

OSSERVATORIO ASTROFISICO DI ARCETRI



Contactless active mirrors

for next generation space telescopes

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Context: next generation space telescope

Next generation of space telescopes will investigate cosmic structures, galaxies formation, exo-atmospheres,...

Science requirements include:

- High contrast
- high angular resolution
- exceptional optical quality and stability



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Active Optics:

crucial technology to meet REQ and reduce mission cost.

BTW these values are beyond our current capacity [





Requirements for ExoEarth Imaging ~10 m diam. ~10 nm surf err ~10 pm stability ~10⁻¹⁰ contrast

Open question:

How can ground base AO contribute to LUVOIR?

E-ELT adaptive M4

6 segments 2.5m diam. 5300+ actuators 10nm Surf Quality incl. phasing





The LATT Project



A 2010 project on Space Active Optics: Testing the **conversion** of the **Adaptive Secondary** into an **Active Space Primary**



Proven with a demonstrator: **17Kg/m², 55mW/act, 19 acts. 1 mm stroke, TRL 5** Laboratory tested. Project ended in 2015.

Prototype recovered in Arcetri under a loan agreement with ESA in 2021.

LATT - working principle





- Reference Body, Aluminum honeycomb
- Voice coil actuators
- Capacitive position sensors
- Thin Zerodur glass shell as optical surface
- magnets on the shell back

The ThinShell "floats" at 100-800um from RefBody



Contactless active mirror?

Voice-coil motors + capacitive sensors provide a contactless actuation mechanism (100-800 um gap from RefBody to optical surface)

A breakthrough in space optics?

- shape error of support have no impact on optical quality
- low freq. deformations are corrected with 0% fitting error
- optical surface is insulated from vibrations from payload

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- mechanical decoupling → Requirements "separation":
 - ultra light-weight structures, simplified components, reduction of mass and cost

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The SPLATT experiment

Goal:

demonstrate that the "floating" optical surface is insensitive to vibrations

Funded by PRIN INAF 2019 - funded in 2021 137 k€, to recover the prototype and set-up the test tower





Vibrations injection

Piezo actuator on elevation arm, ______to inject controlled vibration on the support.

Vibration signal from waveform generator.

Freq range:1 Hz to 120 Hz

seismic accelerom. on the stand and RefBody



Vibration test procedure

We measure the TipTilt with a fast interferometer, when an external disturbance is applied

1. tilt TT_{L} when the ThinShell is pressed against the RefBody by the actuators



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We measure the TipTilt with a fast interferometer, when an external disturbance is applied

- 1. tilt TT_{L} when the ThinShell is pressed against the RefBody by the actuators
- 2. tilt TT_F when the ThinShell "floats" at the working gap (100 um, e.g.)





2. ThinShell floating. Tot force =0

Results of the SPLATT experiment



Tests to be repeated in vacuum chamber, next month...

Resonances @40, 60, 100 Hz but

Vibration rejection > 70% @70-90 Hz

Shell-RefBody coupling mediated by air??



Contactless, floating active mirrors:

could be breakthrough for next generation space telescopes

Physical **decoupling** between mech. support and optical surface:

- → improvement of scientific performances lower vibration foot-print lower thermal foot-print
- → reduction of system-wide requirements ultra light-weight structures, lower manufact. specs,
 - simpler REQ., mass&cost reduction

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Project funded by INAF-PRIN Tecno

137k€ for setting up the **facility** and demonstrate the concept

Grant from INAF was "necessary" aid as startup:

• out of scope for ESA/ASI (TRL and uncertainties)

The team just participated in the ASI "TopicalTeams"



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Results presented at: SPIE2022 ICSO2023

to be presented at ESA-ESTEC nov23

Strong synergy within INAF / ADONI:

- Wavefront sensing
- Control strategy
- Scientific sampling for high contrast

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 Physical decoupling bet
 Active optics for space:

 → improvement of scientific lower vibration foot-p lower thermal foot-p
 Active optics for space:

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 INAF can be a game changer.

 What's our next move?

backup

Laboratory test campaign





Temperature range: -25°C□55°C

Electrostatic locking test



locking pressure: 600 N/m2

Thermo-vacuum test



Tested @ 1e-5mbar

Vibration test



Max acceler.: 10g



WFE comparable with AO DM

Close loop result: sub-nm stability (and sensitivity)

-



Comments in view of LUVOIR:

- value is consistent with LUVOIR REQ
- No GS V<5 needed
 - (no laser GS constellation)
- time scale: << 1 min (relaxed stability REQ system-wide)
 - fast sampling \rightarrow post-facto correction

Frequency response, single frequency test



SPLATT Attenuation

 $T(freq) = T_F / T_L$

 T_{F} : Tilt ampl. shell floating T_{L} : Tilt ampl. shell lifted to RB

Frequency response, sweep test



Sweep measurement:

Excit.Freq variable in time 15 s to capture F range 1-120 Hz

Same processing



Results of the SPLATT experiment



Vibration rejection vs Actuator PID parameters, vs gap, Air vs Helium...

Frequency response, Air vs Helium



Measurement of auto-induced current

