



SST-1M

Single-Mirror
Small Size Telescope



Institute of Physics of the
Czech Academy of Sciences

Observation of Astrophysical Sources with **SST-1M** Telescopes

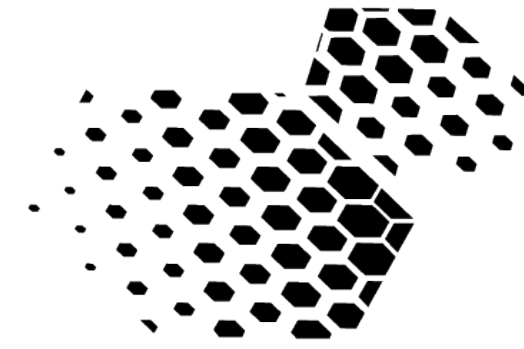
First results

Jakub Juryšek

Institute of Physics of the Czech Academy of Sciences

for the **SST-1M** Collaboration

The SST-1M project



SST-1M

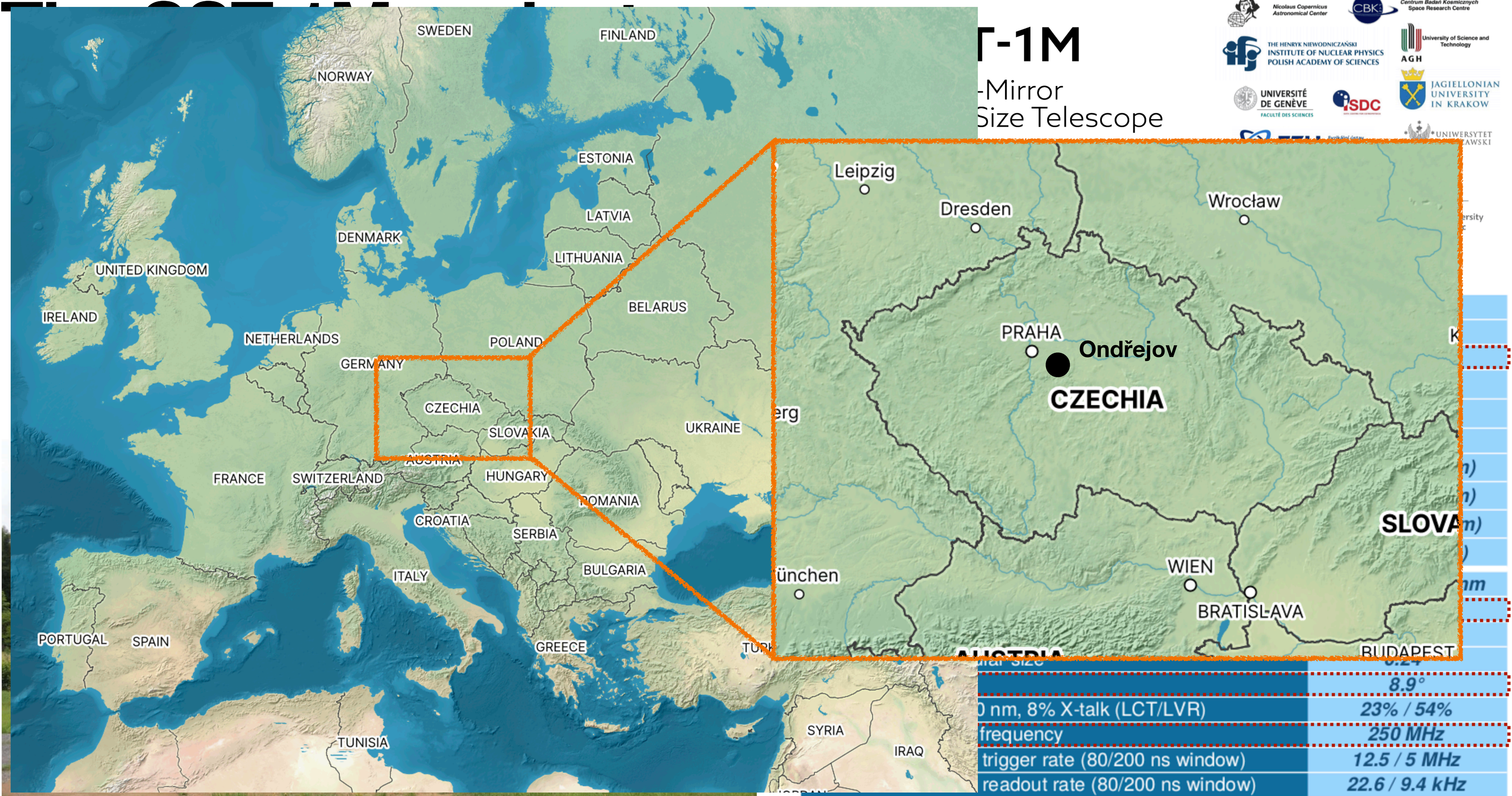
Single-Mirror
Small Size Telescope



- A collaborative effort of 17 institutes in 3 countries
- **Two 4-m IACTs, Ondrejov observatory (500 m a.s.l.), Czech Republic**
- Davies-Cotton optical design, **innovative SiPM camera with fully digital trigger and readout**
- Commissioning phase, atmospheric conditions not optimal in Ondrejov - **‘testing bench’ for the telescopes**
- Operated in **mono and stereo** mode



Optics	Focal Length	5600 ± 5 mm
	f/D	1,4
	Dish diameter	4 m
	Mirror Area (*)	9.42 m ²
	Mirror Effective Area (*)	6.47 m ²
	Hexagonal Mirror facets	780 ± 3 mm
	Mirror PSF D ₈₀ (requirement)	0.082° (8.1 mm)
	Mirror PSF D ₈₀ (measured)	0.028° (2.7 mm)
	Telescope PSF D ₈₀ (required)	0.25° (24.4 mm)
Telescope PSF D ₈₀ (measured) On-Axis	0.082° (8 mm)	
Camera	Camera dimension (R/thickness)	810 mm / 900 mm
	Total pixel number	1296
	Pixel linear size	23.4 mm
	Pixel angular size	0.24°
	FoV	8.9°
	PDE@470 nm, 8% X-talk (LCT/LVR)	23% / 54%
	Sampling frequency	250 MHz
	Maximum trigger rate (80/200 ns window)	12.5 / 5 MHz
	Maximum readout rate (80/200 ns window)	22.6 / 9.4 kHz
Time Spread RMS	< 0.25 ns	



Γ-1M

-Mirror Size Telescope



0 nm, 8% X-talk (LCT/LVR)	8.9°
frequency	23% / 54%
trigger rate (80/200 ns window)	250 MHz
readout rate (80/200 ns window)	12.5 / 5 MHz
Time Spread RMS	22.6 / 9.4 kHz
	< 0.25 ns

SST-1M 'mini array'

- Two telescopes 155 m apart
- **Stereo trigger** managed with **SWAT**, **White Rabbit** synchronisation of the timestamps



SST-1M major milestones

Analysis/science related

- February/November 2022 - **Telescope 2/1 installation**
- March 2023 - **first sst1mpipe prototype**
- April 2023 - **Crab detection in mono**
- June 2023 - **Crab detection in stereo** (no White Rabbit yet)
- July 2023 - **1ES 1959+650 detection**, SST-1M is going extragalactic!
- October 2023 - **First Crab detection in true stereo** with White Rabbit
- November 2023 - **First release of sst1mpipe**
- March 2024 - **Detected Mrk 421** brightening, **first ATel #16533**
- August 2024 - **First extended source detected** in stereo

Mono and stereo performance of the two SST-1M telescope prototypes

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The Astronomer's Telegram <http://www.astronomerstelegram.org>

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ATEL #16533 ATEL #16533

Title: Detection of enhanced very-high-energy gamma-ray emission from Markarian 421

Author: Thomas Tavernier, in the behalf of SST-1M Consortium

Queries: tavernier@fzu.cz

Posted: 15 Mar 2024; 16:55 UT

Subjects:Gamma Ray, TeV, VHE, AGN, Blazar

Analysis of commissioning data from SST-1M : A Prototype of Single-Mirror Small Size Telescope

T. Tavernier,^{a,*} J. Juryšek,^a V. Novotný,^{a,b} M. Heller,^c D. Mandat,^a M. Pech,^a C. Alispach,^c A. Araudo,^{d,e} V. Beshley,^f J. Blazek,^a J. Borkowski,^g S. Boula,^h T. Bulik,ⁱ F. Cadoux,^c S. Casanova,^h A. Christov,^a L. Chytka,^j D. della Volpe,^c Y. Favre,^c L. Gibaud,^k T. Gieras,^h P. Hamal,^j M. Hrabovsky,^j M. Jelínek,^l V. Karas,^d E. Lyard,^m E. Mach,^h W. Marek,^h S. Michal,^j J. Michałowski,^h R. Moderski,^g T. Montaruli,^c A. Muraczewski,^g S. R. Muthyala,^a A. Nagai,^c K. Nalewajski,^h D. Neise,ⁿ J. Niemiec,^h M. Nikořajuk,^k M. Ostrowski,^o M. Palatka,^a M. Prouza,^a P. Rajda,^p P. Schovanek,^a K. Seweryn,^q V. Sliusar,^m Ł. Stawarz,^o J. Świerblewski,^h P. Świerk,^h J. Štrobl,^l J. Vicha,^a R. Walter,^m A. Zagdański^o and K. Ziętarą^o

PoS(ICRC2023)592
 PoS(ICRC2023)741
 Moriond 2024

SST-1M : Commissioning and Preliminary Observation Results

Thomas Tavernier, for the SST-1M Collaboration
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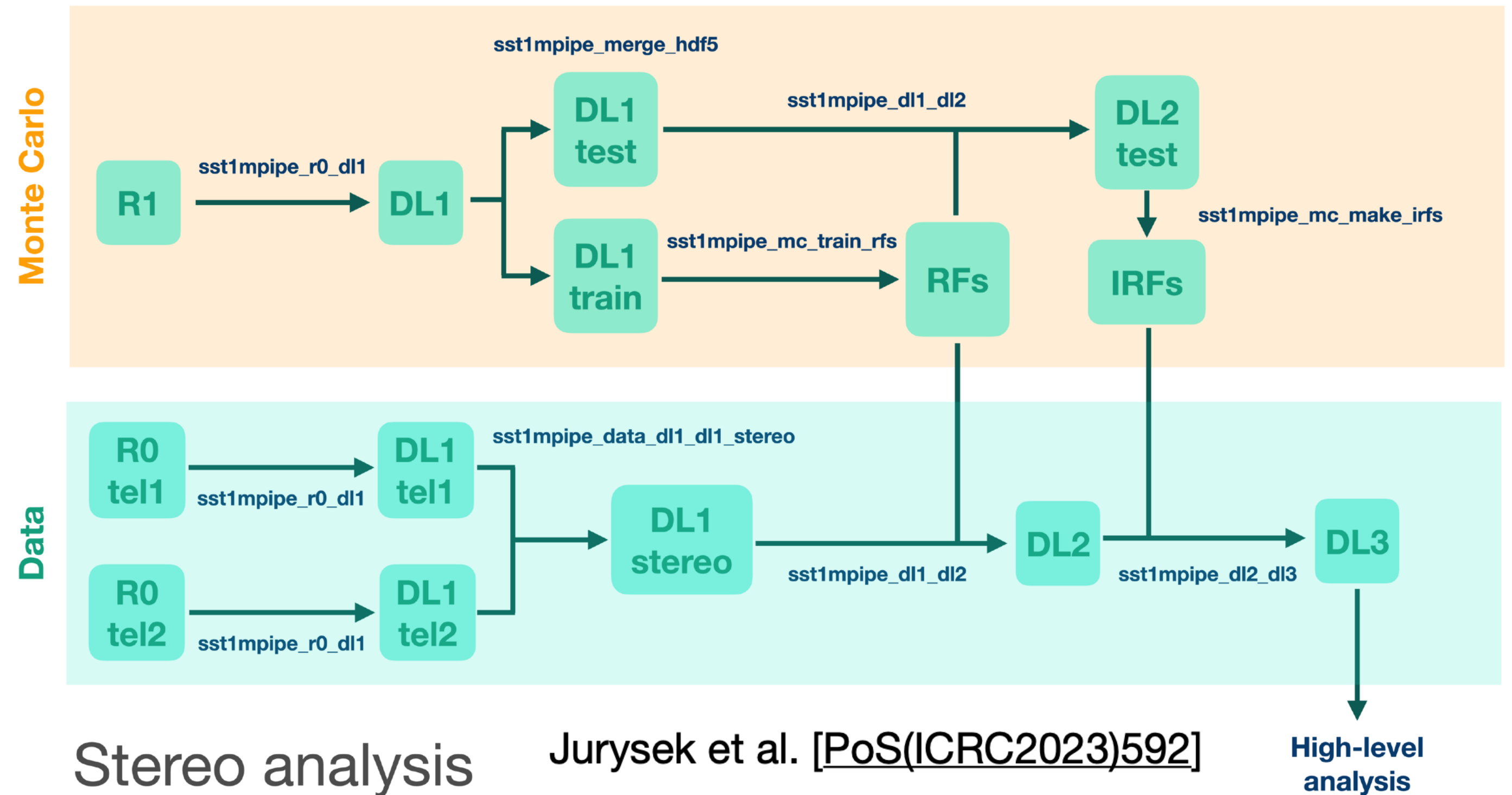
sst1mpipe: v0.4.1. 21 March 2024



Jurysek, Jakub¹ ; Tavernier, Thomas¹; Novotny, Vladimir²; Hamal, Petr³; Heller, Matthieu⁴; Blazek, Jiri¹; Muraczewski, Adam⁵; Muthyala, Srija Reddy¹; Alispach, Cyril⁴; Renier, Yves⁴; Coco, Victor⁴

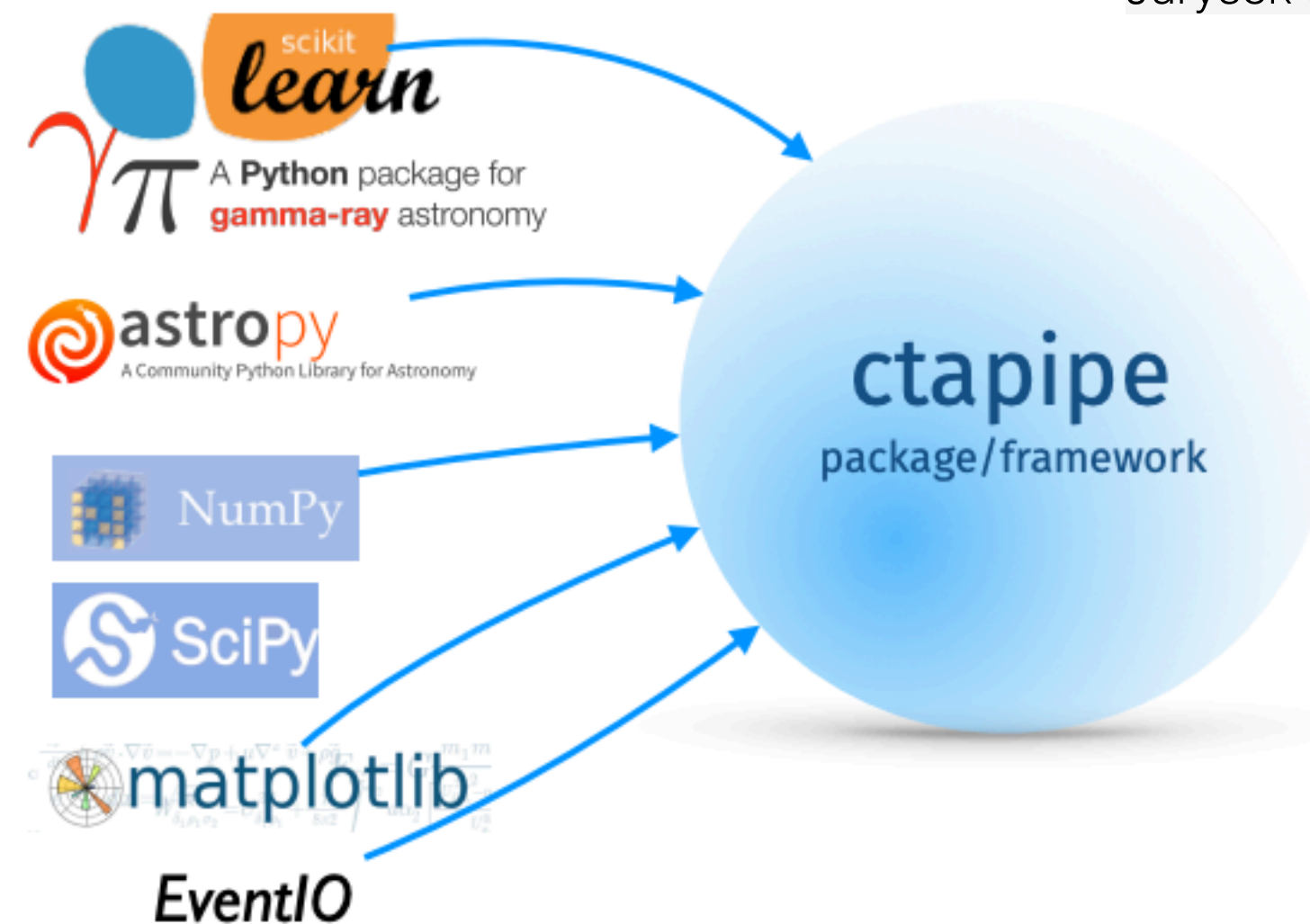
sst1mpipe

- Data and MC analysis pipeline
 - Raw waveform calibration and integration
 - Removing noise pixels and parametrisation of the shower images
 - Random Forest reconstruction
- Heavily based on **ctapipe** libraries, logic behind the analysis chain follows the **Istchain**
- Follows the data models of CTAO, compatibility with GADF for high level data analysis



Jurysek et al. [PoS(ICRC2023)592]

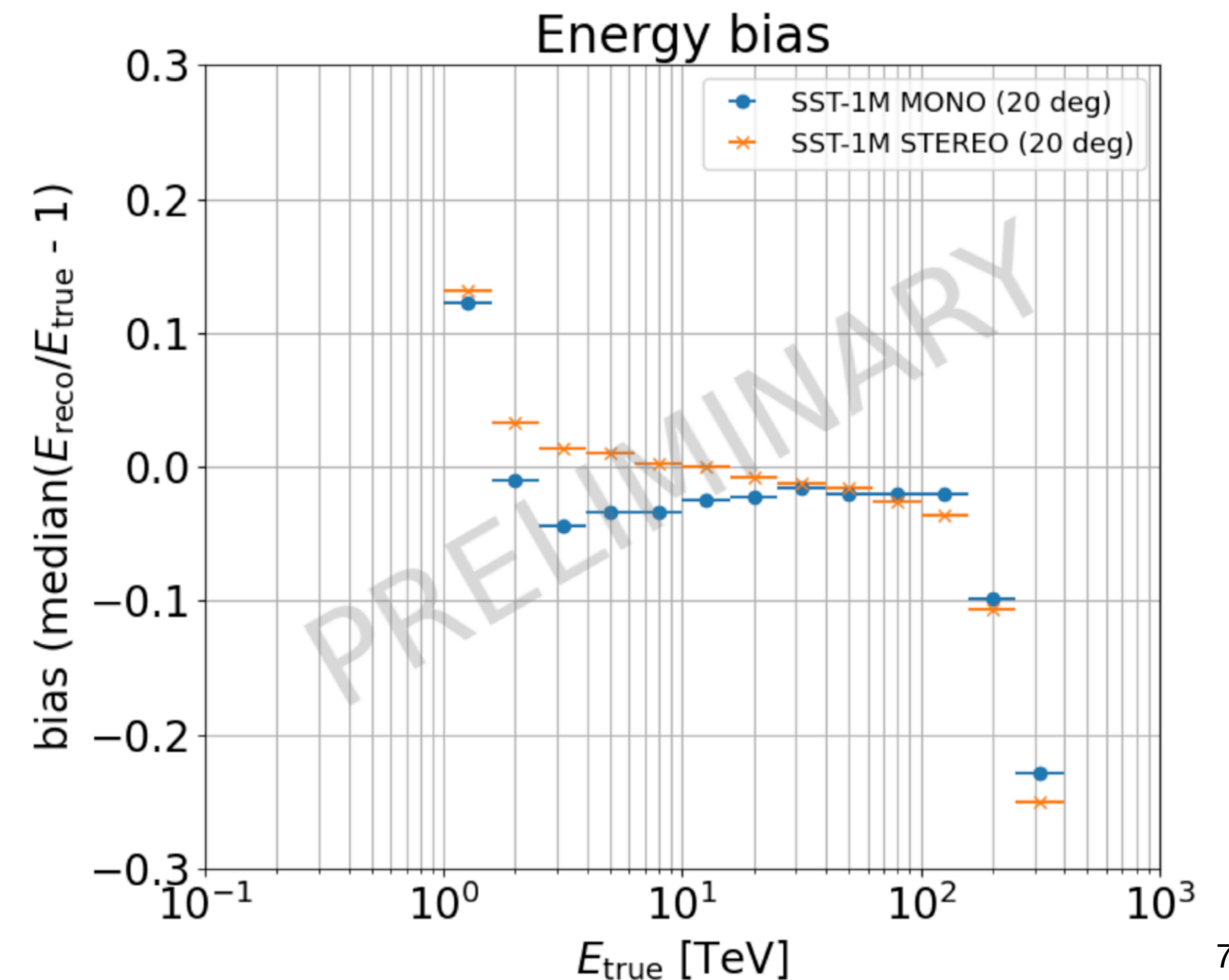
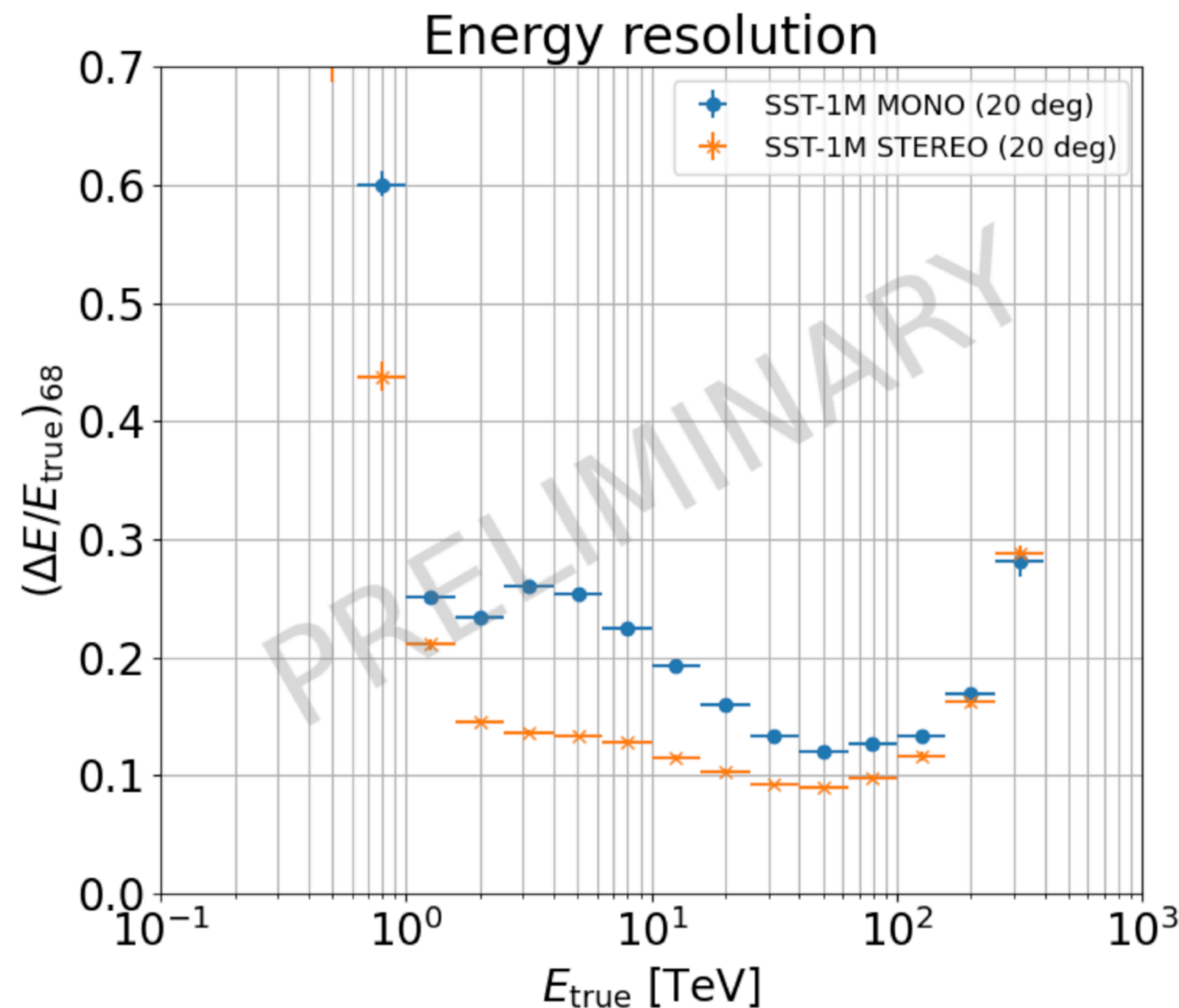
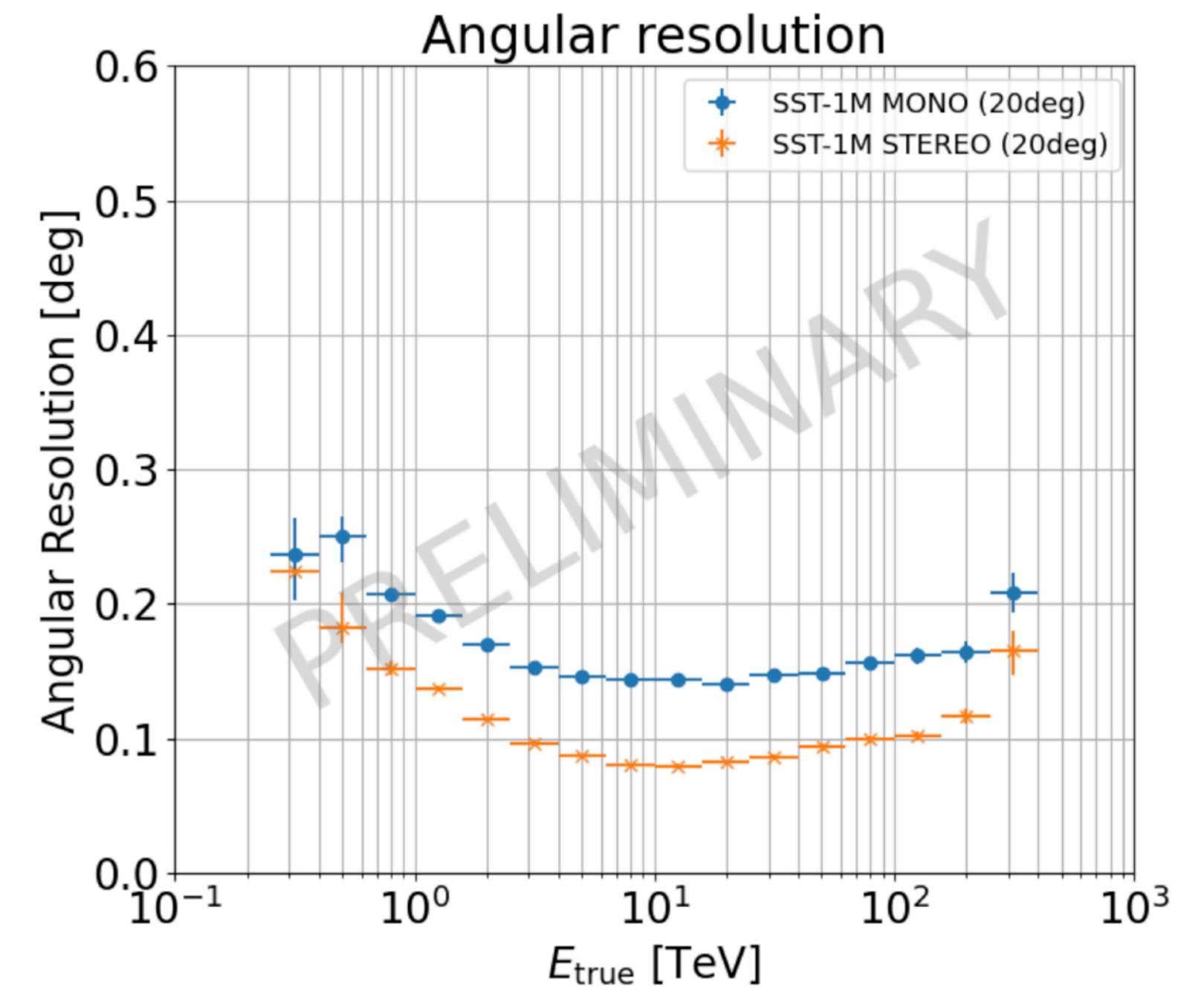
Jurysek et al. (2024). sst1mpipe: v0.4.1. 21 March 2024. Zenodo



SST1-M physics performance

Energy and angular resolution

- Difficult conditions in Ondrejov - low altitude and high NSB both increase the energy threshold (~ 1 TeV@20 deg zenith angle)
- Angular and energy resolution sufficient for physics cases despite that



Monte Carlo

- **Angular resolution:**
 - 0.15 deg (mono)
 - 0.10 deg (stereo)
- **Energy resolution:**
 - 15-25% (mono)
 - 10-15% (stereo)

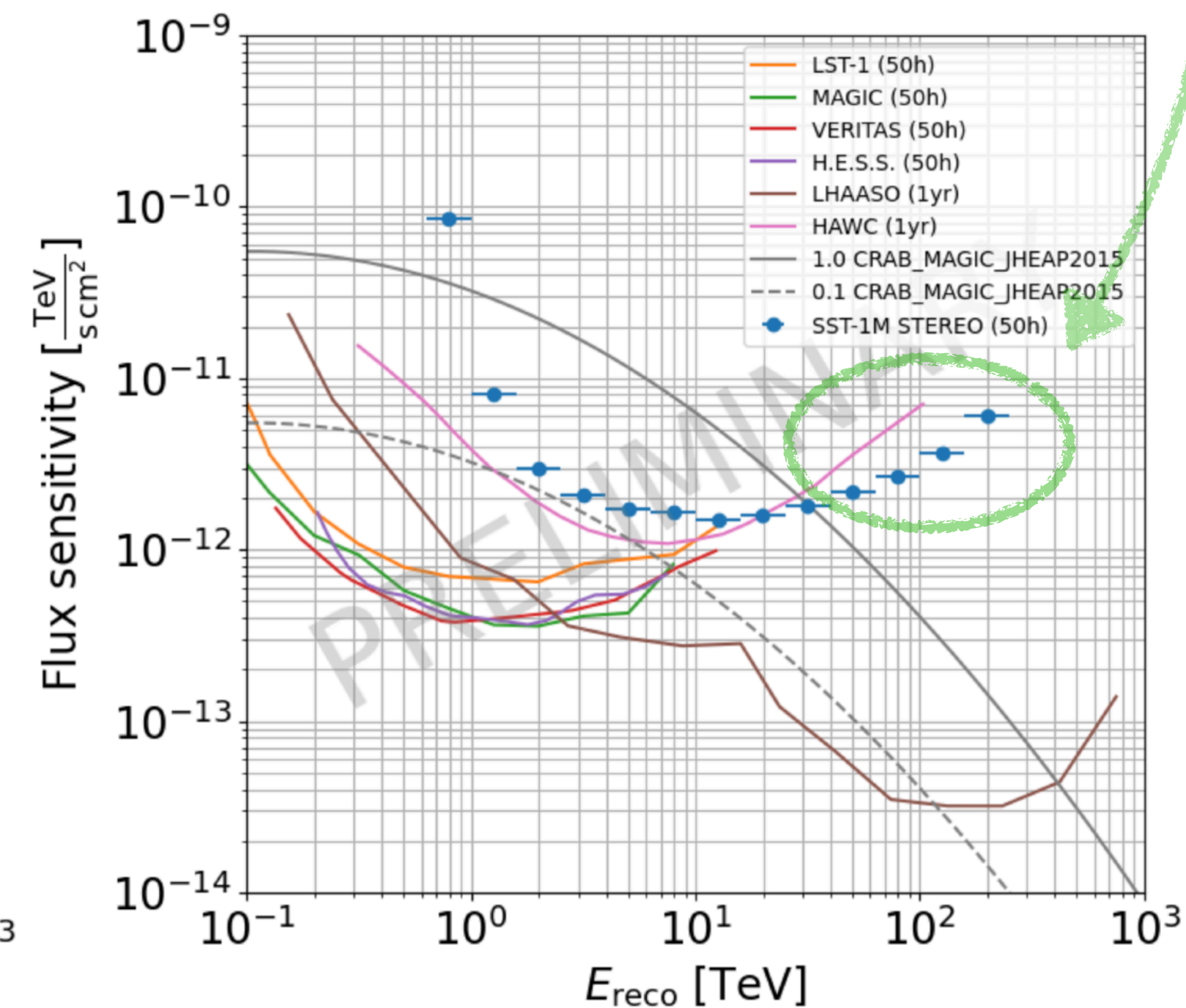
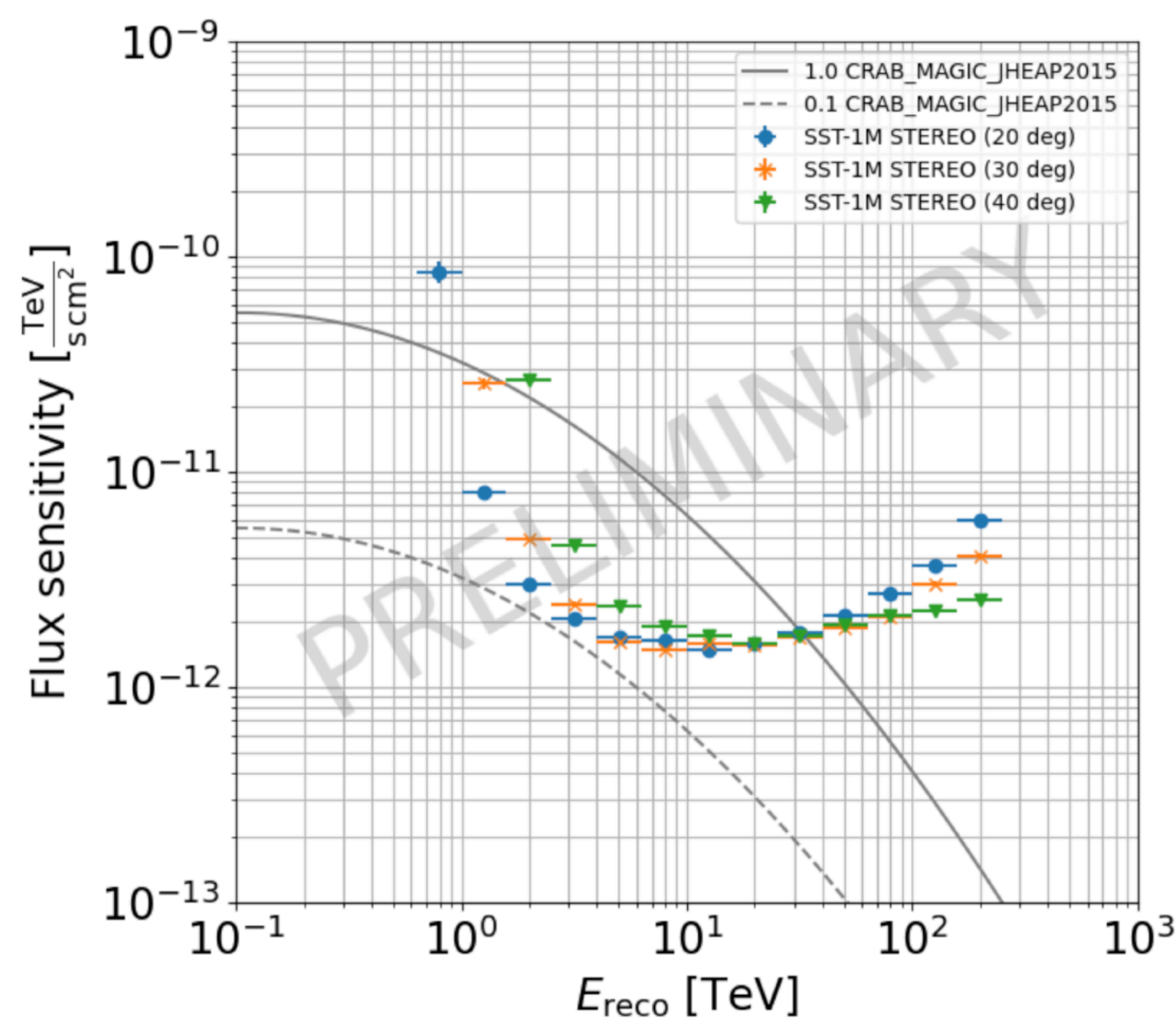
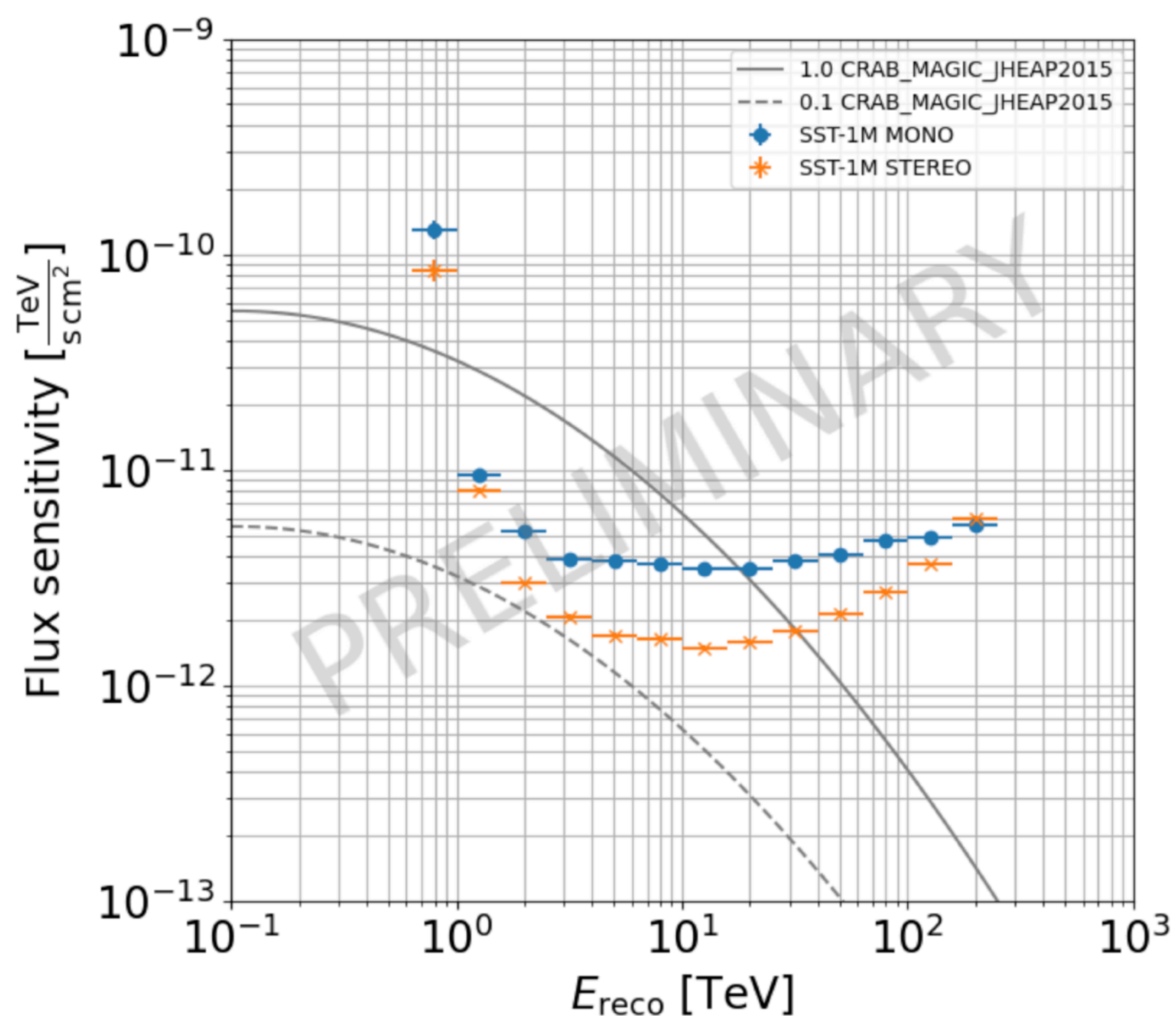
SST1-M physics performance

Differential sensitivity

Galactic PeVatron candidate studies

Monte Carlo

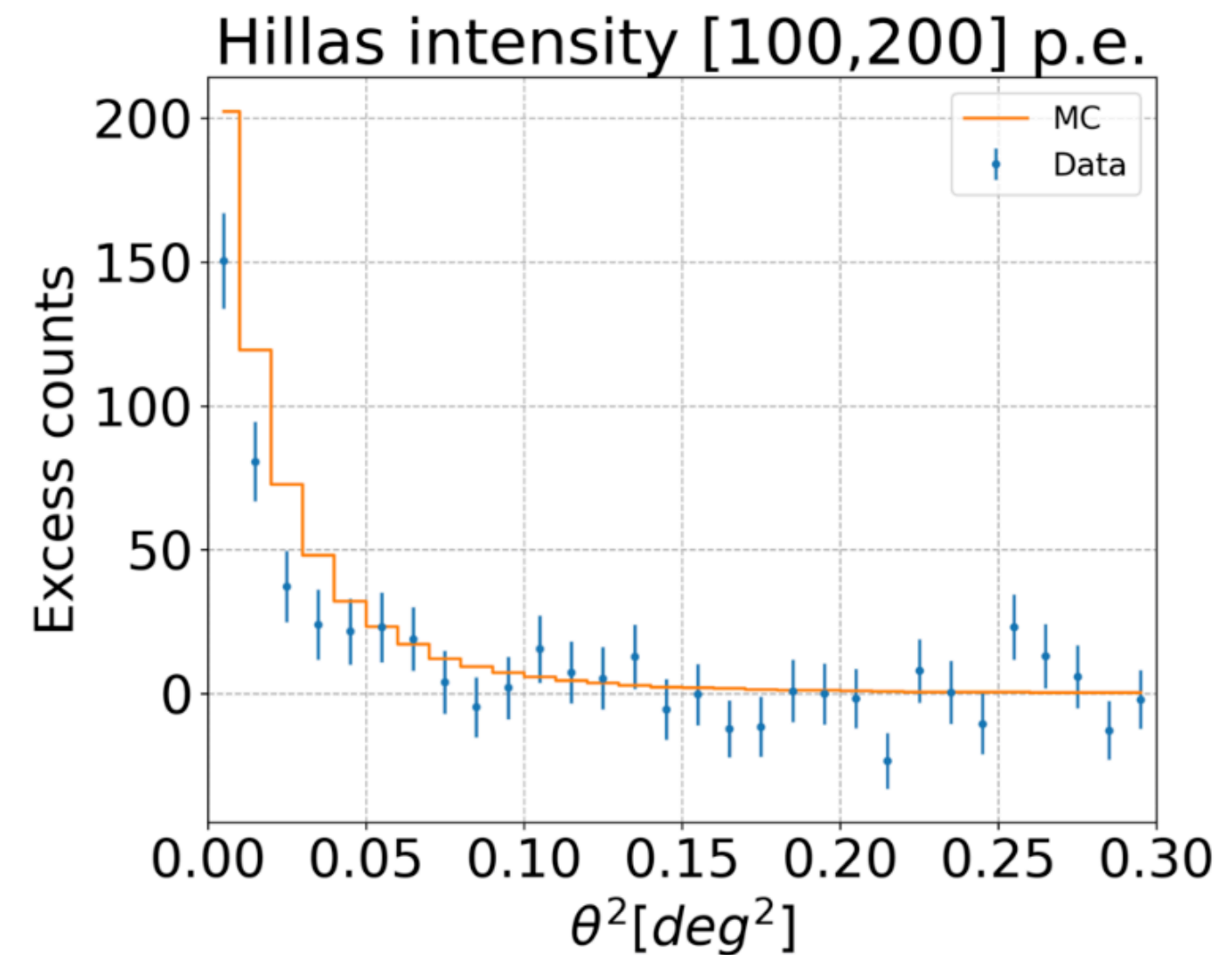
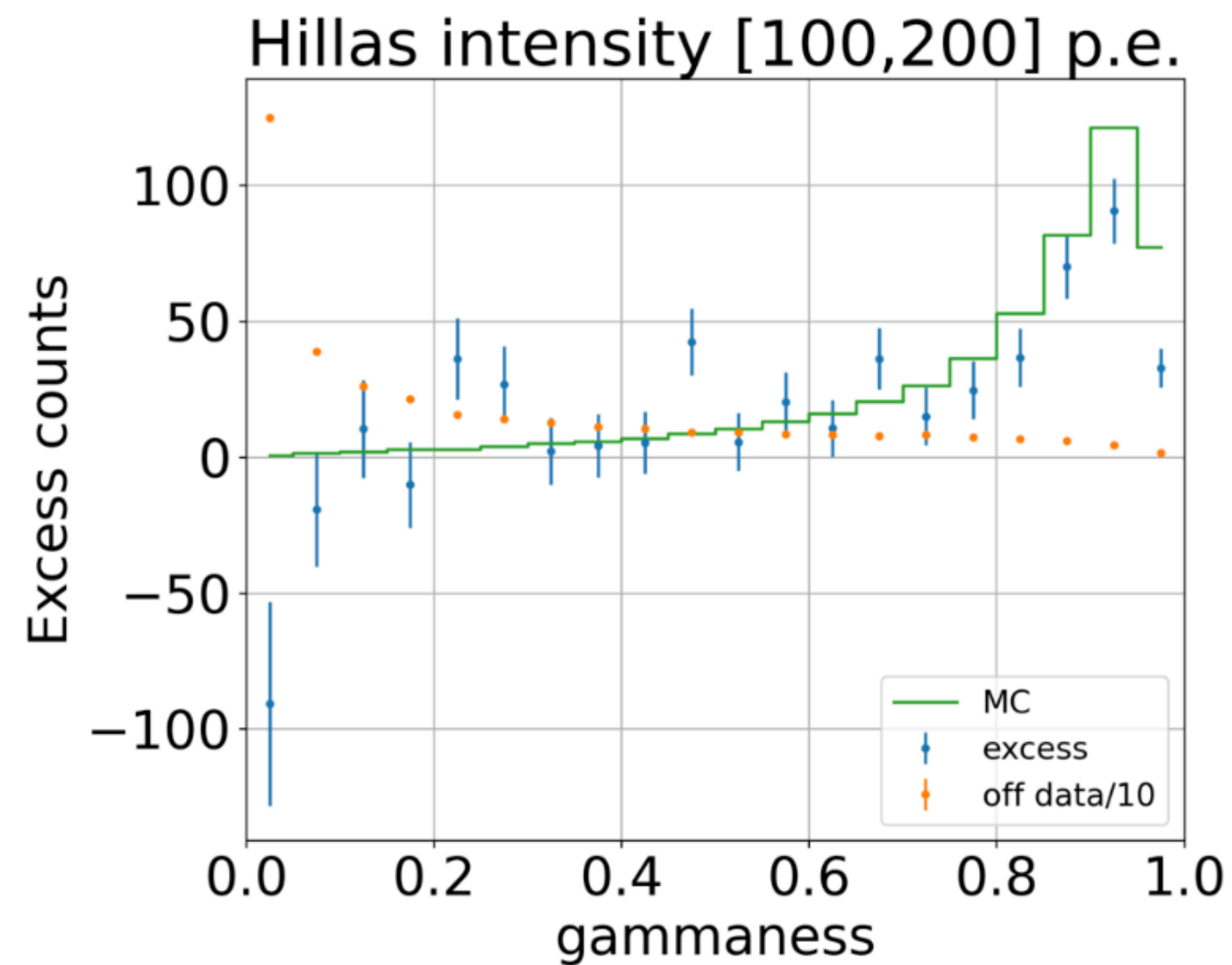
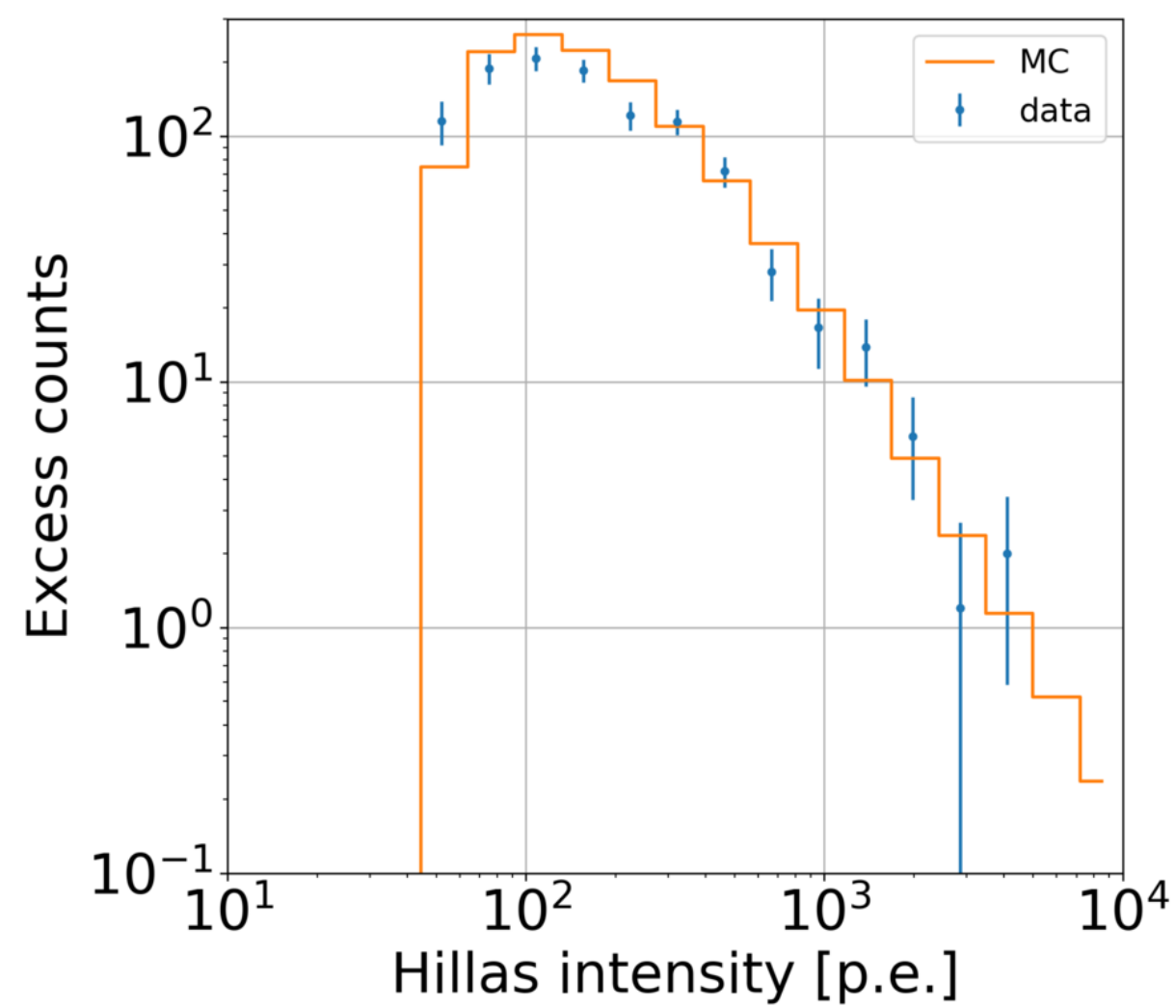
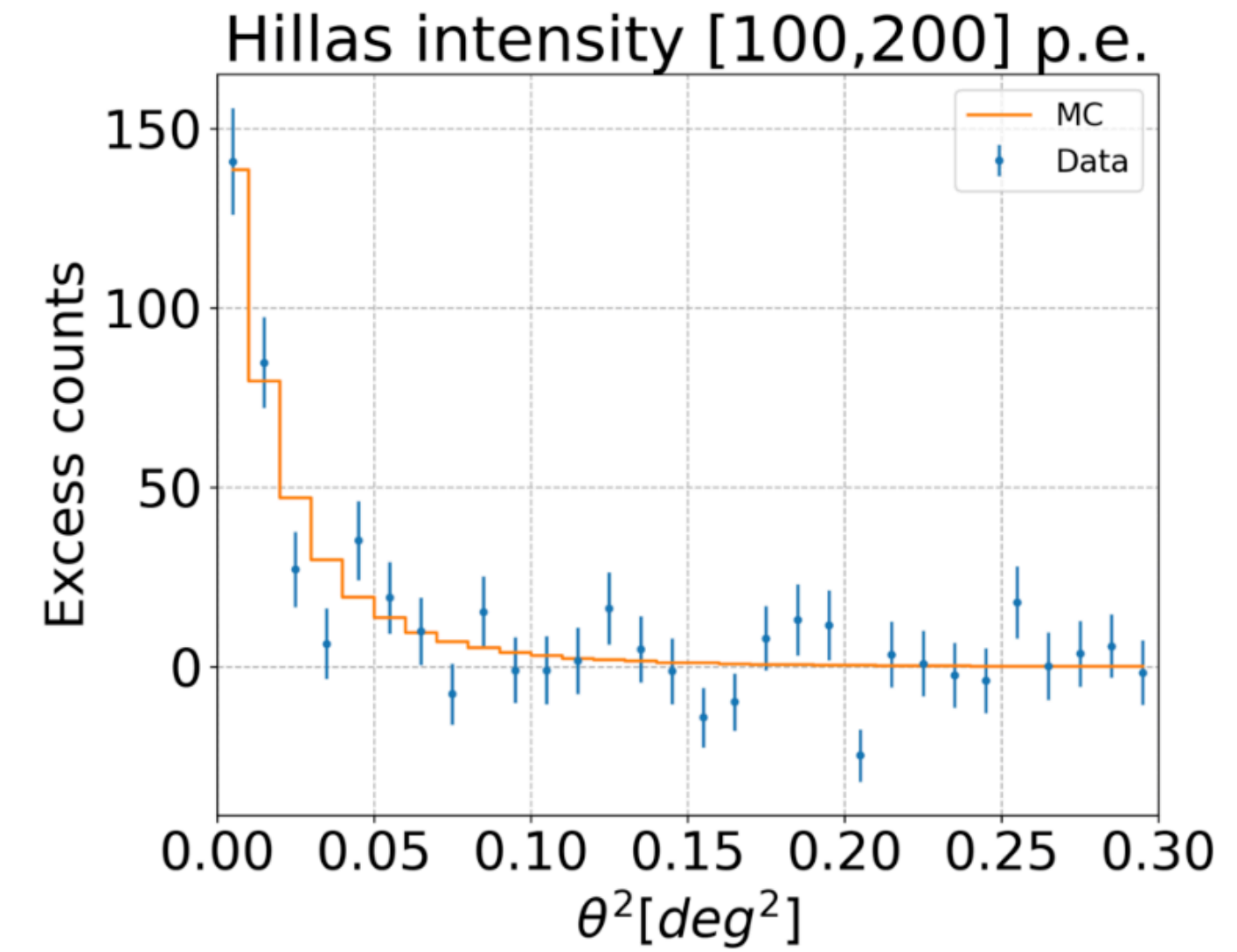
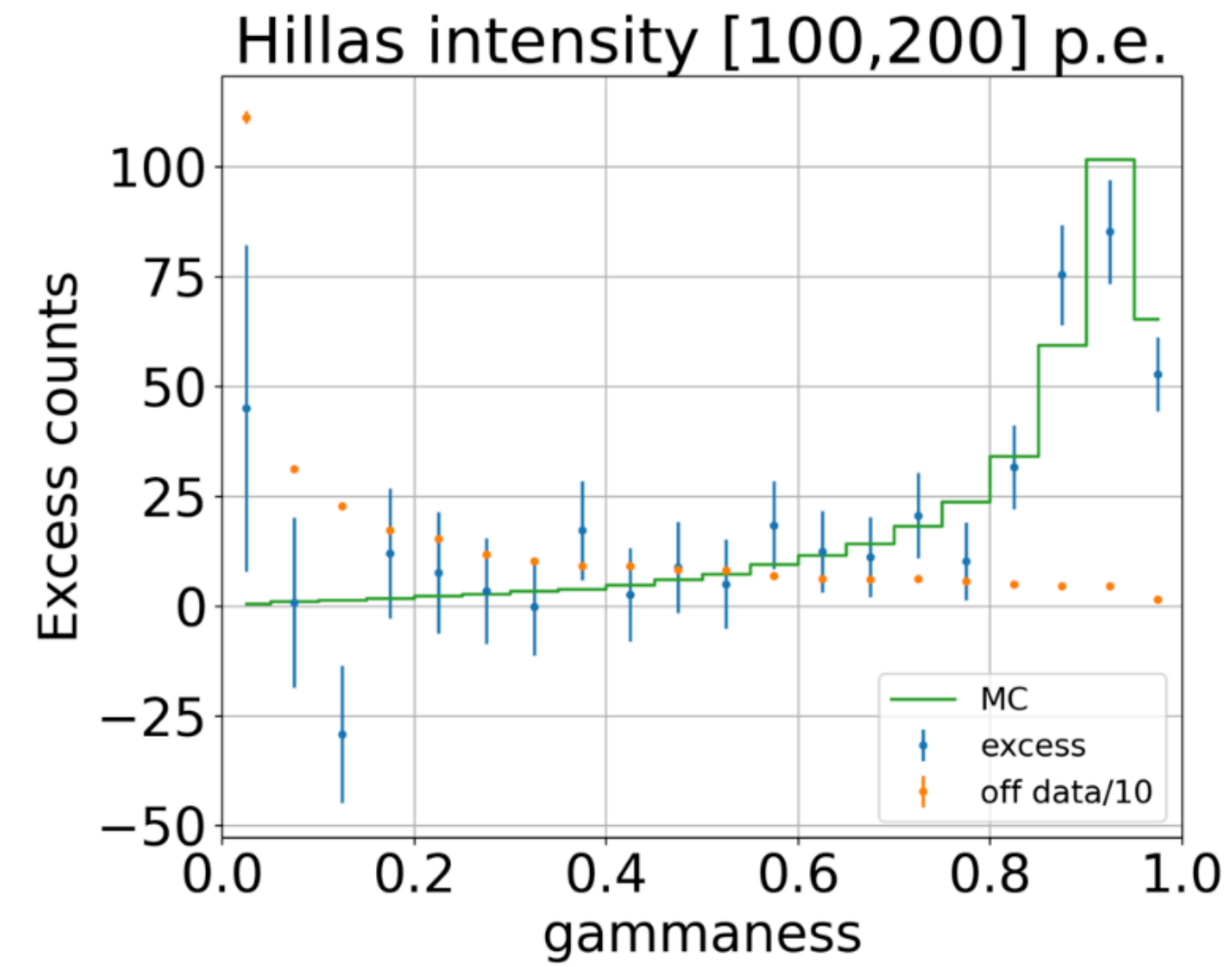
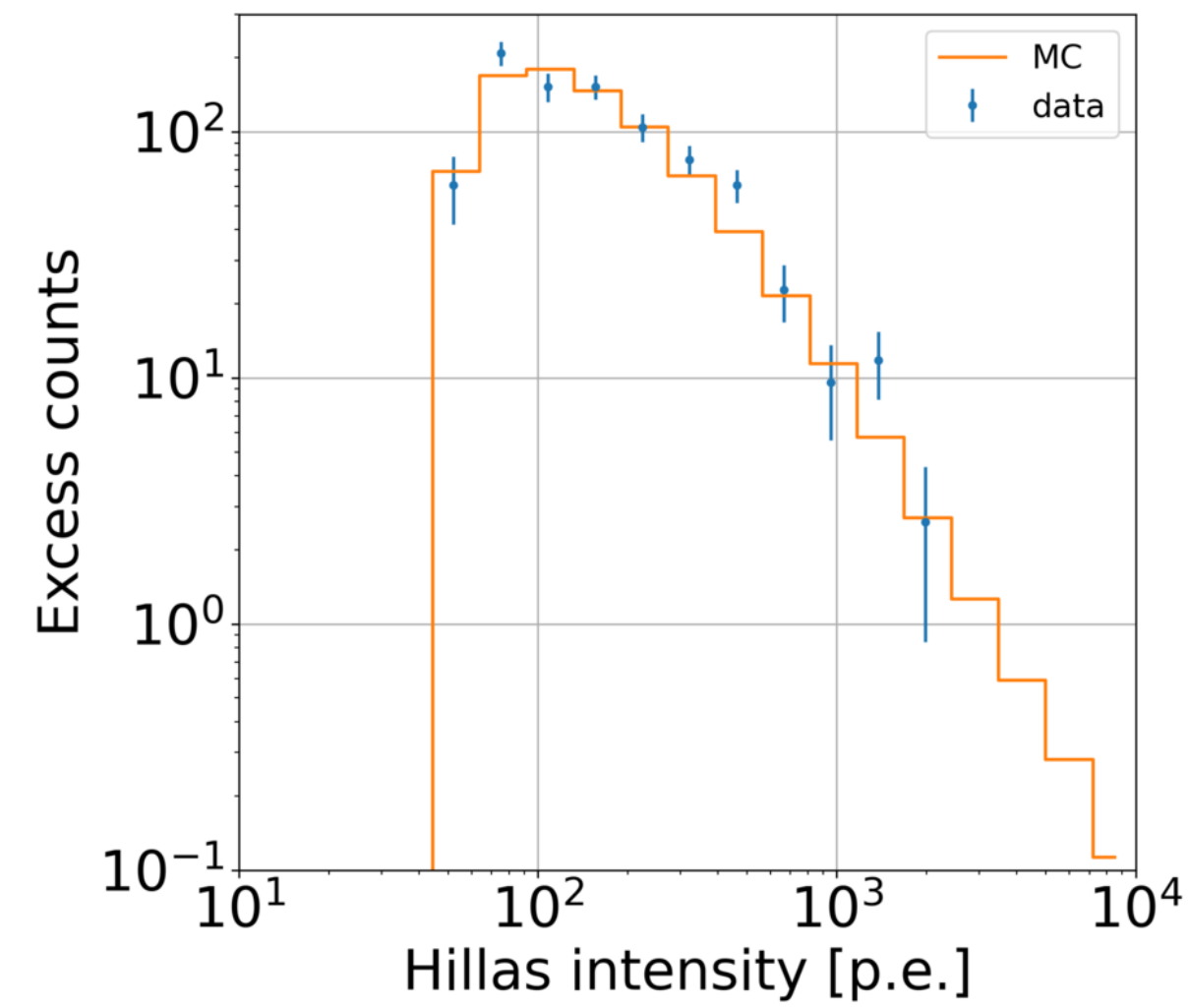
- Significant improvement of stereo over mono reconstruction at all energies (except for the effect of higher energy threshold)
- Given **large FoV** and **low altitude**, SST-1M probes the highest gamma-ray energies among existing IACTs!



MC - data agreement

Crab excess events

Real data
+ MC



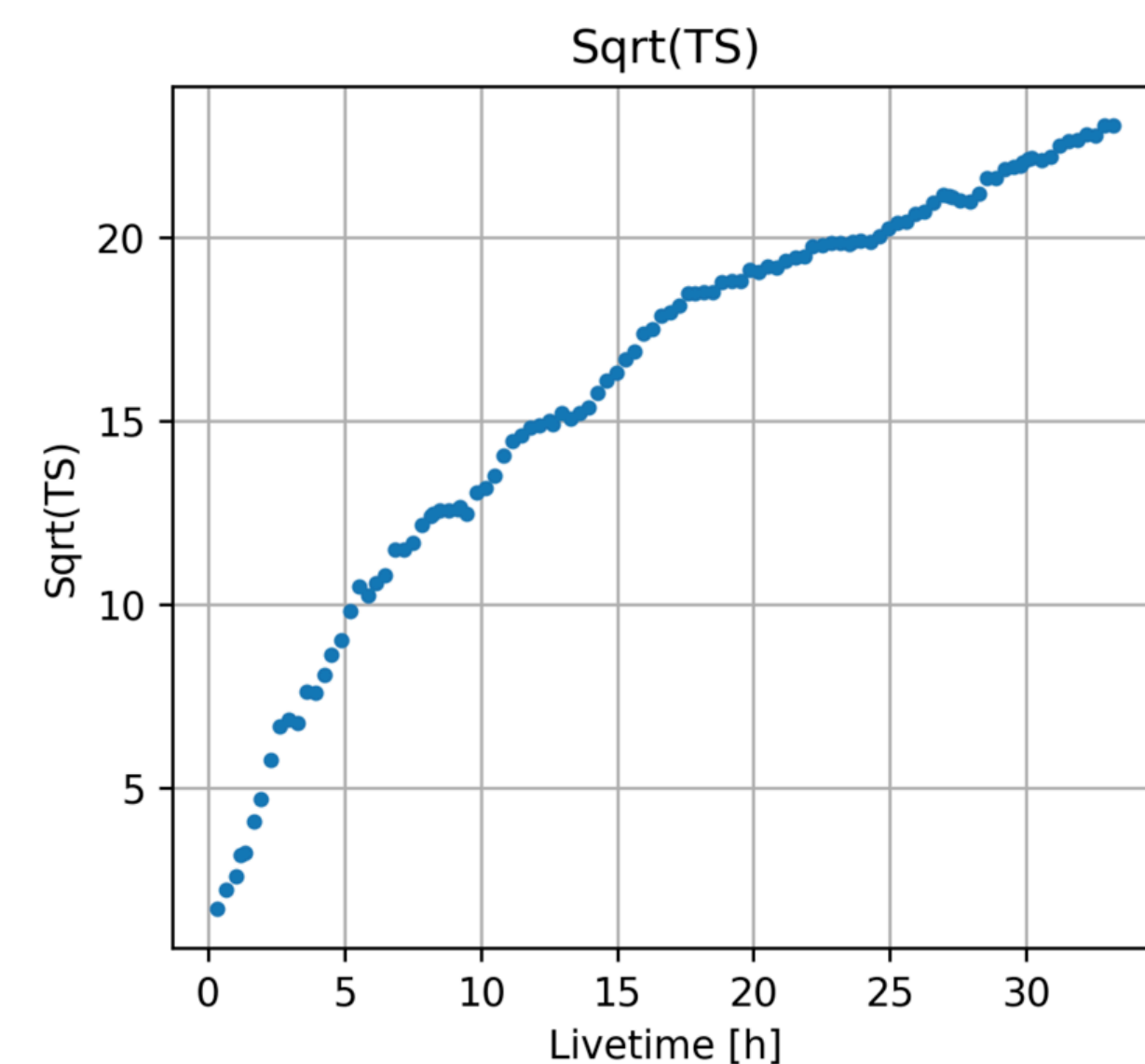
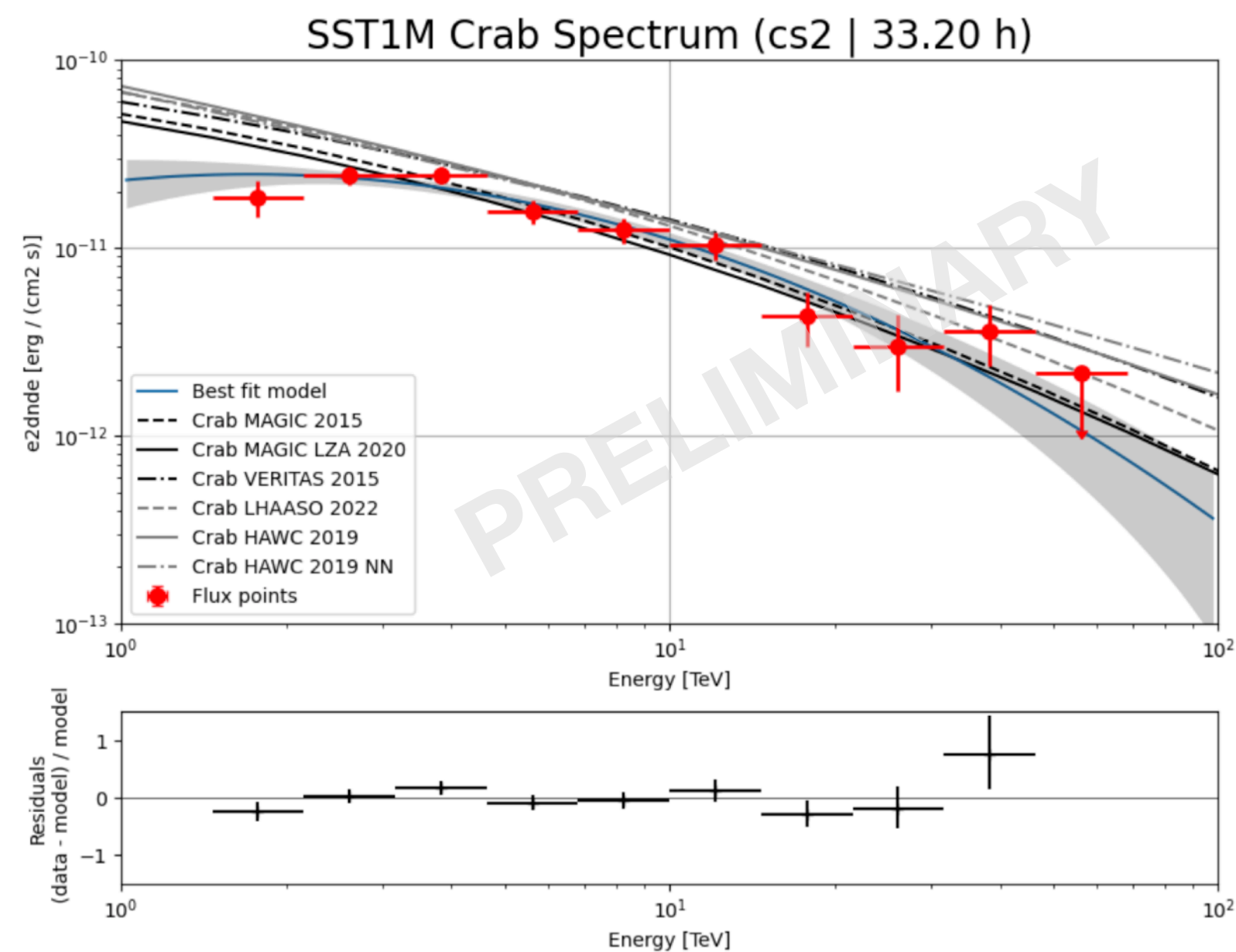
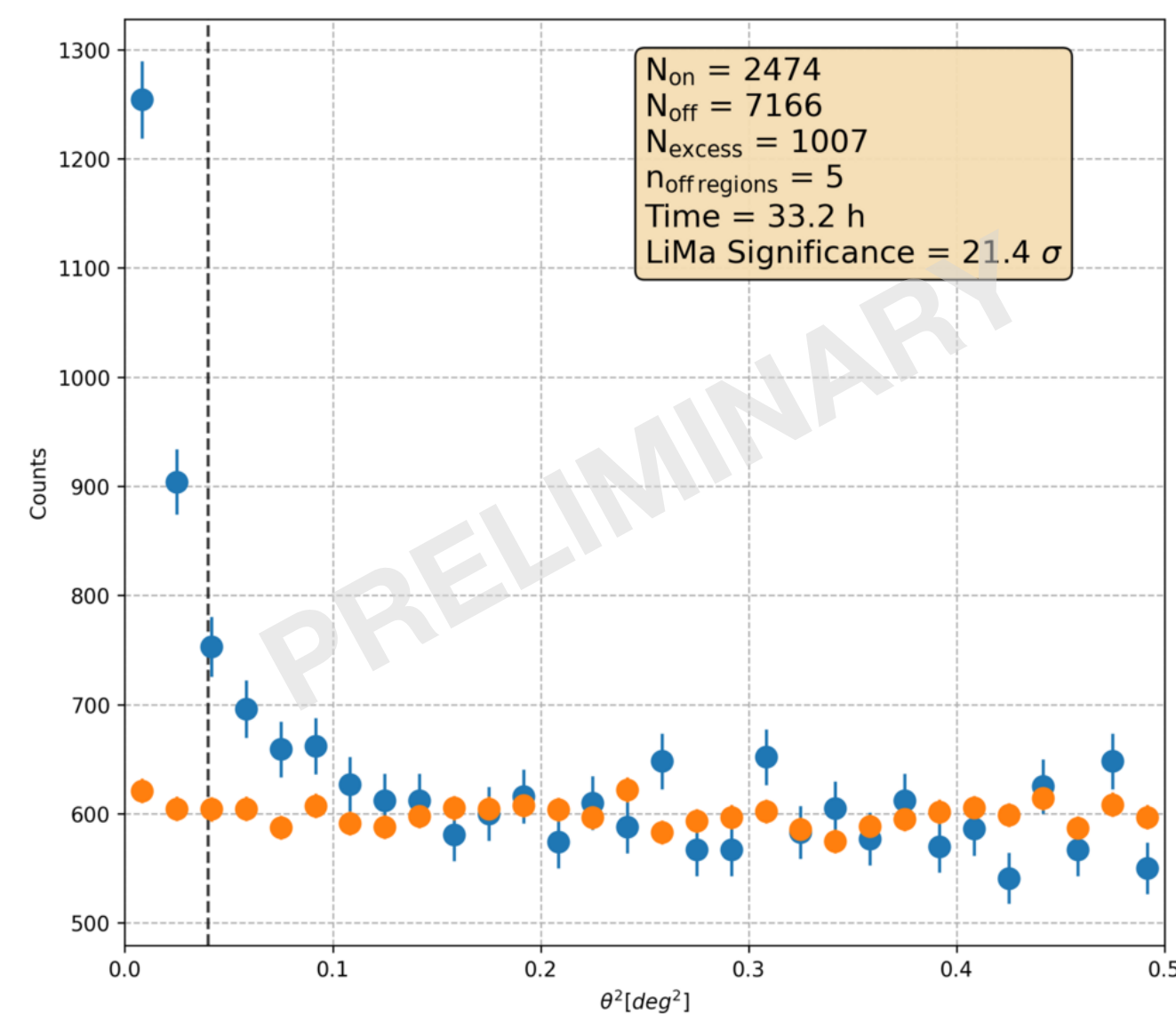
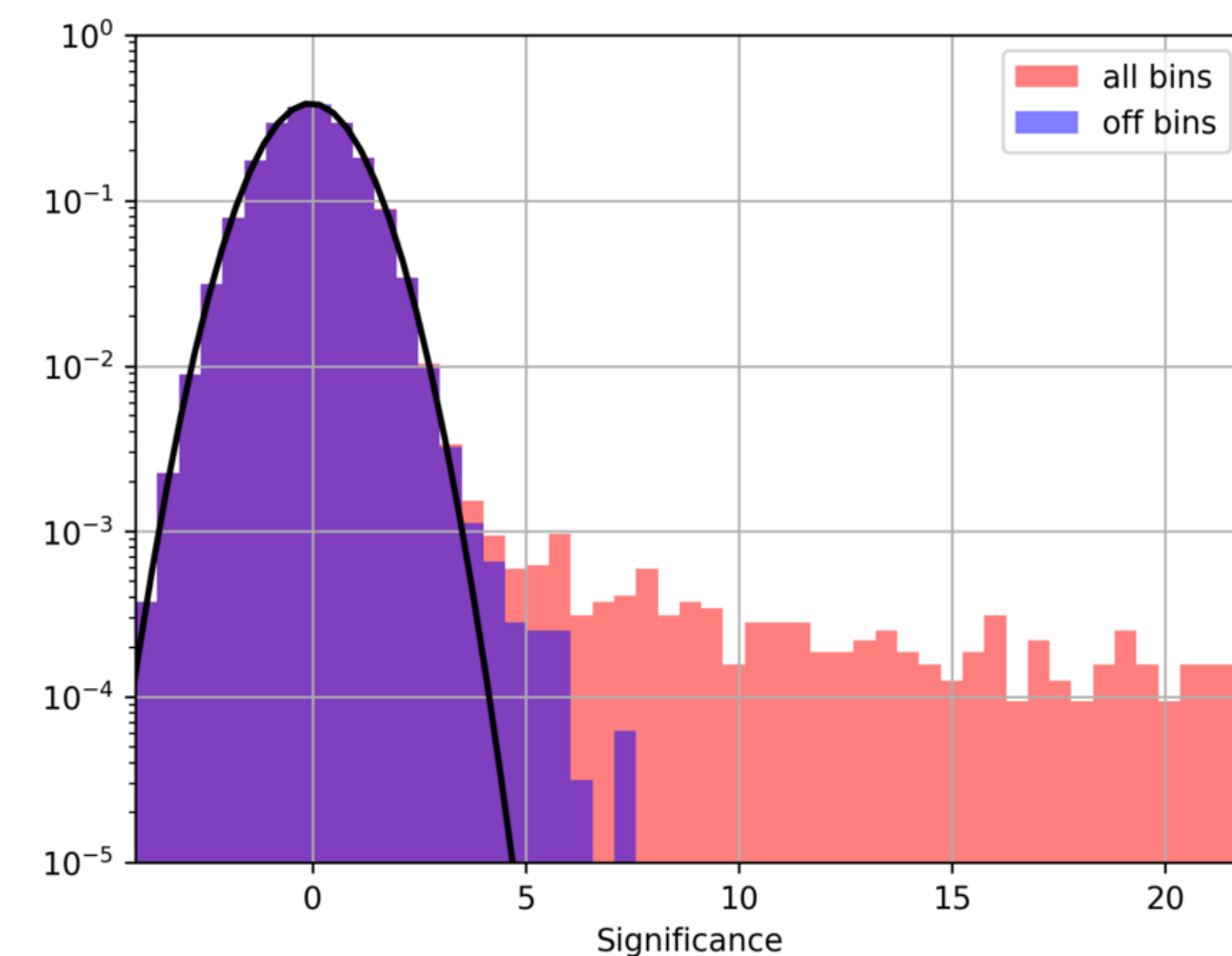
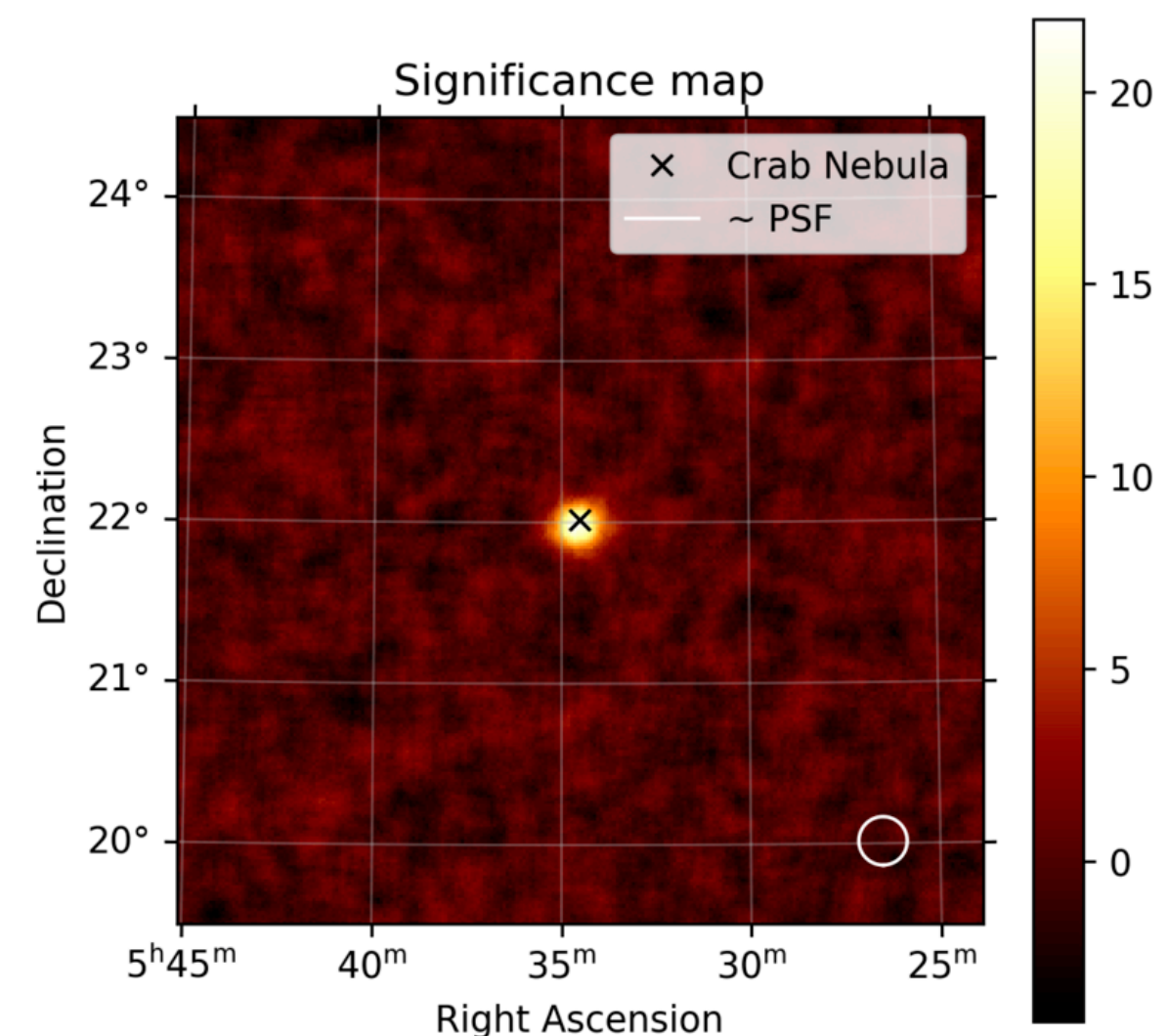
TeV 1

TeV 2

Crab Nebula observation - MONO

Real data

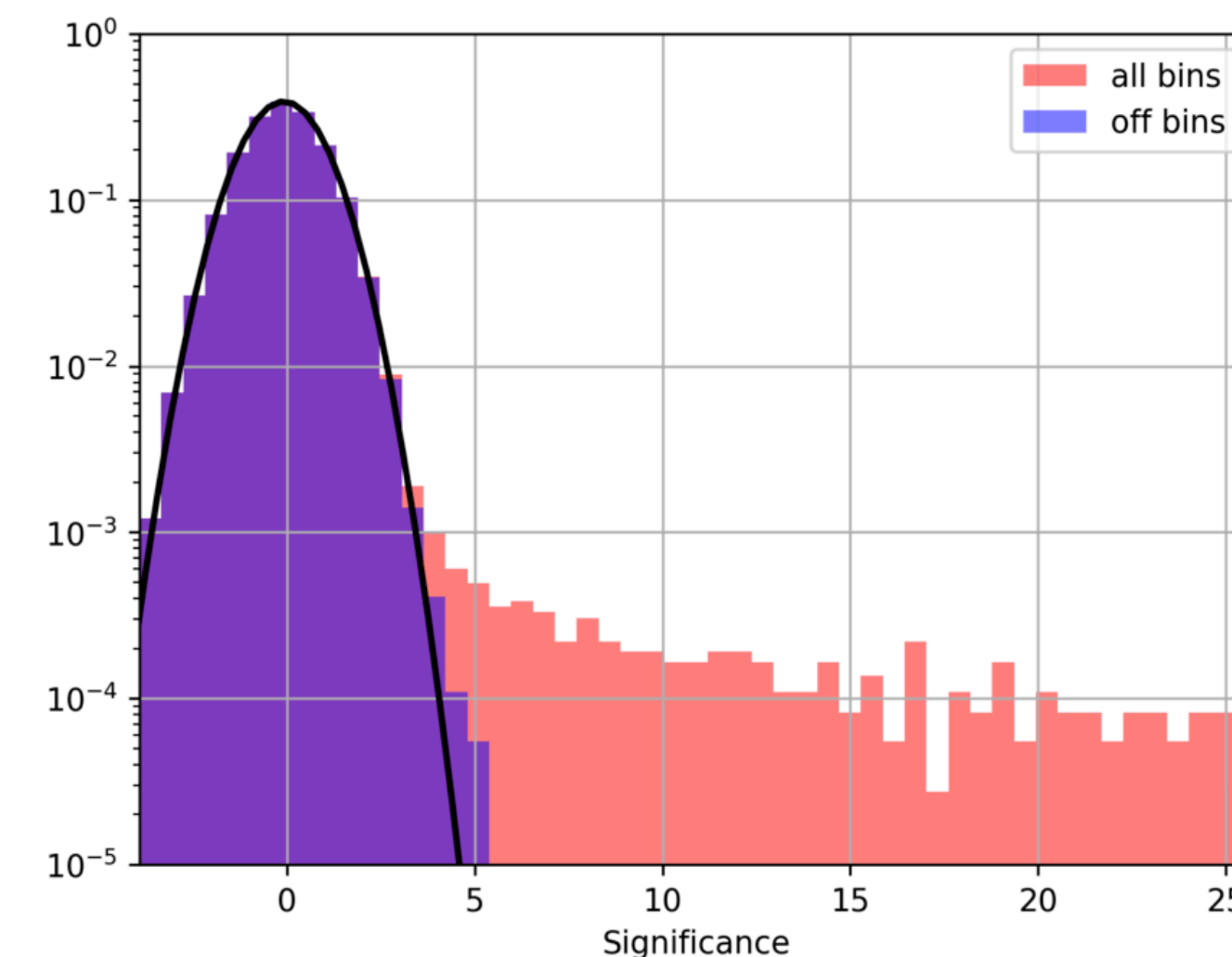
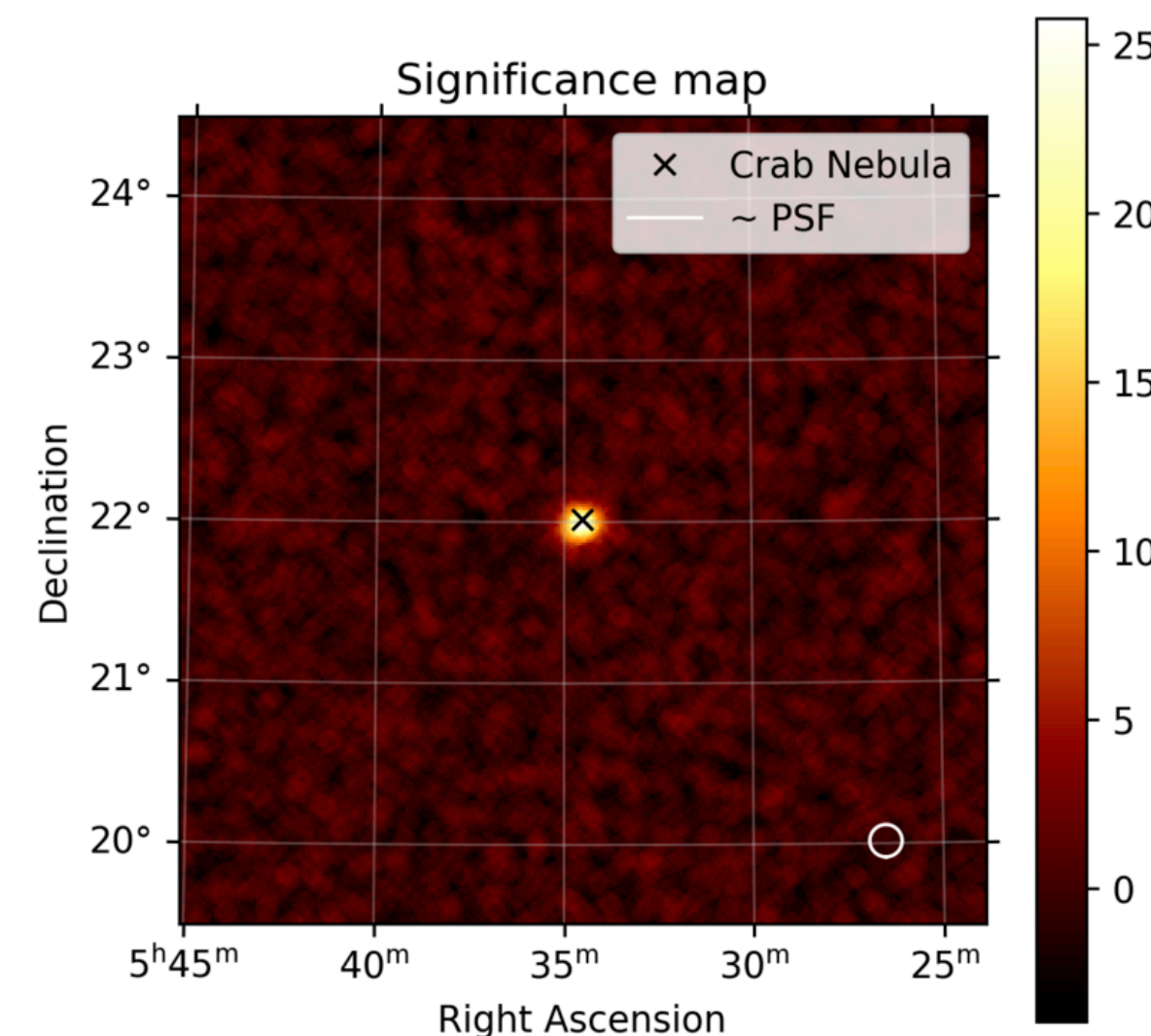
- Obs campaign 2023-2024, zenith angle 25-45 deg, energy threshold $\sim 2\text{-}3$ TeV
- ~ 33 hours of good mono data after quality cuts, **5sigma detection in ~ 2.5 h**
- **Excellent agreement** of the SED with the results of other experiments (**note the tension** between different observatories)



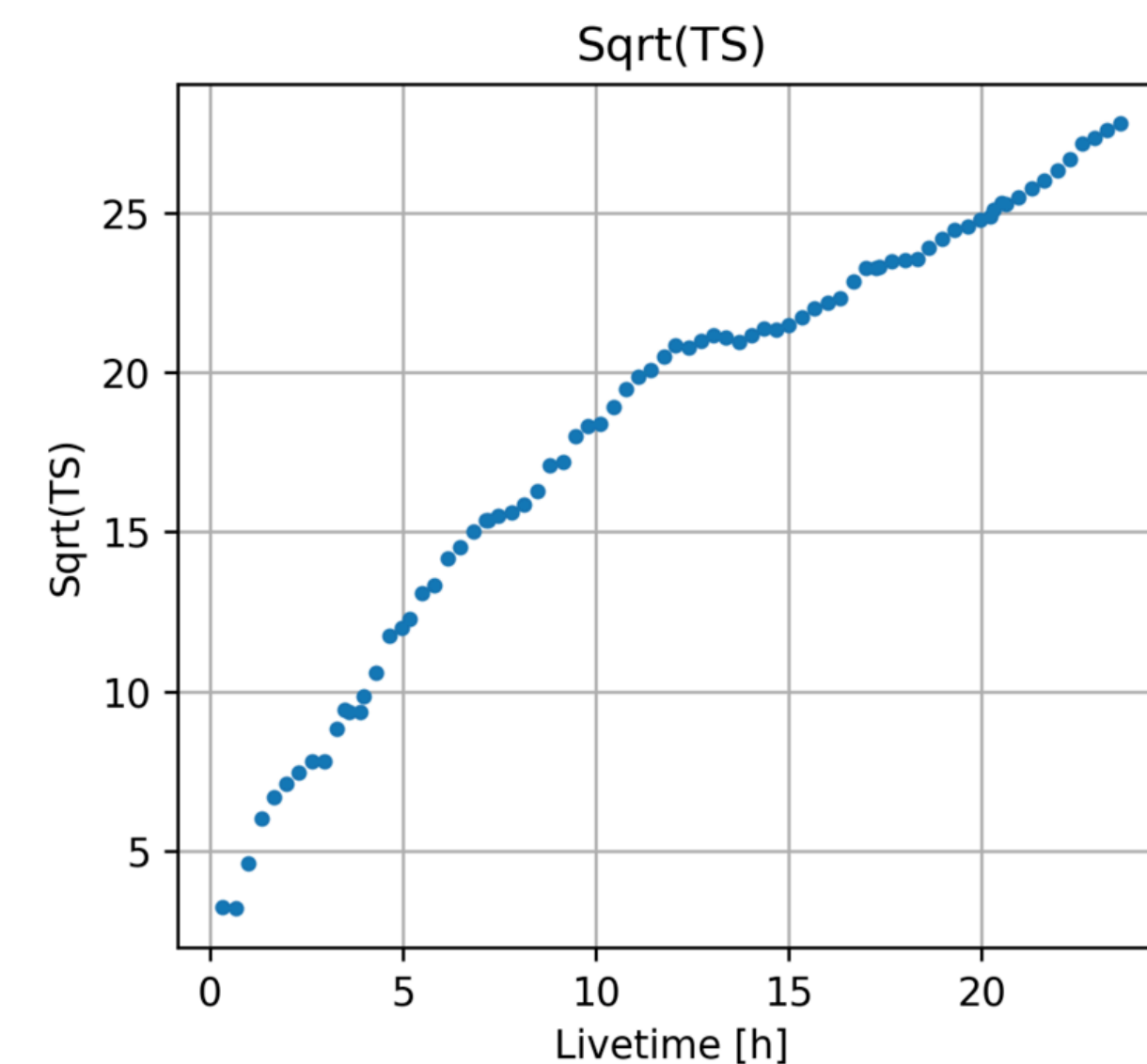
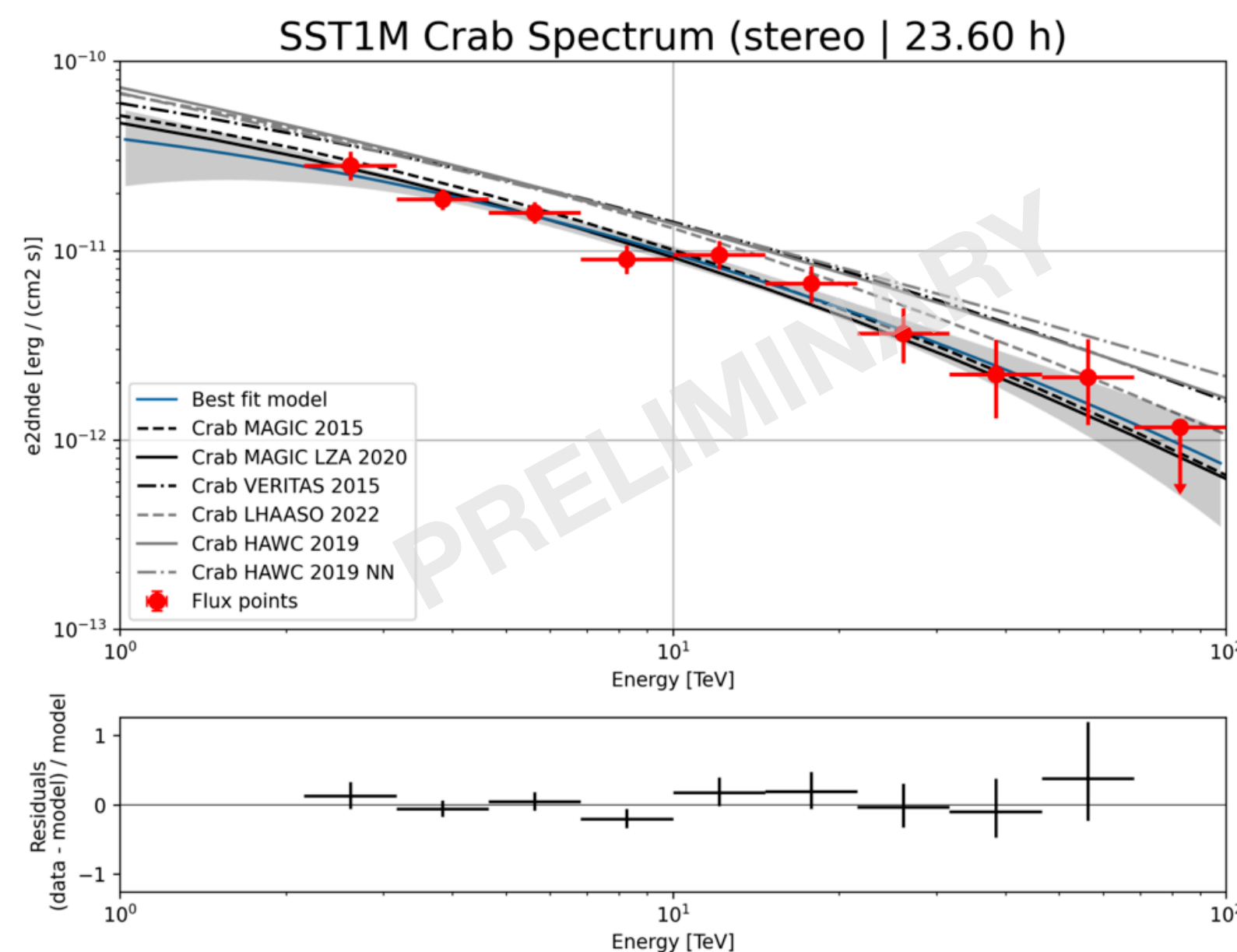
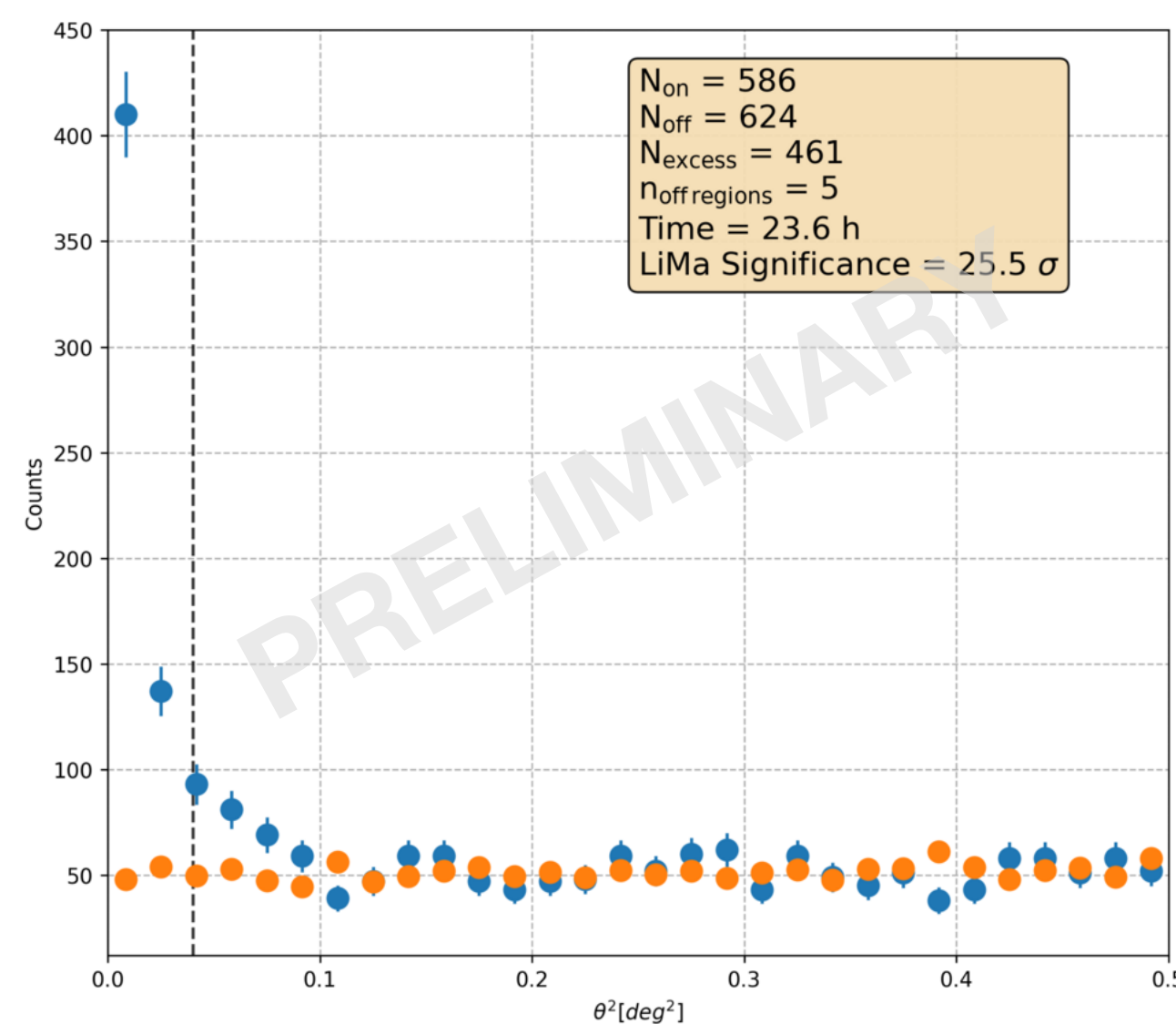
Crab Nebula observation - STEREO

Real data

- Higher energy threshold due to not fully optimised distance between the telescopes
- ~23h of good stereo data after quality cuts, **5sigma detection in ~1.5h**
- Expected performance improvement - angular resolution and sensitivity
- **Remarkable background homogeneity on a scale of a few degrees**



Excellent instrument for observation of extended sources

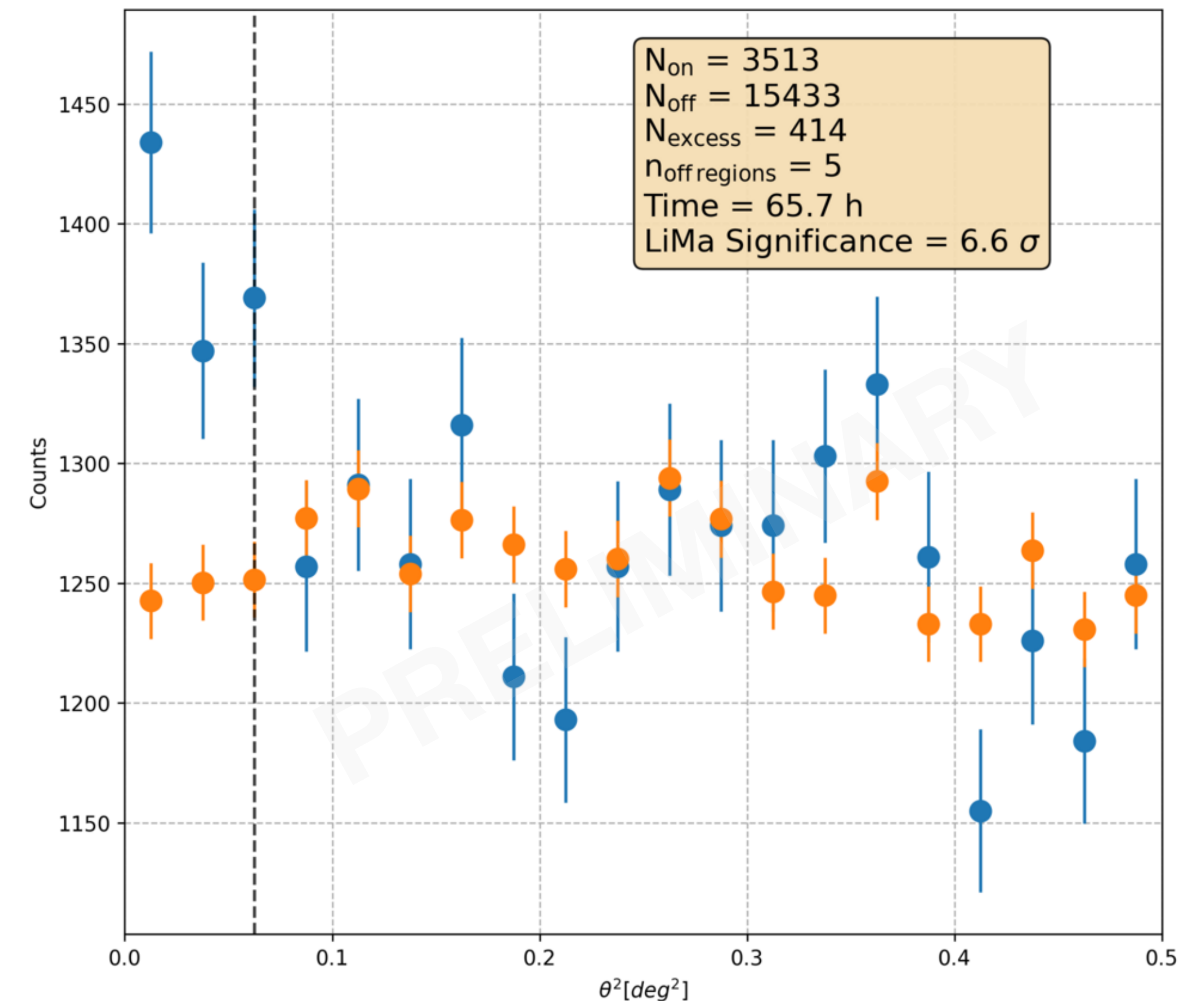


Nearby AGN monitoring

Observation campaign of several bright blazars

Real data

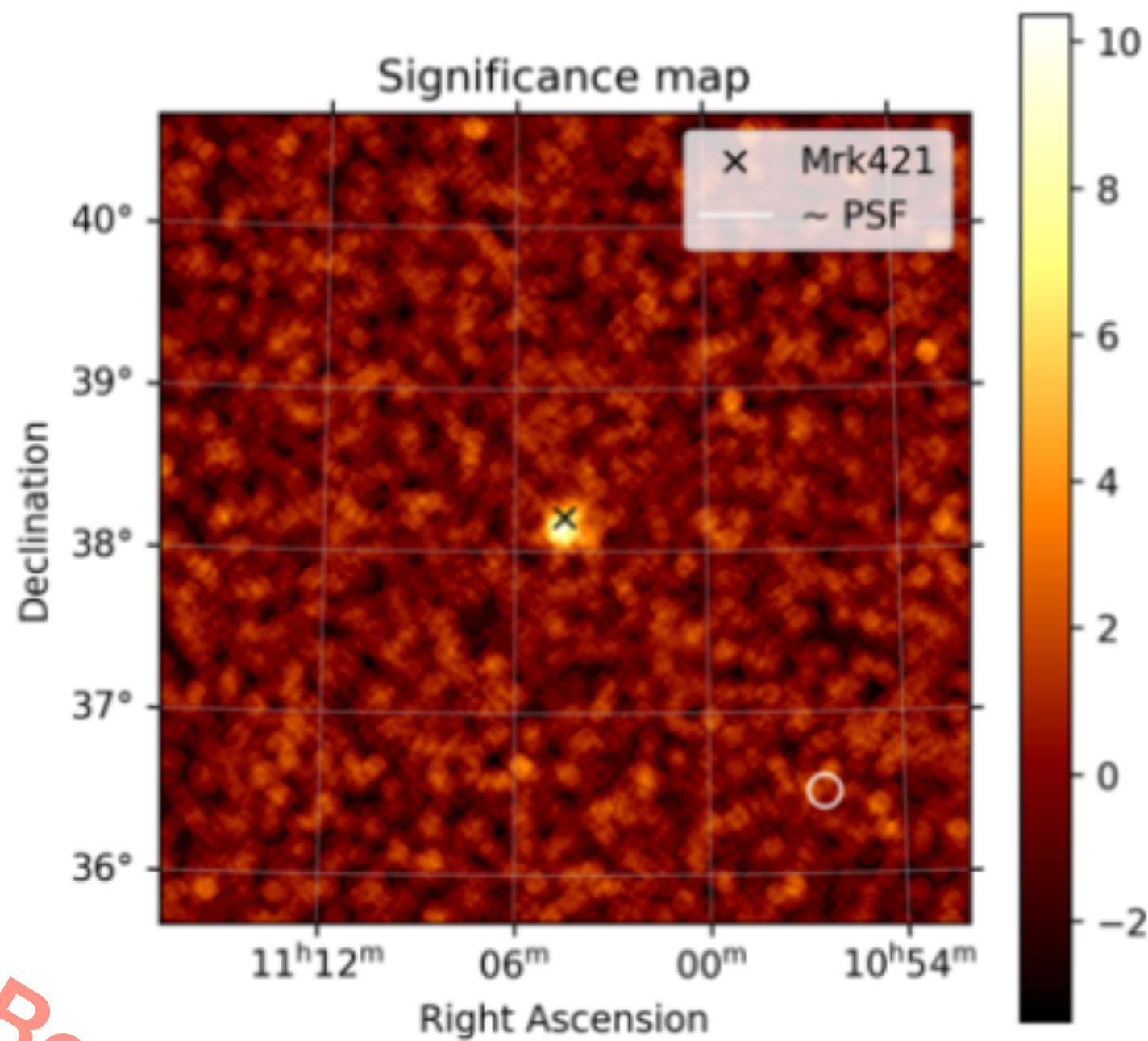
- **1ES1959+650** - first extragalactic source detected with SST-1M (Summer 2023)
- Maintenance of Tel2 camera - Tel1 **mono** only
- Long period of low activity, 5sigma detection in ~20h
- Preliminary analysis using proto-pipeline. No spectral analysis, but **proving SST-1M capabilities for AGN monitoring**



Nearby AGN monitoring

Observation campaign of several bright blazars

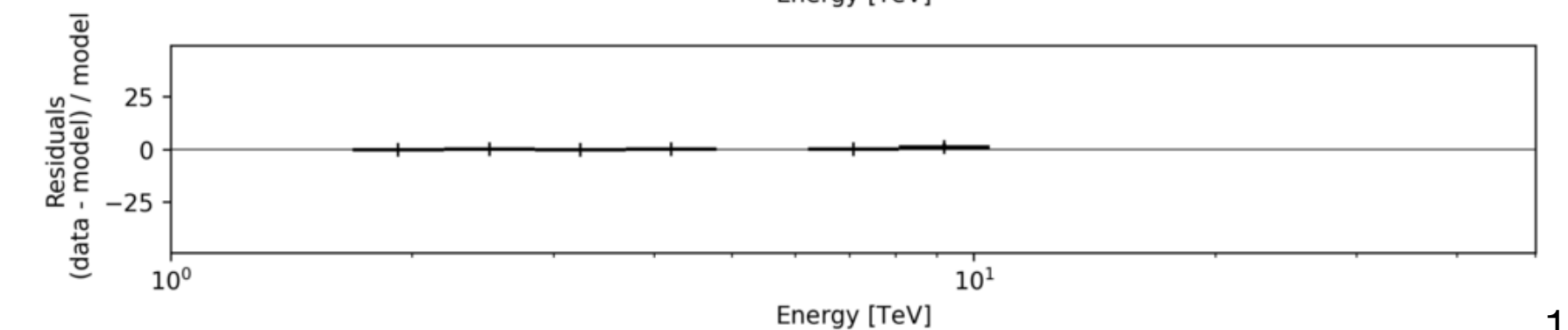
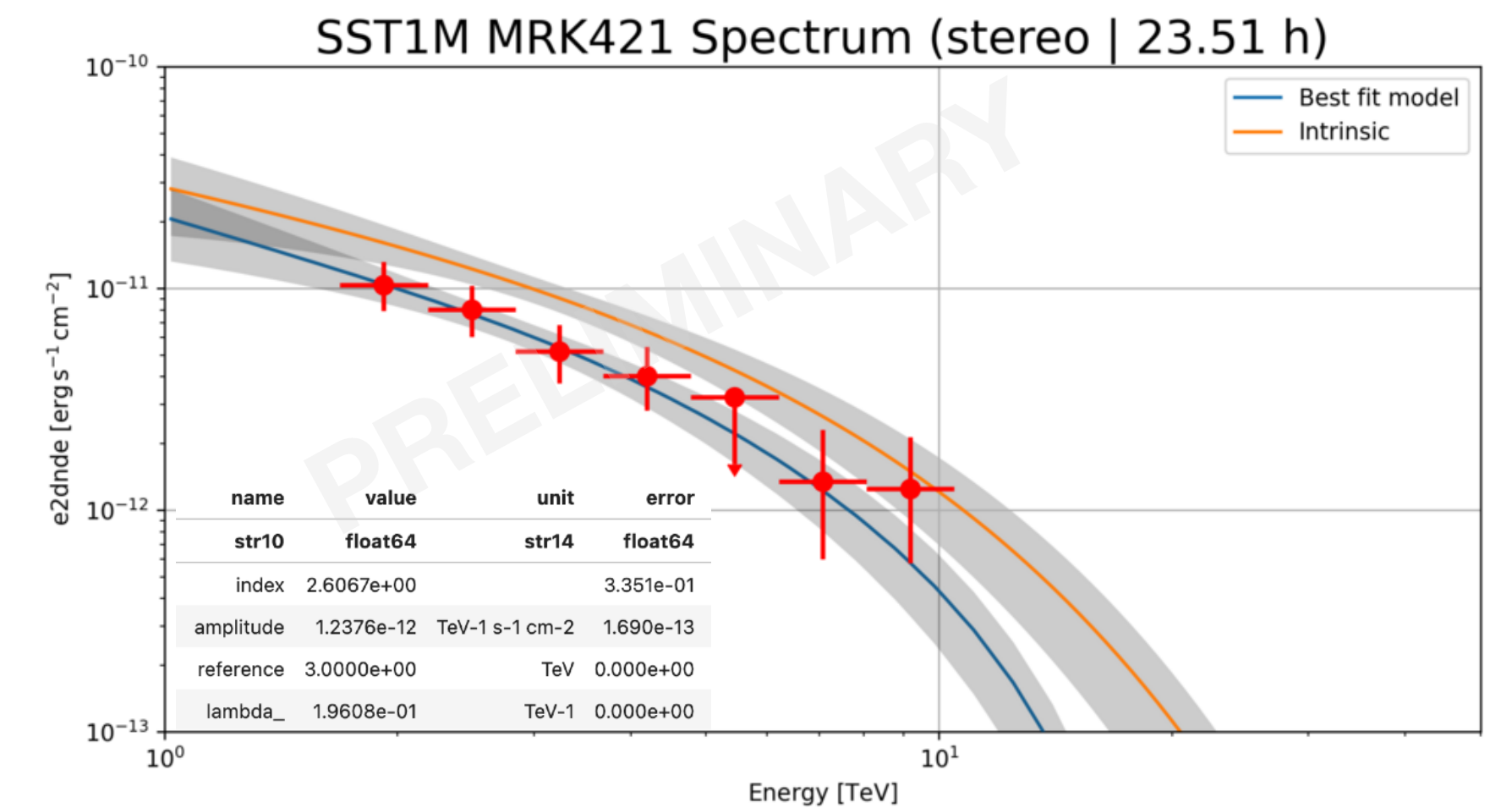
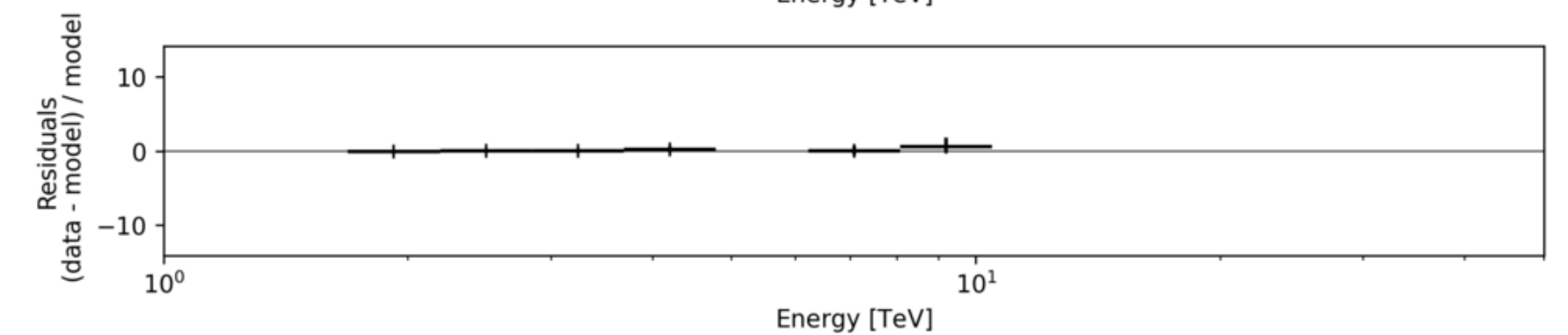
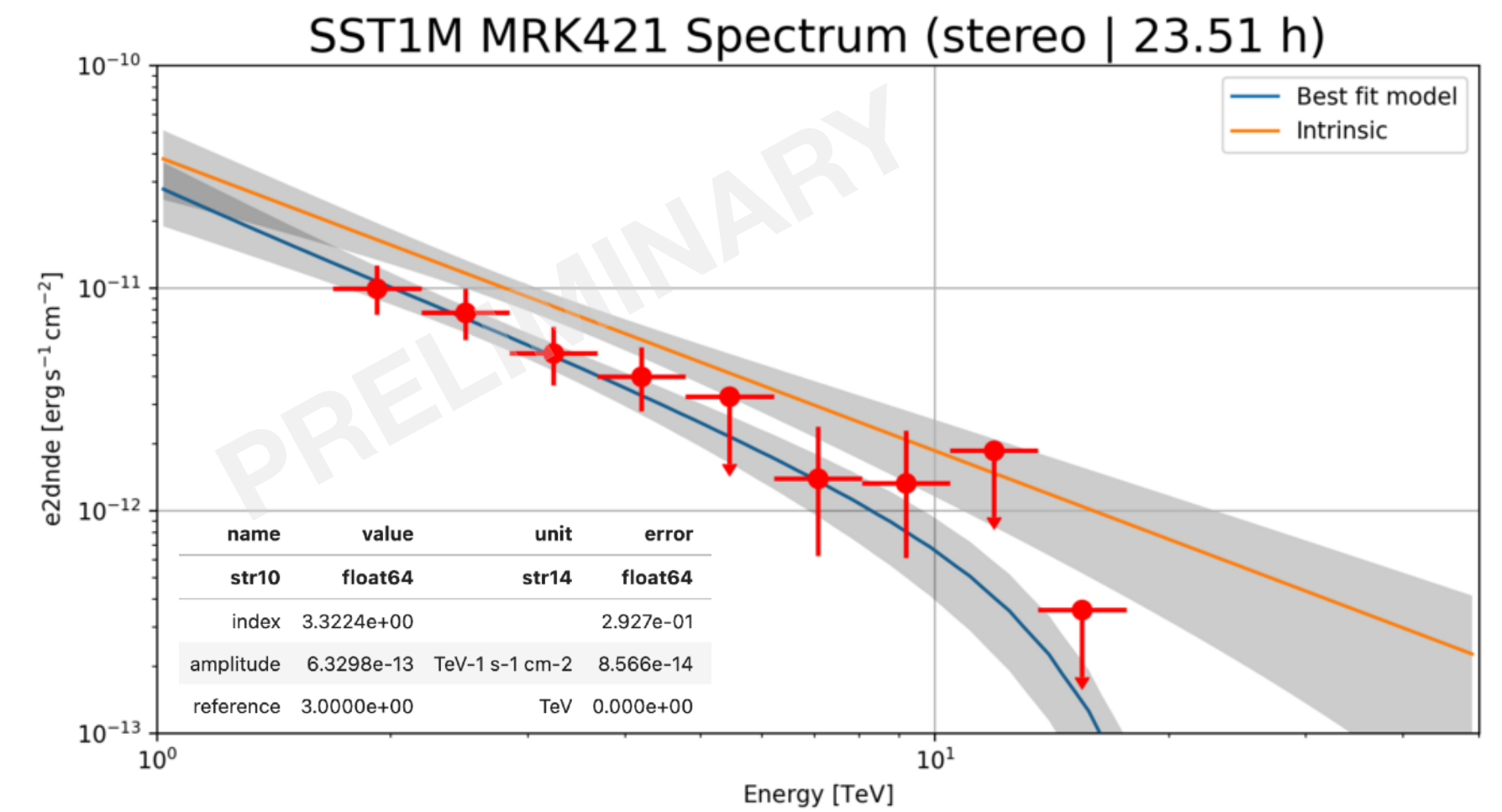
- **Mrk 421** - first extragalactic source detected in **stereo** mode (Spring 2024)
- High state detected on 15 March 2024: ATel #16533



- **Mid-term monitoring (~4 months):**

- integrated SED shows no spectral curvature*
- Sp index from ECPL fit (2.6±0.3) compatible with HAWC 2022 (2.26±0.12)
- last 2sigma fluxpoint at 9 TeV (compare with HAWC 2022)

* DeltaTS=0.03 for intrinsic ECPL over PL, probably because we probe energies already above the $E_{\text{cutoff}} = 5.1$ TeV (HAWC, fixed in the fit)

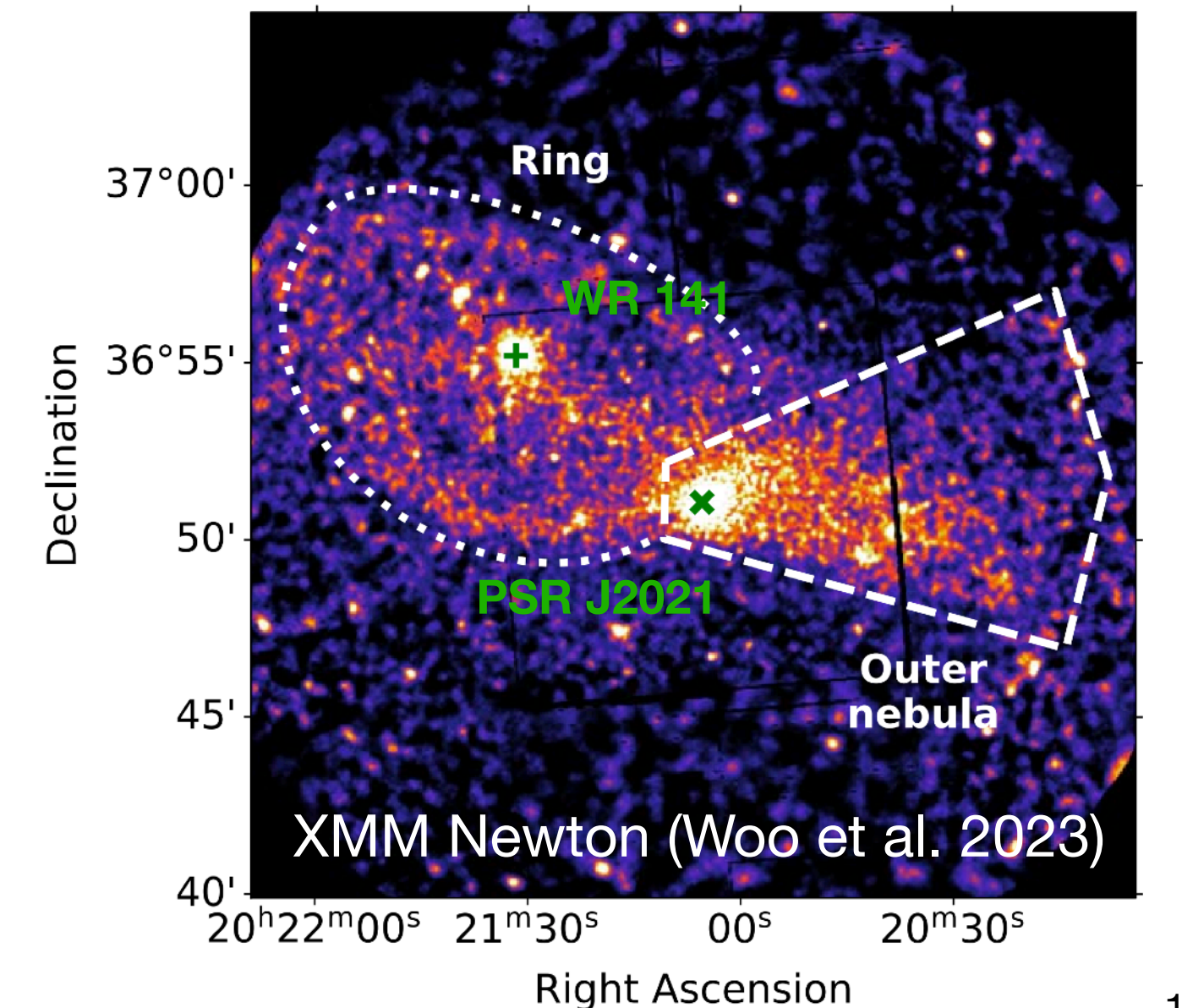
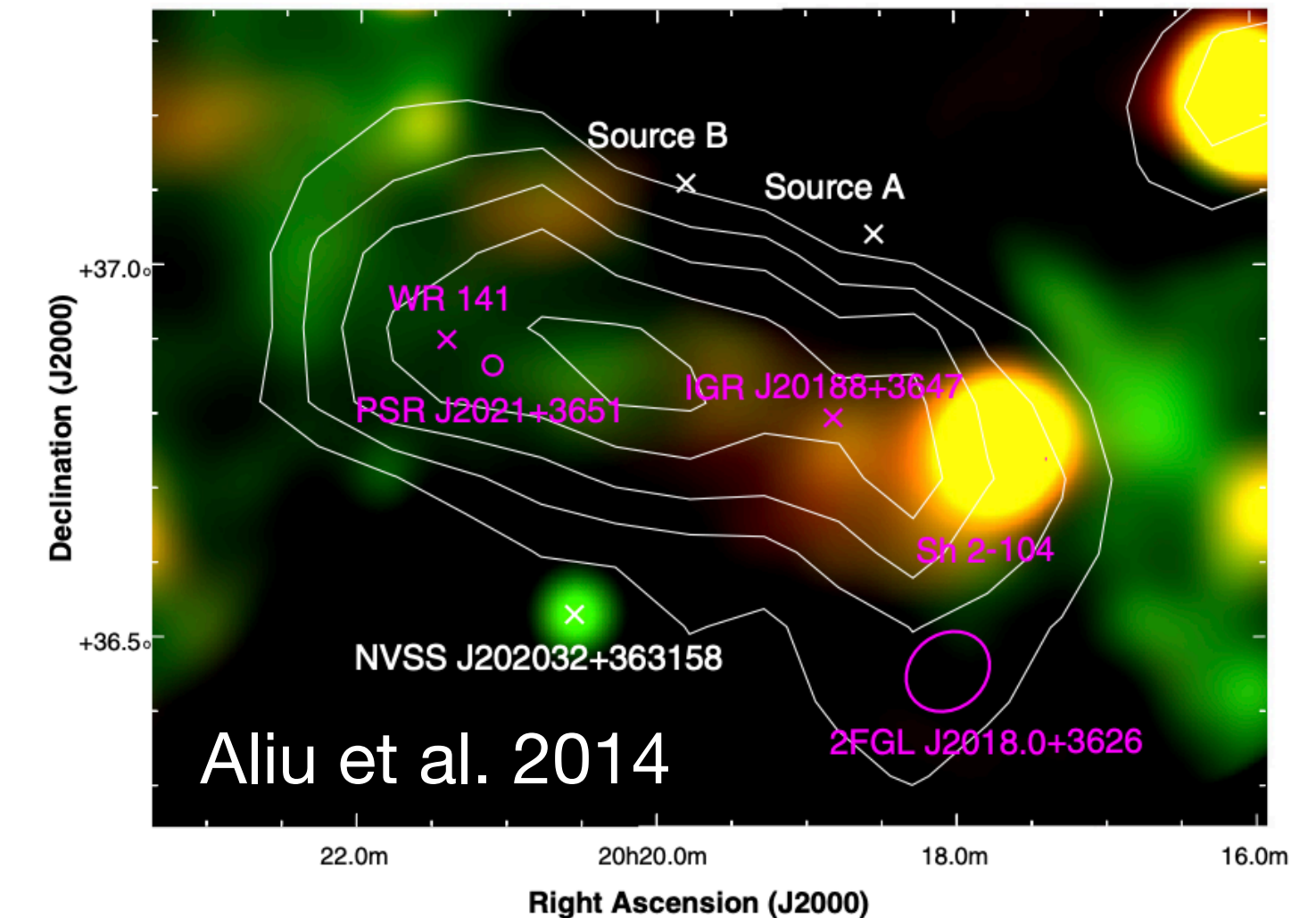


VER 2019+368 (Dragonfly)

SST-1M enters the realm of extended sources

- Discovered by MILAGRO (Abdo et al. 2009)
- Later resolved into 2-3 sources by VERITAS (Abeysekkhara et al. 2014, 2018)
- Photons up to 0.27 PeV (LHAASO 2021)
- Slightly extended: ~ 0.5 deg, highly asymmetric, energy dependent morphology
- **Complex region:**
 - Several radio, X-ray, HE and VHE sources
 - Supernova Remnant CTB 87
 - Two pulsars
 - Sh 2-104 - Star forming HII region
 - G75.2+0.1 - PWN
 - IGR J20188 - fast X-ray transient
 - Wolf-Rayet star WR 141
- **Recent modeling (Woo et al. 2023):**
 - PSR J2021+3651 (17 kyr, 10^{36} erg/s) + PWN (G75.2+0.1/Dragonfly) interacting with SNR reverse shock

Tension between SED from different experiments
There is no detailed study of VHE morphology

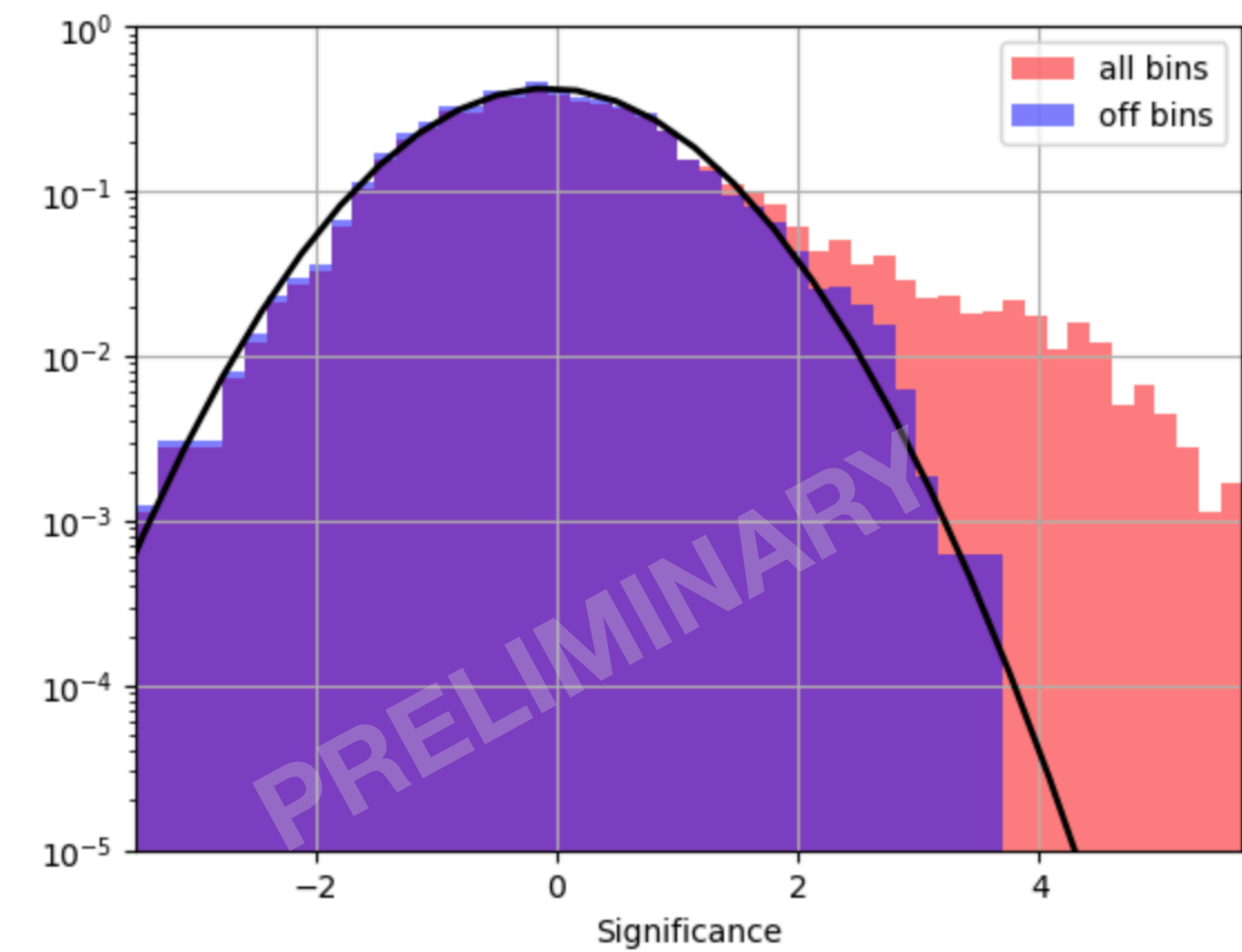
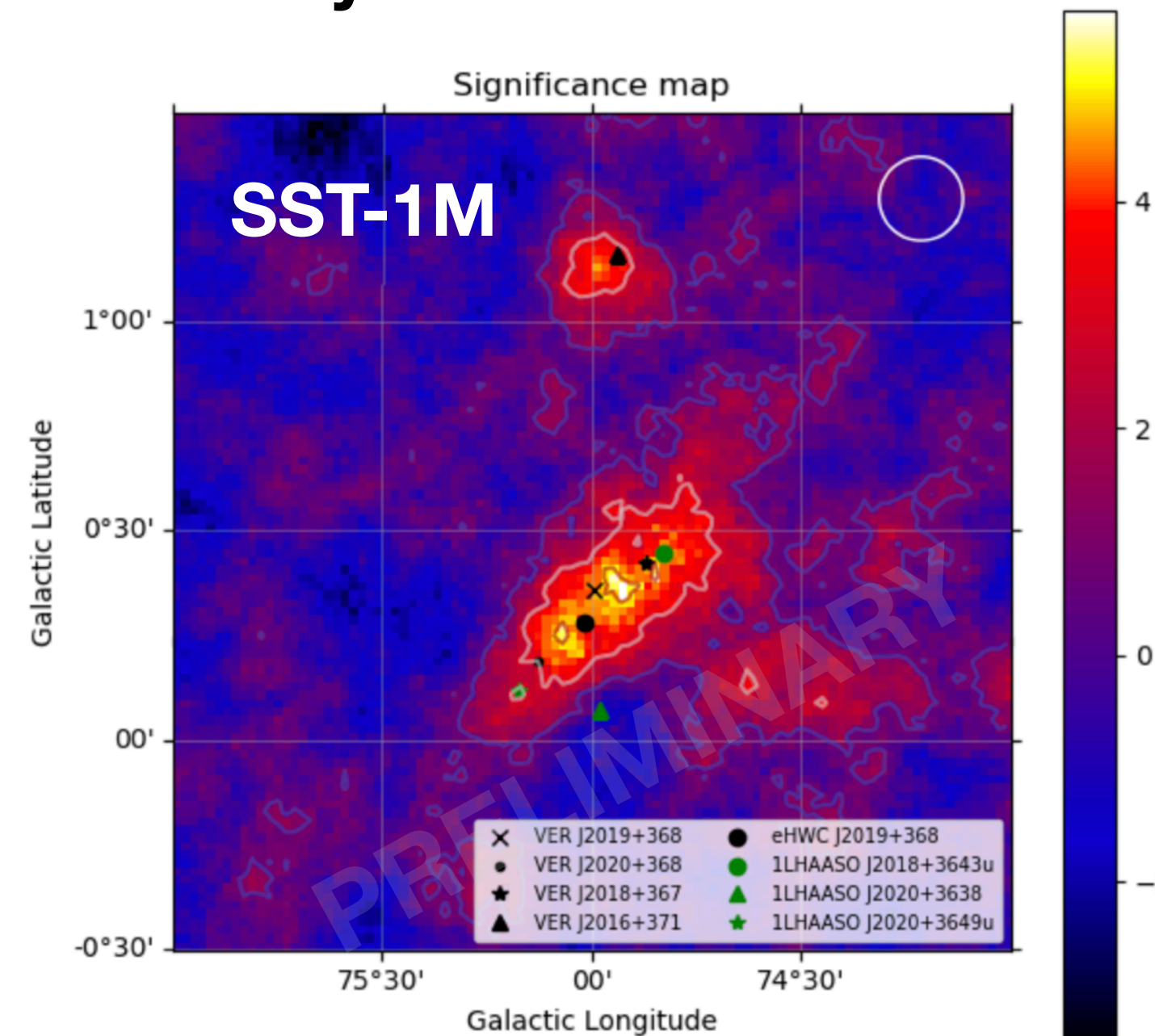
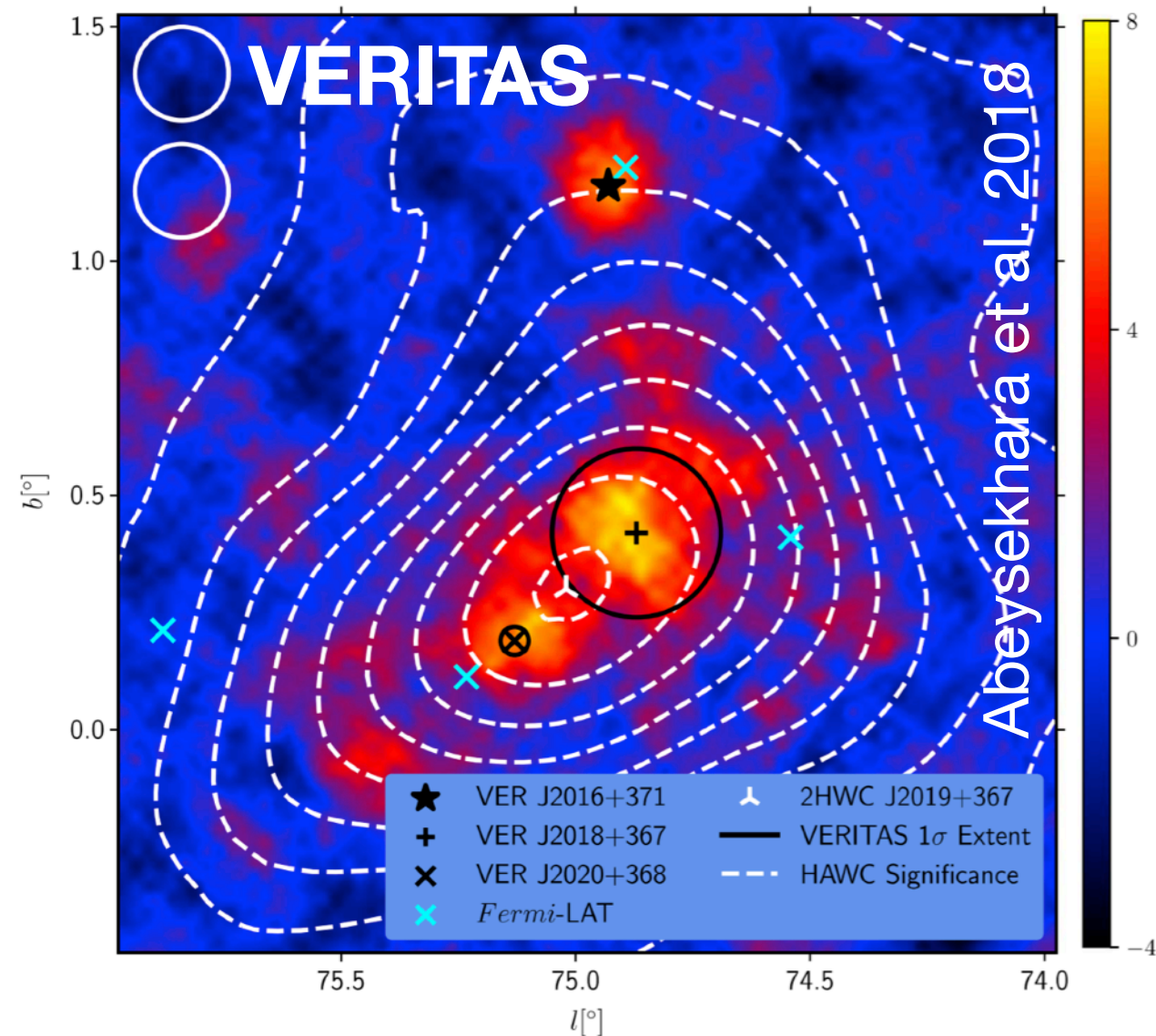
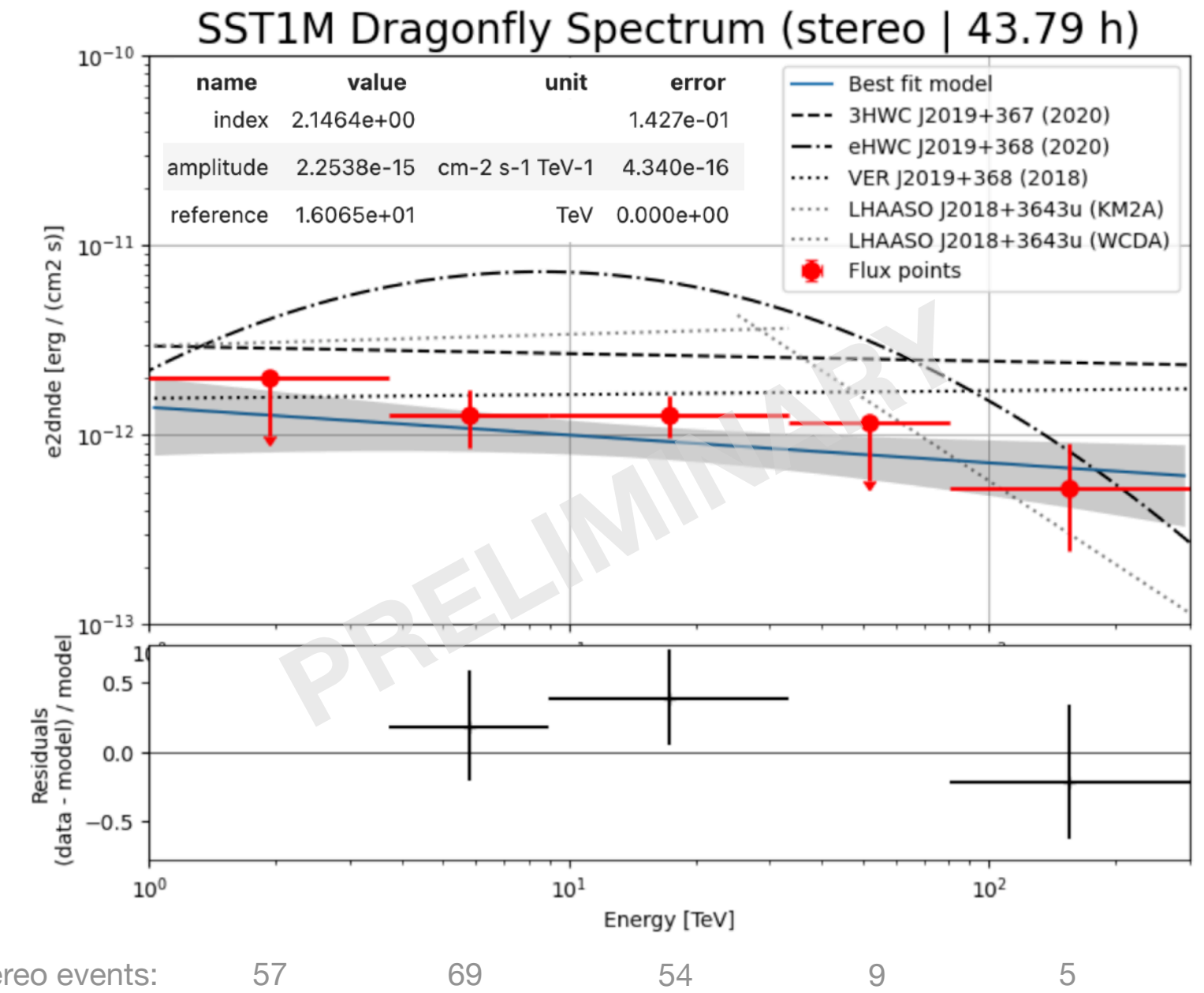


VER 2019+368 (Dragonfly)

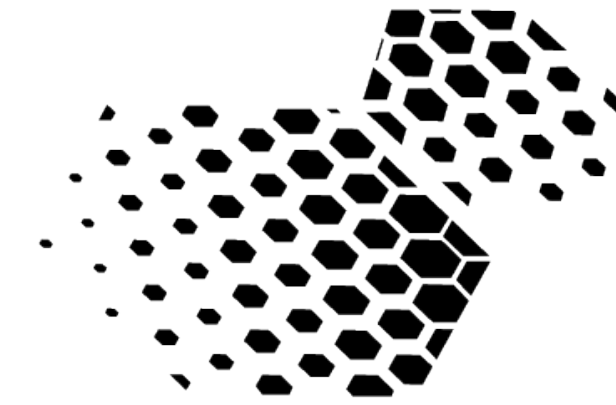
Real data

SST-1M enters the realm of extended sources

- Obs campaign in 2024, zenith angle 5-60 deg, ~44 hours of good stereo data after quality cuts
- **Preliminary results:**
 - Source position and integration region for SED fixed at VERITAS reported values (Abeysekhera et al. 2018)
 - **6sigma (p-value 2e-9)** of PL source over “no source” hypothesis (pre-trial)
 - **CTB 87 and VER J2019 regions clearly resolved**



Conclusions



SST-1M
Single-Mirror
Small Size Telescope

- Commissioning of two SST-1M telescopes continues in Ondrejov, Czech Republic
- **The telescopes operate in stereo mode**
- The telescopes have proven to meet the expected performance
- **First astrophysical sources detected in stereo**, including extragalactic or extended sources
- SST-1M stereo observatory combines **large FoV, good angular resolution** and **low altitude**, which makes it an ideal instrument for morphological studies of extended galactic PeVatron candidates - **unique science case**