

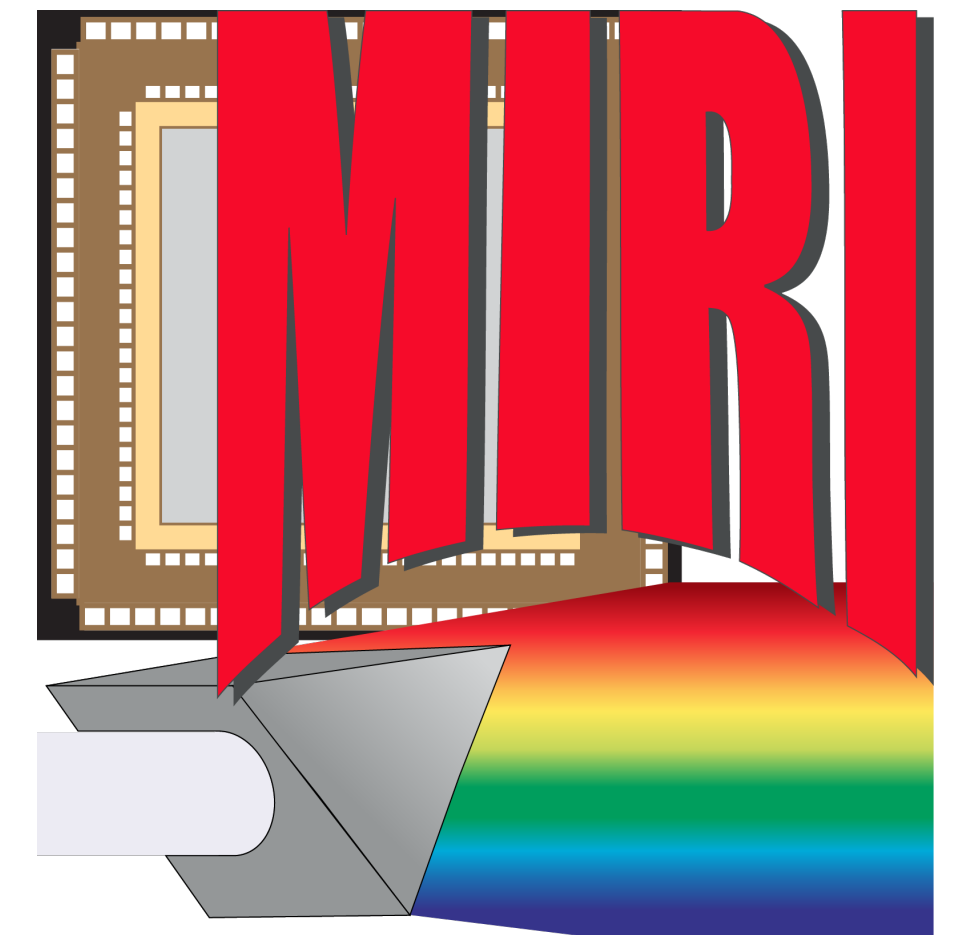
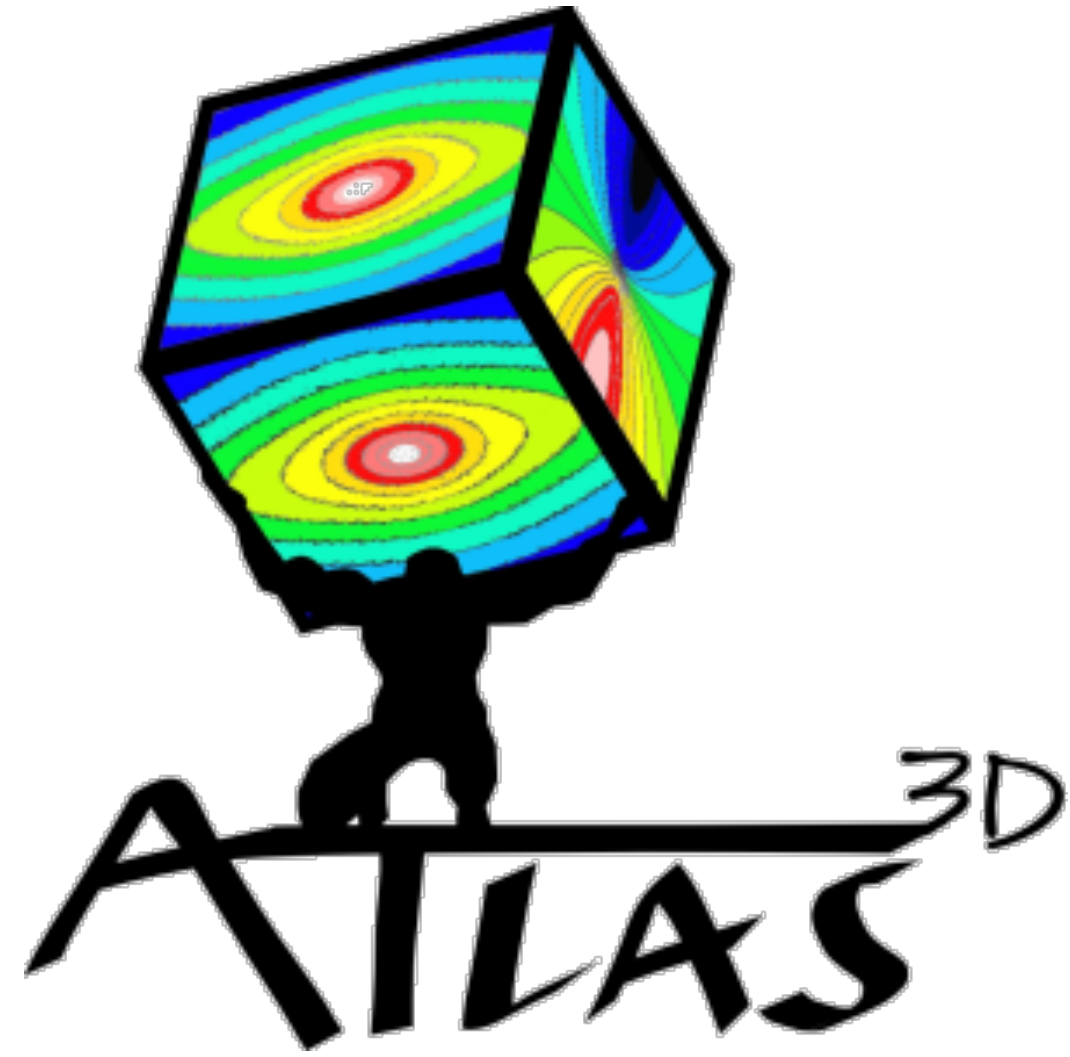
Machine Learning Dynamical Models: Applications to Simulations and JAM

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Integral Field Spectroscopy



Key Questions

- Can Machine Learning be used to Improve Existing Methods?
- Can Machine Learning be used to Create New Methods?

Dynamical Modelling Methods

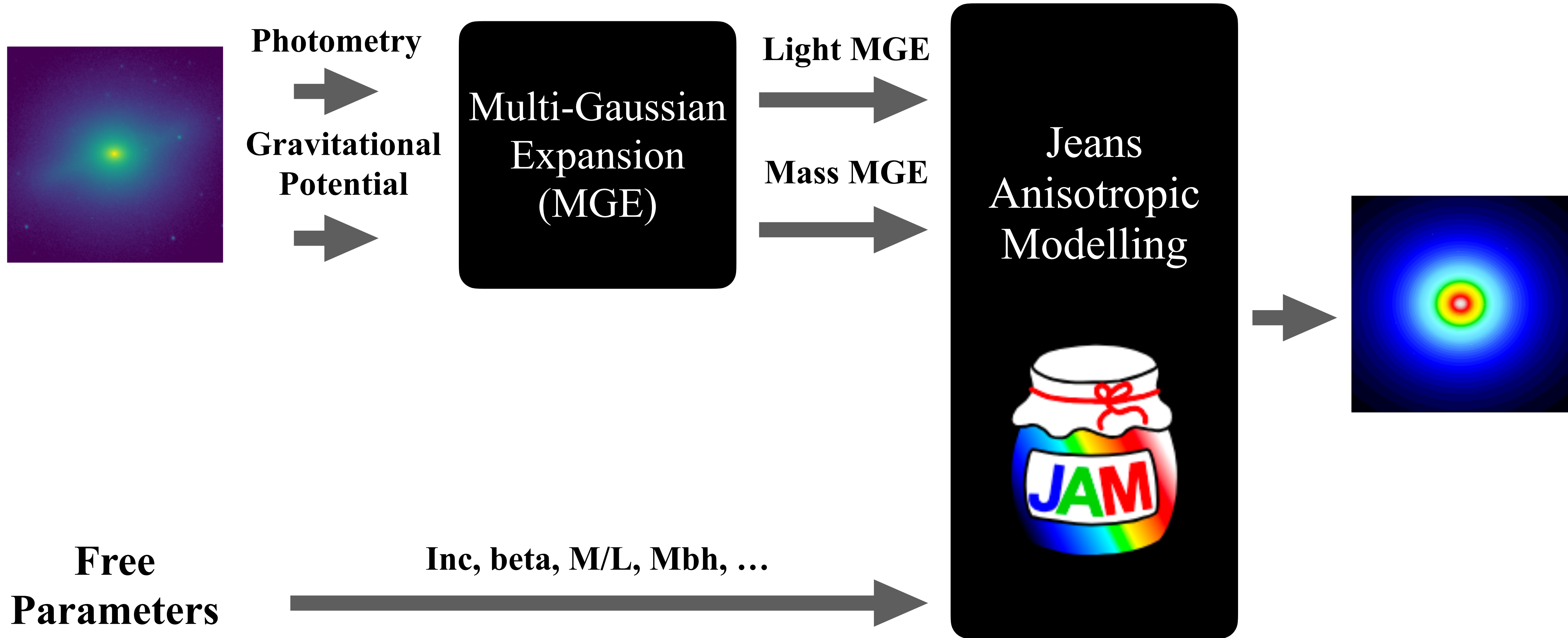
Schwarzschild Modelling

- Axisymmetric or Triaxial
- No anisotropy assumptions
- Uses full LOSVD
- Slow run time ~ 10 hours
- Non-smooth χ^2
- Lots of hyper-parameters

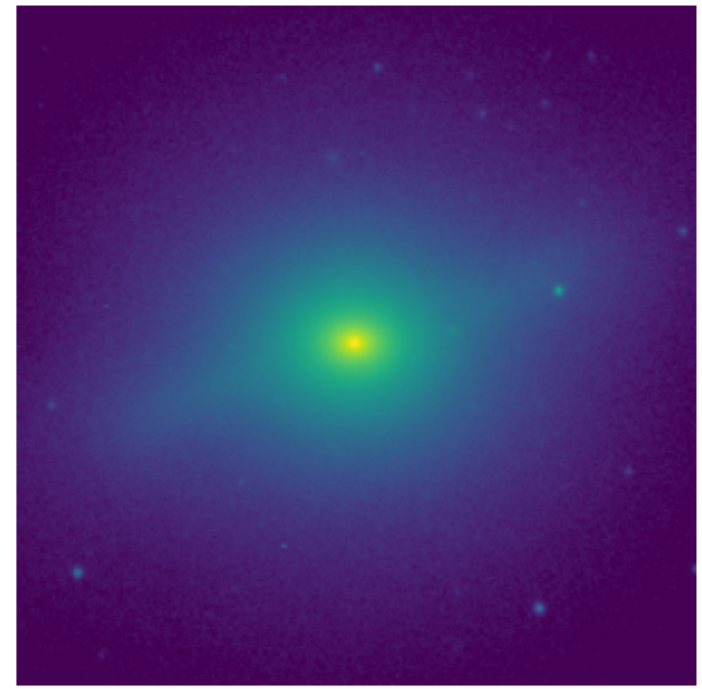
Jeans Modelling (JAM)

- Axisymmetric
- Cylindrically/spherically aligned velocity ellipsoid
- Uses 2nd moment
- Fast run time ~ 10 sec
- Smooth χ^2
- Few hyper-parameters

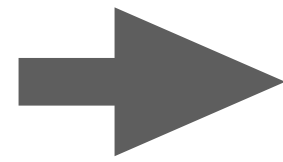
JAM Modelling Pipeline



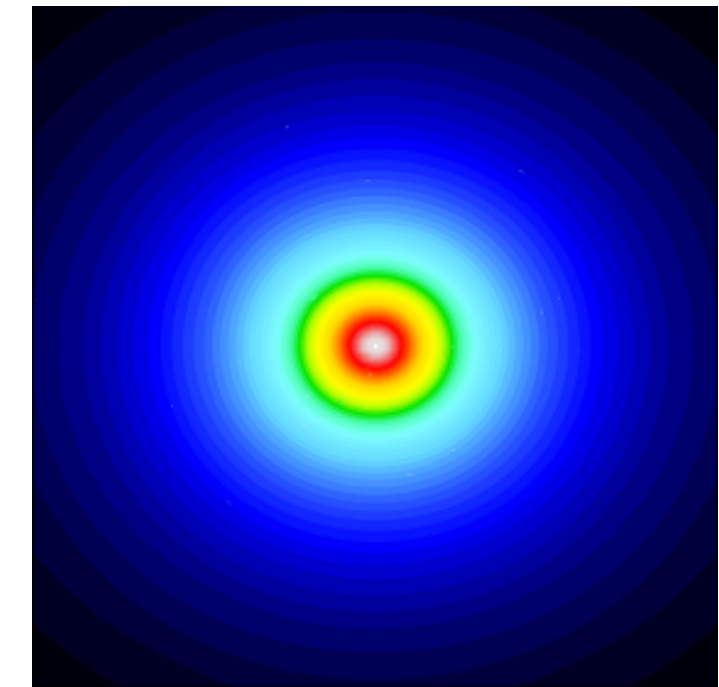
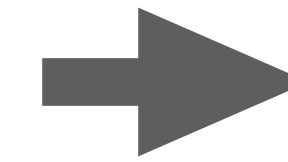
ML Modelling Pipeline



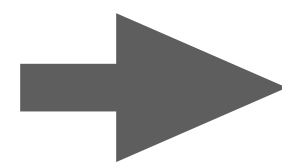
Photometry
→
**Gravitational
Potential**



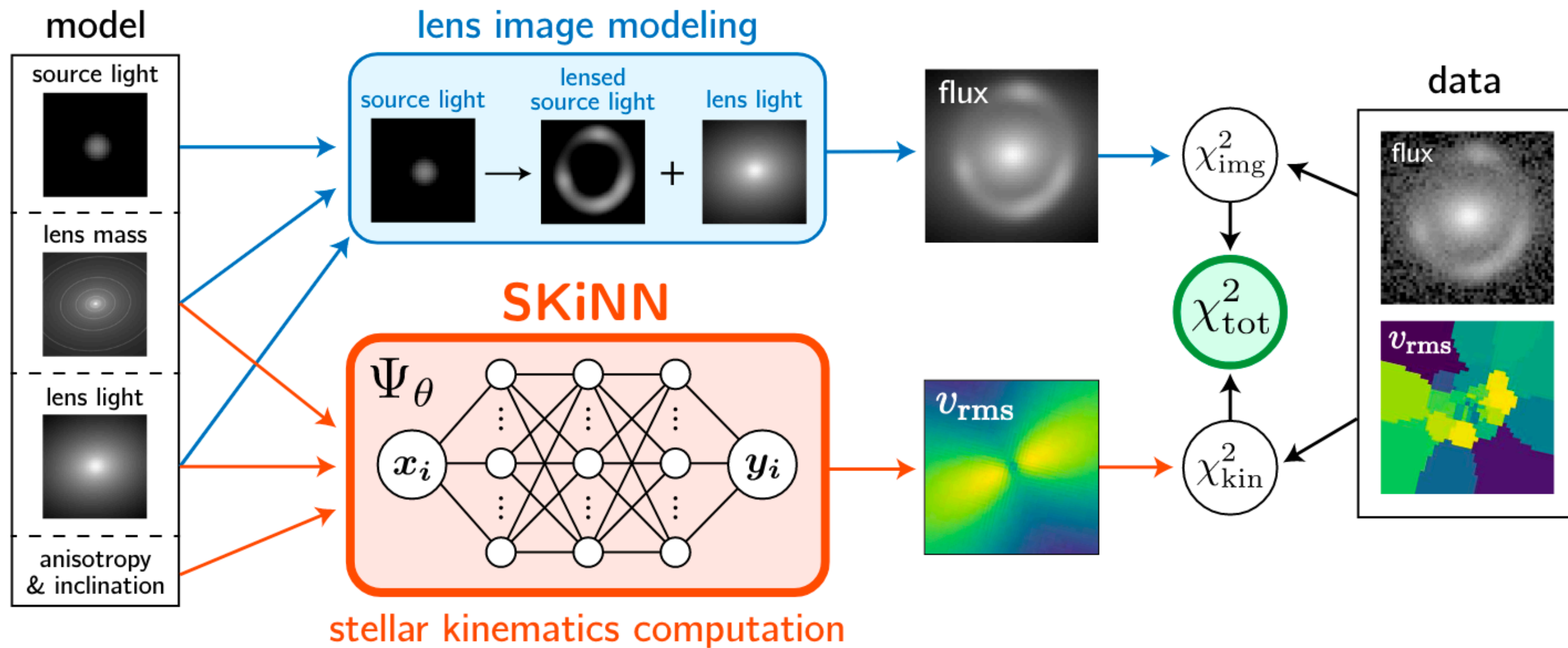
**Machine Learning
Algorithm**



**Free
Parameters** →



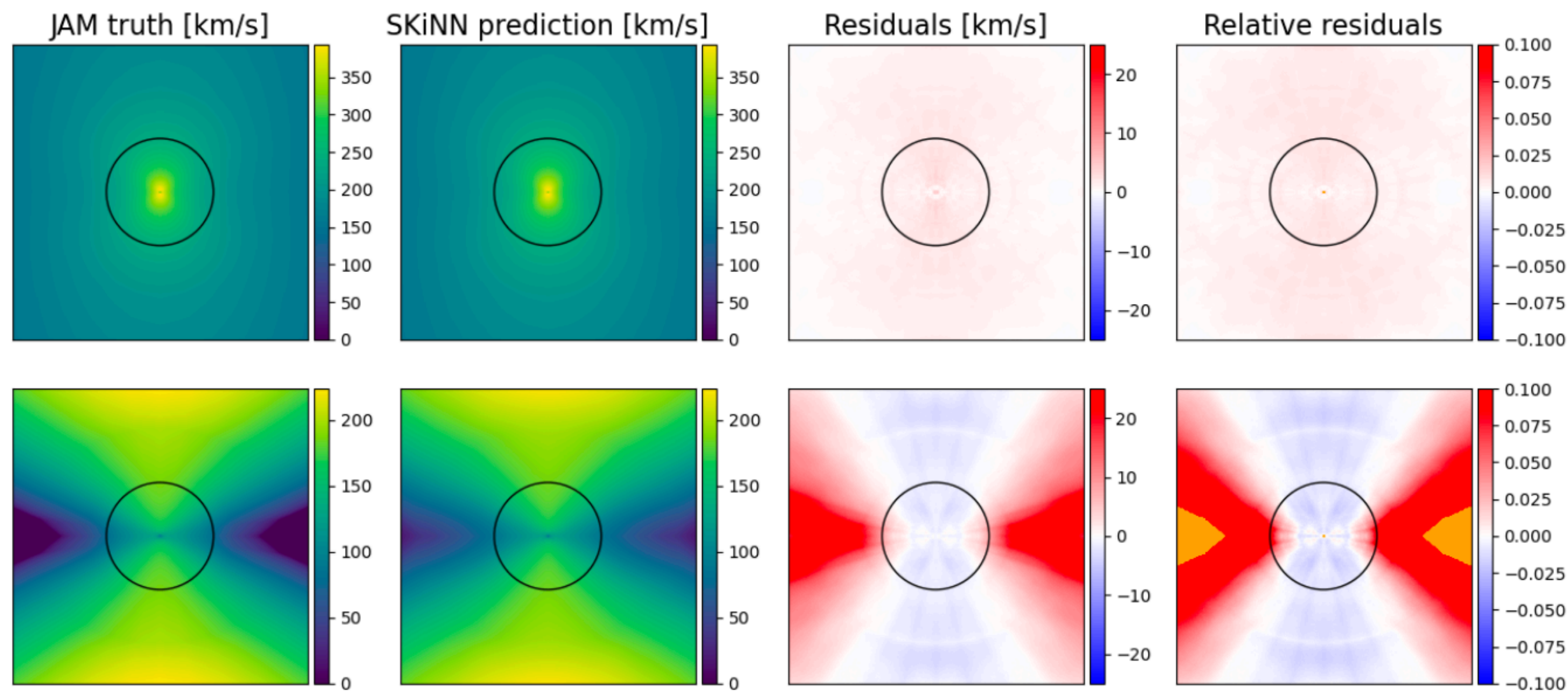
Gomer et al 2023 Architecture



- Joint lensing + kinematics inference
- Kinematics likelihood is a large bottle-neck ($\sim 500x$ slower than lensing likelihood)
- Conditionally Independent Pixel Synthesis (CIPS)
- Hardware requirements: GPU

(Gomer et al 2023, Fig. 1)

Gomer et al 2023 Training Set



Parameter	Description	Training set bounds
θ_E	Einstein radius	[0.5, 2'']
γ	2D PL slope (mass)	[1.5, 2.5]
q_M	Axis ratio (mass)	[0.6, 1.0]
q_L	Axis ratio (light)	[0.6, 1.0]
R_{Sersic}	Sérsic radius (light)	$[0.5\theta_E, \theta_E]$
n_{Sersic}	Sérsic index (light)	[2, 4]
β_z	Anisotropy	[-0.4, 0.4]
i	Inclination	$[\arccos(0.6), 90^\circ]$
Map resolution	0.02''	
Map size	11''	

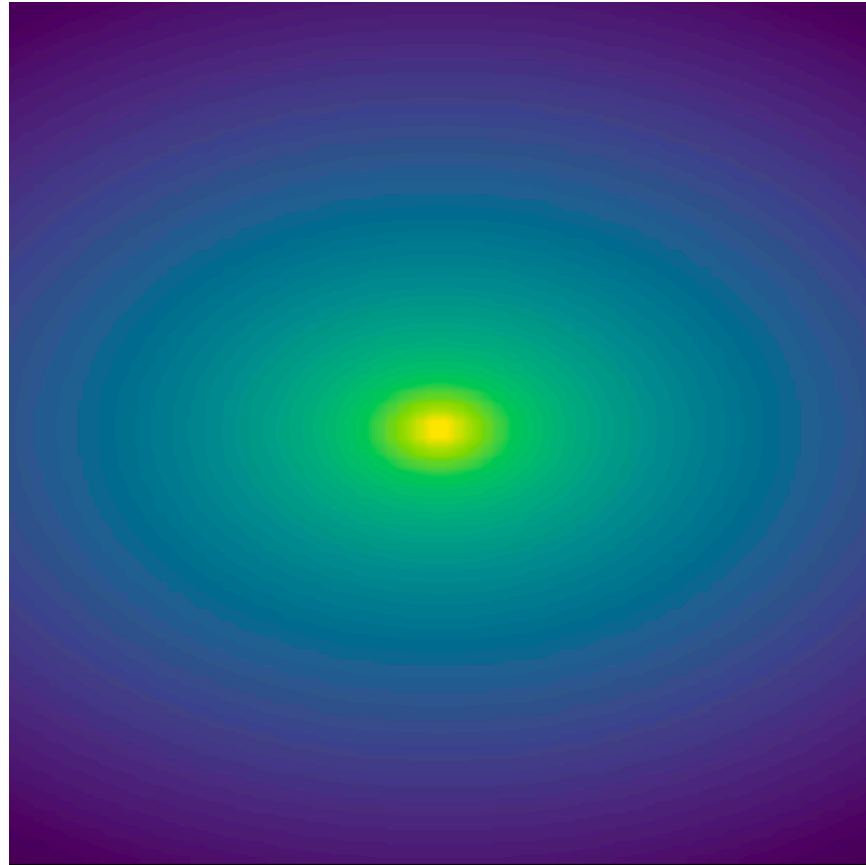
- 9000 Mock Galaxies
 - 4000 training
 - 1000 validation
 - 4000 test
- Uniformly sample Sersic galaxies
- Run time of $\sim 100\text{ms}$ on GPU
- Better than 1% accuracy for most galaxies

(Gomer et al 2023, Fig.2 & Table 1)

ML Architecture Requirements

- Directly use photometry
- Fast run time for MCMC
- Make use of symmetries (axisymmetry, $\overline{v_{\text{los}}^2} \propto M$)
- Compatible with GPU and CPU

Training Set



$$n_{\text{ser}} \in [2, 4]$$

$$R_{\text{ser}} \in [5'', 20'']$$

$$q_{\ell} \in [0.6, 1]$$

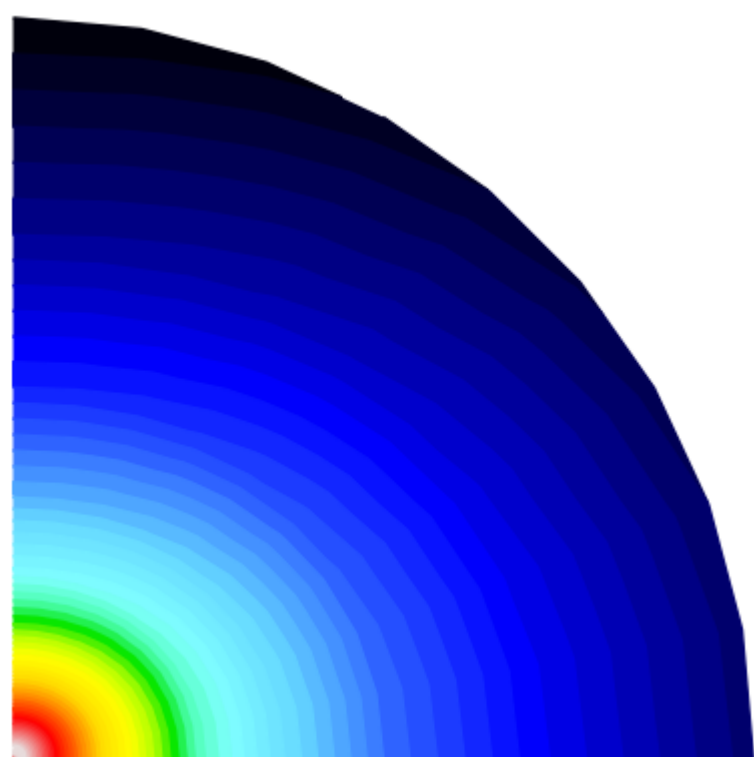
$$\beta_z \in [-0.4, 0.4]$$

$$q_m \in [0.6, 1]$$

$$\text{inc} \in [\arccos(0.6), 90^\circ]$$

$$\theta \in [0^\circ, 180^\circ]$$

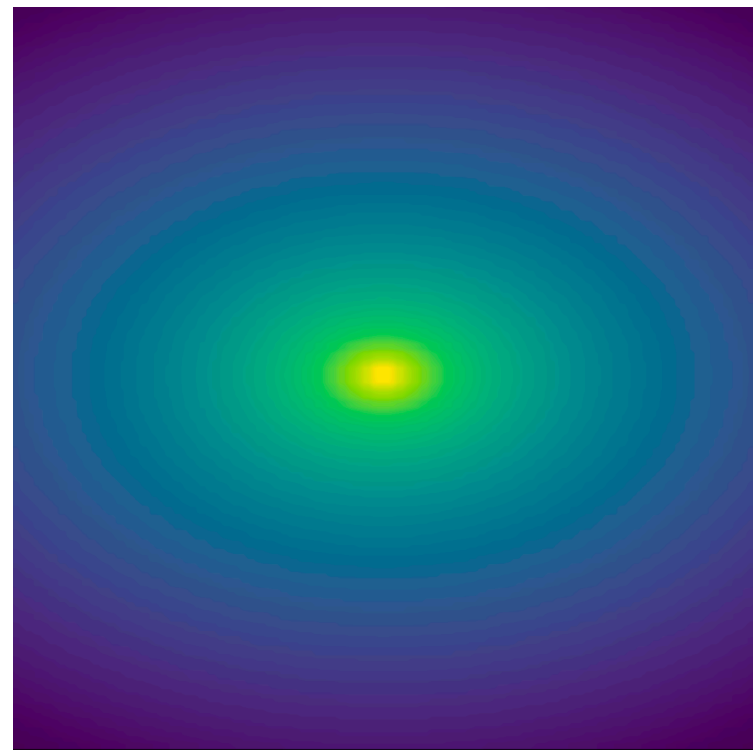
JAM Vrms



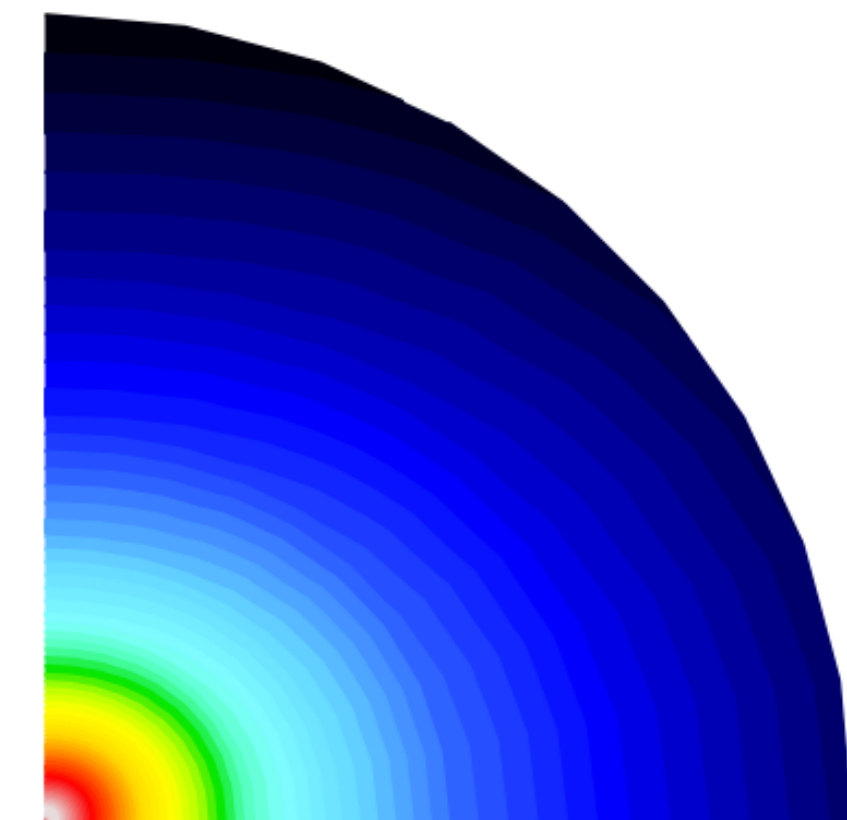
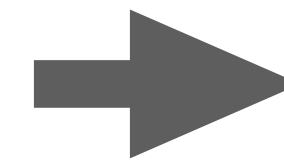
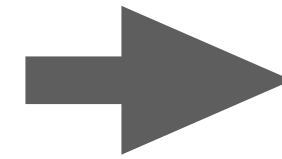
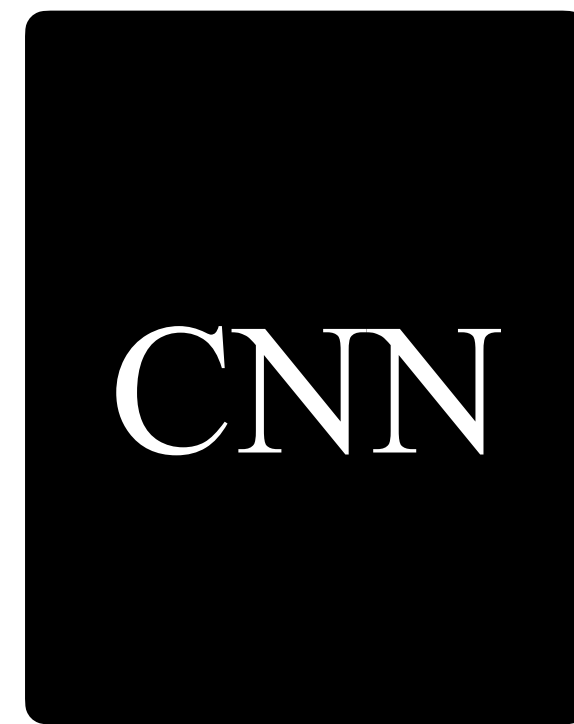
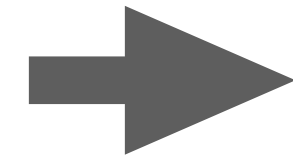
- 10,000 mock Galaxies
 - 8000 training
 - 2000 test
- Randomly sampled parameters while enforcing physical solutions
 - $\beta \leq 0.7\epsilon_{\text{intr}}$ (See Wang et al 2021)

ML Modelling Pipeline

$n_{\text{ser}}, r_{\text{ser}}, q_l$



Photometry

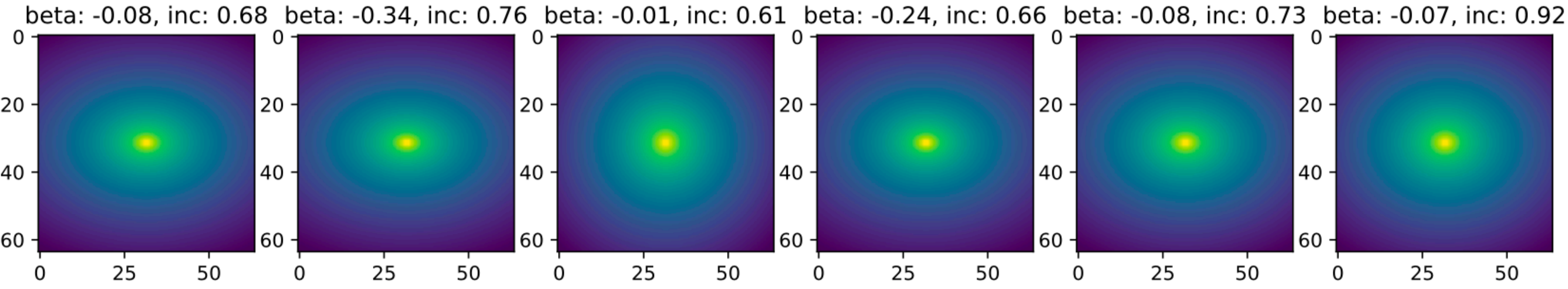


β, Inc, q_m

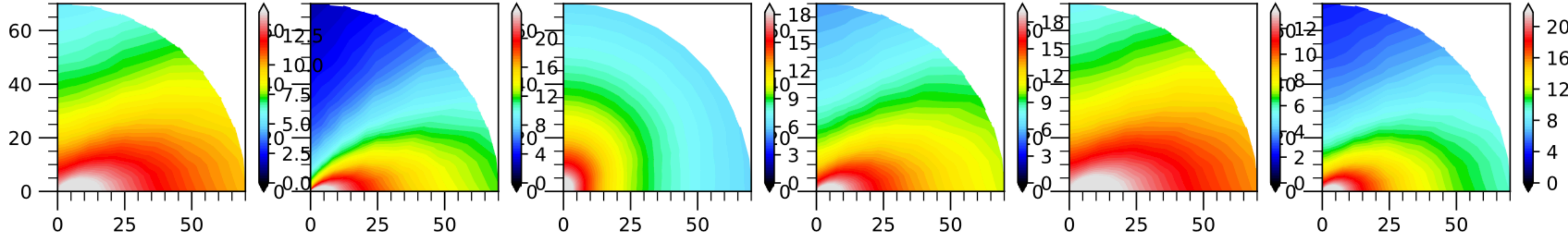


Outcomes: Maps

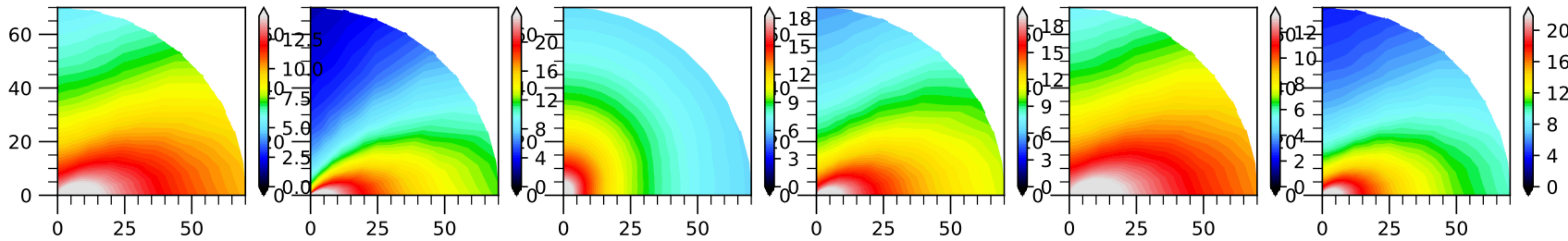
Photometry



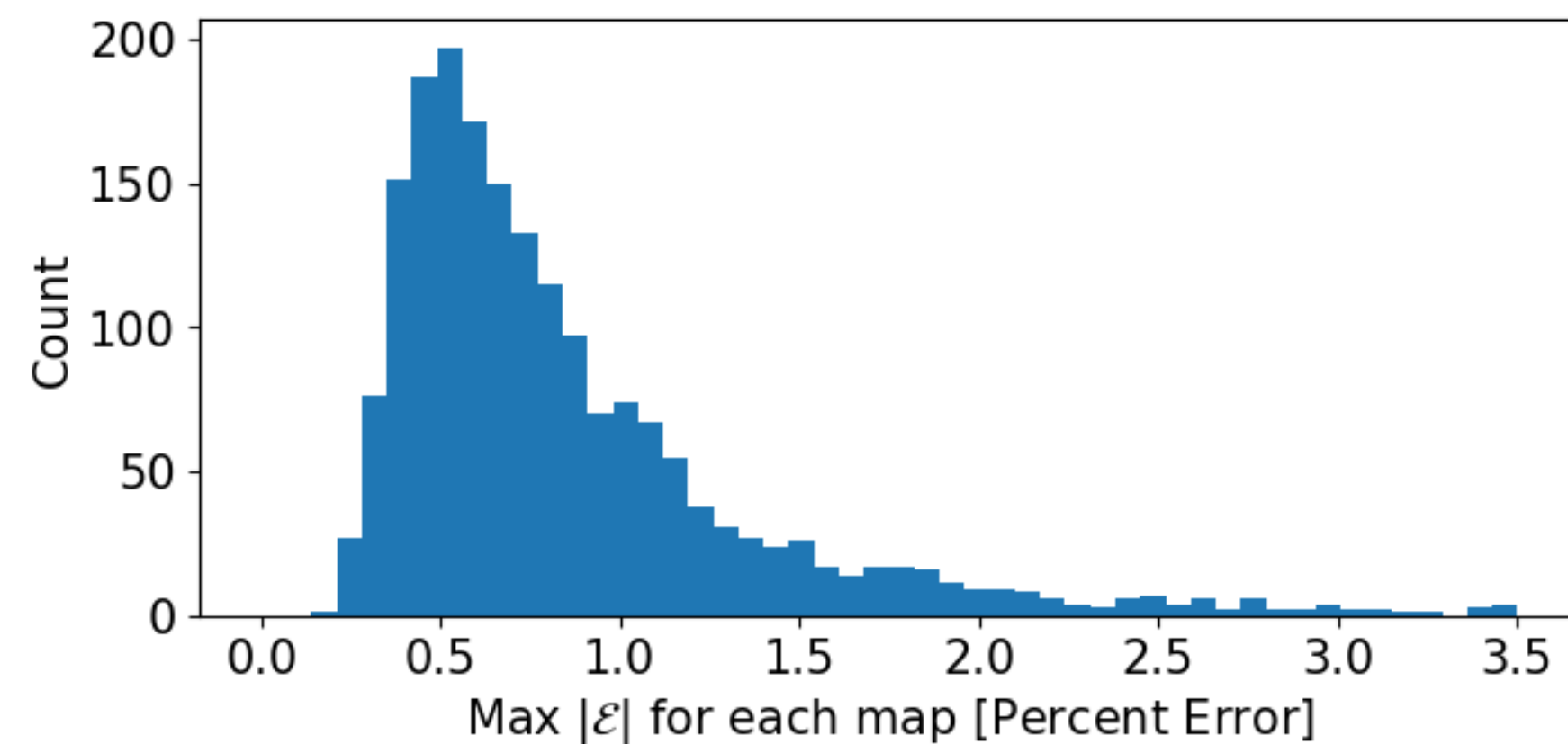
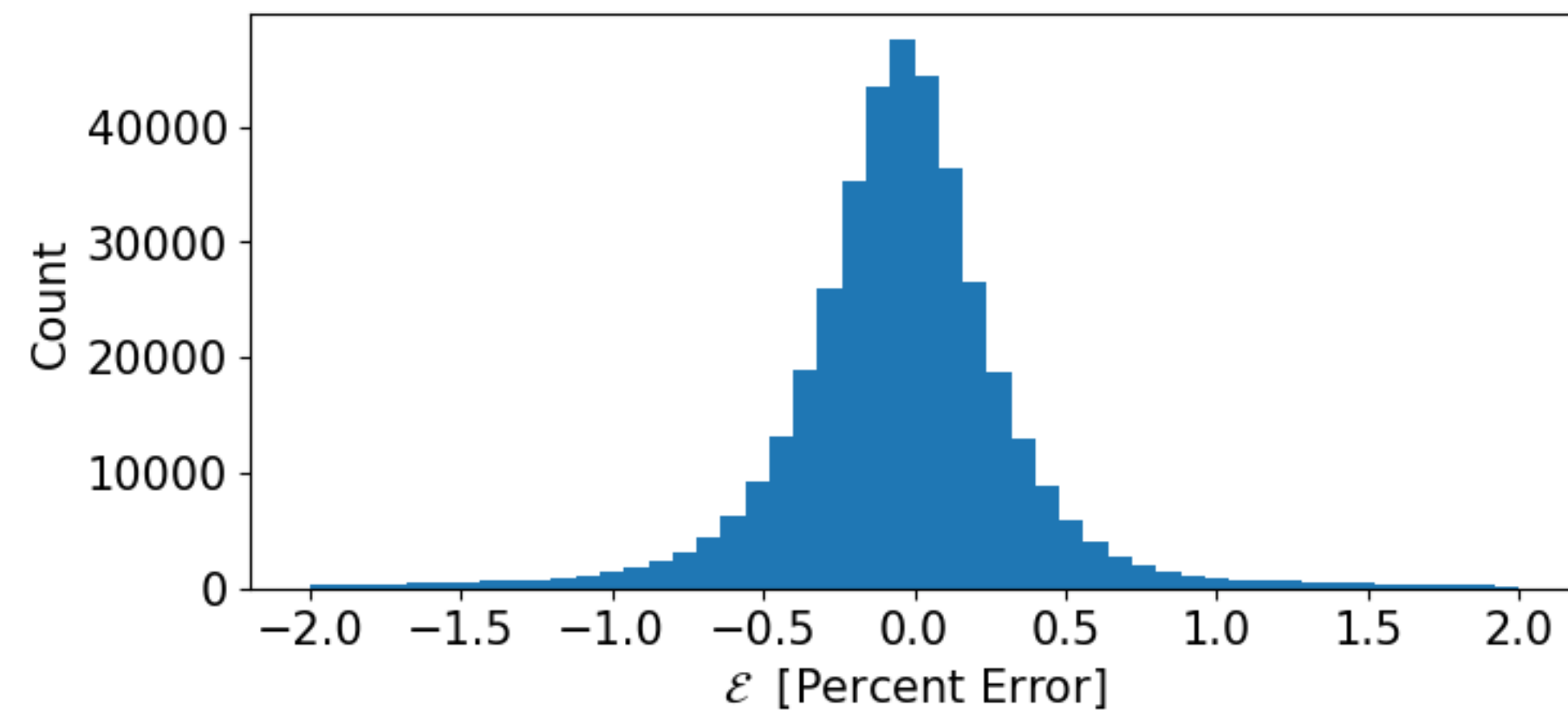
True Vrms



ML Vrms



Outcomes: Accuracy

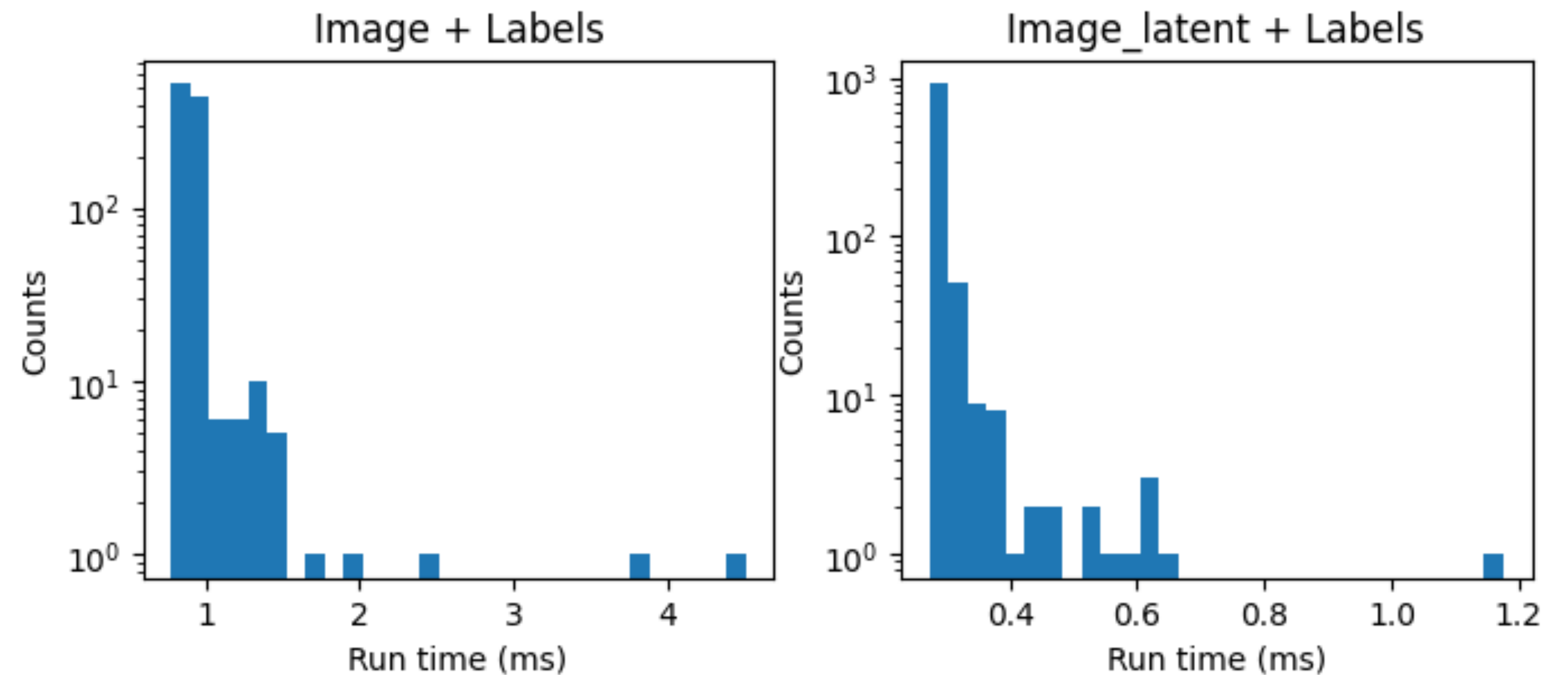


- The test set maps are accurate to within 1% and ninety percent of the maps have a max error less than 2%
- Real observations are accurate to within 6-7 km/s

$$\epsilon = \frac{\text{Model} - \text{Truth}}{\text{Truth}}$$

Outcomes: Timing

- Training set can be generated in ~ 3 hours on a single CPU core
 - 2.5 days for Gomer et al 2023
- Model can be trained on a single NVIDIA V100 GPU in ~ 30 minutes
 - 3 days for Gomer et al 2023



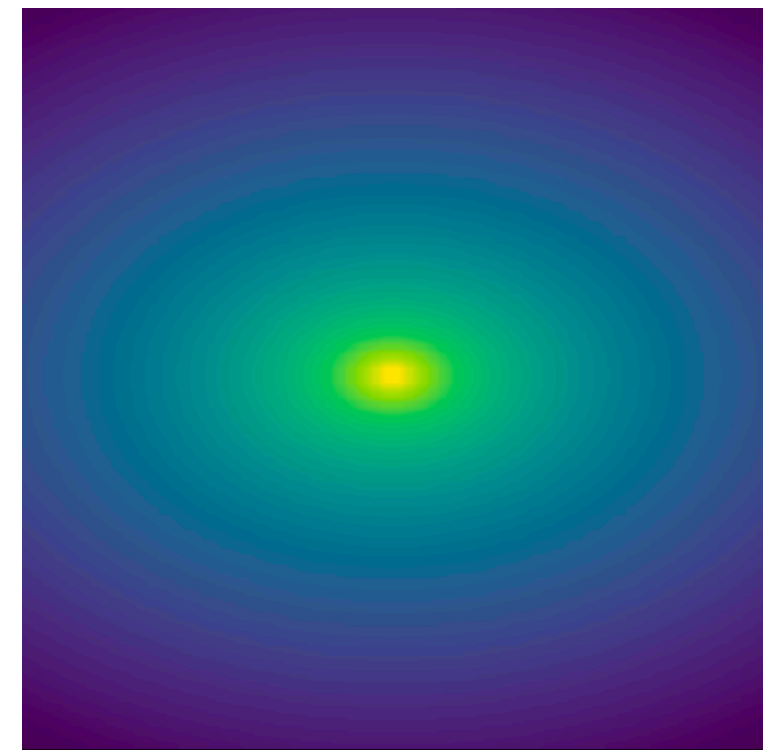
Almost 100x faster than (Gomer et al 2023) and on CPU

Key Questions

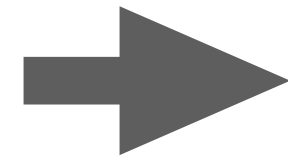
- ~~Can Machine Learning be used to Improve Existing Methods?~~
- Can Machine Learning be used to Create New Methods?

(Preliminary results!)

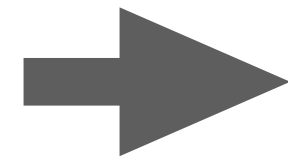
ML Modelling Pipeline



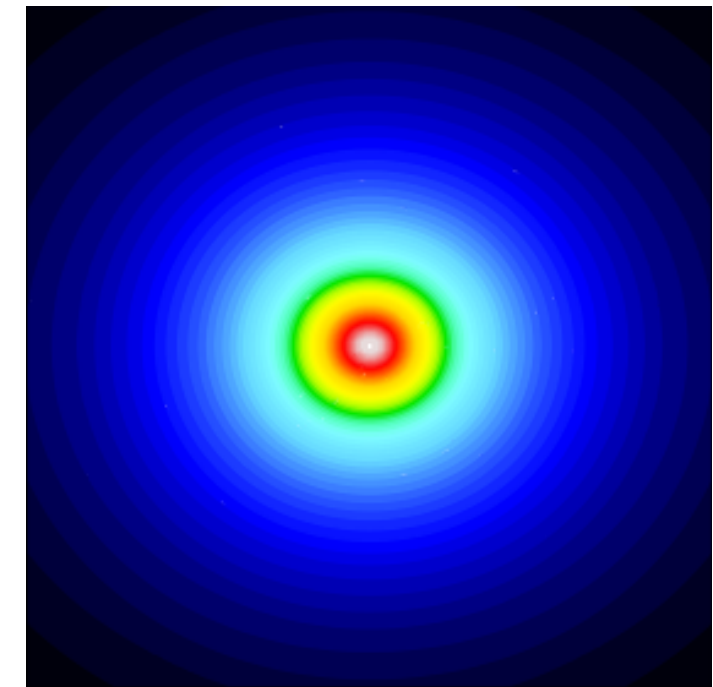
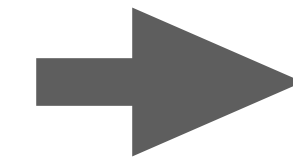
Photometry



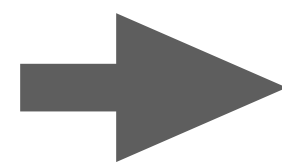
**Gravitational
Potential**



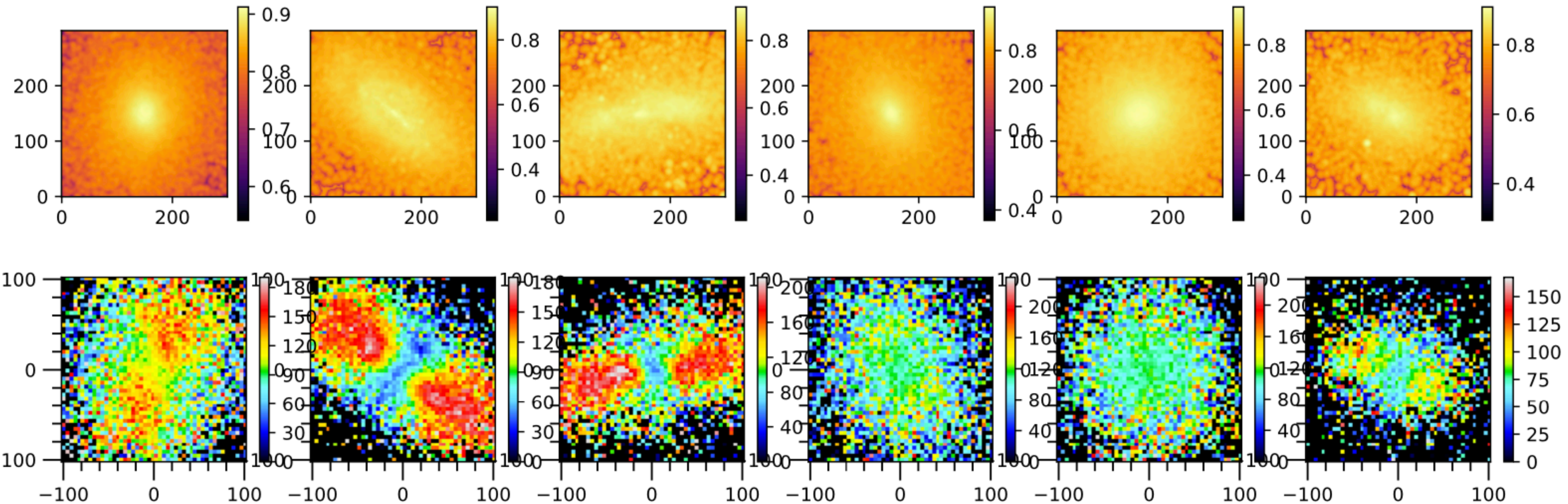
**Machine Learning
Algorithm**



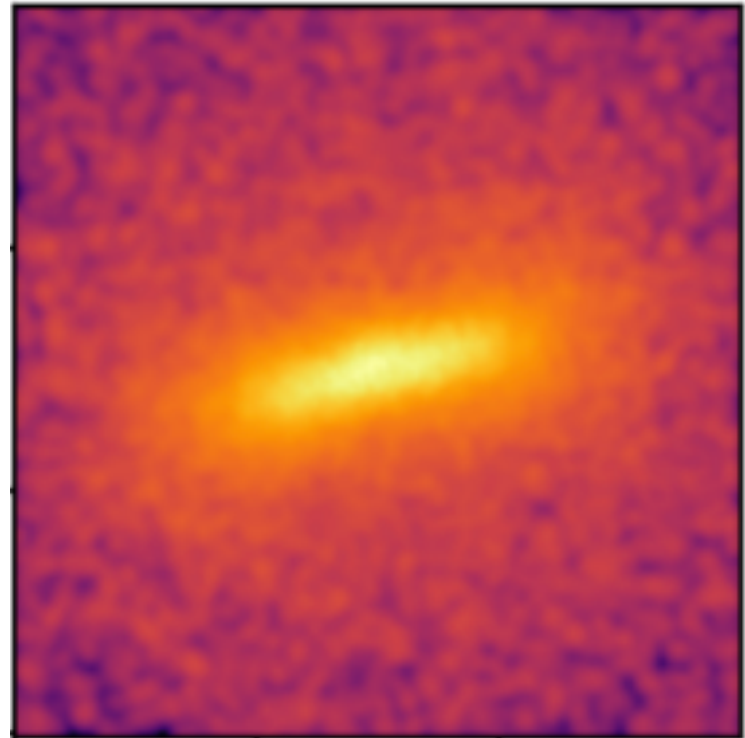
**Free
Parameters**



TNG100 Mock Observations



Training Set



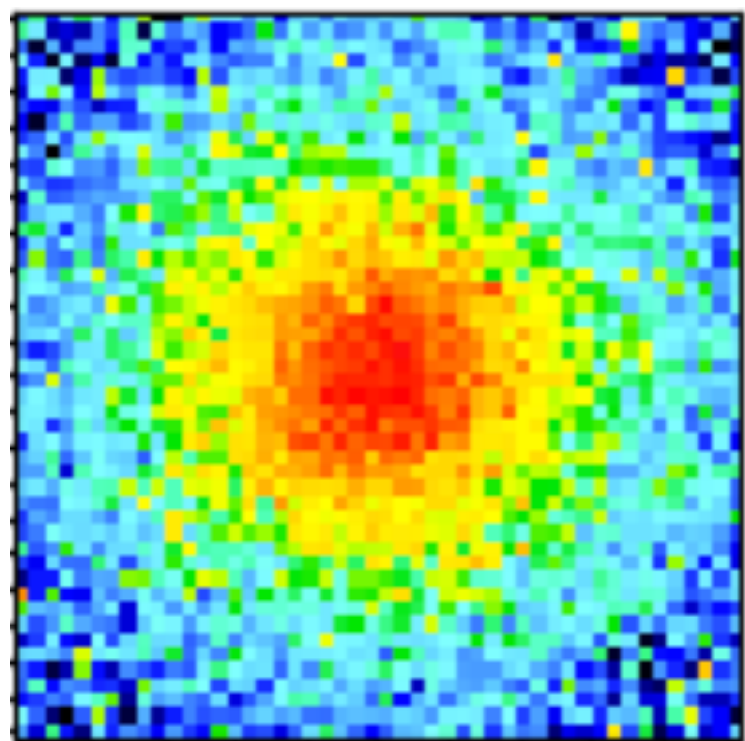
Inc

$$\beta(r < R_e)$$

$$M/L(r < R_e)$$

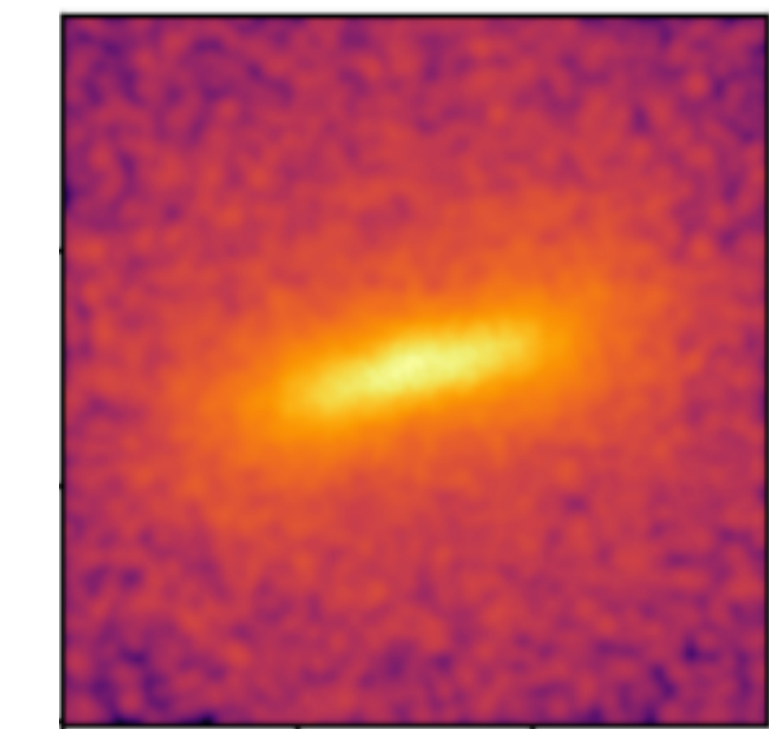
$$f_{\text{dm}}(r < R_e), \gamma, r_s$$

$$\theta \in [0^\circ, 180^\circ]$$

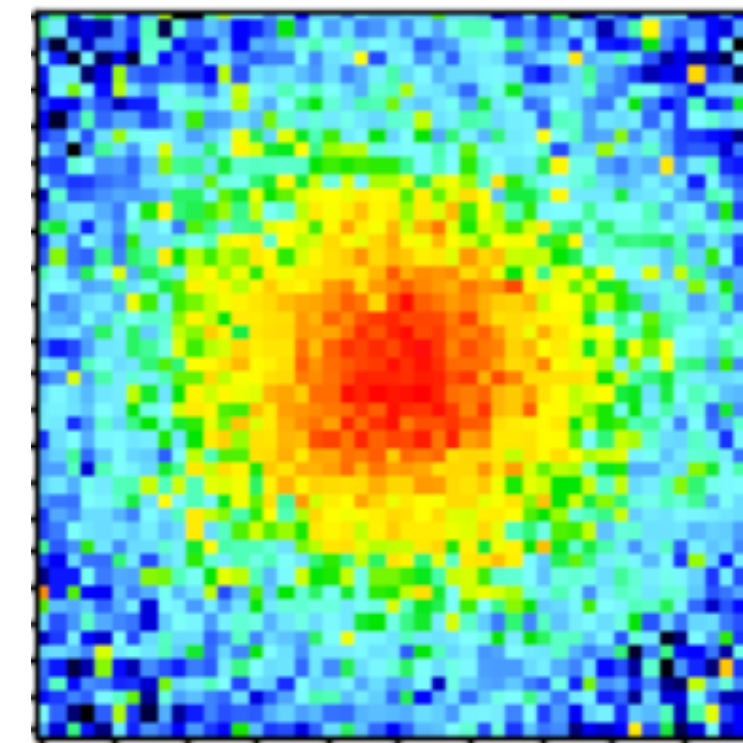
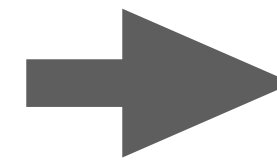
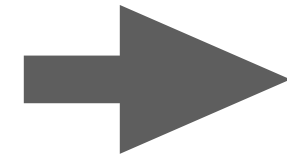
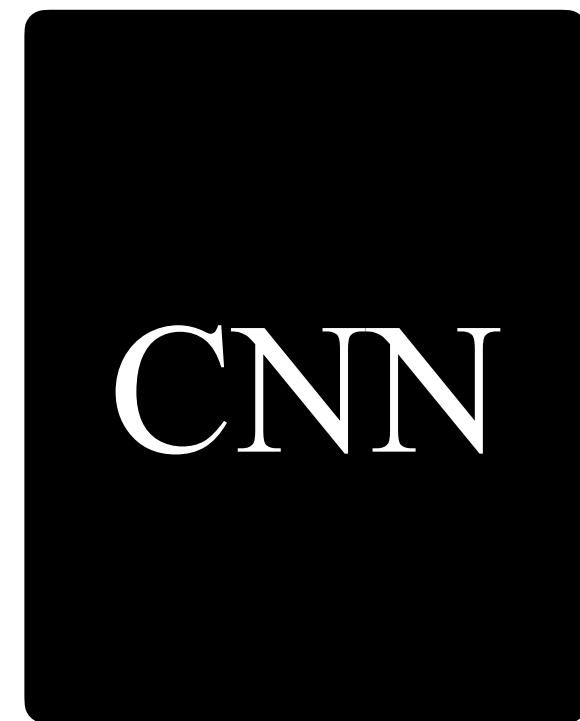
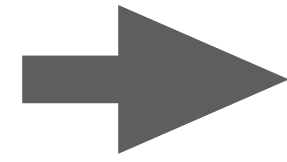


- TNG100 Mock Observations
 - 1.4e6 solar mass resolution for stars
 - Stellar mass within 30 kpc > 5e9 solar mass
 - Total mass < 1e14 solar mass
- ~5000 galaxies
- Mock observations along x,y, and z axis
- Field of view out to 3 Re

ML Modelling Pipeline



Photometry



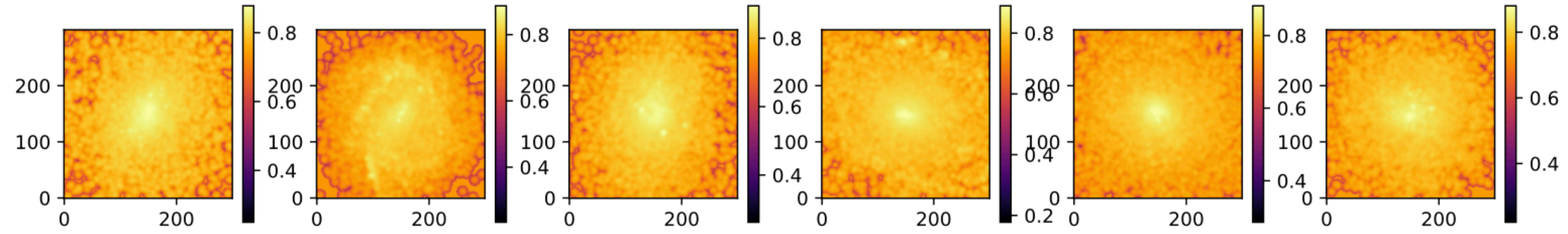
Use non-standard loss to mask empty pixels

β , Inc, M/L ,
 f_{dm} , γ , r_s , θ

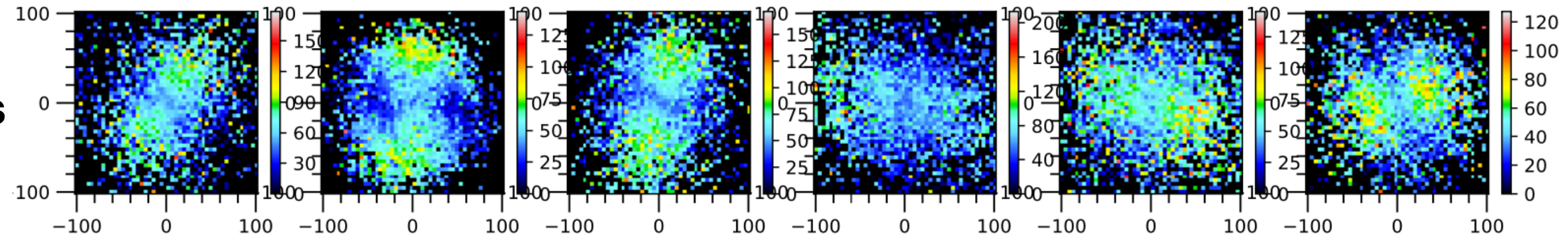


TNG100 Results

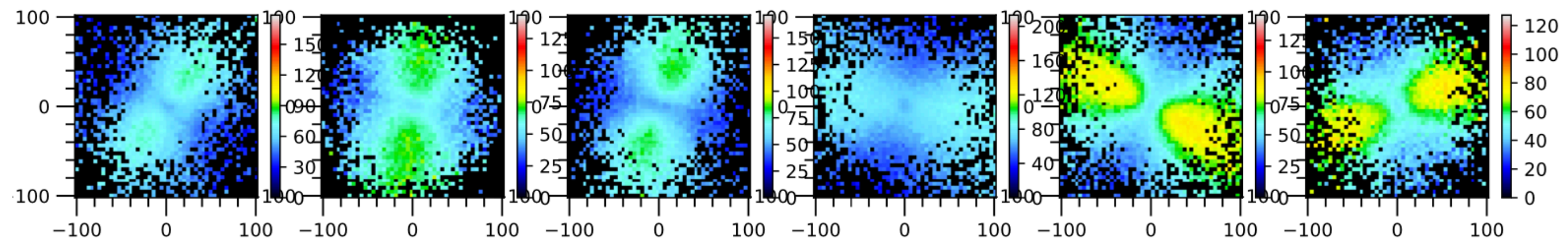
**TNG100
Photometry**



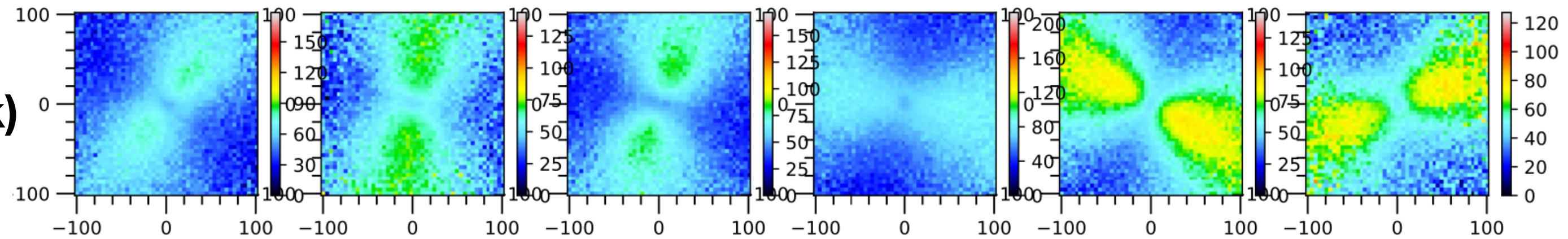
TNG100 Vrms



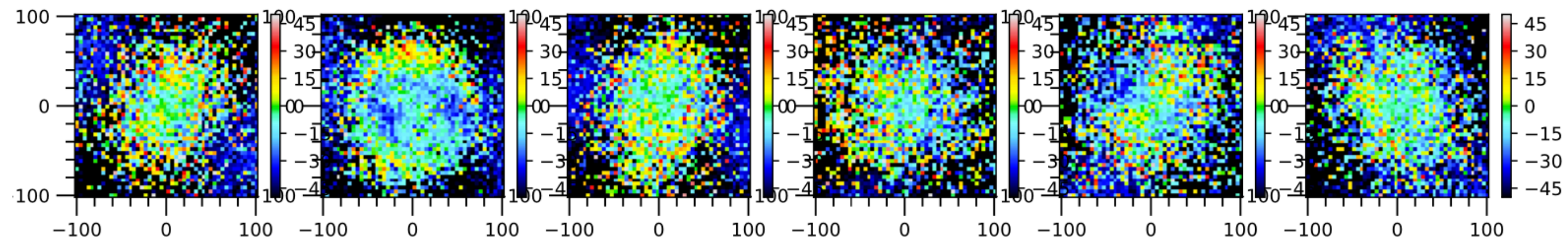
**ML Vrms
(with mask)**



**ML Vrms
(without mask)**



Residuals



Conclusions

- Simple machine learning algorithms can sometimes outperform complicated ones
- Close to 10,000 data points are required in order to machine learn dynamical models (less with more complicated models)
- Machine Learned dynamical models have the potential to outperform existing methods