

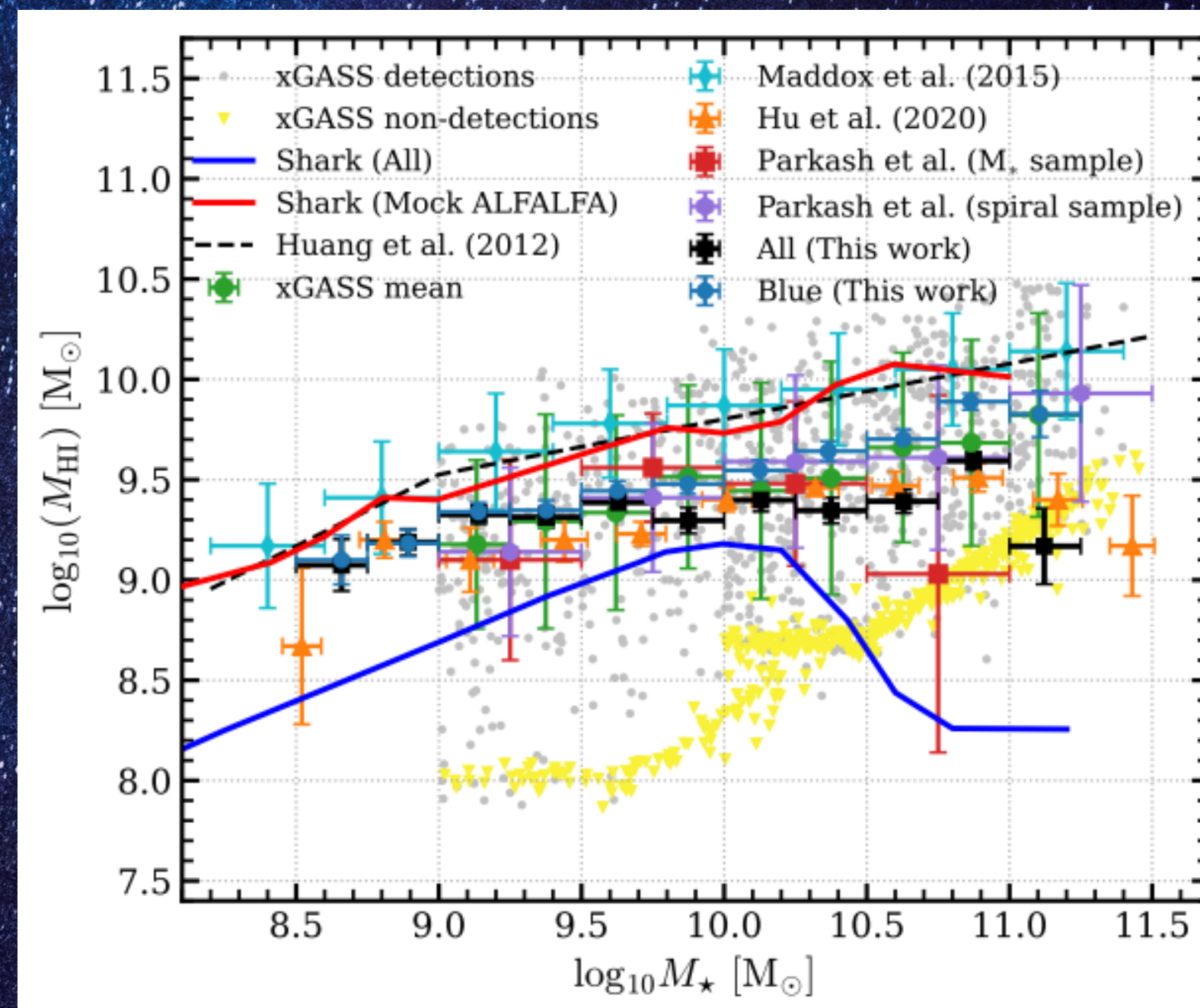
MIGHTEE and CHILES Team Up: new constraints on the evolution of the MHI- M^* scaling relation

ALESSANDRO
BIANCHETTI

SUPERVISORS:
GIULIA RODIGHIERO, ED ELSON

INAF
ISTITUTO NAZIONALE
DI ASTROFISICA

with FRANCESCO
SINIGAGLIA, MATTIA
VACCARI, DJ PISANO
MIGHTEE & CHILES
collaborations

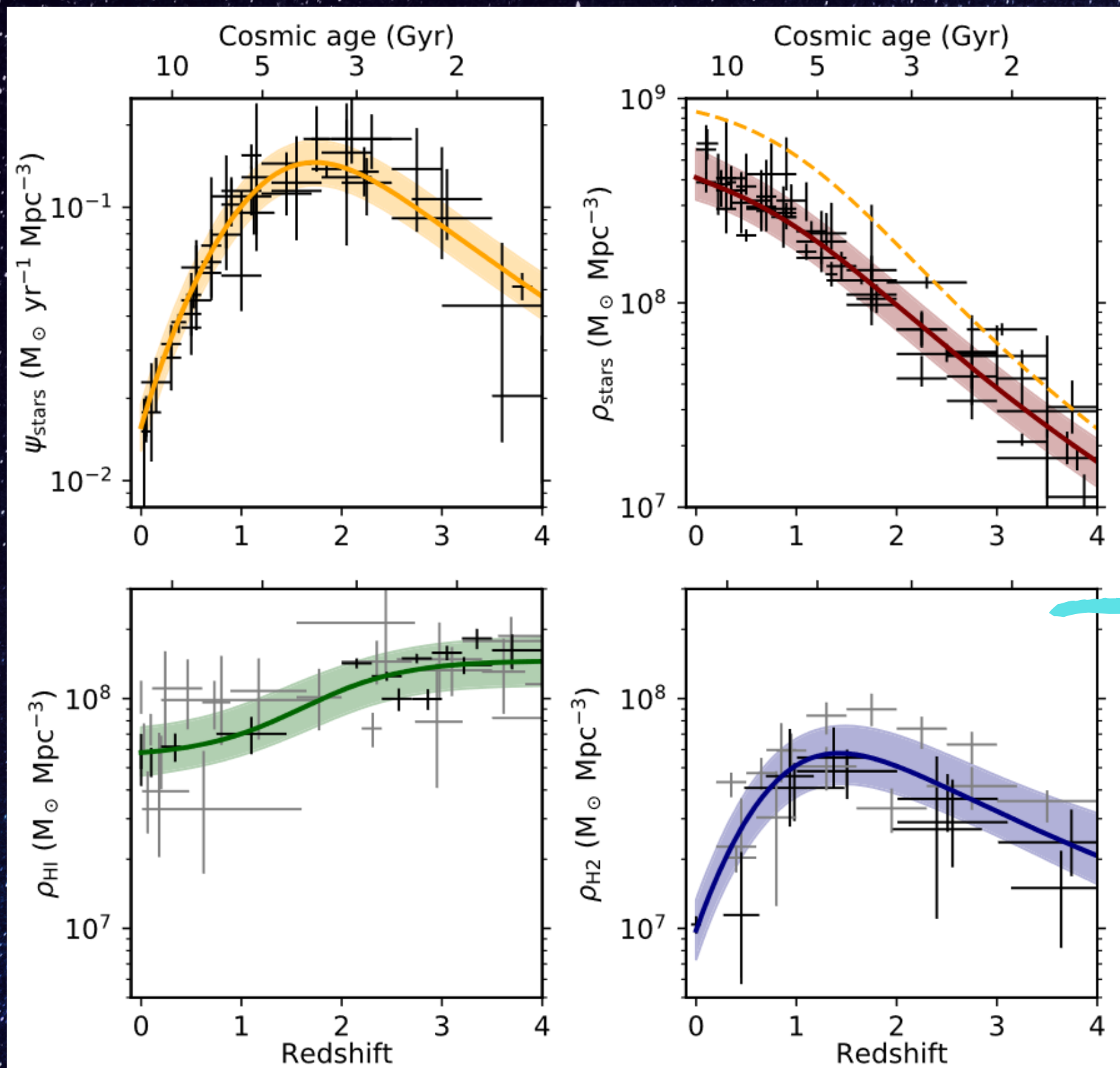


Rhee+22



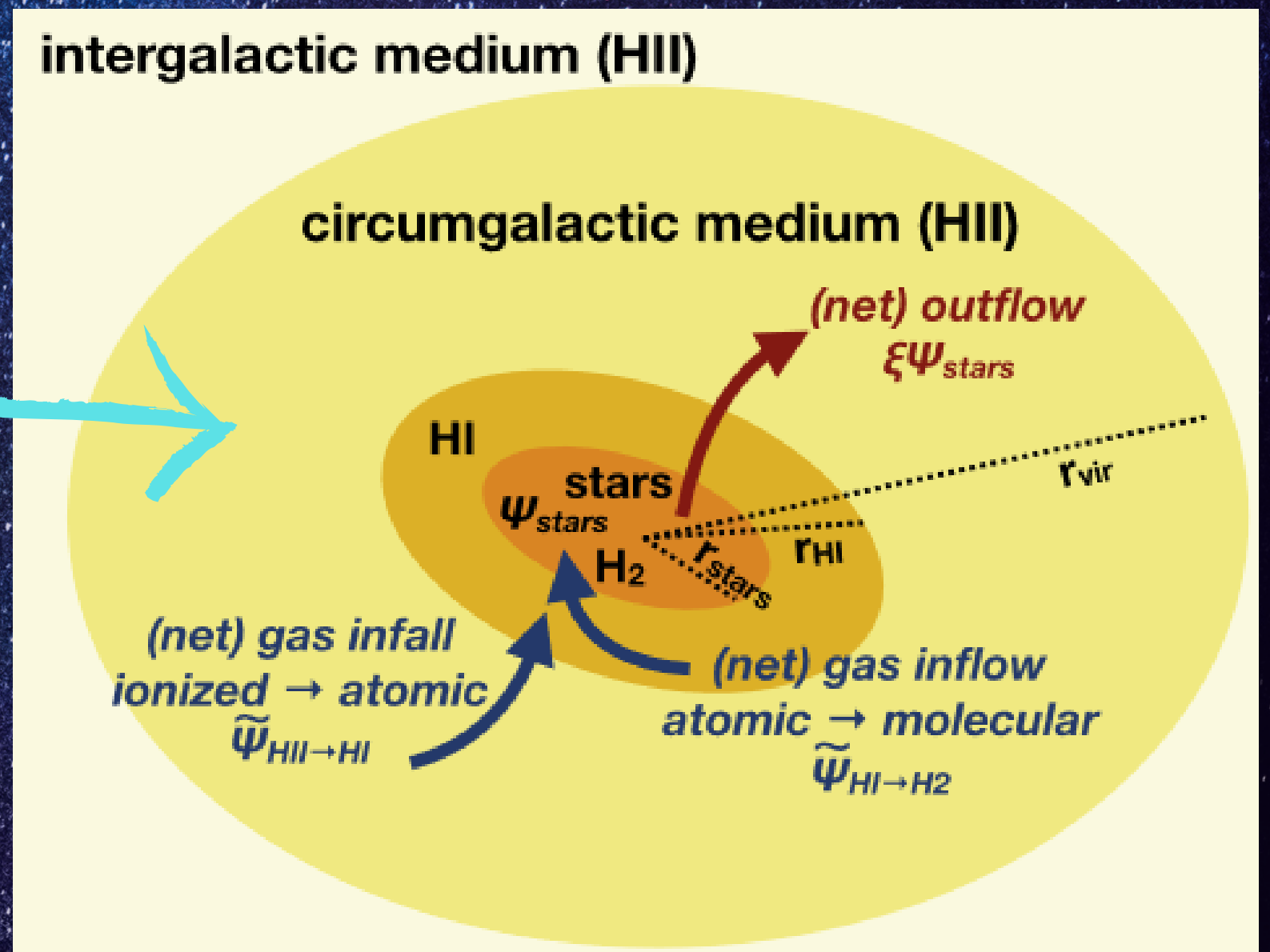
UNIVERSITY of the
WESTERN CAPE

WHY HI?



Walter et al, 2020

- peak of SFR at $z \sim 1-2$
- no evolution of HI mass?
- H₂ density follows the SFR profile
- stellar mass density $>$ total gas density after $z \sim 1.5$

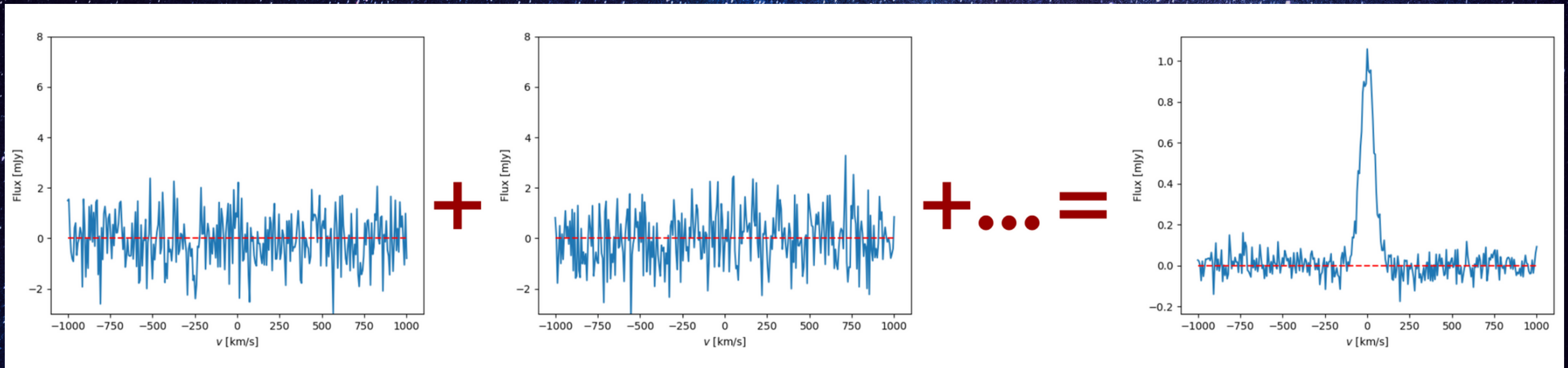


Walter et al, 2020

STACKING

Long term goal: using **stacking** to study the evolution of the scaling relation: **galaxy evolution** - **baryon cycle** - understanding SF efficiency connection with HI budget evolution

At $z > 0.1$ direct detection of the HI emission line is hardly possible: use spectral stacking



HI SURVEYS

Long term goal: using **stacking** to study the evolution of the scaling relation: **galaxy evolution** - **baryon cycle** - understanding SF efficiency connection with HI budget evolution

MIGHTEE HI

(Early Science Data)

3x16 hrs

- $0 < z < 0.58$
- 900-1420 MHz
- beam ~12" (60 kpc at $z=0.4$)
- spectral resolution 200 kHz
- COSMOS, XMM-LSS, ECDFS

[HTTPS://WWW.MIGHTEESURVEY.ORG/](https://www.mighteesurvey.org/)
MADDOX ET AL 2021

CHILES

(VLA)

1000 hrs

- $0 < z < 0.5$
- 946-1420 MHz
- beam ~6" (30 kpc at $z=0.4$)
- spectral resolution 125 kHz
- COSMOS subregion

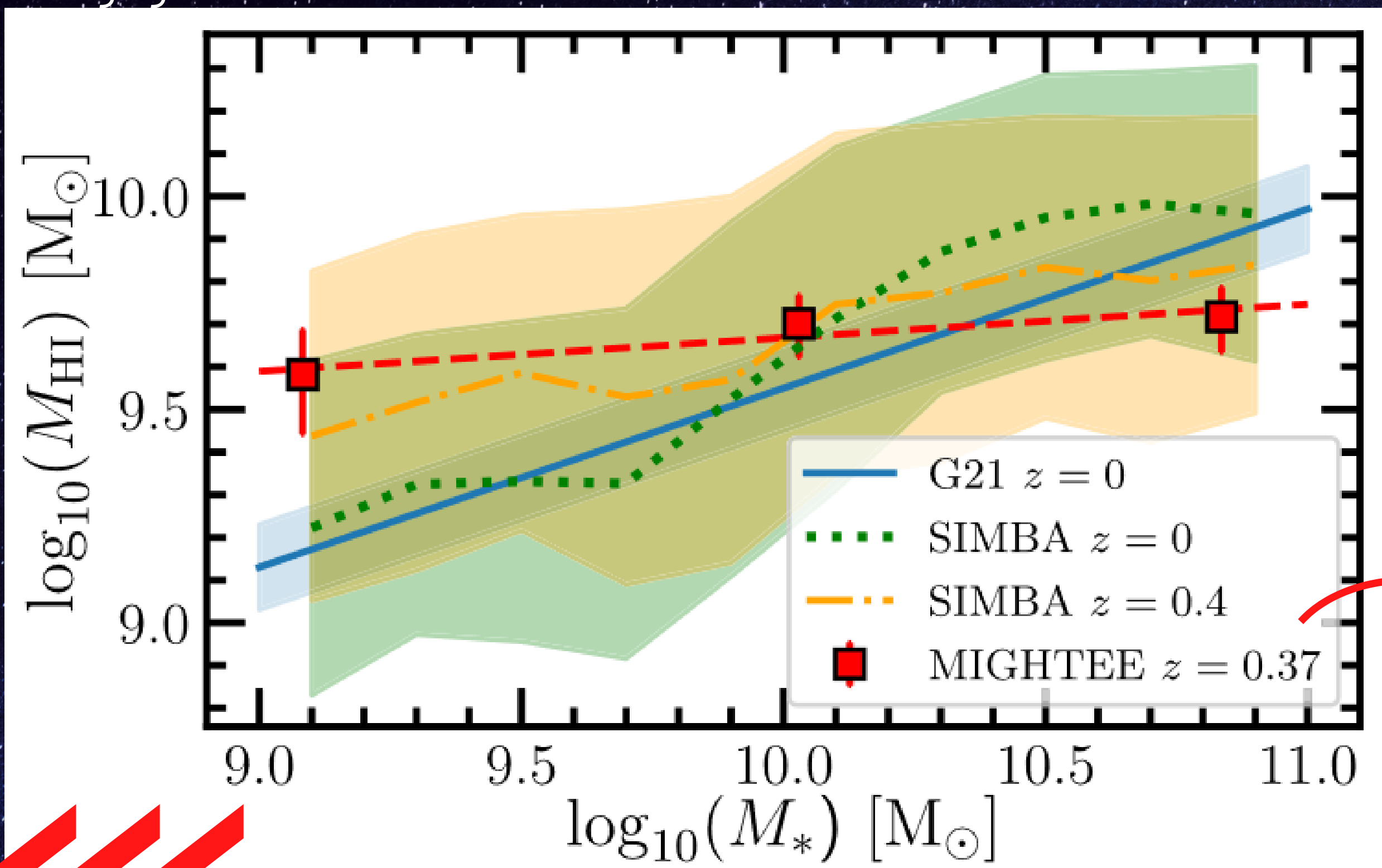
[HTTP://CHILES.ASTRO.COLUMBIA.EDU](http://chiles.astro.columbia.edu)

SUPERMIGHTEE

(MeerKat+uGMRT)

- up to $z \sim 1$
- 250-2500 MHz
- in agenda (2024)

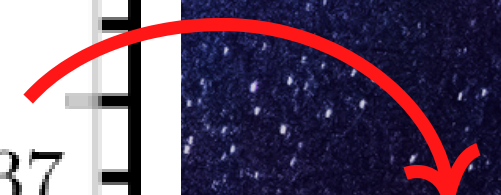
Sinigaglia et al. 2022: what can we work on?



increased statistics to make a finer stellar mass grid
CHILES



can we go to higher redshift?
SUPERMIGHTEE



HI SURVEYS

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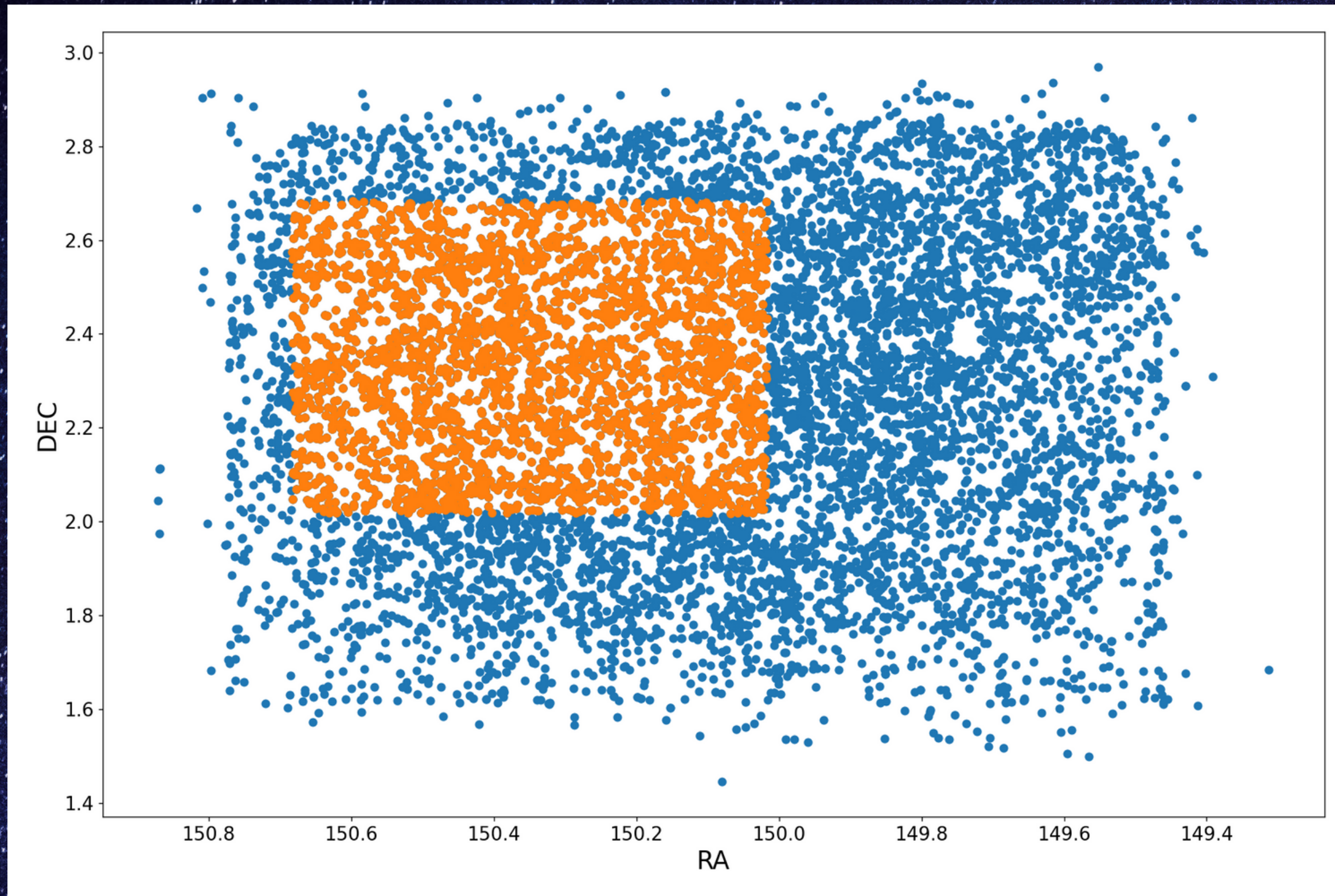
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INPUT CATALOGUE

- spectroscopic redshift catalogue
[UPDATE JUNE 2023, Khostovan et al, in prep.] + DEVILS + DESI
- photometry: COSMOS2015 -> 2020

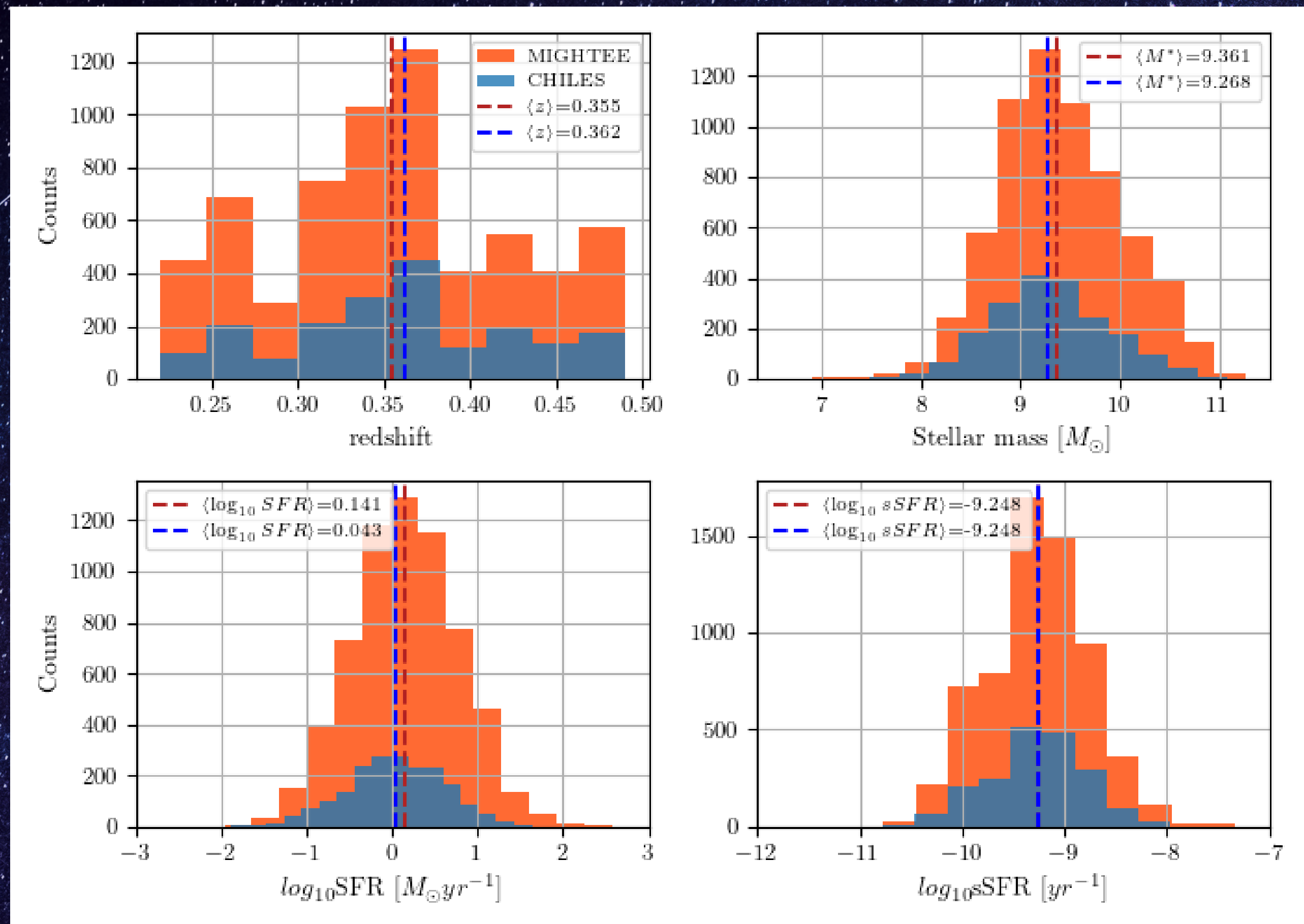
Selection criteria:

- removed photometric outliers
 $(\text{specz} - \text{photoz}) < 0.15 * (1 + \text{specz})$
- $0.22 < \text{specz} < 0.5$
- star-forming galaxies



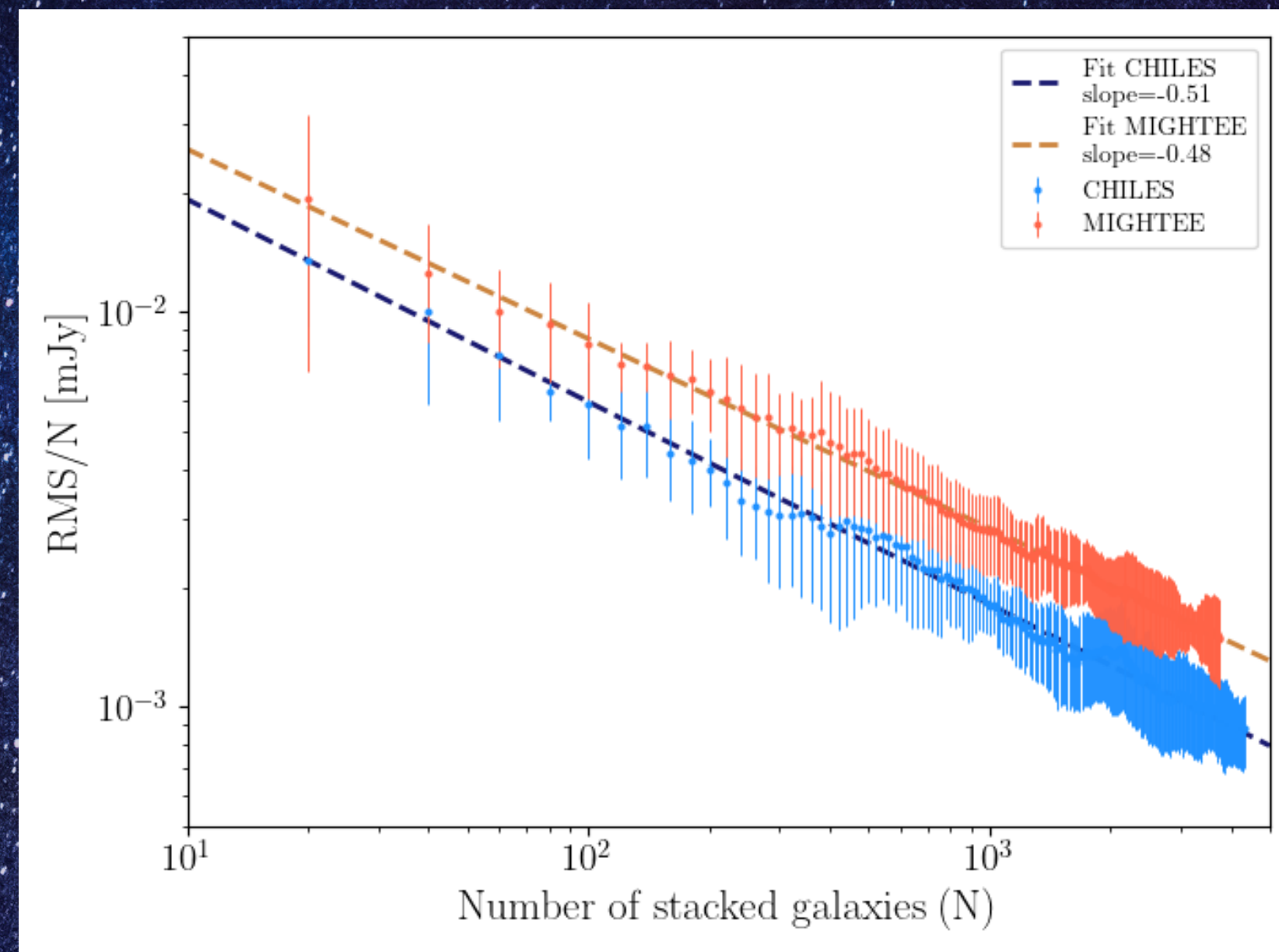
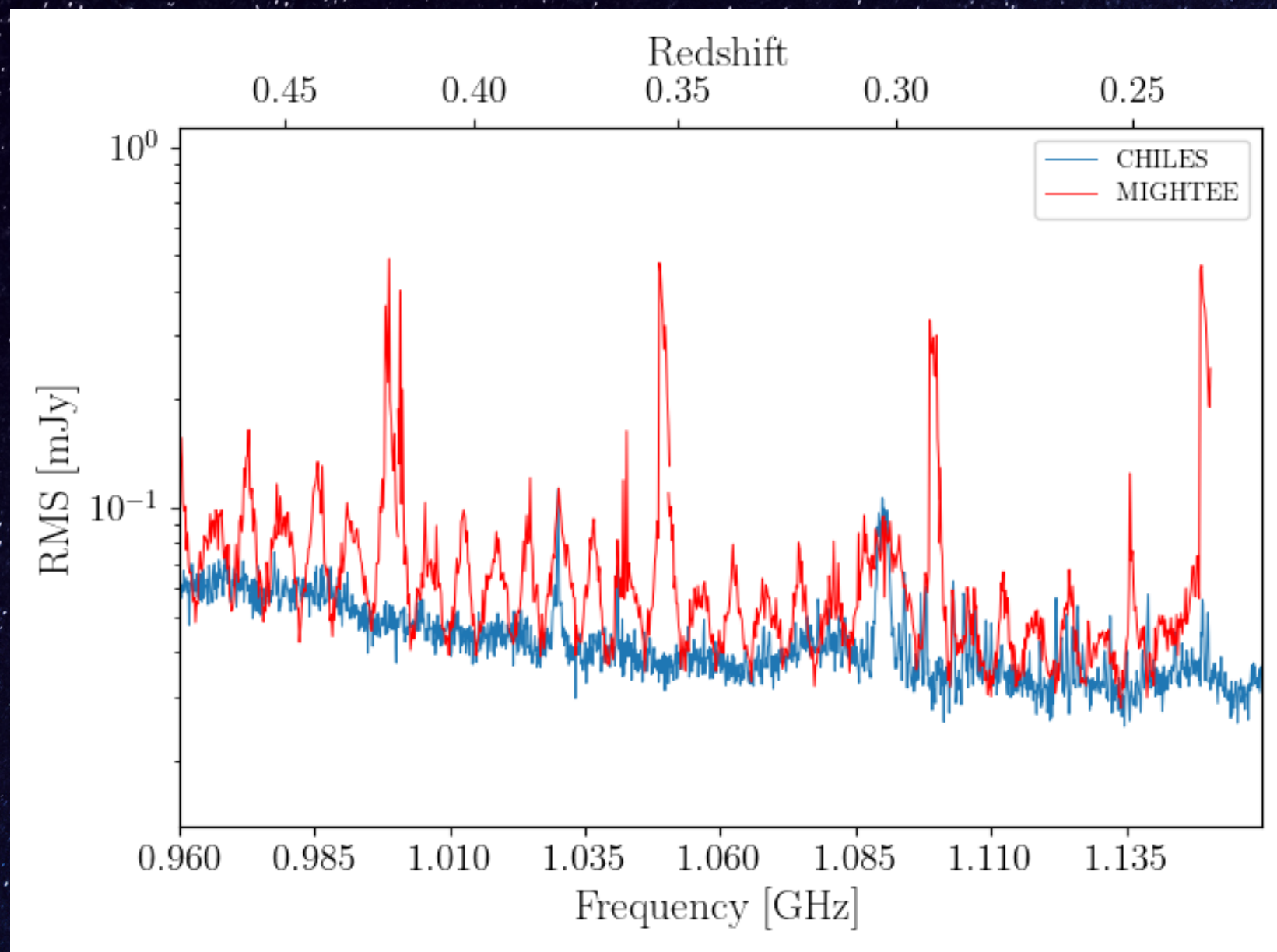
INPUT CATALOGUE CHILES VS MIGHTEE

Are we drawing from the same galaxy population? Is there a bias in the physical properties of galaxies between CHILES and MIGHTEE?



CHILES VS MIGHTEE

SANITY CHECKS AND NOISE DIAGNOSTICS



CHILES and MIGHTEE noise patterns match, **but**:

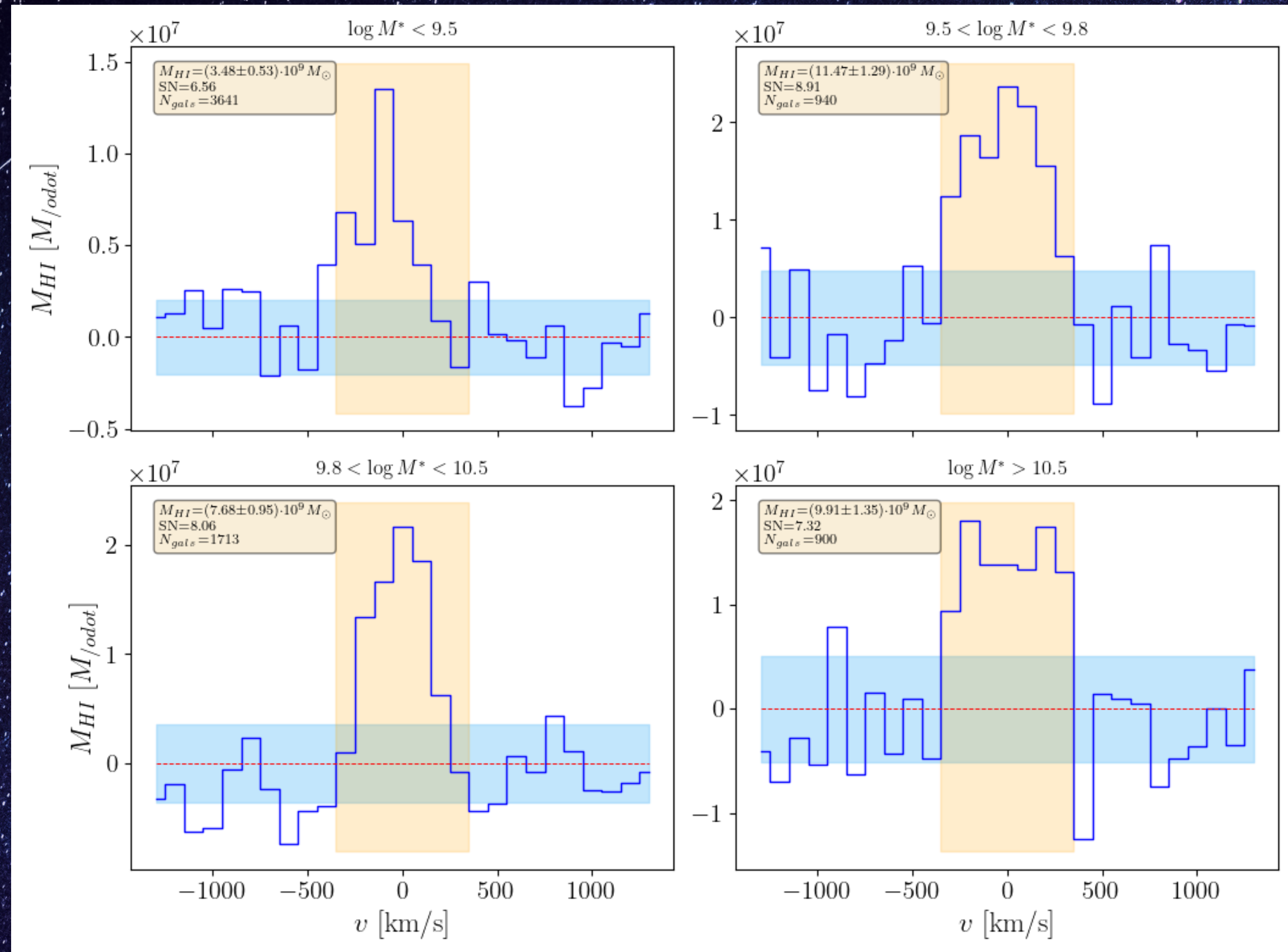
- obs_time $\sim 10^1$ hrs for MIGHTEE
- obs_time $\sim 10^3$ hrs for CHILES

$$\text{RMS} \sim \sqrt{N}$$

$$\text{RMS}/N \sim 1/\sqrt{N}$$

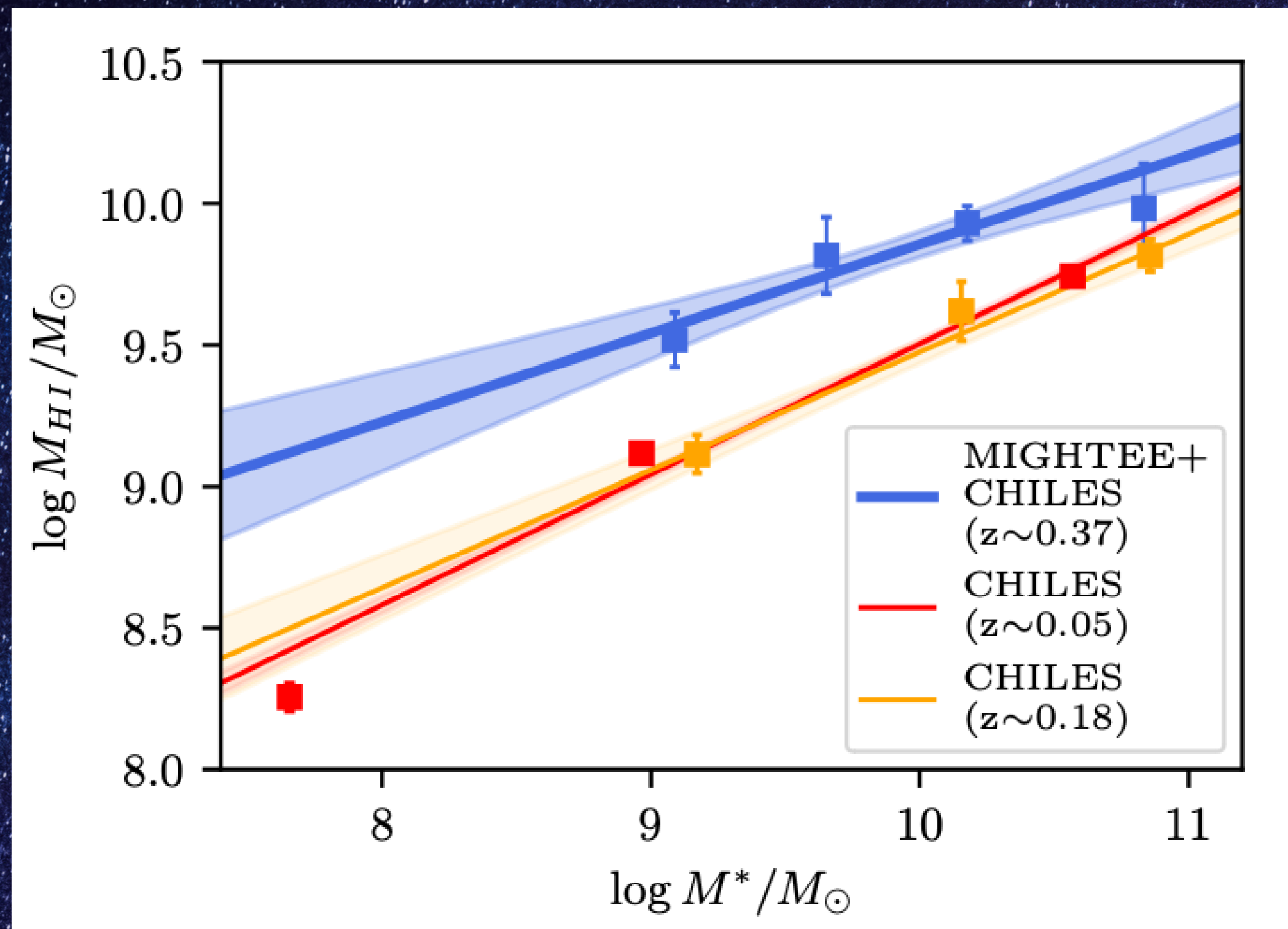
COMBINED STACKS

- ✓ robust S/N everywhere
- ✓ removed issue of the alleged velocity bias of the peak (updated redshift collection)
- ✓ absence of remarkable features outside the central signal
- ✓ no asymmetry in the continuum
- ⚠ continuum slightly negative?

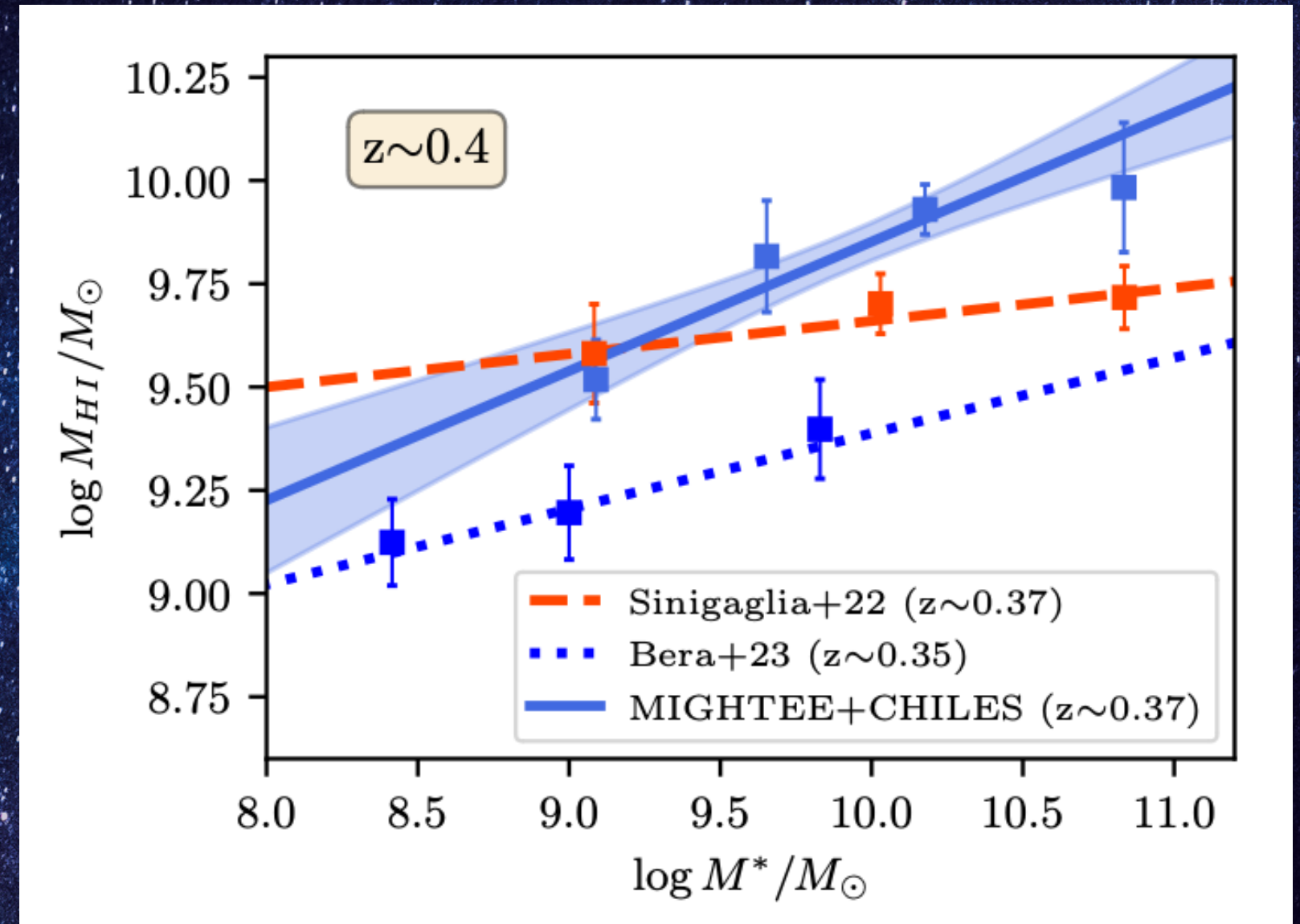


RESULTS

- blue squares: combined stack, MIGHTEE+CHILES ($z=0.37$)
- stacks on CHILES only, two redshift bins
 - $0 < z < 0.1$ $\langle z \rangle = 0.05$
 - $0.1 < z < 0.22$ $\langle z \rangle = 0.18$



STATE OF THE ART



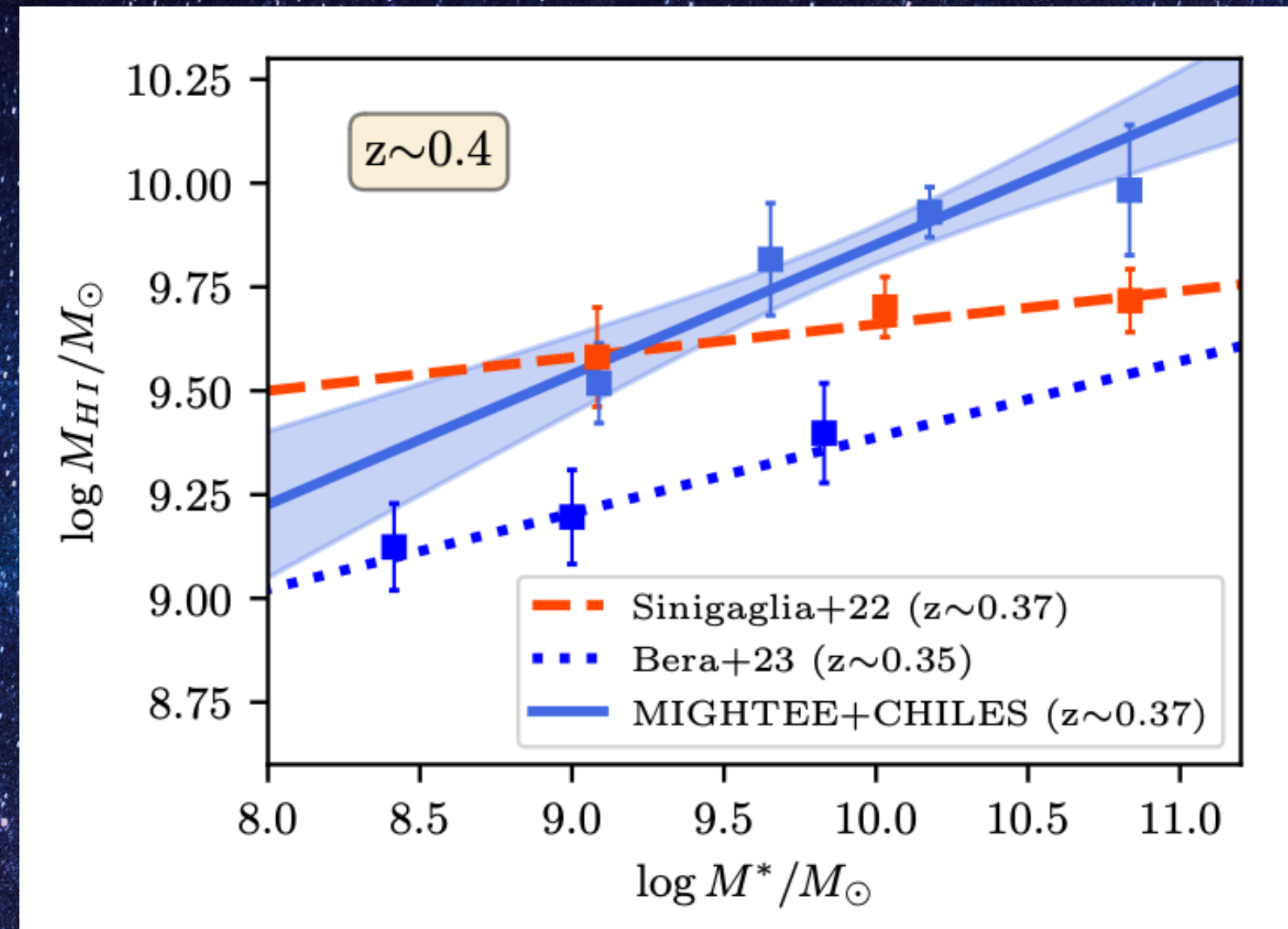
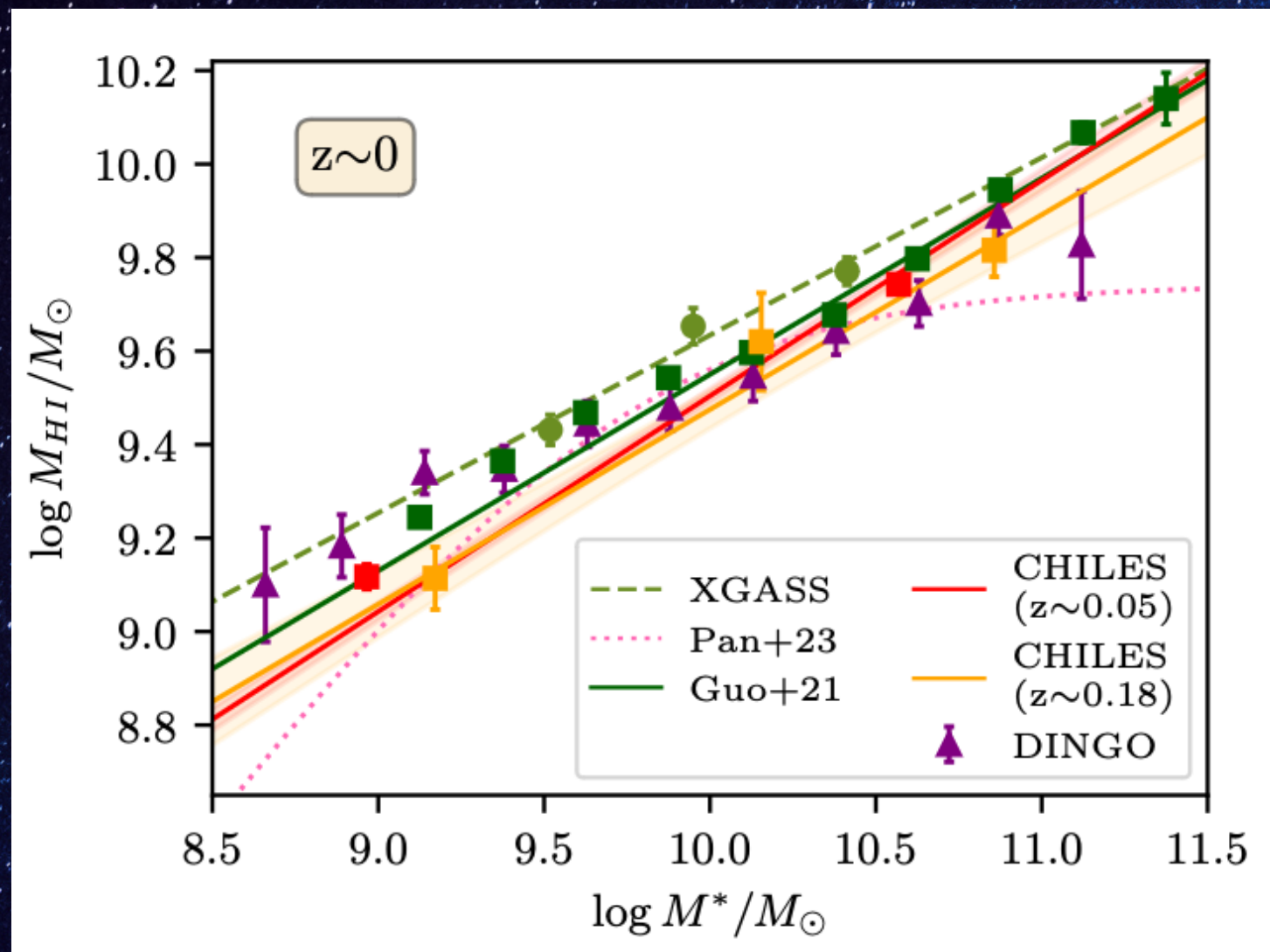
scaling relations at $z=0.4$

- change in tilt wrt previous MIGHTEE result slope independent on redshift
- need for a bent model?

STATE OF THE ART

scaling relations in the local Universe

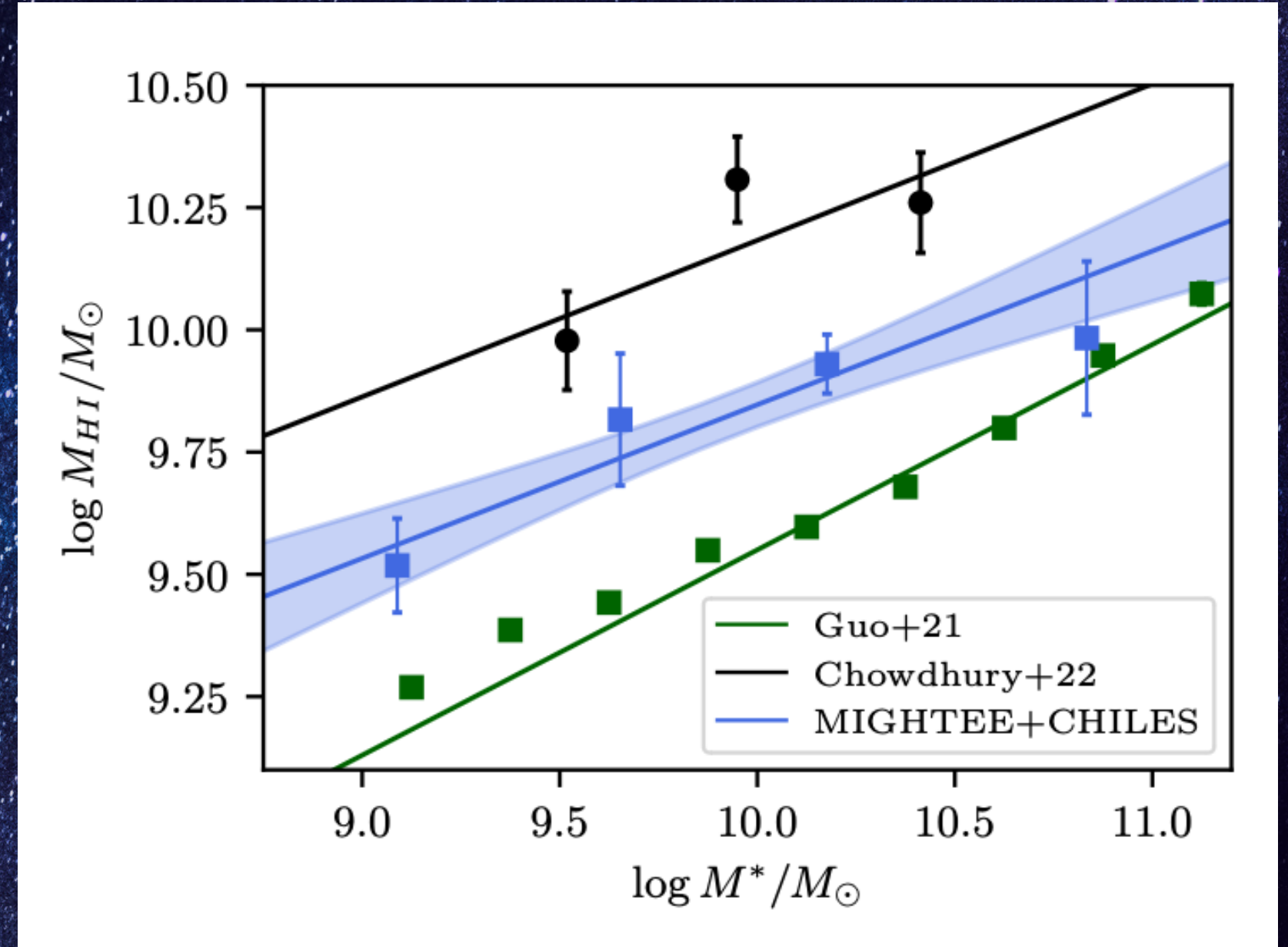
- the two low- z CHILES scaling relations seem compatible
- little to no evolution until $z=0.2$



scaling relations at $z=0.4$

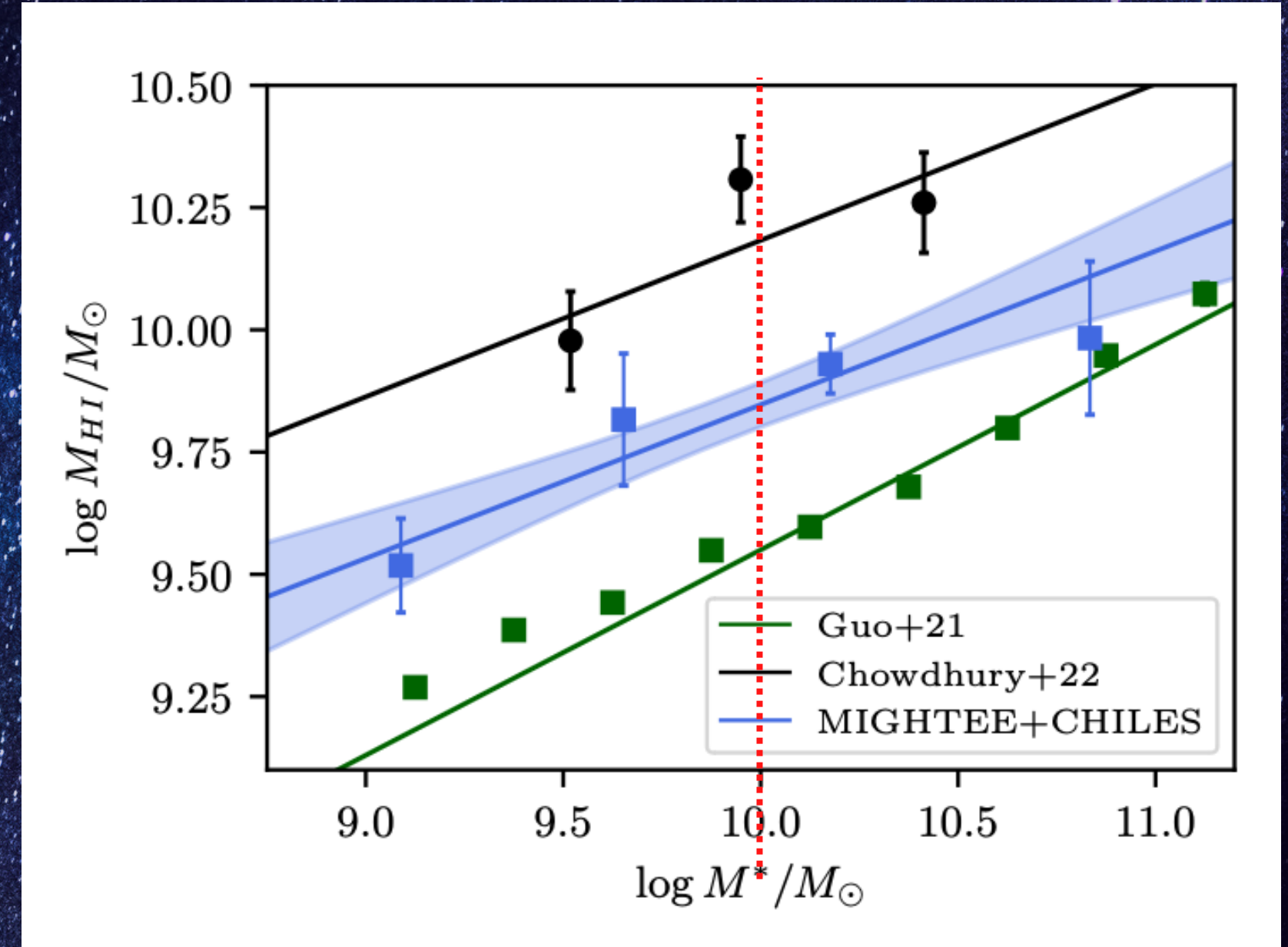
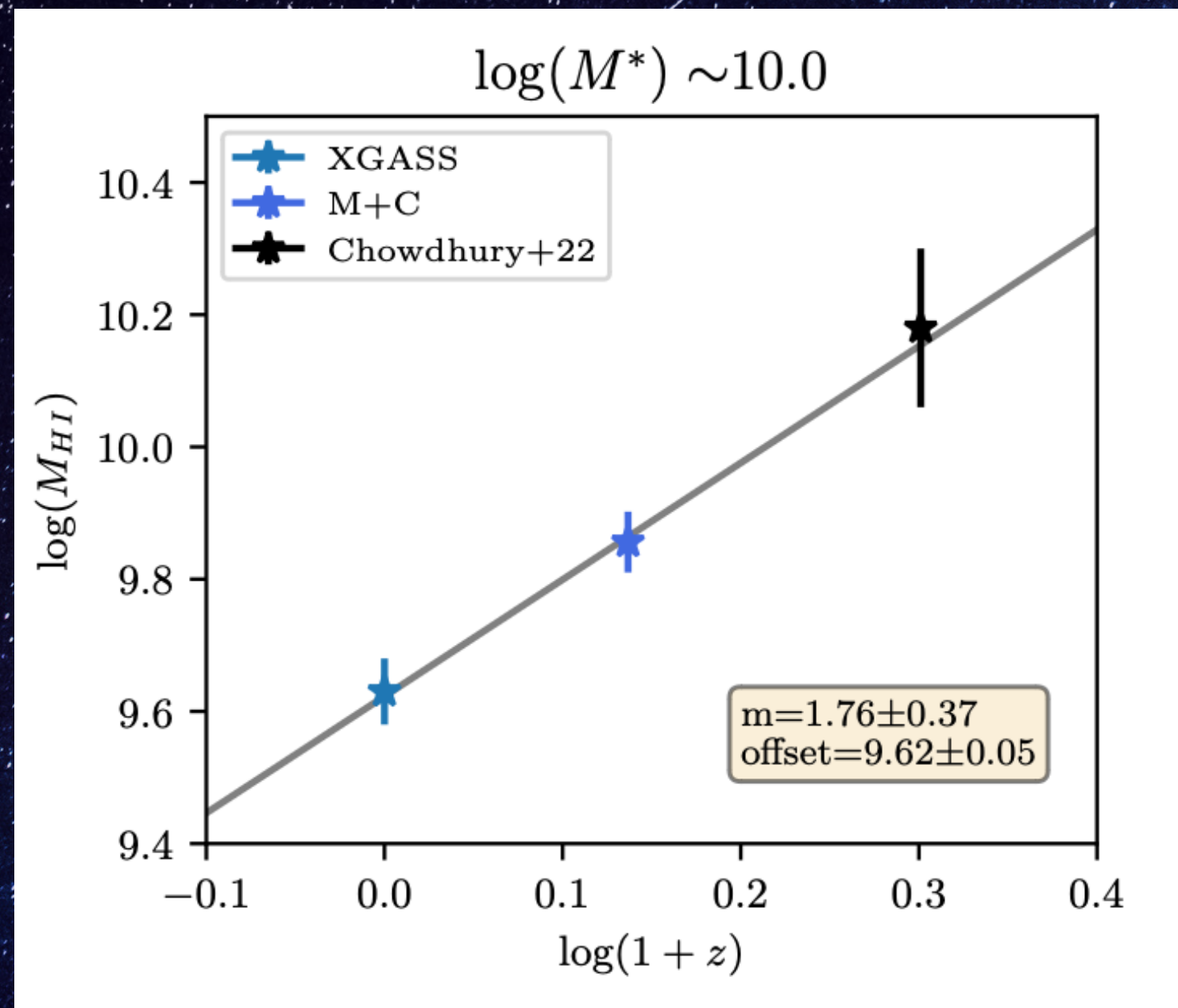
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SCALING RELATION EVOLUTION



- slope constant through redshift
- HI removal/replenishment mechanisms do not vary across stellar mass (at least until $\log M^* 10$)

SCALING RELATION EVOLUTION

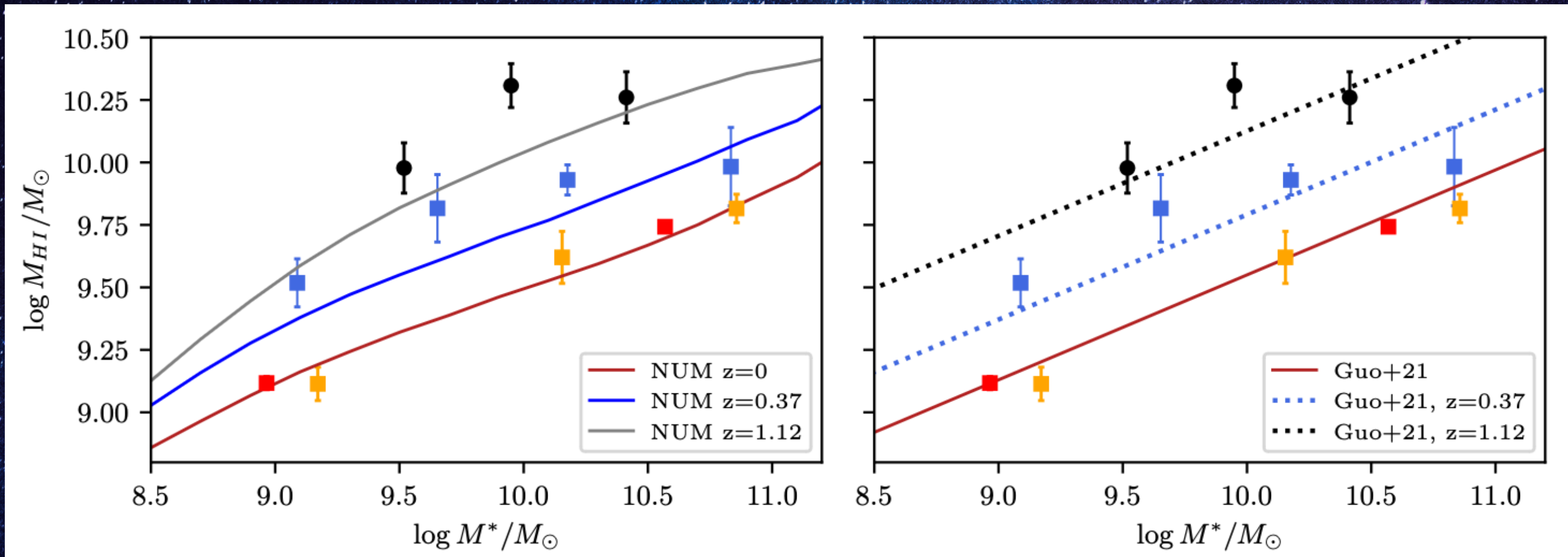


- slope constant through redshift
- HI removal/replenishment mechanisms do not vary across stellar mass (at least until $\log M^* = 10$)
- quantify evolution of M_{HI} as a function of redshift at fixed stellar mass ($\log M^* = 10$)

SCALING RELATION EVOLUTION

comparing with simulations

- NeutralUniverseMachine underpredict HI content
- re-scaling Local Universe scaling relation at $z=0$ provides good match with observations





WHAT'S NEW, WHAT'S NEXT

- stacking technical achievements:
 - MIGHTEE and CHILES stacks generally agree very well
 - the combination of the two ensure great statistics, great for stacking
 - refined RFI masking removed most of the unexplained features
- science questions:
 - do massive galaxies have very efficient Hi replenishment mechanism?
Worth investigating the bent model
- paper in preparation!
- **SUPERMIGHTEE**: pushing our stacking project to $z \sim 1$



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SARAO and UCT are offering a postdoctoral position, mixed MeerKat science operations support and personal research
<https://jobregister.aas.org/ad/de59ca7f>

Attention To: Prof. D.J. Pisano

Subject: SARAO-UCT Postdoctoral Fellowship

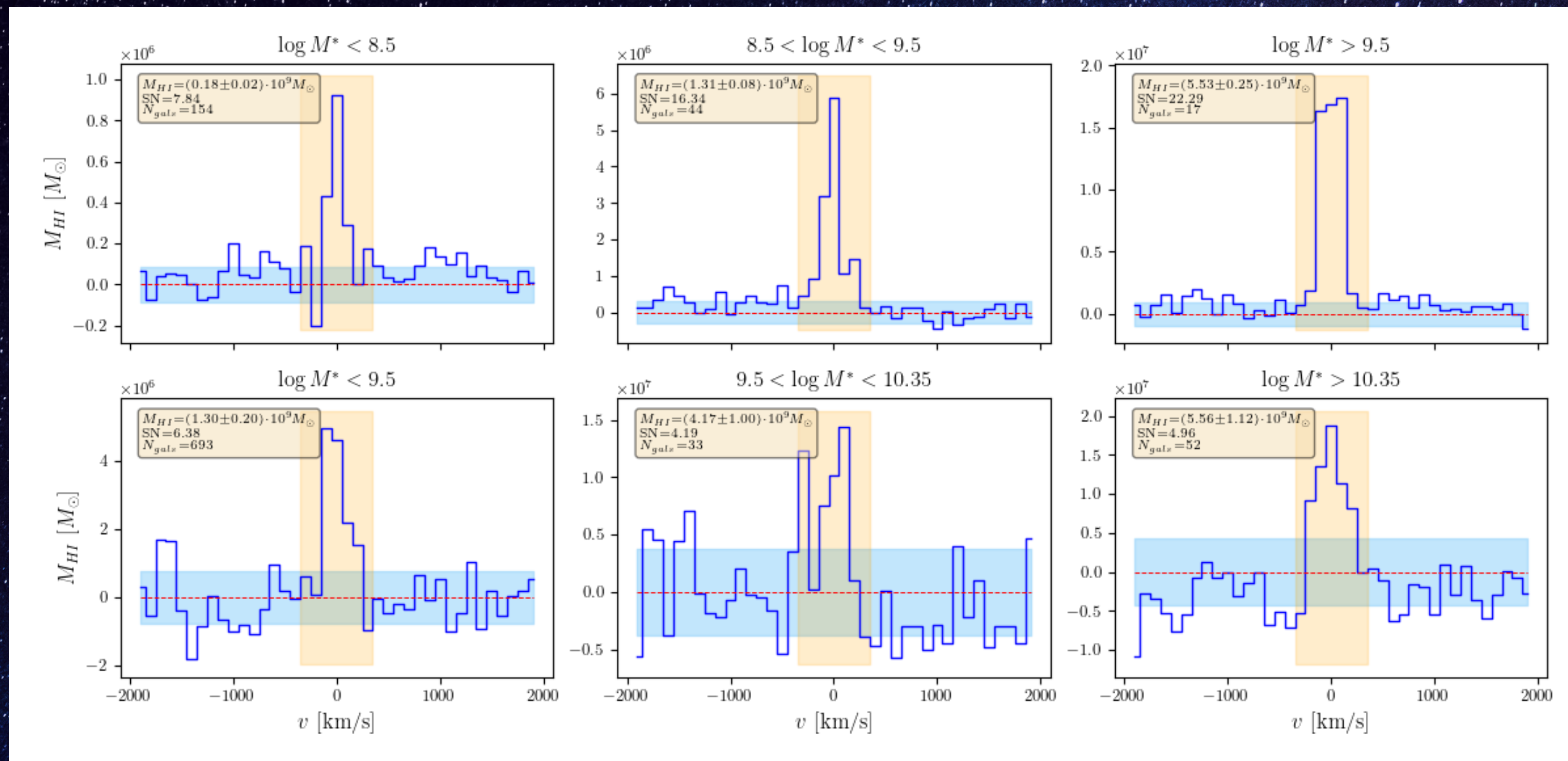
Email: postdoc_applications@ast.uct.ac.za

or contact: dj.pisano@uct.ac.za - mattia.vaccari@uct.ac.za - lucia.marchetti@uct.ac.za

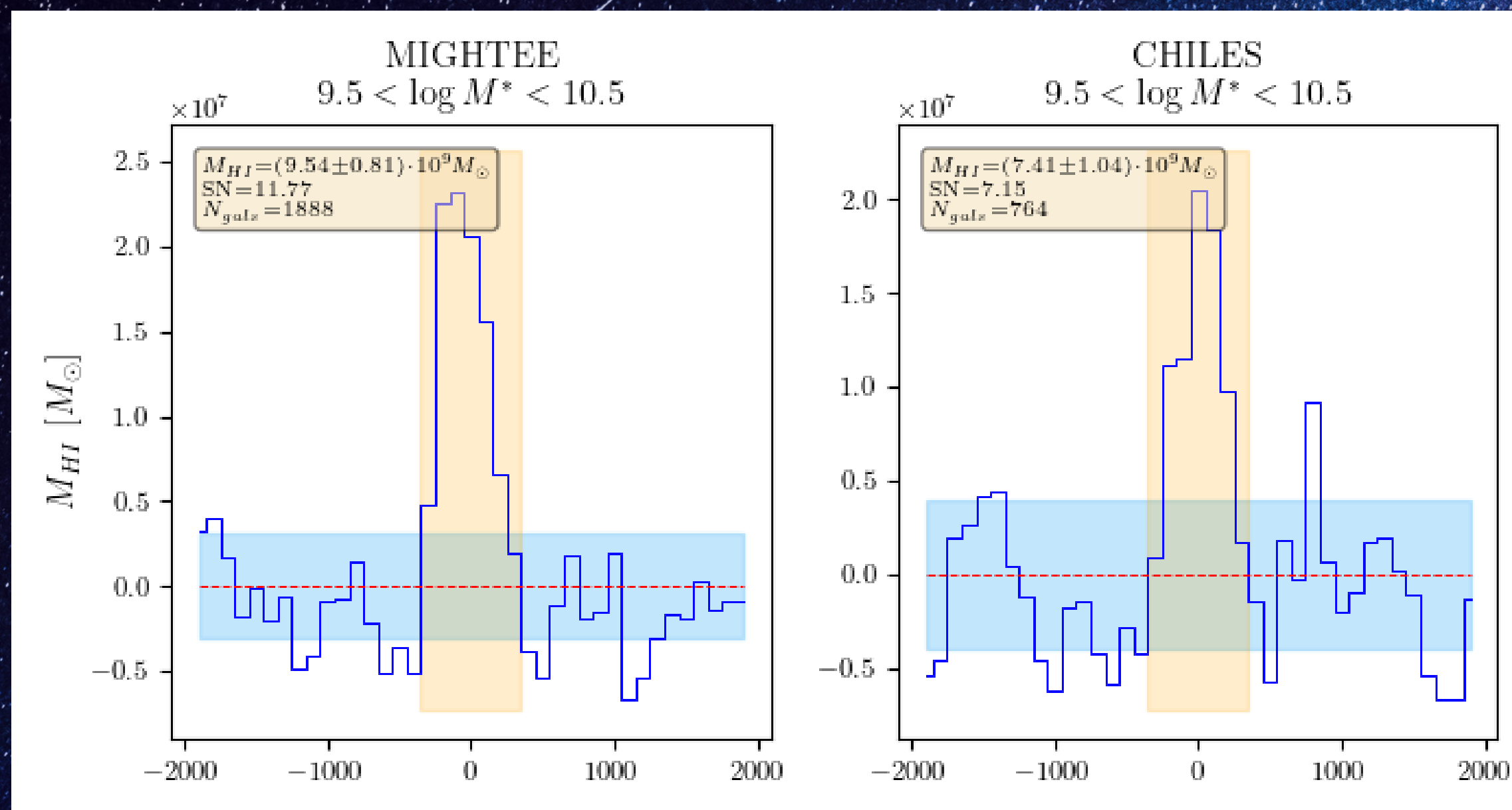


Deadline: January 15th, 2024

BACKUP: CHILES LOW-Z STACKS



BACKUP: MIGHTEE-CHILES COMPATIBILITY



Stacking setup

- RFI masking
- weighting scheme
- extracting aperture = 3xbeam

BACKUP: MIGHTEE-CHILES COMPATIBILITY