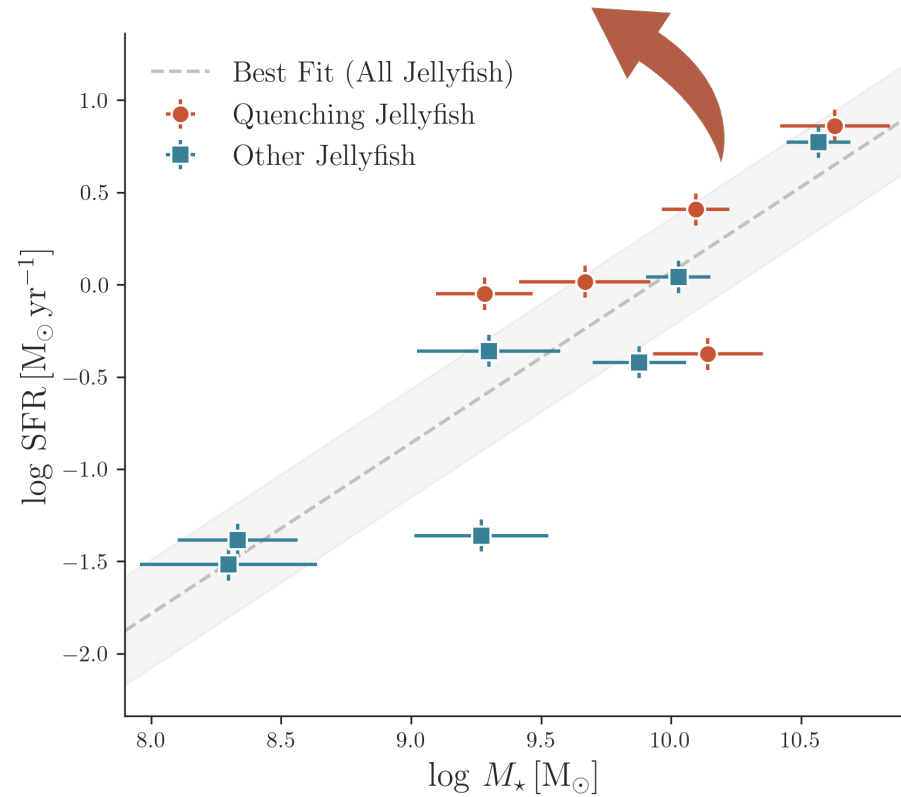


What about the remaining gas?



Star-formation rates

Global SFR enhancement



Main points

We are able to explain how the **quenching patterns** observed in cluster **post-starburst** galaxies take shape in the gas stripping phase.

- A **1D CNN** can reproduce the spectral **classification** expected from **visual inspection**.
- **Quenching** in Jellyfish galaxies happens in **arc-shaped** regions, being **outside-in and side-to-side** at the same time.
- While the outskirts quench, **star-formation** is **enhanced towards the centre**.