

# Cosmology in the machine learning era

Francisco Villaescusa-Navarro



**ML4Astro2**



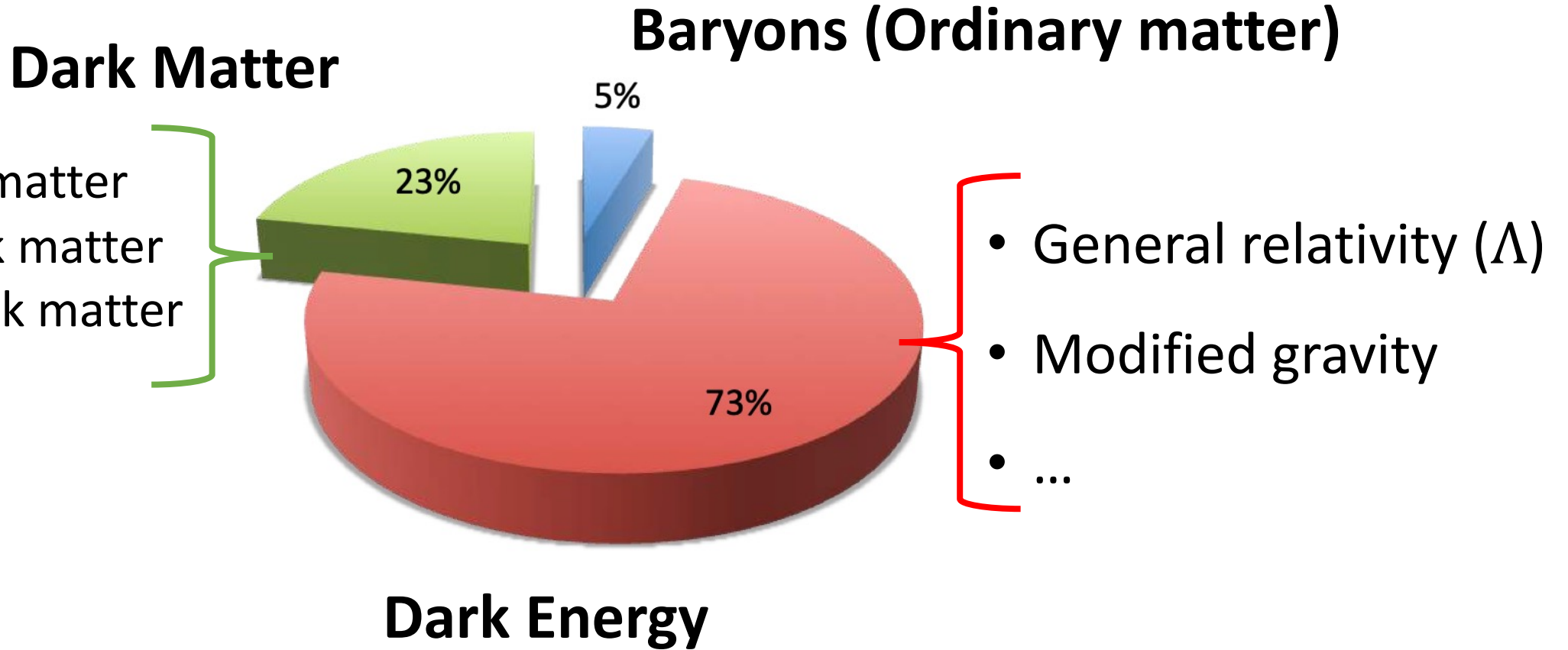
# Cosmology

Branch of Astrophysics that studies the Universe's:

- Origin and fate
- Composition
- Laws
- Structure
- Dynamics

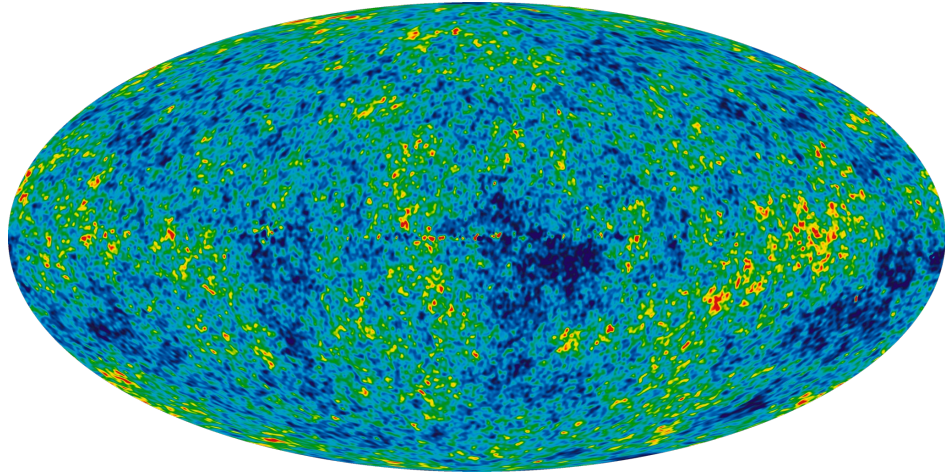


# Our Universe





# The $\Lambda$ CDM model



$$\Omega_b \pm \delta\Omega_b$$

$$\Omega_m \pm \delta\Omega_m$$

$$h \pm \delta h$$

$$w_0 \pm \delta w_0$$

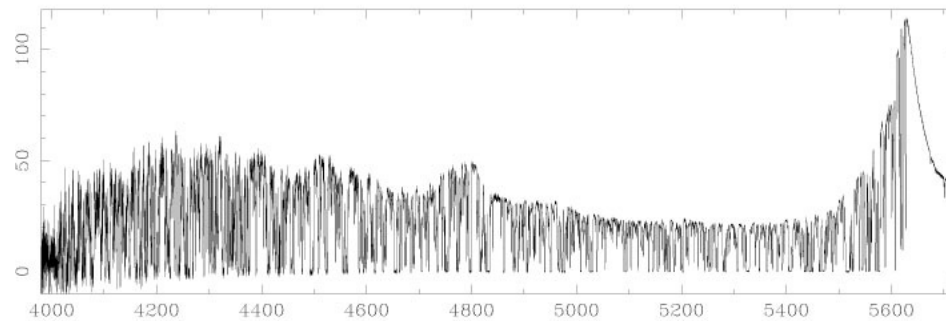
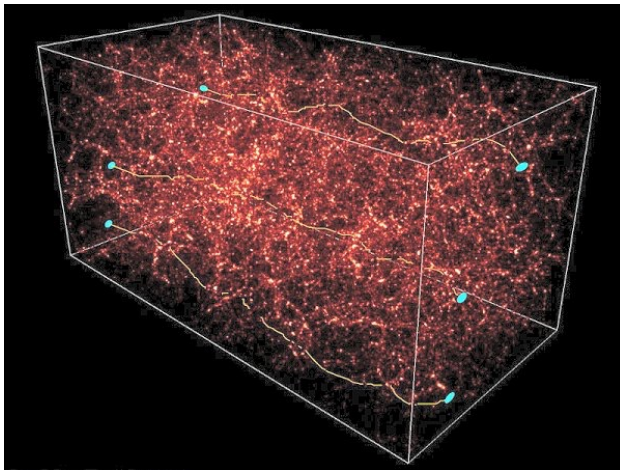
$$w_a \pm \delta w_a$$

$$n_s \pm \delta n_s$$

$$\sigma_8 \pm \delta\sigma_8$$

$$M_\nu \pm \delta M_\nu$$

$$N_{\text{eff}} \pm \delta N_{\text{eff}}$$



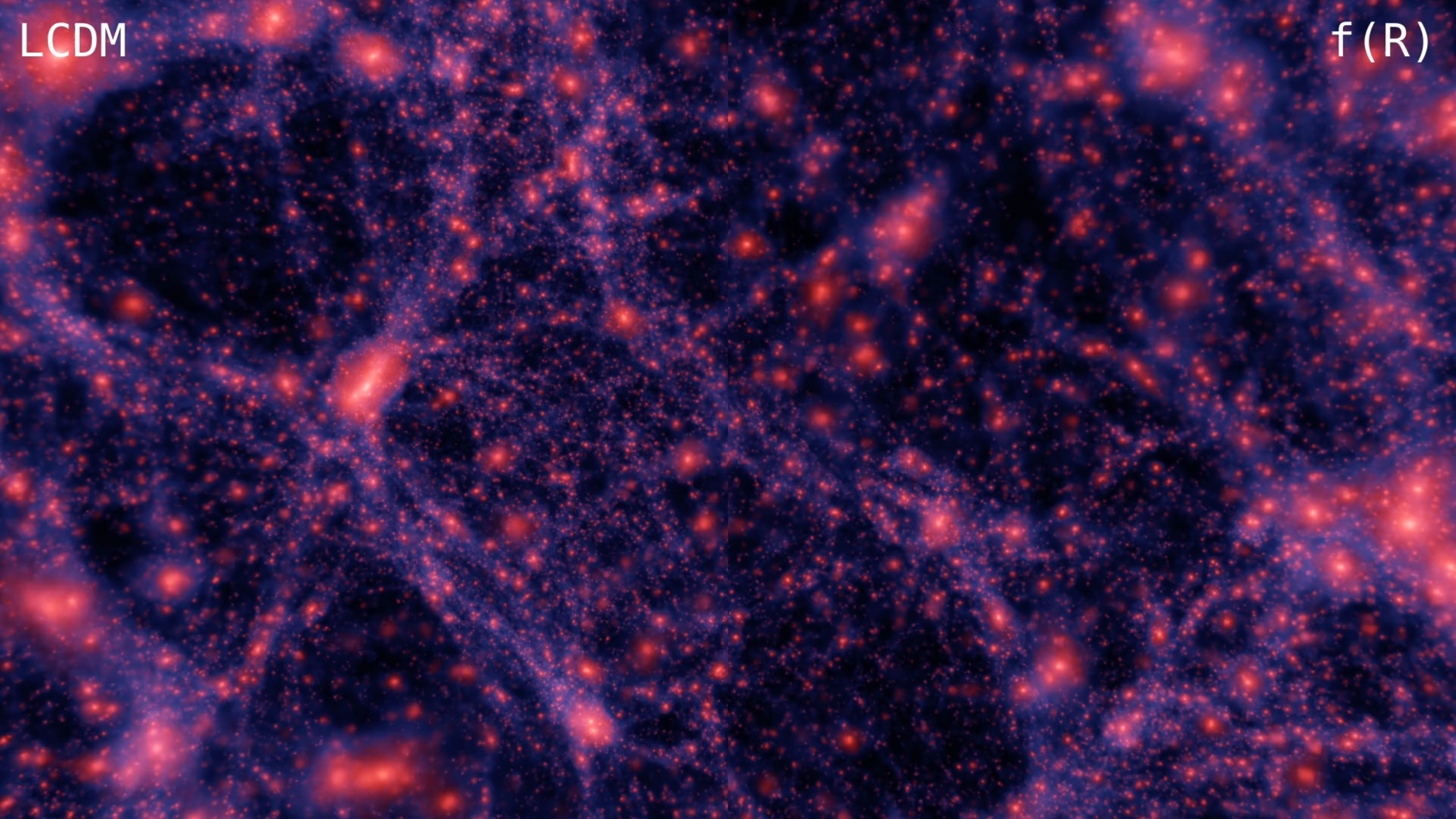






$z = 20.00$





LCDM

$f(R)$



# The Quijote Simulations

(<https://quijote-simulations.readthedocs.io>)

- A set of 85,000 full N-body simulations
- More than 42,000 cosmologies in  $\{\Omega_m, \Omega_b, h, n_s, \sigma_8, M_v, w_0, \delta_b, f_{\text{NL}}, g_{\text{NL}}, f(R)\}$
- 12+ trillion particles over a volume larger than entire observable Universe
- Catalogs with billions of halos, voids (Gigantes), and galaxies (Molino). WL maps (Ulagam)
- 50 Million CPU hours; 1+ Petabyte of data
- 170+ papers written using this data
- All data publicly available (binder & globus)





Generic conclusion:  
Lots of information on  
small scales beyond  $P(k)$

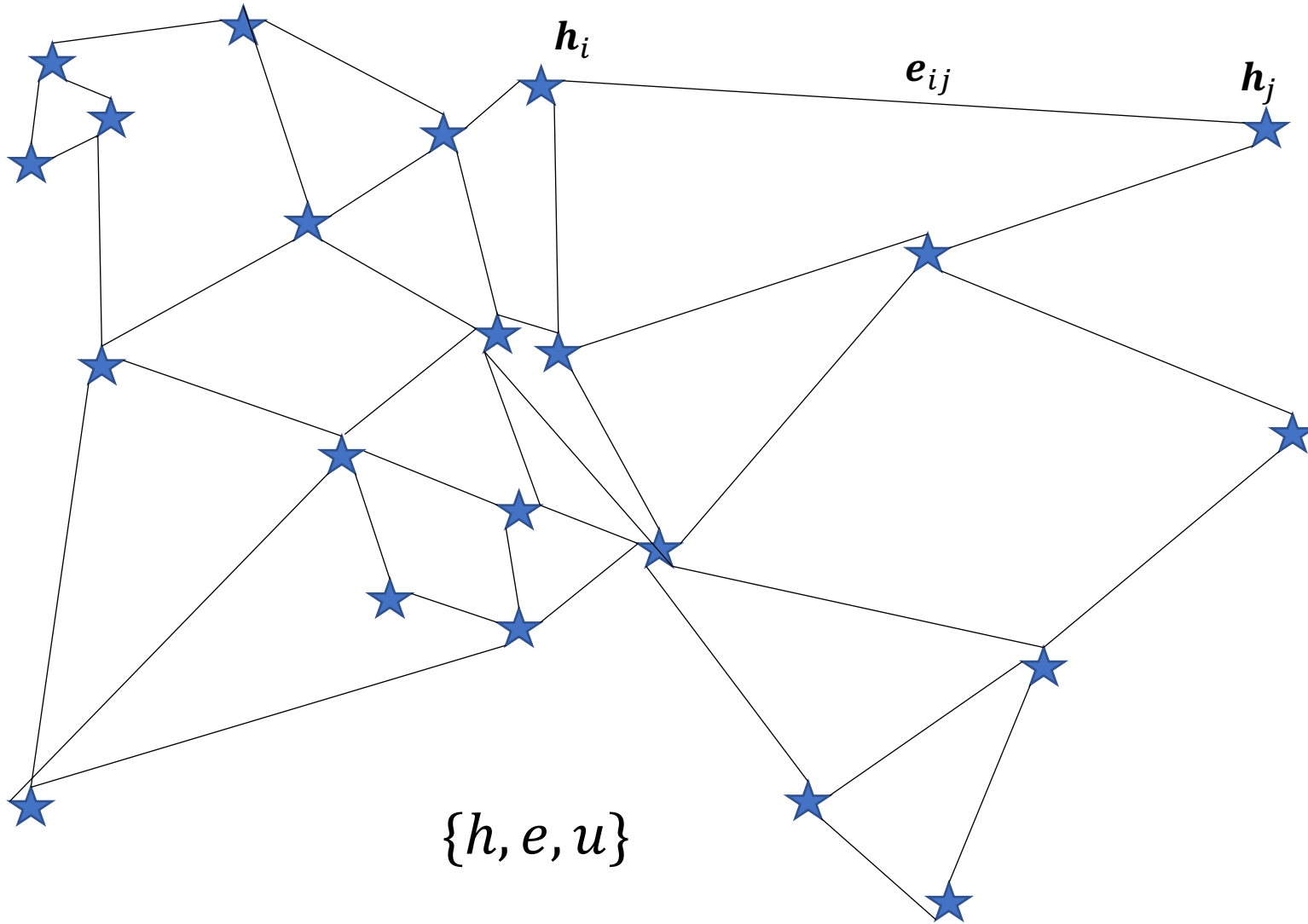
**Benefits:** Lots of information

**Problems:** Non-linearities &  
baryonic effects

$z = 9.94$



# Cosmic Graphs



Pablo Villanueva-Domingo  
(Barcelona)



Natali de Santi  
(Flatiron/Sao Paolo)

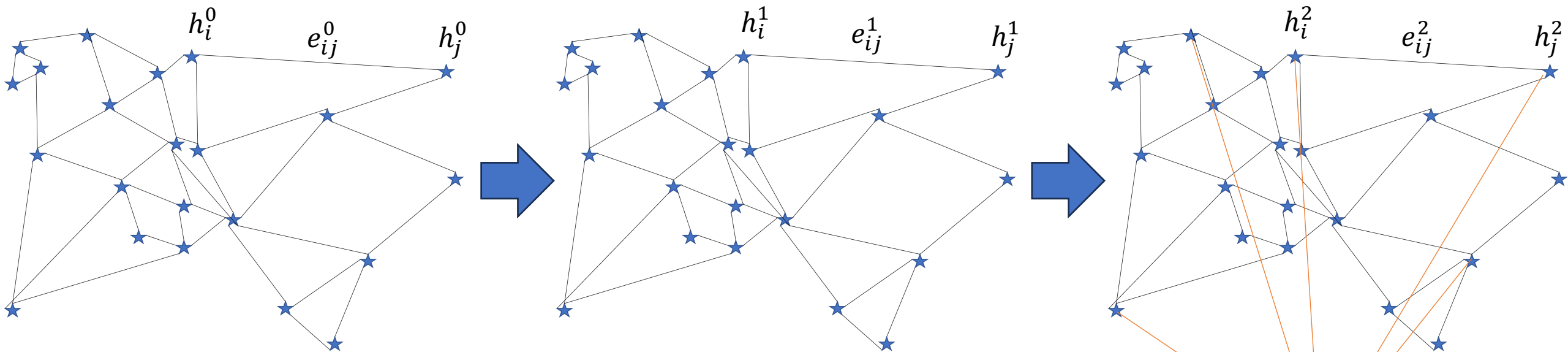


Helen Shao  
(Princeton)

Talk to Farida Farsian



# Graph neural networks



$$\mathbf{e}_{ij}^{(l+1)} = \phi_{l+1}([\mathbf{h}_i^{(l)}, \mathbf{h}_j^{(l)}, \mathbf{e}_{ij}^{(l)}])$$

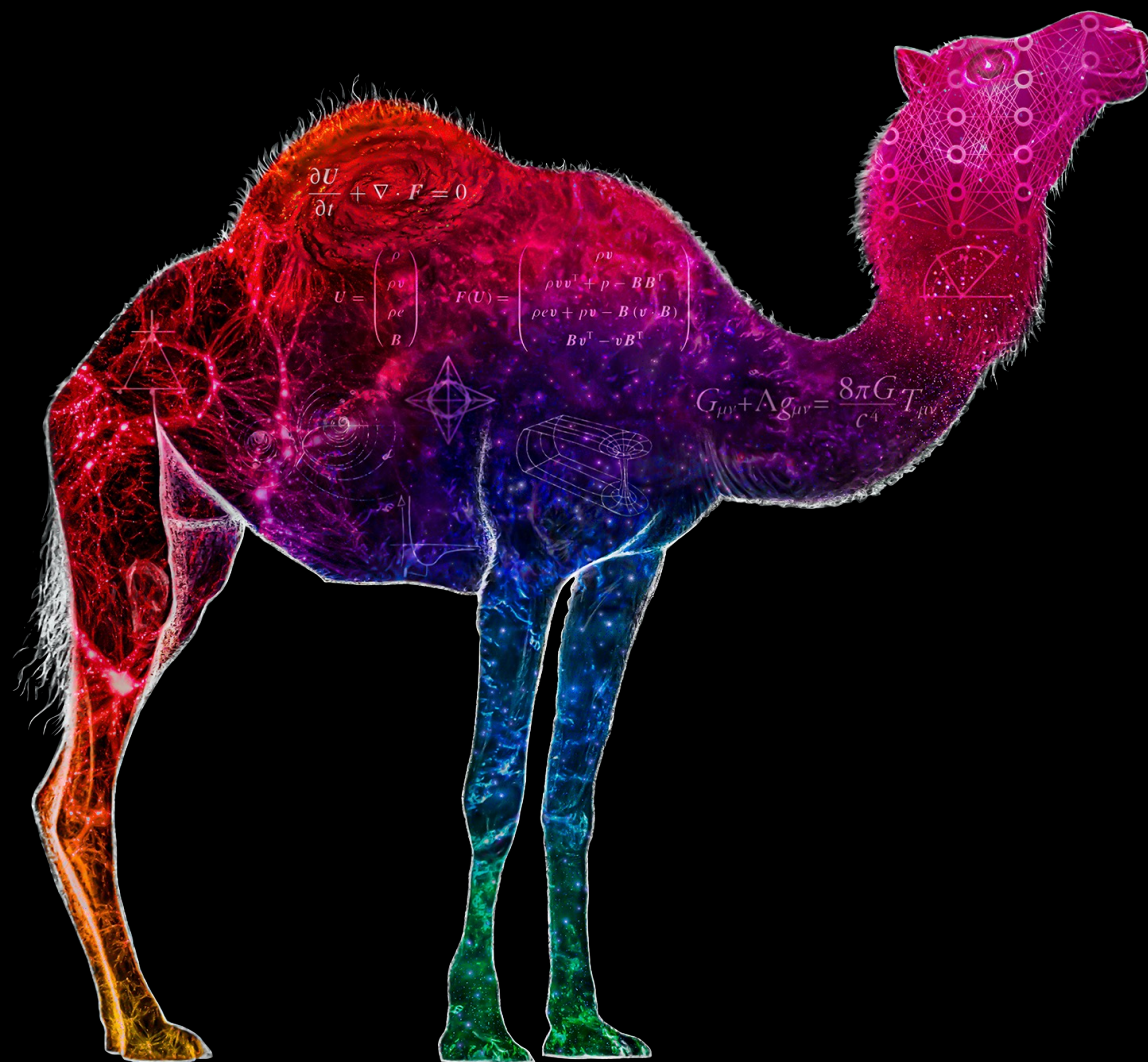
$$\mathbf{h}_i^{(l+1)} = \psi_{l+1}([\mathbf{h}_i^{(l)}, \bigoplus_{j \in \mathcal{N}_i} \mathbf{e}_{ij}^{(l+1)}, \mathbf{u}])$$

$$\mathbf{y} = \xi([\bigoplus_{i \in \mathcal{G}} \mathbf{h}_i^{(L)}, \mathbf{u}])$$

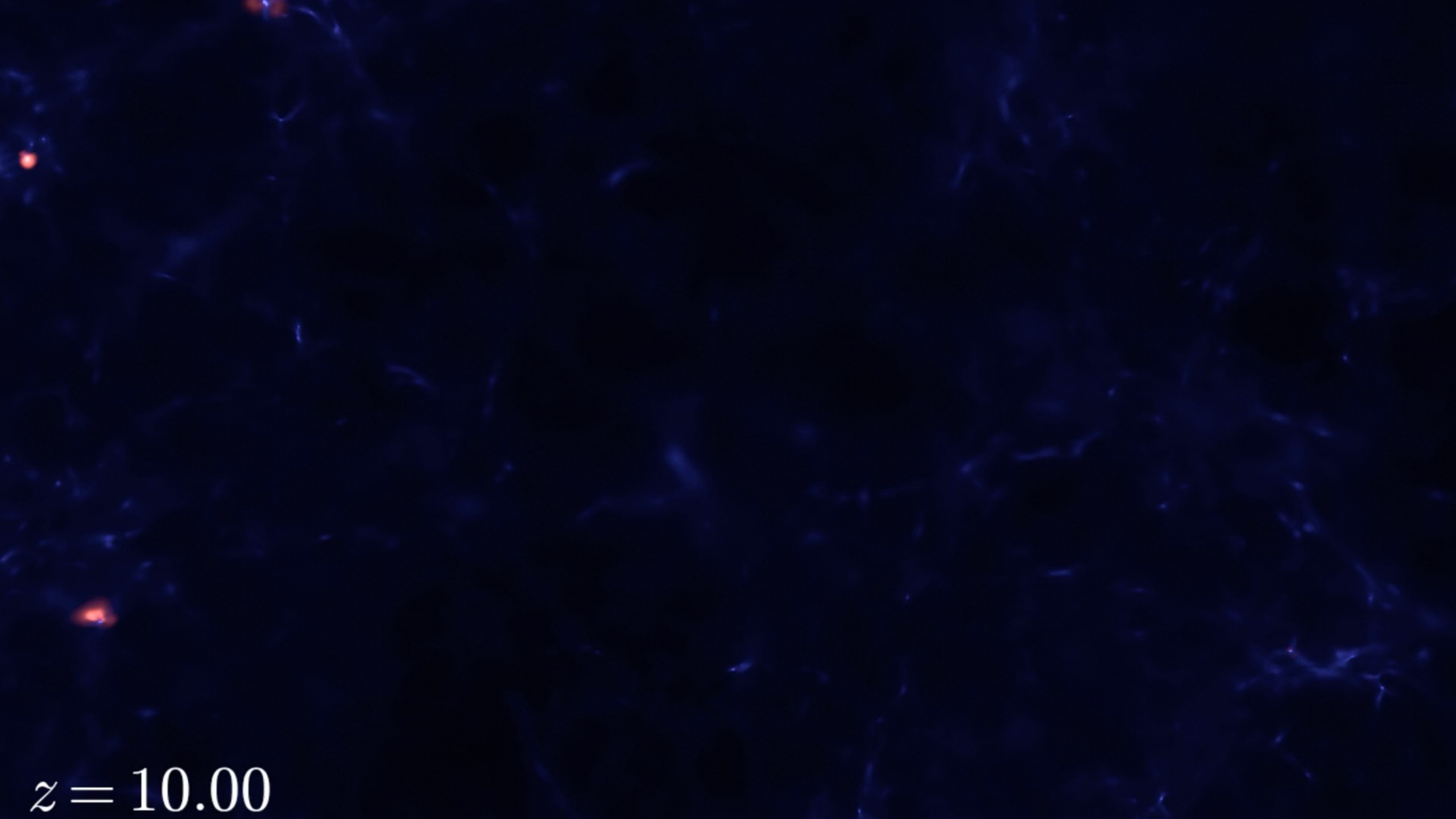
# CAMELS

<https://www.camel-simulations.org>

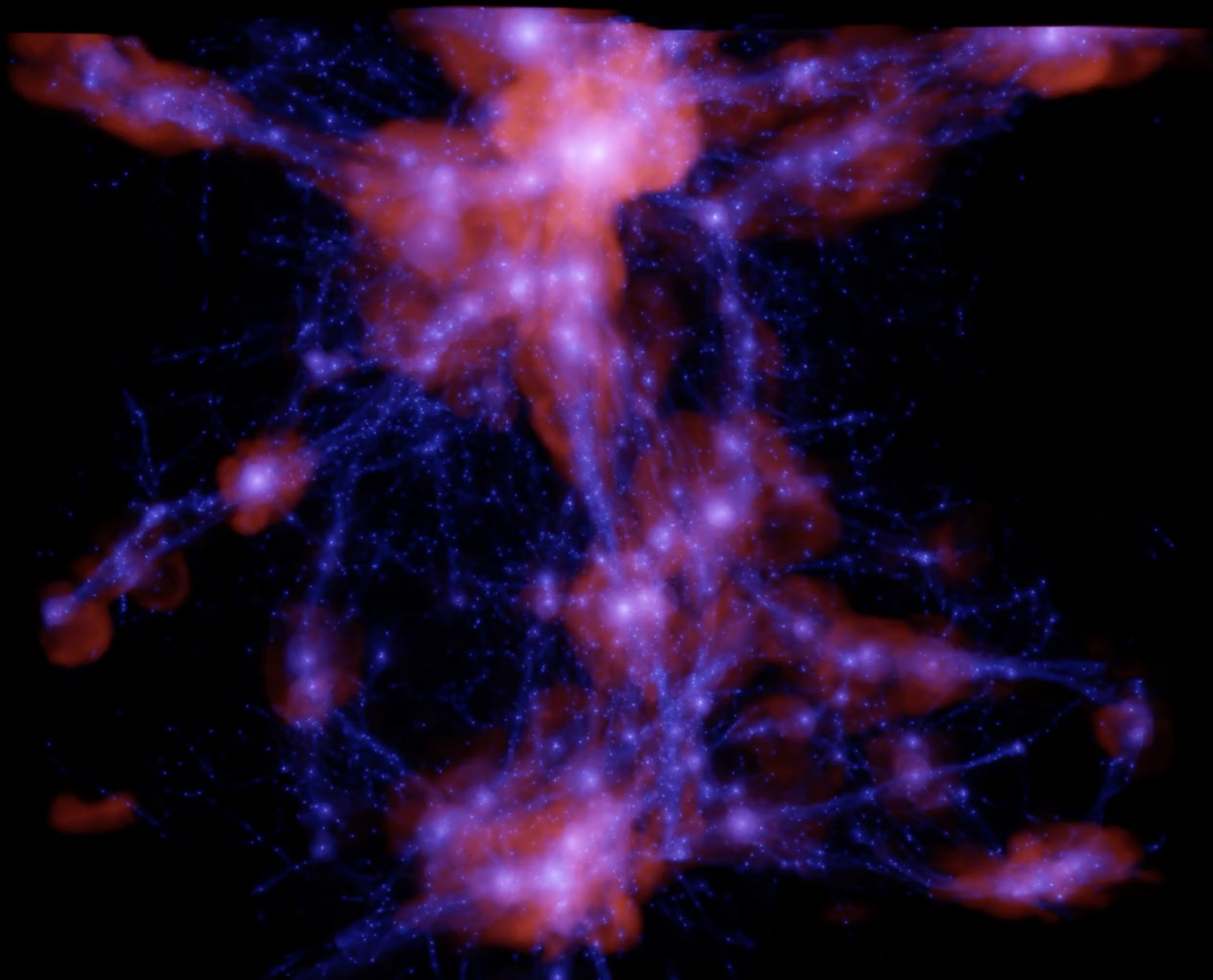
Cosmology and Astrophysics with  
Machine Learning Simulations





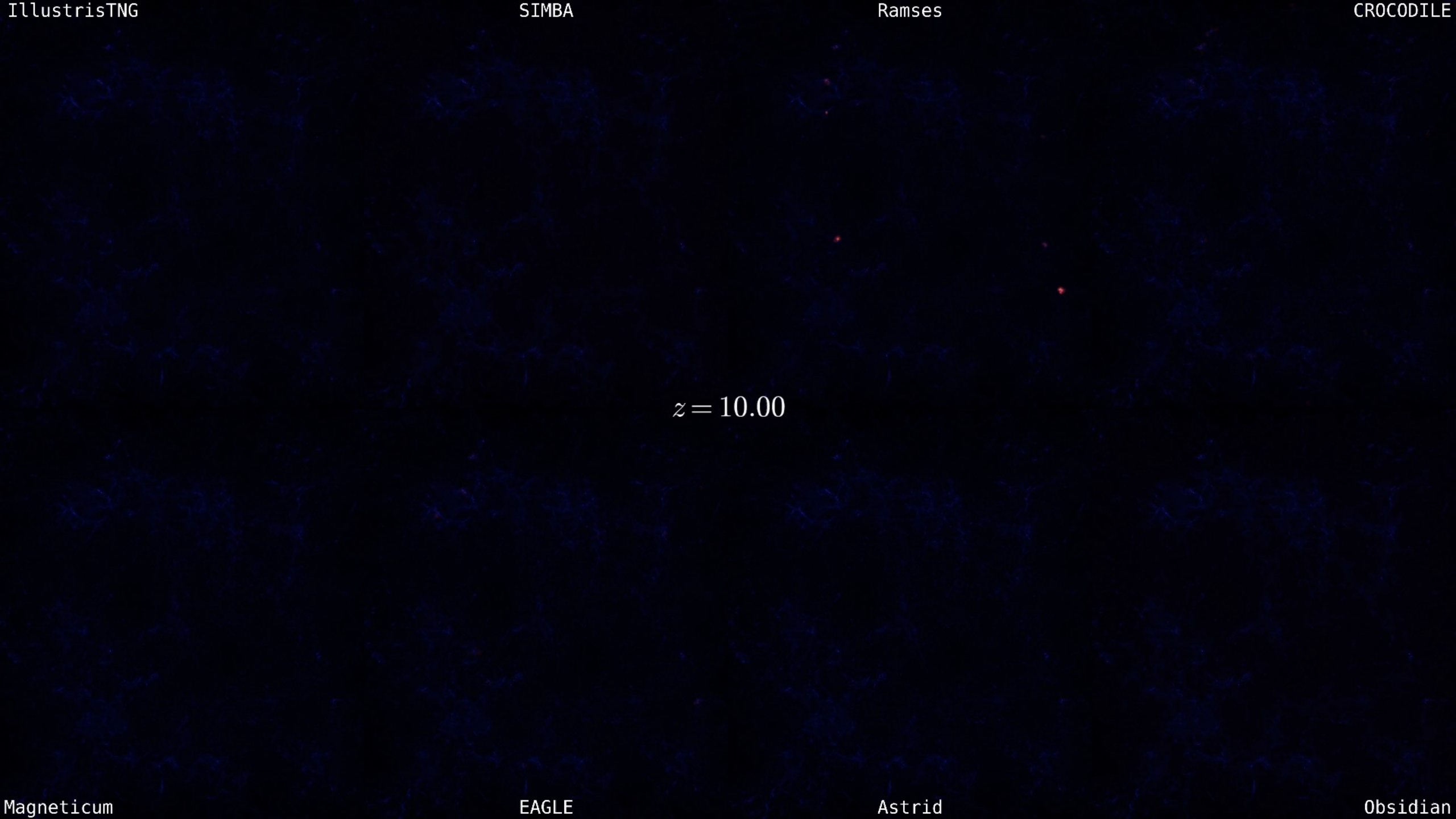


$z = 10.00$









IllustrisTNG

SIMBA

Ramses

CROCODILE

$z = 10.00$

Magneticum

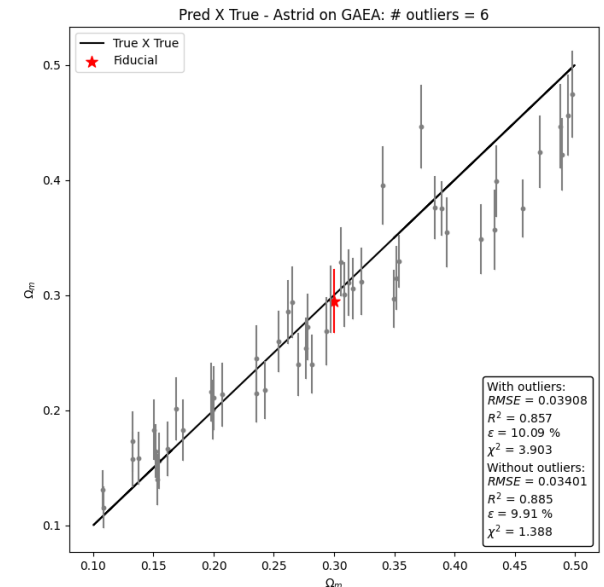
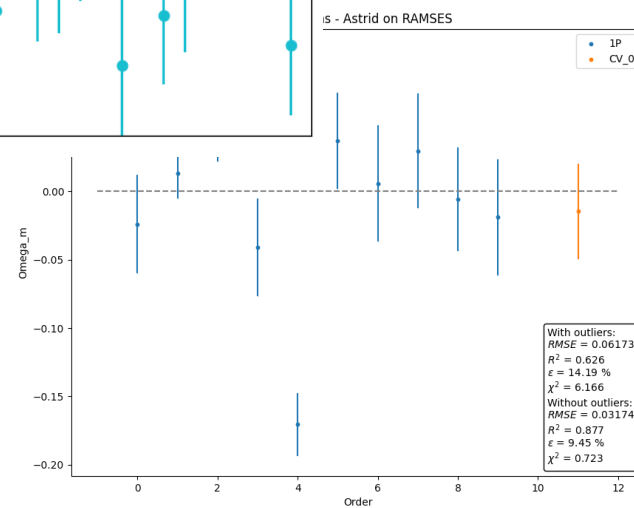
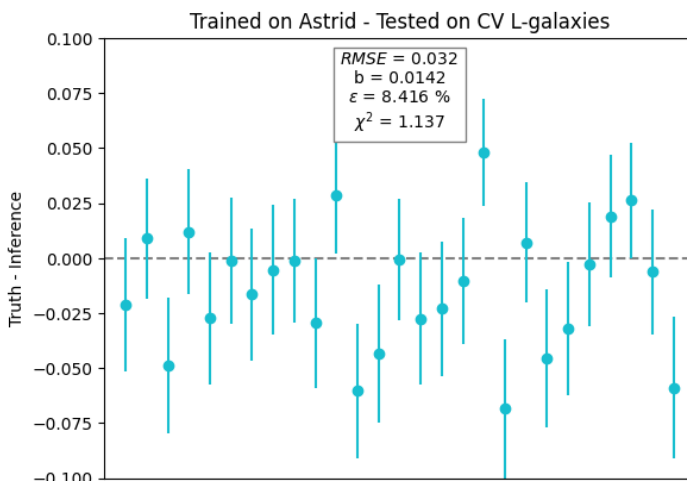
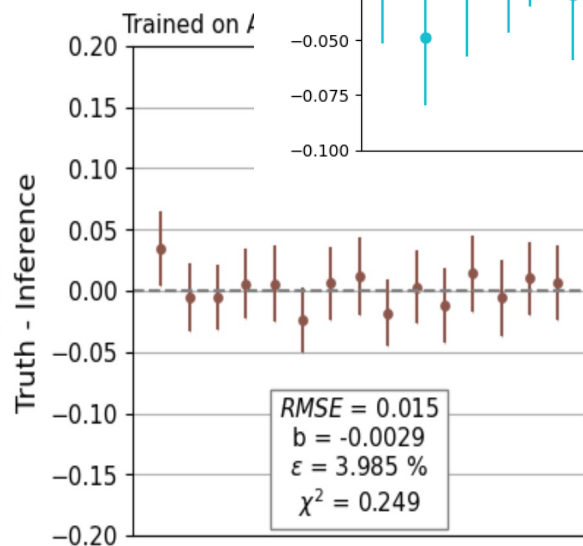
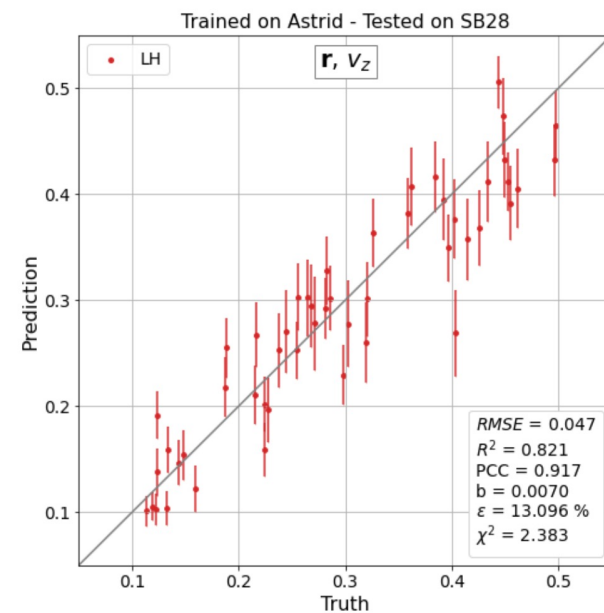
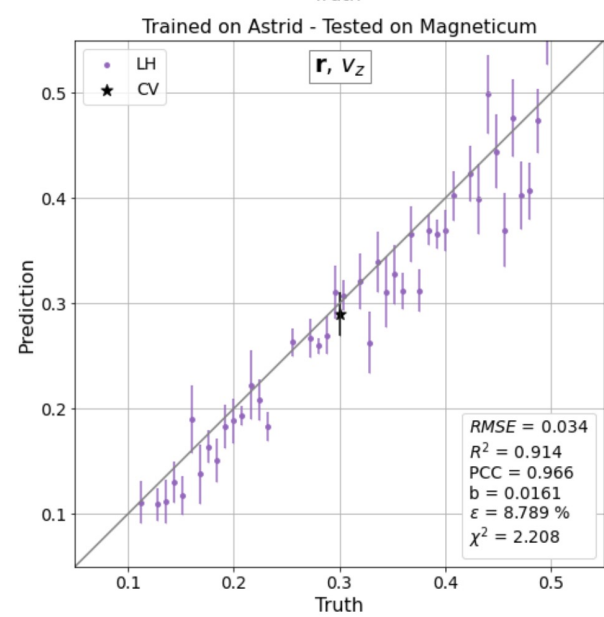
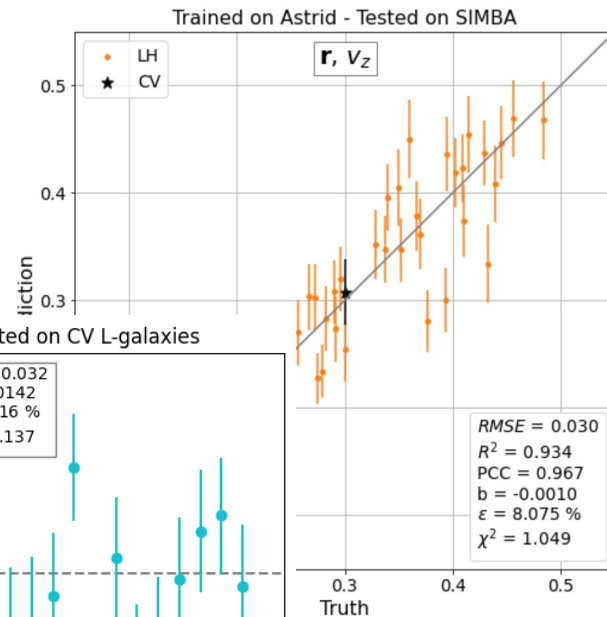
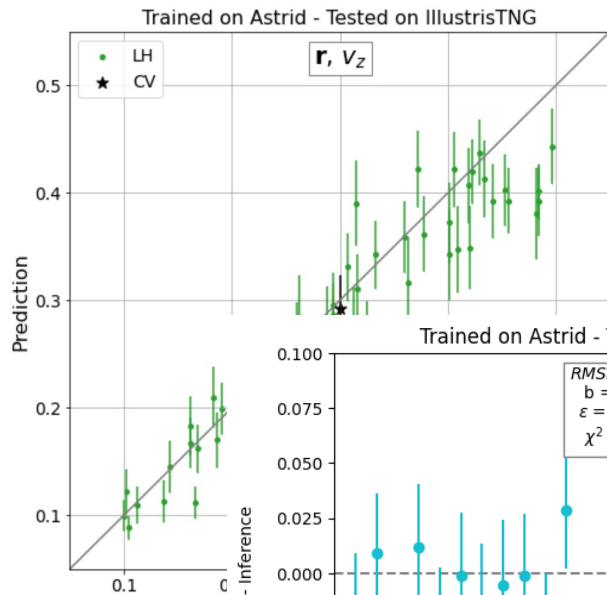
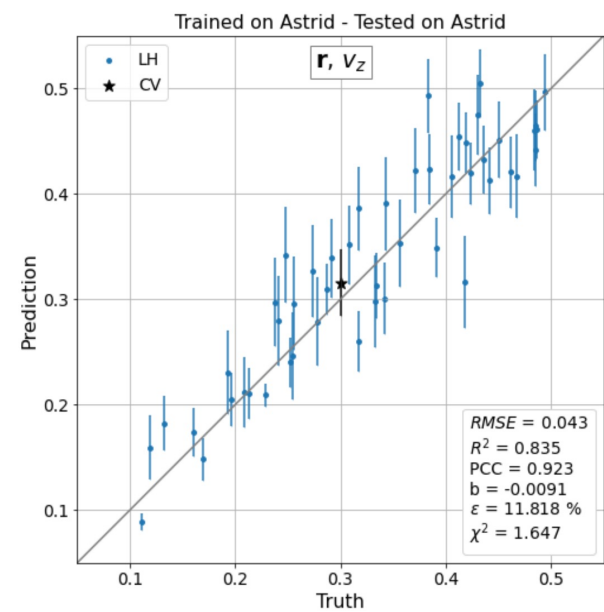
EAGLE

Astrid

Obsidian



# Robust field-level inference with GNNs



# Robust field-level inference with GNNs: Interpretability



Helen Shao  
(Princeton)

$$\mathbf{e}_{ij}^{(l+1)} = \phi_{l+1}([\mathbf{h}_i^{(l)}, \mathbf{h}_j^{(l)}, \mathbf{e}_{ij}^{(l)}])$$

$$\mathbf{h}_i^{(l+1)} = \psi_{l+1}([\mathbf{h}_i^{(l)}, \bigoplus_{j \in \mathcal{N}_i} \mathbf{e}_{ij}^{(l+1)}, \mathbf{u}])$$

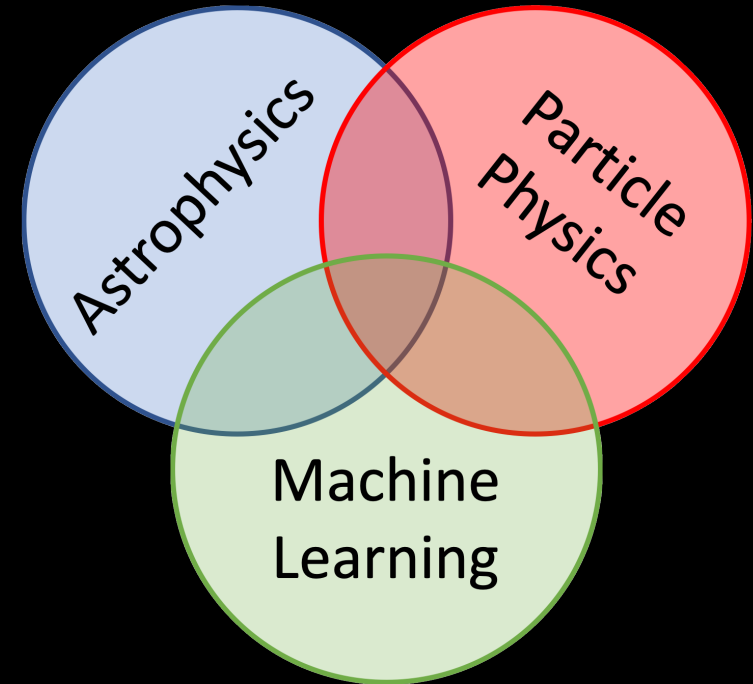
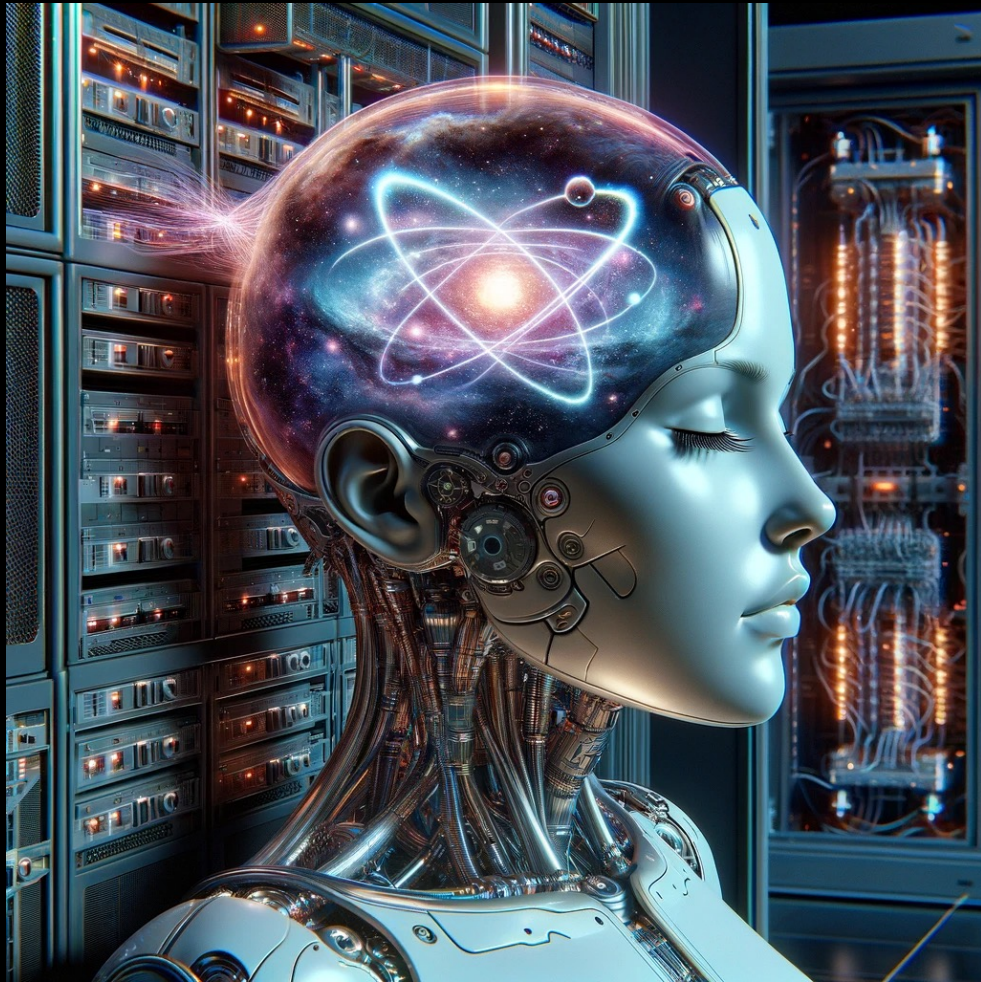
$$\mathbf{y} = \xi([\bigoplus_{i \in \mathcal{G}} \mathbf{h}_i^{(L)}, \mathbf{u}])$$

GNN Component	Formula
Edge Model: $e_1^{(1)}$	$1.32 v_i - v_j + 0.21  + 0.12(v_i - v_j) - 0.12(\gamma_{ij} + \beta_{ij} - 1.73)$
Edge Model: $e_2^{(1)}$	$ 1.62(v_i - v_j) + 0.45  + 1.98(v_i - v_j) + 0.55$
Node Model: $v_1^{(1)}$	$1.21^{v_i} (0.77^{3.29 \sum_{j \in \mathcal{N}_j} e_1^{(1)} + \sum_{j \in \mathcal{N}_j} e_2^{(1)}}) + 0.12$
Node Model: $v_1^{(1)} + v_2^{(1)}$	$0.78 - \sqrt{\log(0.16^{\sum_{j \in \mathcal{N}_j} e_2 + \sum_{j \in \mathcal{N}_j} e_1 - 0.41v_i - 1.05})} + 1.45$
Final MLP: $\mu_{\Omega_m}$	$4 \times 10^{-4} \cdot (-5.5 \sum_{i \in \mathcal{G}} v_2^{(1)} + 2.21 \sum_{i \in \mathcal{G}} v_1^{(1)} +  0.96 \sum_{i \in \mathcal{G}} v_2^{(1)} + 0.82 \sum_{i \in \mathcal{G}} v_1^{(1)} ) - 0.103$



# The DREAMS project

DaRk mattEr with AI and siMulations







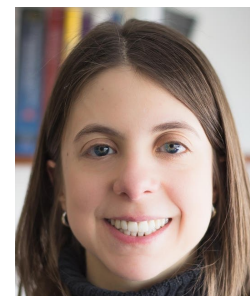
Jonah Rose  
(Florida)



Paul Torrey  
(Virginia)



Francisco Villaescusa-Navarro  
(CCA/Princeton)



Mariangela Lisanti  
(Princeton/CCA)



Sandip Roy  
(Princeton)



Tri Nguyen  
(MIT)



Kassidy Kollmann  
(Princeton)



Alex Garcia  
(Virginia)



Bonny Wang  
(Carnegie Mellon)



Belen Costanza  
(La plata)



Akaxia Cruz  
(CCA)



Nitya Kallivayalil  
(Virginia)



Andrea Caputo  
(CERN)



Mikhail Medvedev  
(Kansas/IAS/ITC)



Mark Vogelsberger  
(MIT)



Arya Farahi  
(UT Austin)



Soumyodita Karmakar  
(New Mexico)



Francis-Yan Cyr-Racine  
(New Mexico)



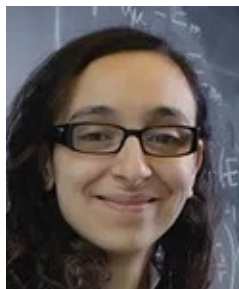
Cian Roche  
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Stephanie O'Neil  
(MIT)



Shy Genel  
(CCA/Columbia)



Lina Necib  
(MIT)



Daniel Angles-Alcazar  
(UConn)



Julian Muñoz  
(UT Austin)



Romain Teyssier  
(Princeton)

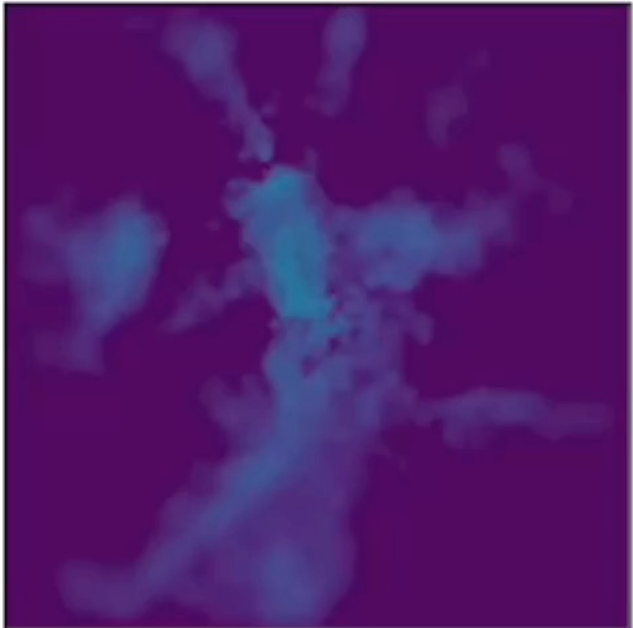
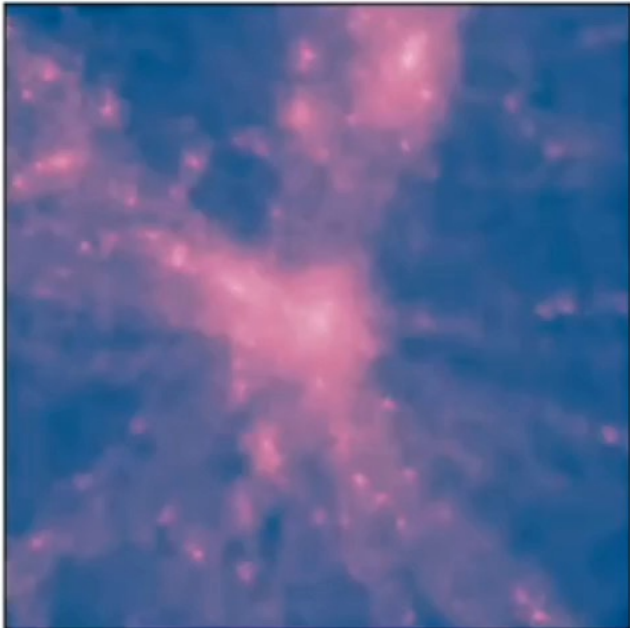


David Spergel  
(Simons Foundation)



Julianne Dalcanton  
(CCA)





Credit:  
Jonah Rose



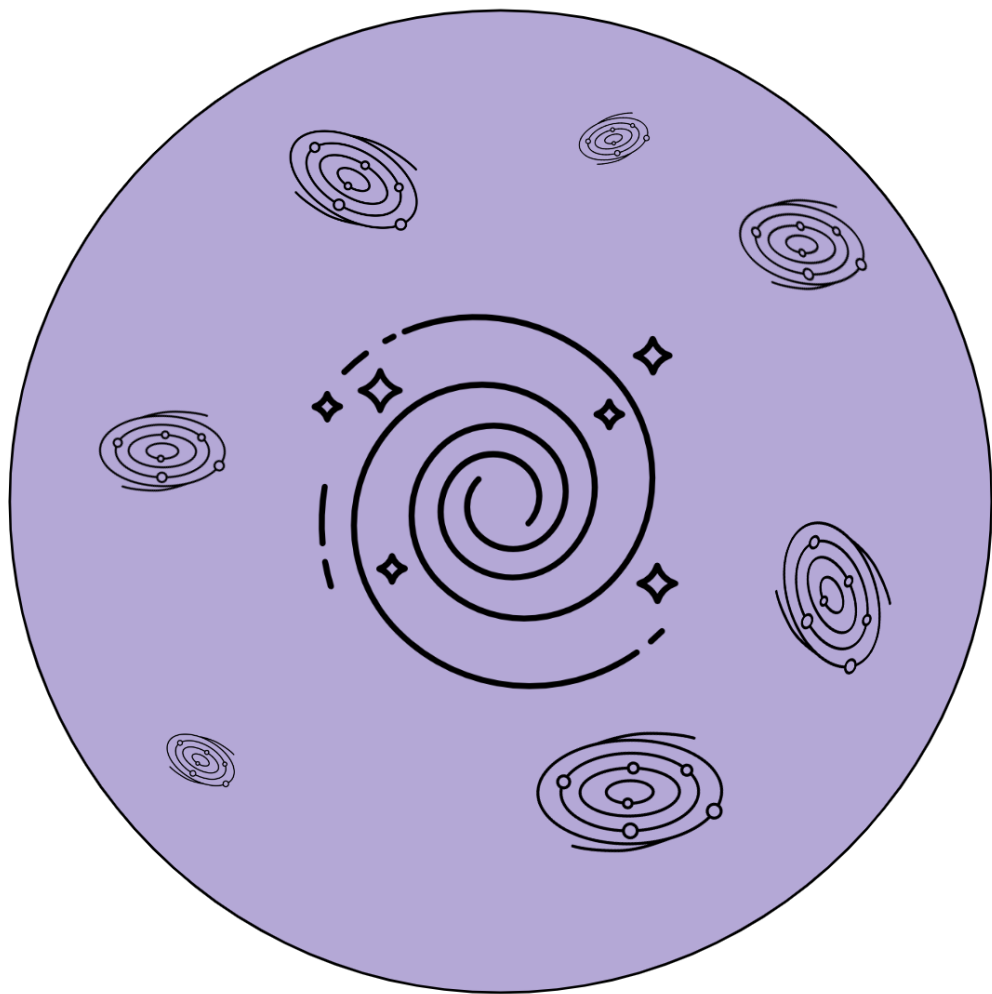


# Neuro HOD



Tri Nguyen  
(MIT)

2407.XXXXX



$$P(\{\vec{p}_i\}_{i=1,2,\dots,N} | \vec{\theta})$$

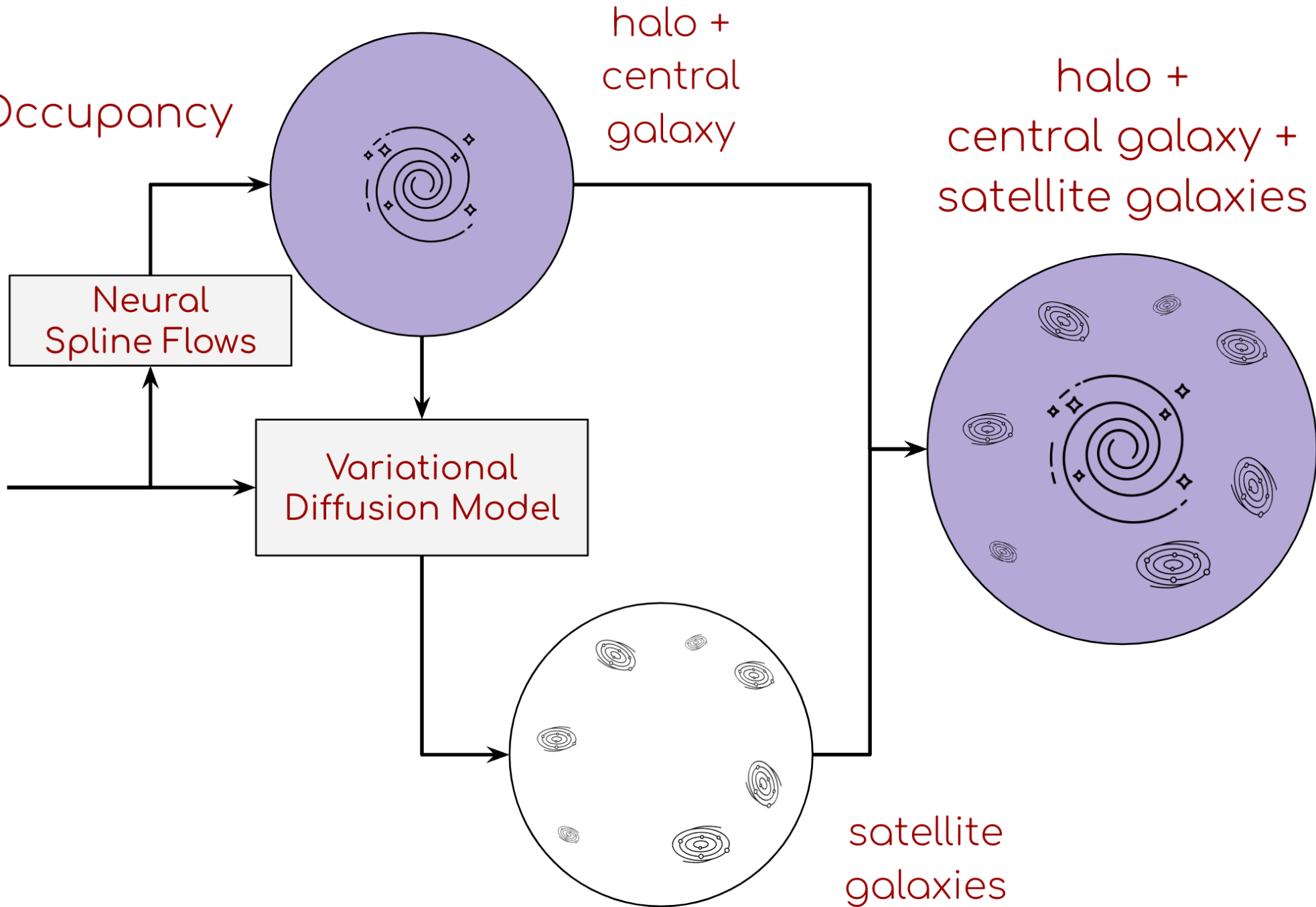
Where

$$\vec{p}_i = \{\vec{x}_i, \vec{v}_i, M_*, L_i, M_i, \dots\}$$
$$\vec{\theta} = \{\Omega_m, \Omega_b, h, \dots SN, AGN, IMF \dots\}$$

# NeHOD

Neural Halo Occupancy  
Distribution

simulation  
parameters  
(e.g. cosmology,  
astrophysics)

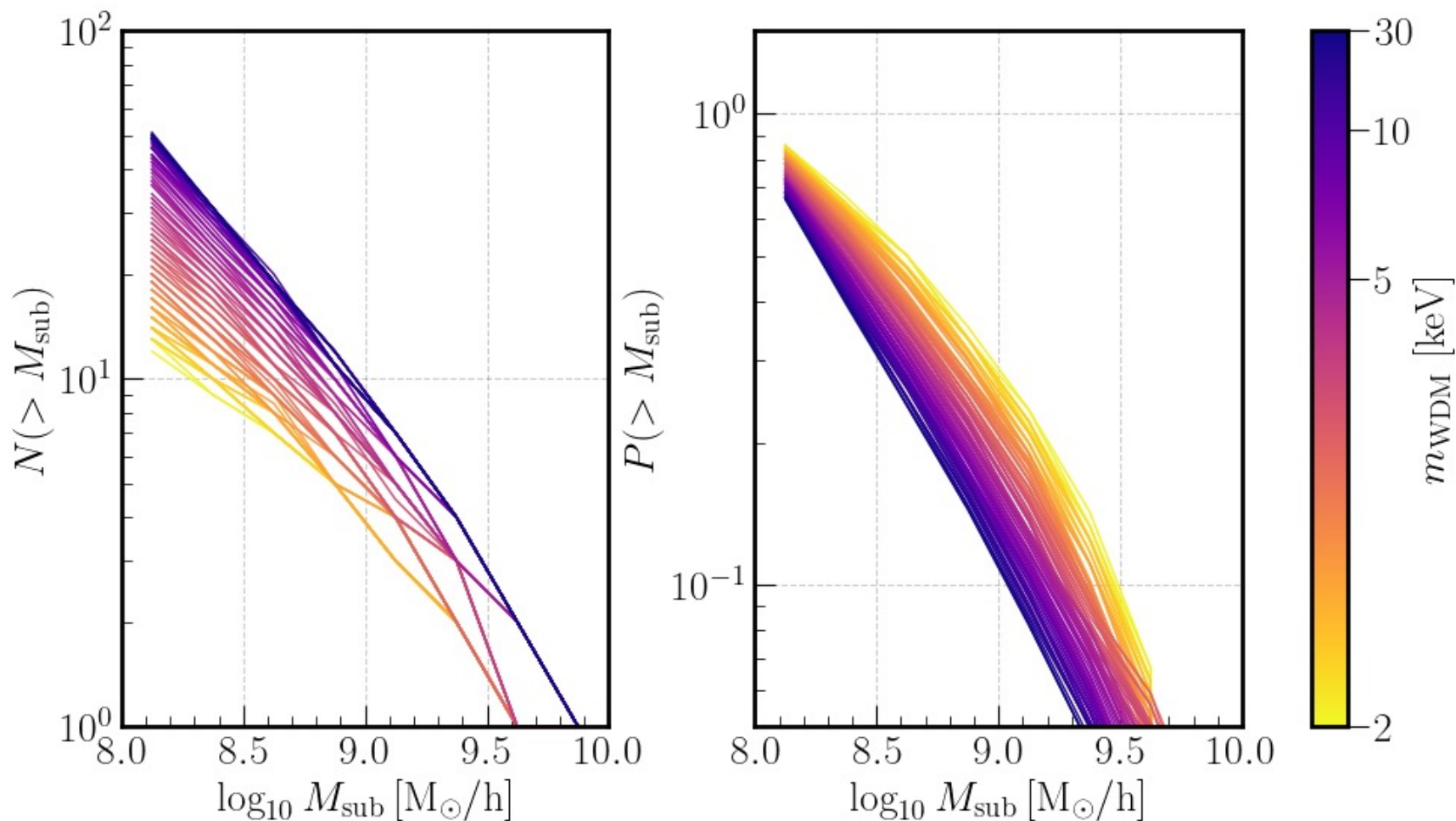




# Neuro HOD



Tri Nguyen  
(MIT)





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# Conclusions

- We want to know what are the laws and constituents of the Universe
- Traditionally, we have used analytic techniques for this
- Simulations allow us to make predictions over broader aspects
- Deep learning is a powerful tool to explore massive amounts of data

