

Inserting L-ISW in the numerical CMBX full covariance

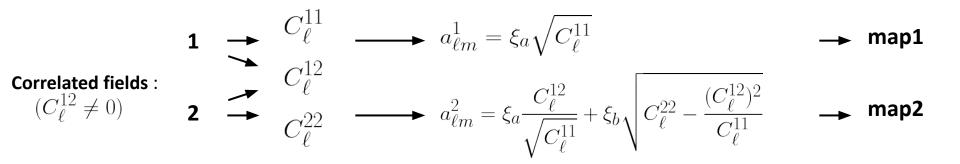
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> Euclid CMBXC-SWG Meeting Milan, 23-24 October 2023

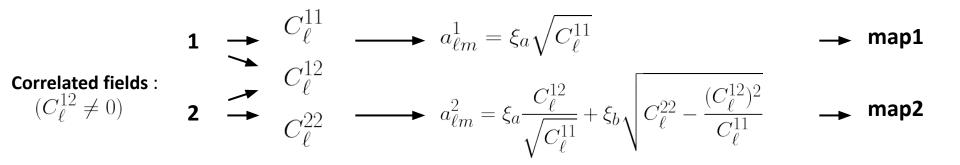
New pipeline for the numerical covariance

- Production of CMB- κ , Galaxy clustering, Weak-Lensing maps with *Flask*
- \succ Production of Late-ISW (L-ISW) maps properly correlated with CMB- κ and Galaxy clustering
- Production of T and E maps (NOT including L-ISW) with CAMB
- Addition of L-ISW maps to T and E maps (NOT including L-ISW) to obtain T and E maps properly correlated with Flask maps
- Lens T and E maps with Flask CMB-Φ
- Computation of the full numerical covariance

L-ISW maps production

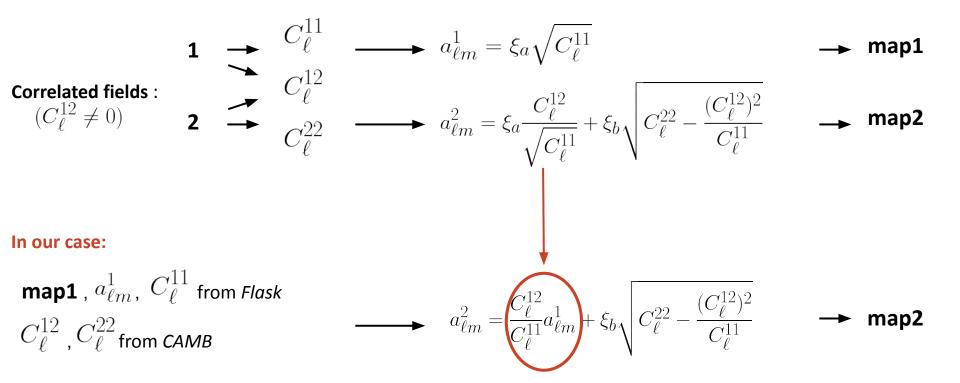


L-ISW maps production

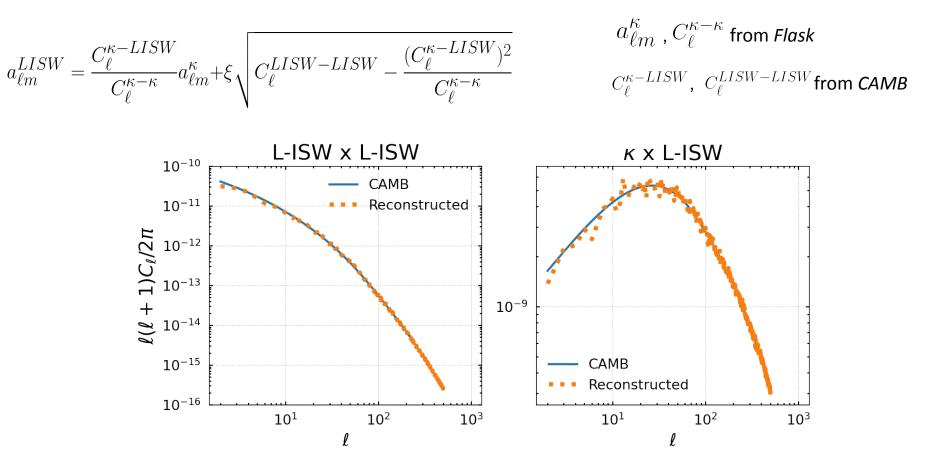


In our case:

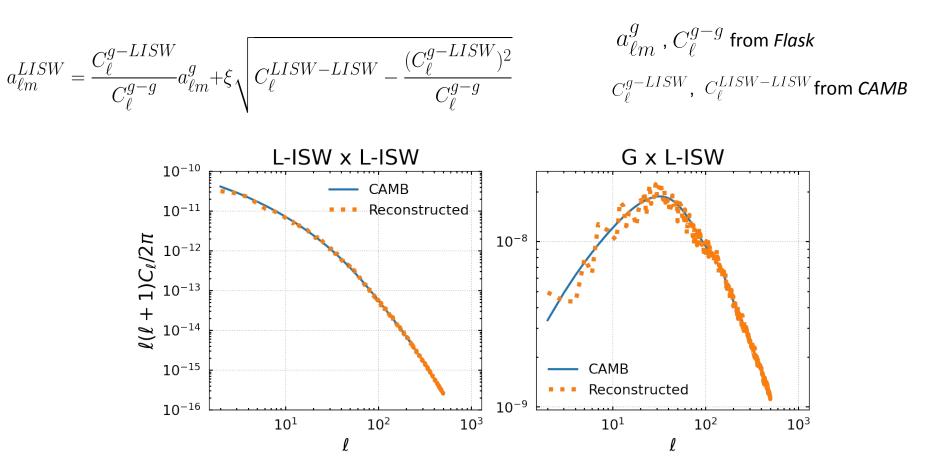
L-ISW maps production



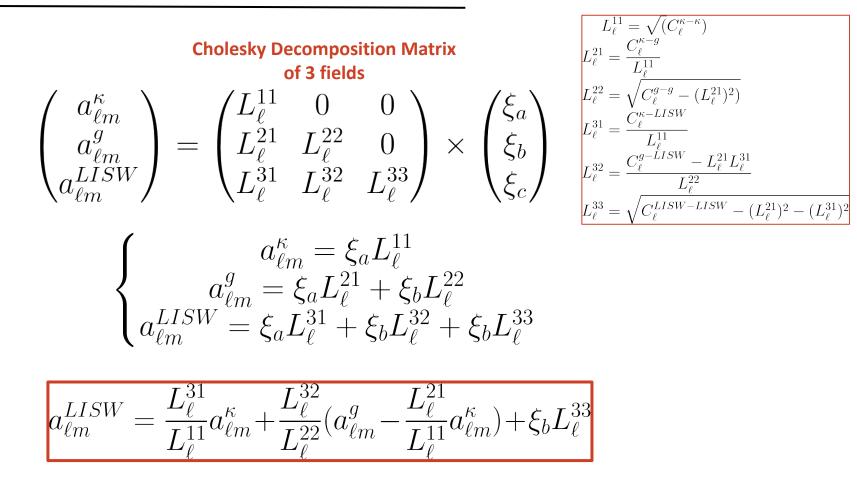
L-ISW maps from CMB-κ



L-ISW maps from non-tomographic Galaxy clustering



L-ISW from 3-fields combination



Outlook

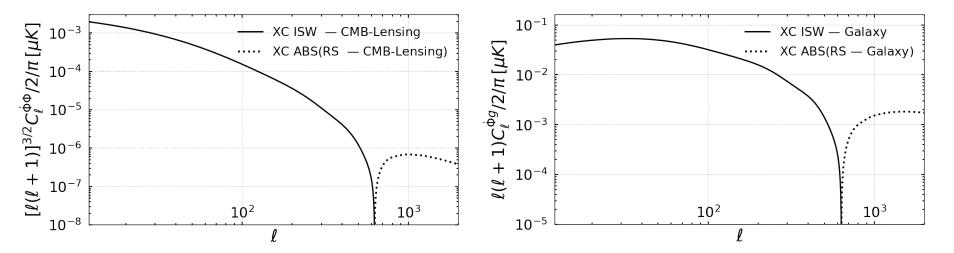
- We are including L-ISW in the numerical covariance matrix to properly consider all the cross-correlation for the joint ISW CMB-Lensing cross Euclid Likelihood analysis
- Our pipeline is working and we are improving it **combining** the information from κ and **g**

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Next steps:

• Producing L-ISW–RS maps, producing C_{ℓ}^{12} , C_{ℓ}^{22} analytically (see Cuozzo+, 2023)



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Next steps:

- Producing L-ISW-RS maps, producing C_{ℓ}^{12} , C_{ℓ}^{22} analytically (see *Cuozzo+*, 2023)
- Reconstruct the L-ISW map from the combination of <u>all</u> the *Flask* maps (N Galaxy bins + CMB- κ)

$$a_{\ell m}^{x} = \sum_{k=1}^{x} \xi_{k} L_{kx} \quad \text{with } \mathbf{x} = \mathbf{1}, \dots \mathbf{N} \quad \longrightarrow \quad L_{kx} = \sqrt{\frac{C_{\ell}^{xk} - \sum_{y=1}^{k-1} (L_{ky})^{2}}{L_{kx}}} \quad \text{if } k = \mathbf{x}$$

$$L_{kx} = \frac{C_{\ell}^{xk} - \sum_{y=1}^{k-1} L_{ky} L_{xy}}{L_{xx}} \quad \text{if } k > \mathbf{x}$$