







Overview on ASTRI activities in Rome

S. Lombardi et al. – INAF-OAR and ASI-SSDC for the ASTRI Project

VHE Rome meeting – 10/05/2023









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Both technical and scientific activities carried out!

In-depth and long-term experience in Cherenkov Telescopes



The ASTRI Mini-Array @ Teide Observatory

- **9 improved ASTRI-Horn-like telescopes**
- **Under construction** at the Observatorio del Teide (Tenerife, Spain, 2390 m a.s.l.), in collaboration with Instituto Astrofísico de Canarias (IAC)
- First 3 operating telescopes by spring 2024; Full array completion within a few years
- **Being developed in all its aspects**, from design/ implementation of all HW/SW components to dissemination of final scientific products
- INAF IAC hosting agreement (~4+4 years)
- Unprecedented performance and wide FoV for observations at TeV/multi-TeV energy scale
- **Important synergies** with other Northern groundbased gamma-ray facilities (LHAASO, HAWC, VERITAS, MAGIC, LSTs, CTAO-N)











The ASTRI Mini-Array Science Operations

First 3/4 years - Core Science: Vercellone, S. et al., JHEAP 35, 1 (2022)

- \succ Array operated as an experiment (not as an observatory)
- Core Science Program developed by ASTRI Science Team
- Pillar Topics: Origin of Cosmic Rays and Fundamental Physics
- Few selected targets / Deep exposures (~200 hr)
- Science beyond VHE astronomy -> Stellar Intensity Interferometry
- \triangleright Possibility of MoUs with other γ -ray facilities for joint observations

Afterwards -> Observatory Science:

D'aì, A. et al., JHEAP 35, 139 (2022) Saturni, F. G. et al., JHEAP 35, 91 (2022)

- Observing proposals from a wider scientific community
- Broader scientific cases, with main focus on the multi-TeV domain:
 - Galactic sources: wide FoV -> multi-target fields
 - Extragalactic sources: survey of a few promising targets at > ~10 TeV
 - Fundamental physics: studies on LIV, EBL, Axion-Like Particles, ...
 - **Direct Measurement of Cosmic Rays**
 - Time Domain and Multimessenger Astrophysics



Mini-Array

Galactic targets:

- SNRs / PeVatrons / SNRs interacting with MCs
- PWNe / TeV Halos
- Gamma-ray binaries



Extragalactic targets:

- Extreme BL Lacs
- Radio galaxies
- Starburst galaxies



Fundamental physics:

- Lorentz invariance violation / Axion-like particles / Hadron beams in AGN jets
- EBL
- Dark Matter







The ASTRI Mini-Array Operation Sites

The ASTRI Mini-Array at Tenerife:

- Telescope Array & auxiliaries (Observatorio del Teide OT)
- Local Control Room @ THEMIS building (OT)
- On-site Data Centre @ IAC Teide Residencia (OT)
- Array Operation Center @IACTEC in La Laguna





The ASTRI Mini-Array in Italy:

- Data Center @INAF in Rome
- Remote Array Operation Centers (any INAF Institute involved)



Main ASTRI activities in Rome

- ASTRI Mini-Array:
 - Software System (management)
 - Data Processing System (responsible)
 - Simulation System (responsible)
 - > Archive System (responsible)
 - \succ Offsite ICT ASTRI Data Center in Rome (responsible)
 - Science Support System (responsible)
 - Transient Handler and Science with ASTRI Mini-Array (responsible)
 - Galactic and Extragalactic Science with ASTRI Mini-Array (relevant contributors)
 - String ASTRI Speaker's and Publication Office (ASPO) (management)

(+ main activities related to the ASTRI-Horn telescope on Mt. Etna)







Main achievements, so far... (I)





ASTRI-Horn Crab Nebula detection:

- Real data taken with ASTRI-Horn (December 2018)
- Analysis using A-SciSoft software package (single-telescope pipeline)
- First detection of a gamma-ray source with a Cherenkov telescope in dual-mirror Schwarzschild-Couder configuration

→ This achievement has represented an important step towards the validation of the dual-mirror optical design for ground-based gamma-ray astronomy applications



Main achievements, so far... (II)







ASTRI Mini-Array performance and IRFs:

- Monte Carlo data from ASTRI-MA-Prod2_20deg (2021)
- Analysis using A-SciSoft software package (end-to-end array MC data pipeline)
- Full assessment of ASTRI Mini-Array performance and generation of IRFs for scientific studies:

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→ The achieved results have demonstrated the significant ASTRI Mini-Array capabilities for observations at TeV / multi-TeV energy scale, in synergy with present- and nextgeneration gamma-ray observatories in the Northern Hemisphere







Main achievements, so far... (III)



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The ASTRI Mini-Array of Cherenkov telescopes at the Observatorio del Teide

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Extragalactic observatory science with the ASTRI mini-array at the Observatorio del Teide

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ASTRI Mini-Array core science at the Observatorio del Teide

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Galactic observatory science with the ASTRI Mini-Array at the Observatorio del Teide

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Backup slides







ASTRI: Astrophysics with Italian Replicating Technology Mirrors

ASTRI-Horn Prototype

INAF-led Project funded by Italian Ministry of Research

End-to-end prototype of the 4-m class telescopes in the framework of CTA Observatory; installed and operational on Mount Etna volcano (Sicily, Italy)

First detection of a gamma-ray source (Crab Nebula) above 5 σ with a dual-mirror, Schwarzschild-Couder Chrenkov telescope (Lombardi et al., 2020)





Array of 9 ASTRI telescopes

INAF-led Project with international partners: Univ. of Sao Paulo/FPESP (Brazil), North-West Univ. (S. Africa), IAC (Spain), FGG, ASI/SSDC, Univ. of Padova, Perugia and INFN

Being deployed at the Observatorio del Teide (Spain) in collaboration with IAC and FGG-INAF.

First 4 years \rightarrow *Core Science*, following 4 \rightarrow *Observatory* Science. Science operation \rightarrow 2025









The ASTRI-Horn telescope



Dedicated to Guido Horn D'Arturo, precursor of the segmented astronomical mirrors technique

SiPM Cherenkov camera:

Performance:



Technological demonstrator with an end-to-end approach

(HW and SW validation through actual Cherenkov observations)

Dual-mirror Schwarzschild-Couder design:

 \blacktriangleright Primary Mirror (M1): \oslash 4.3 m (18 segments) Secondary Mirror (M2): Ø 1.8 m (monolithic) > Optimal PSF ($\leq 0.19^\circ$) across entire FoV \blacktriangleright Post calibration pointing precision \leq 7 arcsec \blacktriangleright Reduced plate scale (37.5 mm/°) \rightarrow use of SiPMs

Fast front-end electronics based on CITIROC ASICs 7×7 mm SiPM sensors (Hamamatsu Photonics) \succ 1344 pixels (21 modules of 8×8 pixels) > Angular pixel size: 0.19° (\geq optical PSF) \succ Field of View: ~8°

Energy threshold: ~1 TeV \blacktriangleright Energy/Angular resolution: $\leq 25\% / \leq 0.15^{\circ}$ \blacktriangleright Sensitivity: 1 Crab @ 5 σ in few hours









Main ASTRI activities in Rome

- ASTRI-Horn prototype:
 - Observation schedules and data taking
 - Data reduction and analysis of real data
 - Monte Carlo production and validation
 - Data archiving

Data exploitation (mainly for validation purposes)







The Data Processing, Simulation, Archive Systems

Data Processing System

Collection of software components in charge of preparing, calibrating, reducing, and analyzing raw Cherenkov data acquired during the observations up to the generation of high-level science-ready data and automated science products. Also in charge of providing suitable data check and calibration products, as well as performing data reduction of the Stellar Intensity Interferometry Instrument data.

Simulation System

Collection of software components in charge of generating all simulated data needed for the ASTRI Mini-Array Project. Design and development phase: MC simulations used to optimize the telescope positions in the array, to fully characterize the (scientific) performance of the system, and to define and test data analysis methods. Commissioning and operation phase: used to provide auxiliary inputs for real data reconstruction (LUTs) and system response (IRFs).

Archive System

Software (and hardware) service that provides storage and organization for all persistent information (data, data products, and metadata) generated for and by the ASTRI Mini-Array systems and defined by the Mini-Array Data Models. The system plays a central role in the whole observing life-cycle of the array, which goes from observation preparation and execution, to data processing and dissemination of the high-level science-ready data and automated science products to the Science Users.



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SPIE Paper 12189-27 (Proc. SPIE2022)

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The Data Processing, Simulation, and Archive Systems of the ASTRI Mini-Array Project

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(DPS Working group)







Data Processing System: main components

• Stereo Event Builder (SEB)

Software component that performs the off-line stereo array trigger of the Cherenkov events recorded by the telescopes of the ASTRI Mini-Array. The array trigger acts on information sent by each triggered telescope (like the trigger timestamps and the telescope status) to select Cherenkov shower signatures using stereoscopic information. (See Germani et al., Paper 12183-66.)

Cherenkov Data Pipeline (CDP)

Software component that performs data reduction of the Cherenkov events observed by the ASTRI Mini-Array. Main goal: reconstruct the physical characteristics of the astrophysical gamma rays (and background cosmic rays) from the raw Cherenkov data collected by the ASTRI Mini-Array telescopes during the observations of the scientific targets. Usage of the A-SciSoft software package.

Intensity Interferometry Data Pipeline (IIDP)

In charge of performing data reconstruction and scientific analysis of the Stellar Intensity Interferometry Instrument (SI3). (See Zampieri/Rodeghiero et al., Paper 12183-14.)

Calibration Software

Collection of software tools that reduces and analyzes all the data produced during the calibrations of the system. It provides suitable calibration factors for the calibration of Cherenkov data.









Data Processing System: Cherenkov data flow



CDP Breakdown stages ; I/O data ; Basic Functionalities ; DPS executables















A-SciSoft(v0.5.6)

https://www.ict.inaf.it/gitlab/astri/data-processing/ascisoft

Status:

Main Applications (so far):

- Data analysis of 2022-2023 ASTRI-Horn data
- Regular MC data reduction and analysis for
 - Mini-Array performance assessment
 - Mini-Array IRF3 production



Reduction of ASTRI-Horn real data (custom analysis, up to DL1c) End-to-end MC pipeline (still, workaround code for IRF3 generation) > External ad-hoc tools for data quality checks, diagnostics (at different DLs)

Detection of the Crab Nebula with ASTRI-Horn (December 2018)

Simulation System: simulation chain





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Archive System: physical units



- To archive
 - Interaction with actors



- **ASTRI Mini-array Archive System**
 - > plays a **central role** in the whole observing life-cycle of the array
 - Shall guarantee long-term data preservation and access
 - > shall manage: observation plans; science data; monitoring/alarm/logging data; system configurations

2 separated physical units (10 Gbit/s connected):

- On-Site Archive System:
 - @Teide Observatory (Tenerife, Spain)
 - Archive System for temporary storage (~1 week) and services
 - > On-site ICT (see Gianotti et al., Paper 12189-53)
- **Off-Site Archive System:** •
 - \geq @Rome (Italy)
 - > Archive System for long-term data storage and services
 - Off-site ICT -> ASTRI Data Center







