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PACE



What are cosmological simulations?

"... a disheartening number of ingredients must be assembled to produce a plausibly complete recipe for galaxy formation"

"What I cannot create, I do not understand"

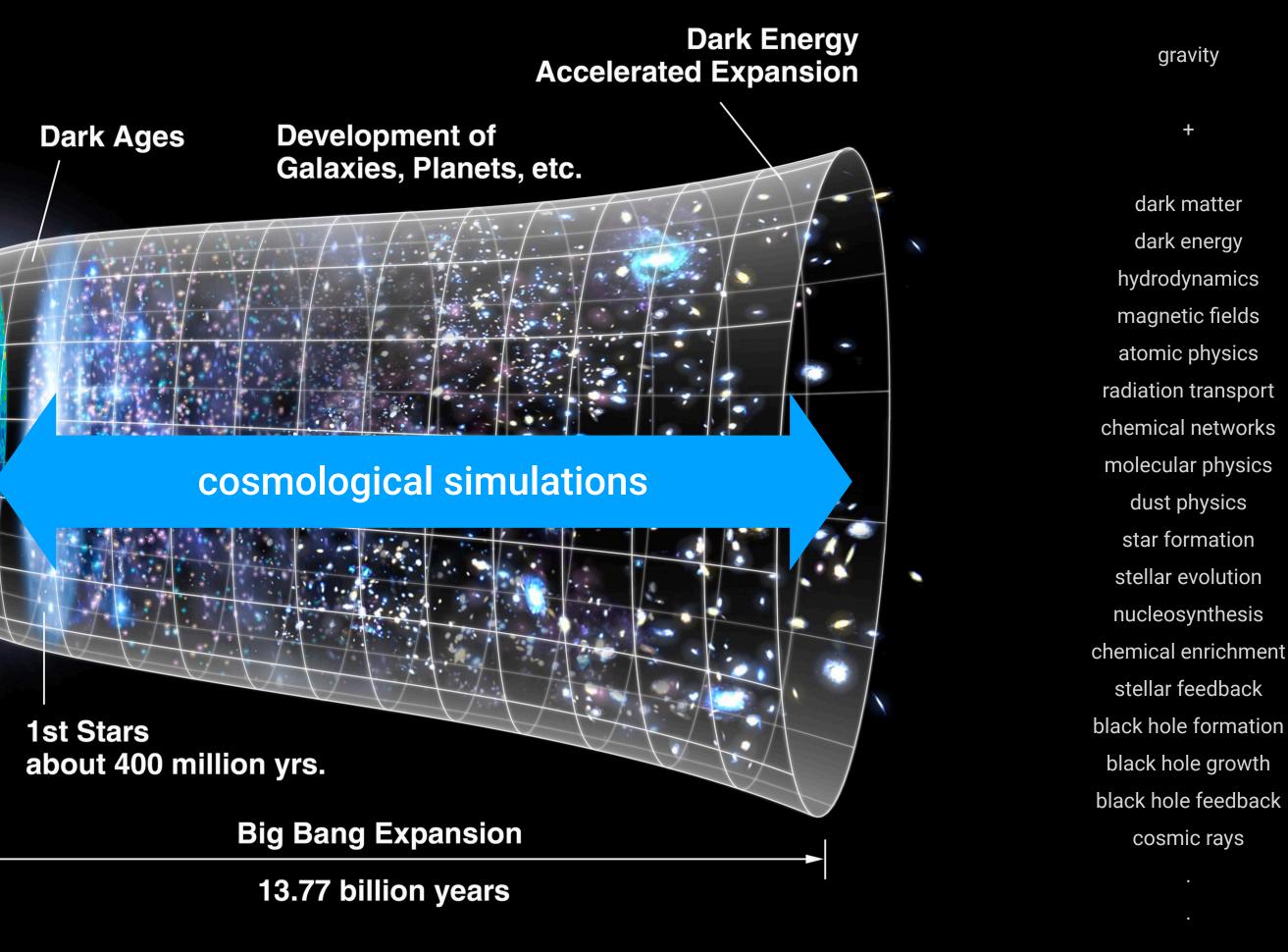
Richard Feynman

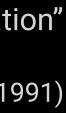
cosmological simulations attempt to build a Universe in a computer using the known laws of physics

Afterglow Light Pattern 375,000 yrs.

Inflation

Quantum **Fluctuations** White & Frenk (1991)

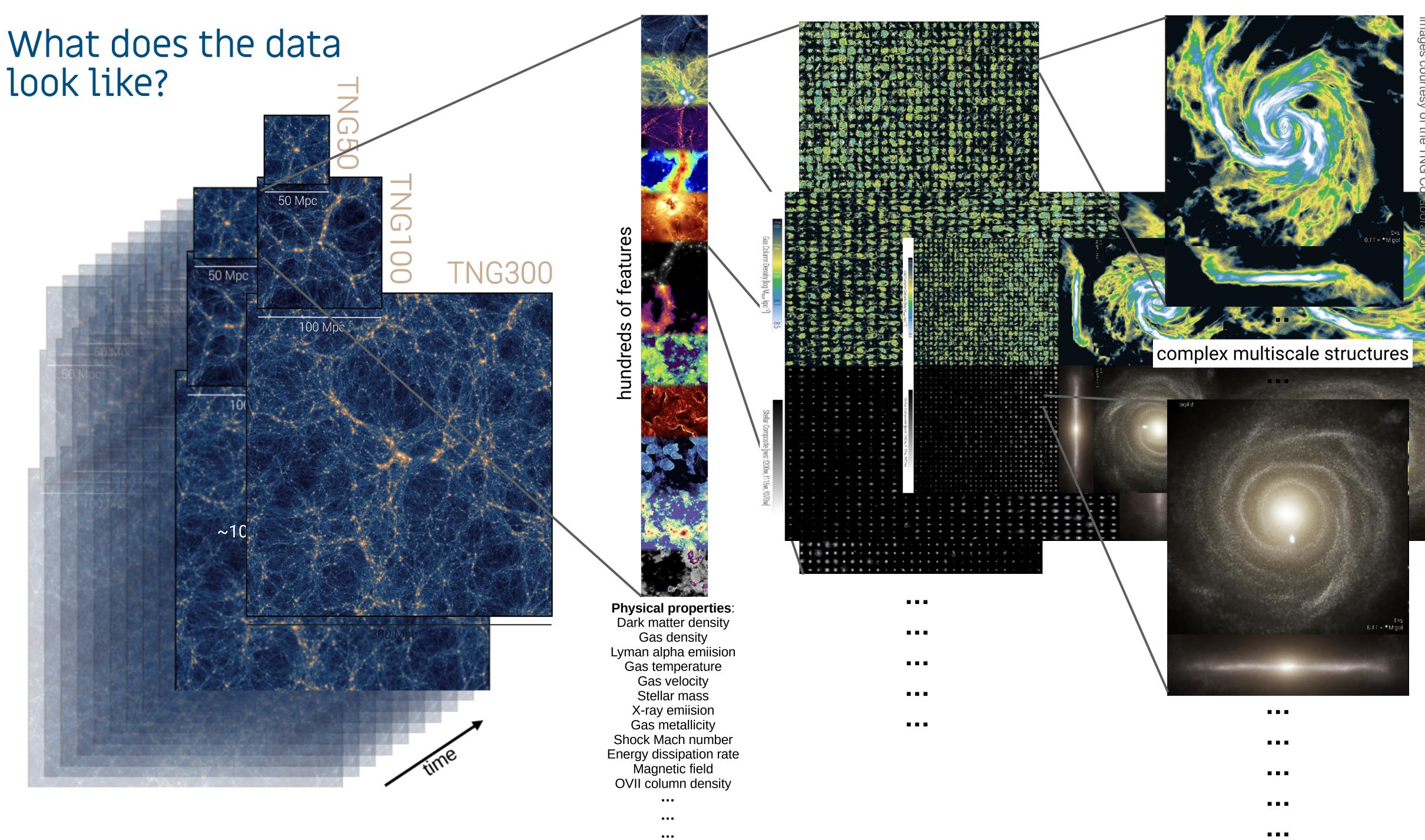


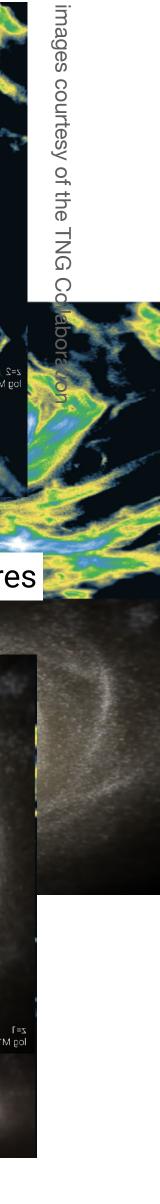


... after running for 2 years on a supercomputer (~100M CPU-hours)

A multi-scale challenge

The TNG Collaboration





If all the information is there, why are these simulations so difficult to interpret?

... let's focus on a single object

The TNG Collaboration

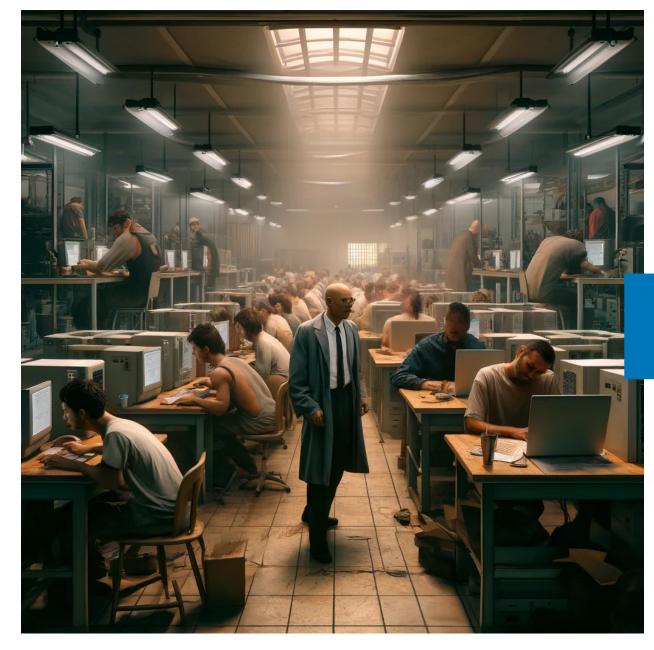
The data volume and complexity problem

Today's largest projects evolve a universe in a box:

- ~100 billion resolution elements
- in a >100-dimensional feature space
- in millions objects (galaxies)
- ~1 petabyte of data (petascale)

a feature space with $> 10^{12}$ dimensions

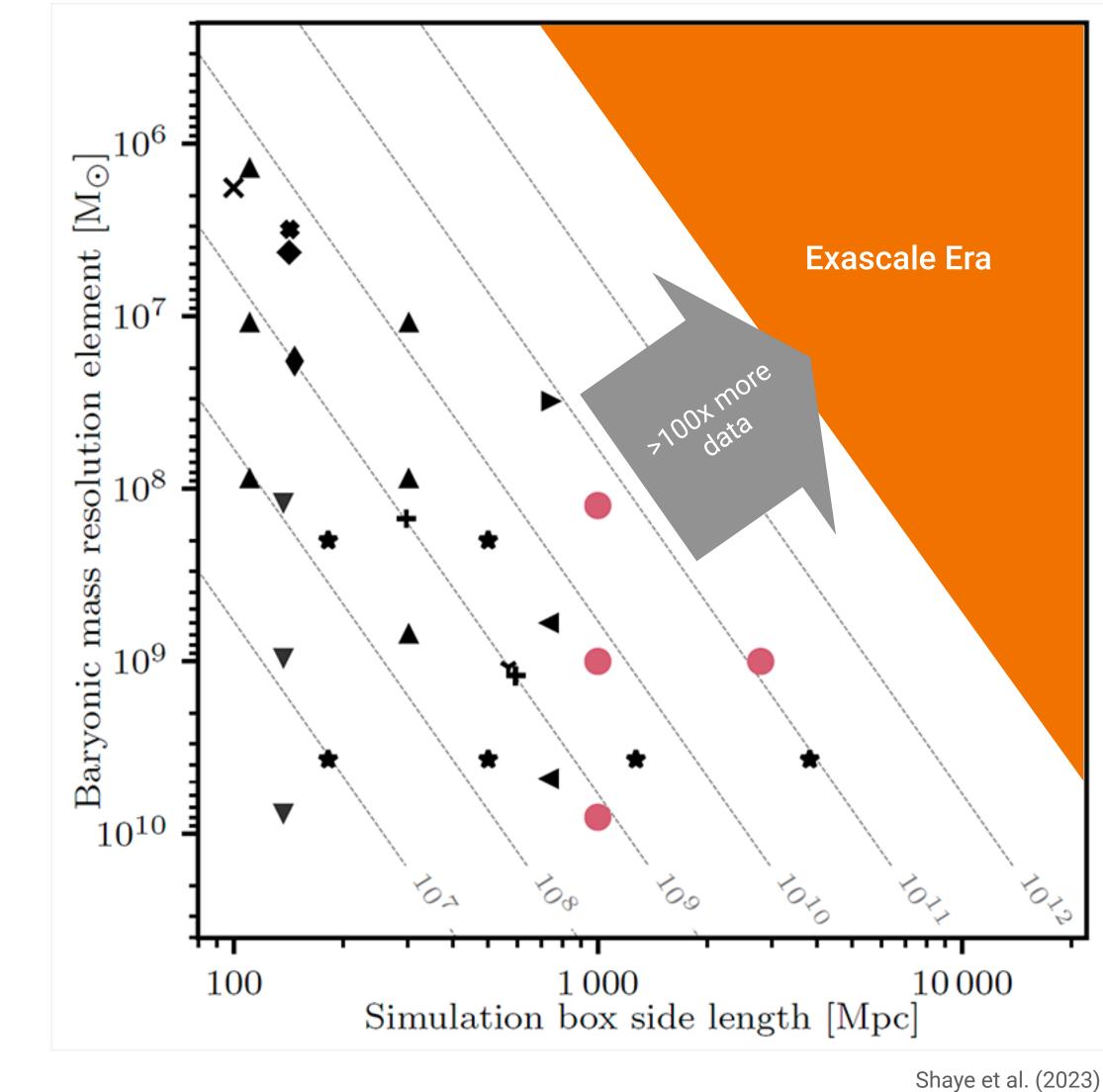
Exascale simulations: >100 times larger datasets -> ML essential





LUMI, Europe's largest supercomputer

Hard-working students at a scientific sweatshop as imagined by DALL-E



The largest cosmological simulations

How can we approach this challenge?

The ideal tool should

- learn a simple representation of all structures in the simulation without labels 1.
- enable exploration and interpretation using interactive visualization for arbitrarily 3. large datasets
- be code-agnostic (can be used on *any simulation* without expertise) 4.

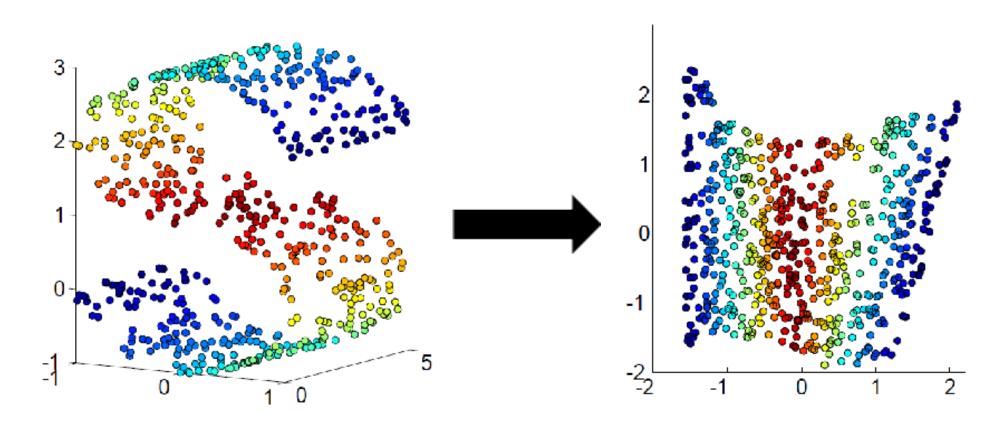


HAL9000 from 2001: A Space Odyssey

The power of Generative Deep Learning

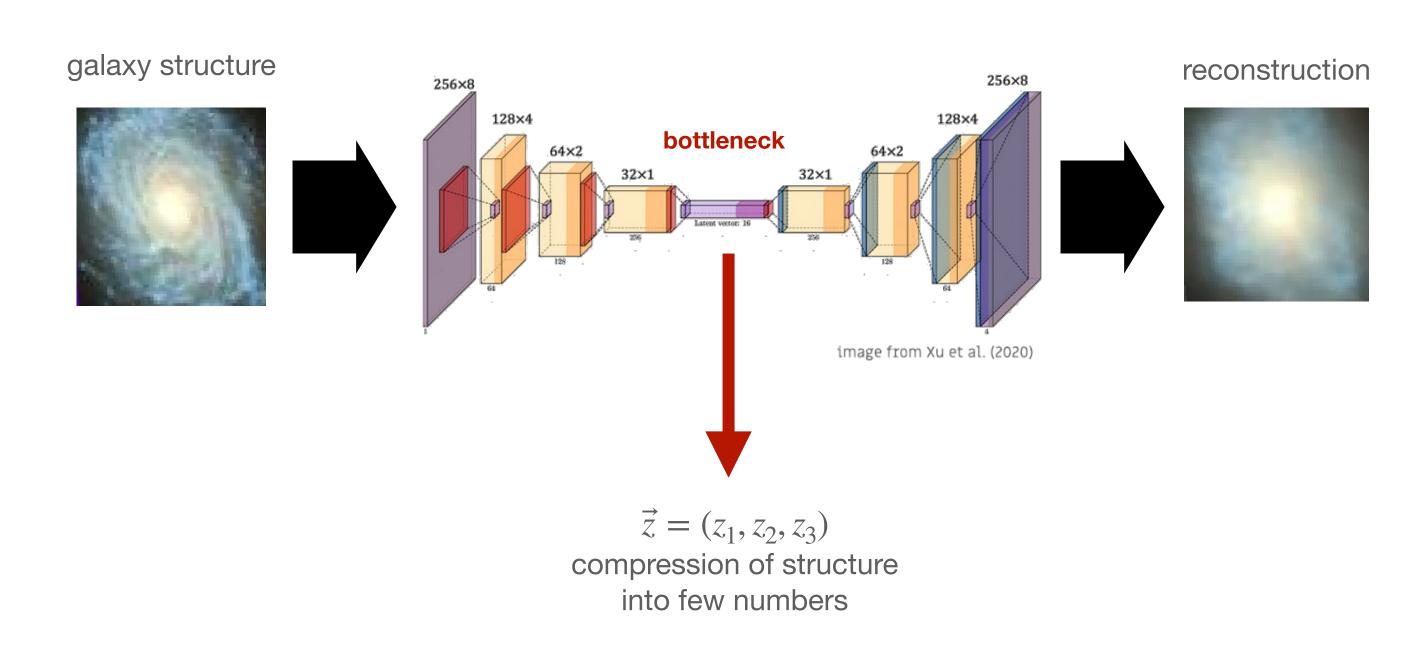
"What I cannot create, I do not understand"

- Learns underlying distribution of dataset from samples without labels
- Projects data onto low-dimensional space, providing a powerful way to explore and interpret it



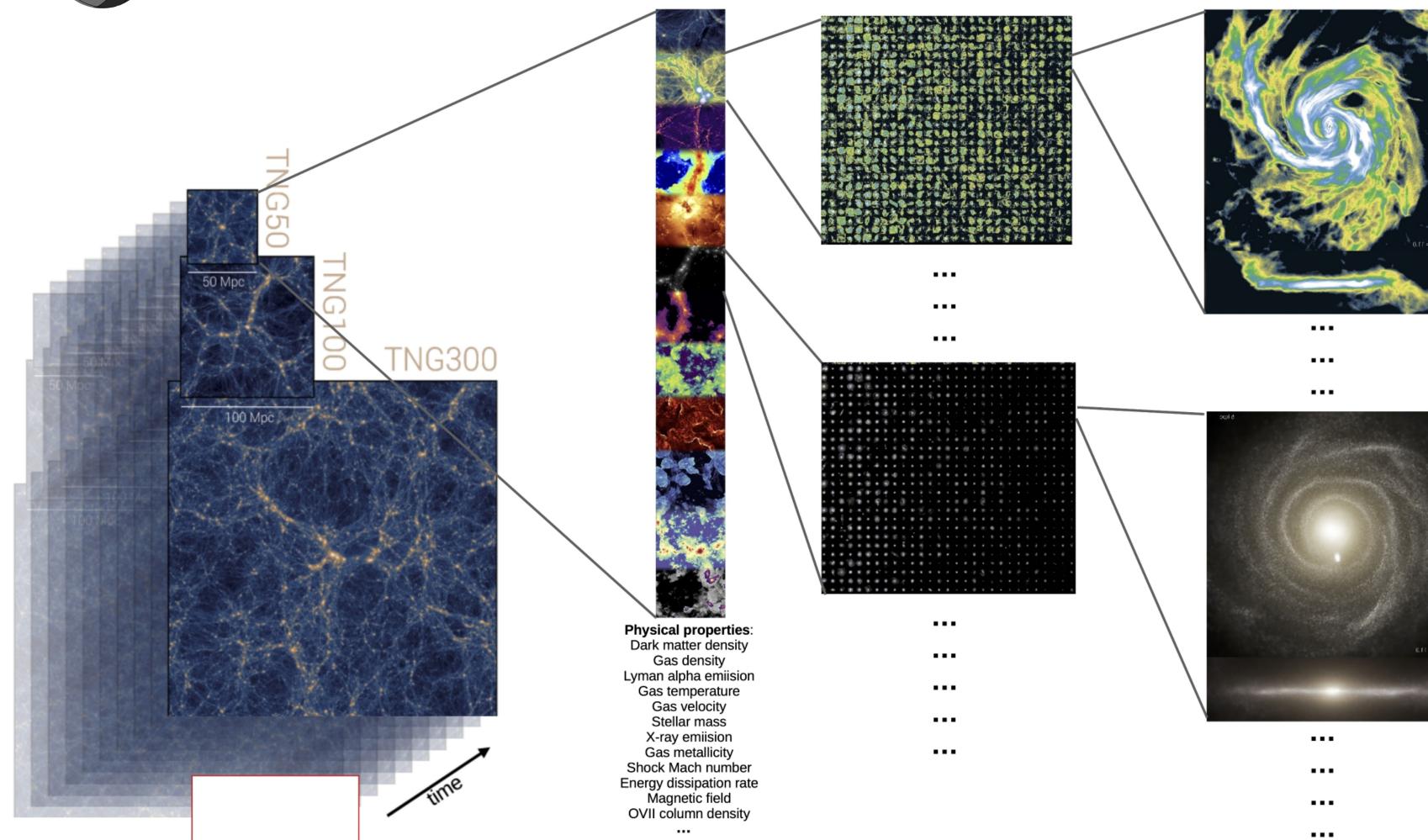
data in physical space

data in intrinsic low-dimensional space Richard Feynman









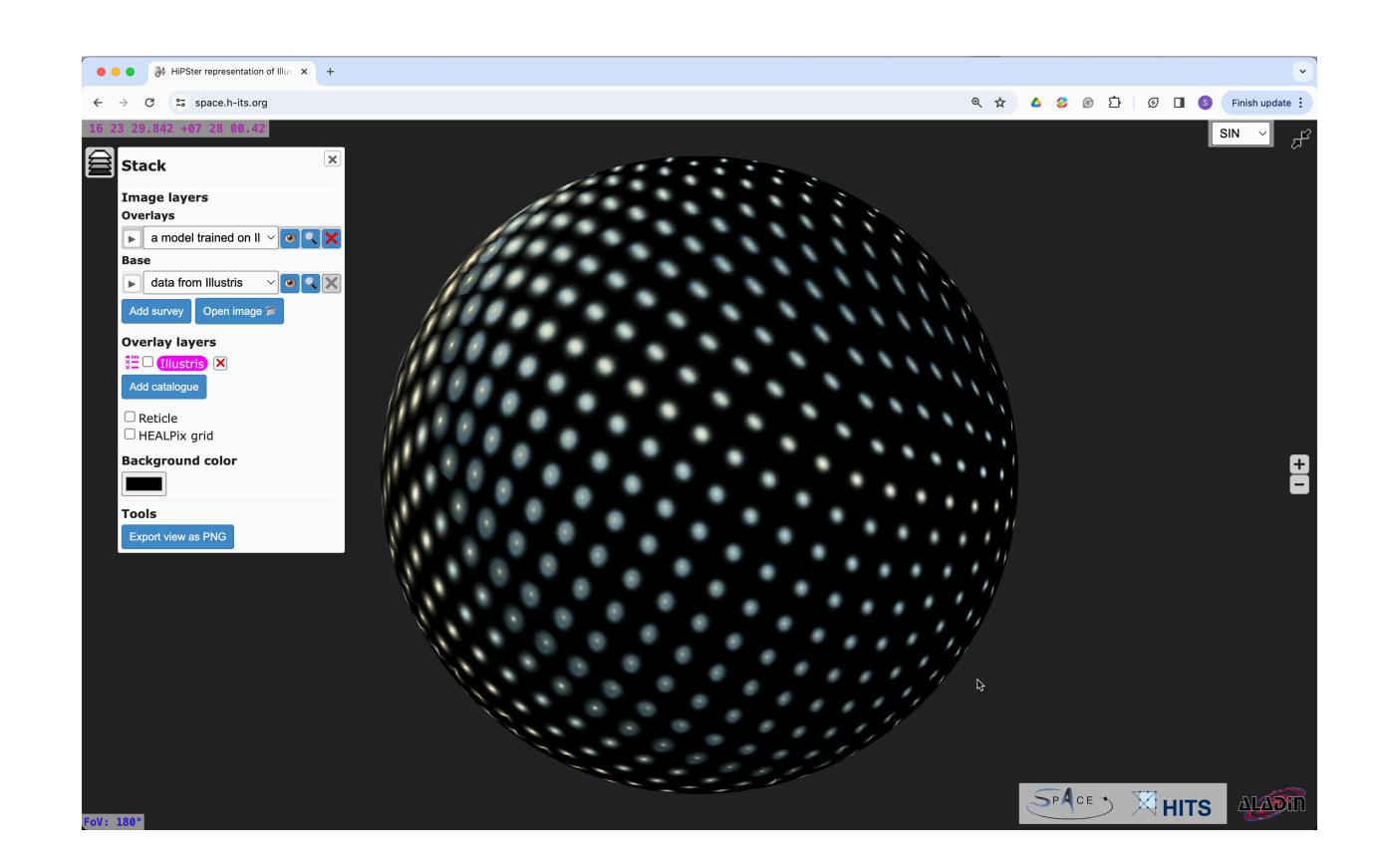
- ...
- •••

ages courtesy of the TNG Collaboratio

...

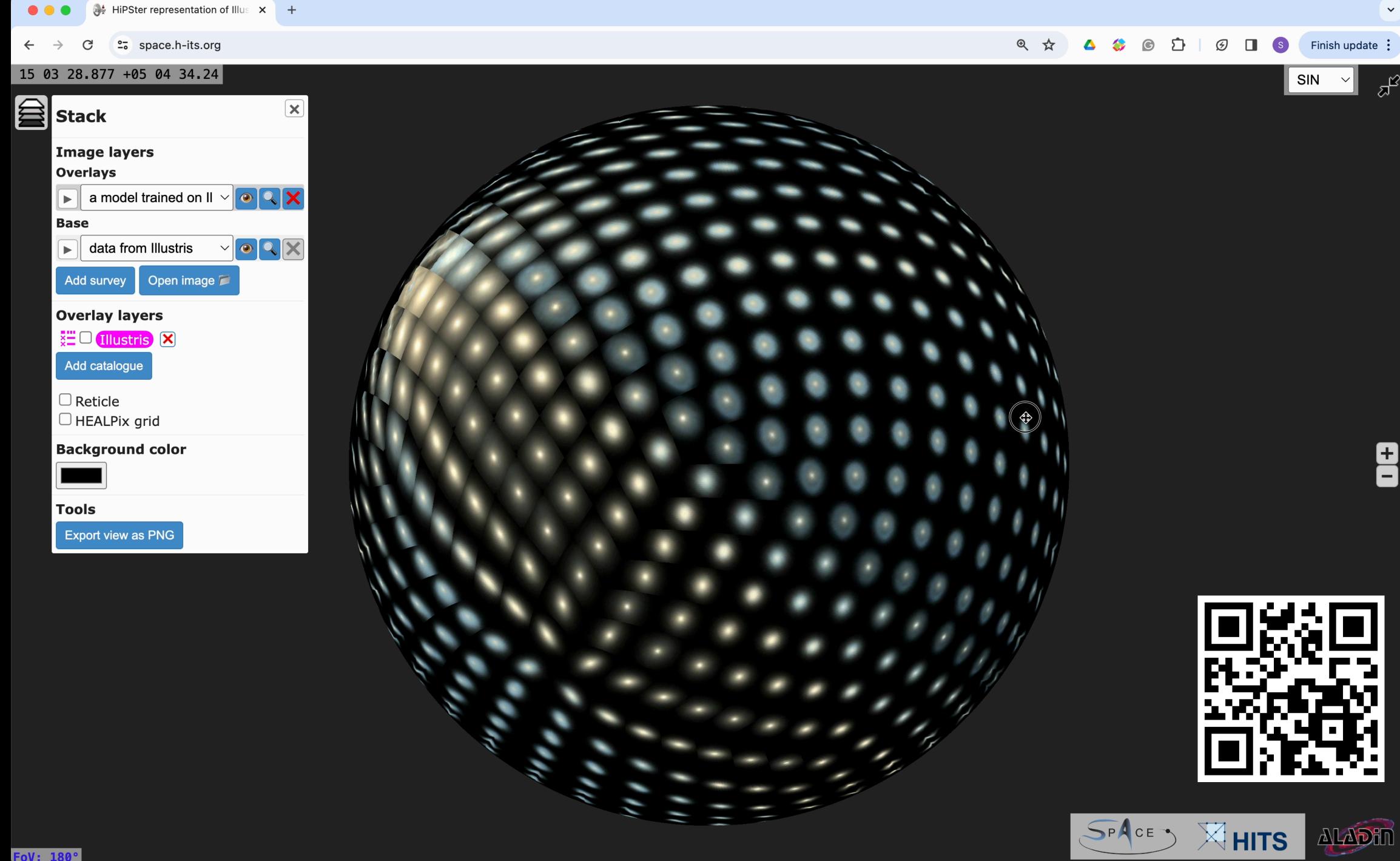
First results

- Trained prototype on •
 - ~51,000 synthetic galaxy images from the IllustrisTNG simulations (Nelson et al. 2019; Rodriguez-Gomez et al. 2019), and
 - ~150,000 real galaxy images from the GalaxyZoo (Lintott et al. 2008)
- Learned the high-level galaxy features (i.e. color, size, morphology, inclination, and interactions) ●



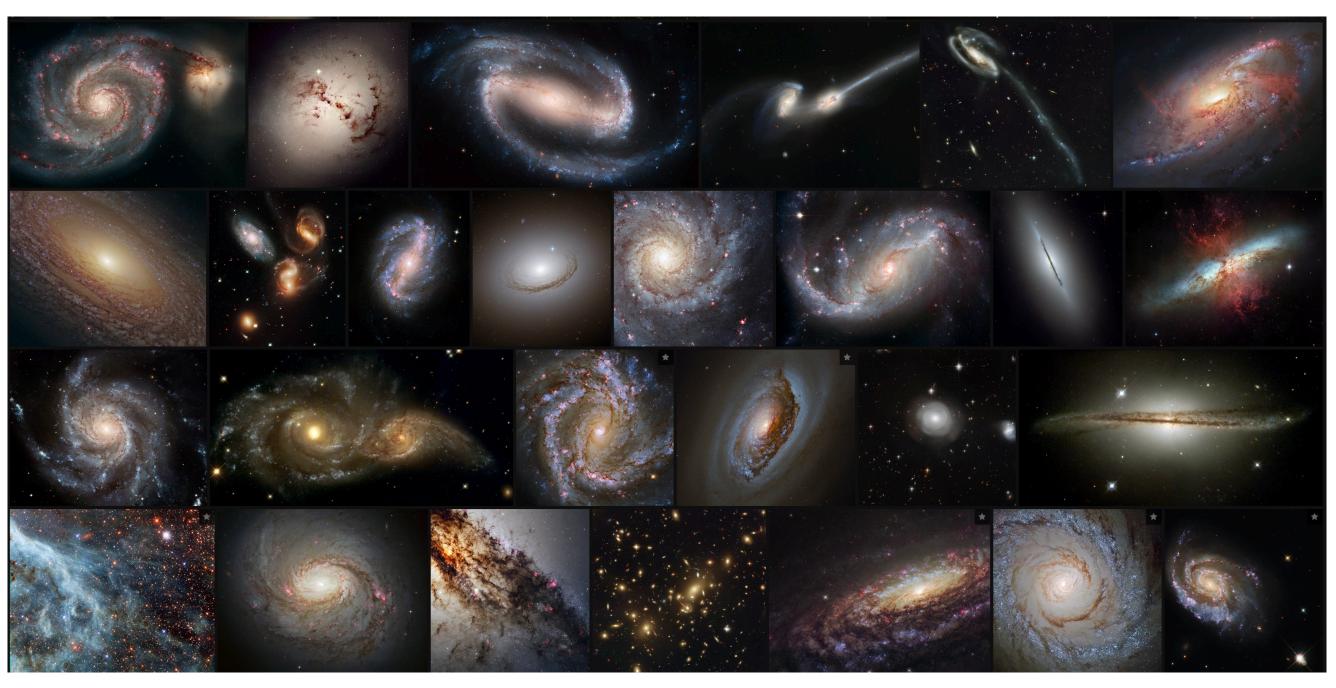
Rule #3: Interpret and visualize models







- Astronomers describe galaxies using single numbers based on intuition (e.g. luminosity, size, shape, etc.)
- But galaxies ARE NOT points!
 - they are complex objects emerging from many physical processes

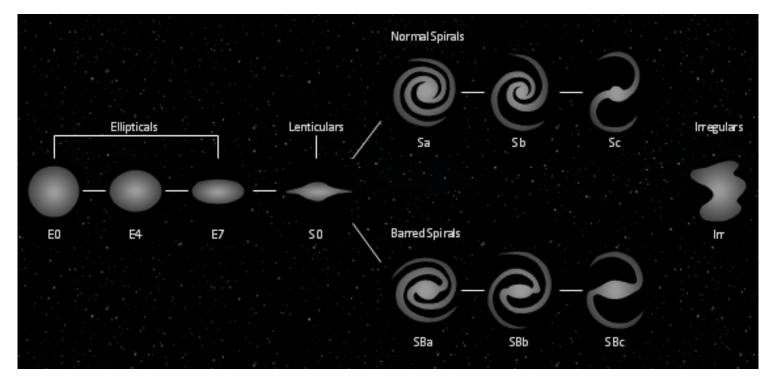


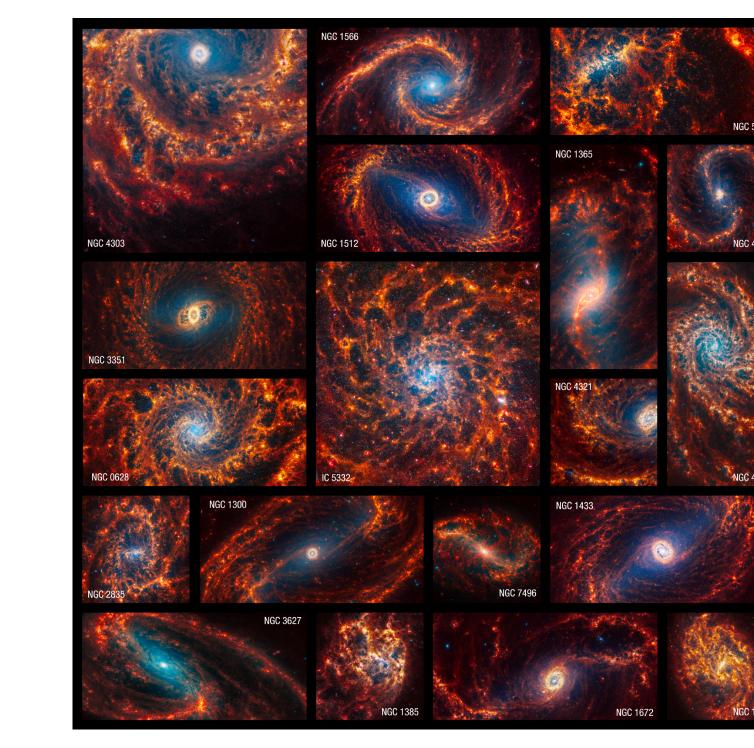
Rule **#1**: Compare to domain reference and broader context





How we have described galaxies for the last 100 years



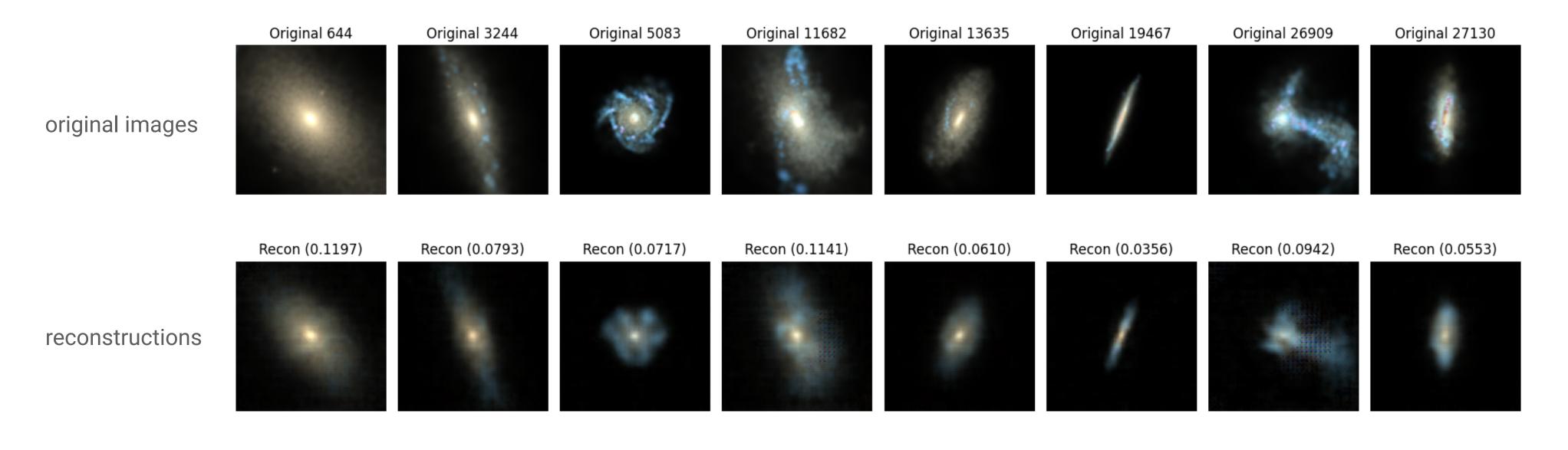


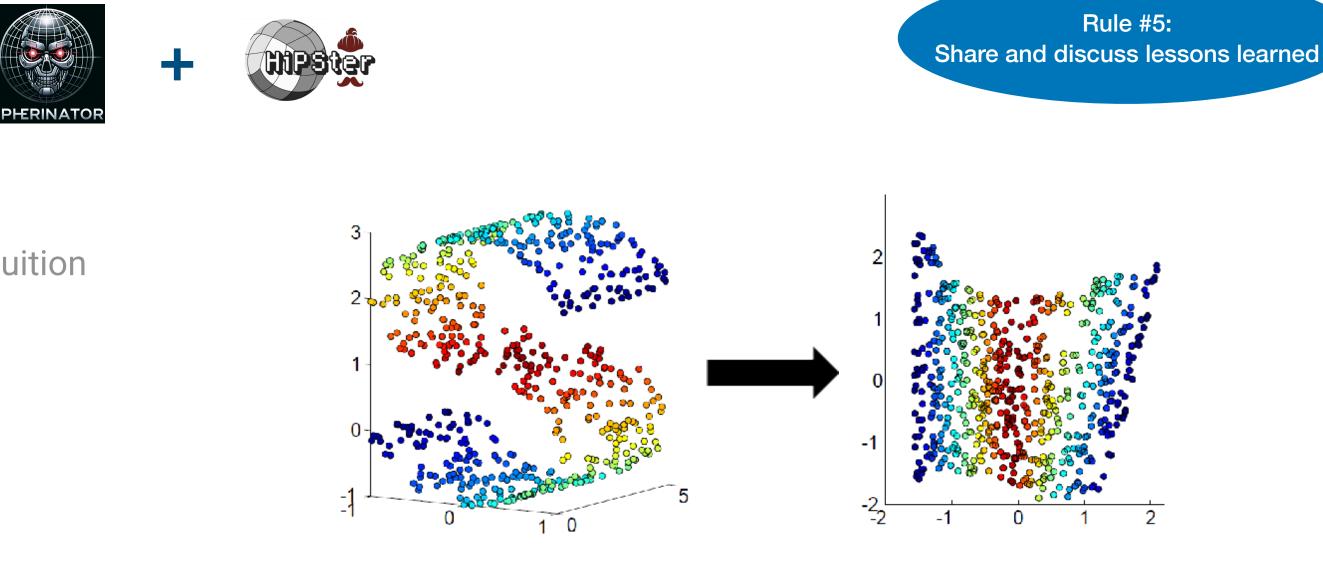
Real galaxies as seen by HST ...

... and JWST



- Astronomers describe galaxies using single numbers based on their intuition (e.g. luminosity, size, shape, etc.)
- But galaxies ARE NOT points!
 - they are complex objects emerging from many physical processes
- Ask the machine: how many numbers are needed to describe a galaxy?





> 10,000-dimensional pixel space

32-dimensional manifold

32 dimensions are enough to capture galaxy structure (at \sim 2% level)



More applications...

- Exploration 1.
 - Correlations between all physical properties of \bullet structures
 - Interpret using known labels (e.g. galaxy morphology ۲ or environment)
- Simulation Based Inference + Model selection 2.
 - Galaxies are chaotic: cannot compare objects directly! \bullet
 - Instead compare structural distributions non-• parametrically
 - Likelihood without need for summary statistics ●

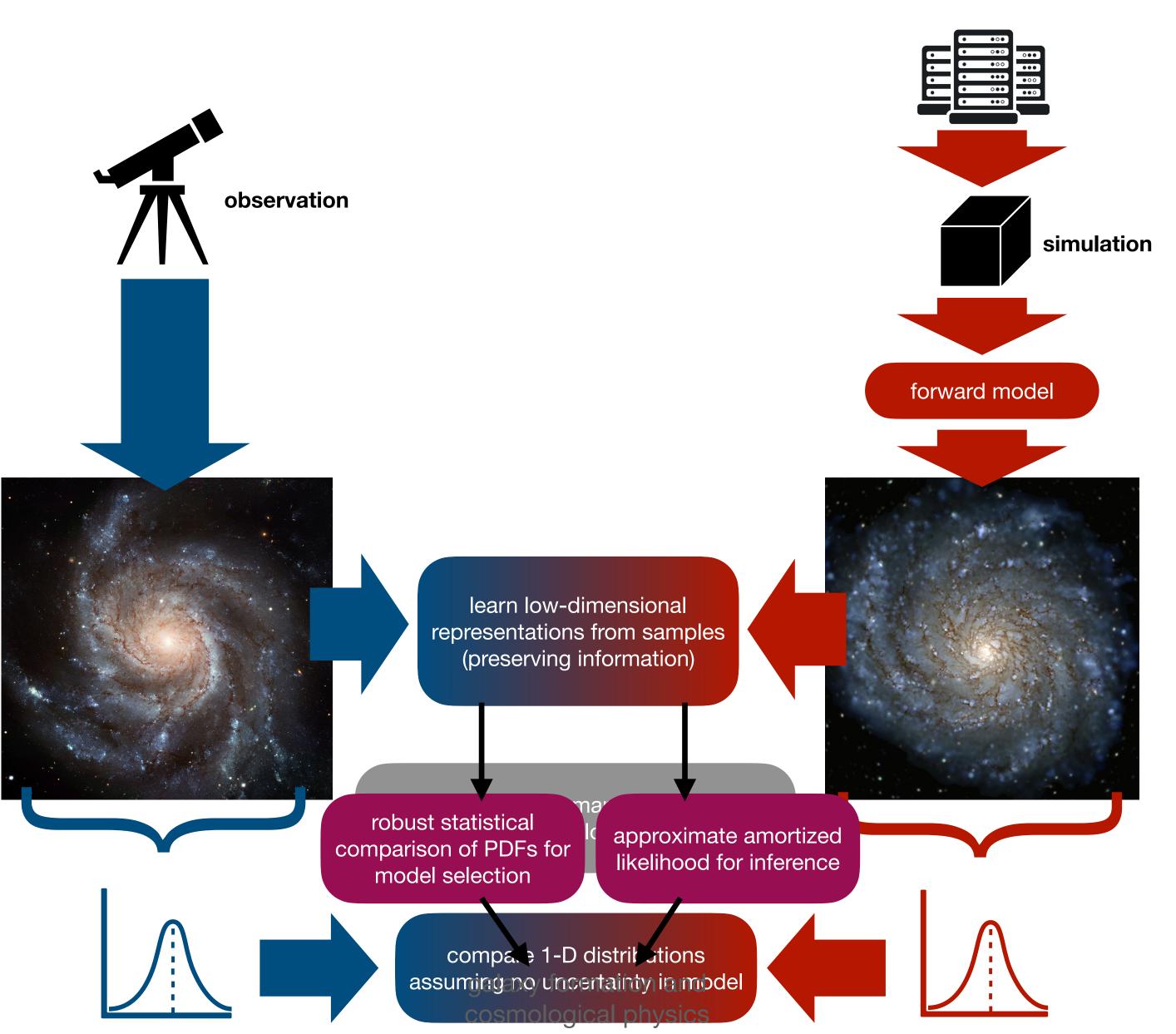






simulation or real?









Try it!

- arXiv:2406.03810 (Polsterer, Doser, Fehlner & Trujillo-Gomez 2023, in press)
- github.com/HITS-AIN/Spherinator

Can we help you explore and interpret your simulations and/or data? sebastian.trujillogomez@h-its.org

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Rule #2: Adopt best practices from ML community

Rule #6: Publish software and data

















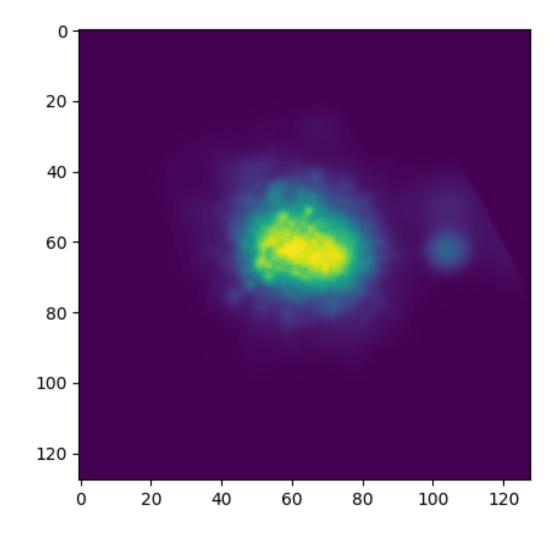


What's next?

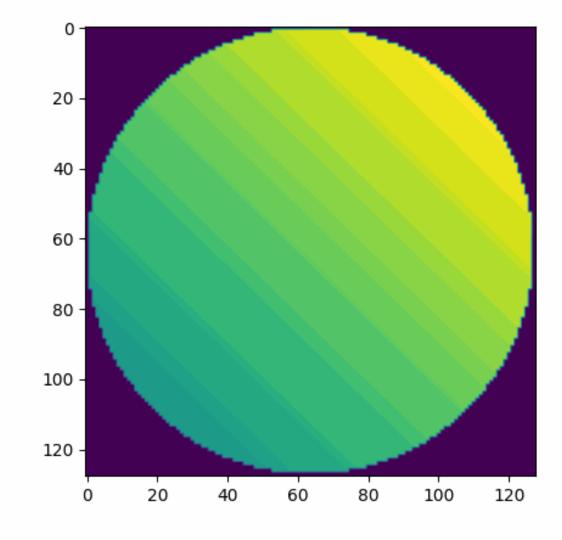
- 1. Physical symmetries: rotational equivariance
- 2. Beyond images to full 3D structures:
 - described using Geometric Deep Learning
 - explore 3D structures interactively



original galaxy image



equivariant representation using Zernike polynomials

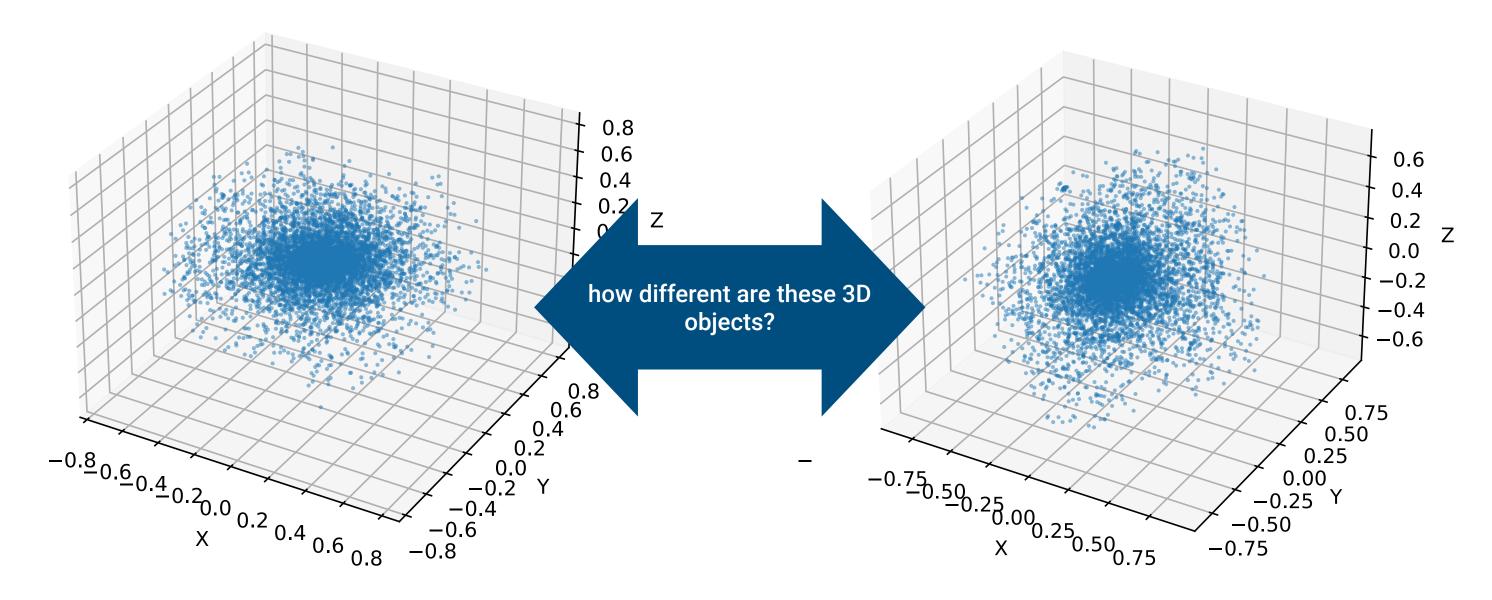




Chazotte et al. (in prep)

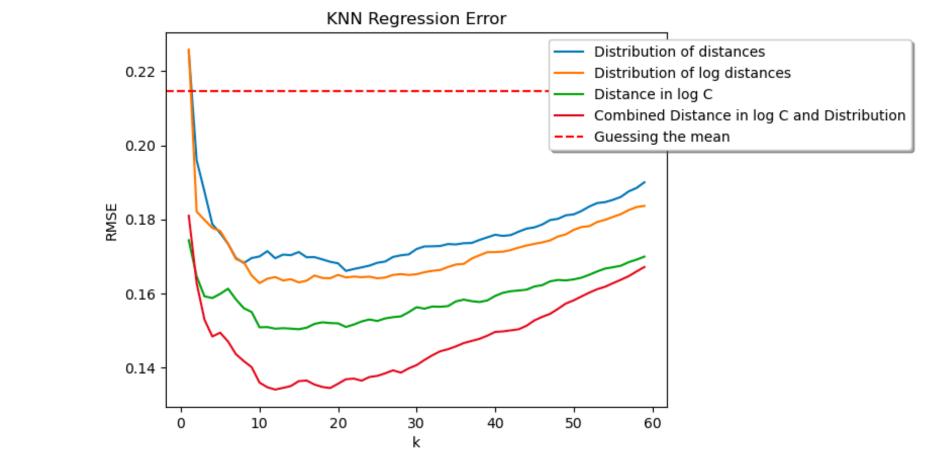


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The dark halos of two simulated galaxies





Shape of the galaxy halo predicts stellar mass (Packer, Villar, Hogg, STG in prep.)

