

The First GRBAlpha and VZLUSAT-2 catalogue: gamma-ray transients and detector sensitivity

Marianna Dafčíková, Masaryk University, Czech republic

Collaborators: J. Ripa, A. Pal, N. Werner, F. Munz, M. Kolar, L. Szakszonova, M. Duriskova, N. Husarikova, et al.

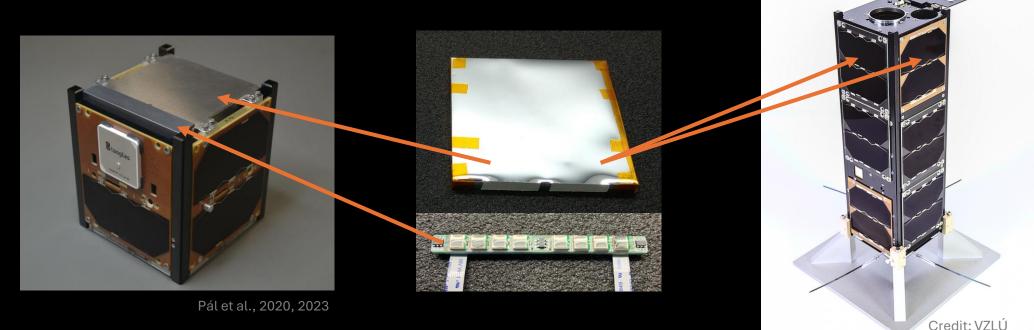


GRBAlpha

- 1U CubeSat
- Launched in March 2021
- 550 km polar orbit
- CsI(Tl) scintillator read-out by 2x4 SiPMs
- Technological experiment for CAMELOT

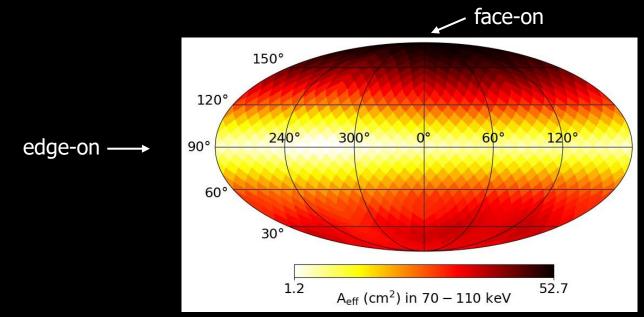
VZLUSAT-2

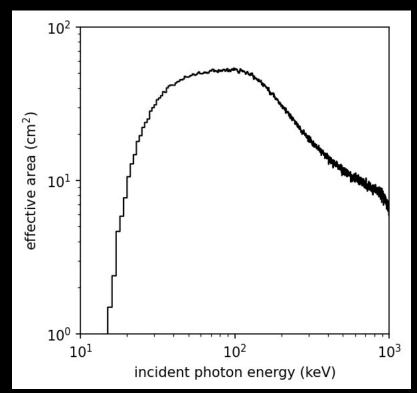
- 3U CubeSat
- Launched in January 2022
- 530 km polar orbit
- Secondary payload: 2 GRB detectors



Observations

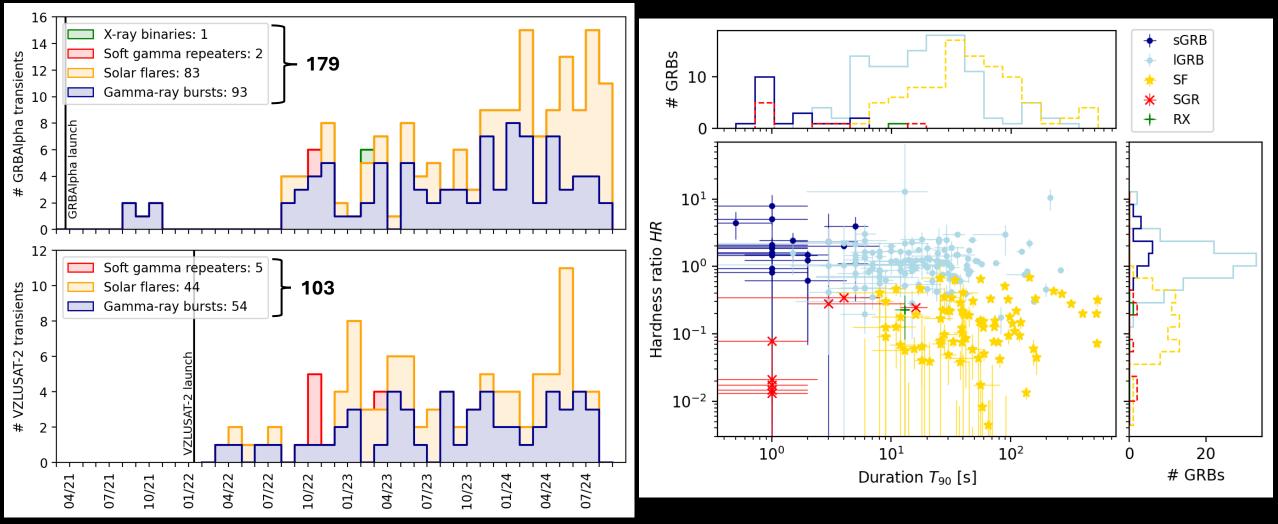
- GRBAlpha operations done by students, nonstop measurements
- VZLUSAT-2 measurements less frequent
- 0.5 and 1 s exposure time, 4 energy bands (70 950 keV)
- No trigger algorithm yet, correlation with other missions (Fermi, Swift, INTEGRAL, Konus, AGILE, CALET, GECAM, AstroSat, FRB + GW triggers)
- 2 detections/week -> 250 detections/year for a constellation





Confirmed detections

https://monoceros.physics.muni.cz/hea/GRBAlpha/ https://monoceros.physics.muni.cz/hea/VZLUSAT-2/



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- GRB 221009A (+ GRB 230307A): peak flux without saturation

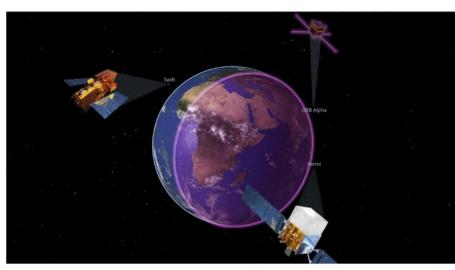
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News By Tereza Pultarova published March 29, 2023

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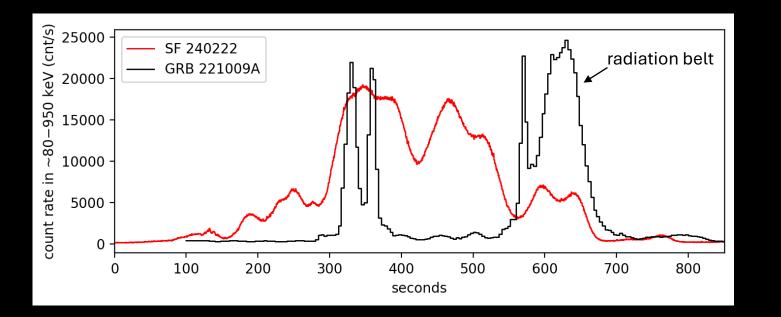
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Tiny cubesat called GRBAlpha played a key role in determining the peak intensity of the brightest gammaray burst ever seen. (Image credit: Francis Reddy/NASA Goddard, University of Maryland)

- GRB 230709B and GRB 230709C: 42 minutes apart
- GRB 231215A: most distant GRB at z = 2.305 (**10.8 Gyr**)
- GRB 221009A (+ GRB 230307A): peak flux without saturation
- X6 class solar flare on 2024-02-22: nearly as bright as the BOAT



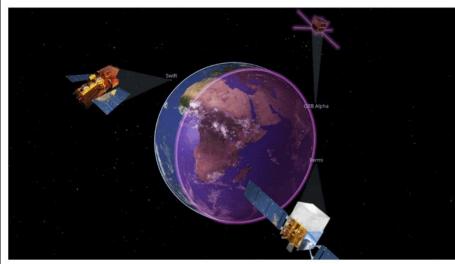
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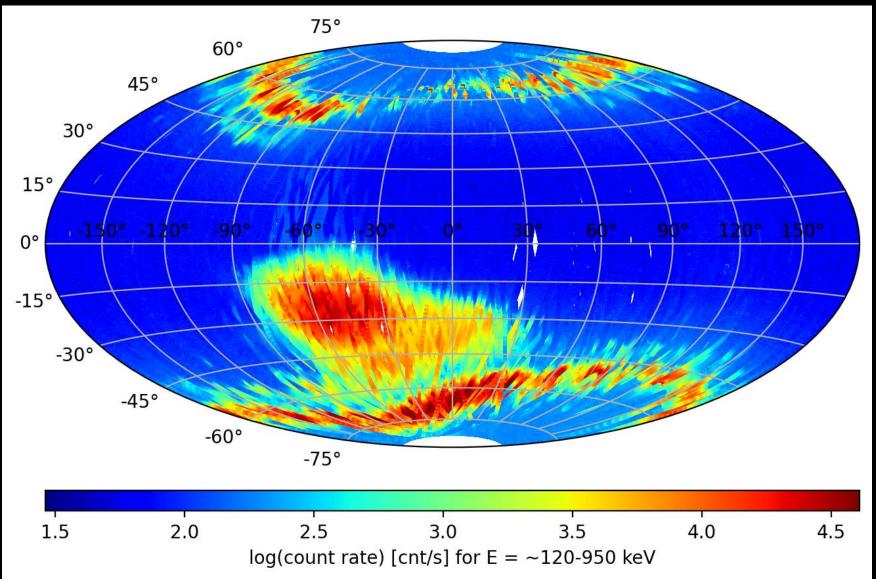
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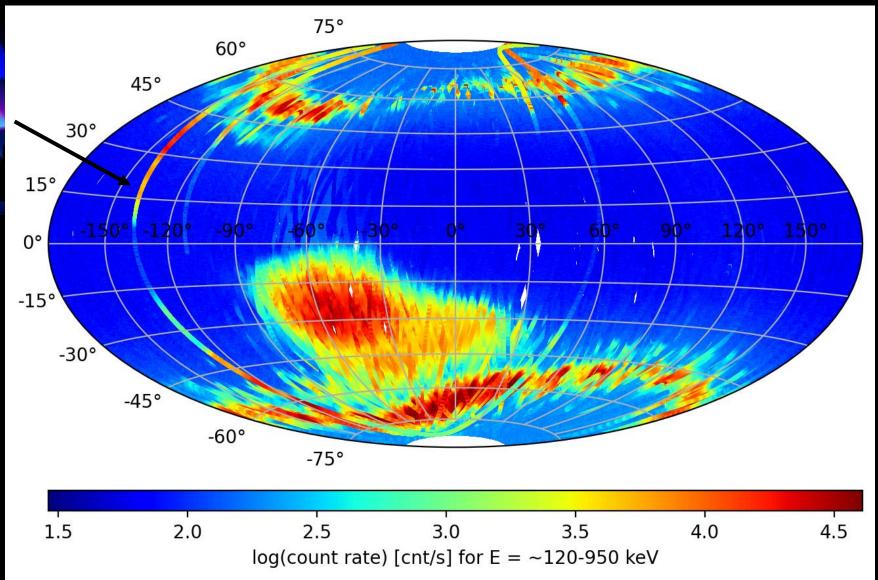
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LEO environment

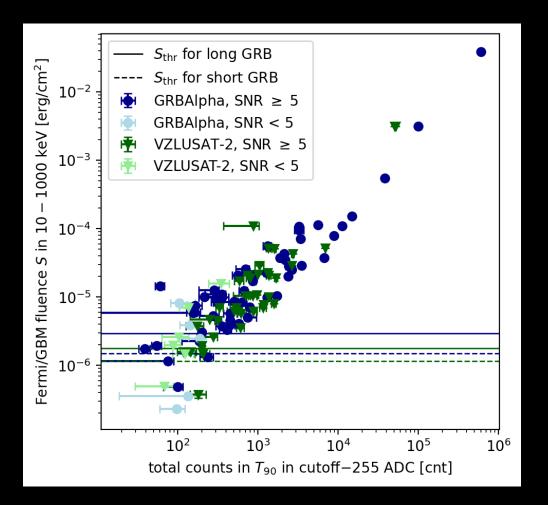


LEO environment

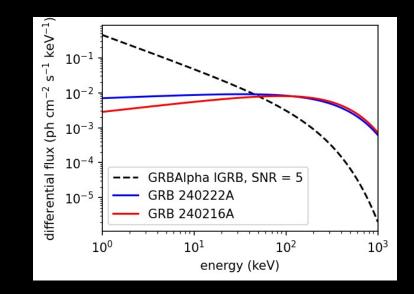
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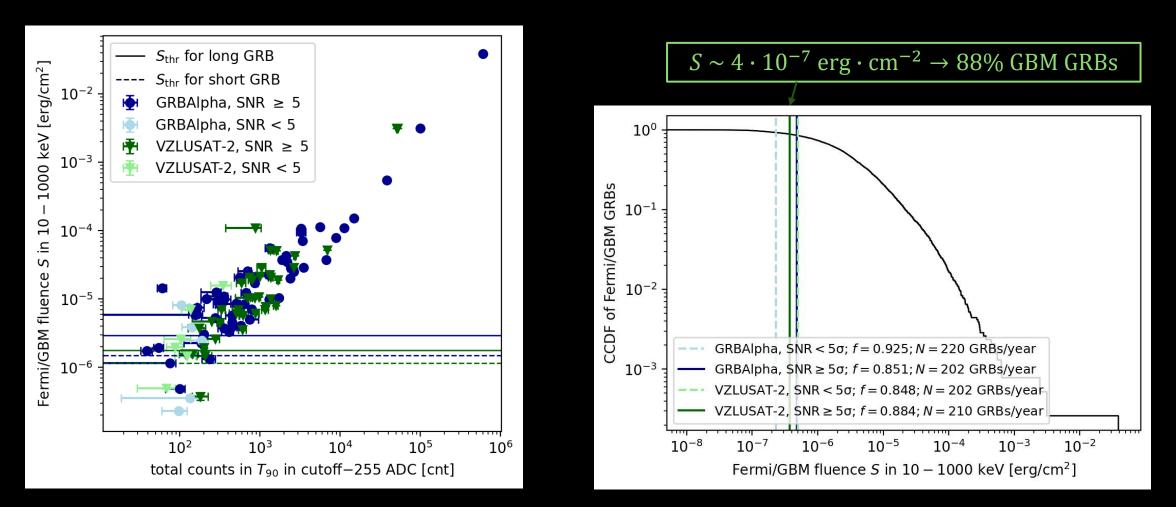
Cross-correlation with Fermi/GBM



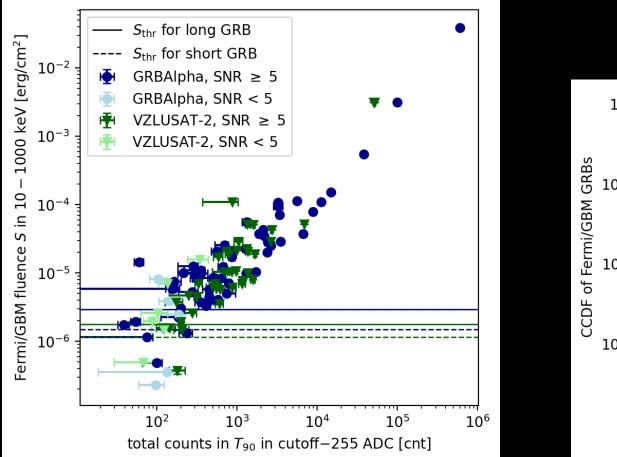
- Unknown GRBAlpha orientation
- No detector response matrix for VZLUSAT-2
- $S_{thr} = 5\sigma$ detection of a typical GRB
- Significant GRBs with $S < S_{thr}$ were harder

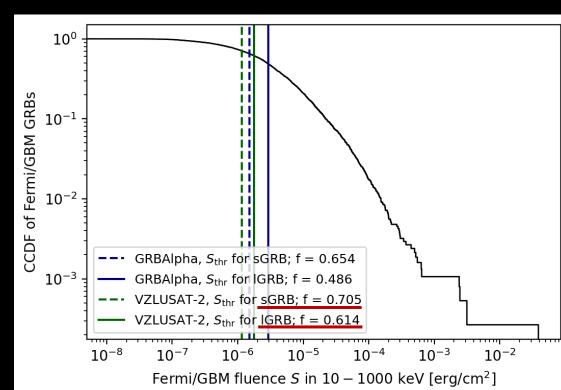


Empirical sensitivity



Theoretical sensitivity





Summary & future plans

- GRBAlpha: 179 detections in 3+ years
- VZLUSAT-2: 103 detections in 2+ years
- Nonstop measurements: 2 detections/week
- Feasible detection of 60 85% of all Fermi/GBM GRBs
- Sensitivity higher for short GRBs -> advantage in search for GW counterparts
- LEO background monitoring, SiPM degradation



More data!

 \rightarrow GRBBeta (2U CubeSat) launched July 9, 2024 More detections and discoveries!

 \rightarrow Trigger algorithm

