



cherenkov
telescope
array



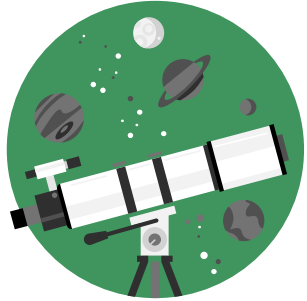
TOR VERGATA
UNIVERSITÀ DEGLI STUDI DI ROMA

INSIGHTS INTO THE PHYSICS OF VERY HIGH ENERGY TRANSIENT EVENTS WITH IMAGING ATMOSPHERIC CHERENKOV TELESCOPES

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INAF / OA Roma

1st VHEGAM Meeting, 15-17 January 2024



WHAT WE DO KNOW

- **Transient** events: $10^{-3}s$ (sGRB) – 10^2s (lGRB).
- **Isotropic** sky distribution.
- $E \sim 10^{50} - 10^{54}$ erg.
- **High-redshift** sources.
- **Prompt** emission (hard-X/soft- γ) and **Afterglow** (X/radio).
- **VHE** emitters (some of them).
- Observed link with **CCSNe** Ib/Ic events and with **GWs**.



WHAT WE DO NOT KNOW

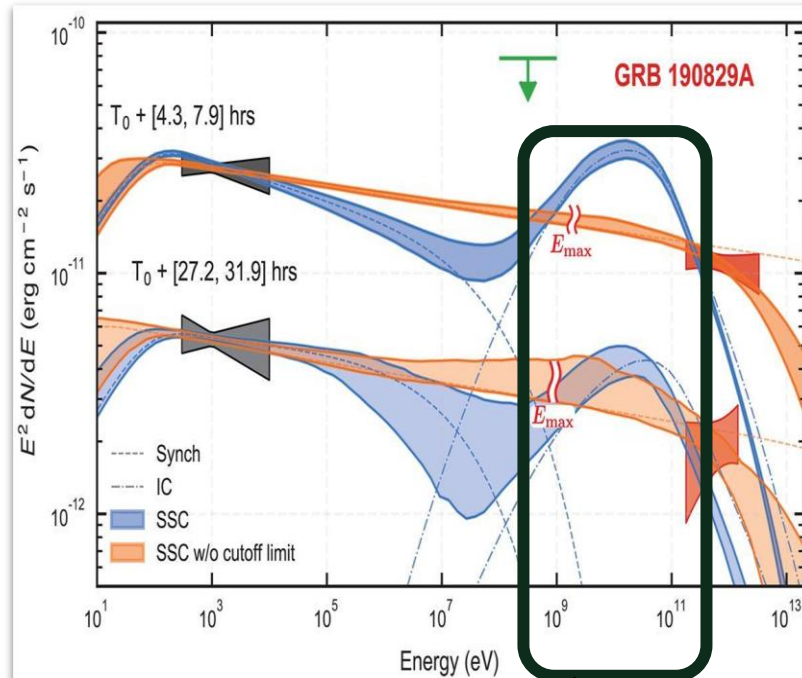
- Which is the emission **model** at work?
- Which are the emission **mechanisms**?
- Do all GRBs have a **VHE component**?
- How GRBs are linked with **GWs events**?
- How to explain the **connection** between GRBs and CCSNe?



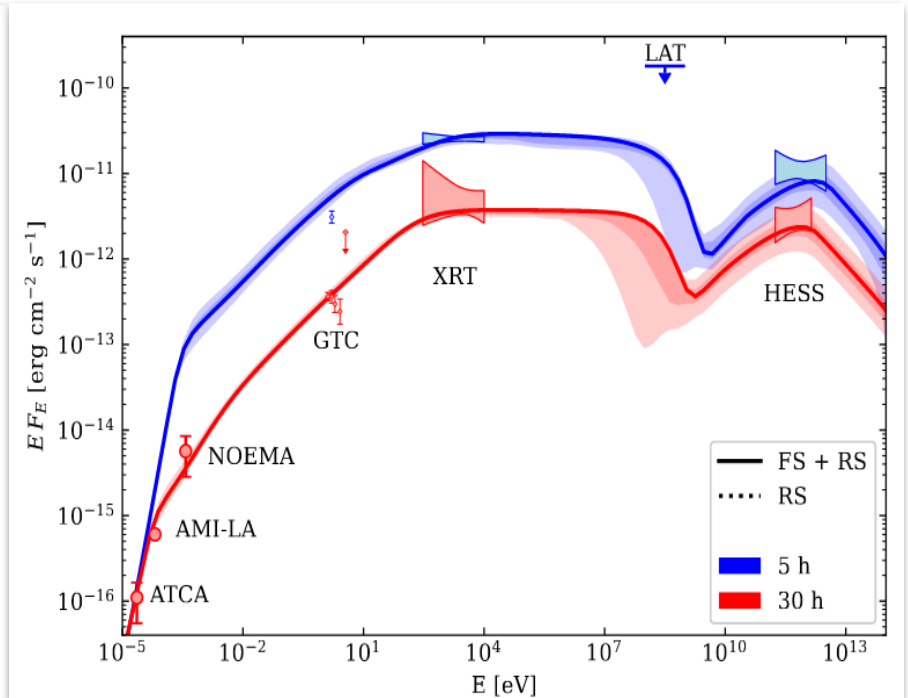
- Which are the emission mechanisms?

An example: Synch vs SSC models for GRB 190829A

HESS Coll. 2021, Science, 372, 6546



Salafia et al. 2022, The Astrophysical Journal Letters, 931.2, L19

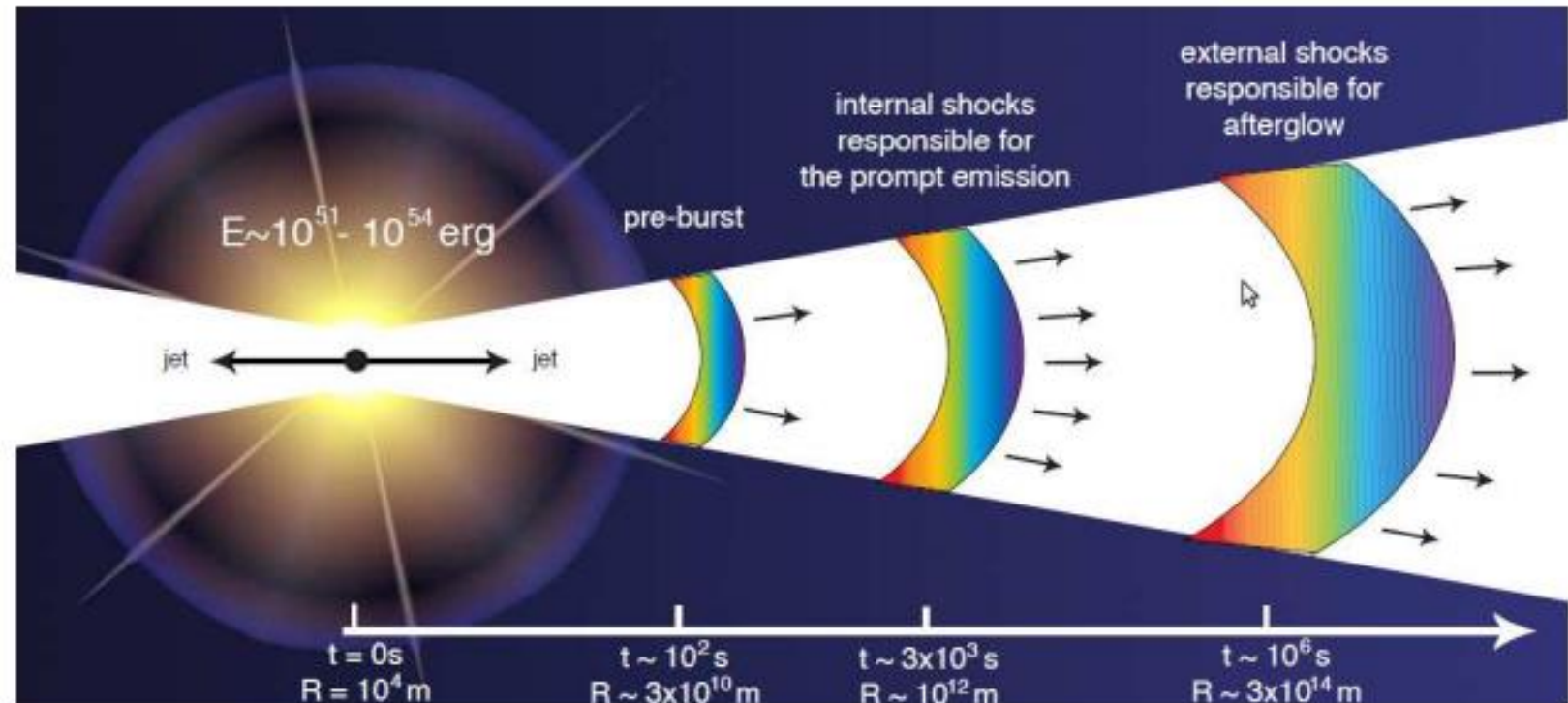


Crucial region to discriminate between models we expect CTA/LST to explore.



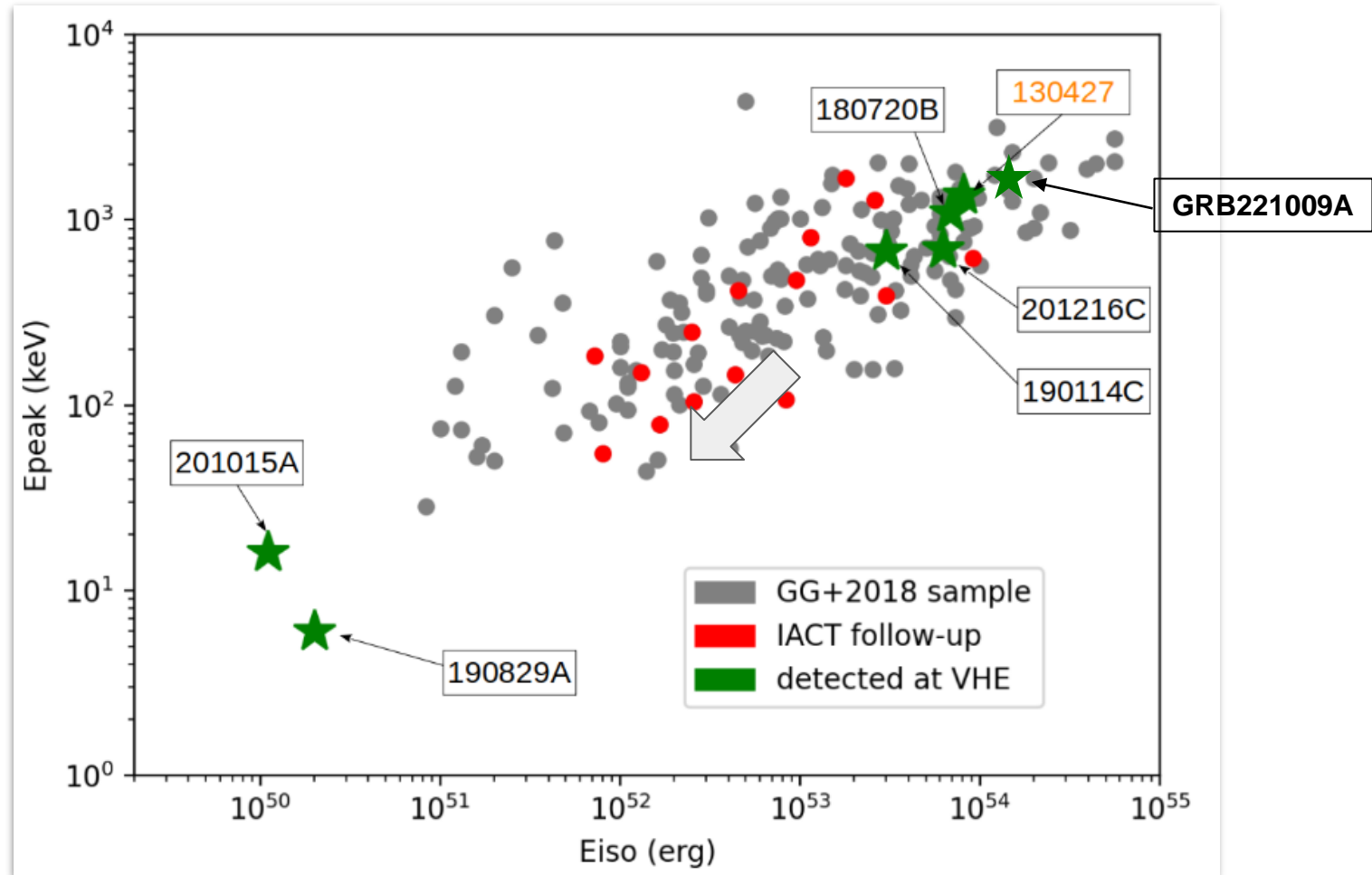
An example: the fireball model

- Which is the emission model at work (internal/external shock model or magnetic reconnections)?



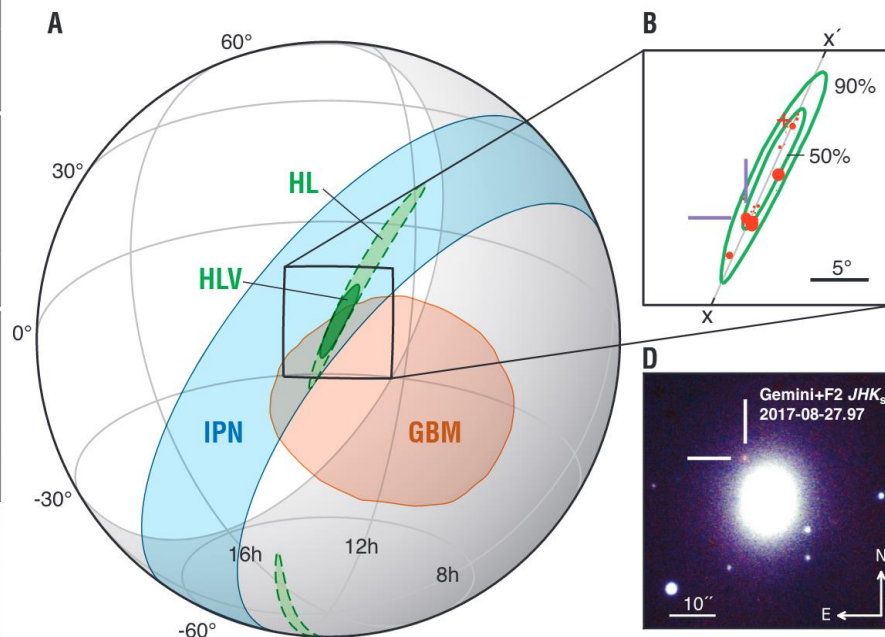
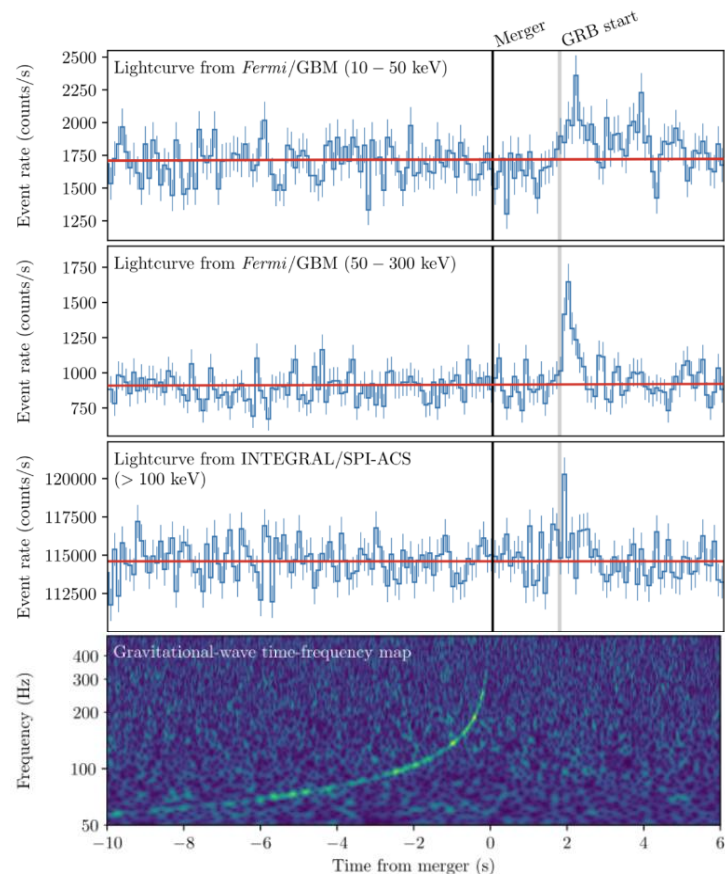


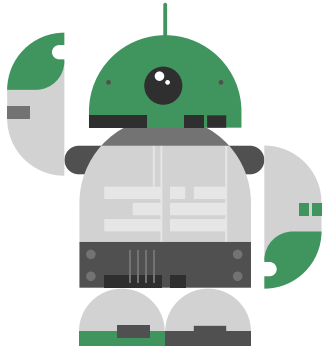
- Do all GRBs have a VHE component?





- **How GRBs are linked with GW events (off-axis GRB)?**





WHAT WE DO

- Use **broadband** observational **data**, including the VHE data from **MAGIC and CTA/LST1** (follow up campaigns, data analysis).
- Investigate different **emission scenarios** and **physical processes** (discrimination of leptonic/hadronic emission in prompt and afterglow emission).
- Investigate **the link between GRBs and GWs** (multimessenger event/optical counterpart from GRB /follow up of event of interest).

THANK YOU

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