Dynamo of the solar interior Powering the decadal cycle and Its comparison to stellar magnetic cycles

VERSITY

OF OSLO

simulations

observations

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Solar-type stars Activity



Differential Rotation 🔭





- Not so easy to measure for solar-like stars
- Precision decreases for slow rotators

[Garcia & Ballot 2019 for a review]

Solar Dynamo: 2.5D Mean-field view



[Adapted from Noraz et al. 2022a]

[see also Parker 1955b, Babcock 1961, Steenbeck et al. 1966, Leighton 1969, Sanchez et al. 2014, review Charbonneau 2020]



Global turbulent models of Solar-like dynamo



Many different results...

Can we reach a consensus on solar-type dynamo modeling?

Normalizing Prot by τ_{conv}

[See+ 2019

see also Wright+ 2011, Reiners+ 2014, 2022]



New insights from simulations



Numerical setup:

- Code : ASH (global)
 3D MHD spherical (r,θ,φ)
- → convection is explicitly resolved, magnetic retro-action on the flow ←

MHD



15 models of solar-type [Brun et al. 2022]

- from 0.25 Ω_{\odot} to 5 Ω_{\odot}
- from 0.5 M_{\odot} to 1.1 M_{\odot}
- Resolution 769 x 256 x 512



Rotational transitions

[Brun+ 22, Noraz+ 24]



Karak+ 18, Hindman+ 20...]



Energy tranfers in dynamo solutions see Starr & Gilman 66 Brandenburg+96 Rempel 06 Brun + 22Nuclear generation **Internal energy Potential energy** Viscous Buoyancy Pressure work heating (compression & expansion) work Ohmic **Differential rotation**, convection, **Kinetic energy** heating meridional circulation ...

Lorentz force work (dynamo + mag. retro-action)

Magnetic energy

Large scale dipole, magnetic ribbons, flux-tubes,...

Powering the Differential Rotation





Powering the Dynamo [Brun+22]

	Powering
	magnetism
Slow rotator	~0.1% L★
Solar-like	~0.1-3% L★
rotator	
Fast rotator	~1% L★



1 % of the solar luminosity : ~ 4×10^{31} erg/s !

Comparison to solar observations



Observational constraints: Topology [Noraz+ 24, obs. from See et al. 2015]



Magnetochronology of old solar-type stars



- Offset of the simulations, likely because we probe deeper, and ZDI partially miss flux
- The large-scale **decreases**, agrees with observational trends,

minimum around the solar? \rightarrow stalling of the spin-down?

We need further constraints for the [see also Brandenburg & Giampapa 2018,



Coupling global codes with the Surface



[MHD Dyablo code Delorme et al. 2023] [DISPATCH Volley-ball Popovas et al. 2022]

WholeSun







WP1: Dynamo, Convection

WP2: Flux Emergence

WP3: Eruptions, Flares & Jets

WP4: Heating & Coupling of Atmosphere

WP5: Solar-Stellar Connection

WPX: Generation of Next-Gen Codes

Convective conundrum : Towards more turbulent regimes

The amplitude of giant **convection** cells in global convection **simulations** is **stronger than the one observed**. [Hanasoge+ 2012,16]

 \Rightarrow **Rossby** number achieved is **too large** \rightarrow transitions to "anti-solar"



Hotta et al. 2022 report a strong impact of the small-scales magnetism, on the large-scale convection, with R2D2 code

But No Cycle...



Non-local convective transport



Conclusion: Take-home messages

- Scenario for the Sun's life [Strugarek+ 17, Brun +22, Noraz +24] young (fast-rotating) - short cycle – no global polarity reversa solar age - prograde equator - decadal cycle, older - stationnary dynamo - anti-solar profile?
- **Possible mechanism** of the solar cycle: **Prey-predator** mechanism deep-within the convection zone,
- Differential rotation can be sustained by means of 10% of the stellar luminosity Magnetism by means of 1% : cyclic reversals + surface eruptive events,

 Such results are reproduced with other numerical methods, [Strugarek+ 17,18, EULAG code, Küker 2023, Rayleigh code] quentin.noraz@astro.uio.no @Norastraz





Perspectives

• These are only simulations:

Not yet at the solar turbulence regime, but good qualitative **agreement with observations**, Rossby trends are robust, but the exact solar one is questioned: see *Convective conundrum*, [see Hanasoge+12,16, Hotta+ 23, Warnecke+ 24]

- Need for new constraints for the interior dynamics

 > Helioseismology with Rossby modes
 [see Gizon+21, Bekki+22]
 see Yuto Bekki's talk this morning
- Need to further understand the link with surface activity (Active Regions)
 see Hannah Schunker's talk this morning (Atmospheric response)
 see Evangelia Deliporanidou's talk this morning

