



INSPIRE onboard GRAPHIUM:
**exploring MeV gamma-ray sky
with a small satellite**

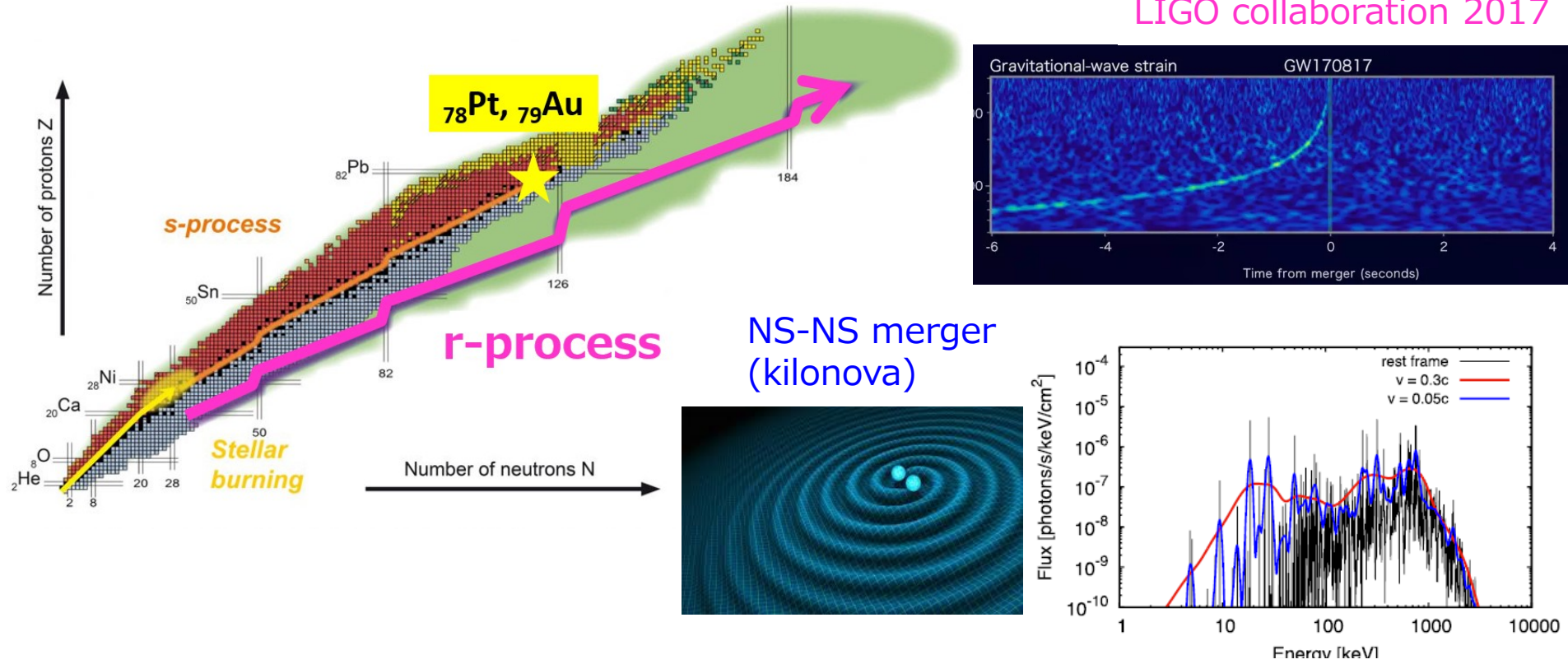
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On behalf of

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MeV γ rays as nucleosynthesis probe

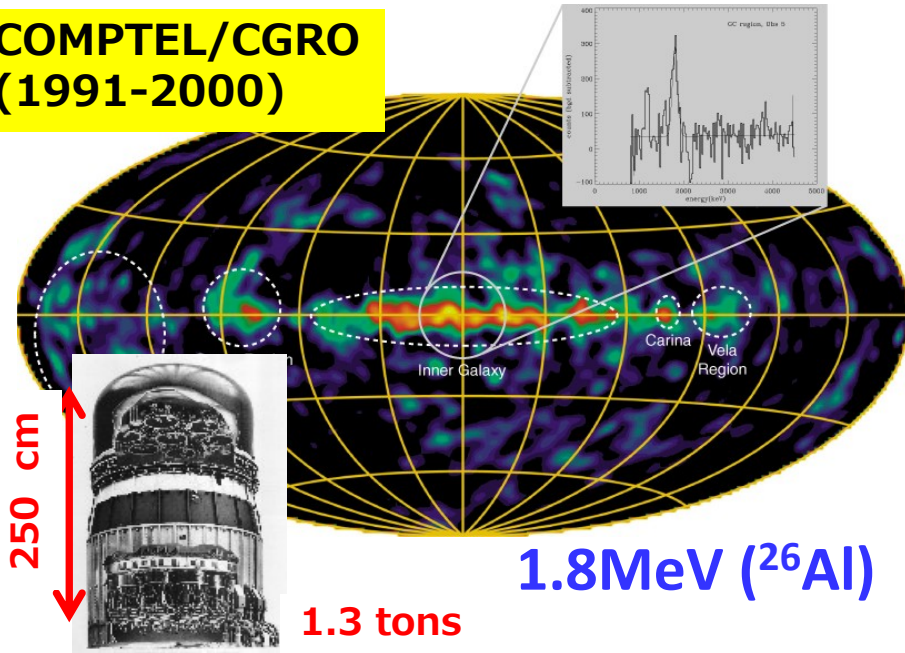
LIGO collaboration 2017



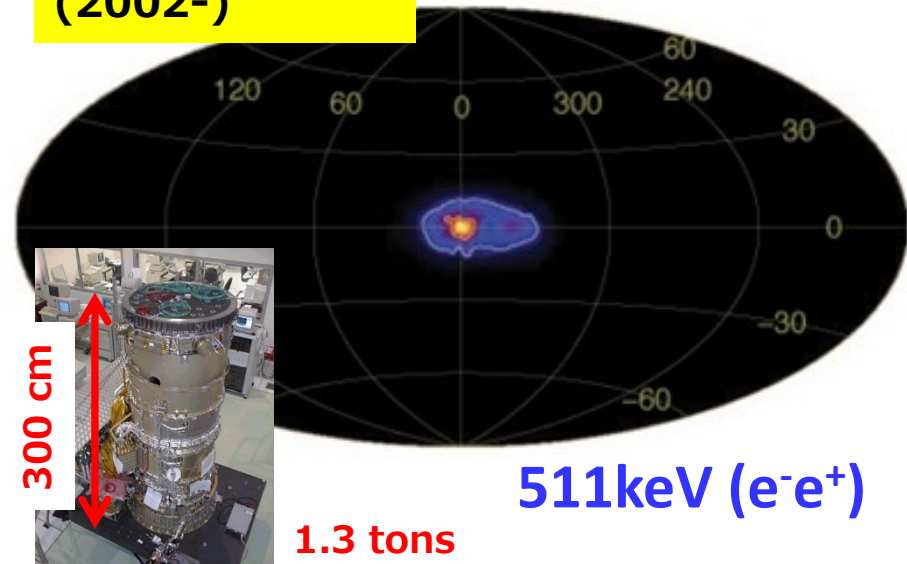
- Typical nuclear reactions: Q value ~ 0 (MeV) *Hotokezaka et al. 2016*
- Heavy elements than ^{26}Fe are thought to be produced via NC by *s*-process, but origin of Pt, Au ++ (rare metals) remains mystery
- "Kilonova" is a candidate - also emit line γ rays at 30 keV – 3 MeV ?

MeV observations in the past

COMPTEL/CGRO
(1991-2000)



SPI/INTEGRAL
(2002-)

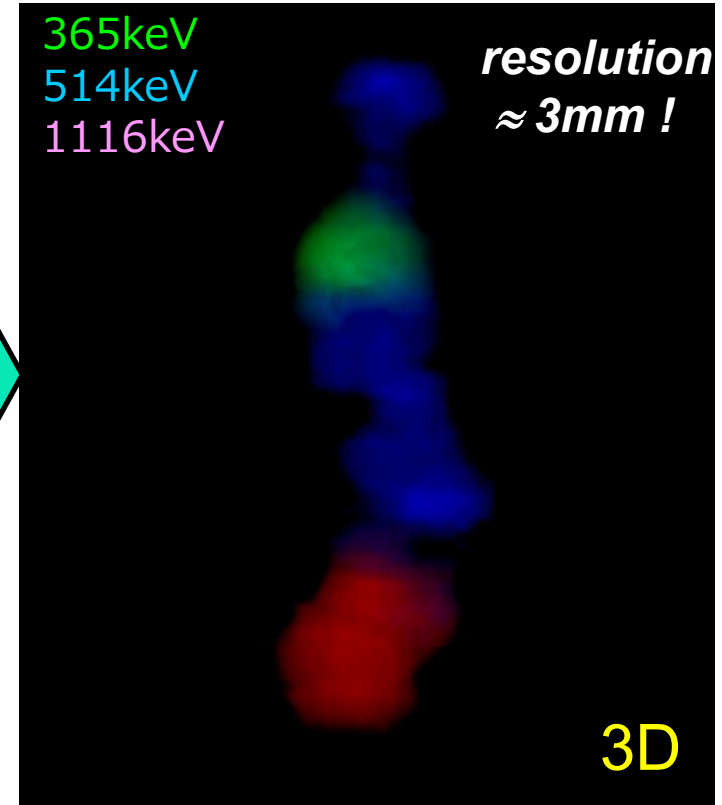
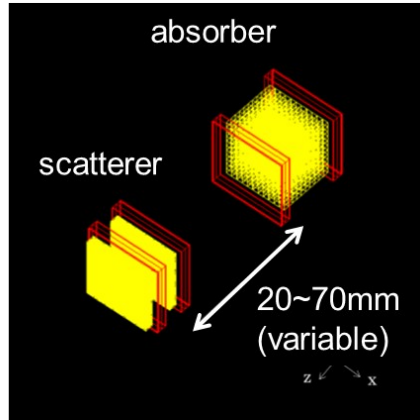
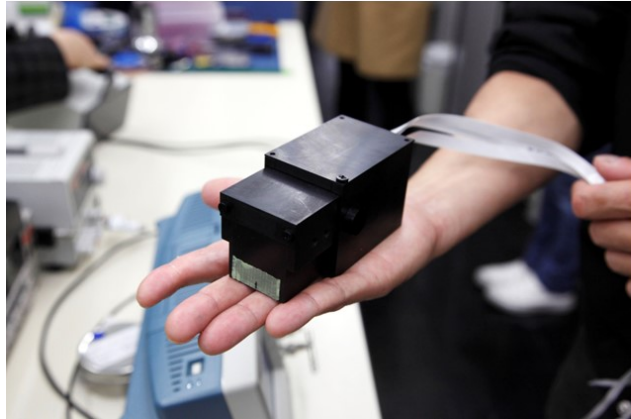


- First MeV survey by COMPTEL/CGRO resulted in many discoveries, like 511 keV emission from GC, 1.8 MeV (^{26}Al) from Gal. plane
- “Dark age” after 2000’s – many difficulties in terms of cost, man power, rockets etc

→ *What if we can do MeV observation with a 50-kg class, very small satellite?*

Challenge 1 : palm-sized CC

Kishimoto et al. 2017, *Sci. Rep.*



Palm-sized CC w/
3D position-
sensitive scintillator

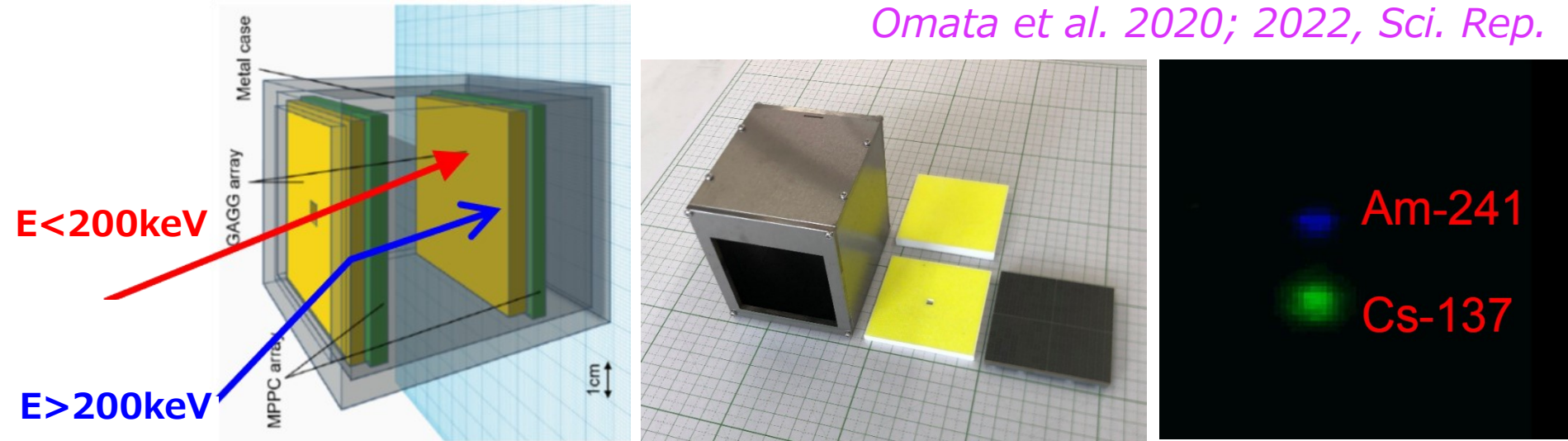


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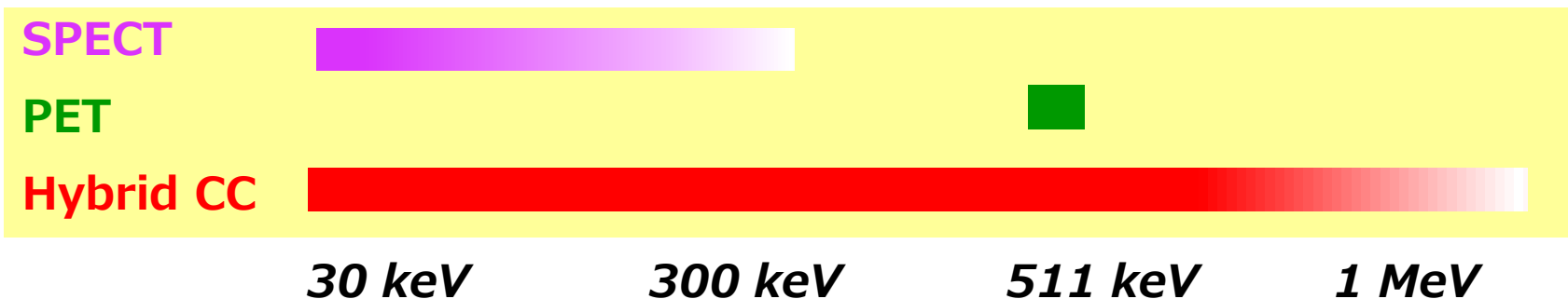
- **3D *in vivo* CC imaging of a mouse**
 - ✓ Accumulation of ^{131}I (throat), ^{85}Sr (bone), ^{65}Zn (stomach)
w/ activity 1MBq \sim 1ng/ml
- **Reduced weight and size of gamma-ray imager**

Challenge 2 : wide-band imaging

Omata et al. 2020; 2022, Sci. Rep.



- A scatterer with an “active pinhole” ($5 \times 5 \text{ mm}^2$) in the center:
 - ✓ $E < 200 \text{ keV}$ as a **pinhole camera**
 - ✓ $E > 200 \text{ keV}$ as a conventional **Compton camera**
- Extend the imageable energy range down to **30 keV!**



Challenge 3: small satellites

2021: HIBARI



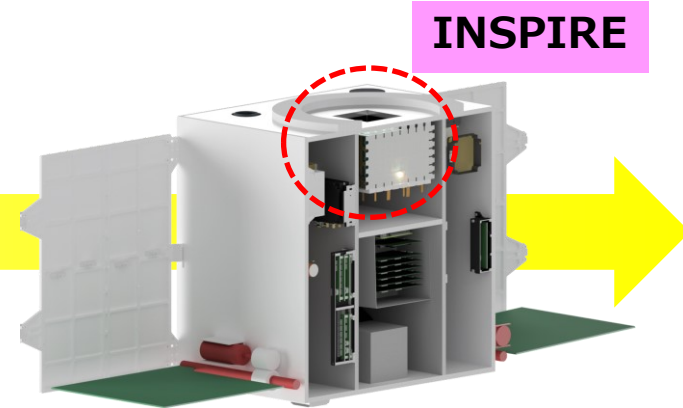
Demonstration of variable shape control

2025: Petrel



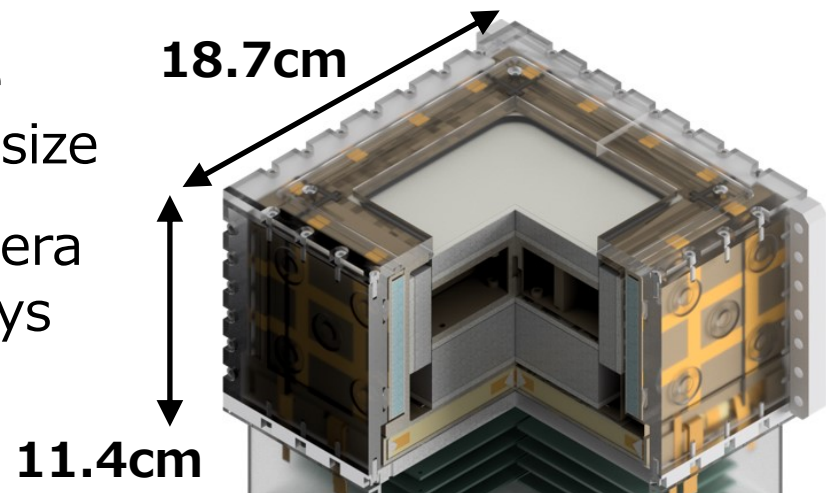
UV telescope and Earth monitor

2027: GRAPHIUM

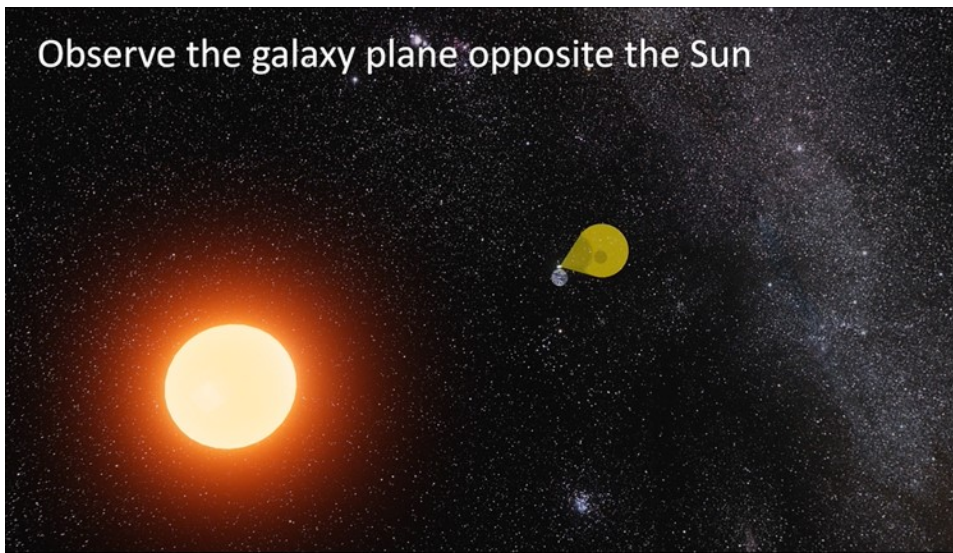
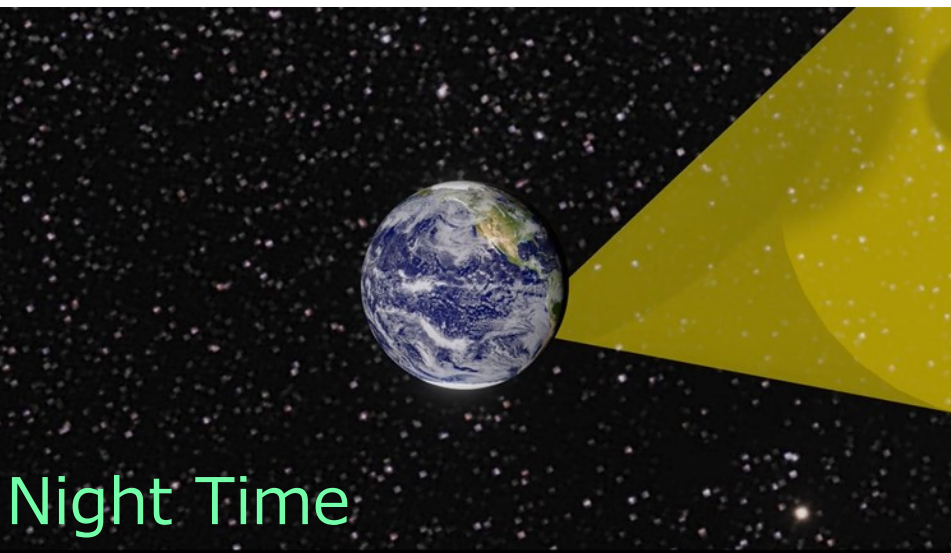
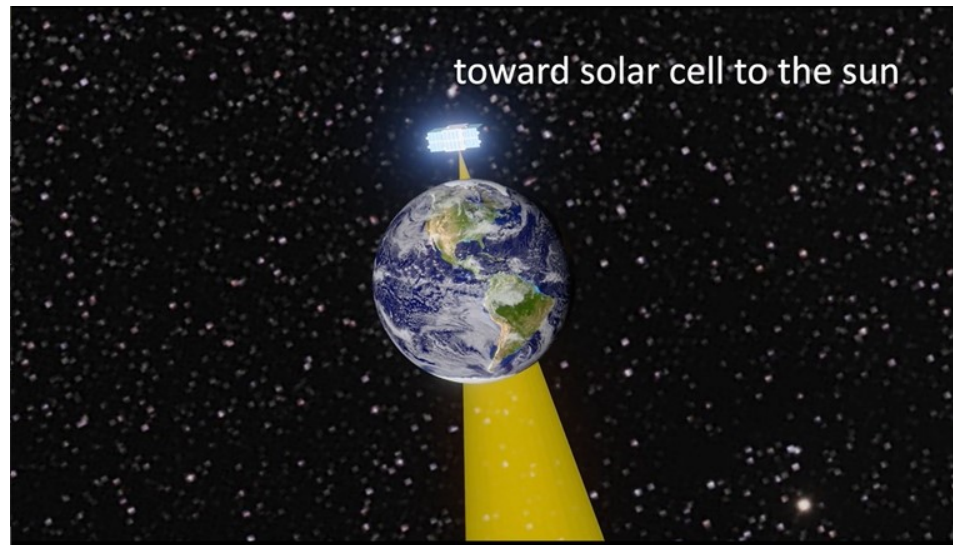
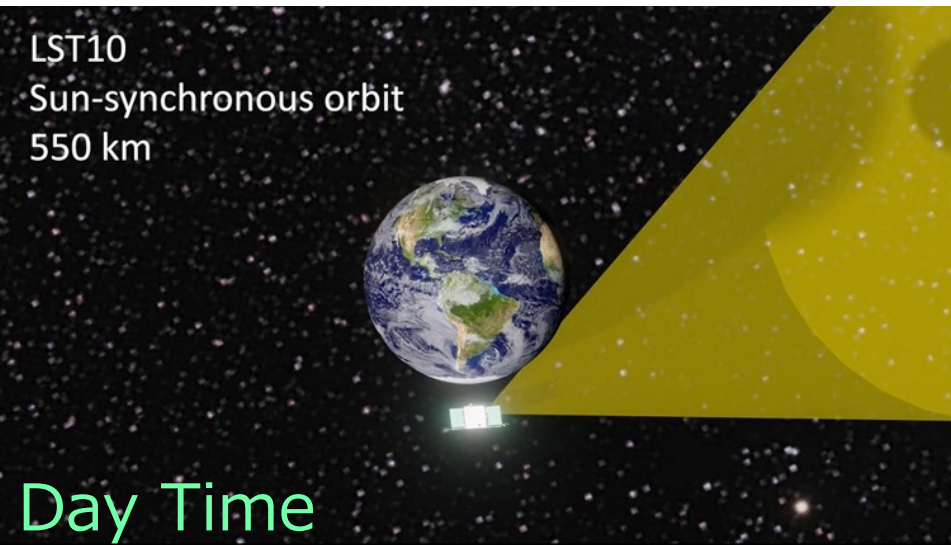


MeV all-sky survey + FF demonstration

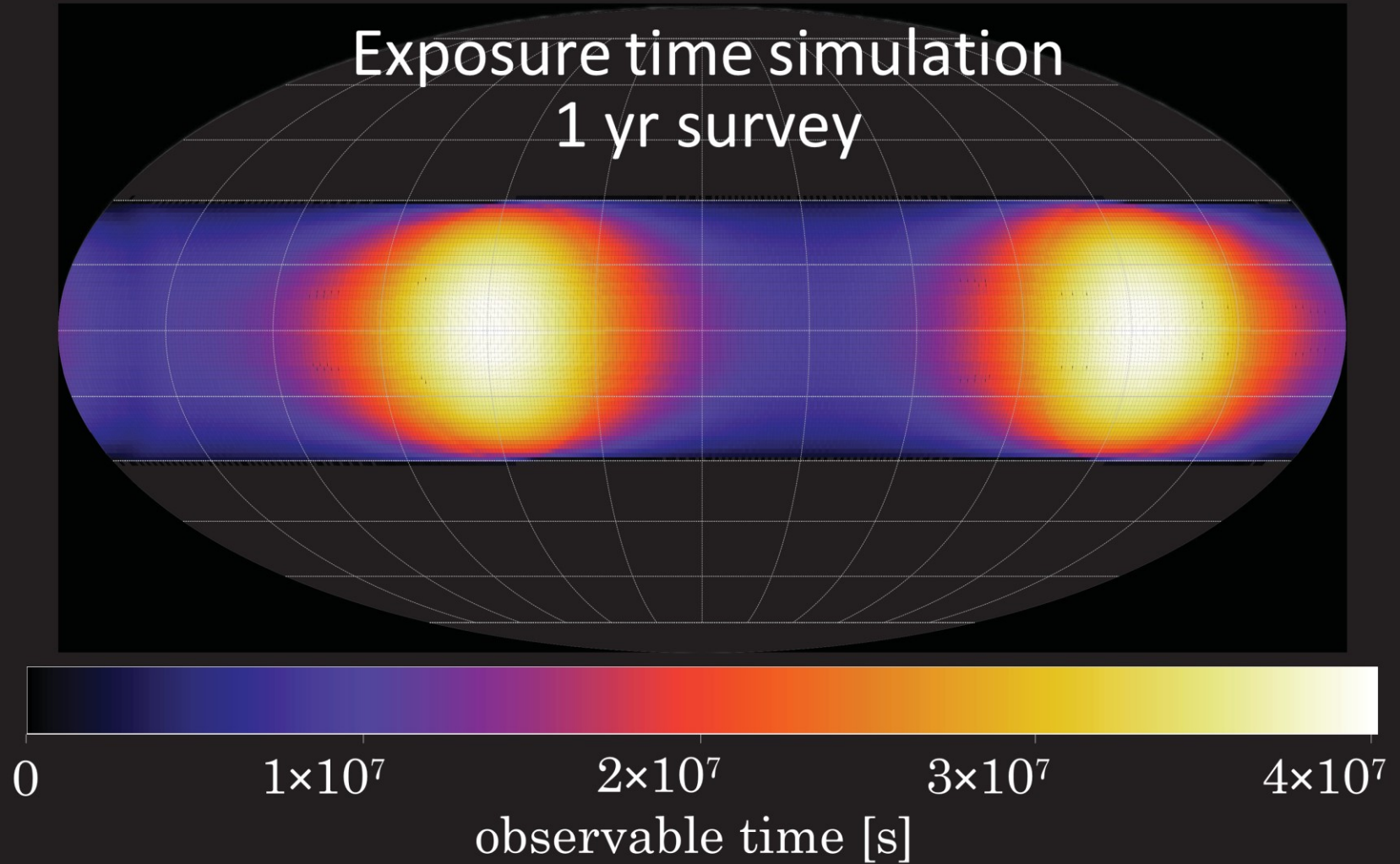
- 3rd Series of Tokyo Tech satellite w/ **75kg and 50x50x50 cm** in size
- **INSPIRE**: Hybrid Compton camera to monitor **30 keV – 3 MeV** γ rays w/ a single detector



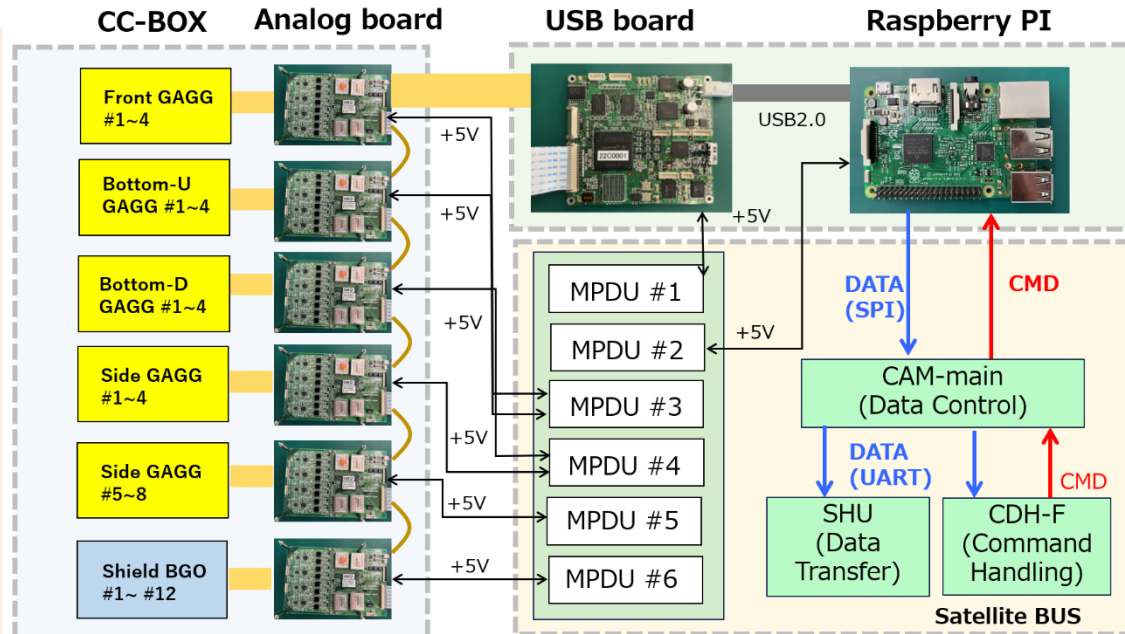
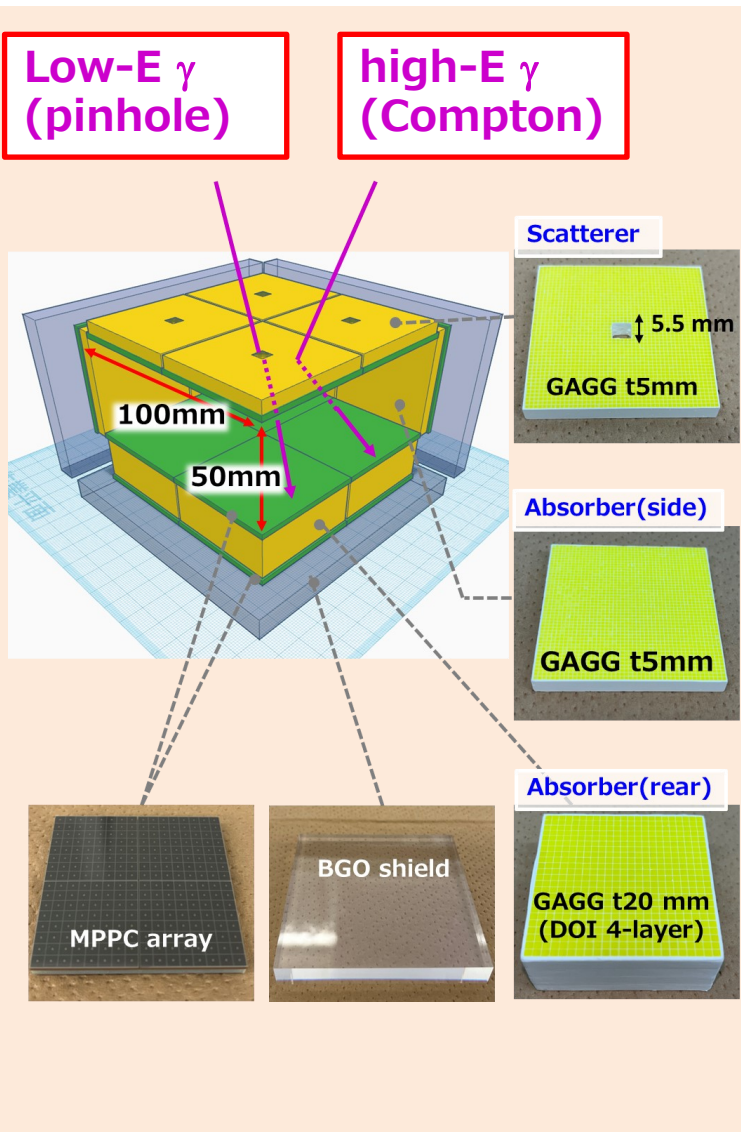
Observation Simulation



Exposure Map

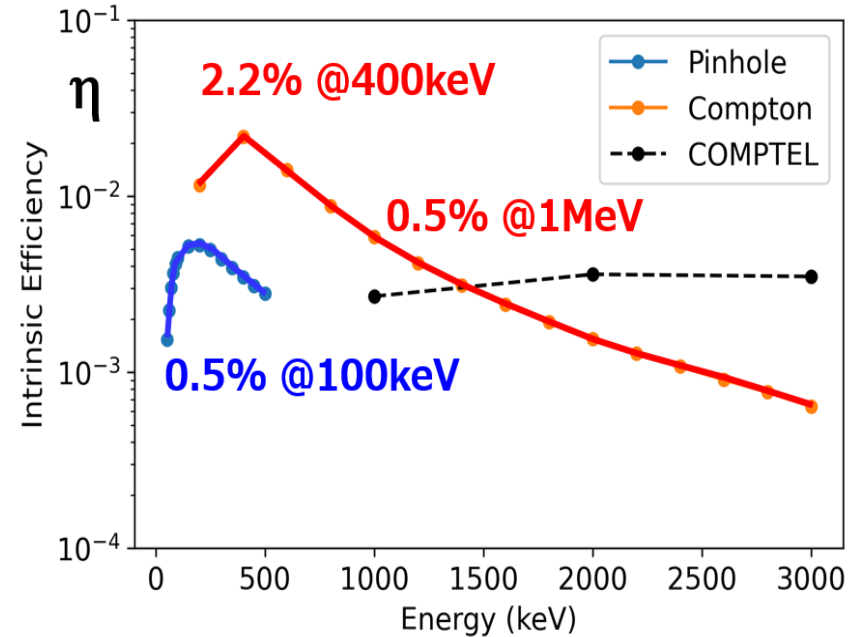
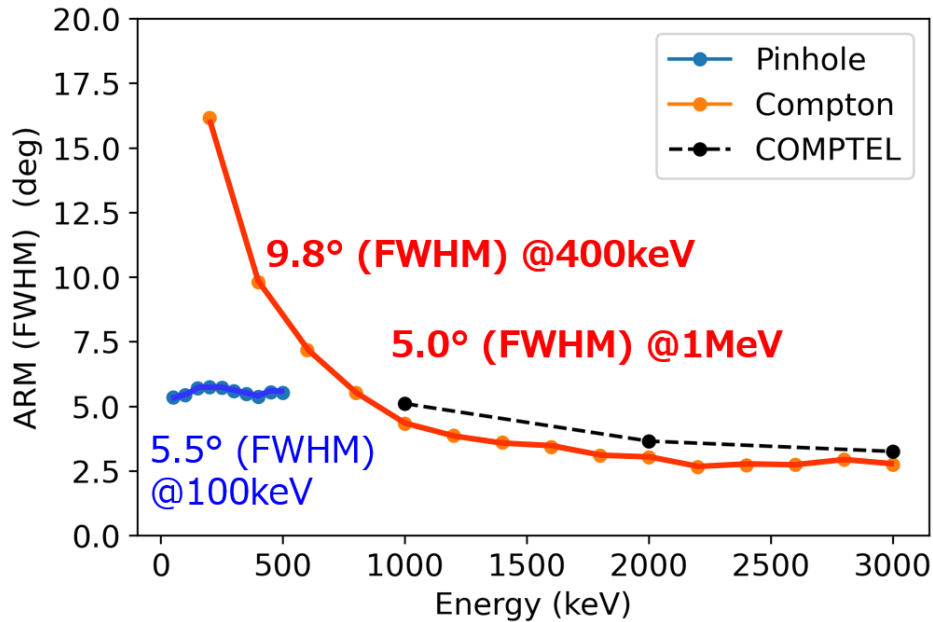


INSPIRE: system configuration

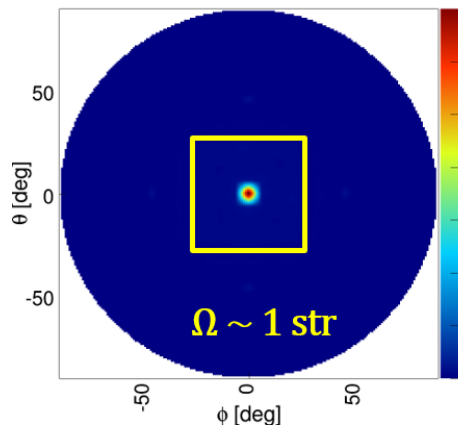


- 3D position sensitive Ce:GAGG array + 16x16ch MPPC array
- Simultaneous X and γ -ray imaging:
Pinhole (30-200 keV)
+ CC box (150-3,000 keV)
- Compact DAQ system :
 total power \sim 18W, weight \sim 10kg

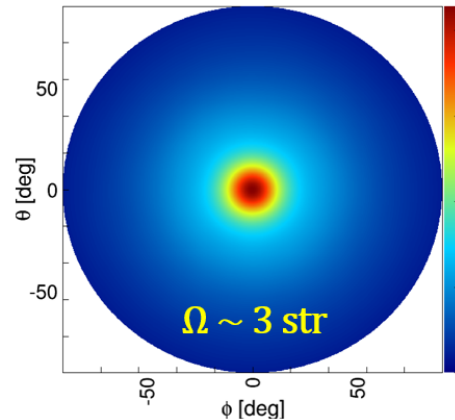
$\Delta\theta$ (ang. res.) and η (int. eff.)



Pinhole 100keV



Compton 400keV

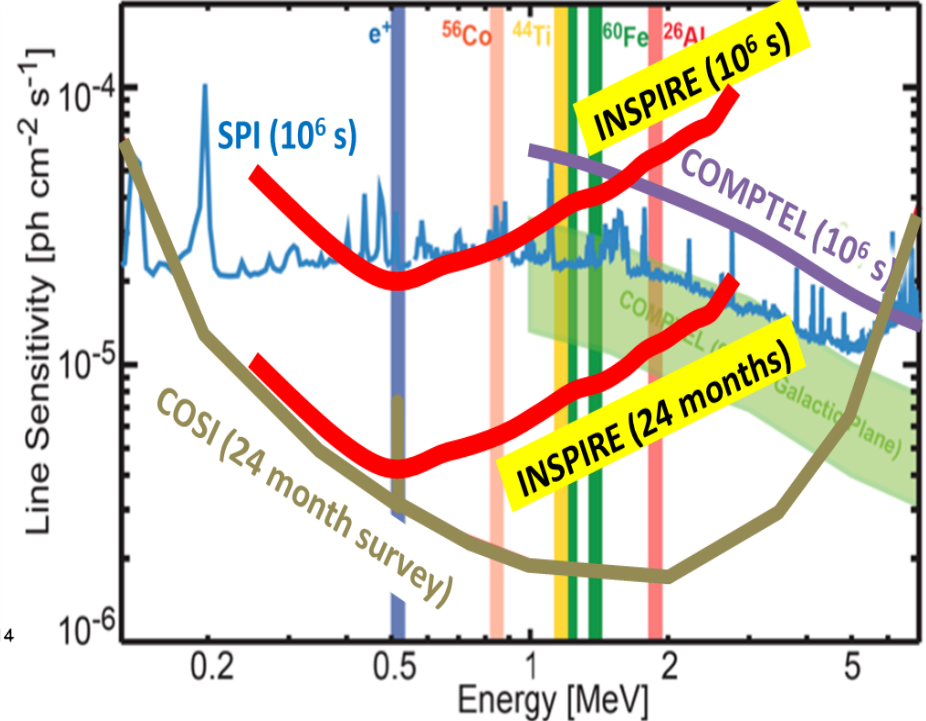
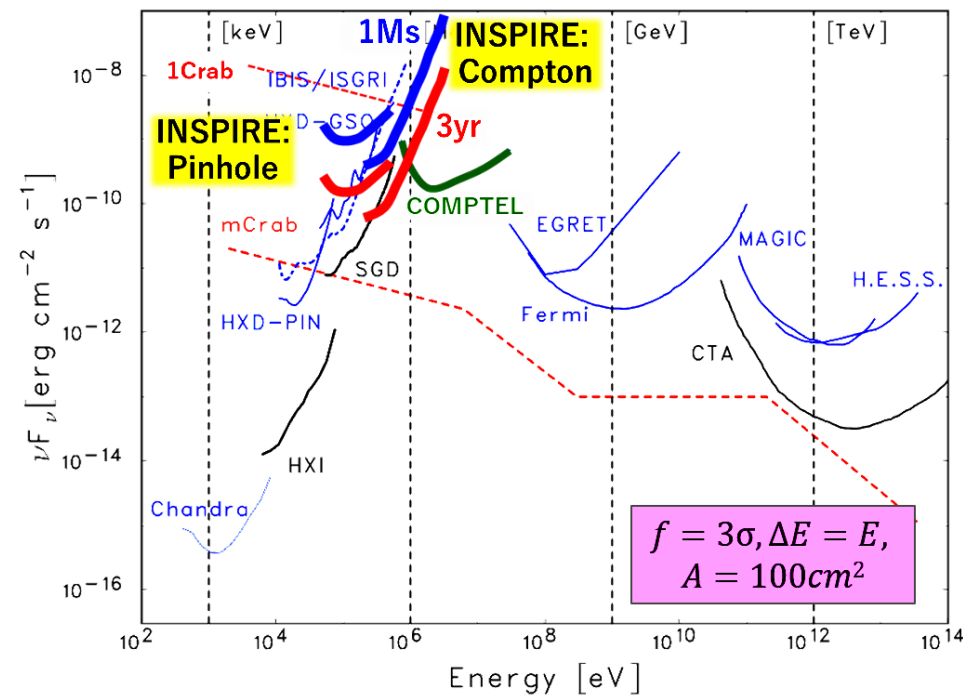


- 3D position sensitive GAGG
- ✓ Typical angular resolution $\Delta\theta \sim 5$ deg (FWHM)
- ✓ Wide FOV w/ $\eta > 0.5\%$
 $d\Omega \sim 1$ str (pinhole)
 ~ 3 str (Compton)

Continuum & line Sensitivity

Takahashi et al. 2012

Tomsick et al. 2021



- **For continuum obs.**

- ✓ Possible targets: Crab, Cyg X-1, Cen-A, NGC4151, AGN flare

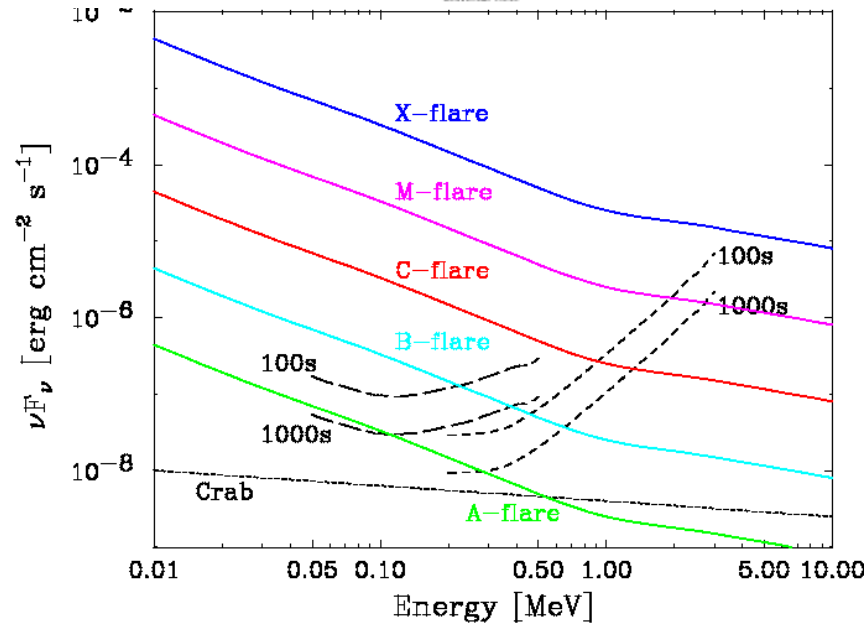
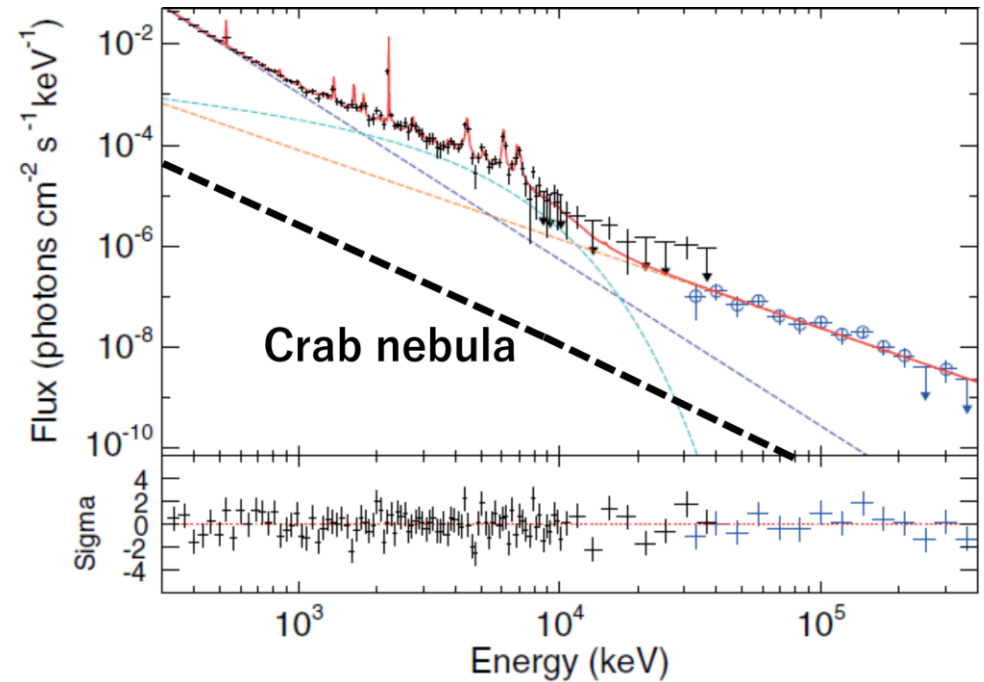
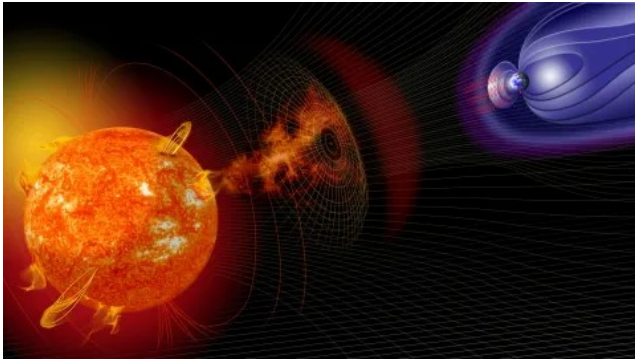
- **For narrow line obs.**

- ✓ Line sensitivity is almost equal to that of COMPTEL at 1.5 MeV, but can extend sensitivity well **below 1 MeV**

- ✓ $E < 0.5$ MeV sensitivity **almost comparable w/ COSI**

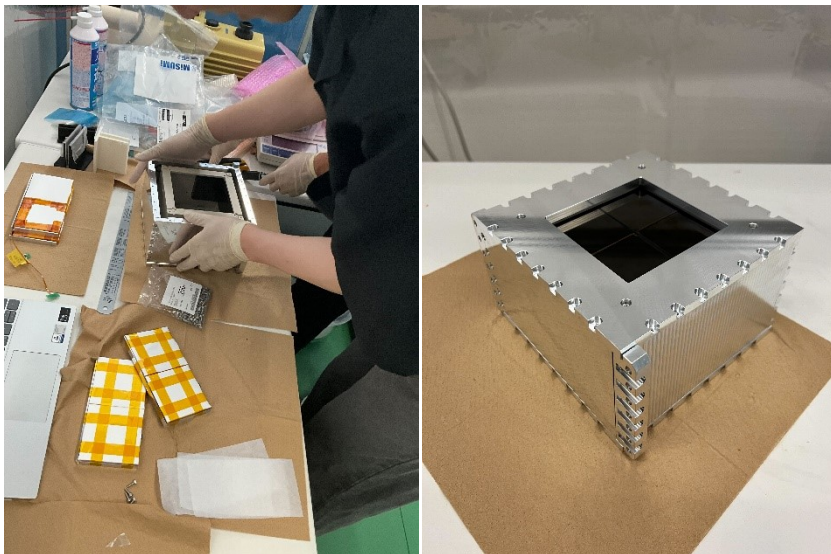
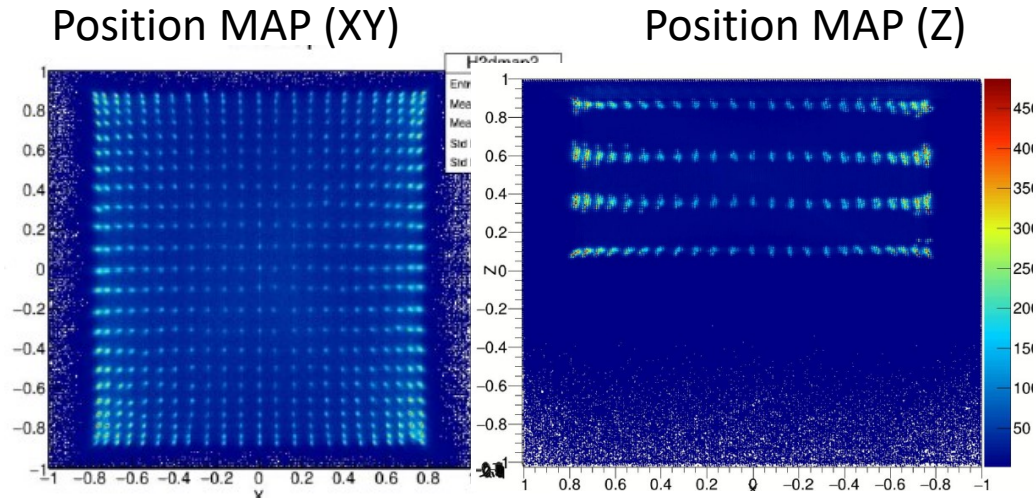
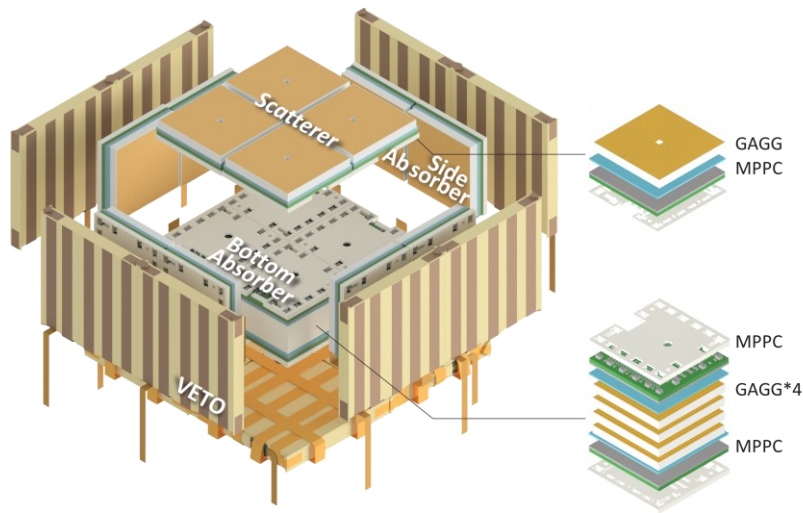
Observation of Solar flare

Ackermann et al. 2012



- Very bright and frequent :
 - ✓ $\sim 5/\text{day}$ for C-class flare
 - ✓ $\sim 0.7/\text{day}$ for M-class flare
- Various de-excitation lines w/ along with non-thermal bremsss:
 - e^-e^+ (511keV), ^{56}Fe , ^{24}Mg (0.5-2MeV)
 - NC (2.2MeV), ^{12}C (4.4MeV)++

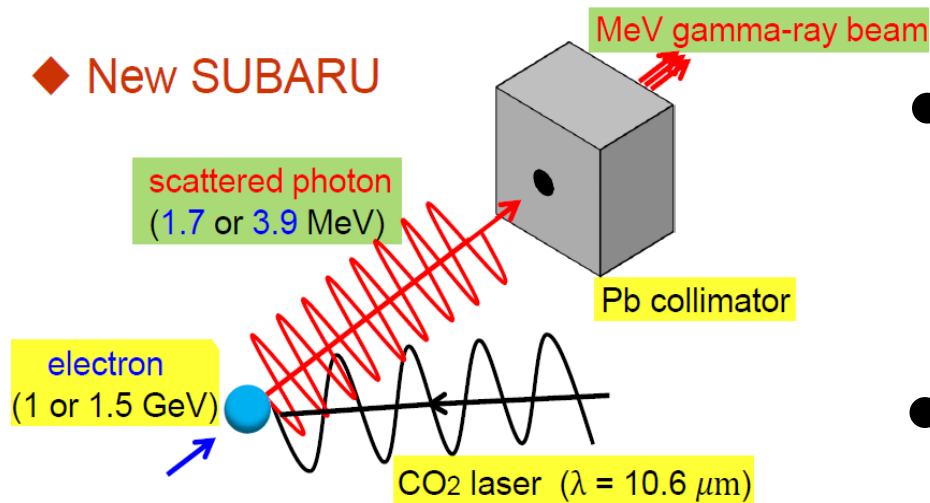
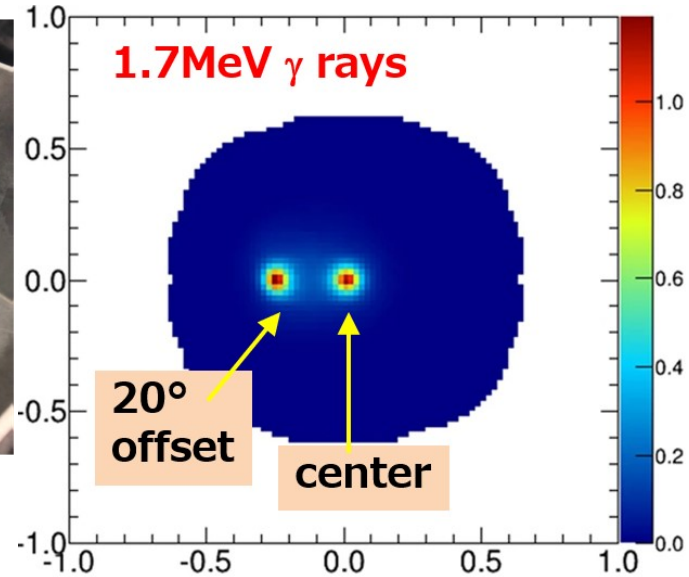
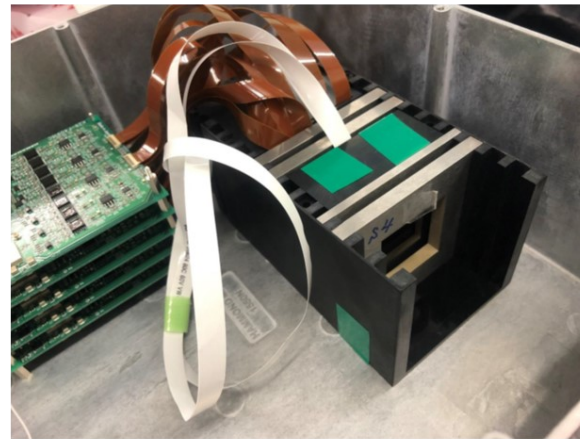
EM fabrication and testing



- $\frac{1}{4}$ scale of full-FM sensor (1 unit of hybrid CC implemented)
- $\Delta E/E \sim 5.5\%$, $\Delta\theta \sim 5^\circ$ (FWHM) @1MeV as expected
- Confirmed detector performance between -40°C and $+85^\circ\text{C}$ w/ thermal vacuum testing

MeV gamma imaging

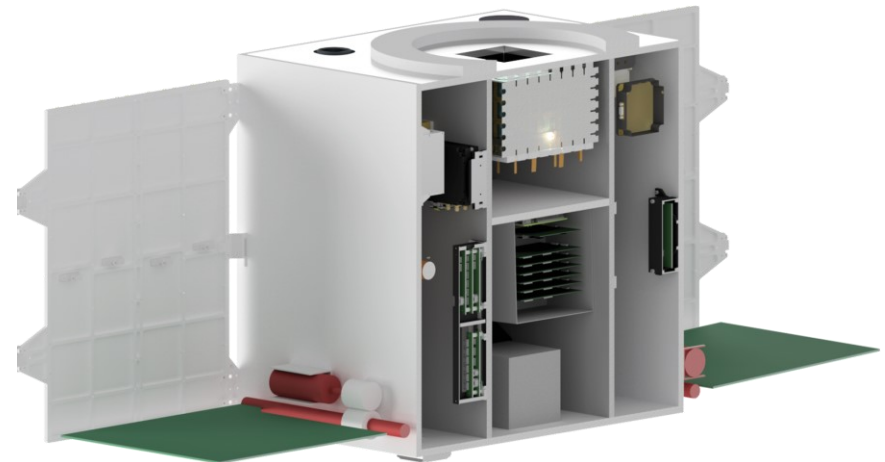
Hosokoshi et al. 2019, Nature Sci. Rep.



- Experiment was performed at New SUBARU, which provides **monochromatic γ -ray beam** by IC of GeV electrons from SP8
- Measured angular resolutions:
 $3.4 \pm 1.1^\circ$ (FWHM) @ 1.7 MeV
 $4.0 \pm 0.5^\circ$ (FWHM) @ 3.9 MeV

Summary

- We are developing “**GRAPHIUM**” as a new challenge of **small satellites for frontiers in space science**
- Despite limited resources (i.e., weight, size, budget :-), GRAPHIUM provides important contribution to **MeV astronomy**, that is stagnated over 30 years
- Also **a compact MeV camera** is highly useful in various other fields, such as **nuclear medicine and atmospheric science**



Kataoka et al. 2024, NIM-A