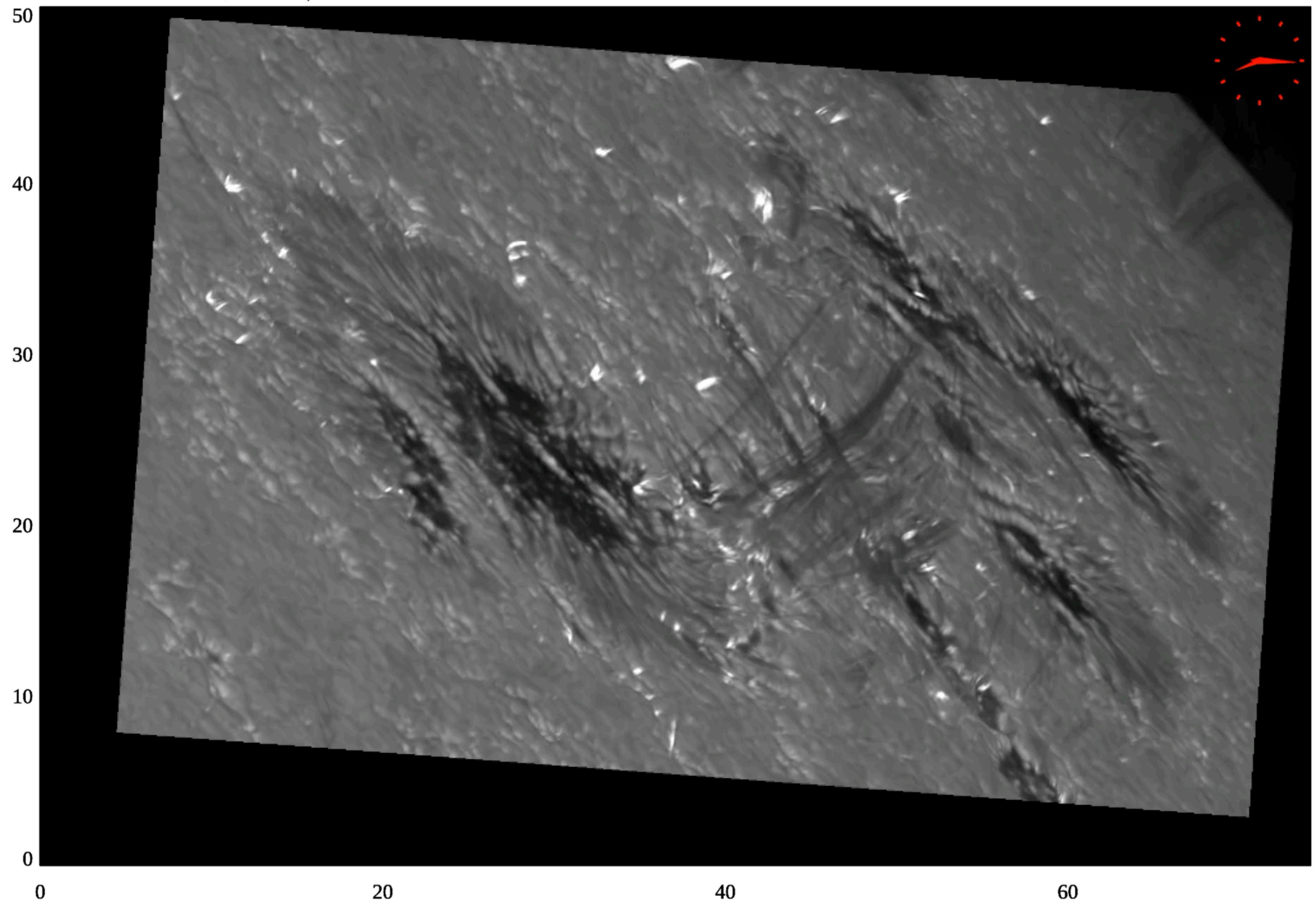


Rosseland  
Centre  
for Solar  
Physics

# Impact of small-scale photospheric magnetic reconnection events on the upper atmosphere

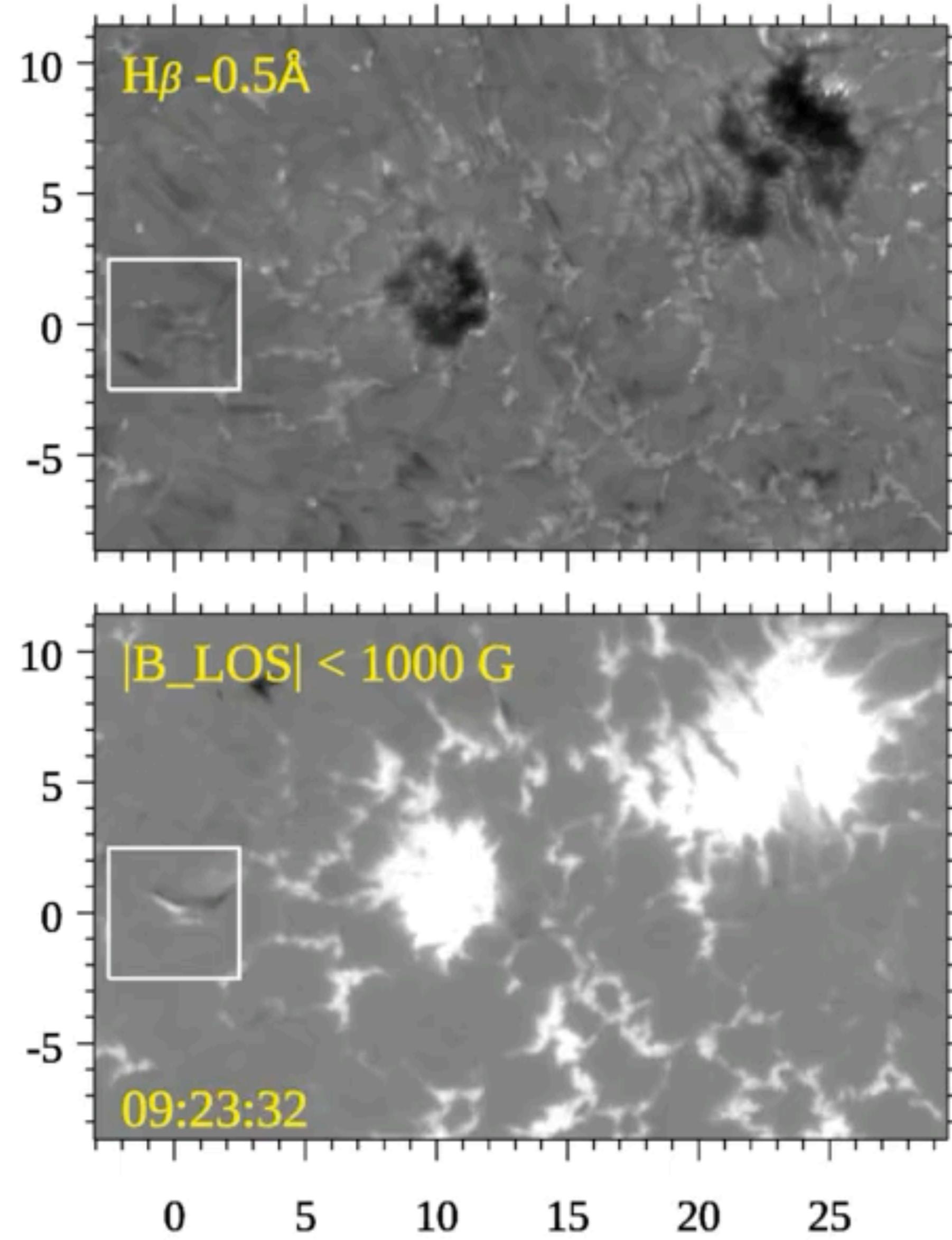
Luc Rouppe van der Voort  
Rosseland Centre for Solar Physics (RoCS), University of Oslo

In collaboration with Jayant Joshi (IIA Bengaluru), Aditi Bhatnagar (RoCS Oslo) & Mats Ola Sand (RoCS Oslo)

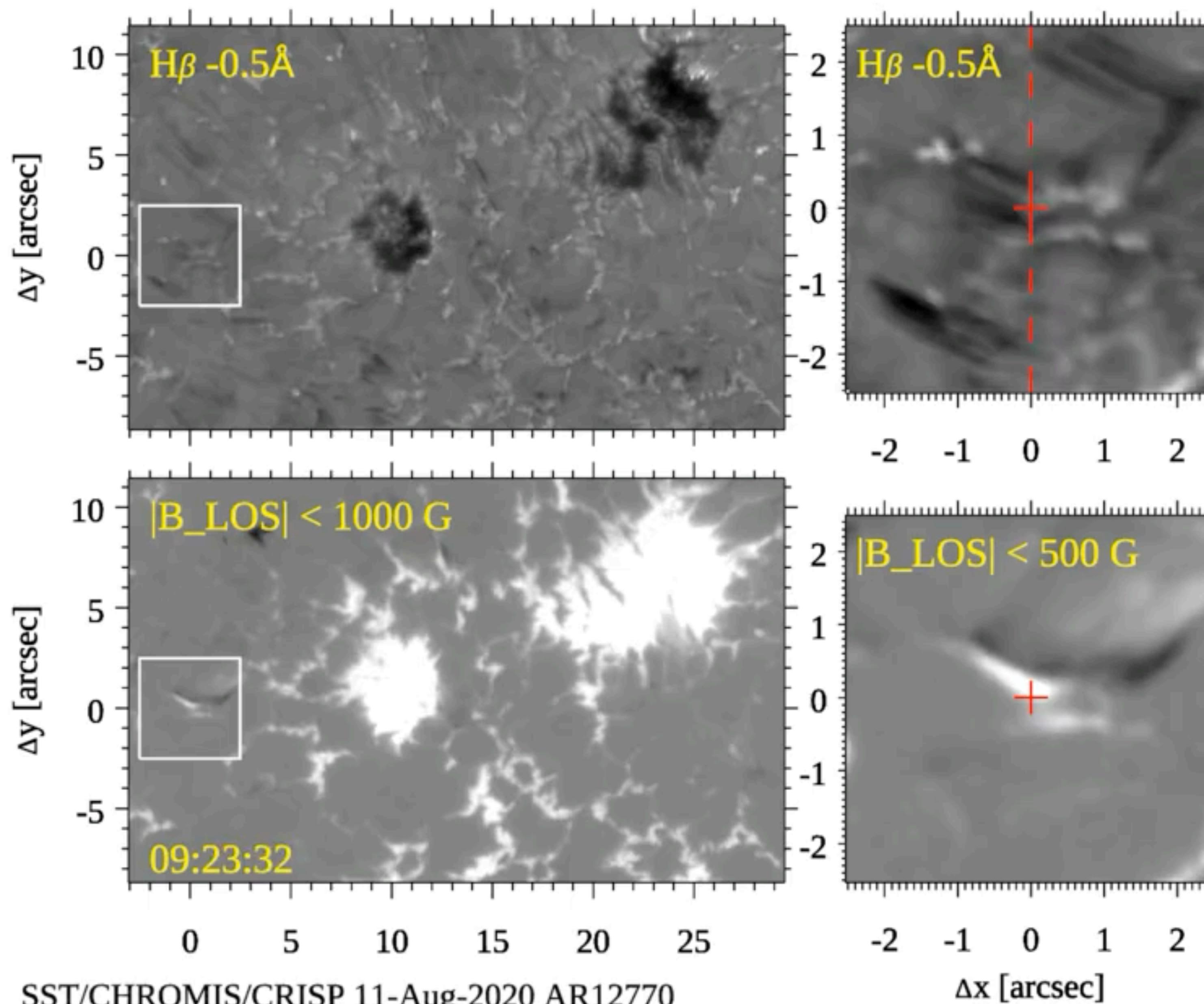


**Ellerman bombs:**  
magnetic reconnection  
in low solar atmosphere

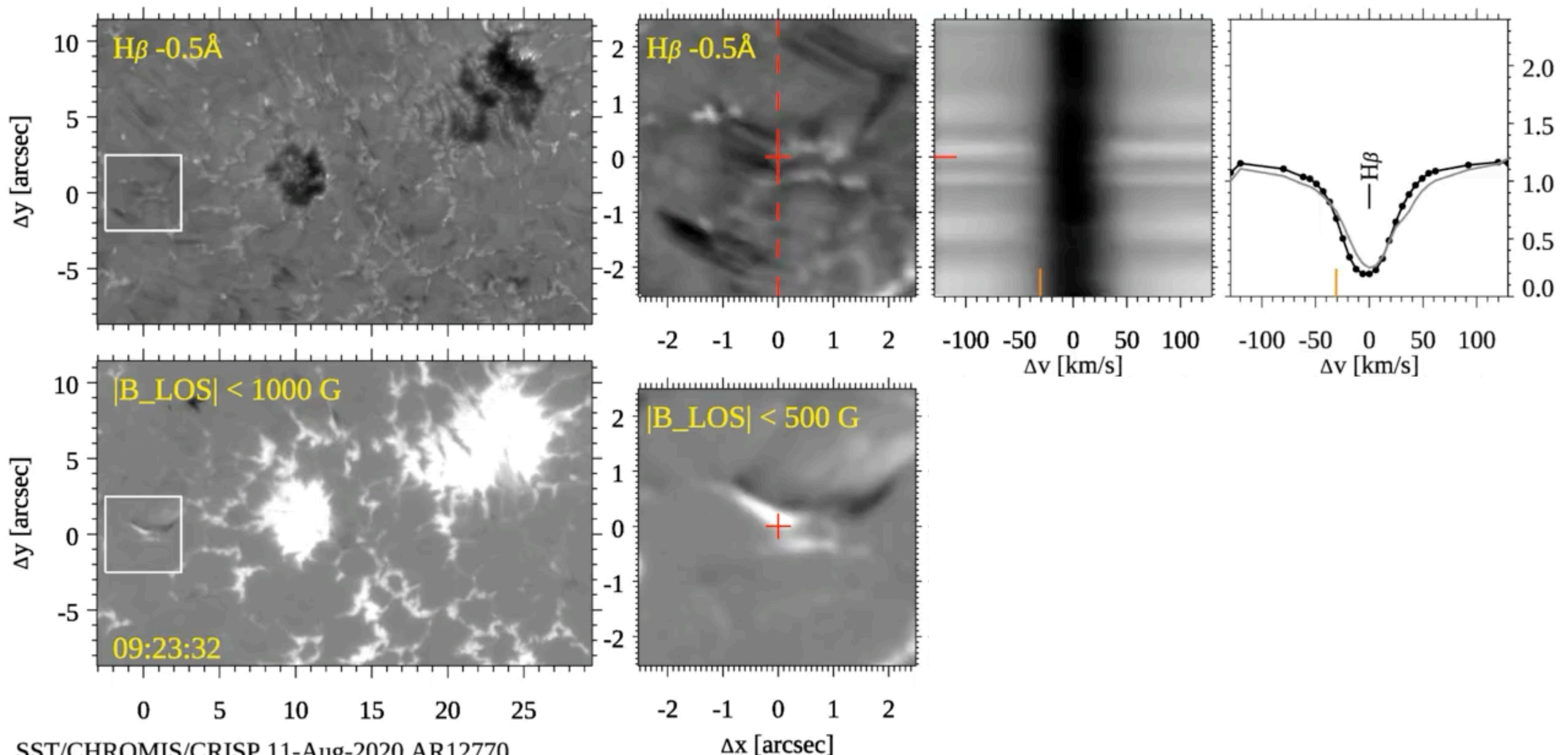
# Ellerman bombs: magnetic reconnection in low solar atmosphere



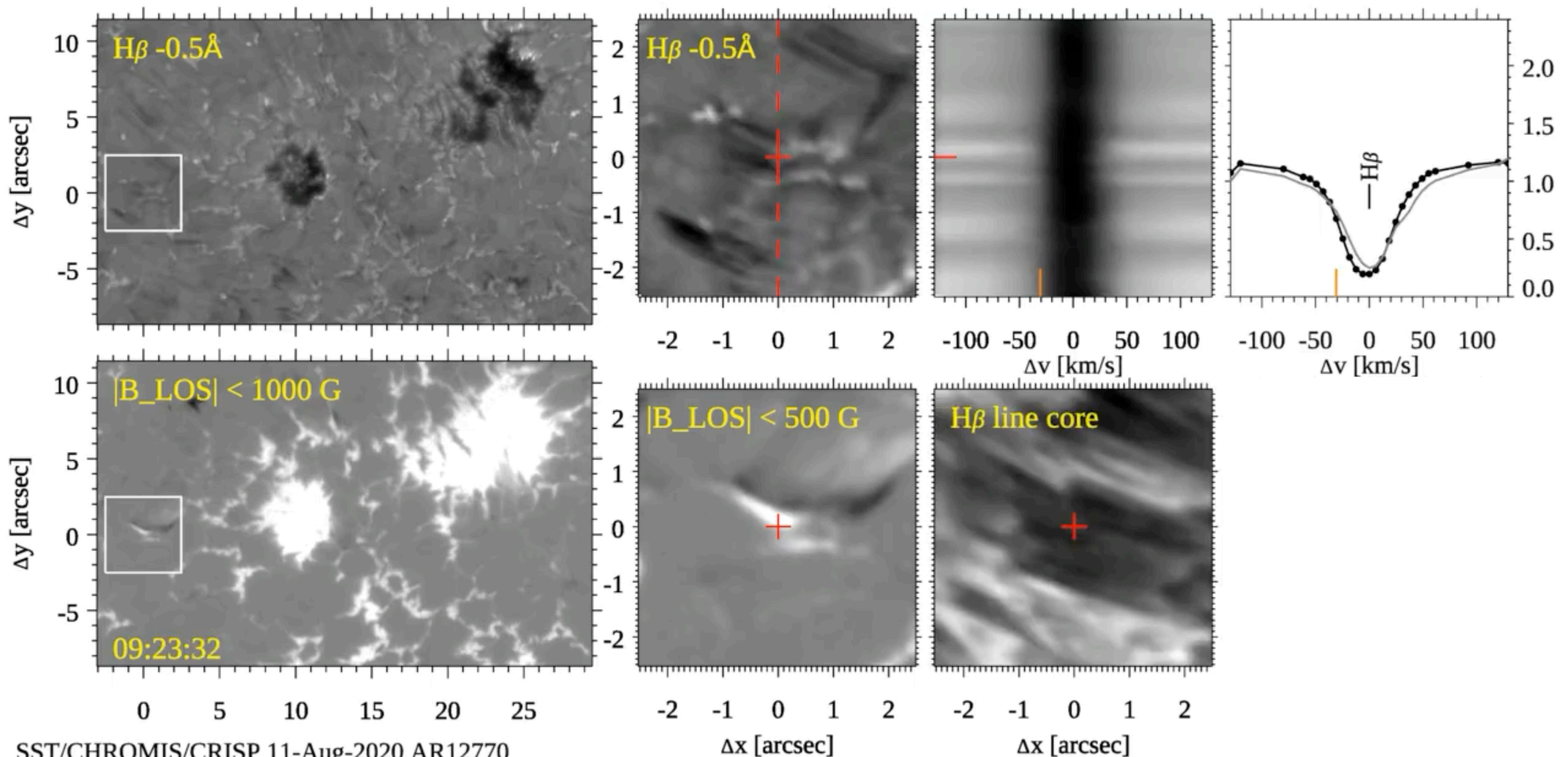
# Ellerman bombs: magnetic reconnection in low solar atmosphere



# Ellerman bombs: magnetic reconnection in low solar atmosphere

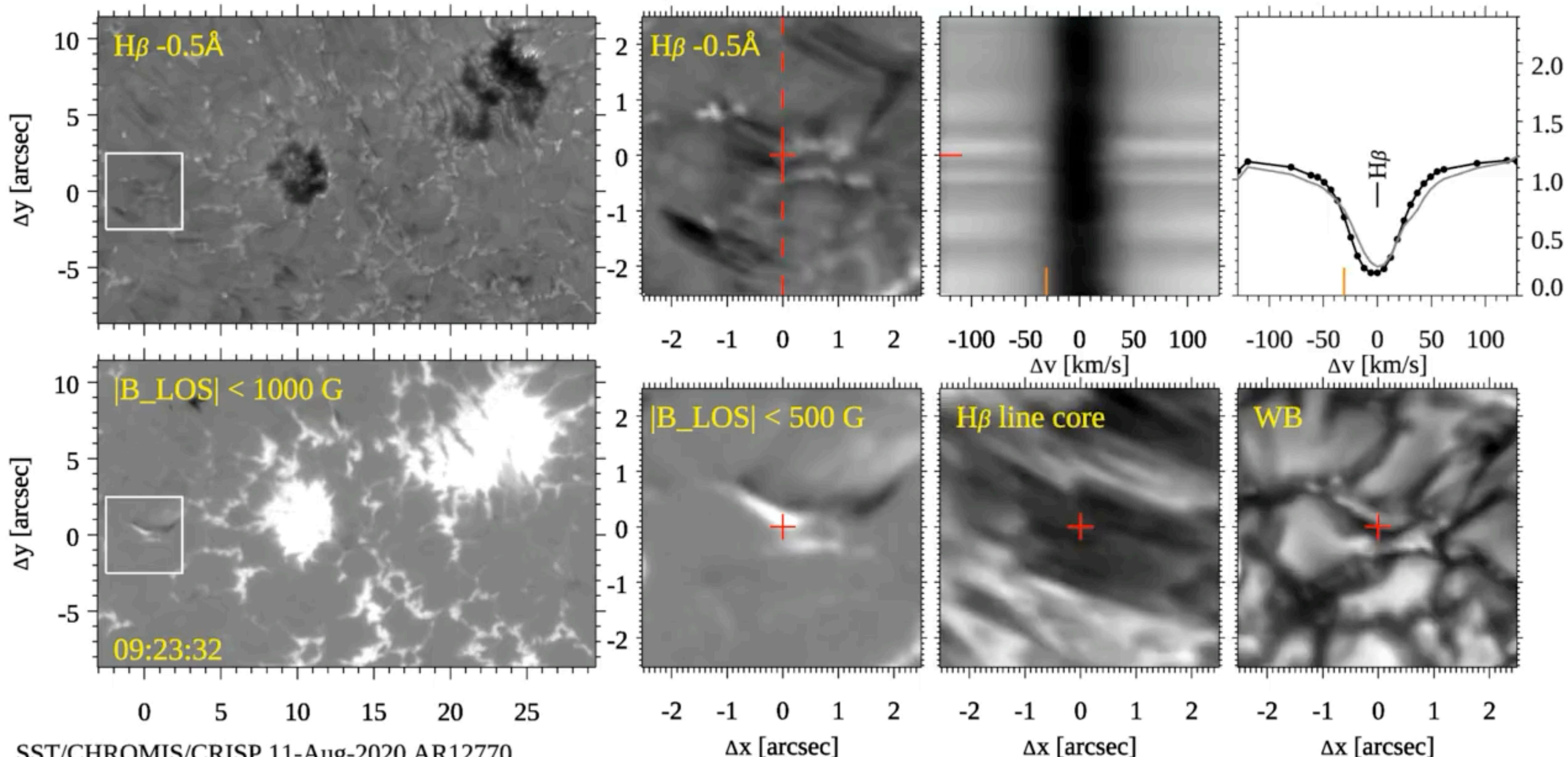


# Ellerman bombs: magnetic reconnection in low solar atmosphere

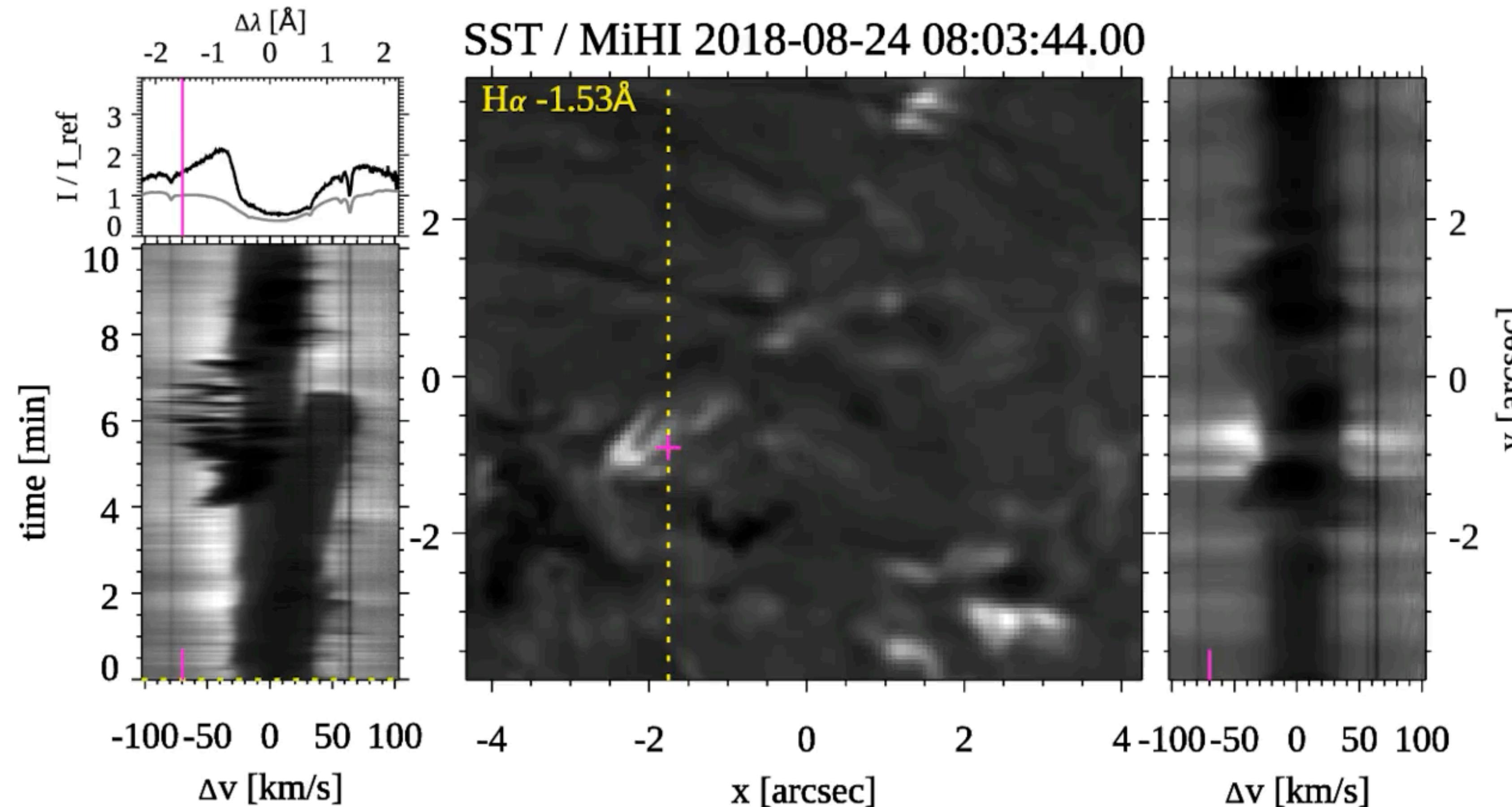


# Ellerman bombs: magnetic reconnection in low solar atmosphere

[http://tsih3.uio.no/lapalma/gallery/textbook\\_EB\\_hbeta\\_11Aug2020.mp4](http://tsih3.uio.no/lapalma/gallery/textbook_EB_hbeta_11Aug2020.mp4)

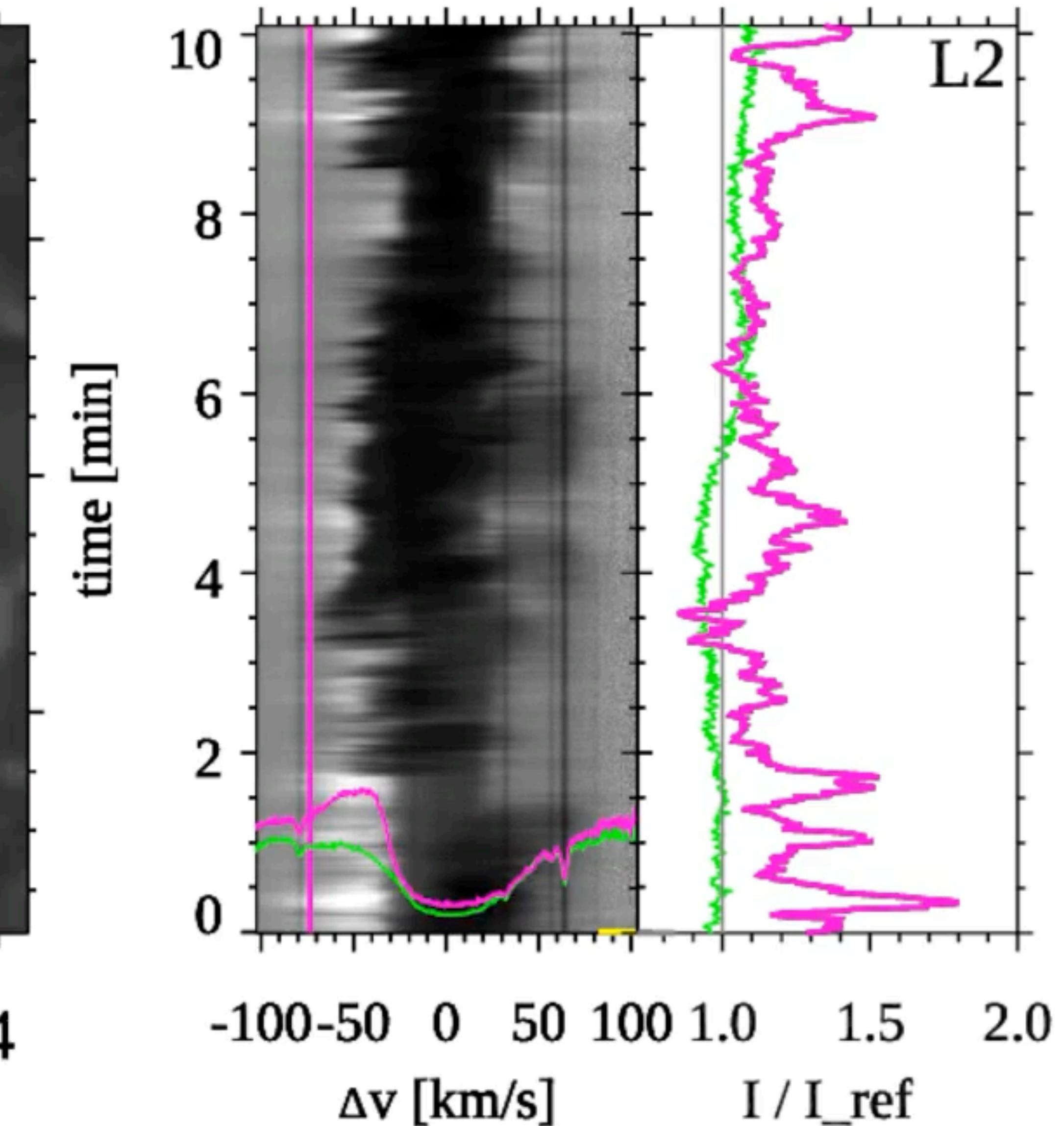
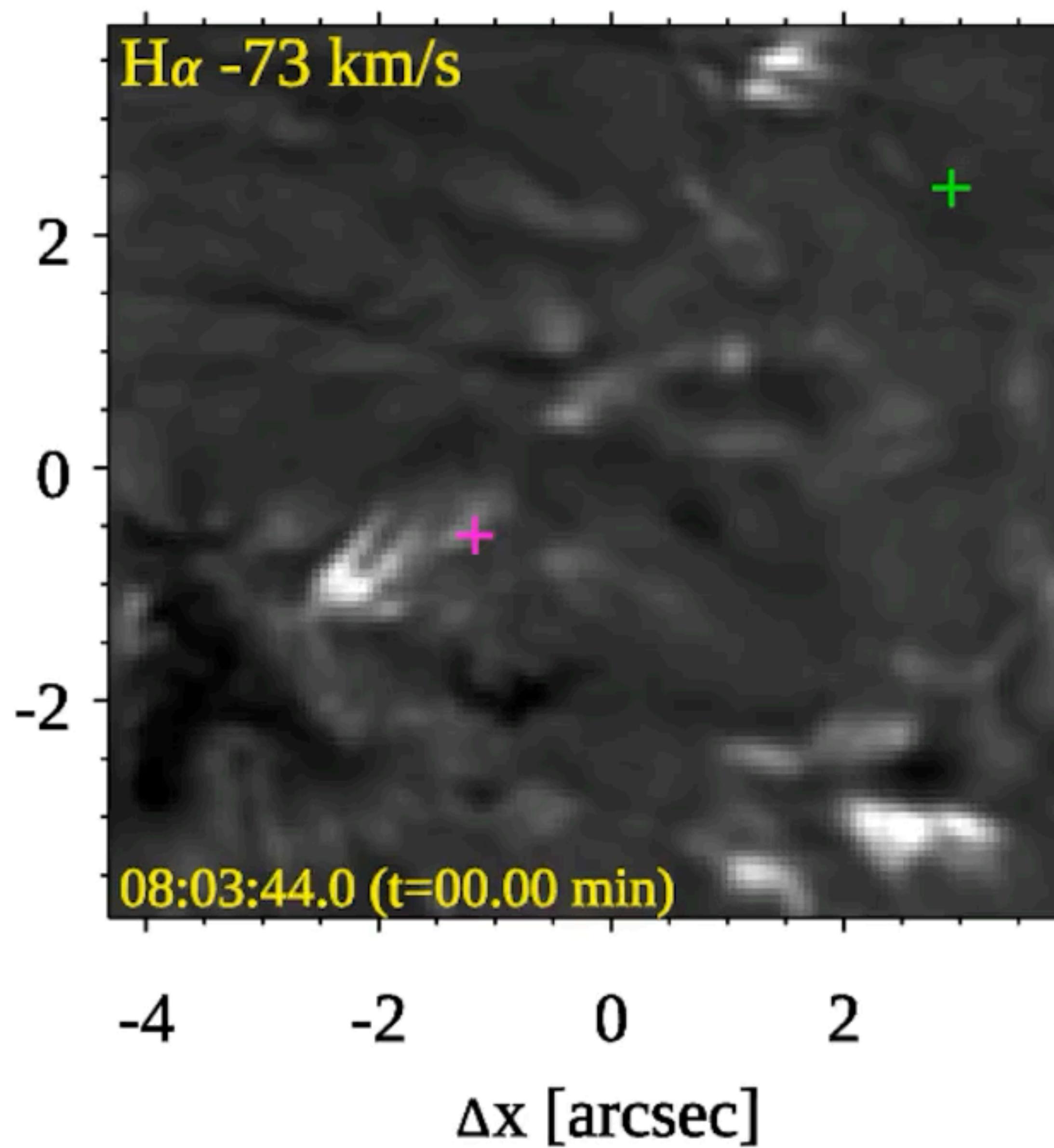


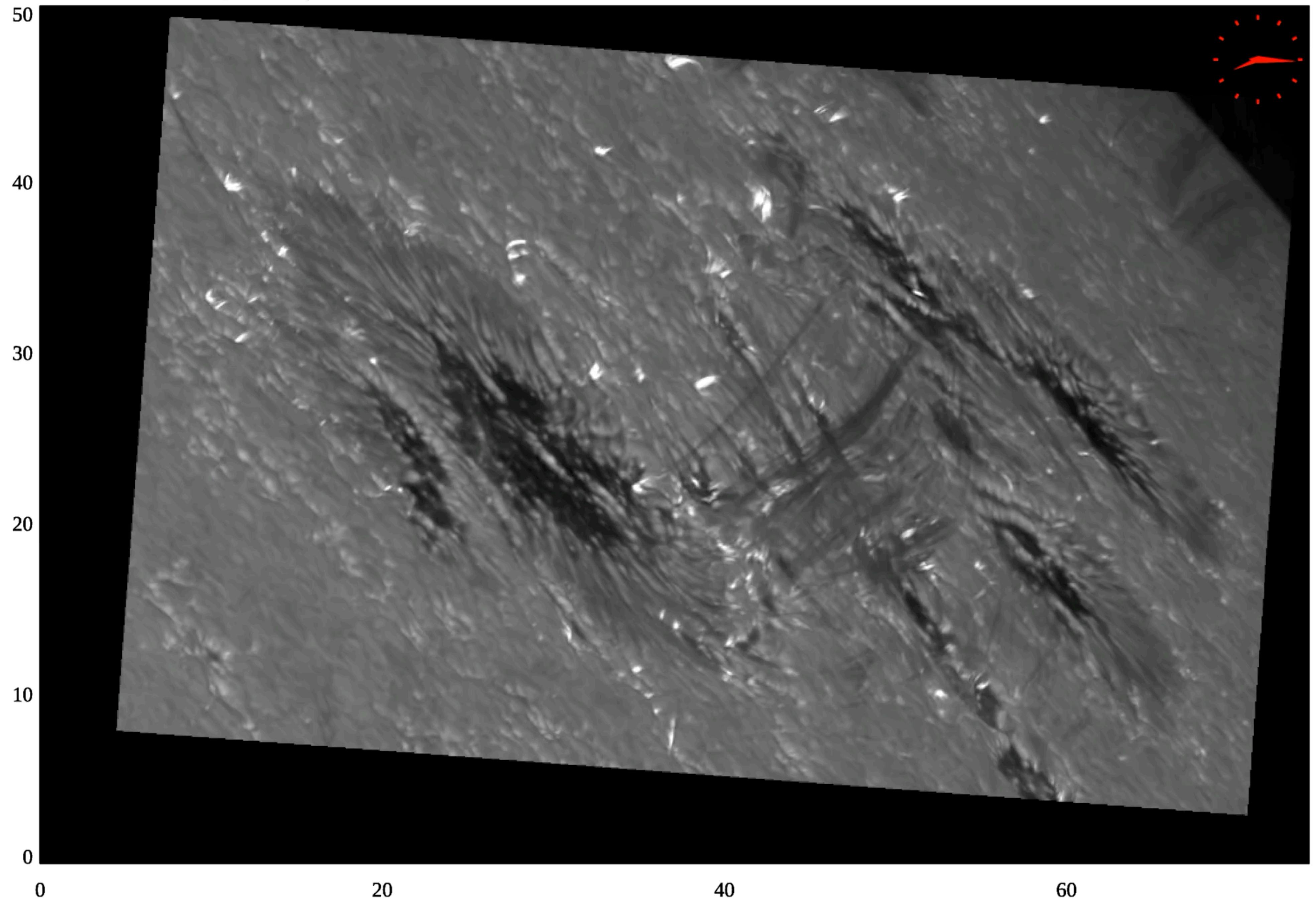
# Unique observations of Ellerman bombs with MiHI instrument



**$H\text{-alpha } 6563 \text{ \AA}$**   
 $8.6'' \times 7.7''$   
 $4.5 \text{ \AA}$  or  $\pm 102 \text{ km/s}$   
1.33 s cadence  
10 min duration

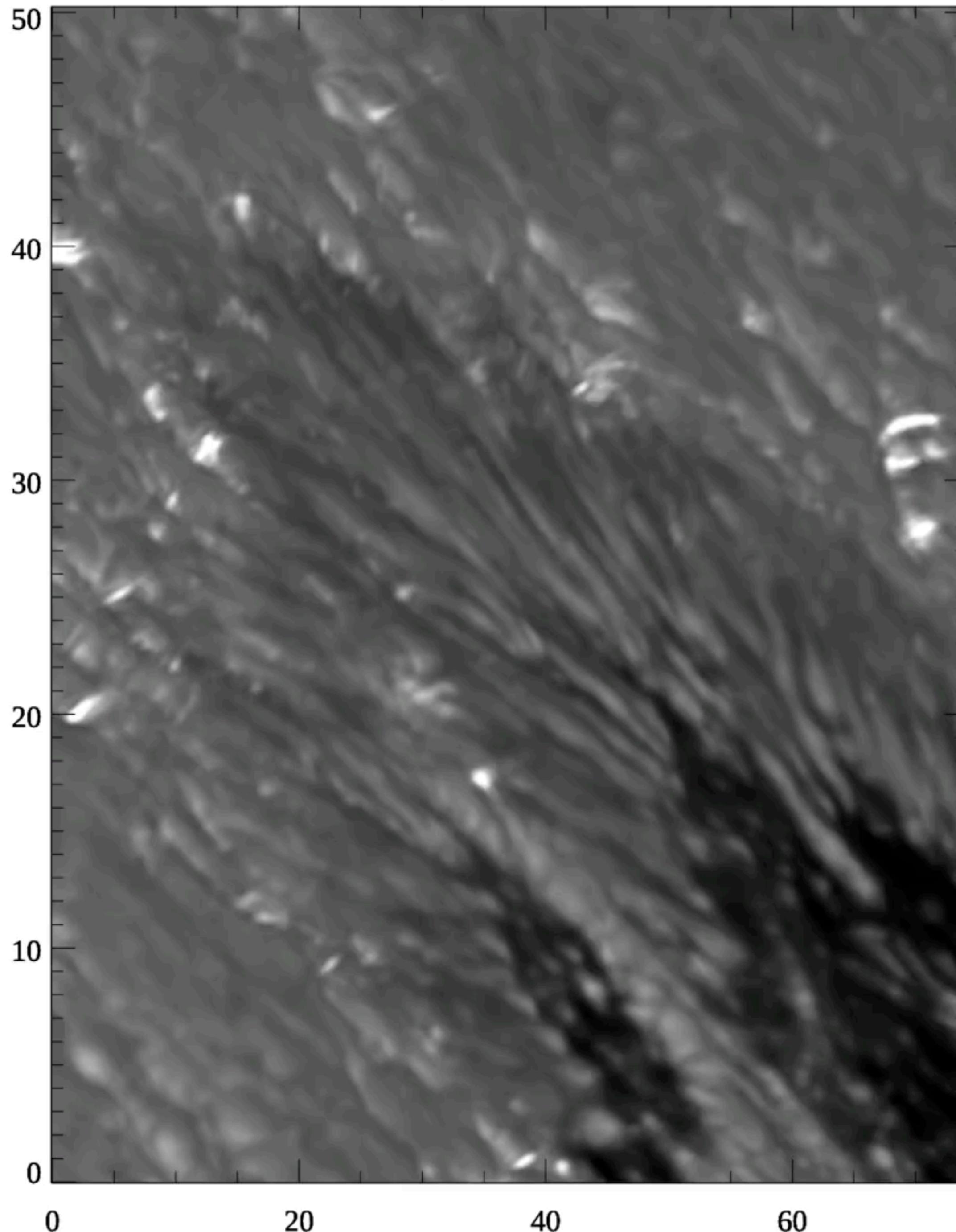
# Plasmoid-like blobs emerging from Ellerman bombs: FWHM sizes 0.1 - 0.4"



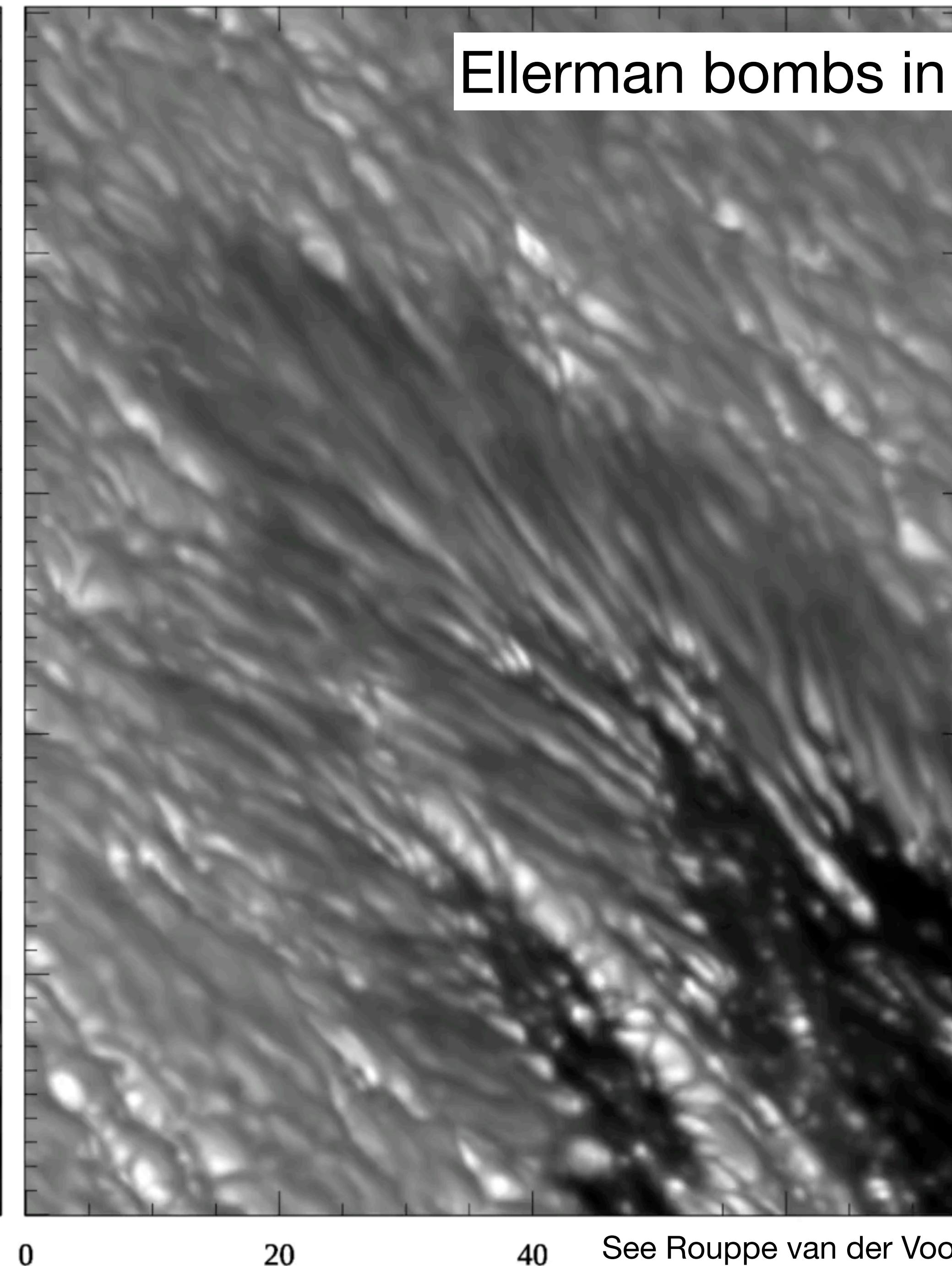


**Ellerman bombs:**  
magnetic reconnection  
in low solar atmosphere

SST / CHROMIS 2024-05-24 H $\beta$  -0.7 Å



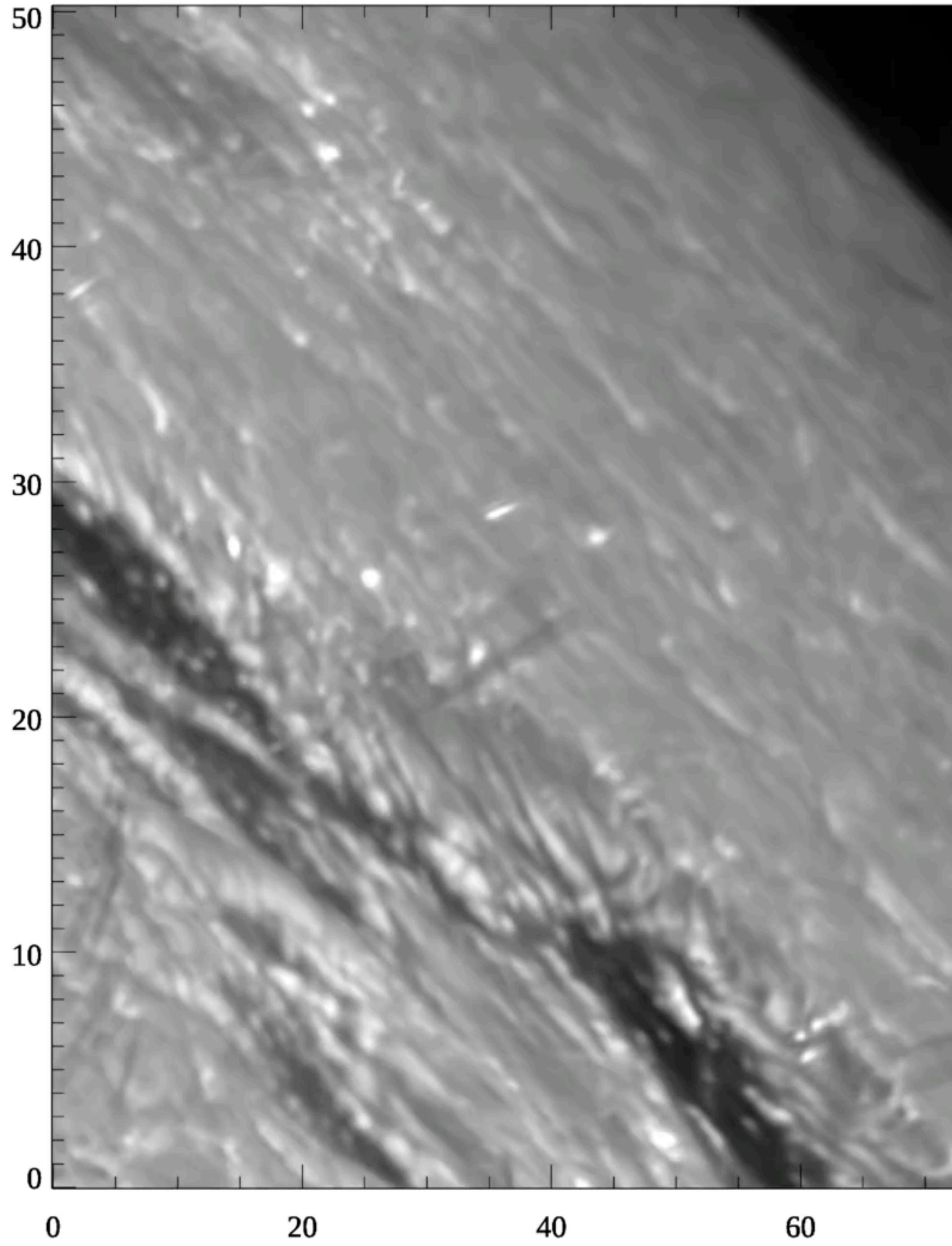
WB 4846 Å



08:15:28

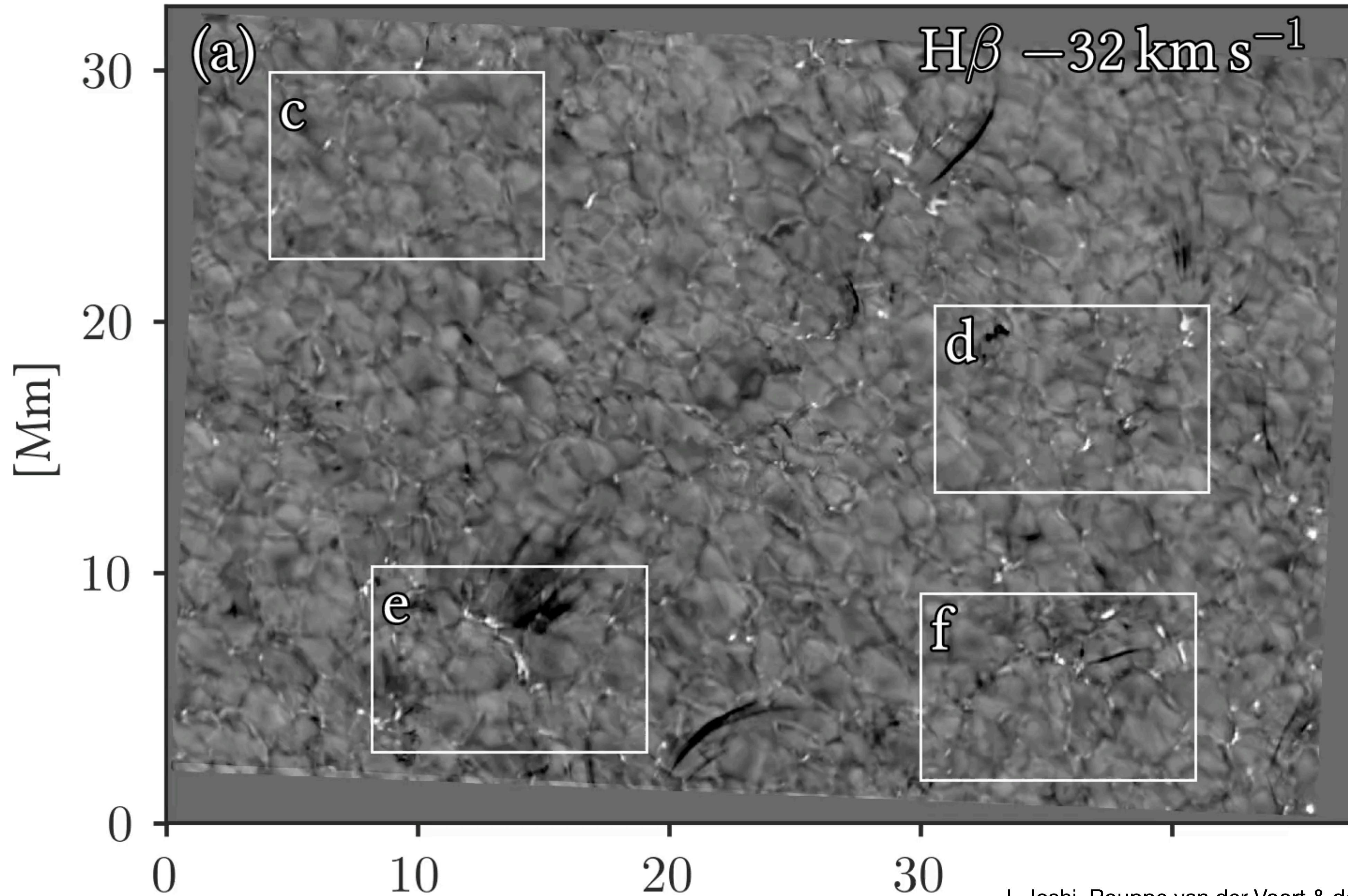
Ellerman bombs in the penumbra

SST / CHROMIS 2024-05-24 H $\beta$  -0.7 Å

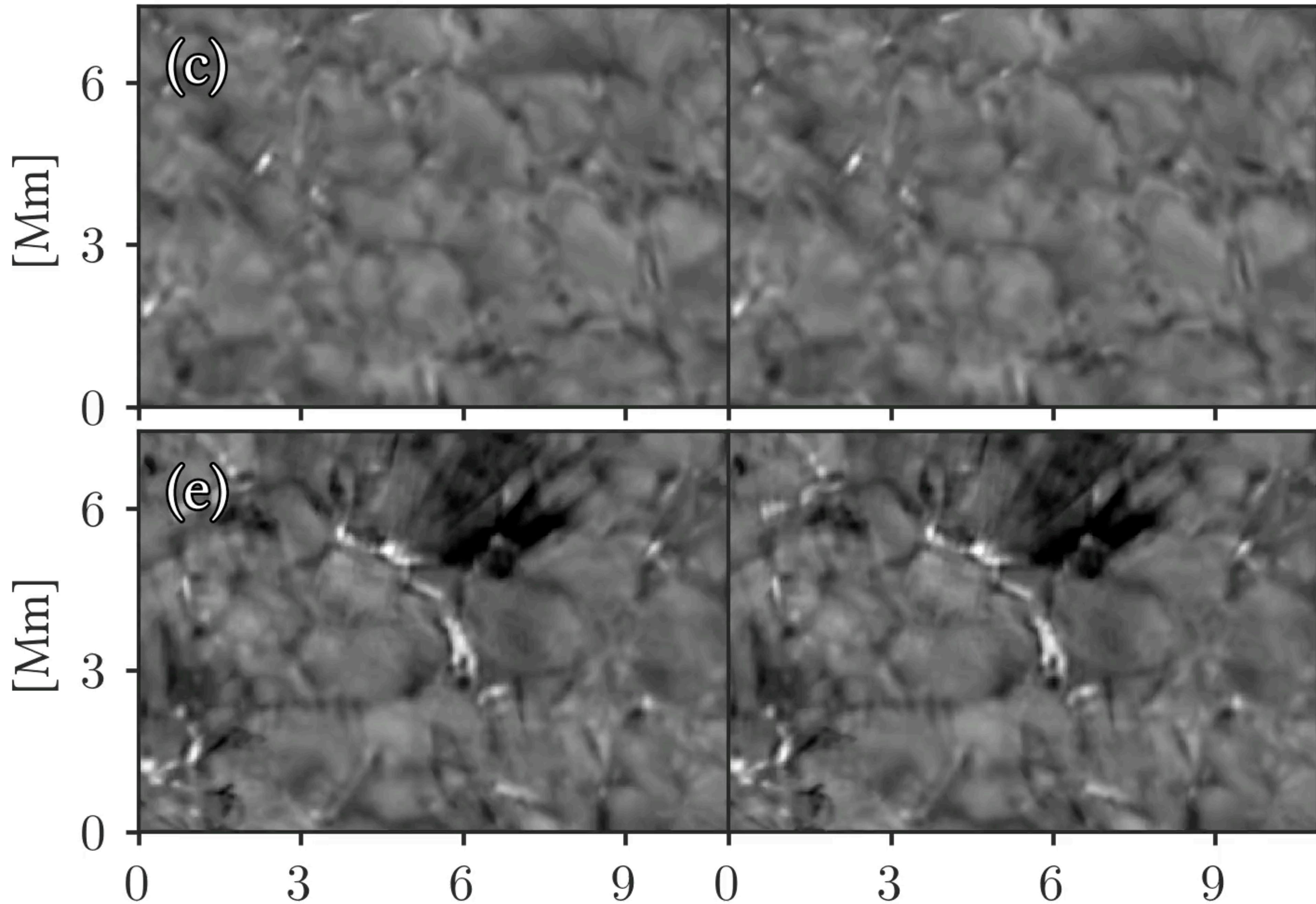


Ellerman bombs in the Quiet Sun

2019-06-06 08:41:14 UT



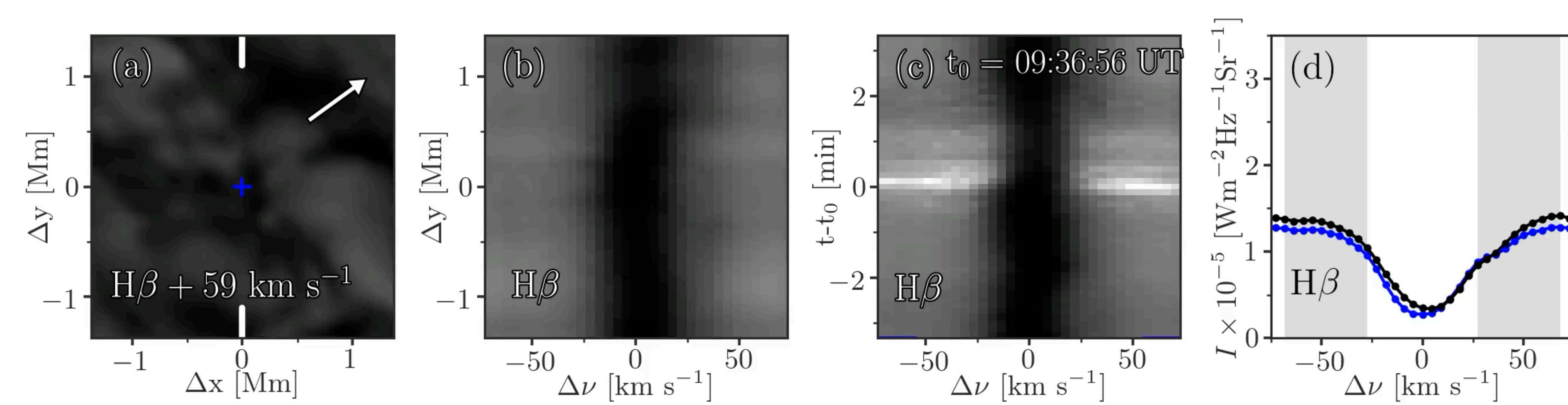
Large numbers of  
Ellerman bombs in  
quiet Sun:  
>120 QSEBs



Large numbers of Ellerman bombs in quiet Sun:  
→ >120 QSEBs  
→ >750,000 over whole Sun

QSEB at  $\mu=0.76$

Flame, strong H $\beta$  wing

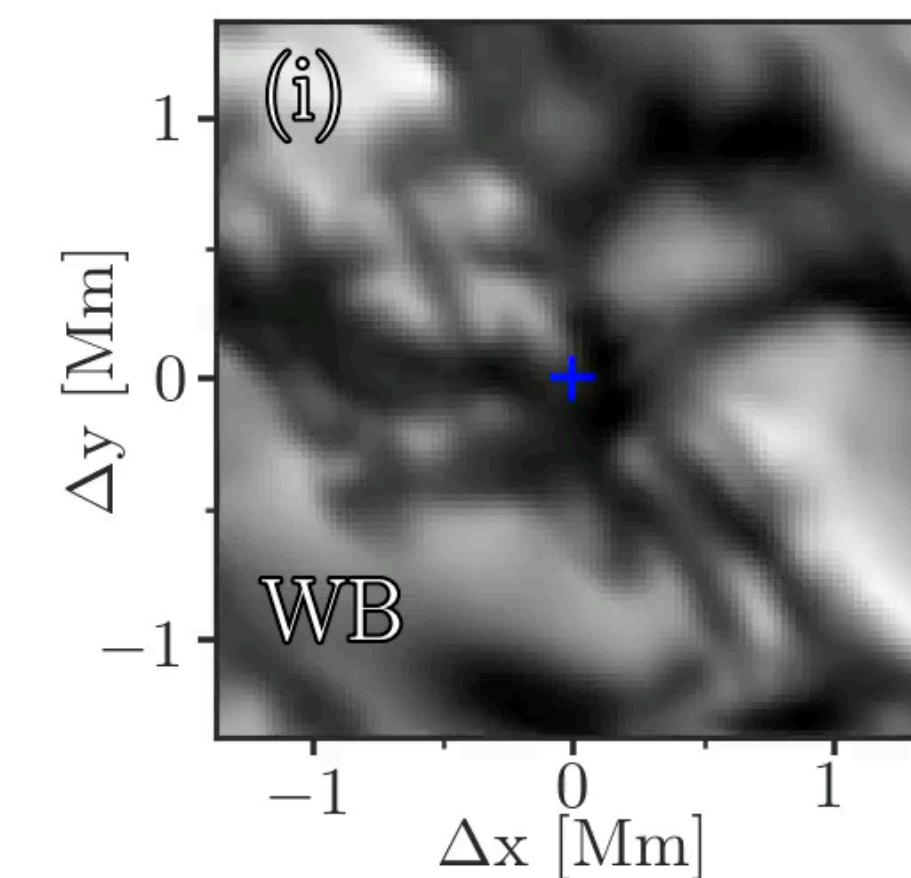
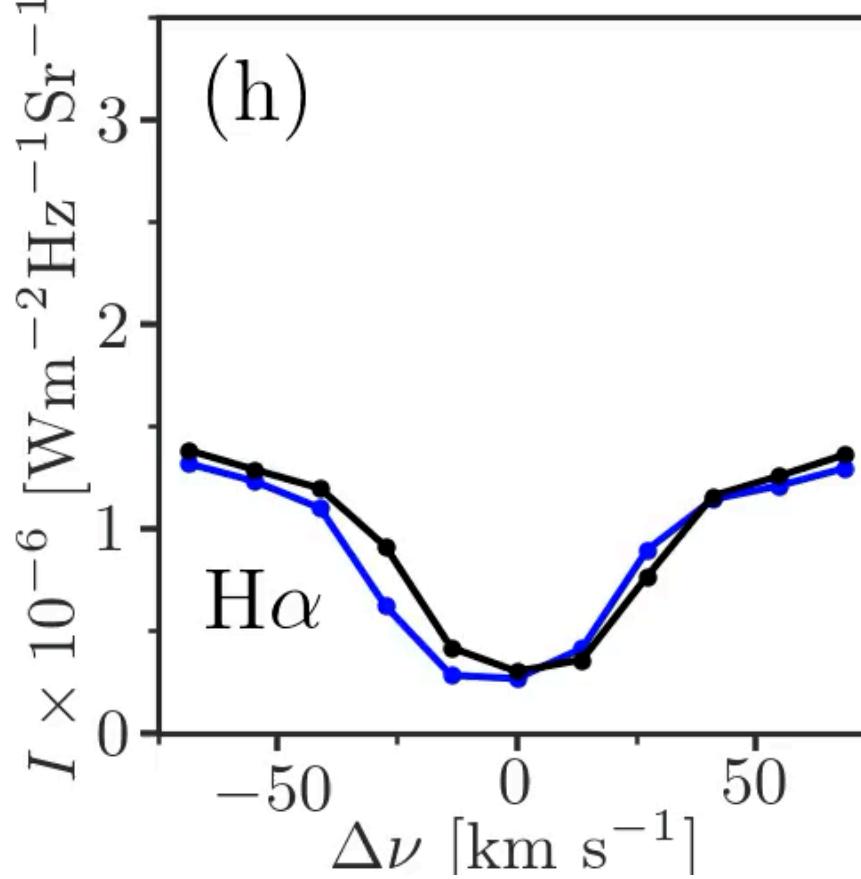
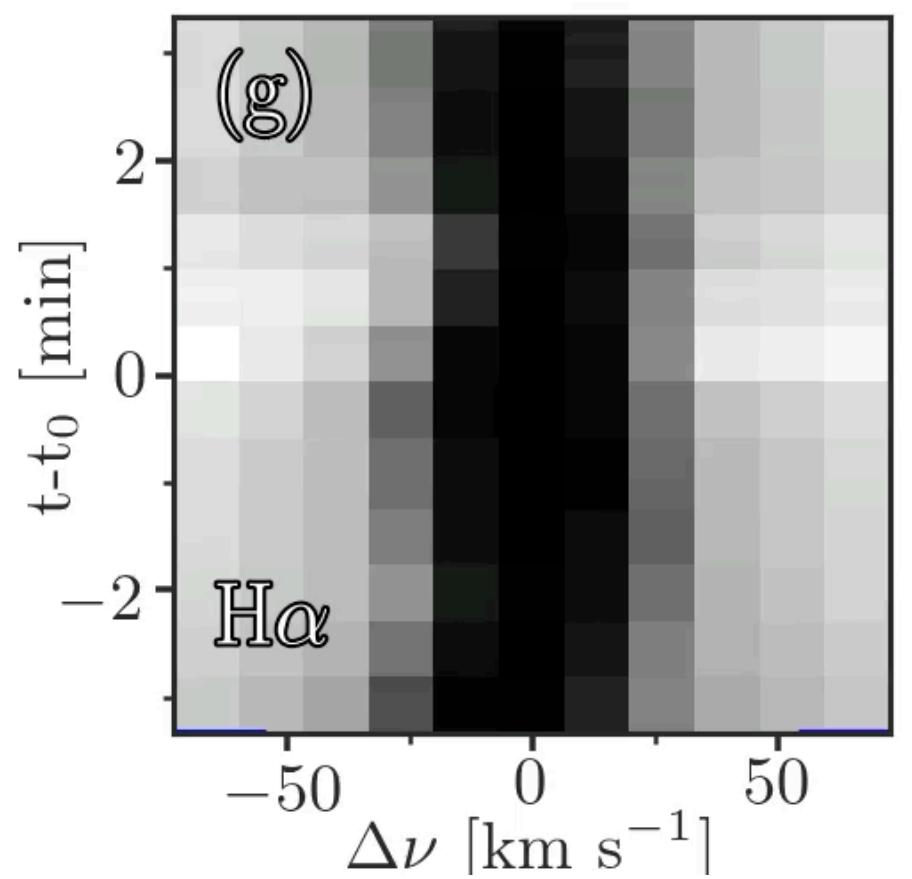
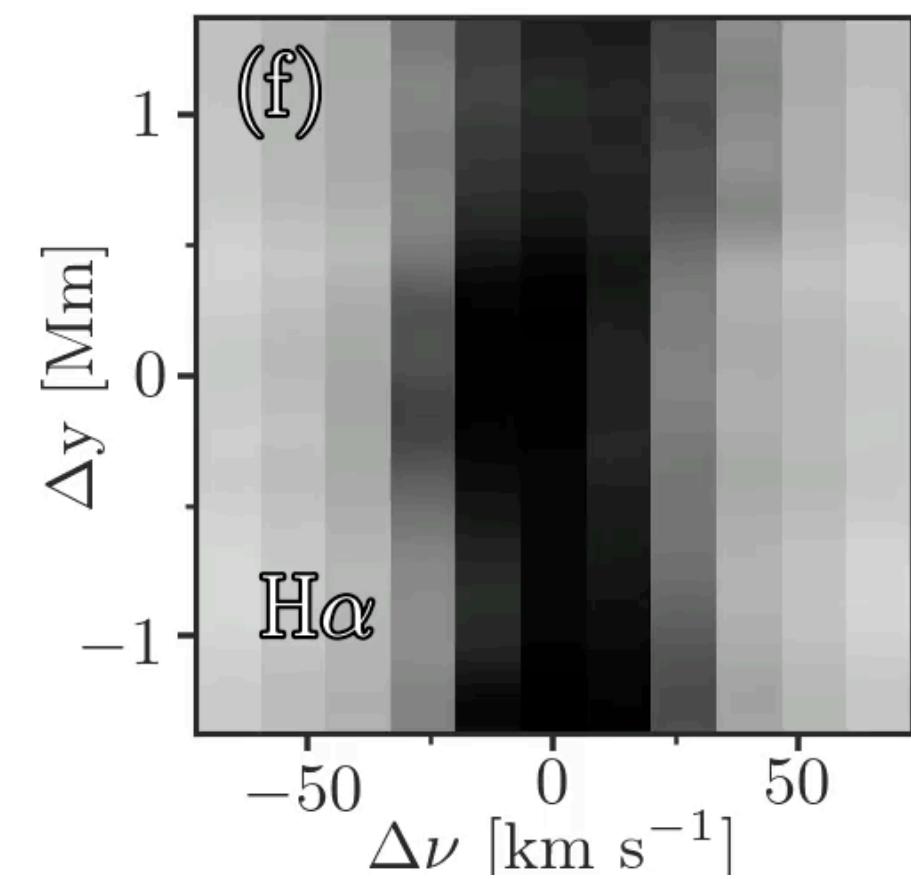
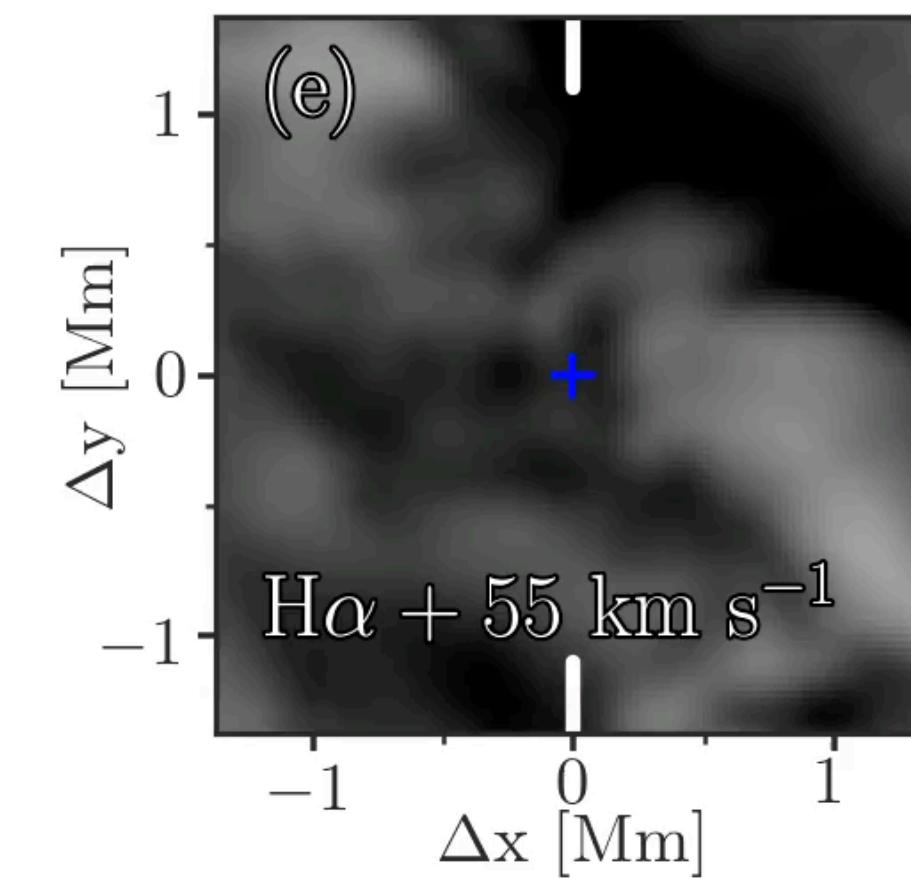
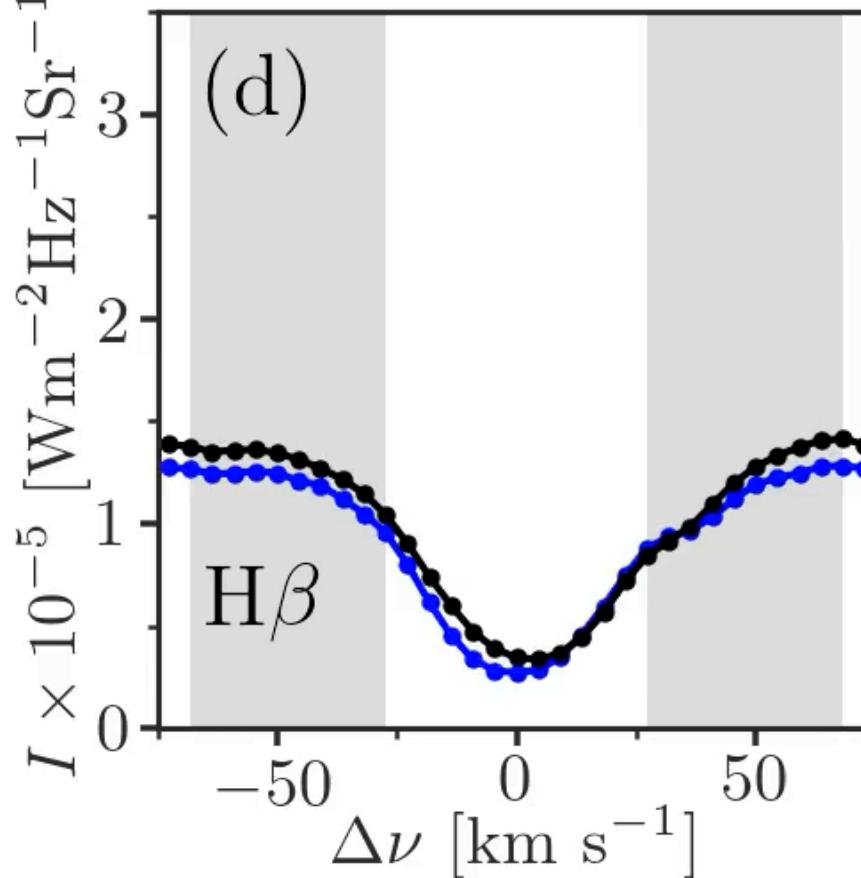
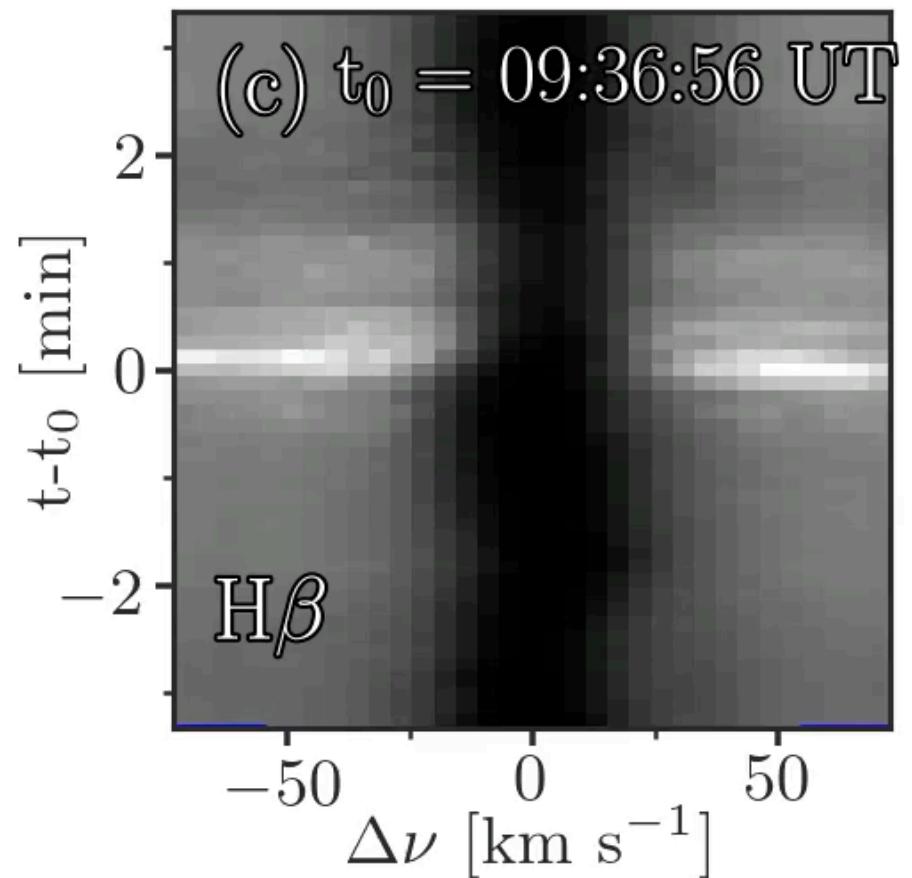
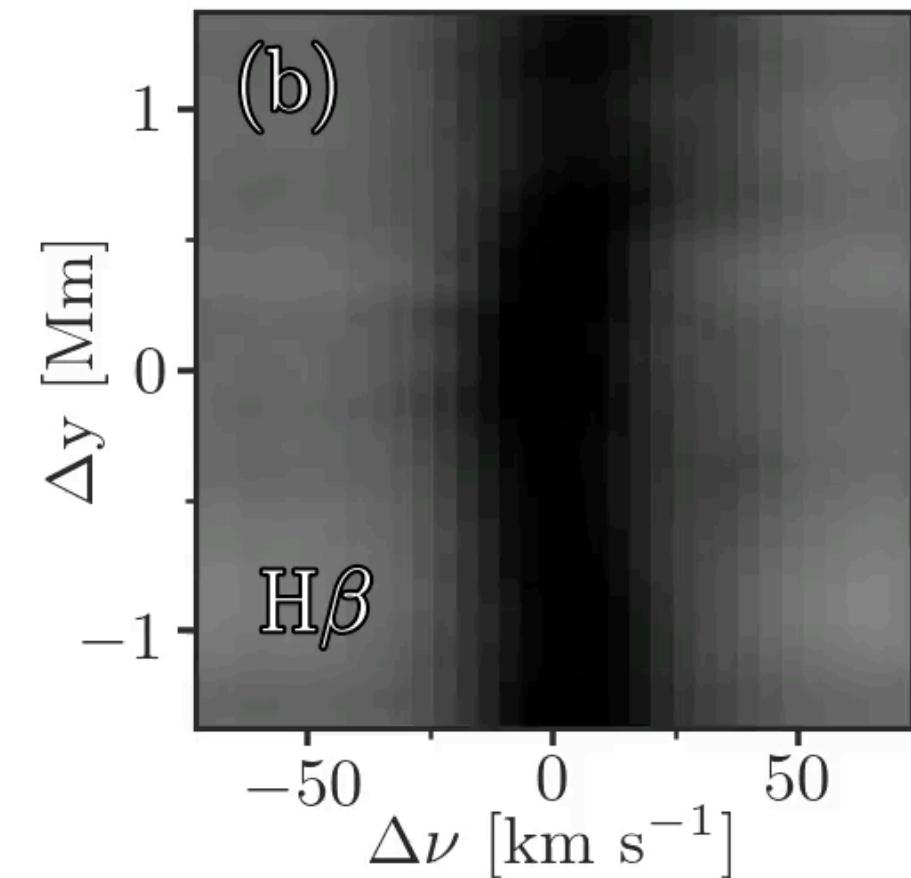
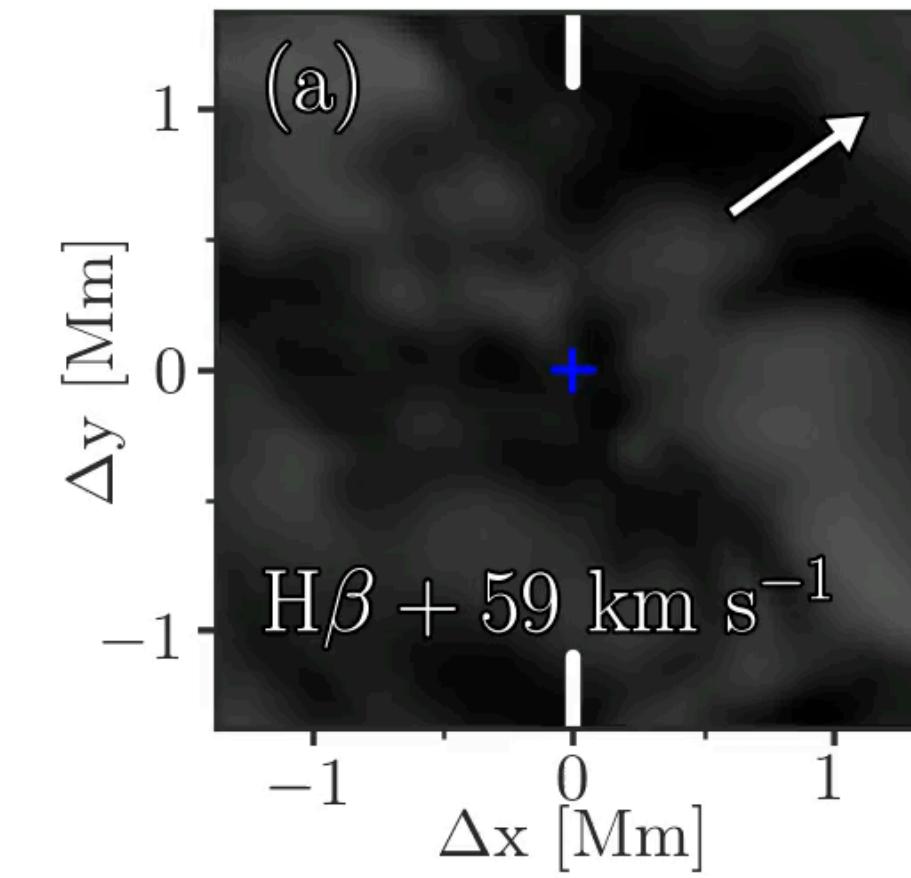


QSEB at  $\mu=0.76$

Flame, strong H $\beta$  wing

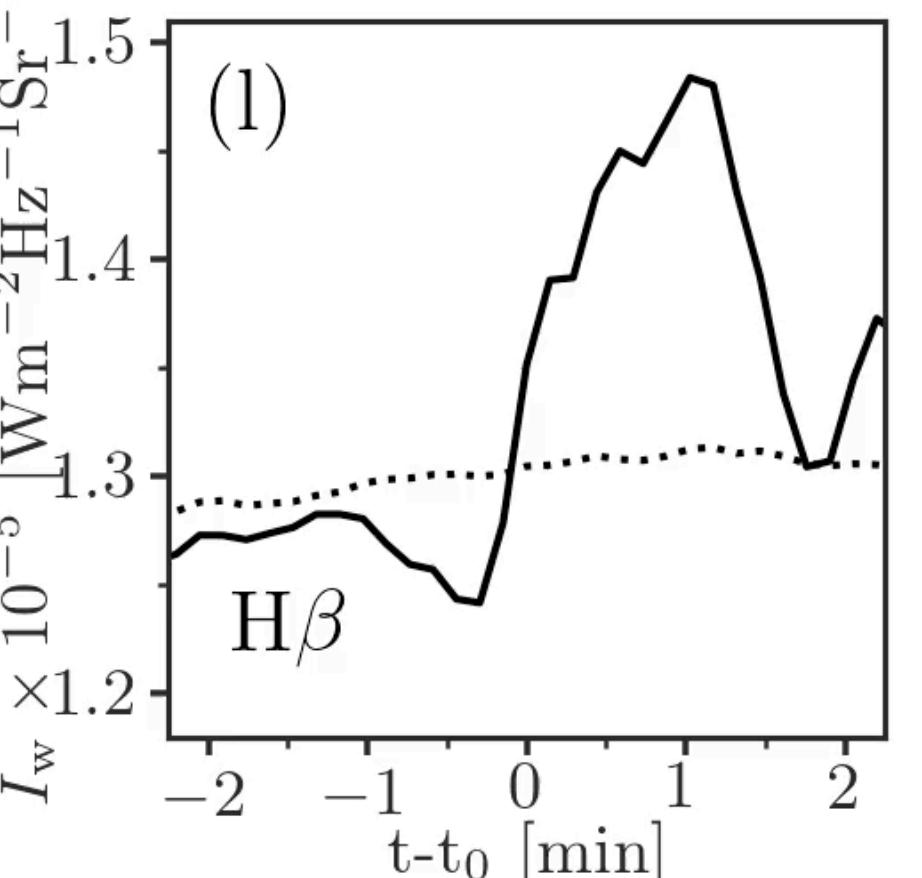
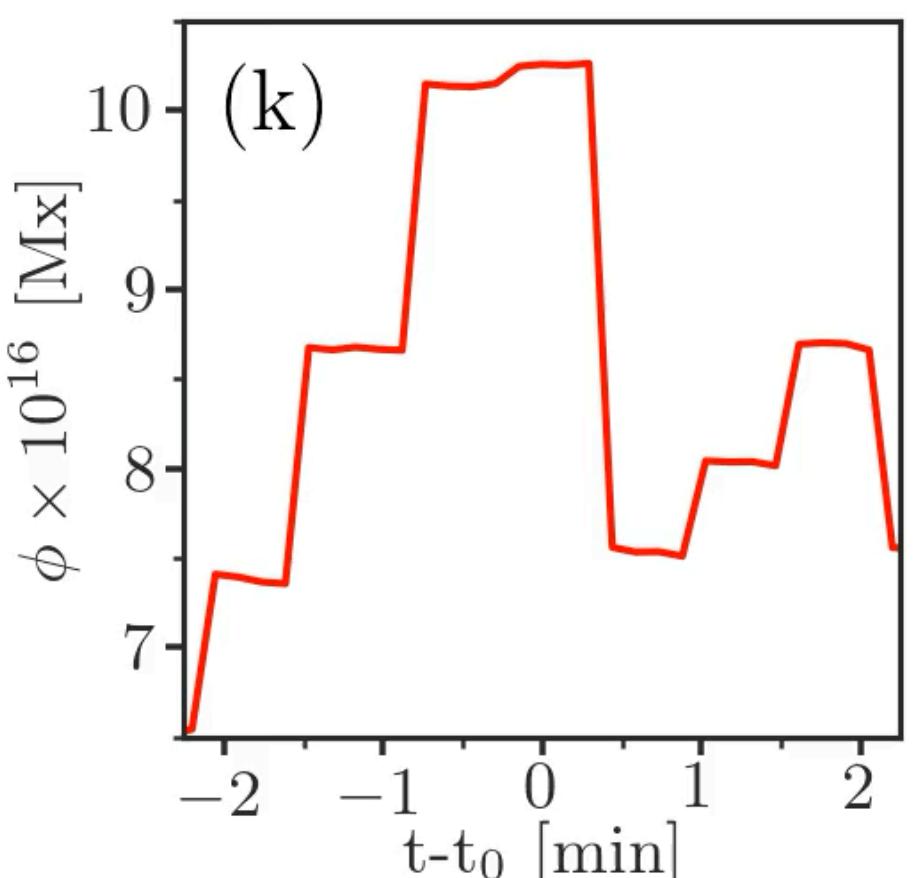
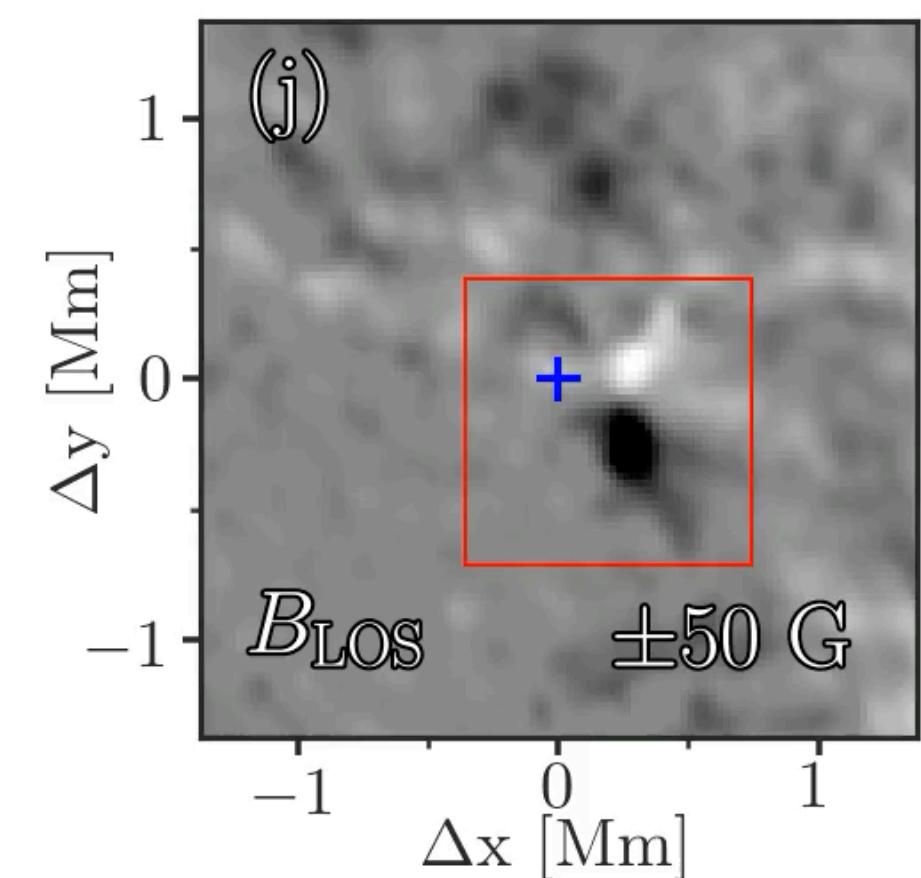
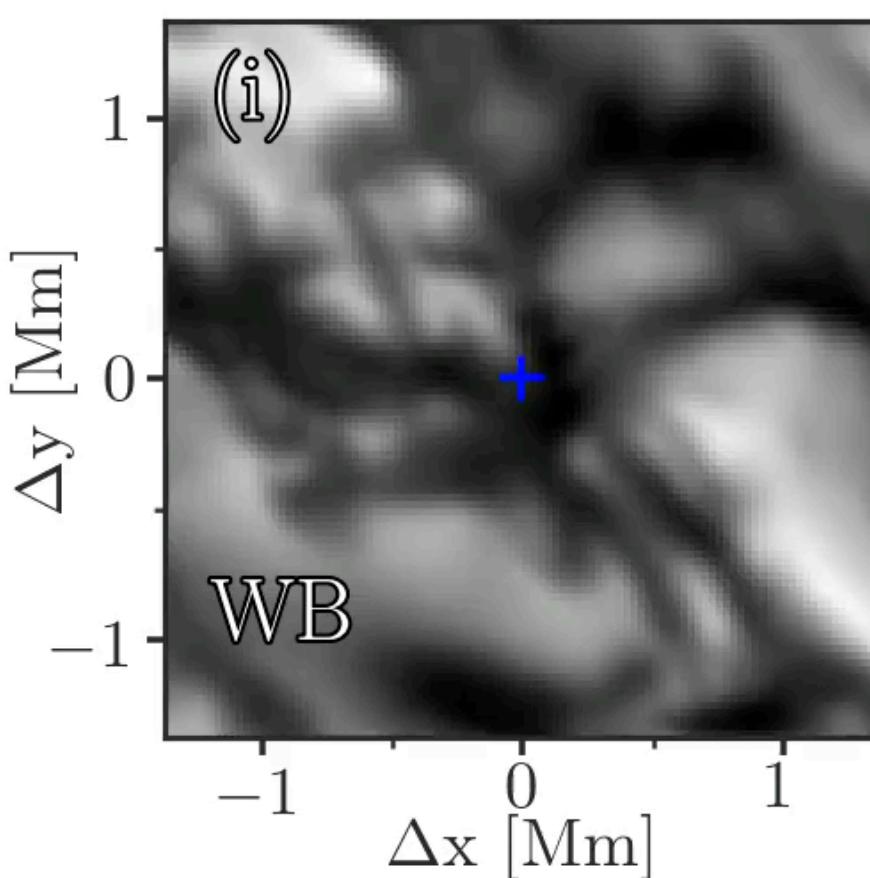
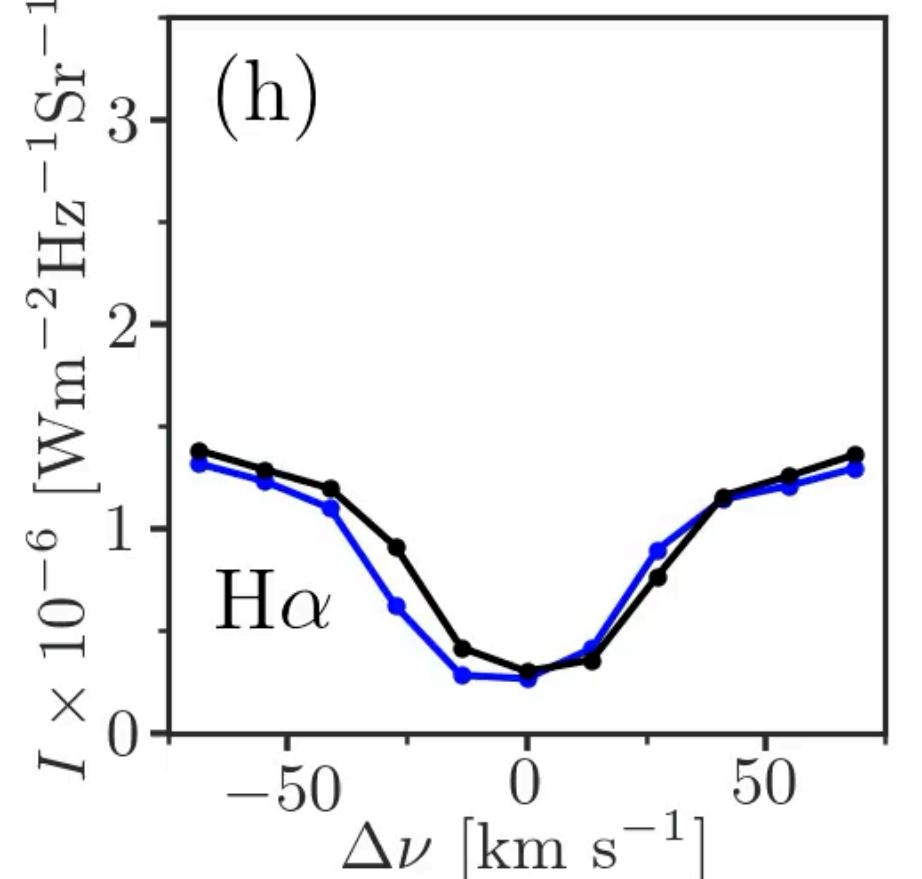
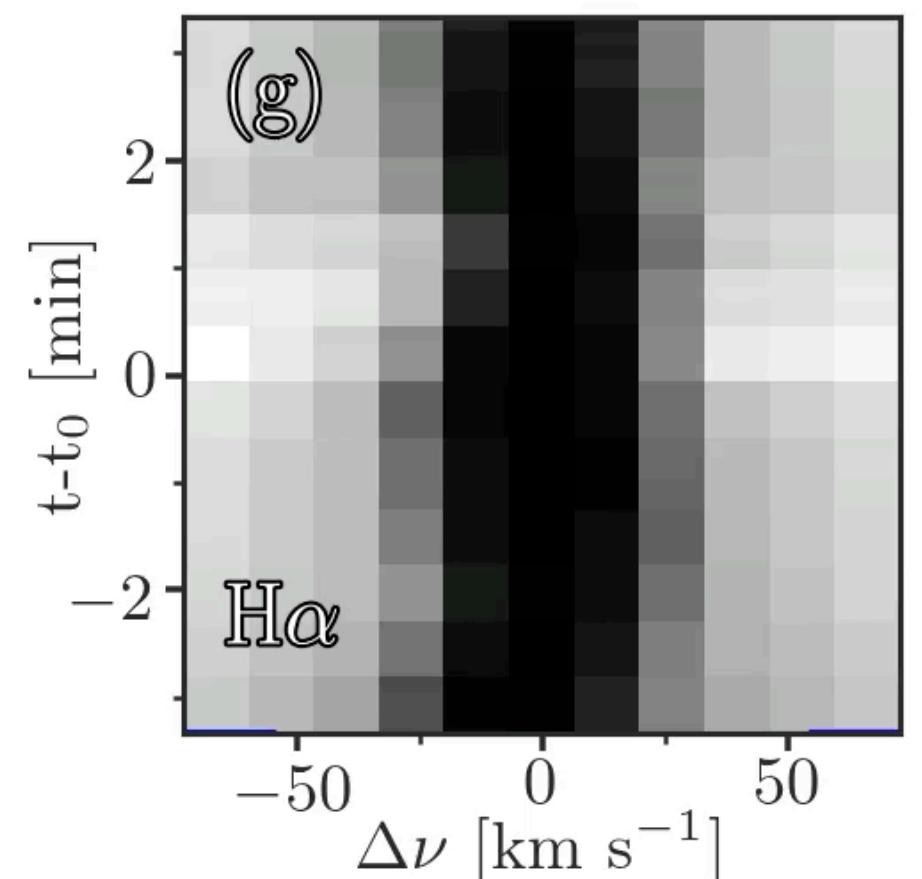
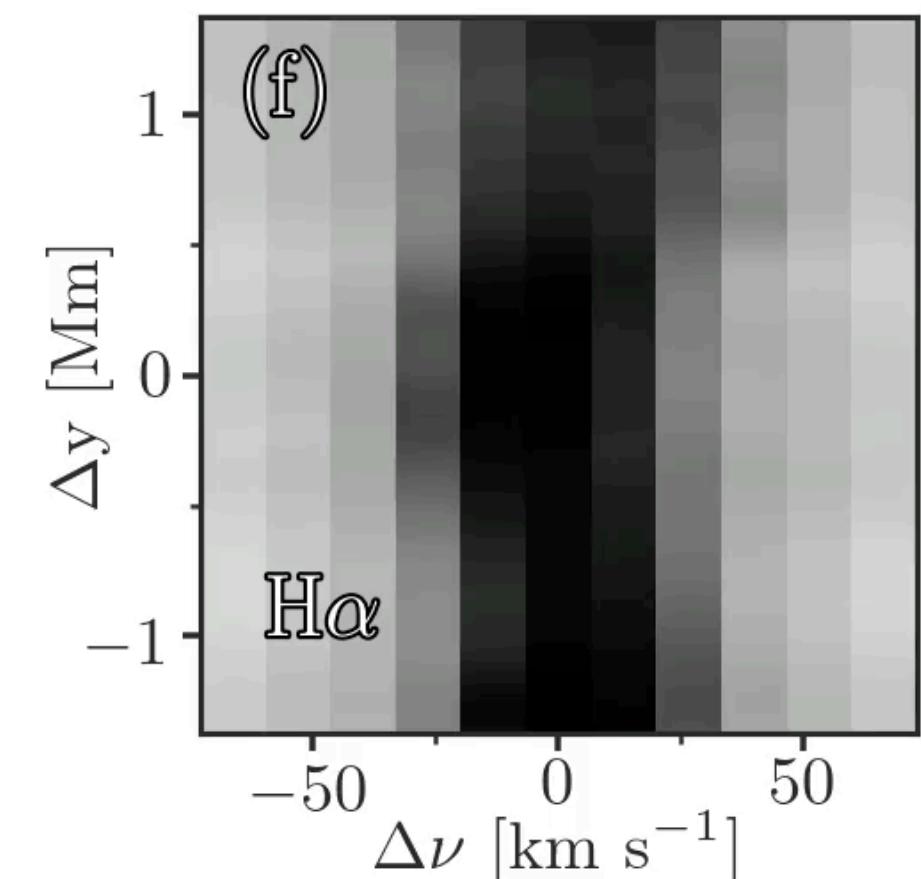
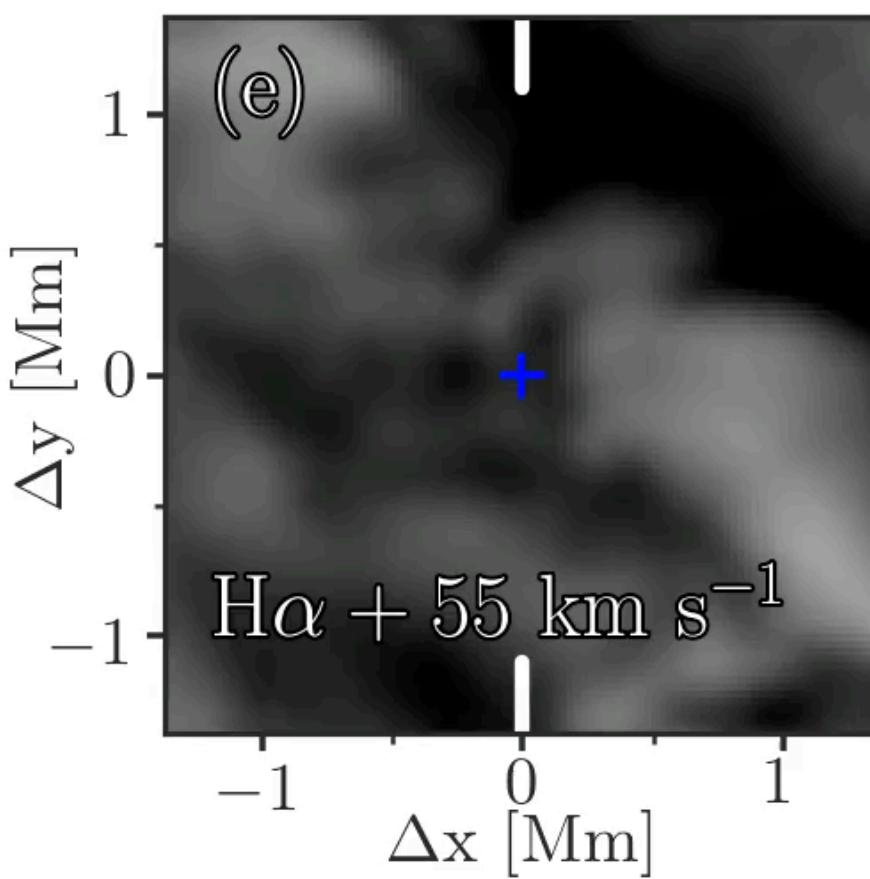
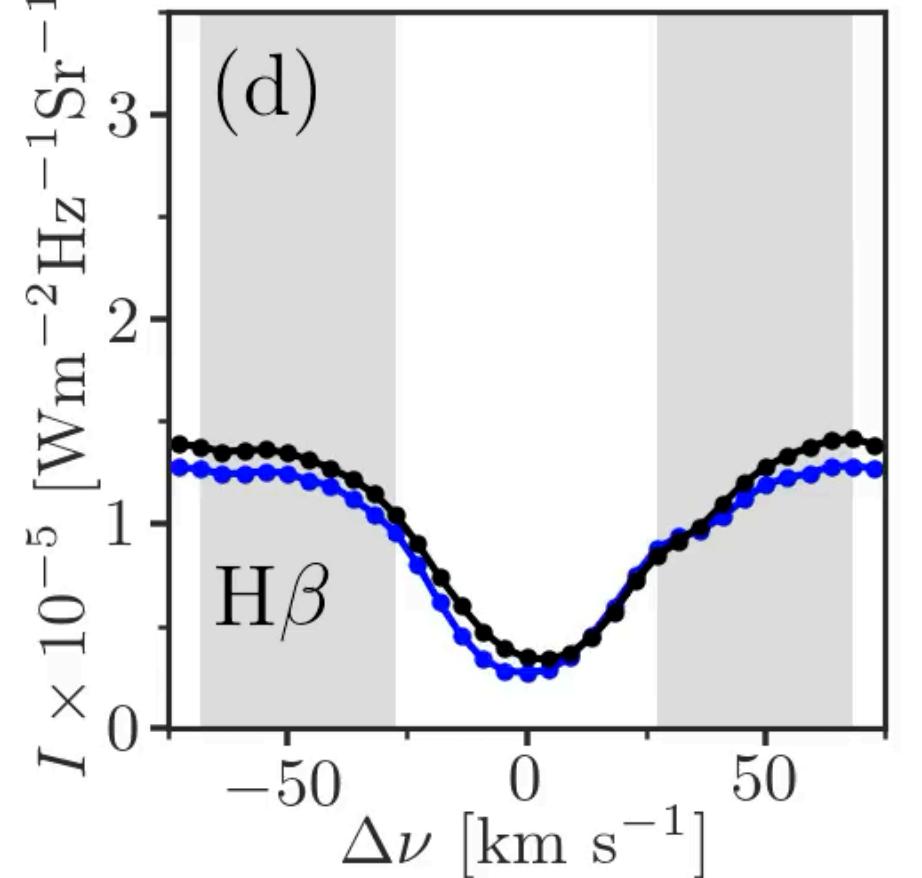
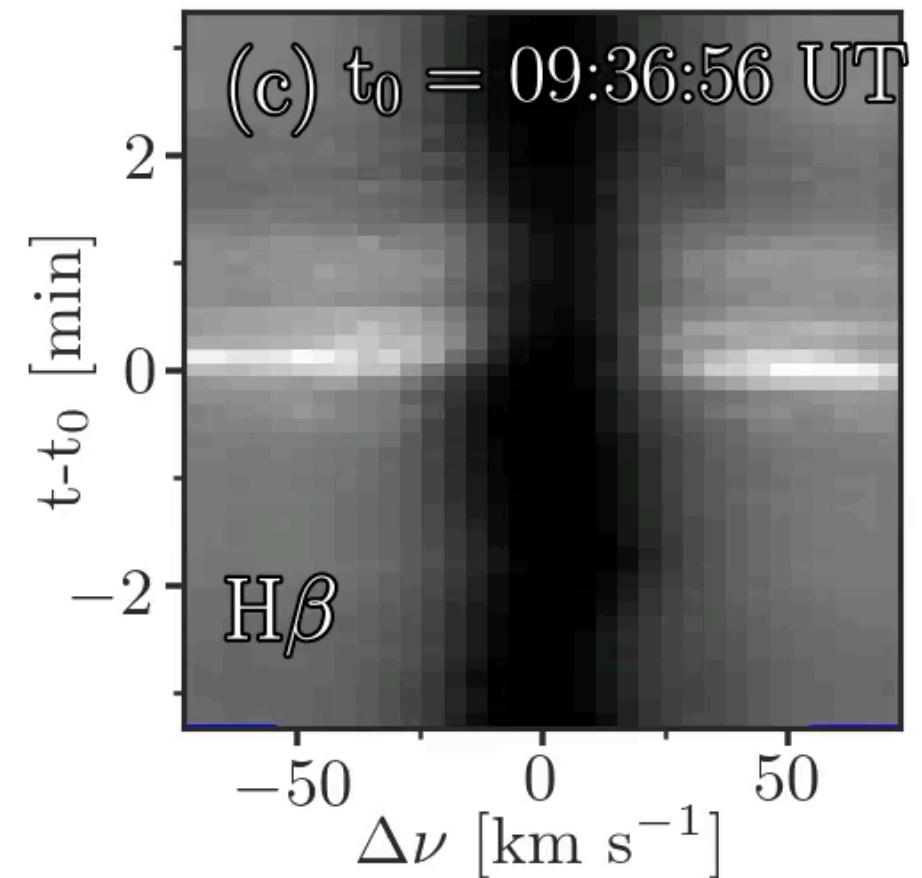
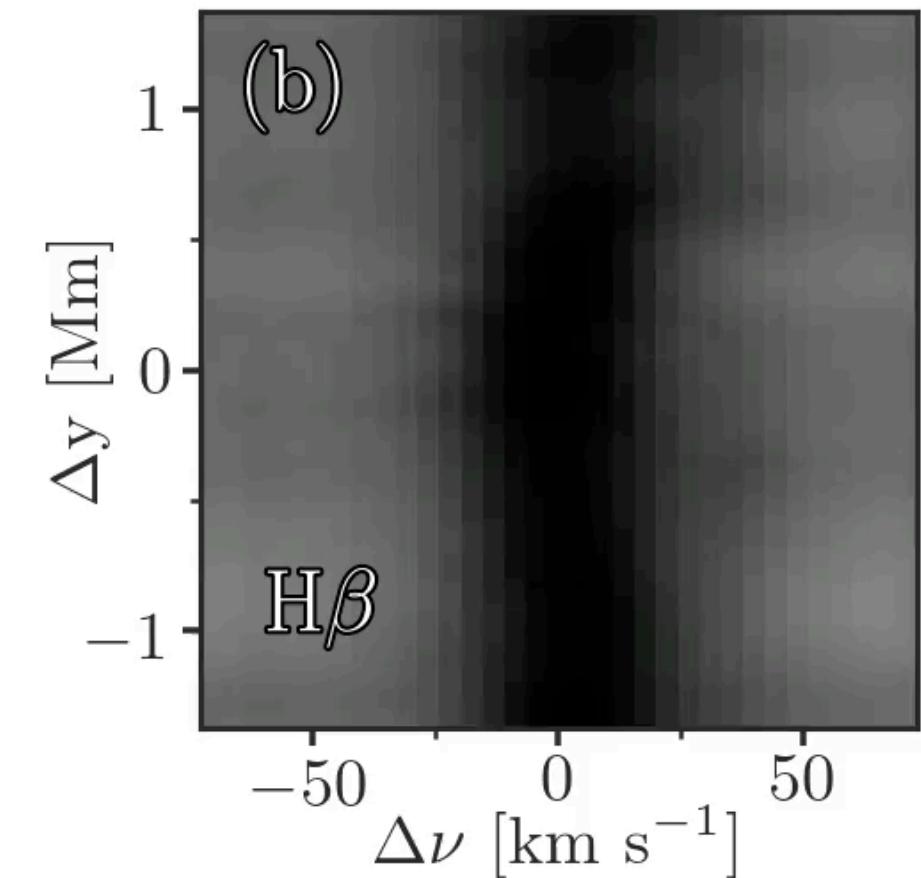
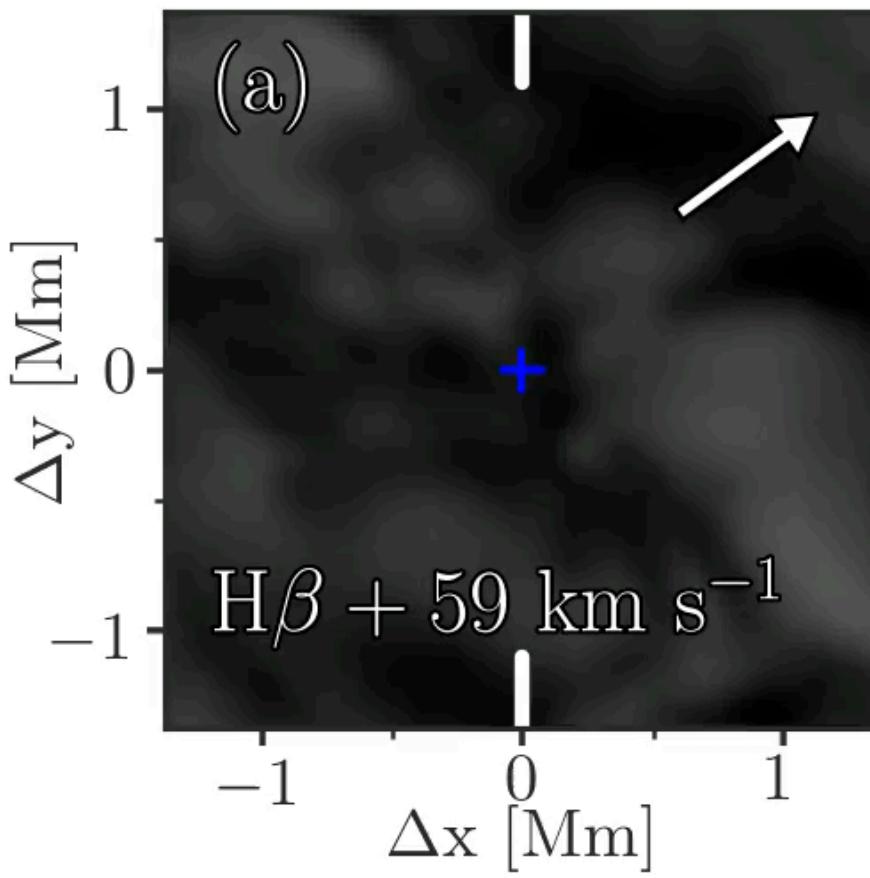
Weak in H $\alpha$  wing

Absent in WB cont

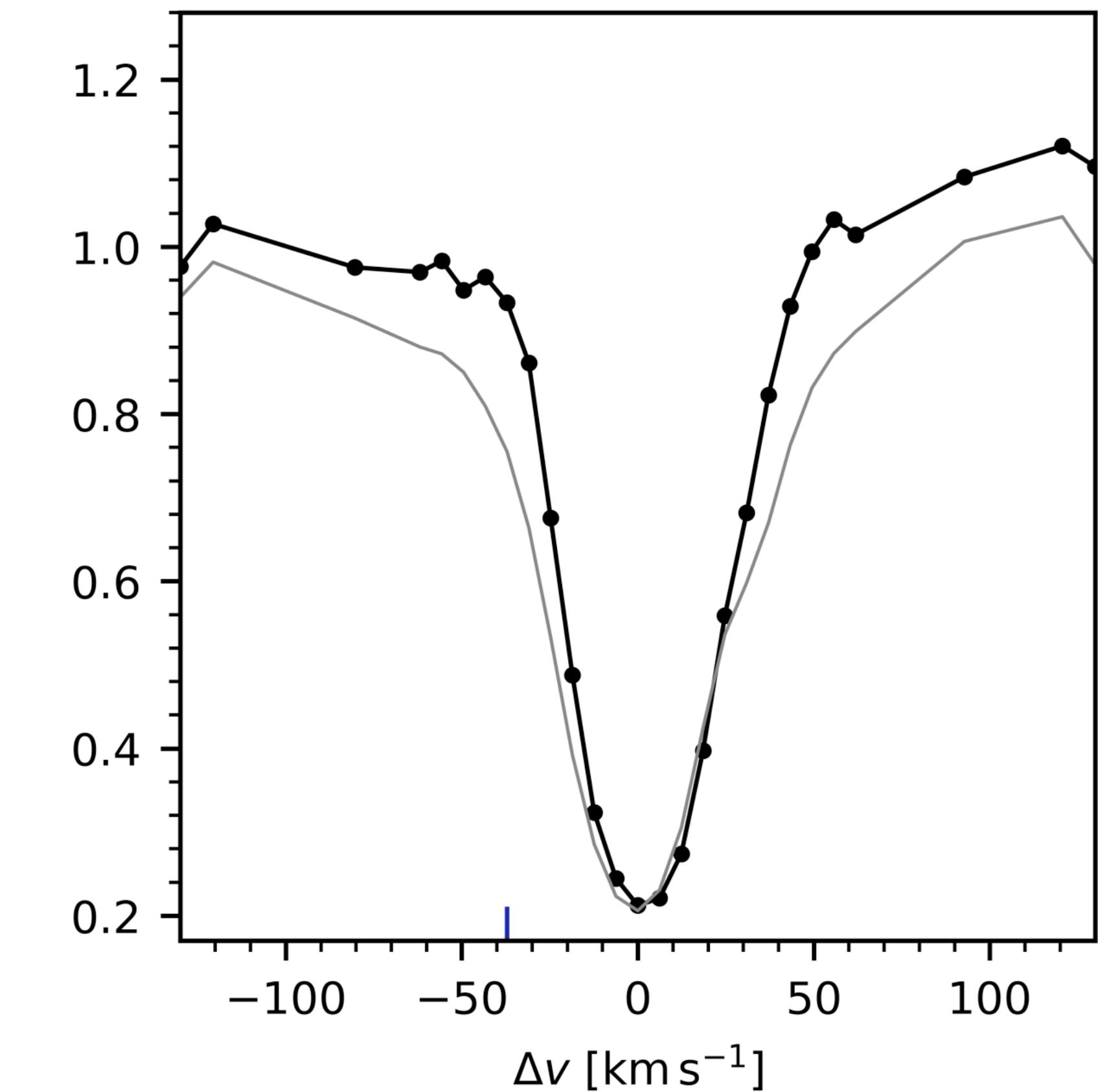
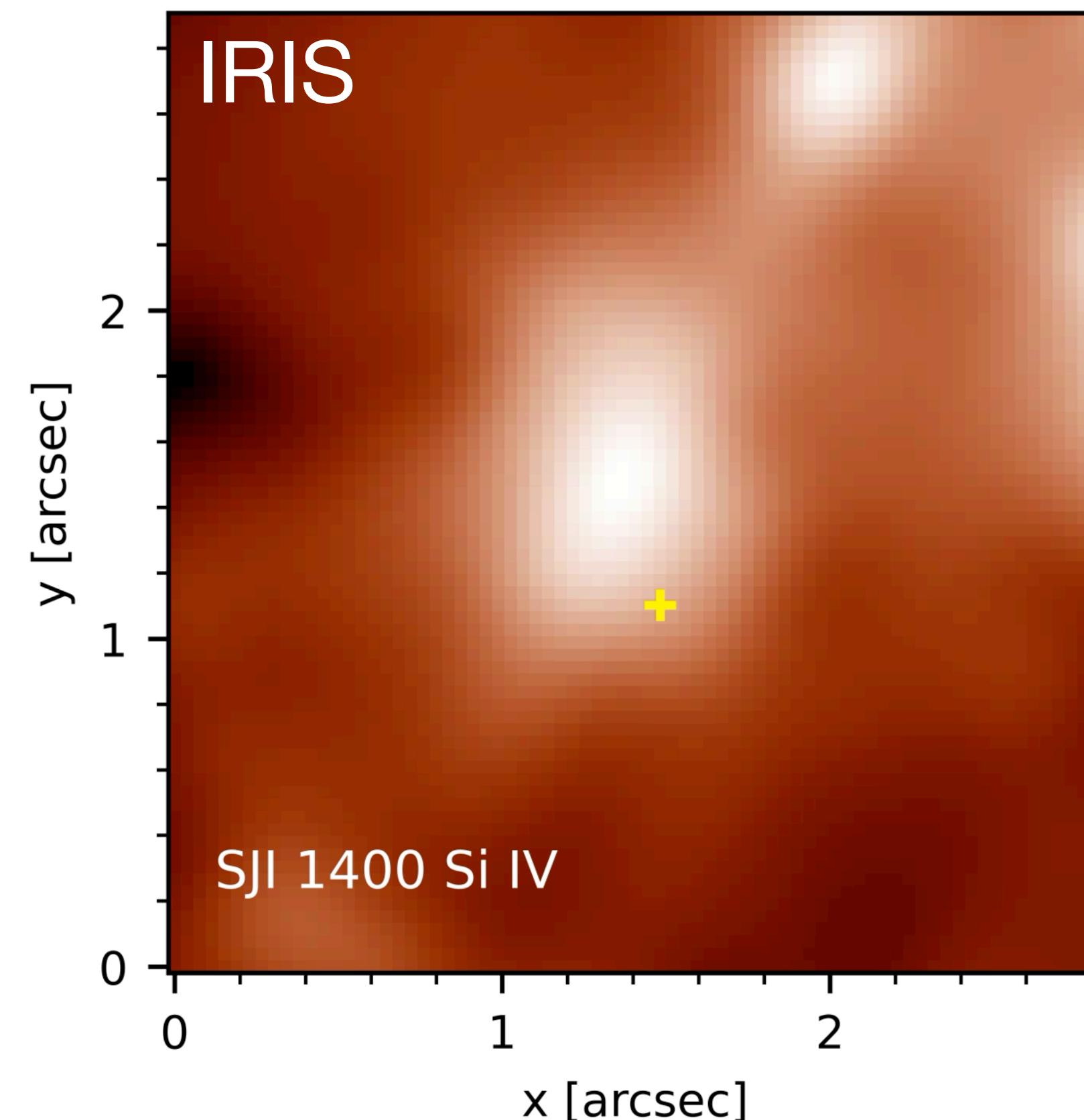
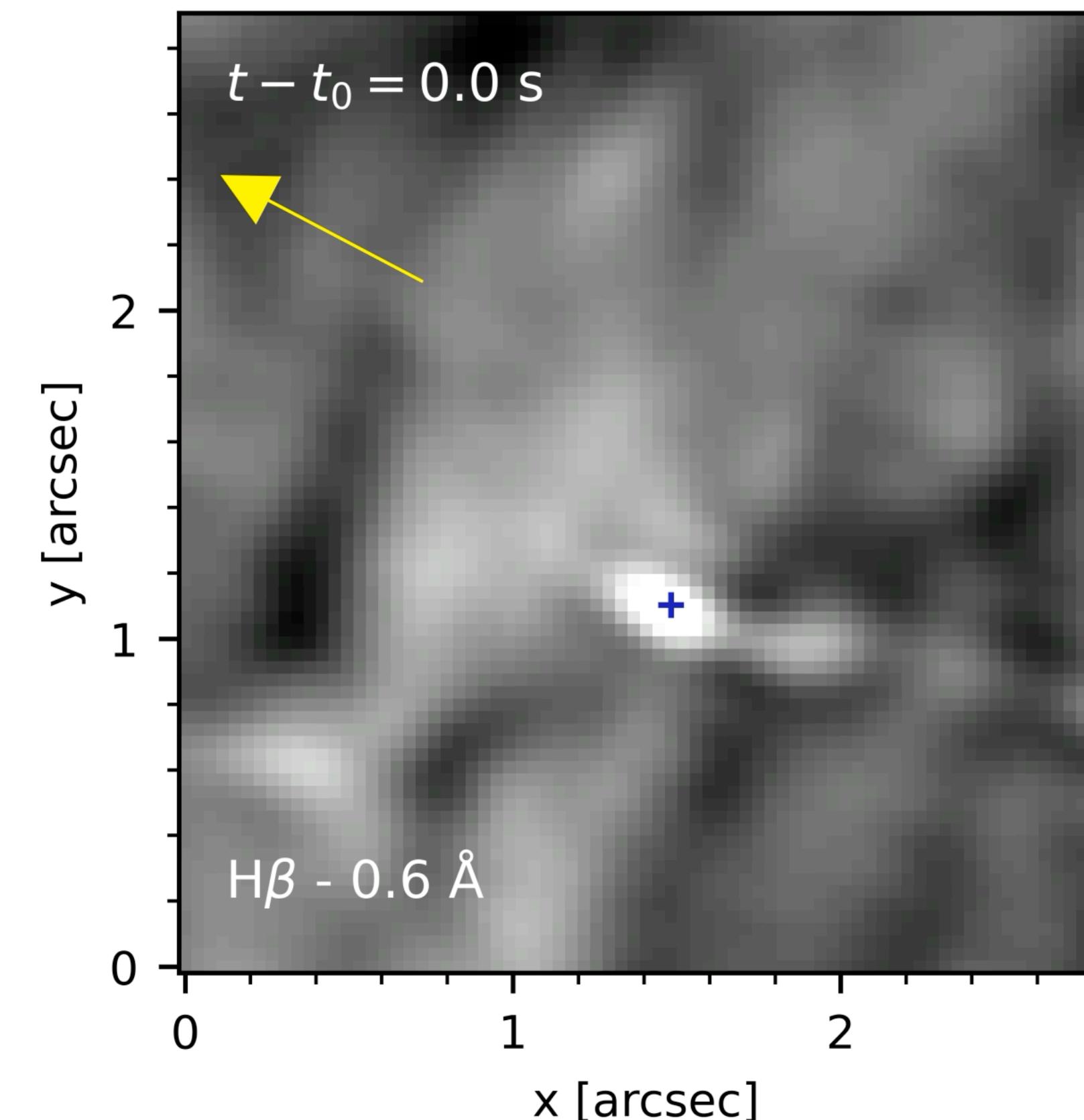


QSEB at  $\mu=0.76$

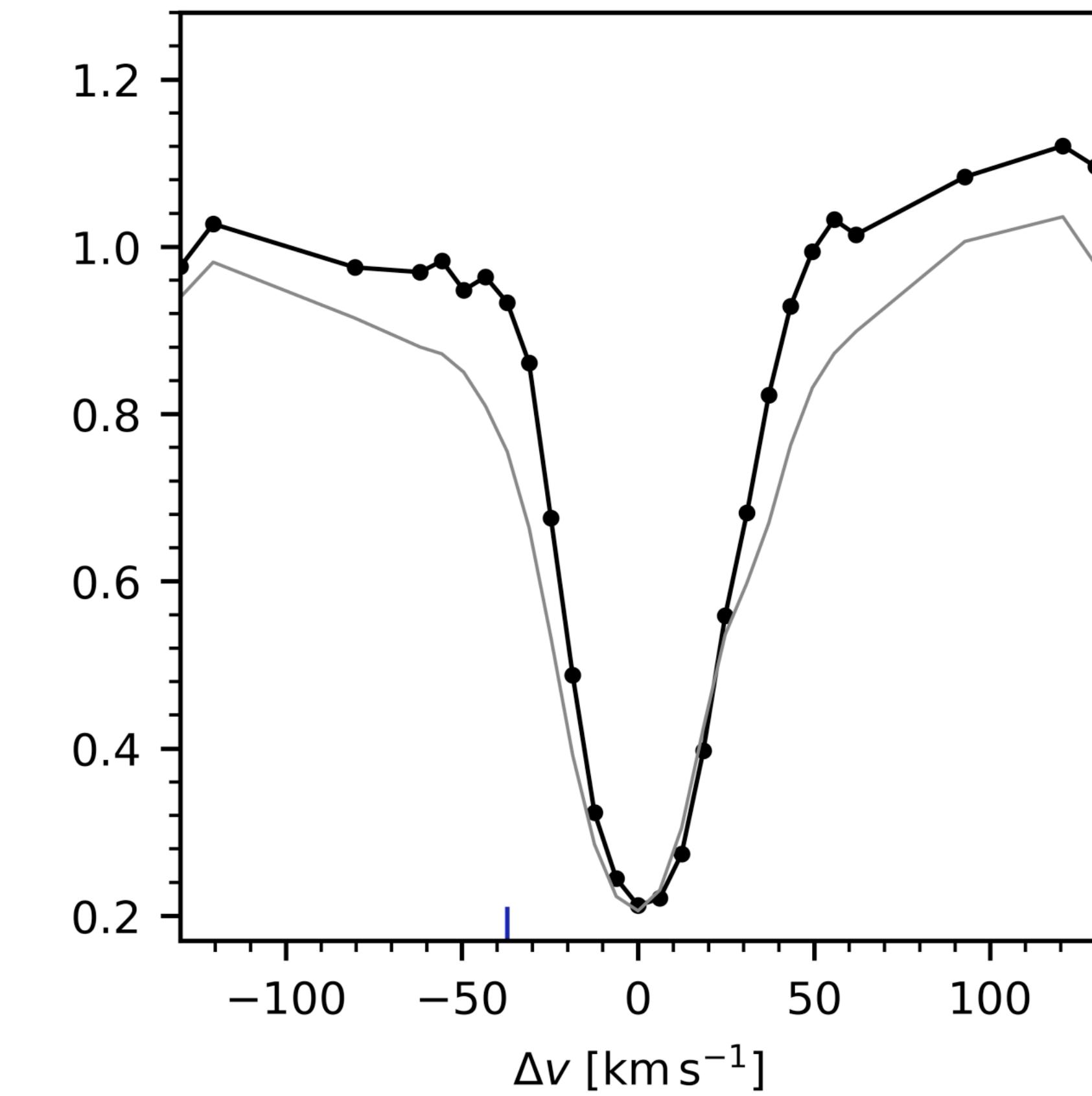
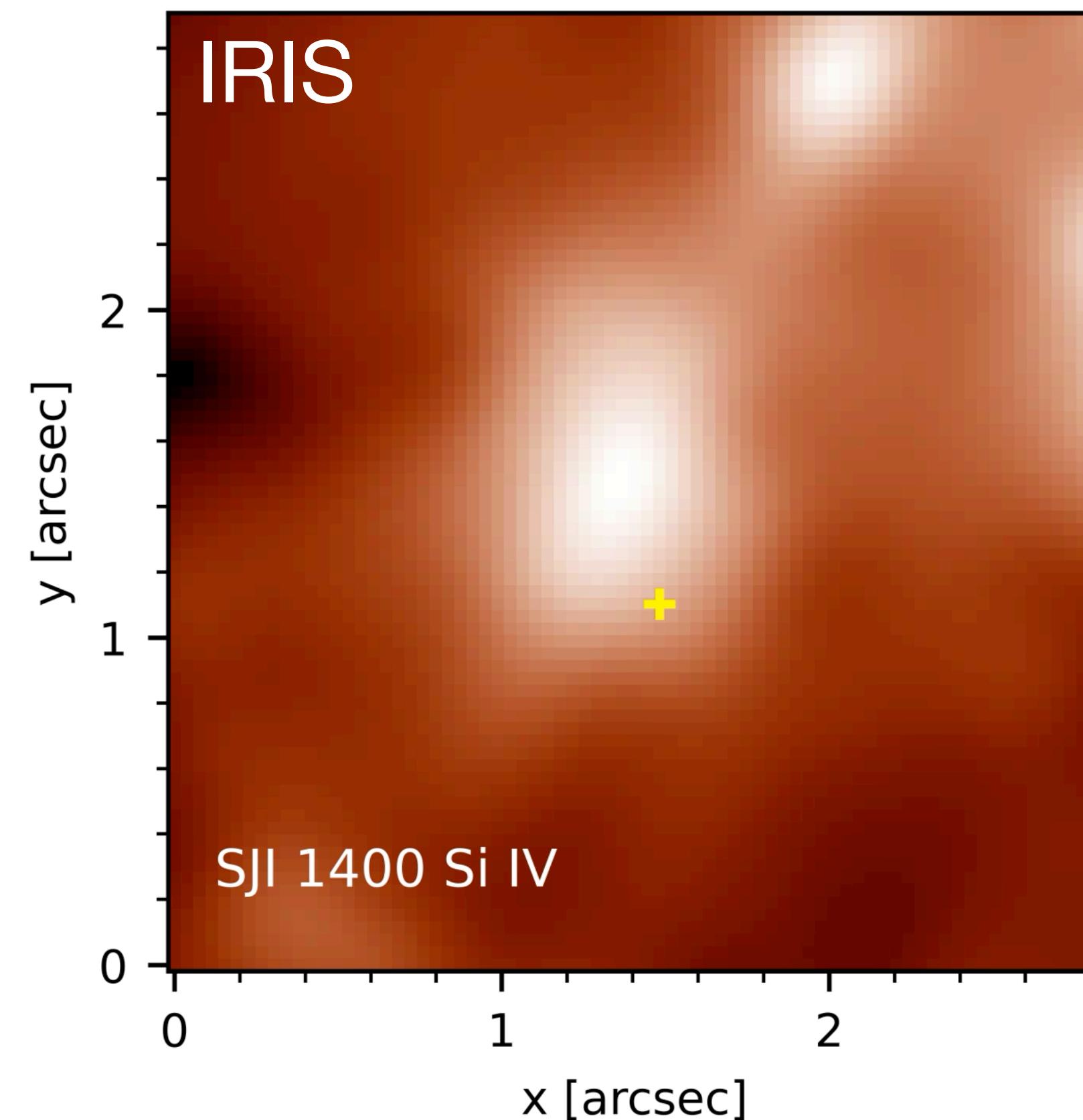
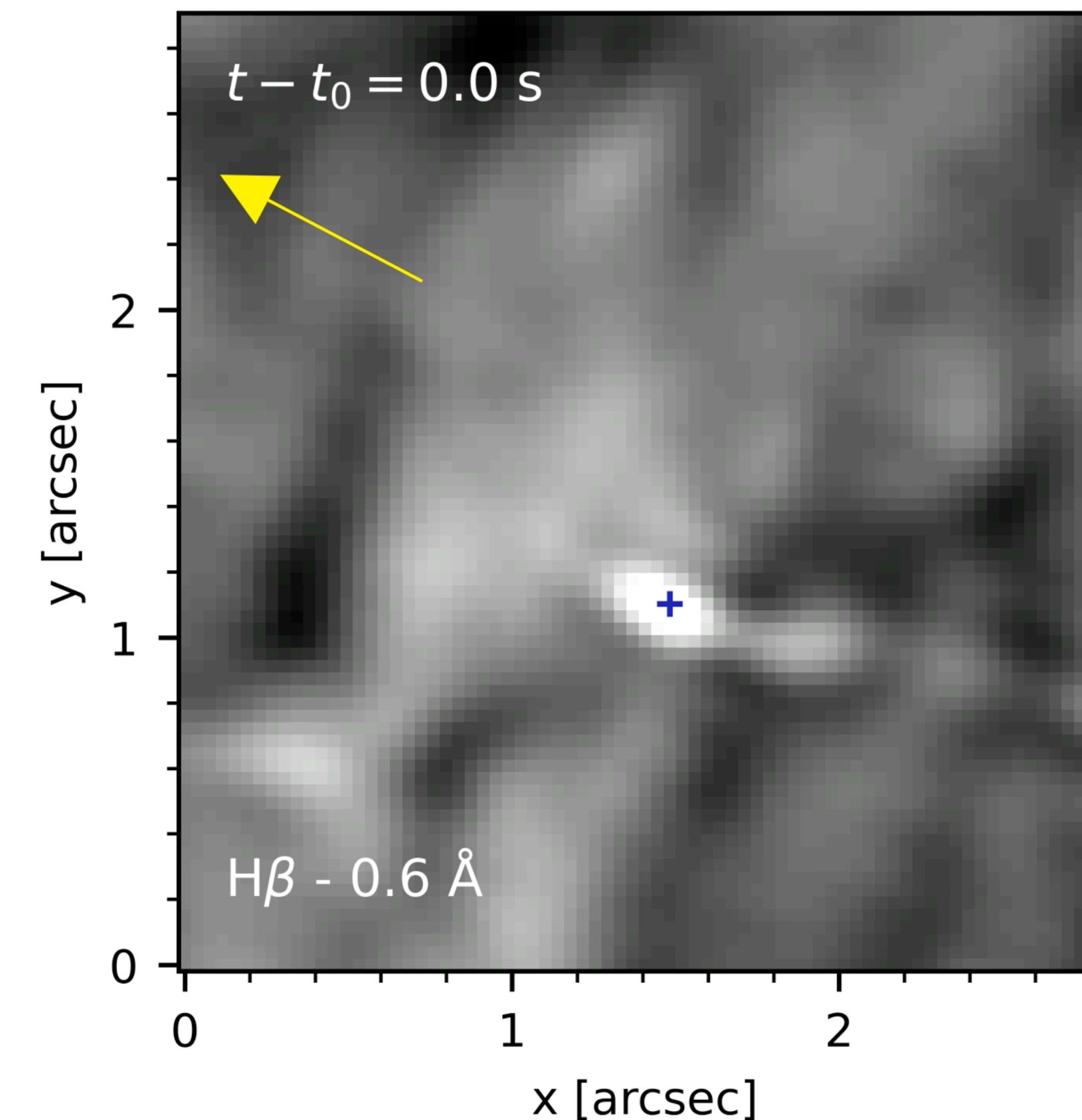
Flame, strong H $\beta$  wing



# Impact of Quiet Sun Ellerman bombs on upper atmosphere?



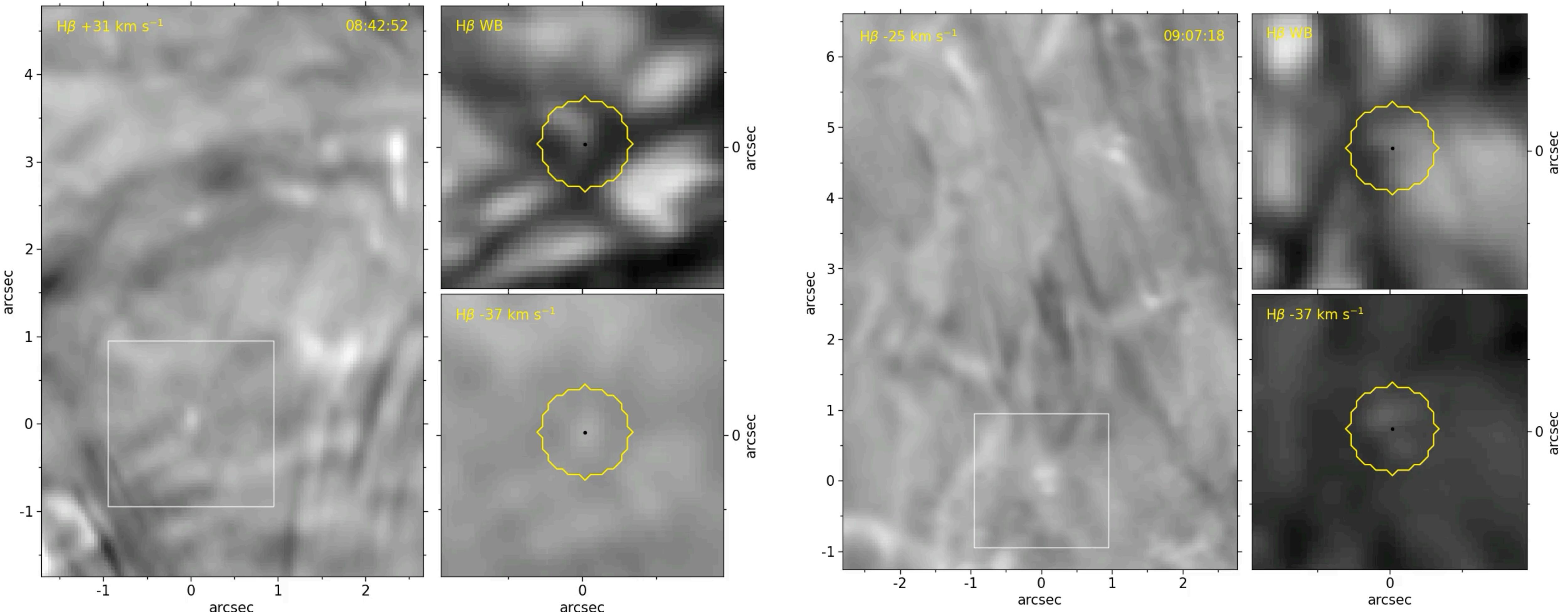
# Impact of Quiet Sun Ellerman bombs on upper atmosphere?



Longer-lived QSEBs (> 1 min):

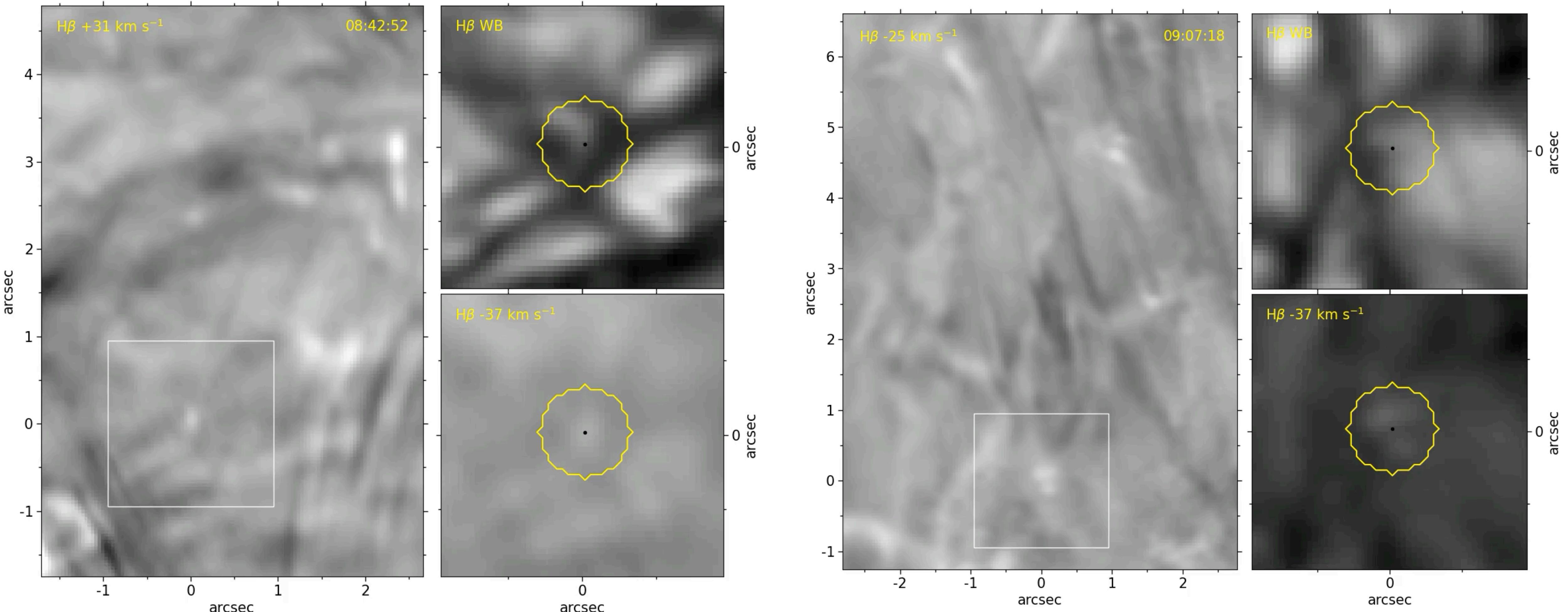
Some (15%) can be associated with brightenings in SJI 1400 and have Si IV 1394Å emission  
→ *impact on overall atmosphere is limited*

# Connection between Quiet Sun Ellerman bombs and spicules?



**Some** suggestive cases of QSEBs with spicule activity

# Connection between Quiet Sun Ellerman bombs and spicules?



**Some** suggestive cases of QSEBs with spicule activity

....but: majority of spicules have no QSEB....and majority of QSEBs have no spicules

## Conclusions

- Ellerman bombs can be found in many different environments: Active Region, sunspot penumbra, and Quiet Sun
- Ellerman bombs in Quiet Sun are ubiquitous: >750,000 at any time
- Impact on upper atmosphere is limited
  - Some have signatures in transition region diagnostics
  - Some seem to be associated with spicule activity