## Exploring the Galactic content in s-process elements

## High precision space spectroscopy with Gaia RVS



RVS = Radial Velocity Spectrograph

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Gaia/RVS is **SPACE spectroscopy f** ground-based spectroscopy

- Continuous observations for years: 34 months for DR3 (~25 000 h of continuous observations) 60 / 120 months for DR4 / DR5
- High number statistics : hundreds of thousands of high SNR (>150) DR3/data
- Stable conditions (no atmosphere)
- Very good control and modeling of systematics
- Extremely homogeneous treatment

Parametrization quality comparable to ground-based surveys of higher spectral resolution and wavelength coverage.

#### CU8/GSPspec: 5.6 million stars parametrized in Gaia DR3



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## 5.6 million stars with chemo-physical parameters







#### → AGB Stars in GSP-spec/DR3

~130,000 AGB with High-Quality atm. parameters



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#### $\rightarrow$ AGB Stars in GSP-spec/DR3

- ~18,000 stars with [Ce/Fe] AND High-Quality parameters + Abund.
- ~ 3,500 stars with [Nd/Fe] **AND** High-Quality parameters + Abund.





Vacuum Wavelength (nm)

#### → AGB production of s-process elements

Higher Ce and Nd abundances for more evolved AGB stars of similar metallicity.



#### → AGB production of s-process elements

- Comparison with FRUITY & Monash models : C/O<1 confirmed
- These stars have M<~3Msun
- Too large Ce & Nd enhancements predicted by Monash



Circles: FRUITY models Diamonds: Monash

- → **AMBRE Project** (de Laverny+14, Worley+12,...)
  - Automatic parameterisation of ESO archived spectra



- ~8700 UVES and FEROS spectra of MS  $\rightarrow$  AGB stars

PbI line @368.35nm (Contursi+24b)





#### → ~700 stars with Pb abundances: ~35 % giants but No AGB Pb-rich



→ Some Pb-enriched stars: Already known CEMP stars + new candidates



→ [Pb/Eu] and [Pb/alpha] trends with metallicity: ~ flat Galactic Chemical Models (Trieste): lower Pb production in metal-poor AGB + important contribution from high-rotating massive stars



- Large sample of AGB parameterised thanks to Gaia/GSP-spec
- 2nd peak *s*-elements: Ce & Nd abundances from Gaia/RVS
  - [Ce/Fe] & [Nd/Fe] very well correlated with AGB evolutionary stage
  - Stars are more enriched in Ce & Nd when more advanced on the AGB

SUMMARY

- Favoured production site : low-mass AGB
- Good agreement with FRUITY predictons
- 3rd peak *s*-element: Pb abundances from the AMBRE Project
  - ~700 Galactic stars with Pb abundances
  - Galactic chemical models: lower AGB production at low [M/H]?

## **Spectroscopic analysis of Gaia spectra**

## **SUMMARY**

- The Gaia future is bright:
  - only <sup>1</sup>/<sub>4</sub> of the data analysed in DR3!
  - RVS data SNR increasing with time : Better-quality parameters + Abund.
- Much larger chemo-dynamical catalogues to come:
  - 5.6 million stars with chemo-physical parameters in DR3 (2022)
  - $\circ$  ~8 / ~20 times more stars in DR4 / DR5
  - $\circ$  New chemical species to come

# Gaia/GSP-spec offers high-precision parametrization with unprecedented high number statistics

## **<u>Gaia/GSP-spec</u>** Frequently Asked (and not-Asked) Questions

- When I query the Gaia catalogue, my sources do not have GSPspec parameters.
  Please check that you query the right table: AstophysicalParameters
- When I plot GSPspec parameters, there is a cloud of points...
  - Please check the quality flag chain (including atm. param AND abundances)
  - Filter out results with quality lower than desired
  - Check also the published uncertainties and the SNR (rv\_expected\_sig\_to\_noise)
- GSPspec parameters have to be calibrated
  - See recommendations in Recio-Blanco et al. 2023 & 2024
  - For a given species with published abund., others have assumed [X/Fe]=0