# UFOs in AGNs with EmulatoRs: UAGNER

(Linkeed to the scheda UAGNER – Mauro Dadina )

**Goal**: test and validate, both on technical and physical basis on real CCD data, machine learning based emulators (XRADE but not only) of radiation driven disk winds (DW) in AGN.

How: X-ray data analysis of different kinds of AGN in the local Universe  $\rightarrow$ 

- **1)** Testing of the basic DW model (using DW spectral tables);
- 2) testing the emulator XRADE (Matzeu et al. 2022);
- **3) (added!)** testing different scenarios (magneto-hydrodynamic models, MHD, Fukumura et al. 2010, 2015; Fukumura, MD et al. 2022)

Members: Mauro Dadina (INAF/OAS), Gabriele Matzeu (DIFA UniBO and INAF/OAS... now ESA)

## Proposed timetable and status:

1) within end of June 2022 (+ ~10/12 months due to funding availability/programm starting in February 2023): definition of the sample of 2/3 Seyfert and radio-loud galaxies; Done

*i) IRAS F11119+3257* → ULIRG, UFO detected with NuSTAR (Tombesi et al. 2017)

Paper almost submitted and based on new XMM-Newton+NuStar data (Lanzuisi et al. in prep): Tested the radiation driven wind (DW) model table at the base of XRADE. To be done: test XRADE, WINE (Luminari et al. 2021) and MHD models (Fukumura et al. 2010, Fukumura, MD, et al. 2022).



Lanzuisi et al. to be sumitted: Confidence contour of the incl. angle vs. outflow velocity (left panel) and of the Edd. Ratio vs mass outflow rate as obtained for IRAS F11119+3257 using the disk wind Model (Sim et al. 2008, 2010; Matzeu et al. 2022)

*ii)* 3c 120 → BLRG, UFO detected using Suzaku data (Gofford et al. 2013)

Paper in preparation using XMM-Newton+NuSTAR data (Dadina et al. in prep): Tested the radiation driven wind model table at the base of XRADE: to be done: test XRADE and MHD model before submission (this work will be done in close contact with Gabriele Matzeu and Keigo Fukumura)

*iii)*  $MCG-03-58-007 \rightarrow Seyfert 2 galaxy, UFO detected using XMM-Newton data (Matzeu et al. 2019). DW model already tested in Braito et al. 2022 (not within the context of this project). To be done: test XRADE and MHD model before submission (this work will be done in close contact with Gabriele Matzeu and Keigo Fukumura)$ 

2) within end of September 2022 (+10/12 months) first results of data analysis using XRADE; Done for the testing of the DW model (see above). To be done: testing XRADE

3) within end of November 2022 (+10/12 months) simulations of the obtained spectral set-up with the Athena/X-IFU and XRISM-Resolve response matrices; delays in the production of high-resolution tables for the DW model due to problems in purchasing high performance computer (costs increases and time). Some available highresolution table already used to evaluate the newAthena capabilities in coordination with the Athena Science Redefinition Team of ESA.

4) within summer/winter 2023 (+10/12 months) publication of the first results. Some publications almost ready

#### **Forseen path:**

**1)** Within the end of spring 2024: submission of the papers on IRAS F11119 and 3c120 using DW models

2) Within the summer/autumn 2024: submission of the papers on MCG-03 based also on MHD models

3) Within the end 2024: publication of the XRADE results if not done in the DW papers

**4) 2024:** preparation of the first Xrism proposal using high-resolution DW and MHD tables (to be verified the availability of high-resolution tables of the DW model, see below).

5) added... using machine learning tech. also for MHD models? (strong collaboration with Keigo Fukumura of James Madison University)

## **Critical issues:**

a) Hugely increased costs for computing power (strong delay – months- in purchasing the needed computing power. Just arrived as of November 2023)

b) Lost of collaborators – namely Gabriele Matzeu- due to other projects (Athena/XIFU) programmatic issues

## **Raw financial statement:**

- Computing: ~8500 Euro
- Travels: ~1600 Euro
- Visits to INAF/OAS: ~1600 Euro (Gabriele Matzeu, Keigo Fukumura)
- Still available for the second year: ~6000 Euro (expected to be travels)