



UNIVERSITÀ  
DEGLI STUDI  
DI PADOVA



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# Towards a TeV blazar sequence and its physical interpretation

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Ilaria Viale

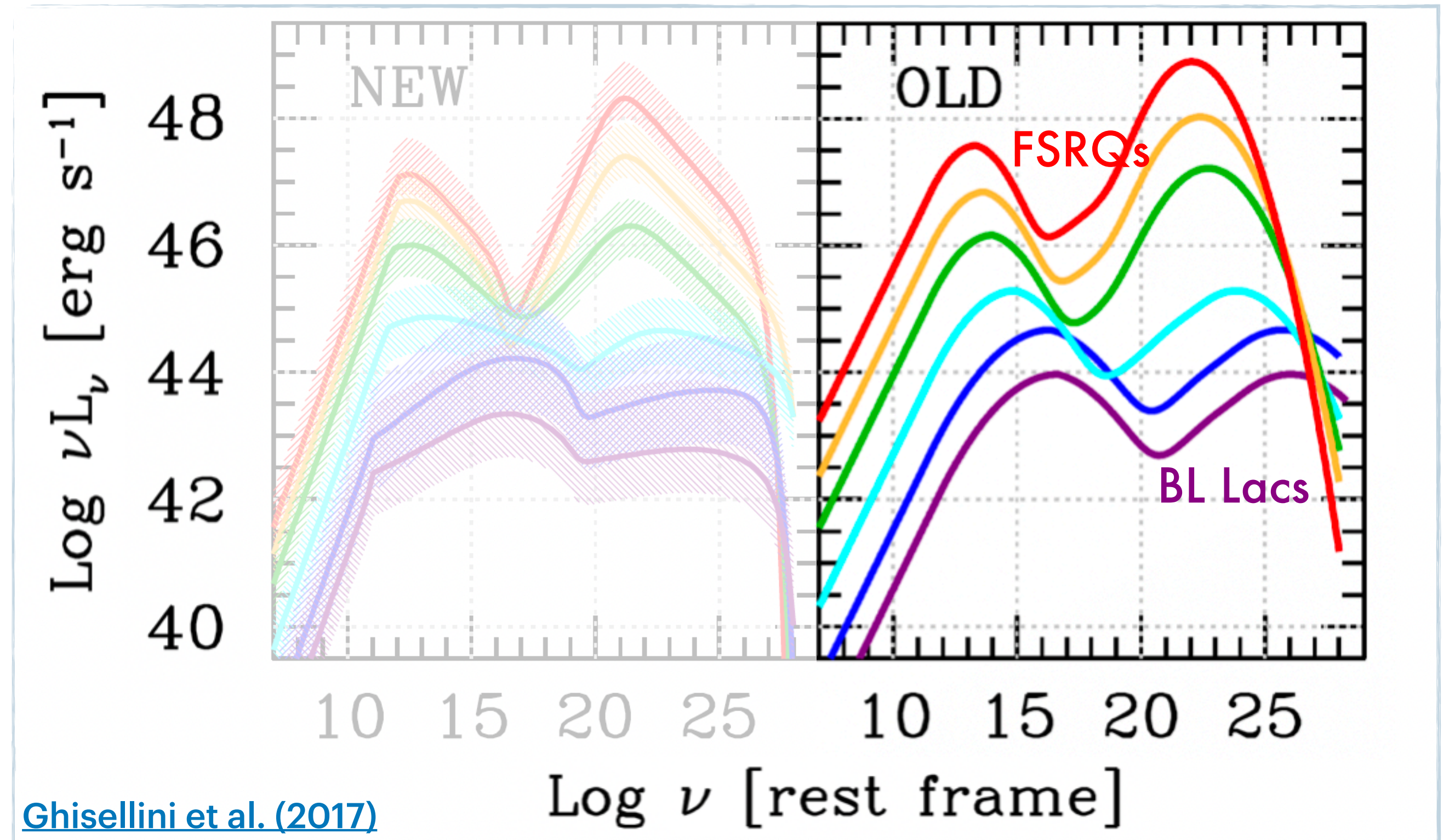
Gamma 2024

Milano, 2 - 6 Sep 2024

# The blazar sequence

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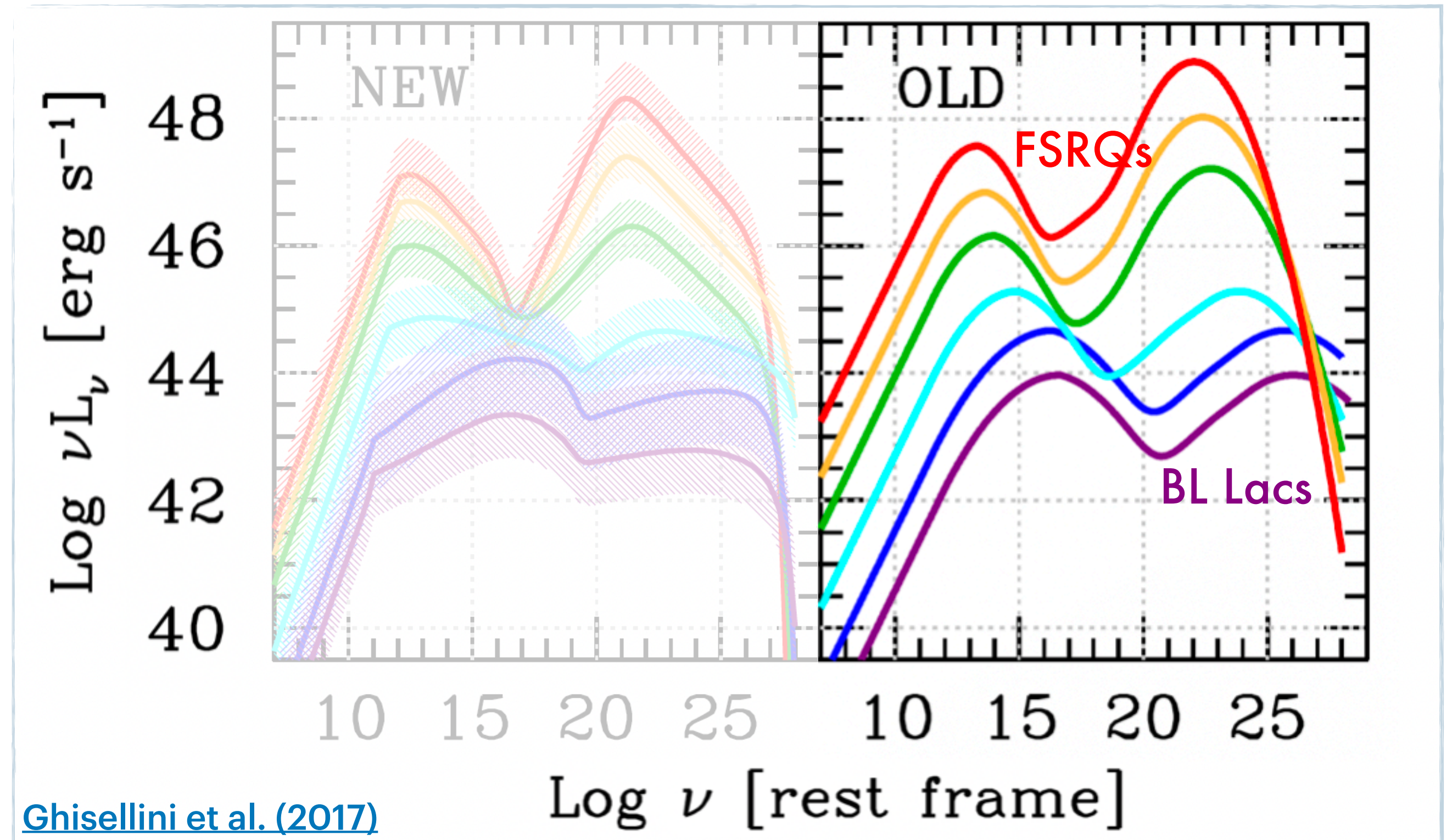
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- Correlation between peak frequencies of bumps
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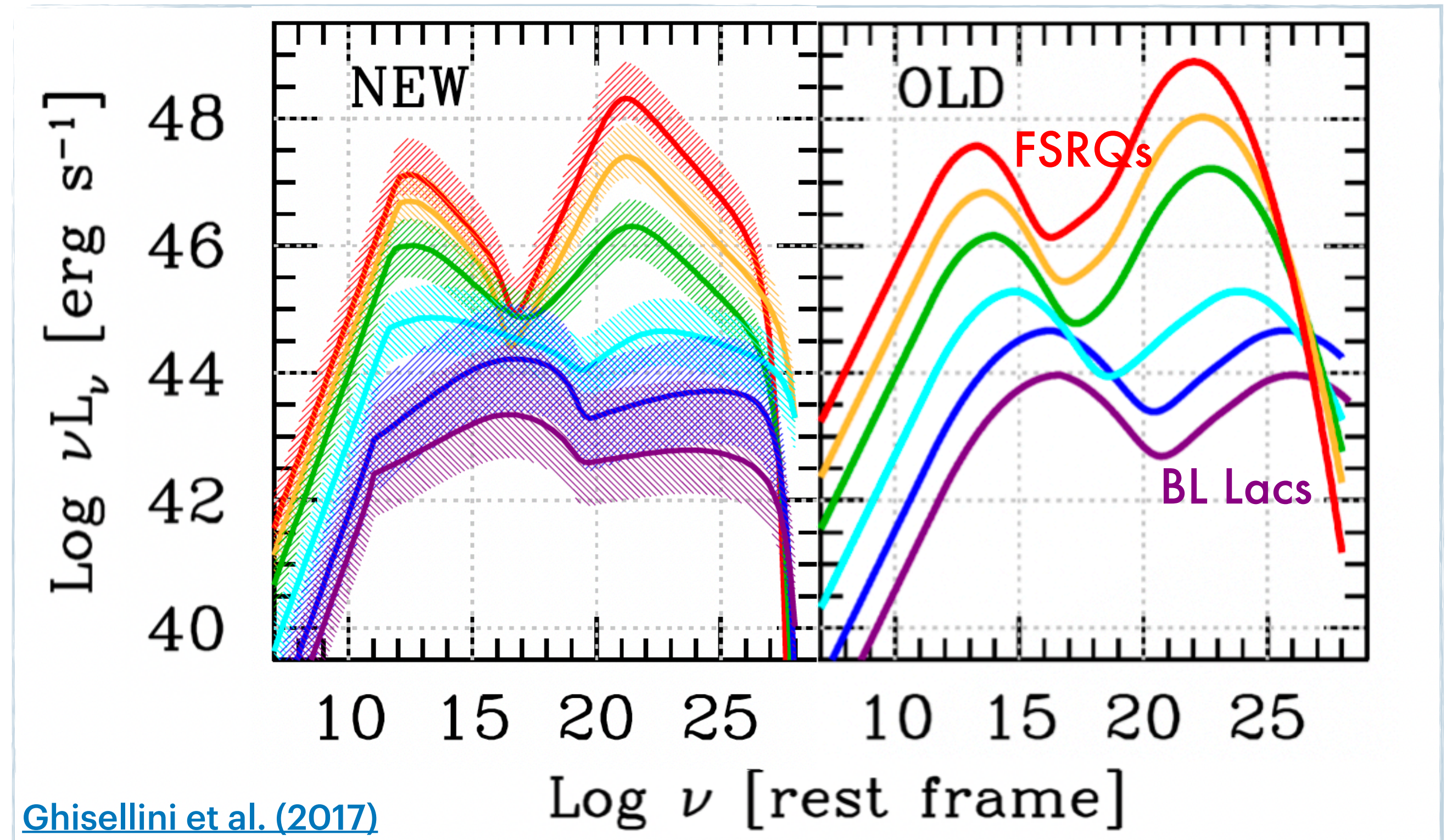
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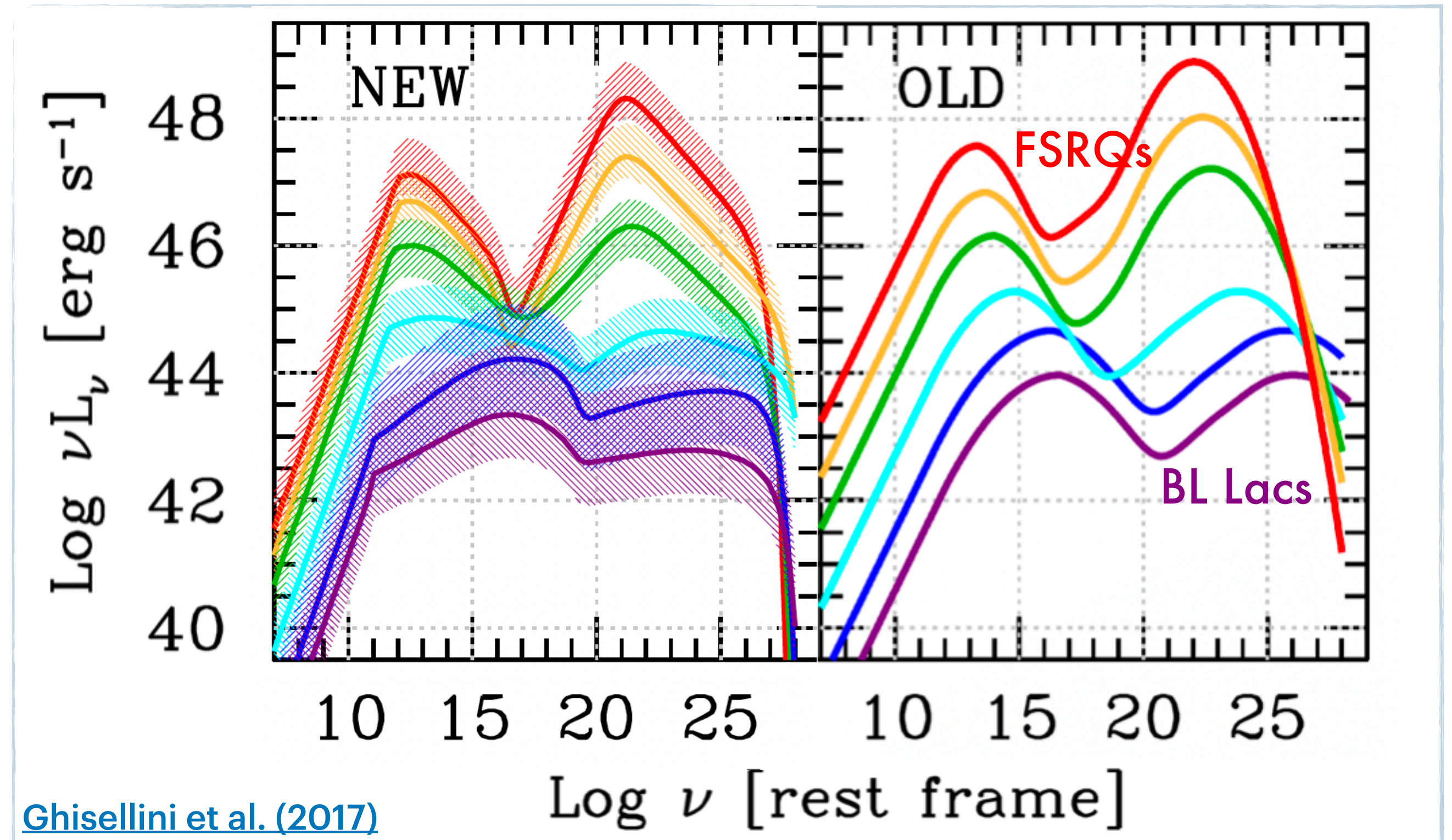
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- FSRQs: stable synchrotron peak frequency
- BL Lacs: peaks correlate with luminosity

# First studies of a TeV blazar sequence

## TeV-detected sources and the sequence:

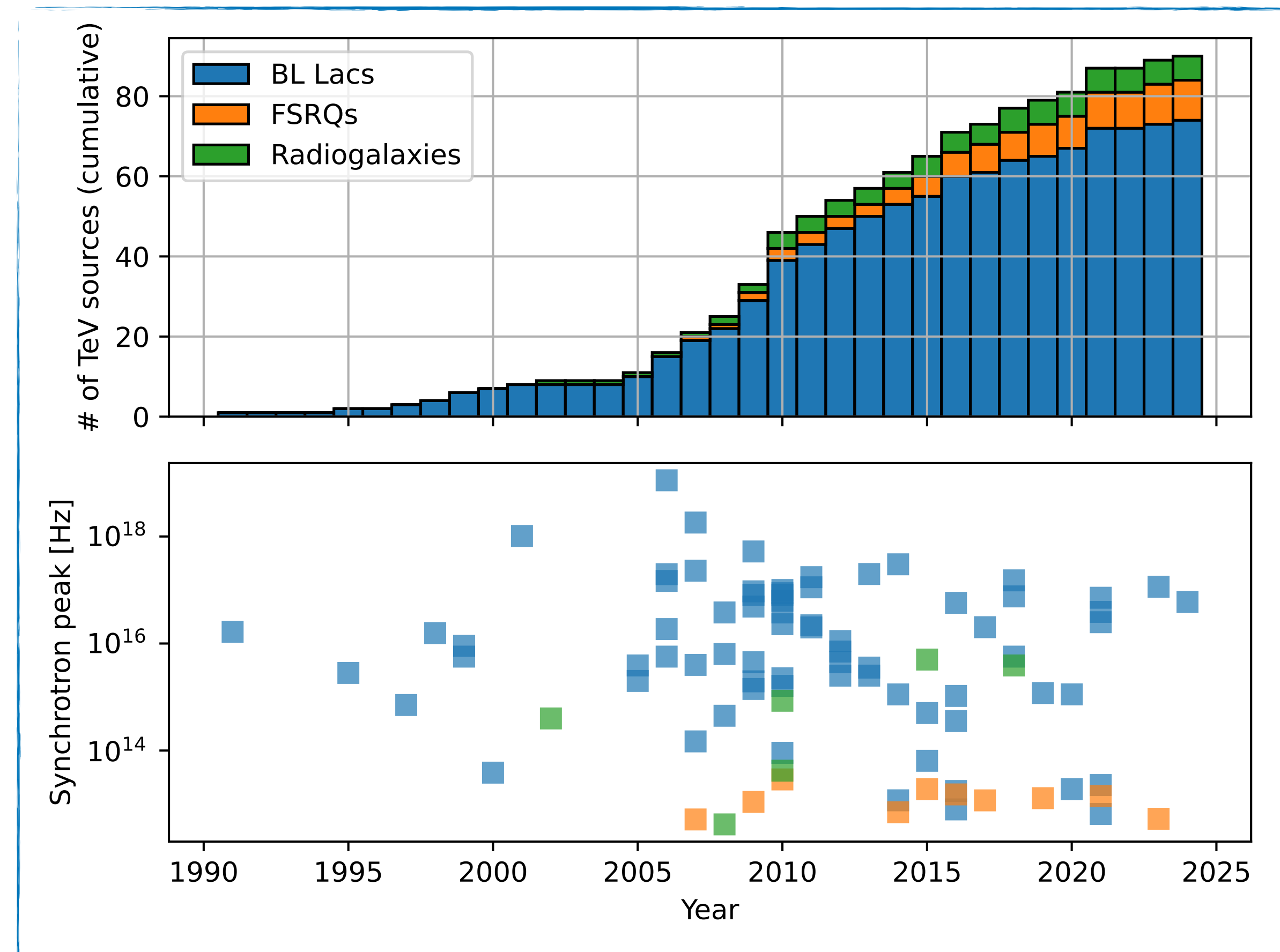
[Prandini & Ghisellini \(2022\)](#)

- 81 TeV blazars from TeVCat (at Jan 2022)
- Same luminosity bins as in 2017 sequence
- No strong differences in SED wrt Fermi blazars
- Larger X-ray luminosity than in Fermi blazars

## Quiescent vs flaring activity in TeV blazars:

[Ouyang et al. \(2023\)](#)

- 48 sources in quiescent state, 21 in flaring state
- Anti-correlation between luminosity and  $\nu_{syn}$  present only during flares



# This work

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## Goal

Which is the role of TeV sources in the blazar sequence?

Is the blazar sequence physically driven?

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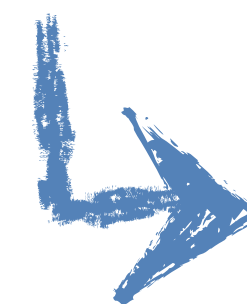
Which is the role of TeV sources in the blazar sequence?

Is the blazar sequence physically driven?

## To do it we ...

Concentrated only on TeV BL Lac sources

Modeled their emission in framework of Synchrotron Self Compton model

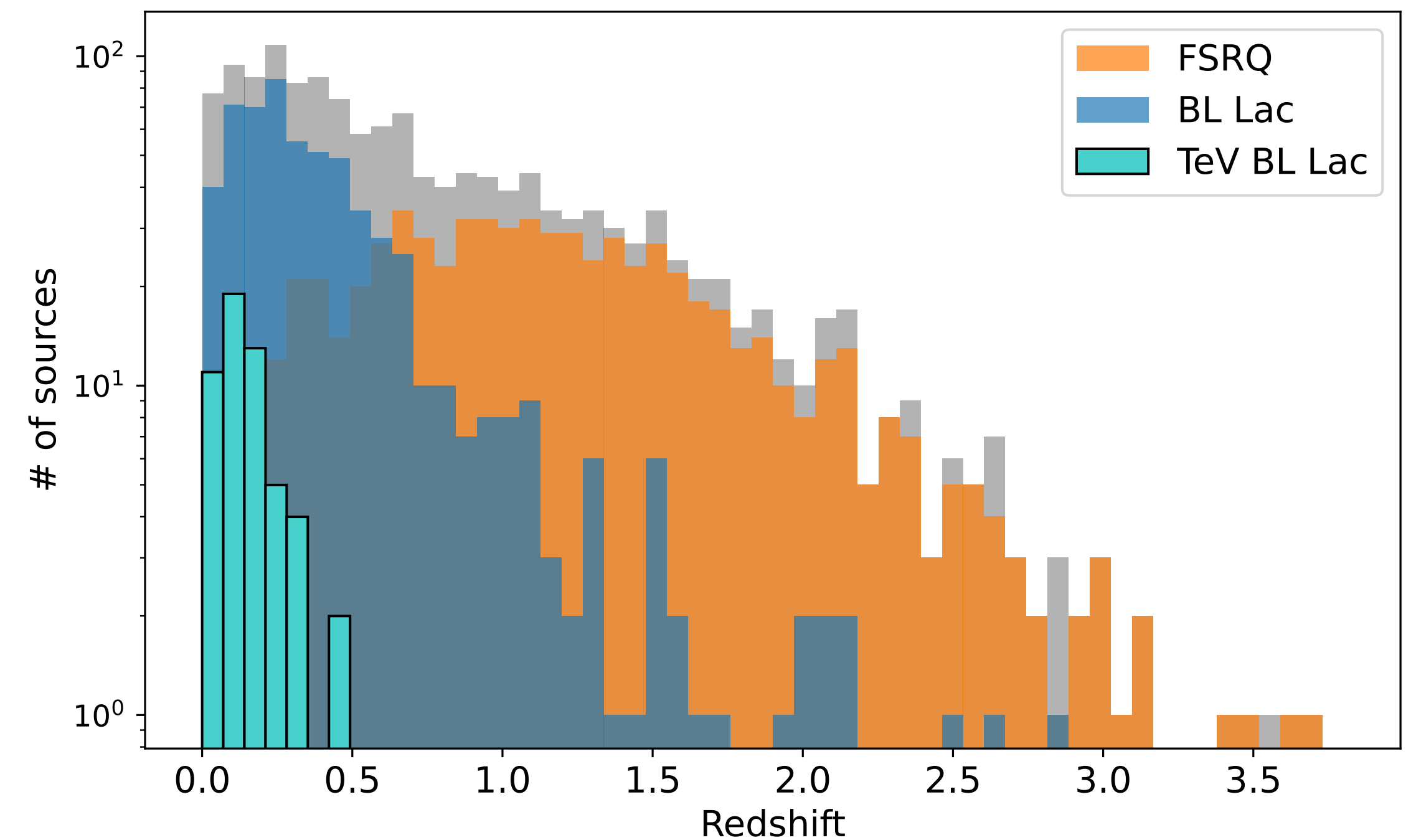
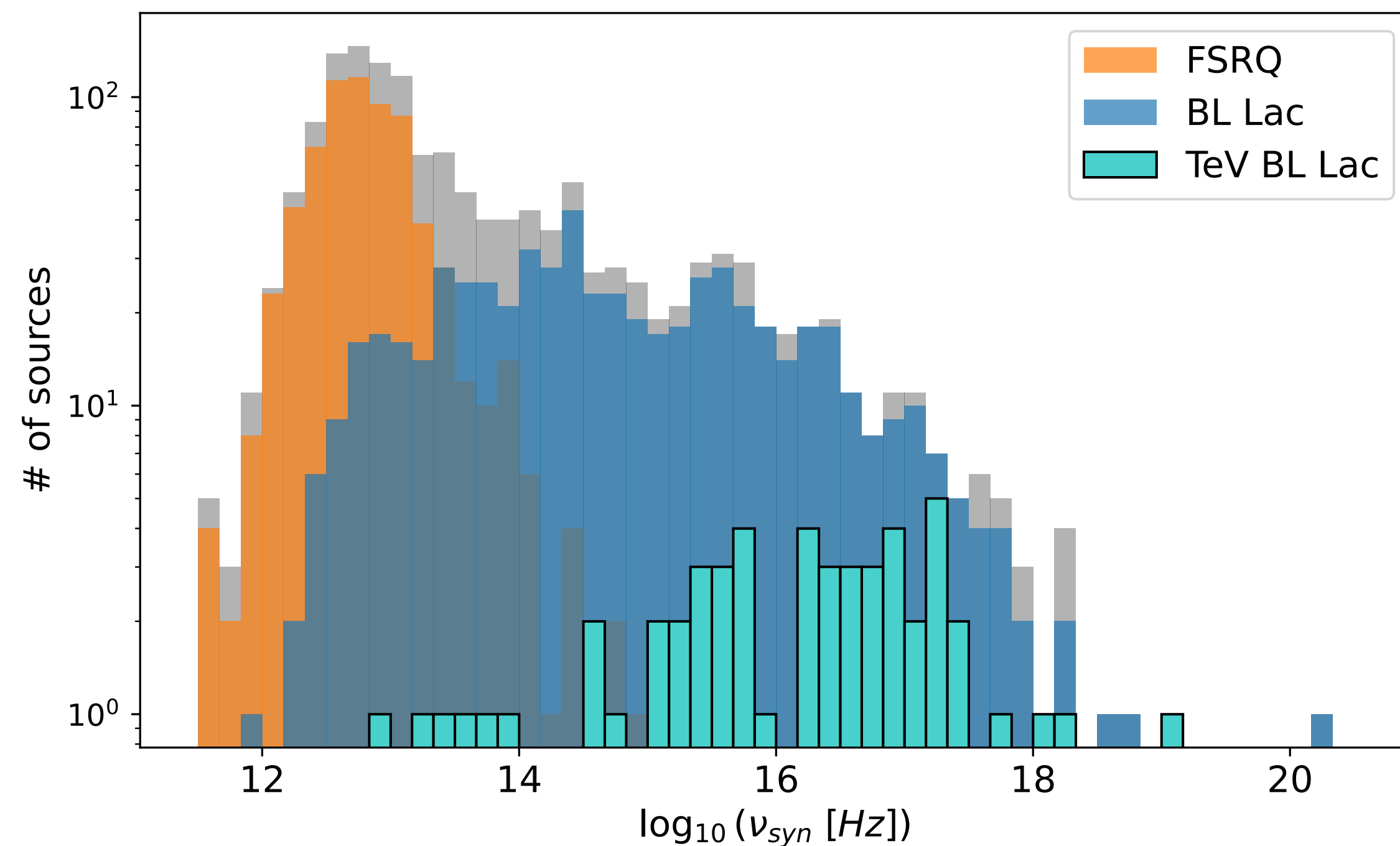


Is there a trend in the model parameters?



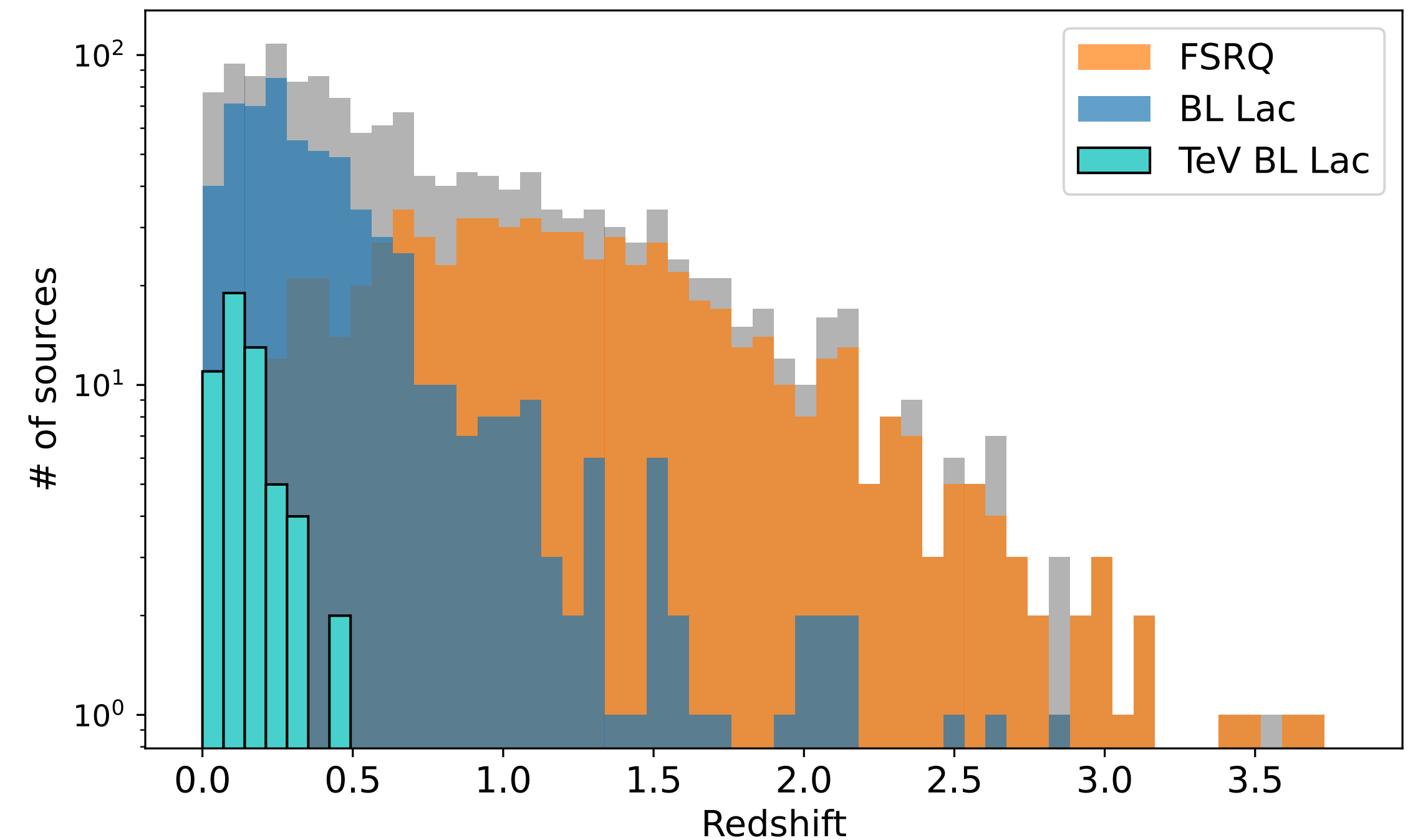
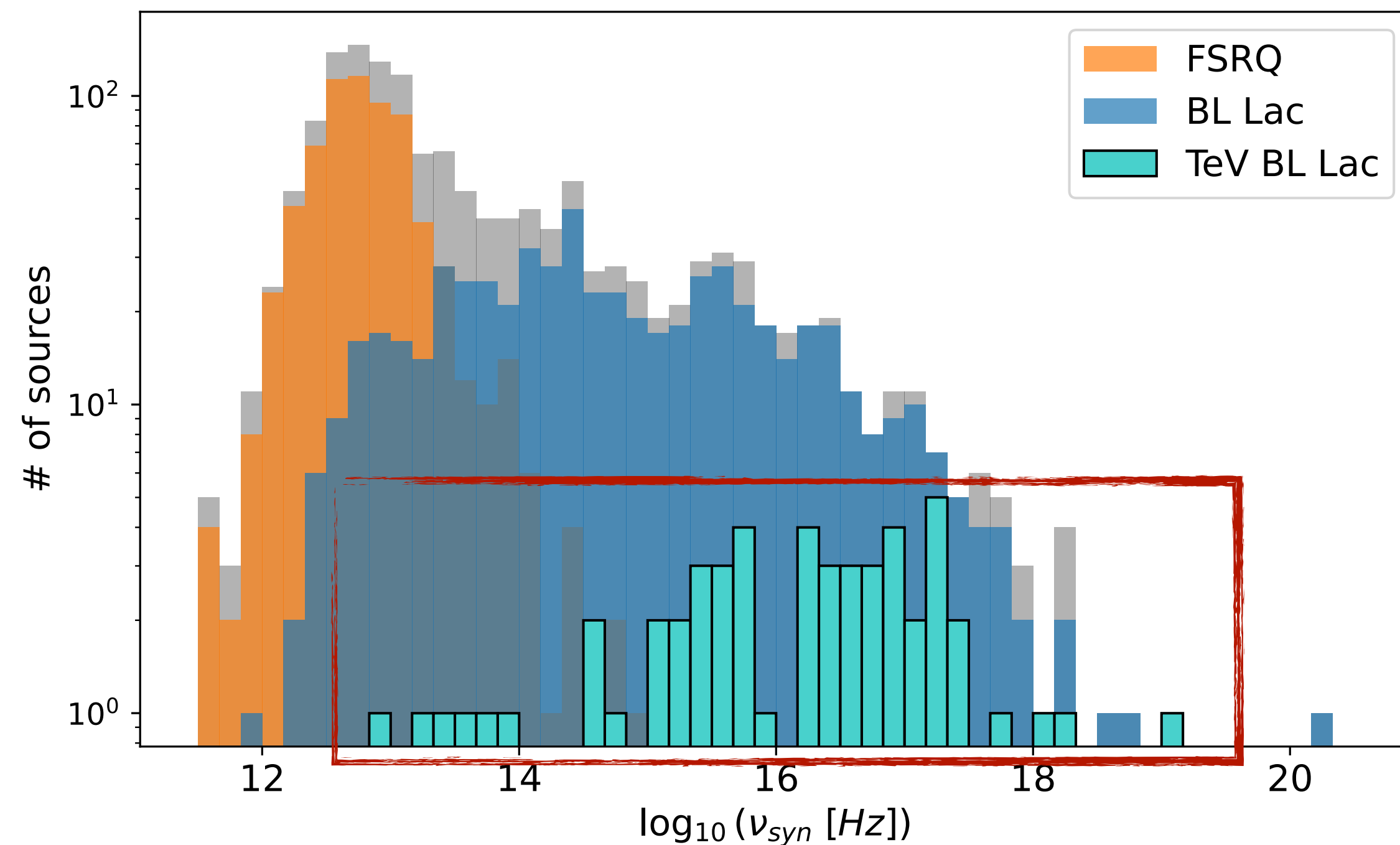
# Our sample

- 84 TeV-detected blazars at TeVCat to date, data from 4LAC-DR2
- Concentrate only on BL Lacs with known redshift → **total of 56 sources**
- $\nu_{syn}$  bins used instead of luminosity bins
- Selection of one representative source for each bin

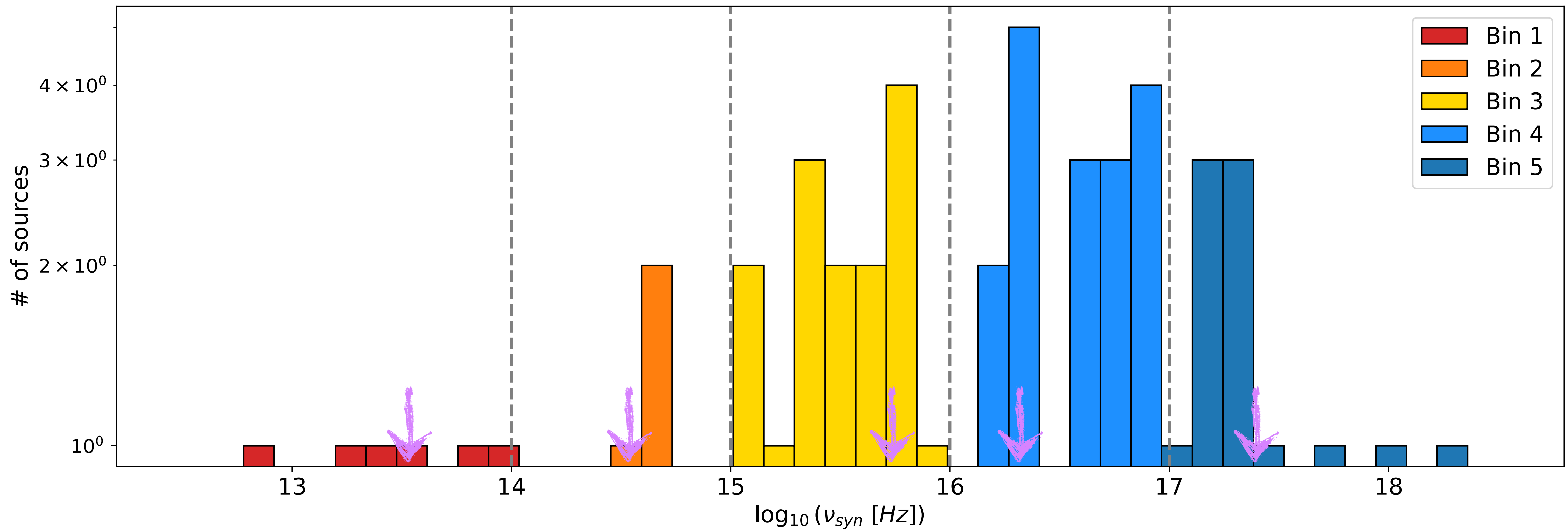
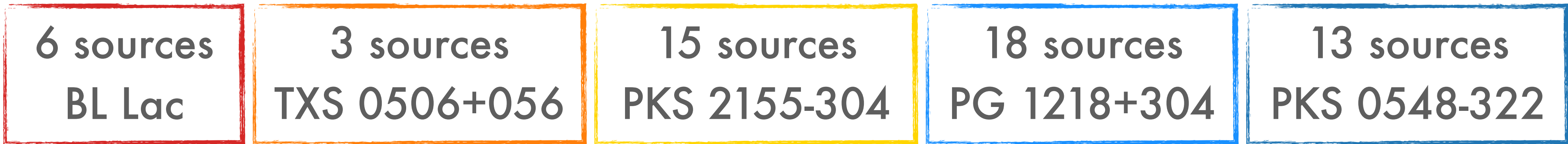


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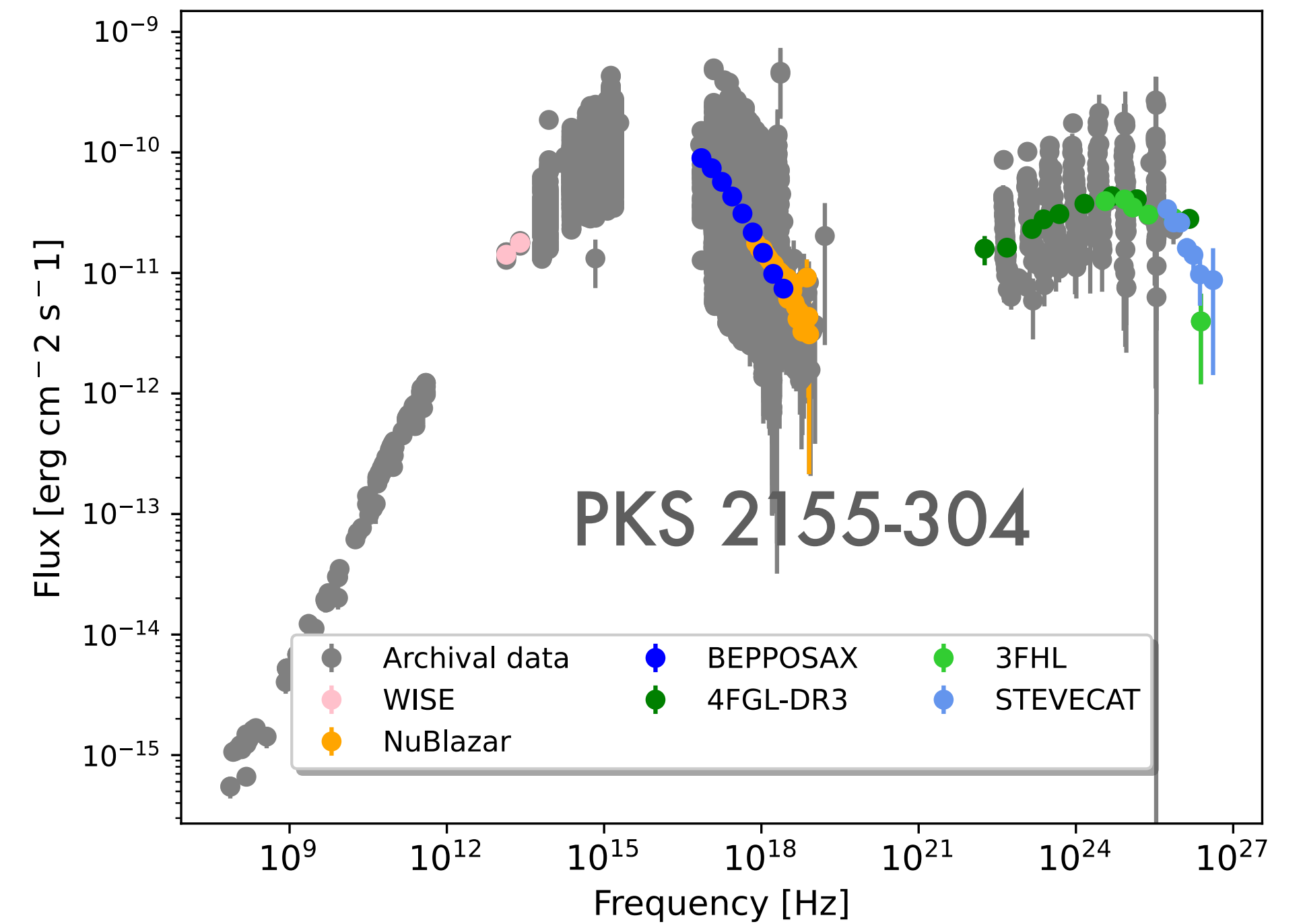


# Data selection

Selection done source-wise

For each selected source:

- Data retrieved from *MMDC + STeVEC*  
[Sahakyan et al. \(2024\), in prep.](#) [Grèaux et al. \(2023\)](#)
- Average state of activity

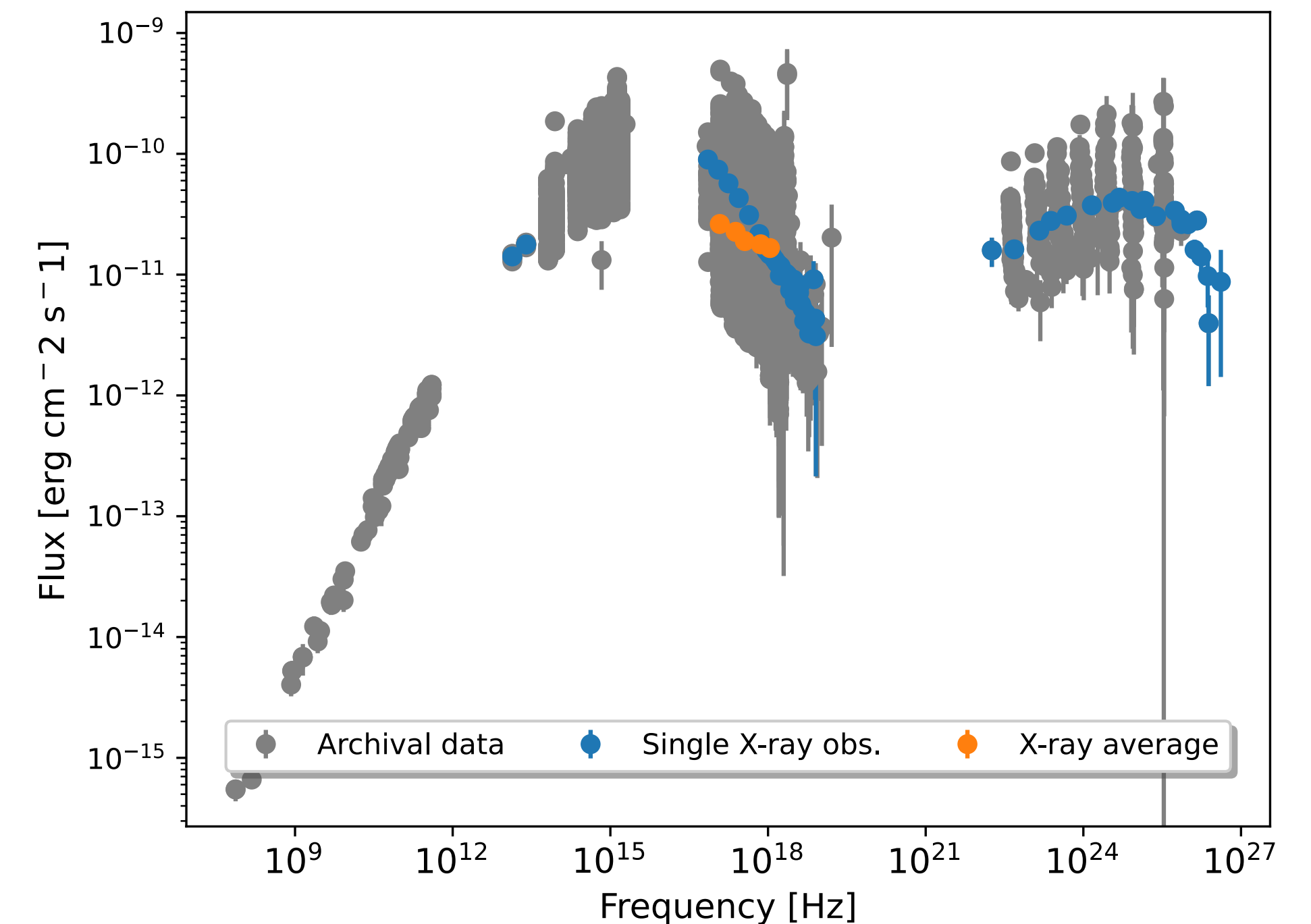
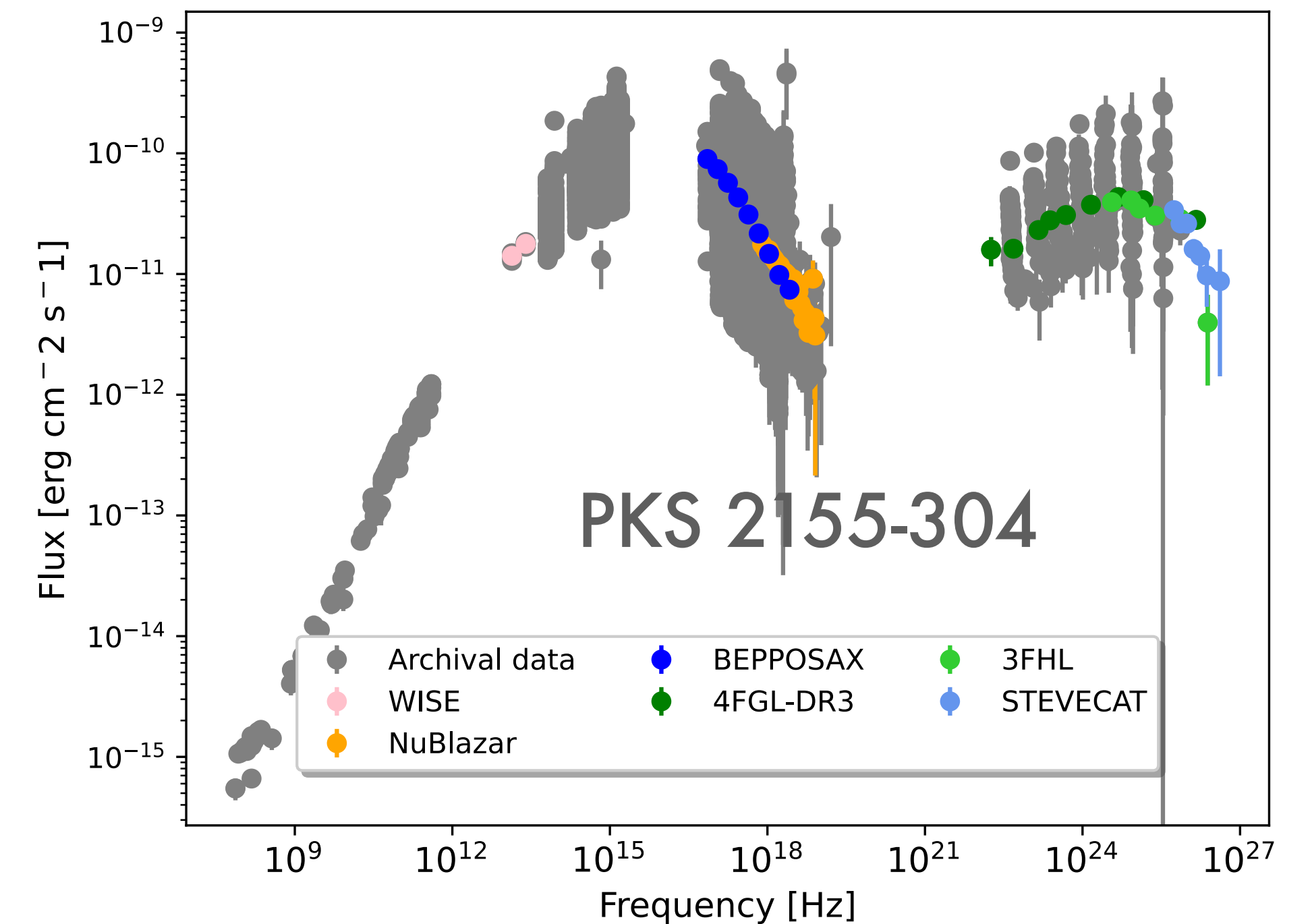


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- Average state of activity
- Large **spread** in **X-ray** data
- Two selections:
  - Single observations
  - Average data from catalog *OUSXB*  
[Giommi et al. \(2019\)](#)



# Model

Selected sources modeled with **Synchrotron Self Compton model**

- **Aim** of finding trend in model parameters describing the sequence

Concentrated only on HBL as starting point,  $\nu_{syn} > 1 \times 10^{15}$  Hz

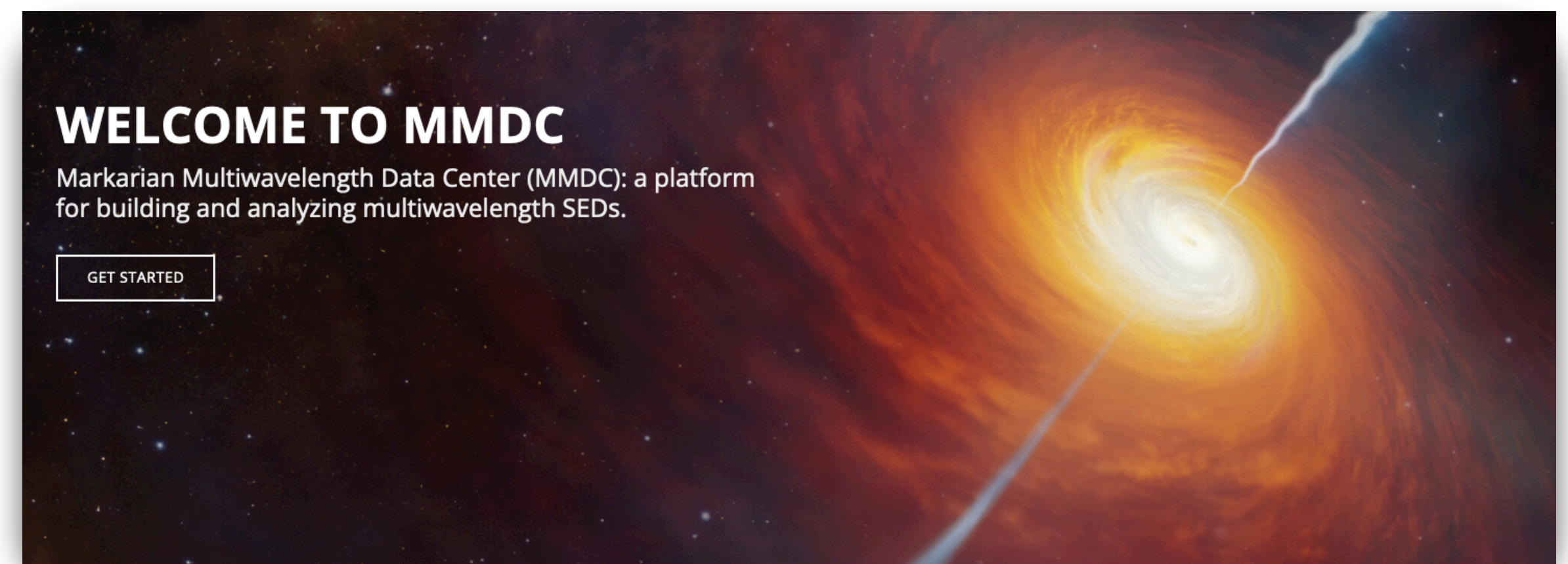
- Difficult to describe sources with lower  $\nu_{syn}$  with pure SSC

Modeling software based on

**Convolutional Neural Network:**

[Bégué et al. \(2023\)](#)

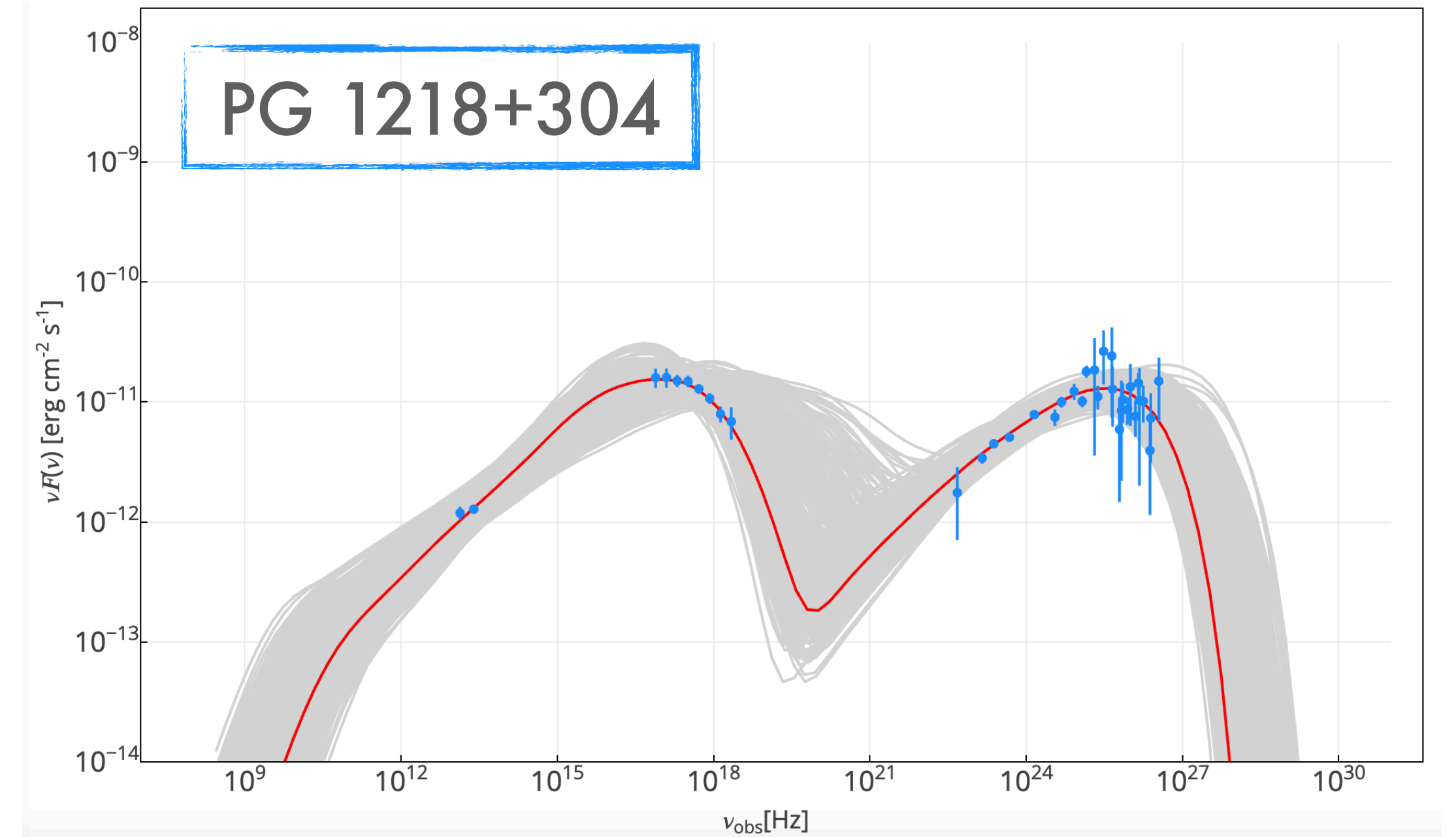
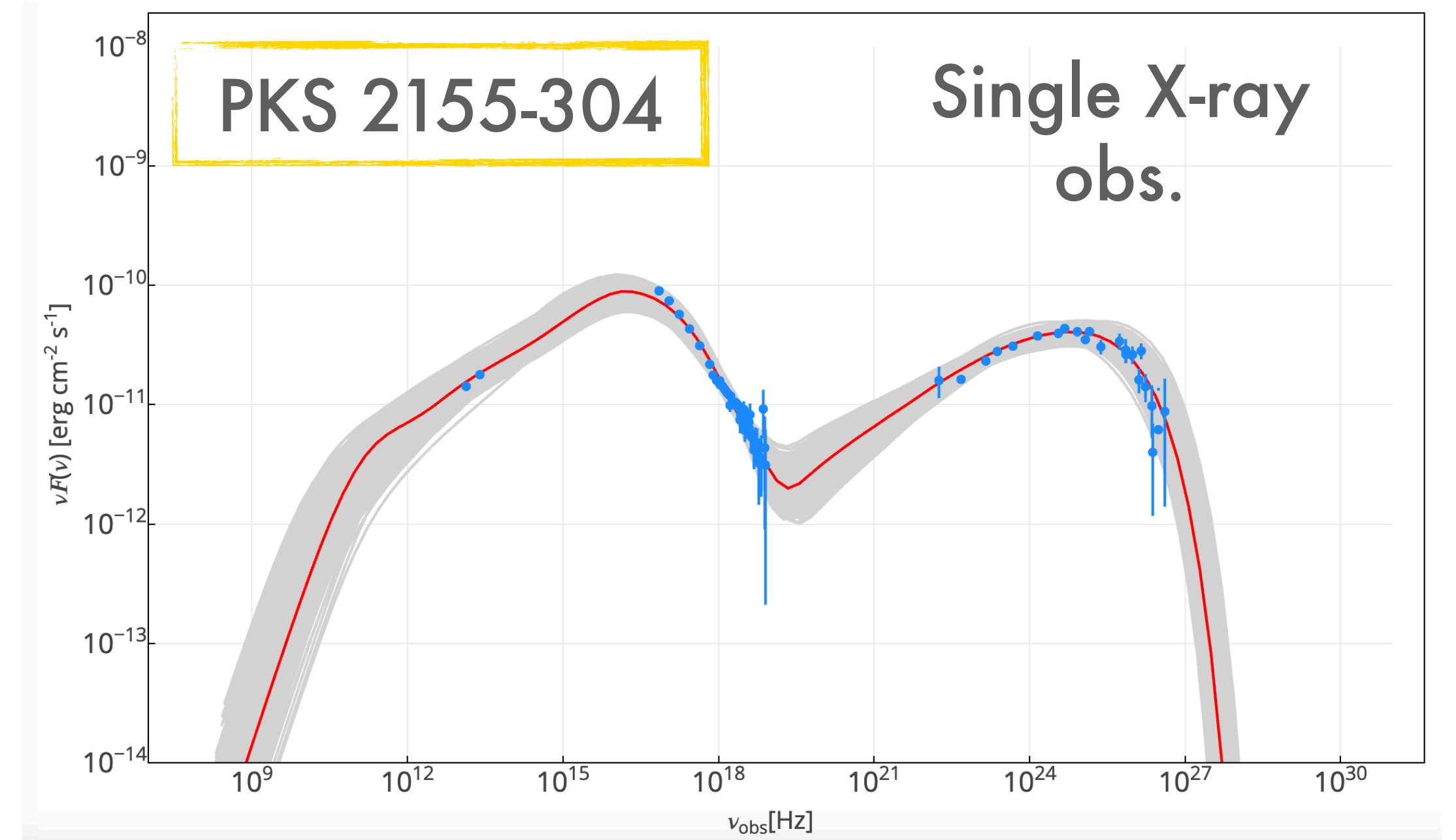
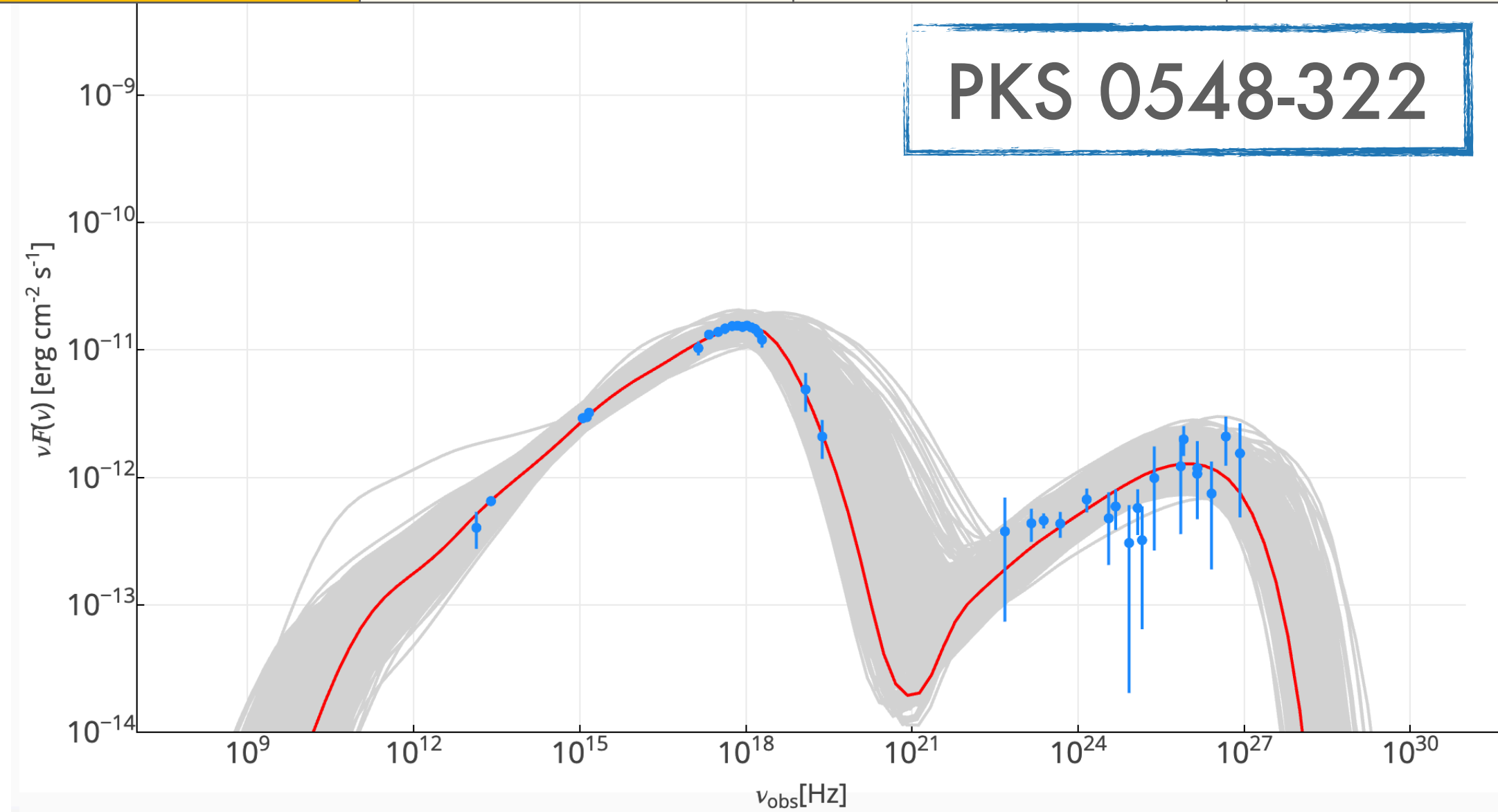
- Allows to scan the whole parameter space
- Available through MMDC platform



# Results

## Preliminary

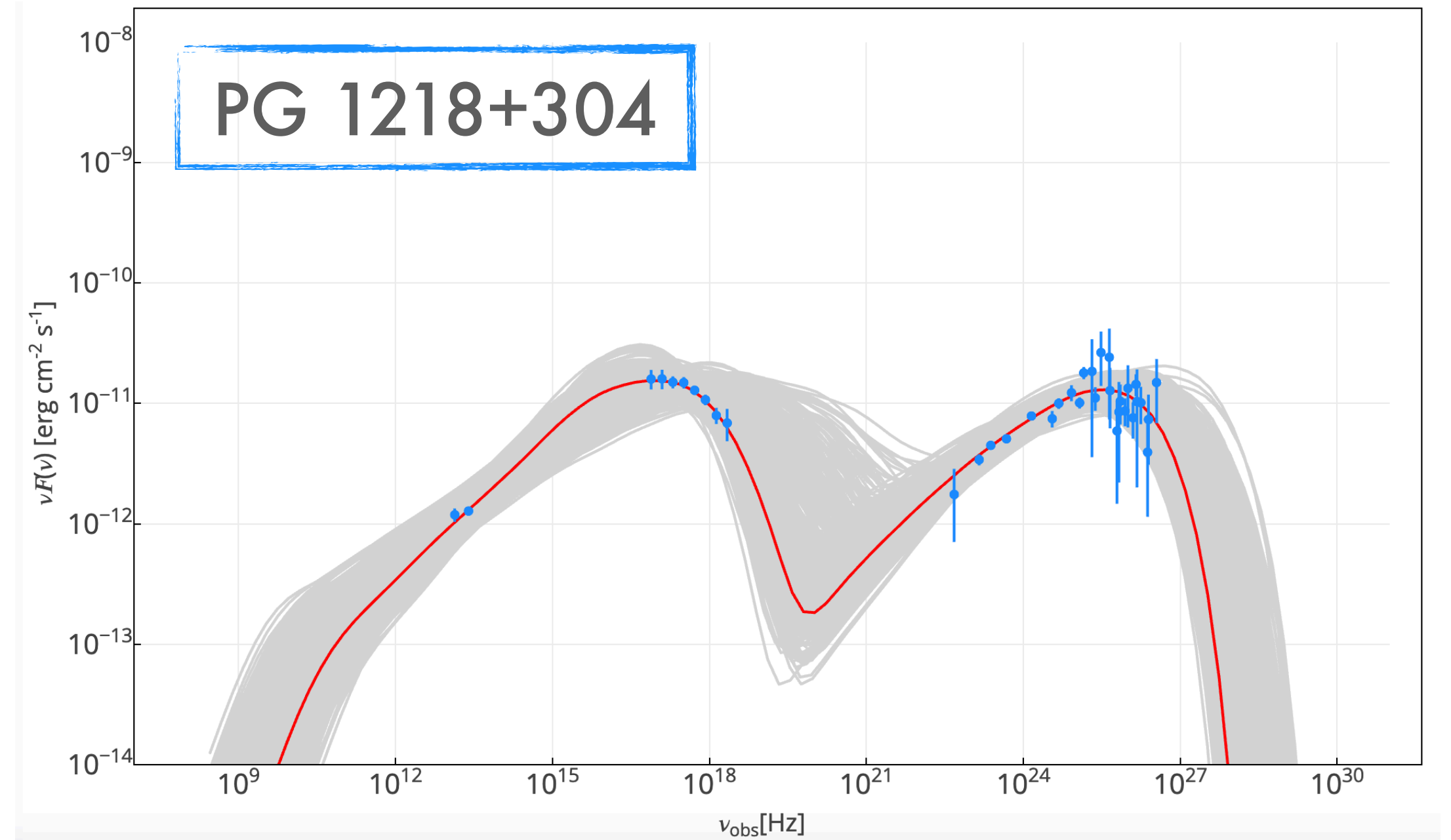
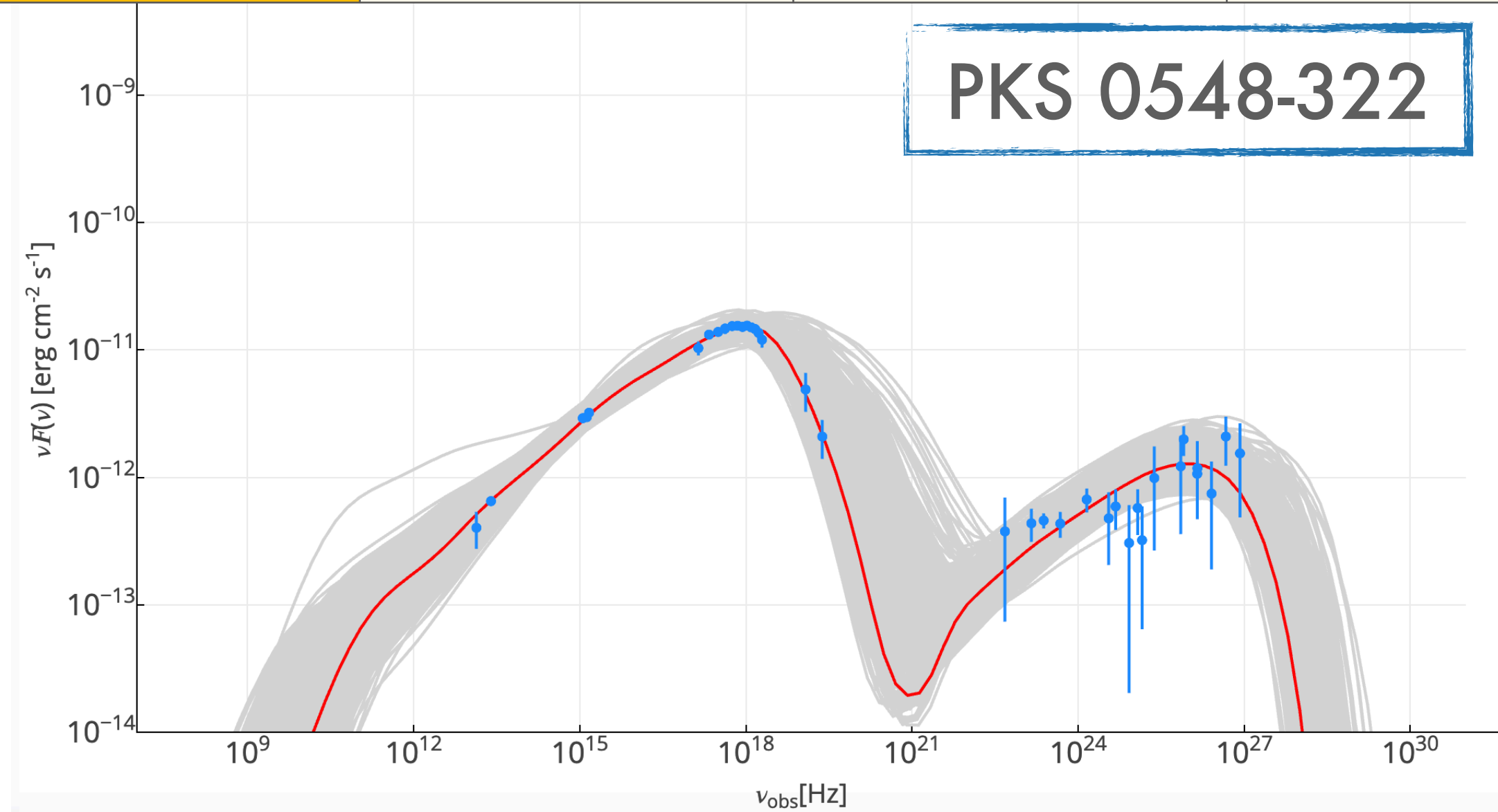
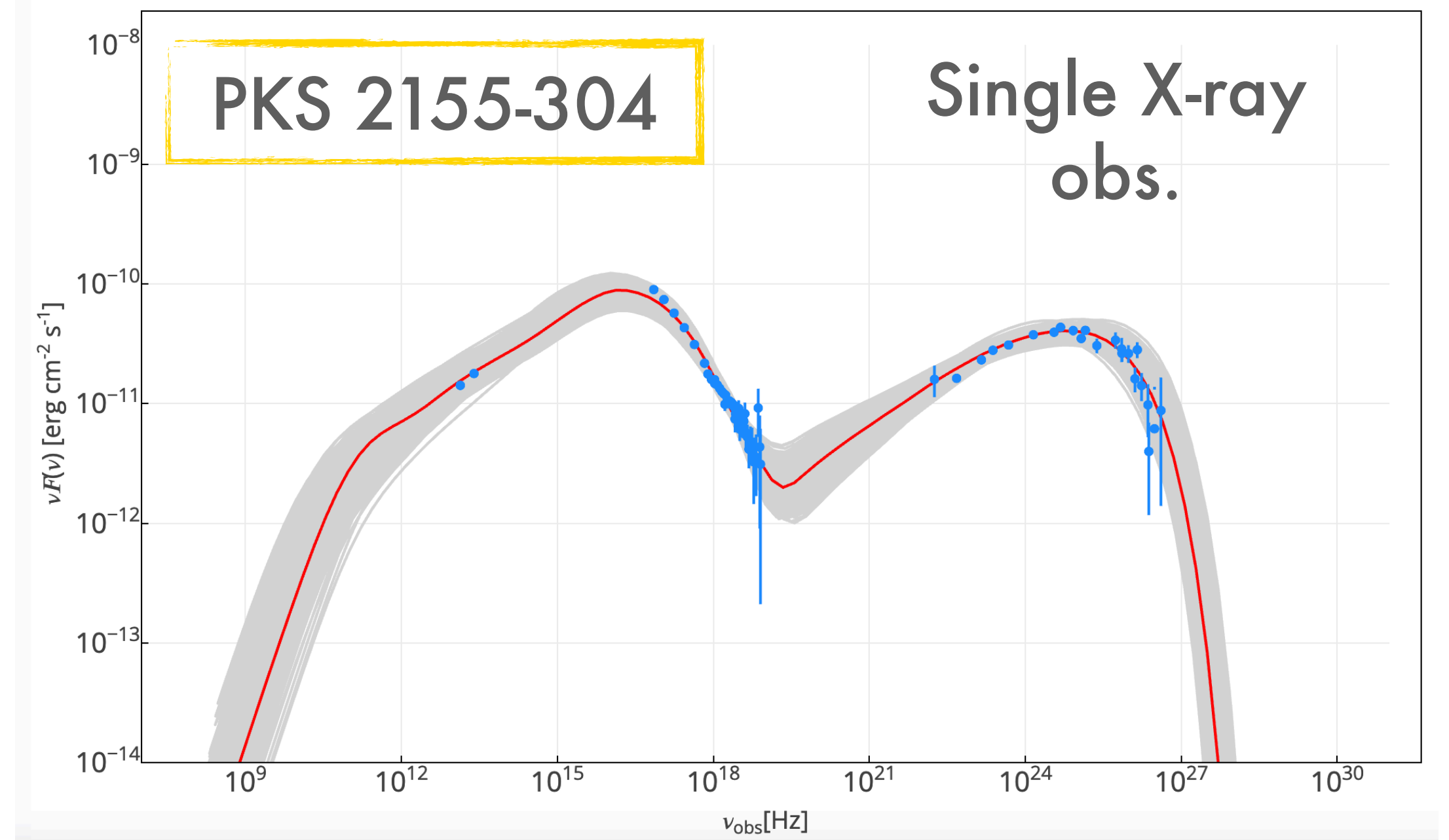
	PKS 2155-304	PG 1218+304	PKS 0548-322
t_var [s]	4.86E+04	8.57E+04	1.05E+04
p1	2.34	2.1	2.16
log(L_e [erg/s])	44.77	44.48	43.39
log( $\gamma_{\text{max}}$ )	5.28	5.66	5.79
$\delta$	43.06	29.71	49.25
log(B [G])	-1.51	-1.61	-1.6



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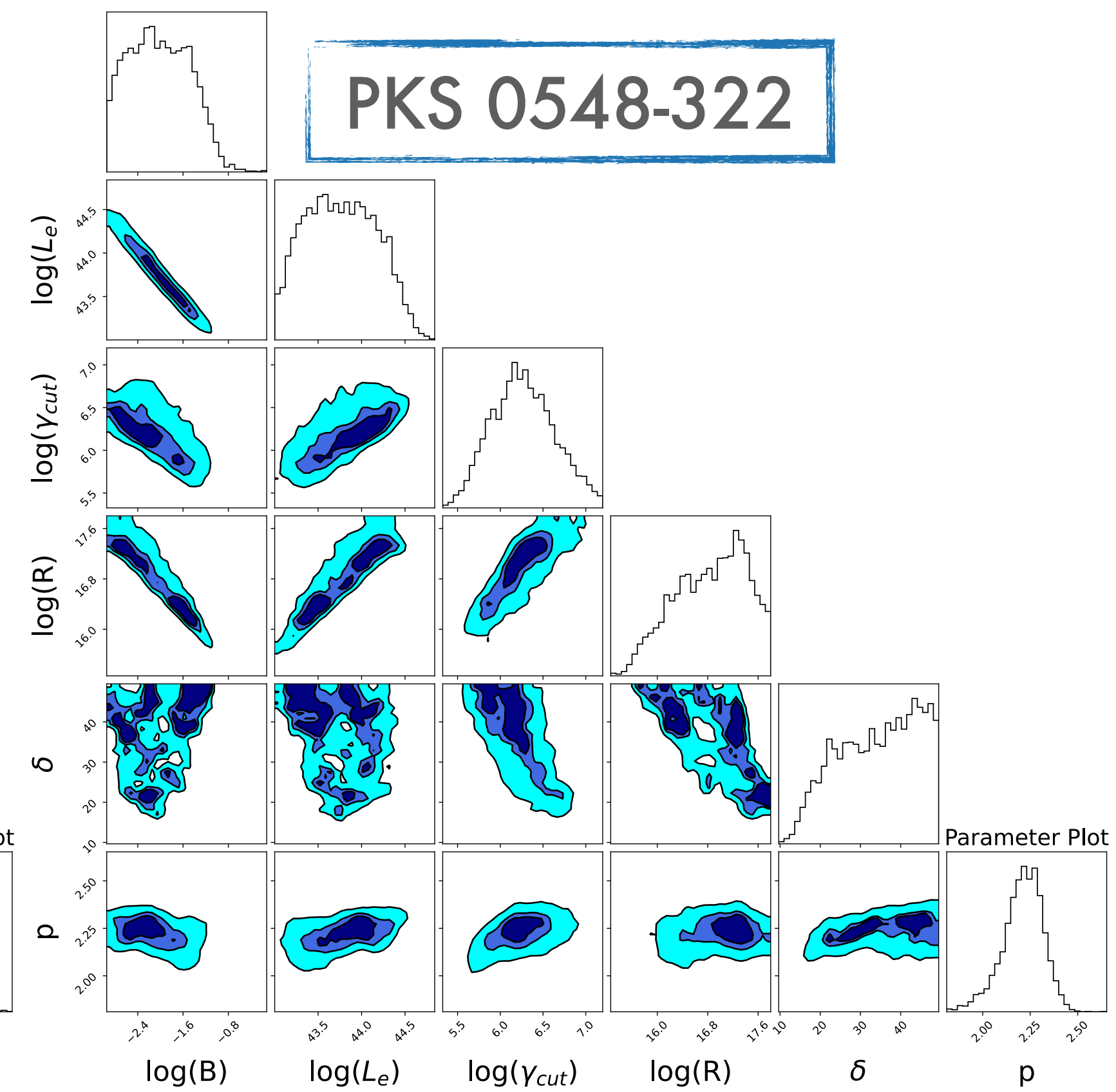
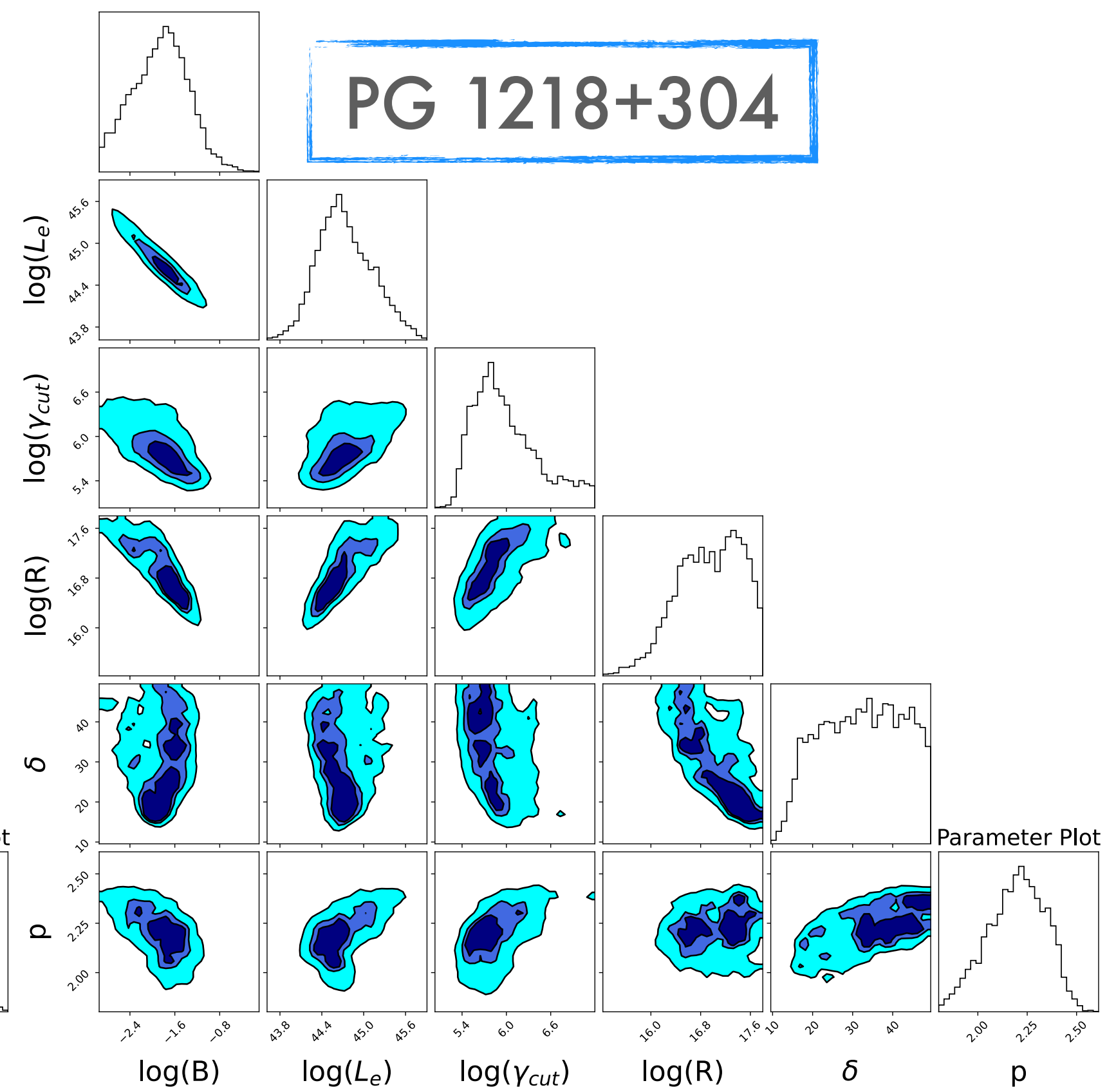
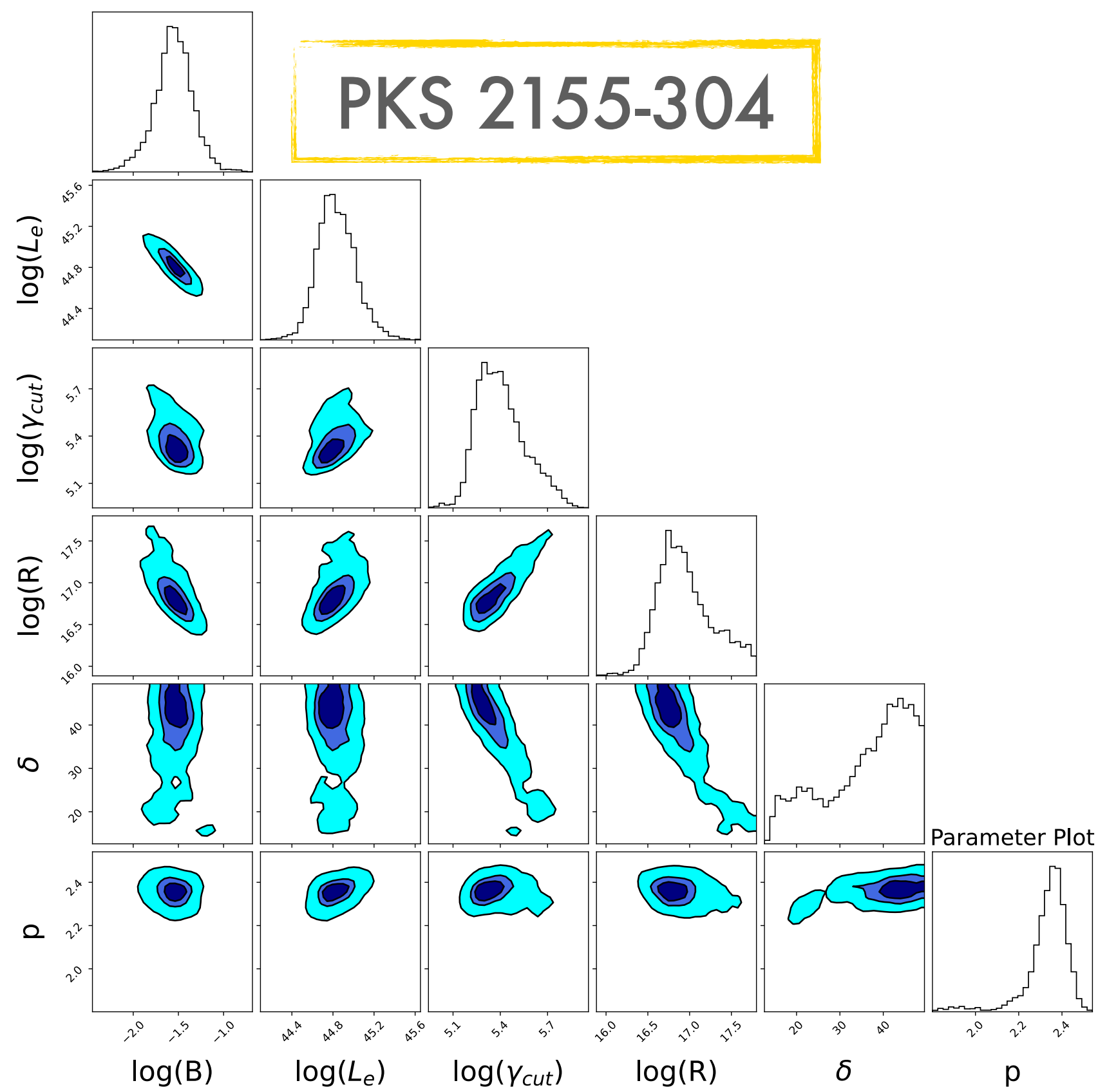
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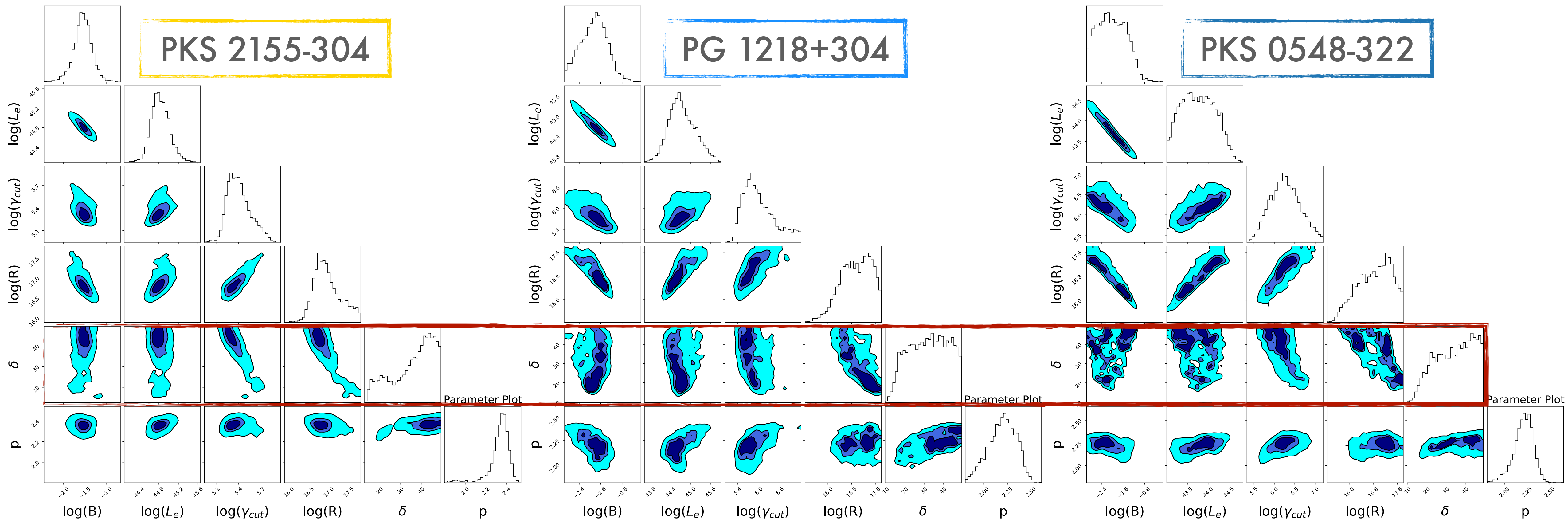
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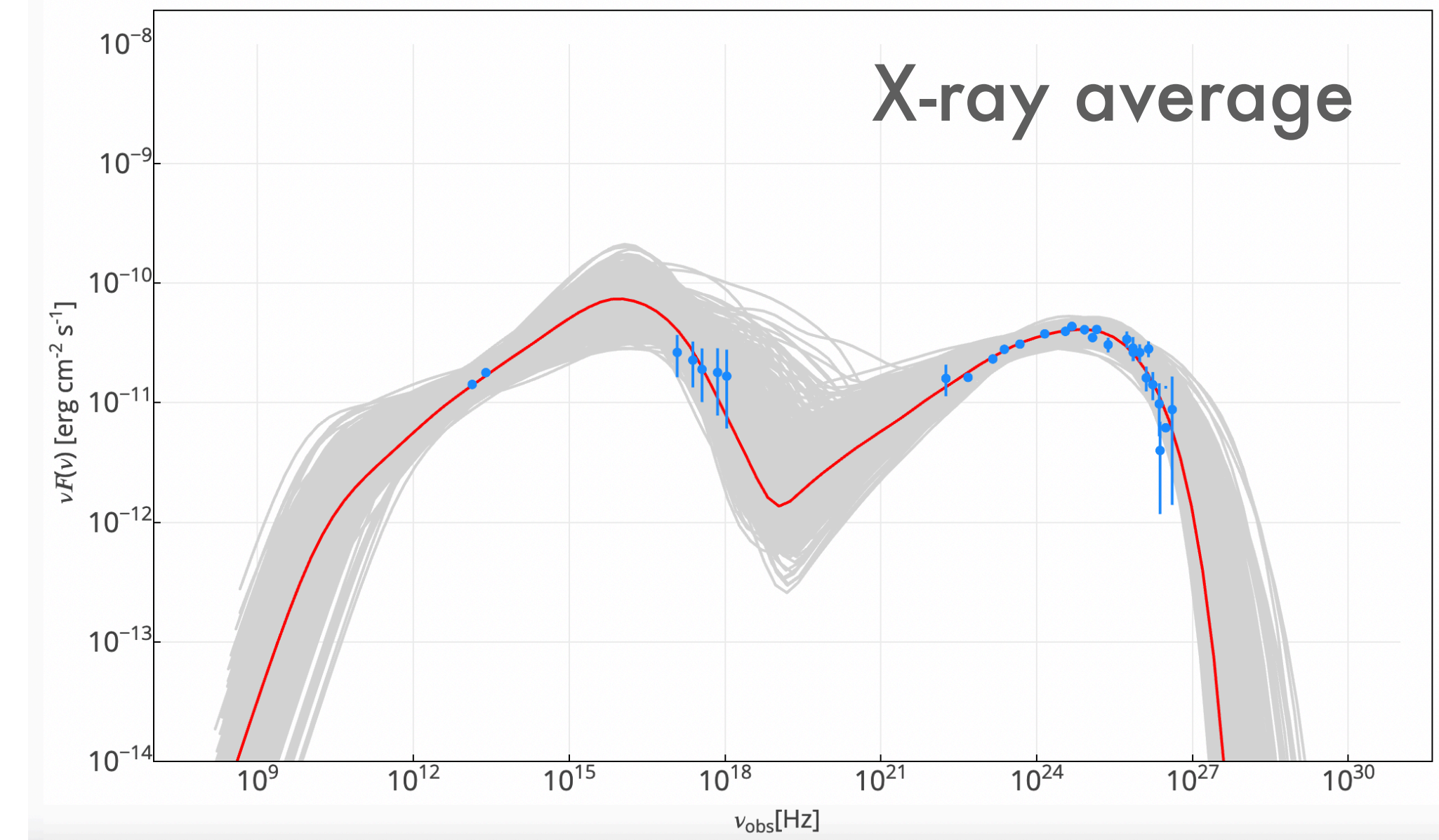
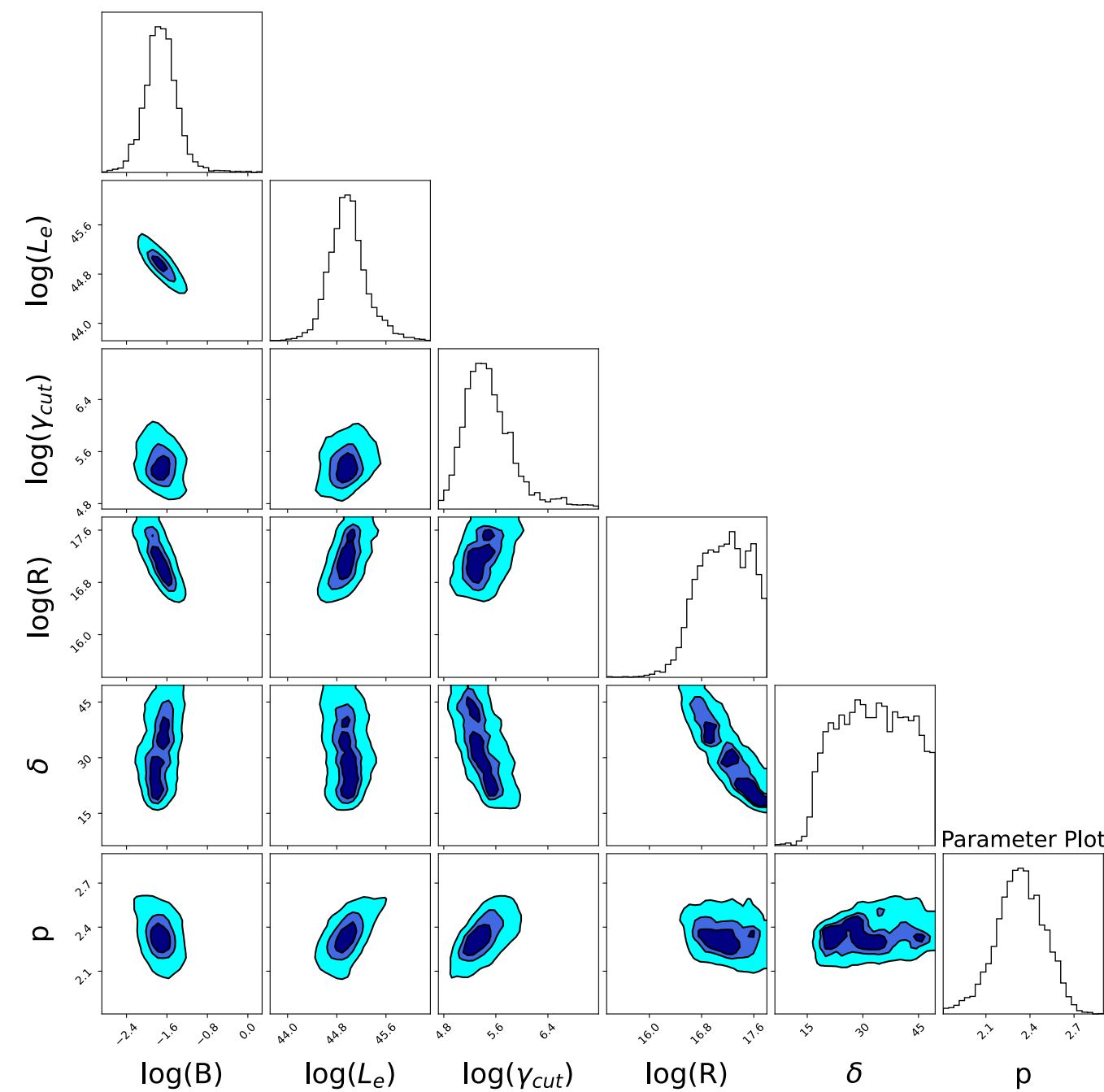
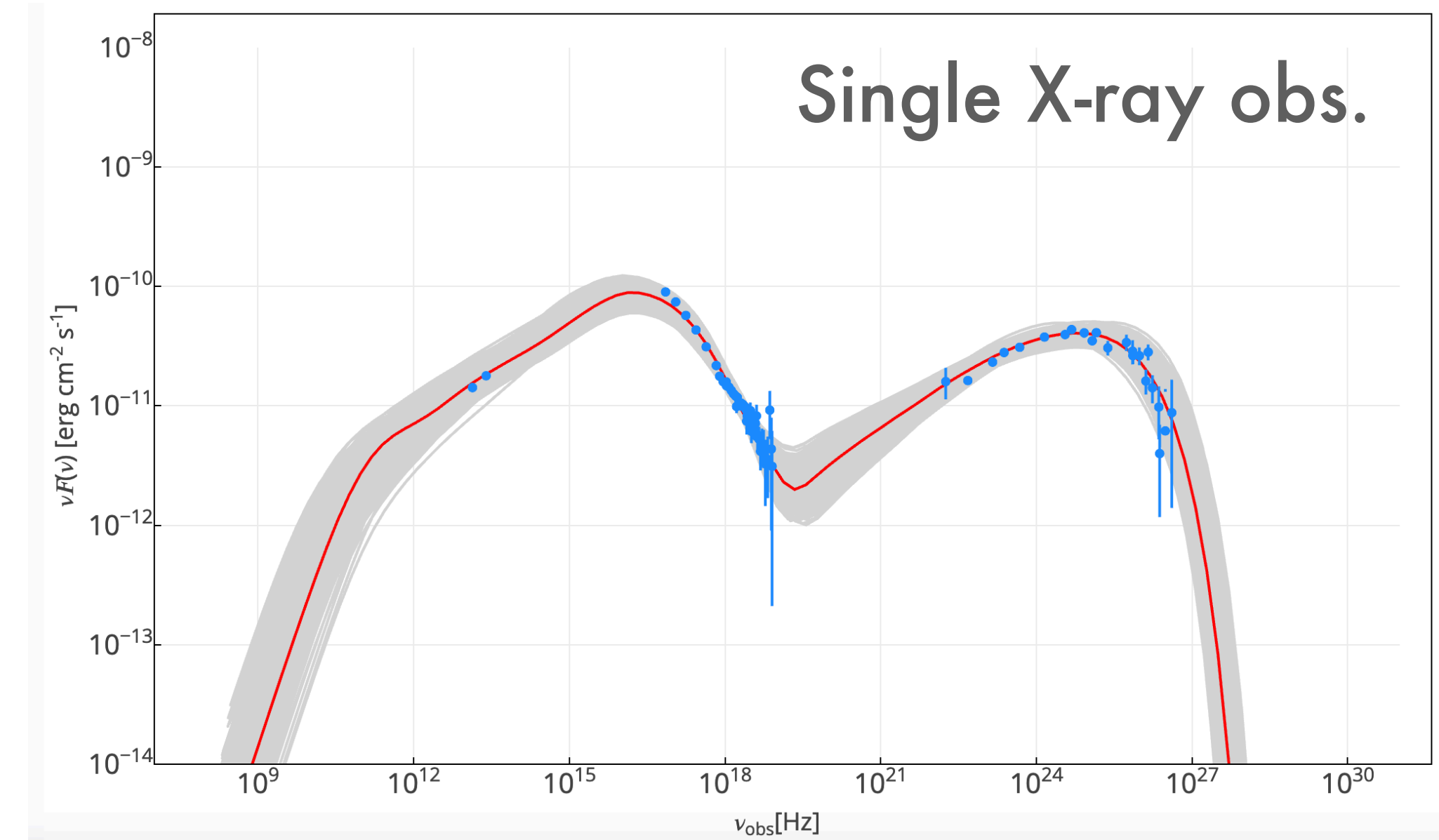
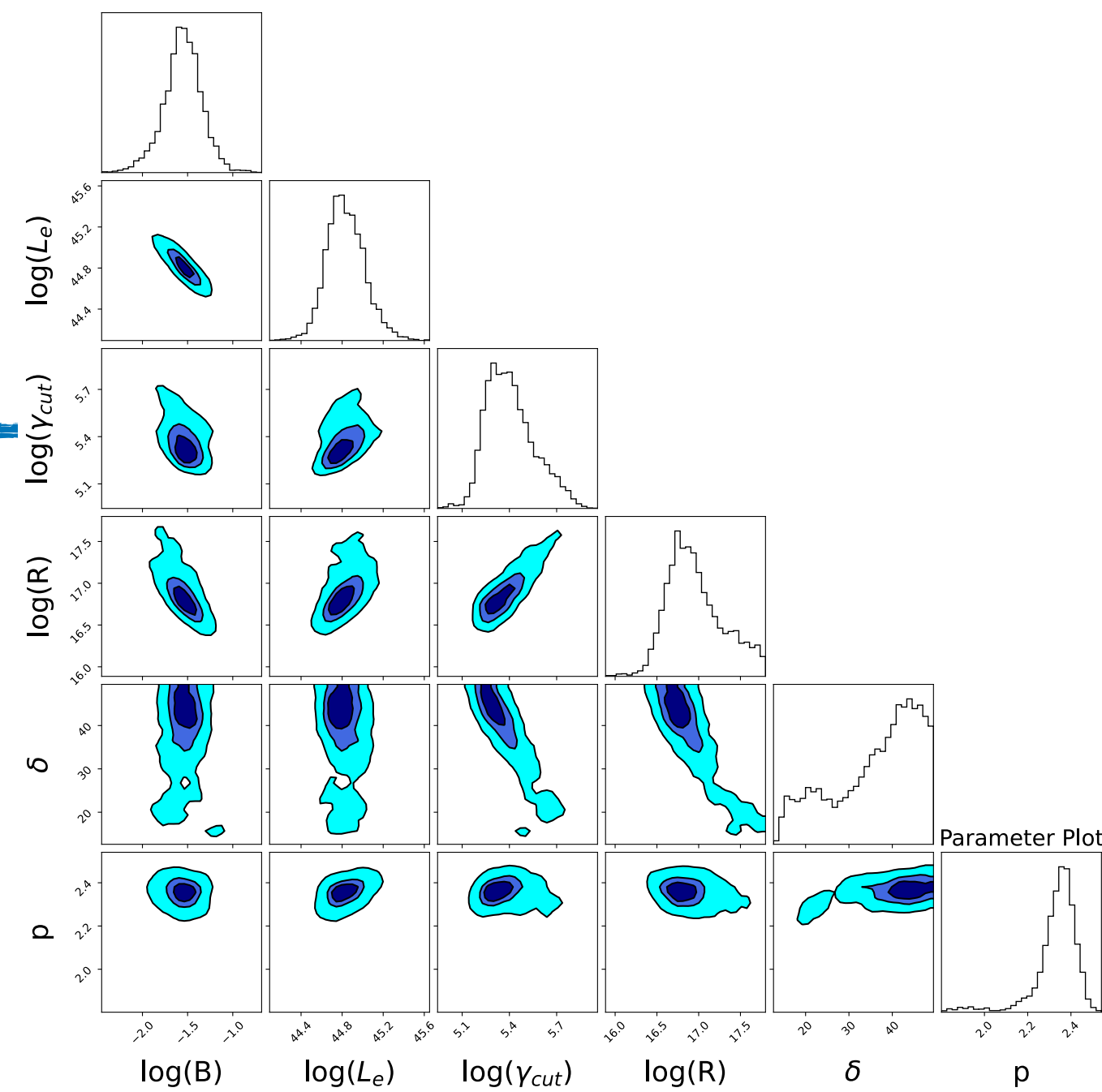
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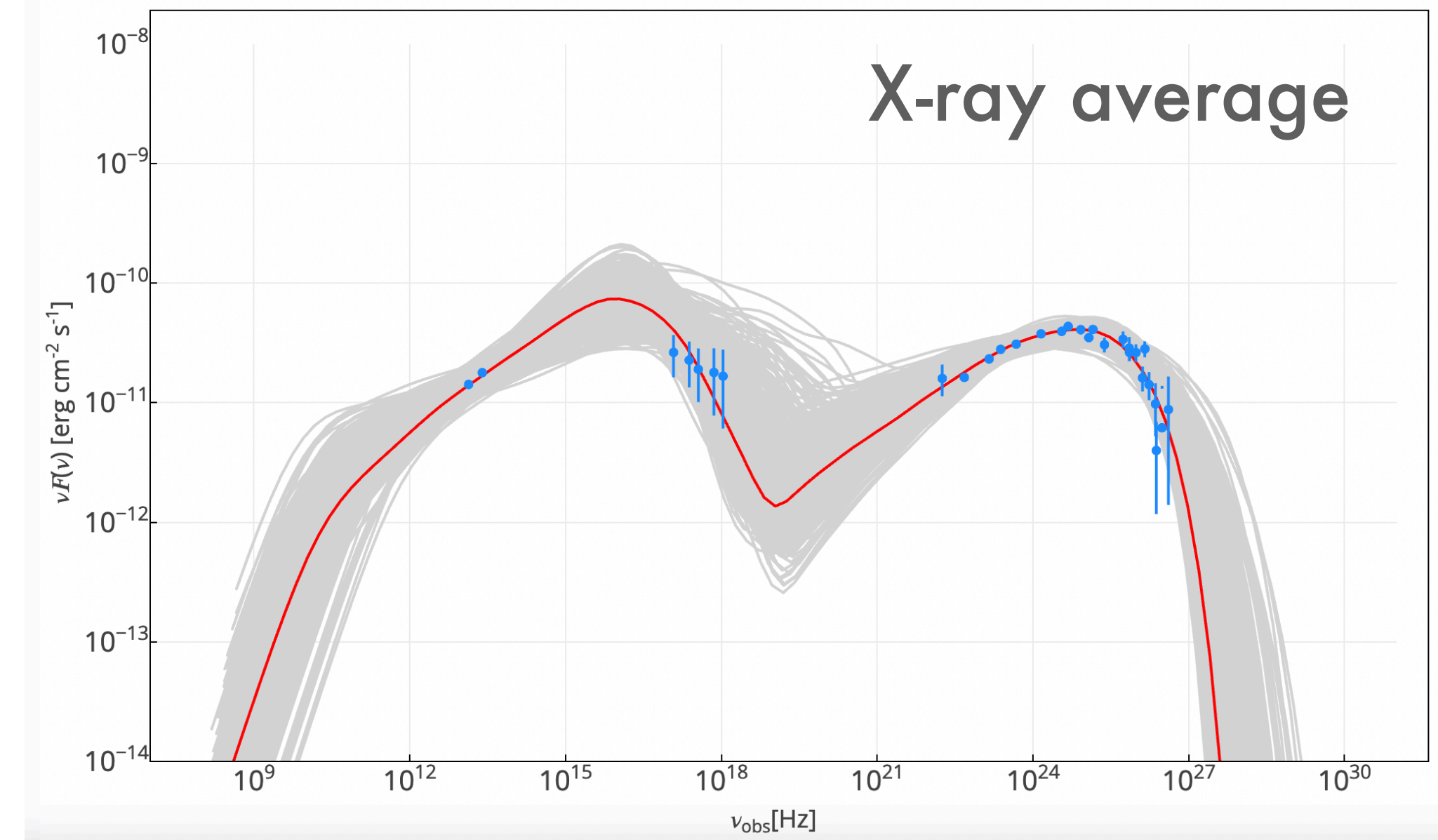
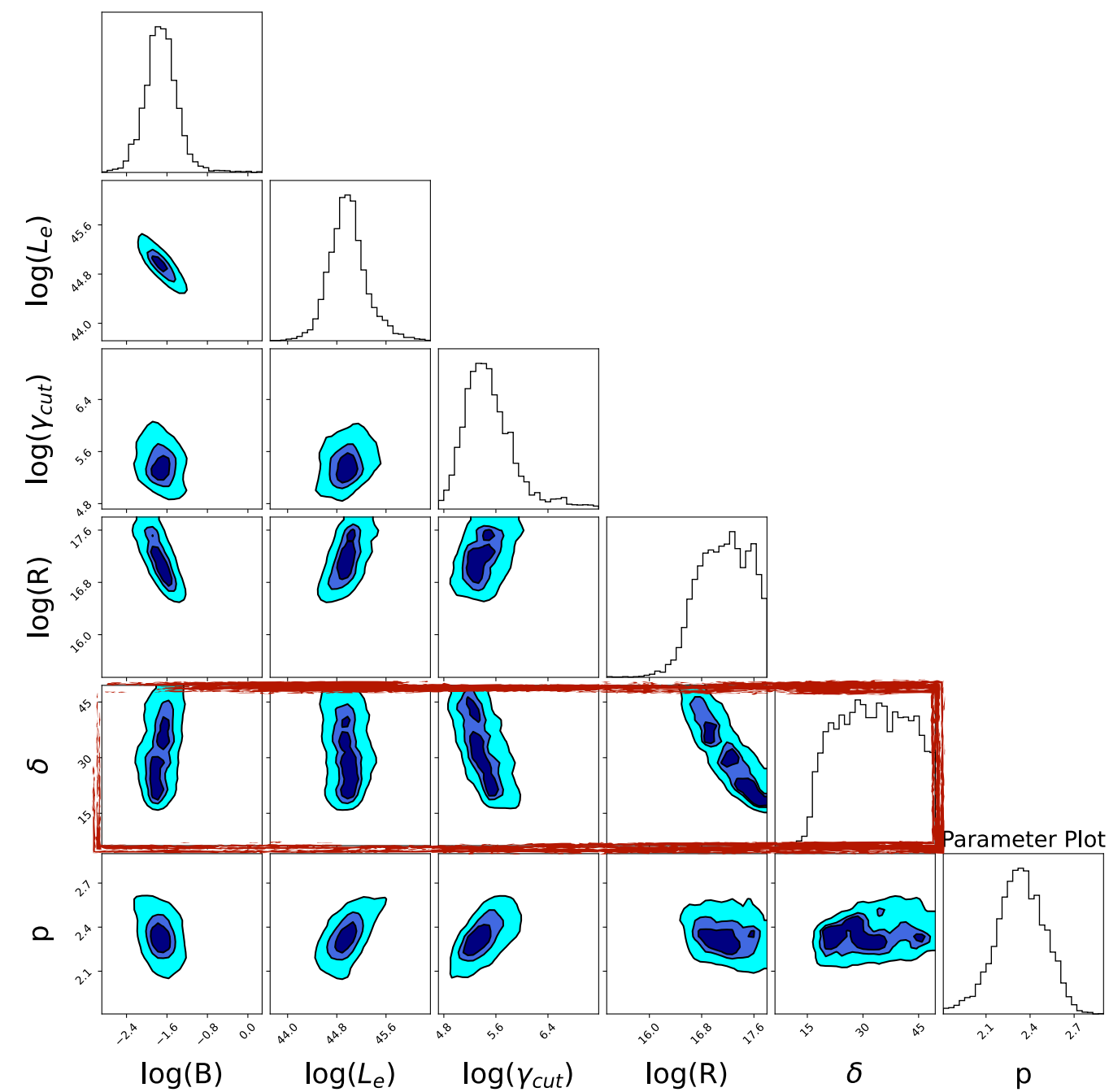
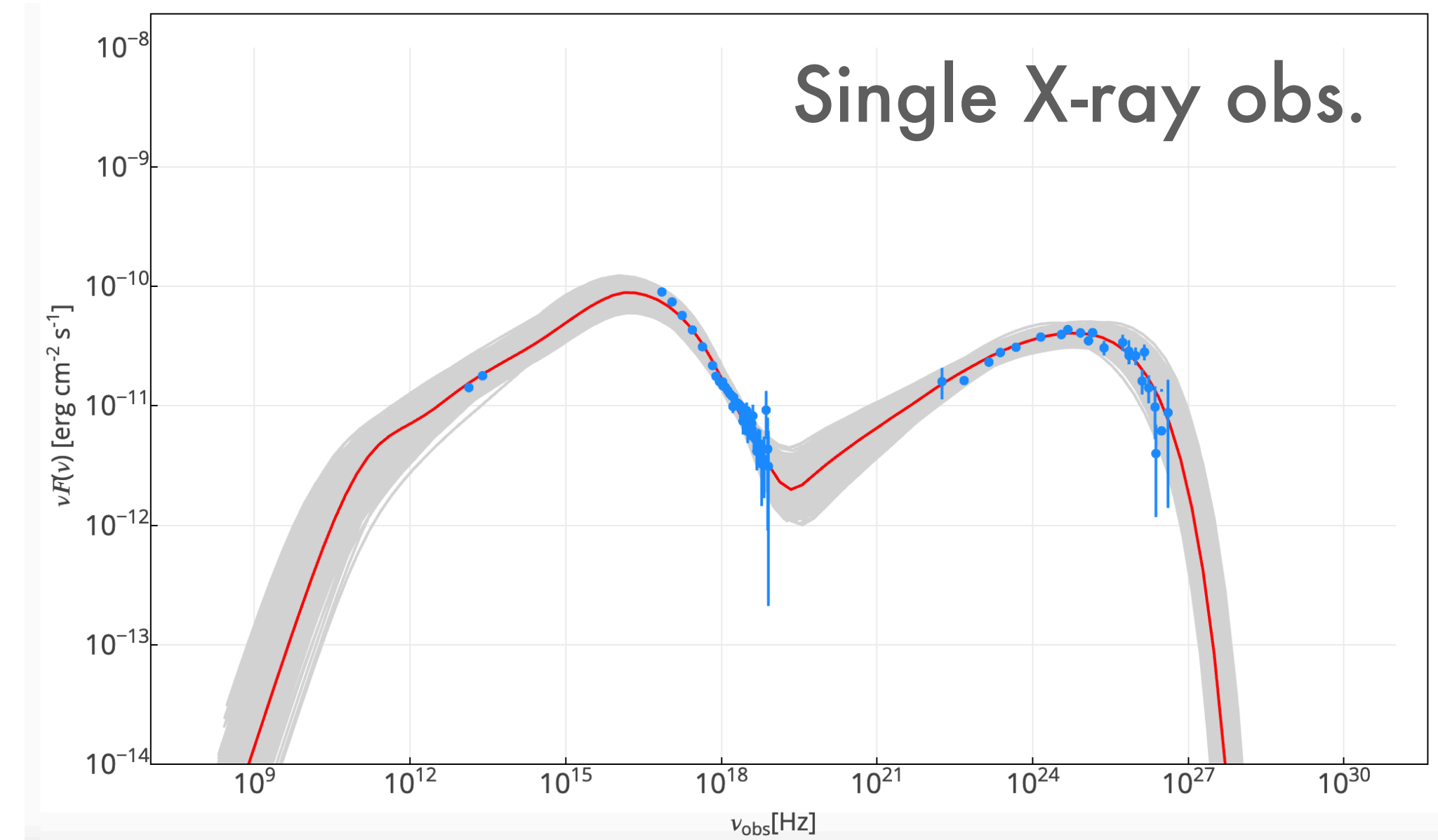
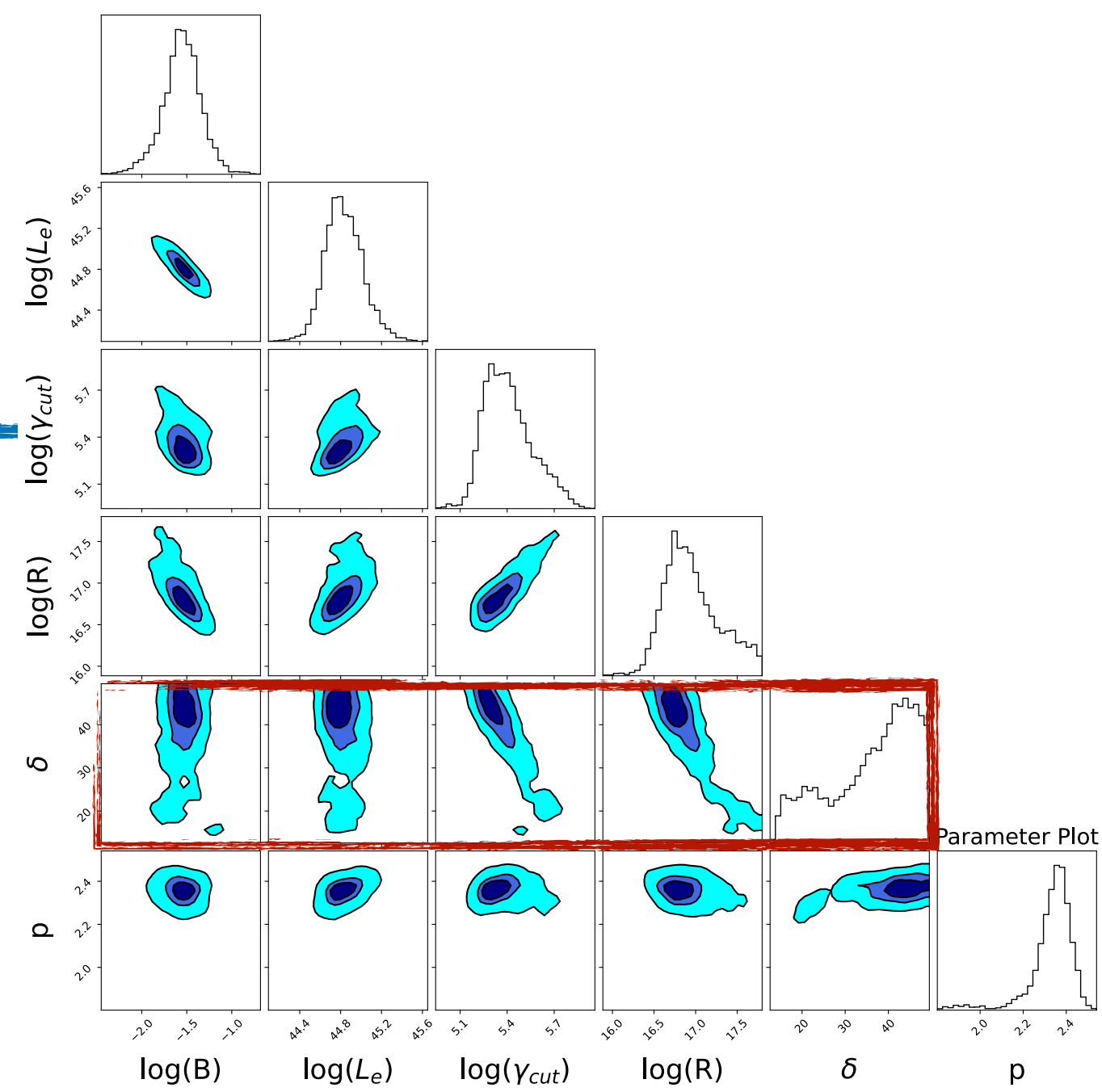
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- Non-trivial data selection, especially in X-rays:
  - Useful in the future...
    - Continuous monitoring of sources in X-rays
    - Obs. strategy of simultaneous X-ray and  $\gamma$ -ray data
    - Online tools filtering observations based on the obs. Period
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  - **Work in progress:** investigation of different  $\delta$  values for each source and interpretation of modeling results
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**Stay tuned for  
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**Thank you!**

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