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NP Transition Matrix code G. La Mura, G. Mulas, R. Saija, M. A. Iatì, C. Cecchi-Pestellini

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ICSC Italian Research Center on High-Performance Computing, Big Data and Quantum Computing

Missione 4 • Istruzione e Ricerca









Scientific Rationale

- The *NP Transition Matrix* code envisages the problem of modelling the scattering of radiation by a distribution of particles with complex geometric and optical properties.
- The code is well suited to investigate scattering processes in **aerosols** (e. g. Sindoni et al. 2006), in the **interstellar medium** (Cecchi-Pestellini et al. 2010) and in all conditions where the shape of particles (and its deviation from spherical geometry) has a large impact.
- The calculation includes **dynamical** and **thermal effects** of the radiation particle interaction (Borghese, Denti, Saija 2007), for proper evaluation of the scatterers' physical state.
- The presence of non-trivial geometry has important implications on the derivation of differential **interaction cross-sections** and in their result as integrated effects (Saija et al., 2003).
- Detailed modelling of absorption and scattering (extinction), as well as polarization properties are possible (Borghese, Denti, Saija 2007).









Technical Objectives, Methodologies and Solutions

- GOAL:

- model radiation scattering / absorption by particles with arbitrary shape and optical properties
- solve the analytical problems arising in absence of spherical symmetry
- account for thermal and mechanical effects of the radiation particle interaction

- APPROACH:

- field expansion in **polar spherical harmonics**
- construction of the **Transition Matrix** from boundary conditions

- PECULIARITIES:

- model complex particles through an arbitrary number of spherical components
- allow stratification of materials for more realistic grain simulation
- RESULT:
 - connect analitically incident and scattered field for every interaction angle (as opposed to numerical solutions based on *Discrete Dipole Approximation*)
 - compute extinction and polarization for both near- and far-field cases (addressing problems such as chirality and chemistry in dust cavities)
 - model forces, torques, dust acceleration, centrifugal stress, evaporation



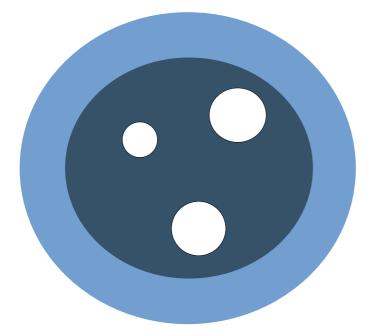






Technical Objectives, Methodologies and Solutions

Ζ



 10^{-10}_{-10} 1^{-10}_{-2

Example 1: Spherical particle with layered coating and internal cavities. Different color shades represent changes of material and refractive index

Example 2: representation of a dust particle as aggregate of spherical components. Different color shades represent different coating / material

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Timescale, Milestones and KPIs

The project structure can be narrowed down to three main stages:

Profiling (estimated completion: 3 months) **KPIs** • Identification of bottle-necks Evaluation of resource requirements Legacy result reproducibility Representation of data structures ▶ - Parallelization (estimated time: \geq 5 months) Transition of core functions to C++ Computing time reduction **Resource** optimization Definition of parallel architecture for calculation steps Migration of data structures to parallel architecture - Practical application (estimated time: 3 months) Customizable configuration system Scatterer complexity System requirements • Interstellar dust properties from extinction models Optical tweeners control









Accomplished Work, Results

- ORGANIZATION OF LEGACY CODE:

- Existing implementation in (old-fashioned) FORTRAN
- Scattering solution for a single spherical unit (with internal stratification)
- Scattering solution for clusters of units
- Radiation trapping (optical tweeners)

- SET UP OF DEVELOPMENT ENVIRONMENT:

- Source code management in *gitLab*
- Migration to C++
- Addition of inline documentation (*doxygen*)

- TEST CASE COLLECTION:

 Choice of pre-computed models for profiling activity and consistency checks Communication among units handled through physical files











Next Steps and Expected Results (by next checkpoint: April 2024)

- Profiling activity
 - Definition of the model test suite (published reproducible results)
 - Identification of bottle-necks and parallelisation strategies
 - Choice of parallel architecture (e.g. MPI for wavelength parallelism, OpenMP for T-matrix, porting to GPUs)

- Code porting

- Refactoring of data structures and core functions (using FITS or netCDF instead of binary I/O)
- Optional integration with algebraic libraries (LAPACK, ScaLAPACK, MAGMA upon convenience, but with fall-back portable internal routines)
- Consistency checks between state-of-the-art and ported results

- Interface development

- Improvement of configuration files (human readable formats)
- Implementation of GUI-oriented setup (depending on time requirements)
- Introduction of diagnostic tools (for inspection of binary data)

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