



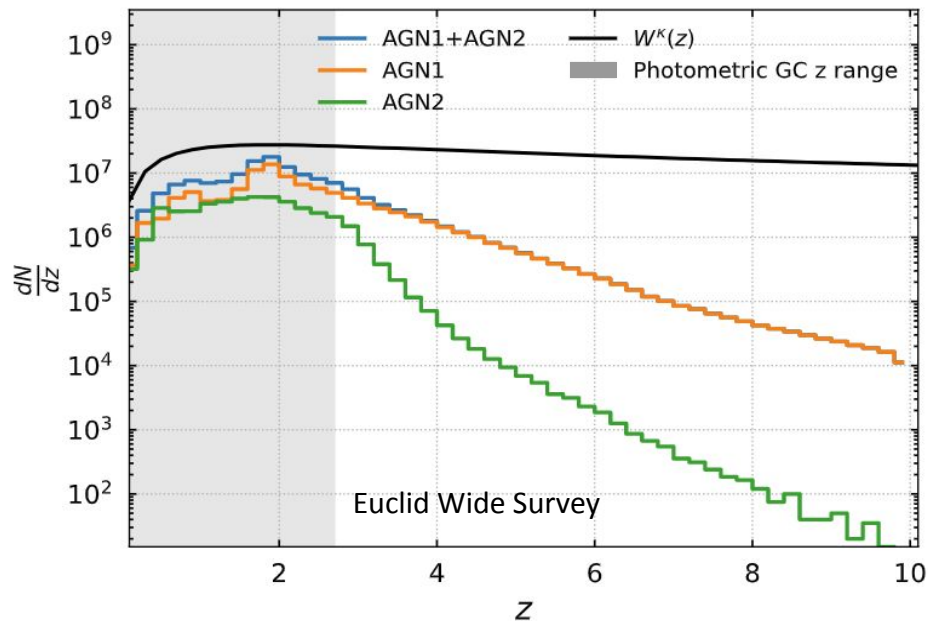
Cross-correlations with the Euclid AGN sample

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in collaboration with L.Bisigello

Euclid CMBXC-SWG Meeting
Milan, 23-24 October 2023

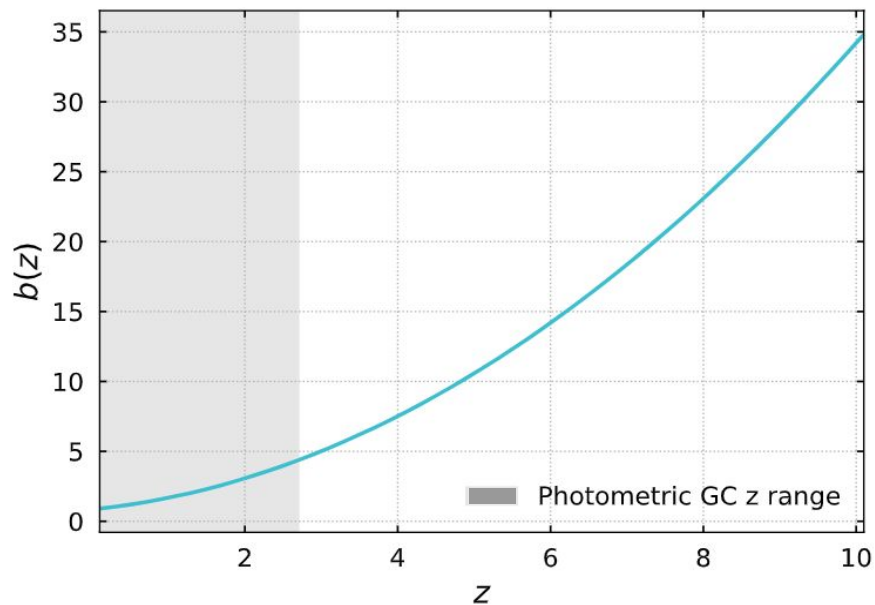
Euclid AGN $N(z)$ and $b(z)$

AGNs taken from *Bisigello+, 2023* "Euclid preparation. TBD. Selecting active galactic nuclei using observed colours"



AGN1: unobscured AGN systems ($\sim 20 \times 10^6$ objects)

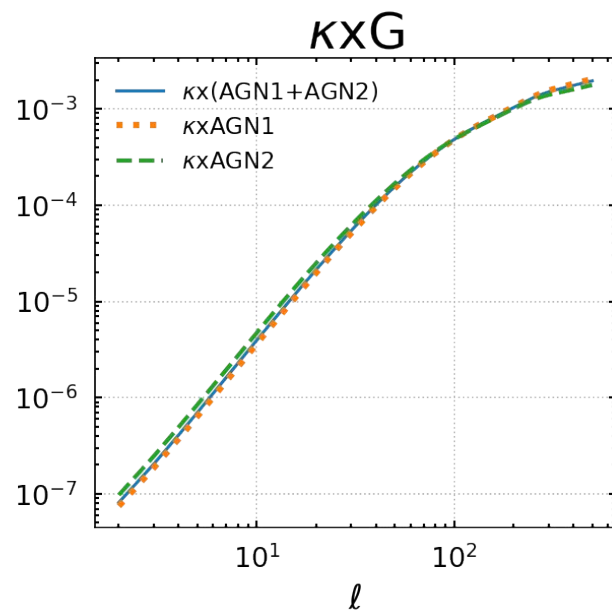
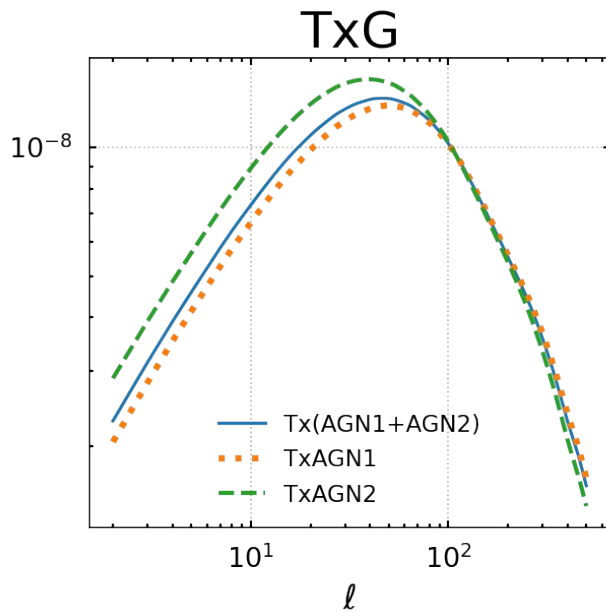
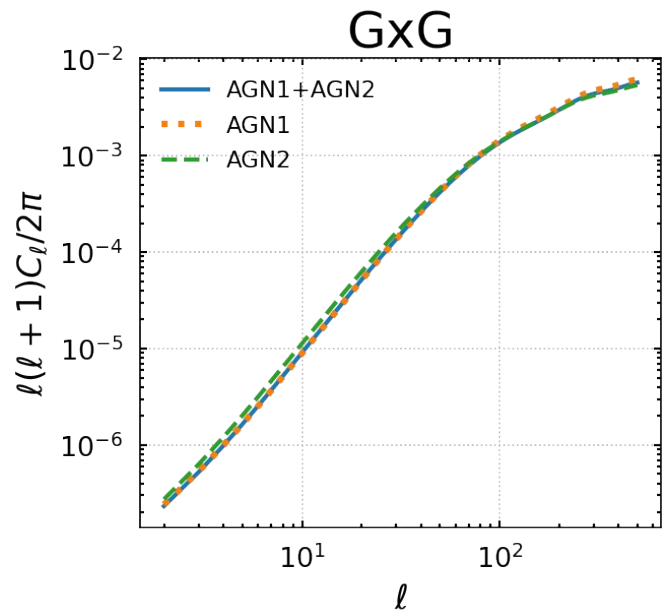
AGN2: obscured AGN systems ($\sim 8 \times 10^6$ objects)



Bias model constrained from **eBOSS** quasar auto correlation function (*Laurent+, 2017*):

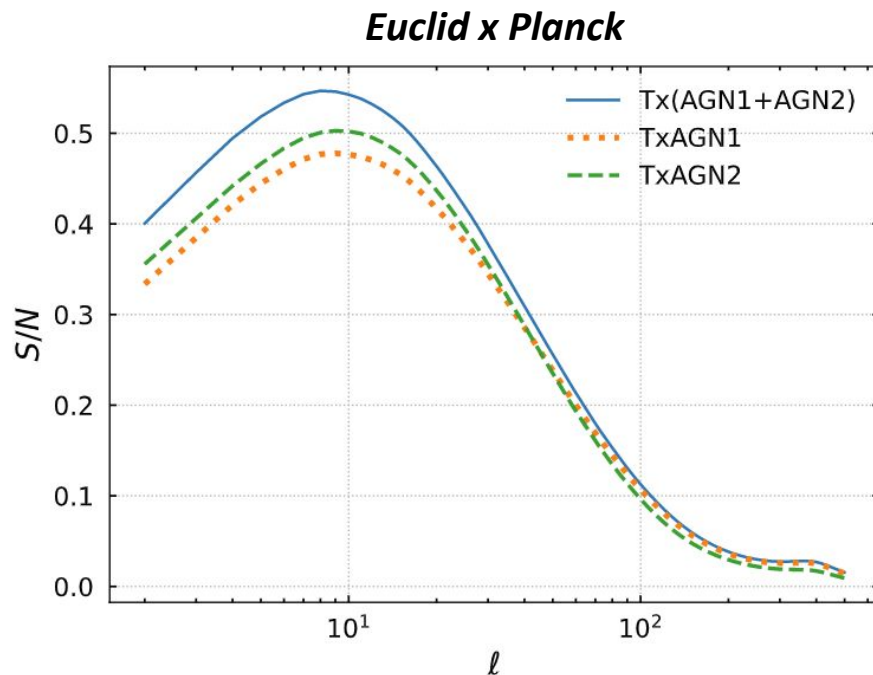
$$b(z) = [0.278((1+z)^2 - 6.565) + 2.393]$$

Euclid AGN: auto- and cross-spectra



Signal-to-noise estimation: T_g

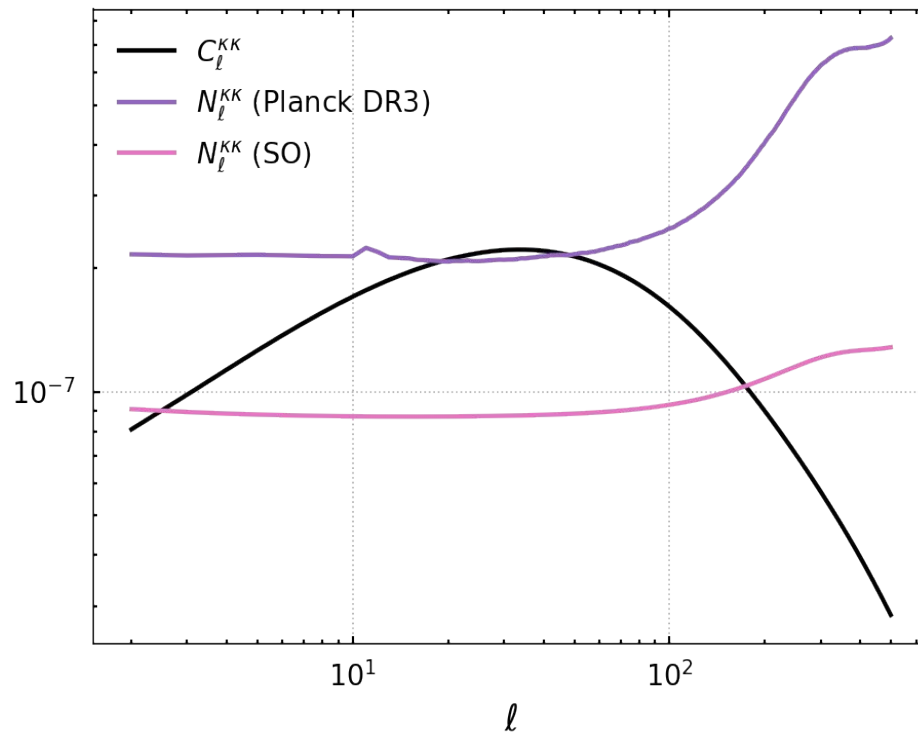
f_{sky} Euclid x Planck : 0.36



S/N total

$T_x(\text{AGN1} + \text{AGN2})$	$T_x\text{AGN1}$	$T_x\text{AGN2}$
3.28	2.97	3.00

Signal-to-noise estimation: κg



f_{sky} Euclid x Planck : **0.36**

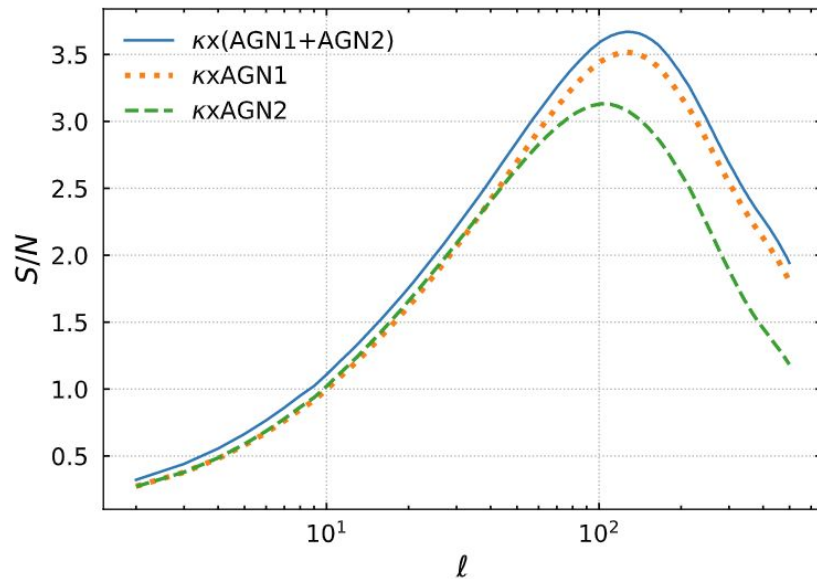
f_{sky} Euclid x SO : **0.25**

Signal-to-noise estimation: κg

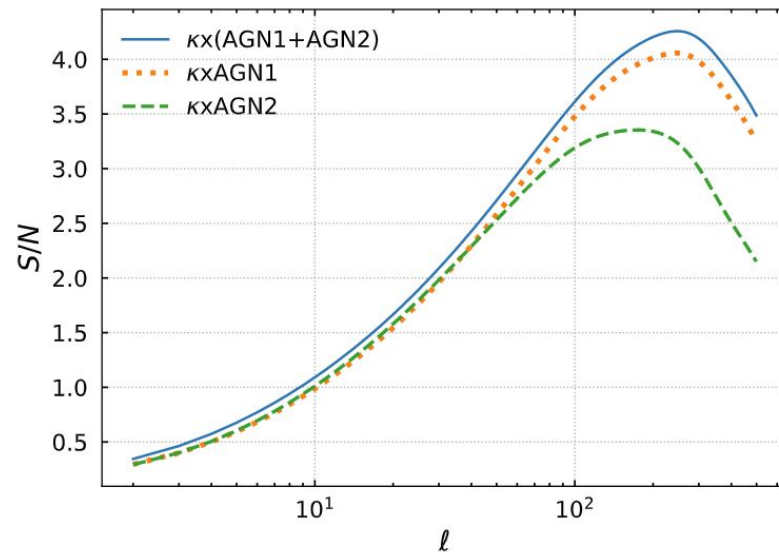
f_{sky} Euclid x Planck : **0.36**

f_{sky} Euclid x SO : **0.25**

Euclid x Planck



Euclid x SO



$\kappa x (\text{AGN1} + \text{AGN2})$	$\kappa x (\text{AGN1})$	$\kappa x (\text{AGN2})$
66.67	63.05	50.83

$\kappa x (\text{AGN1} + \text{AGN2})$	$\kappa x (\text{AGN1})$	$\kappa x (\text{AGN2})$
93.53	88.39	67.86

NB: S/N comparable to what can be obtained with Photometric galaxy clustering

Primordial Non-Gaussianity:

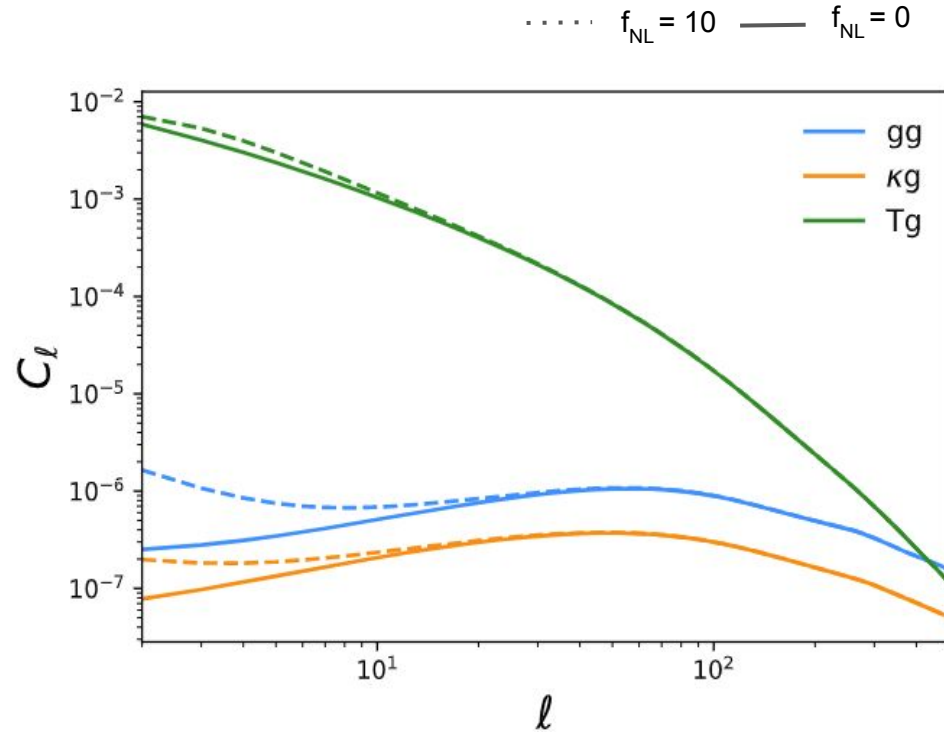
PRELIMINARY!

Local type non-Gaussian correlations between short-scale modes (which form halos) and long-scale modes in the primordial potential induce a **scale-dependent galaxy bias**:

$$\mathbf{b}_g(\mathbf{z}) \rightarrow \mathbf{b}_g(\mathbf{z}) [1 + f_{\text{NL}} \boldsymbol{\beta}(\mathbf{k}, \mathbf{z})]$$

$$\boldsymbol{\beta}(\mathbf{k}, \mathbf{z}) = 3 \frac{(\mathbf{b}_g - 1)}{\mathbf{b}_g} \frac{\Omega_{\text{m},0} \delta_{\text{c}}}{\mathbf{k}^2 \mathbf{T}(\mathbf{k}) \mathbf{D}(\mathbf{z})} \frac{\mathbf{H}_0^2}{\mathbf{c}^2}$$

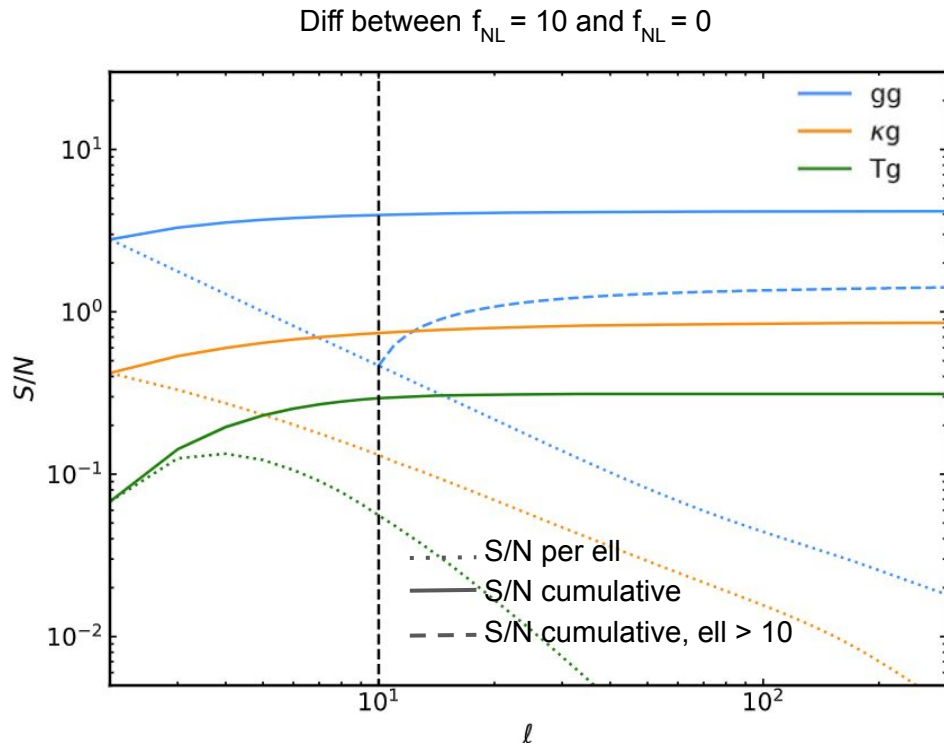
NB. in this analysis we tested only the AGN1 sample



Primordial Non-Gaussianity:

PRELIMINARY!

χ^2 analysis with only f_{NL} as free parameter: $\ell_{\text{min}}(\text{kg, tg}) = 2$, $\ell_{\text{min}}(\text{gg}) = 10$, $\ell_{\text{max}} = 500$ and lensing noise reconstruction from Planck



Outlook

- AGN represent an interesting datasets for cross-correlations
- Comparable S/N to Photometric galaxy clustering
- Complementary high-z dataset
- Competitive constraints on f_{NL} from scale-dependent galaxy bias

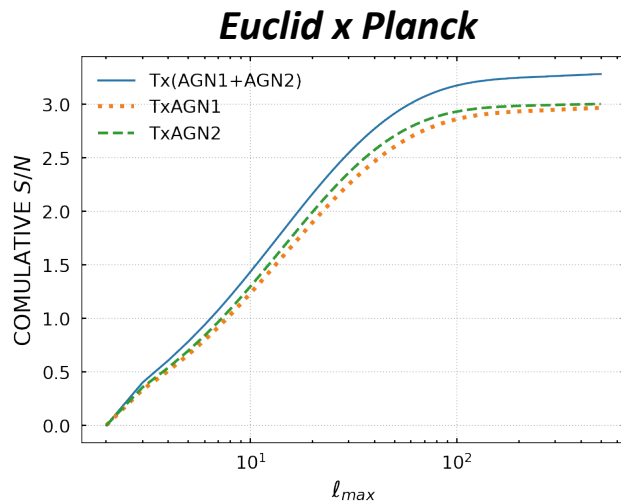
Ongoing activities:

- Improve AGN-bias recipes
- $\sigma_8(z)$ at high-z
- ISW science constraints
- ...

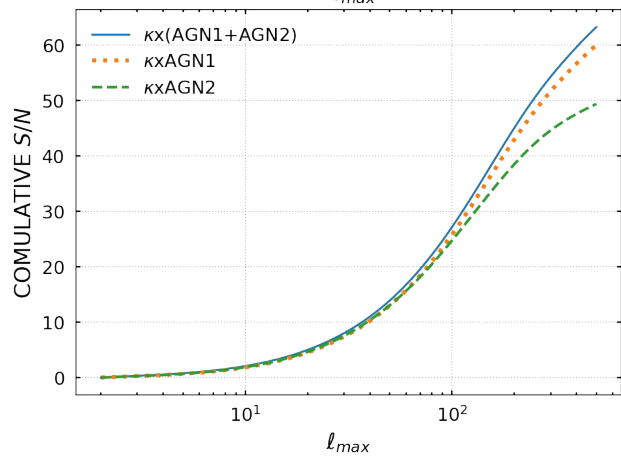
Backup

Cumulative signal-to-noise estimation

T x G



κ x G



Euclid x SO

