

# First Galaxies with JWST

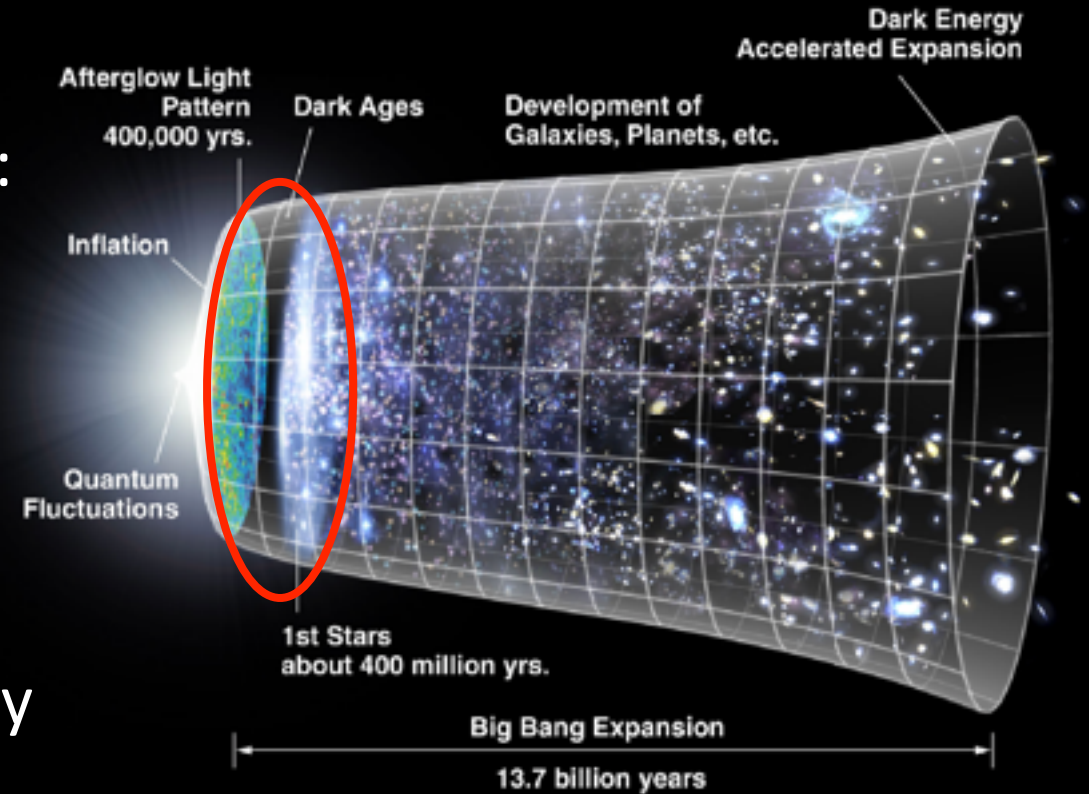


Adriano Fontana  
Istituto Nazionale di Astro Fisica

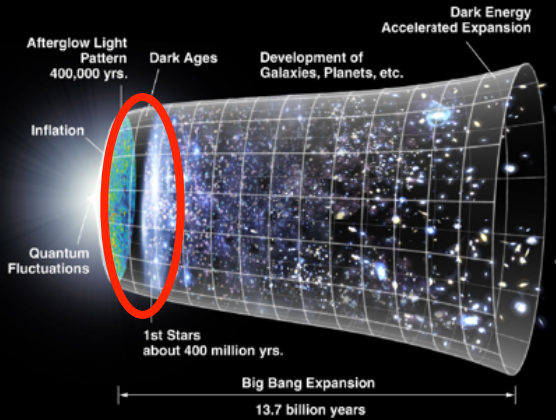
# Why is it so interesting to study the high- $z$ Universe?

Physical conditions were completely different:

- Universe much denser
- Primordial/Very low metallicity
- “Warmer” CMB
- No UVB
- Density fluctuations much smaller than today



# Why is it so interesting to study the high- $z$ Universe?



There is room to show evidence of “*New Physics*”

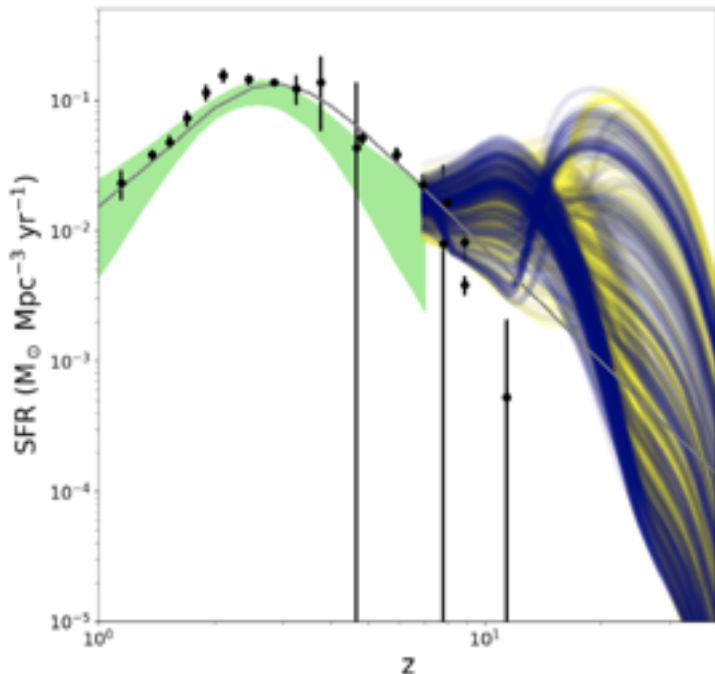
e.g.:

Non-standard  $\Lambda$ -CDM:

- redshift-time relation different? (more time to build structures?)
- WDM? Lower abundance of small structures

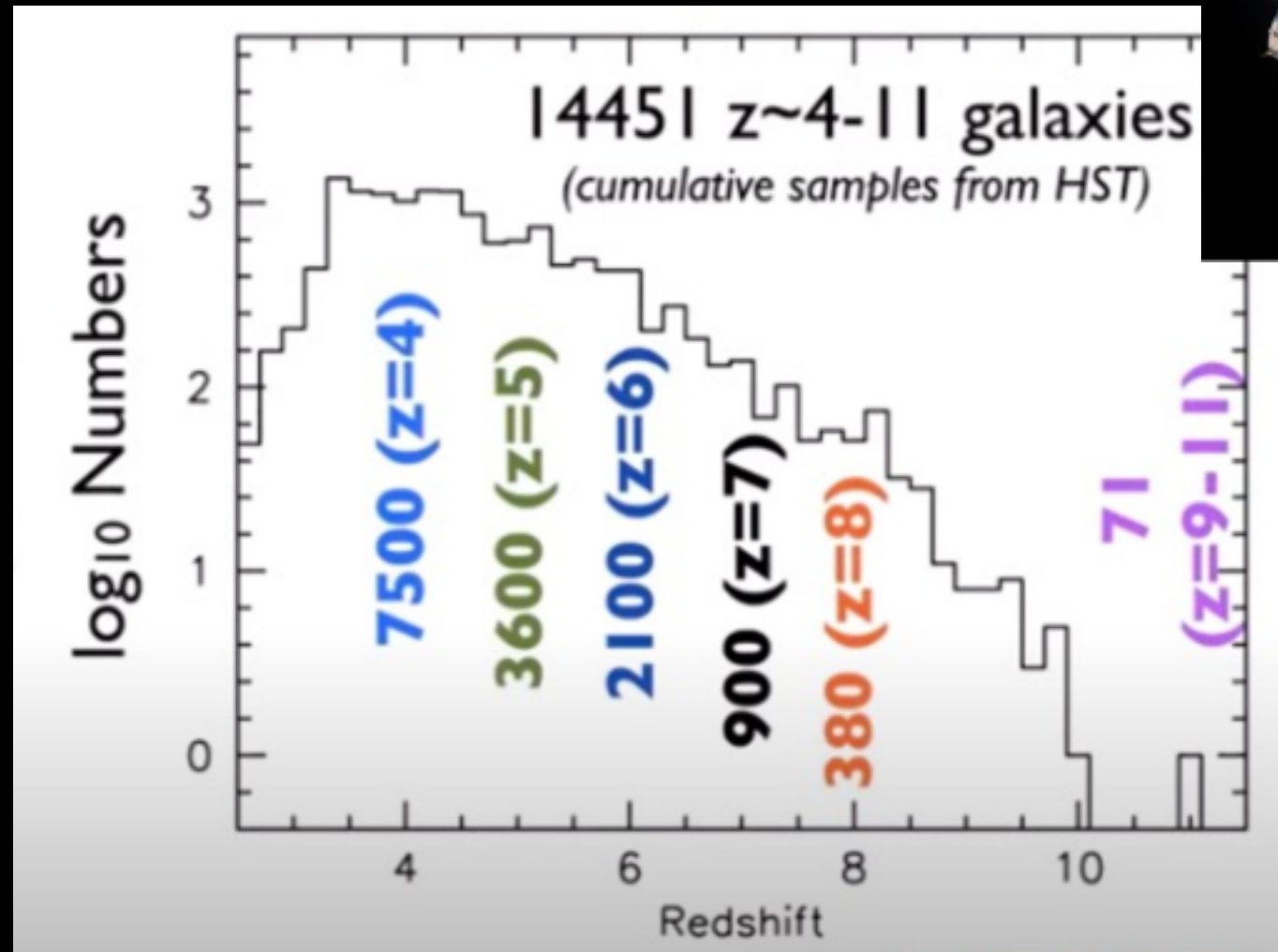
New constituents?

- Primordial BH trigger (additional) star-formation

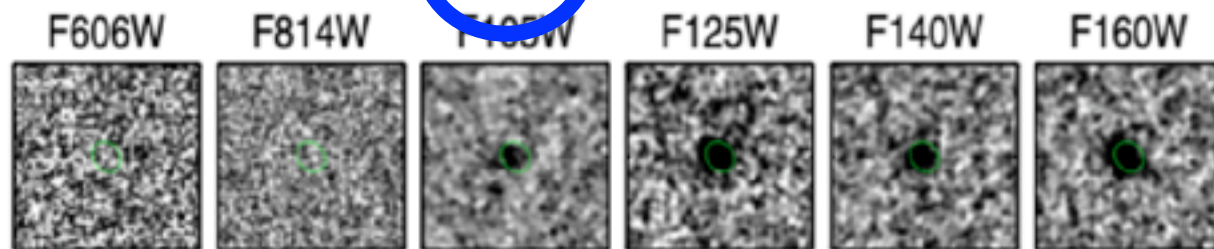




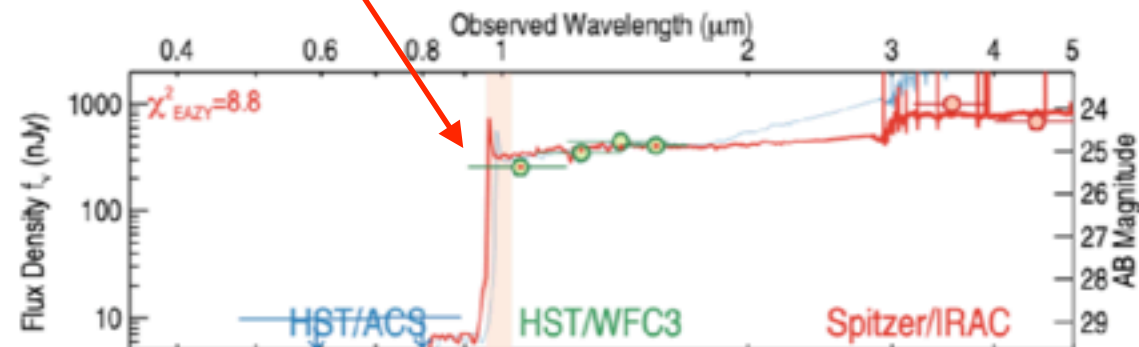
# The picture at the end of the HST era...



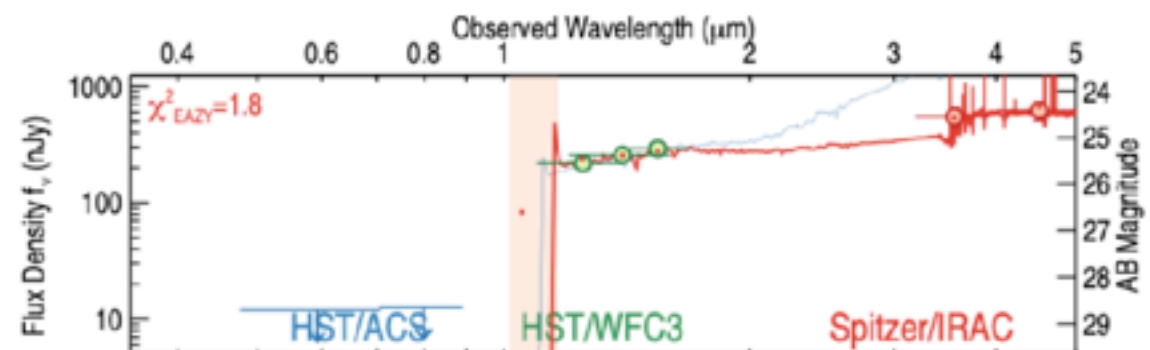
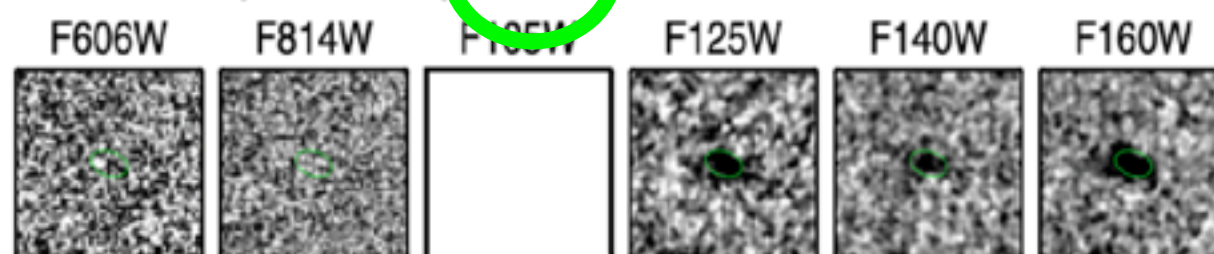
ID=13433, m=25,  $z=6.9$



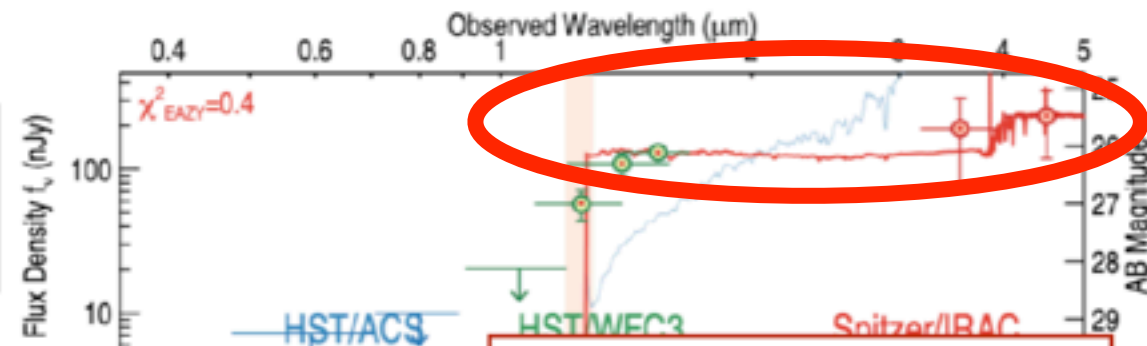
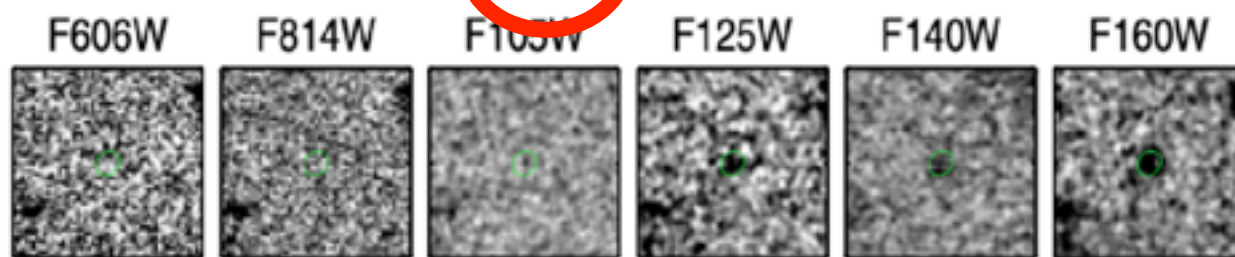
Lyman  $\alpha$  (121.567nm)  $\rightarrow$  1.2 $\mu$ m @  $z=9$



ID=7364, m=25.2,  $z=8.4$



ID=26816, m=26.1,  $z=9.4$



ID=26890, m=26.1,  $z=9.0$

NIRCam/NIRSpec coverage

On July 14, the first JWST *raw* data were made public...!  
A number of *public* programs were made available.  
Including the data from the *GLASS* program, a US/Italian collaboration.  
PI: Tommaso Treu (UCLA)







Marco Castellano, Emiliano Merlin, Diego Paris, Paola Santini,  
Antonello Calabrò, Sara Mascia, Laura Pentericci, Eros Vanzella (INAF)  
Andrea Correnti, Matteo Correnti, Gianluca Polenta (ASI SSDC)

[Submitted on 19 Jul 2022 ([v1](#)), last revised 26 Sep 2022 (this version, v2)]

## Early results from GLASS-JWST. III: Galaxy candidates at $z \sim 9-15$

Marco Castellano, Adriano Fontana, Tommaso Treu, Paola Santini, Emiliano Merlin, Nicha Leethochawalit, Michele Trenti, Uros Mestric, Eros Vanzella,



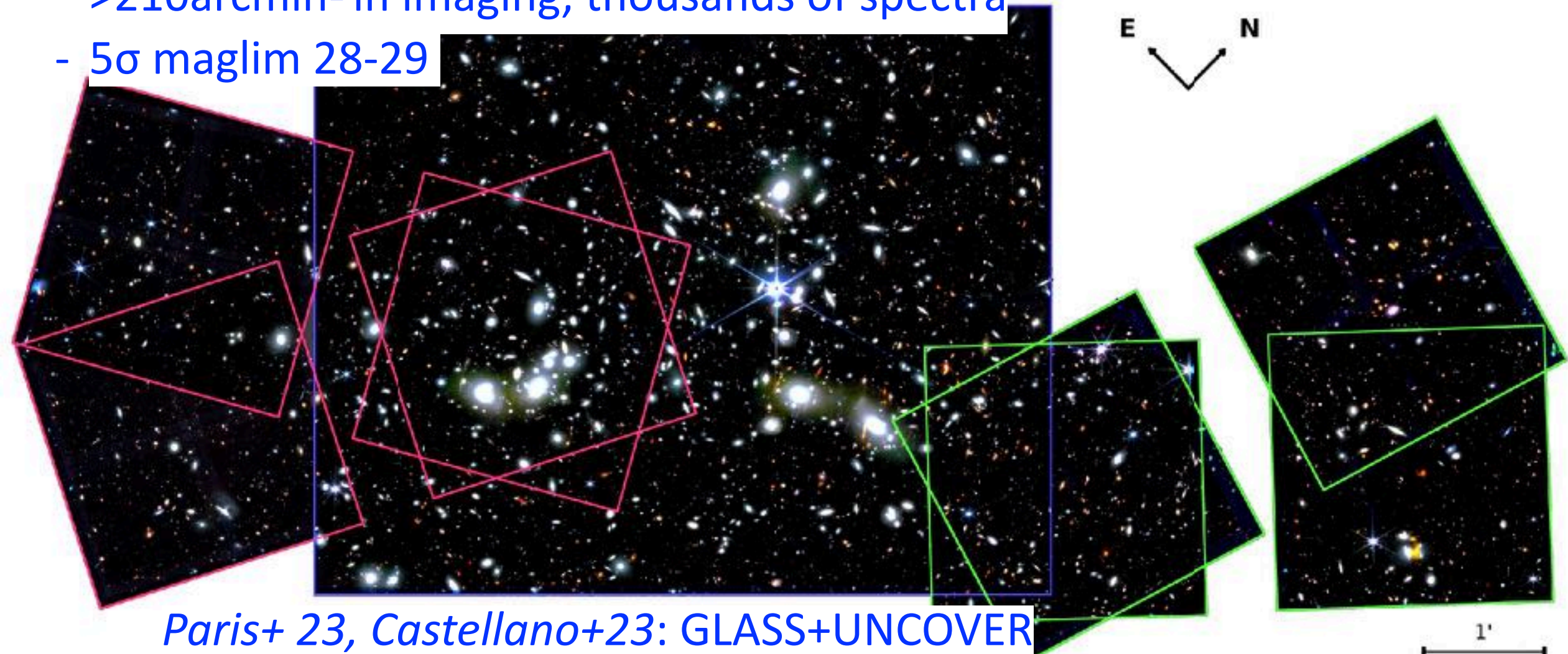
GLASS Deep Field: the *deepest* field obtained so far with JWST





In only 10 months, a spectacular data set has already been collected:

- 12 **public** deep fields (images and spectra)
- each with at least 7-8 bands + HST data
- $>210\text{arcmin}^2$  in imaging, thousands of spectra
- $5\sigma$  maglim 28-29





What did we get?  
A lot of surprises!

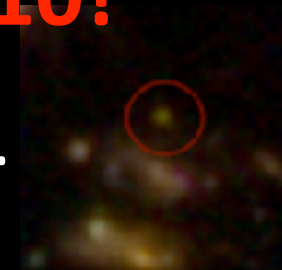




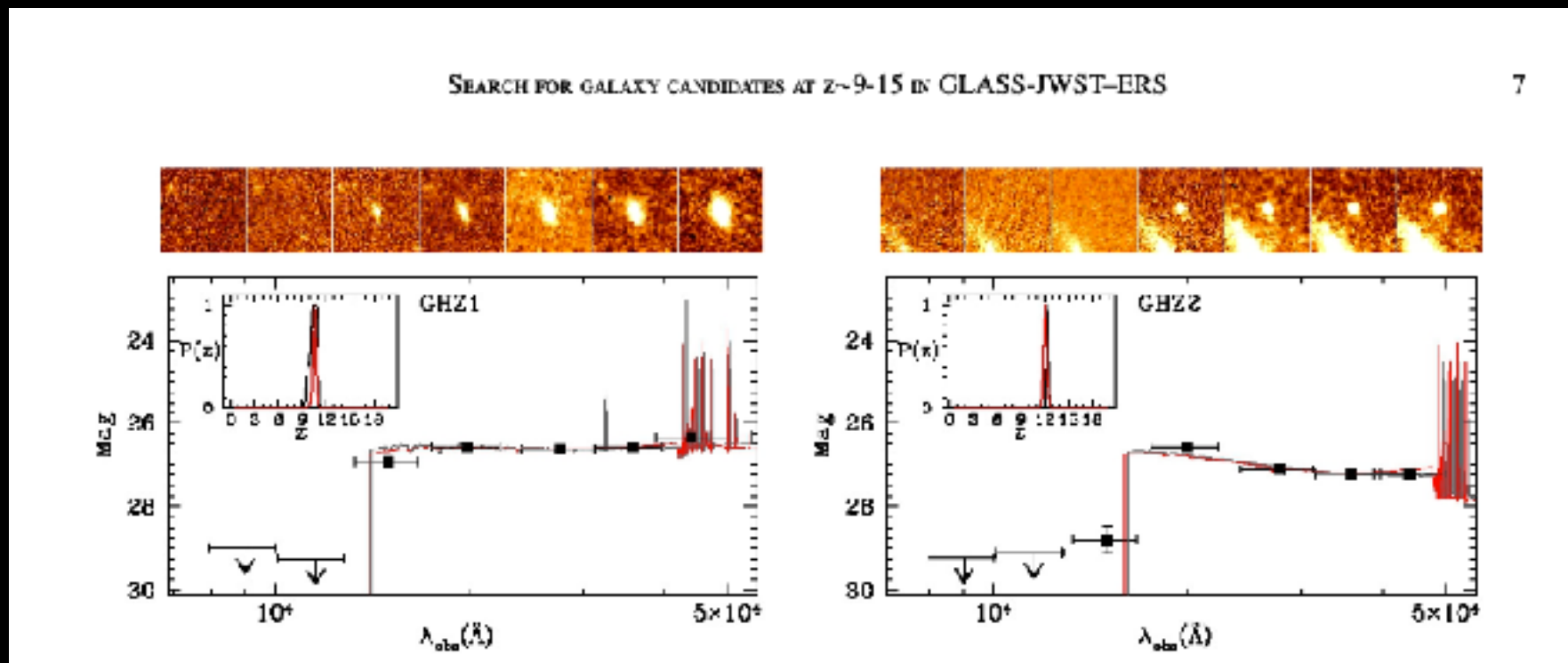


# Surprise#1: Too many UV bright galaxies at $z > 10$ !

Two bright & robust candidates in GLASS at  $z=10$  and 12...  
...We were expecting 0.1!

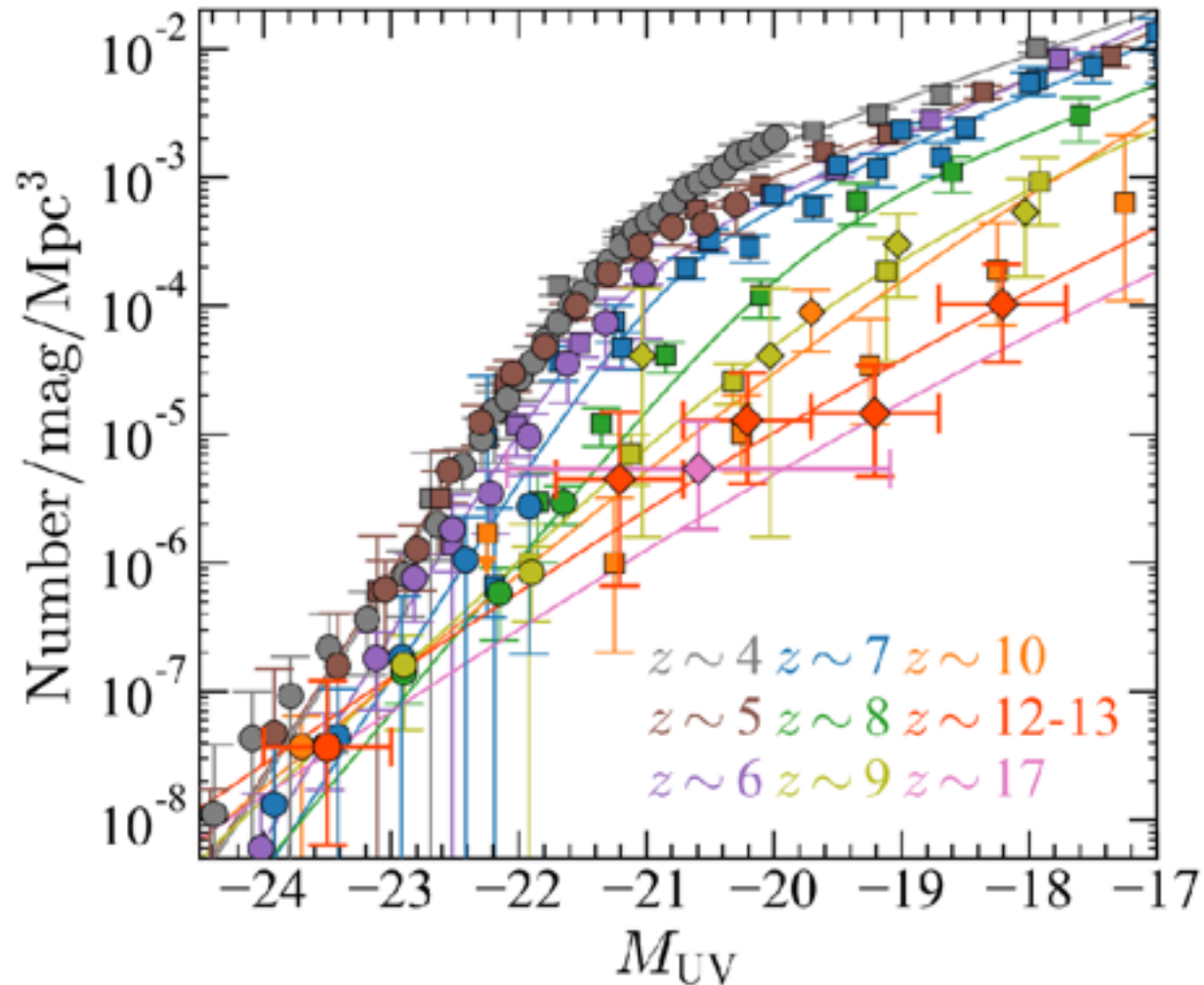


Castellano, AF+23

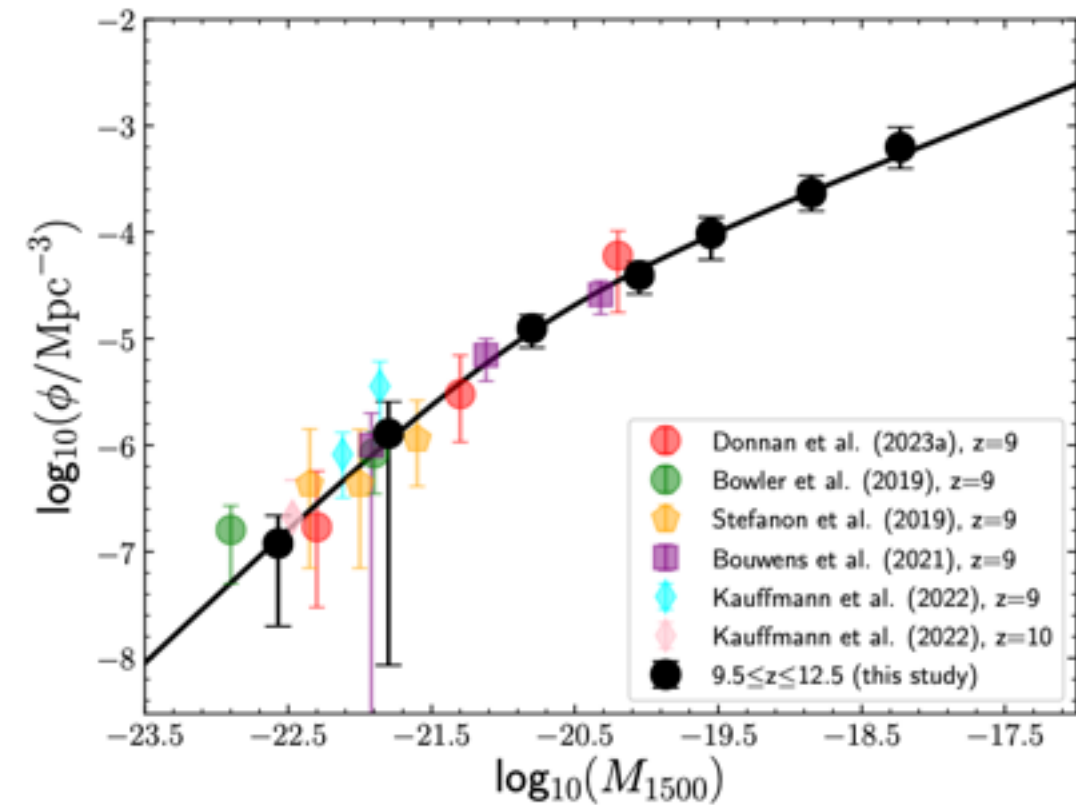


Later on discovered that GLASS is overdense (Castellano+23)

The observed evolution from  $z \sim 8$  to  $z \sim 12$  is marginal,  
especially for bright galaxies



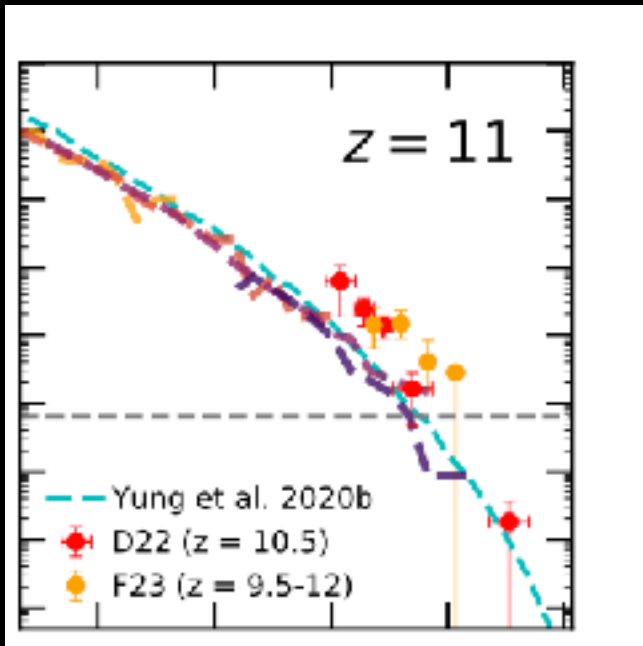
Harikane+22



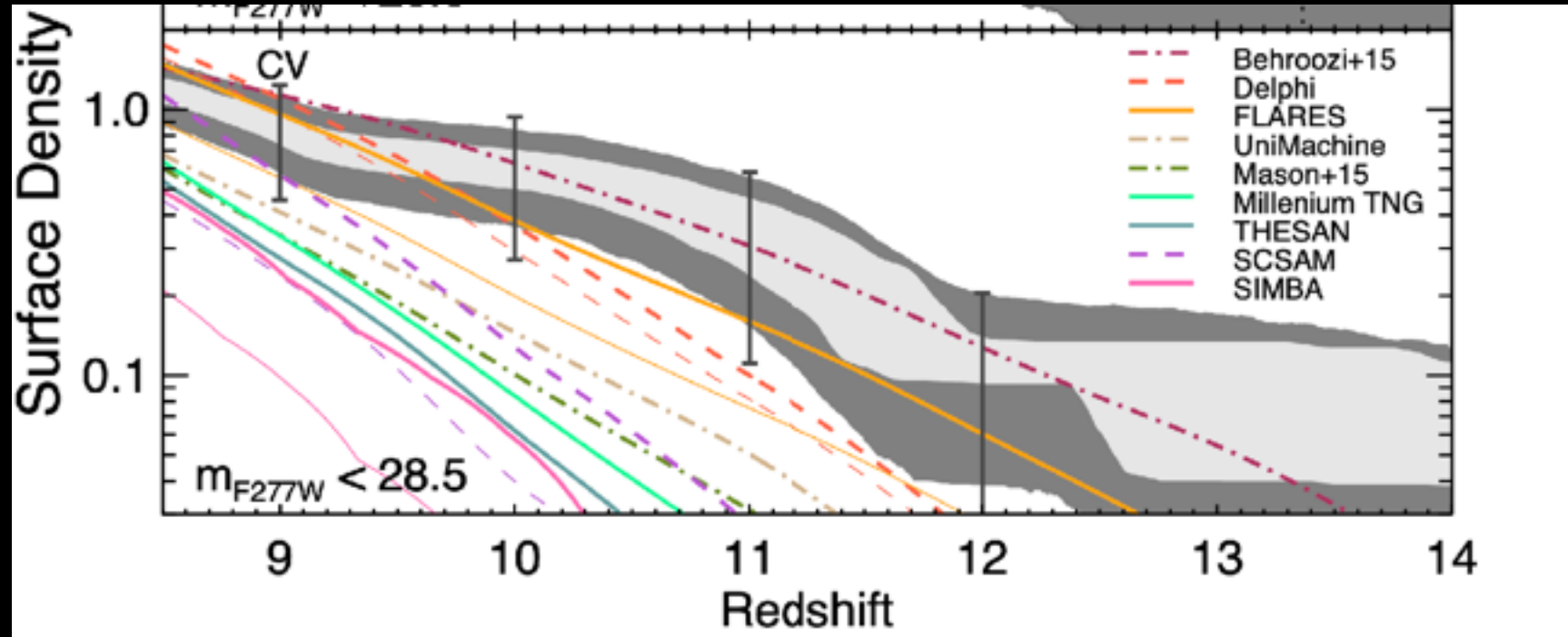
McLeod+23



The number of bright galaxies is way larger than predicted  
by *any* theoretical model



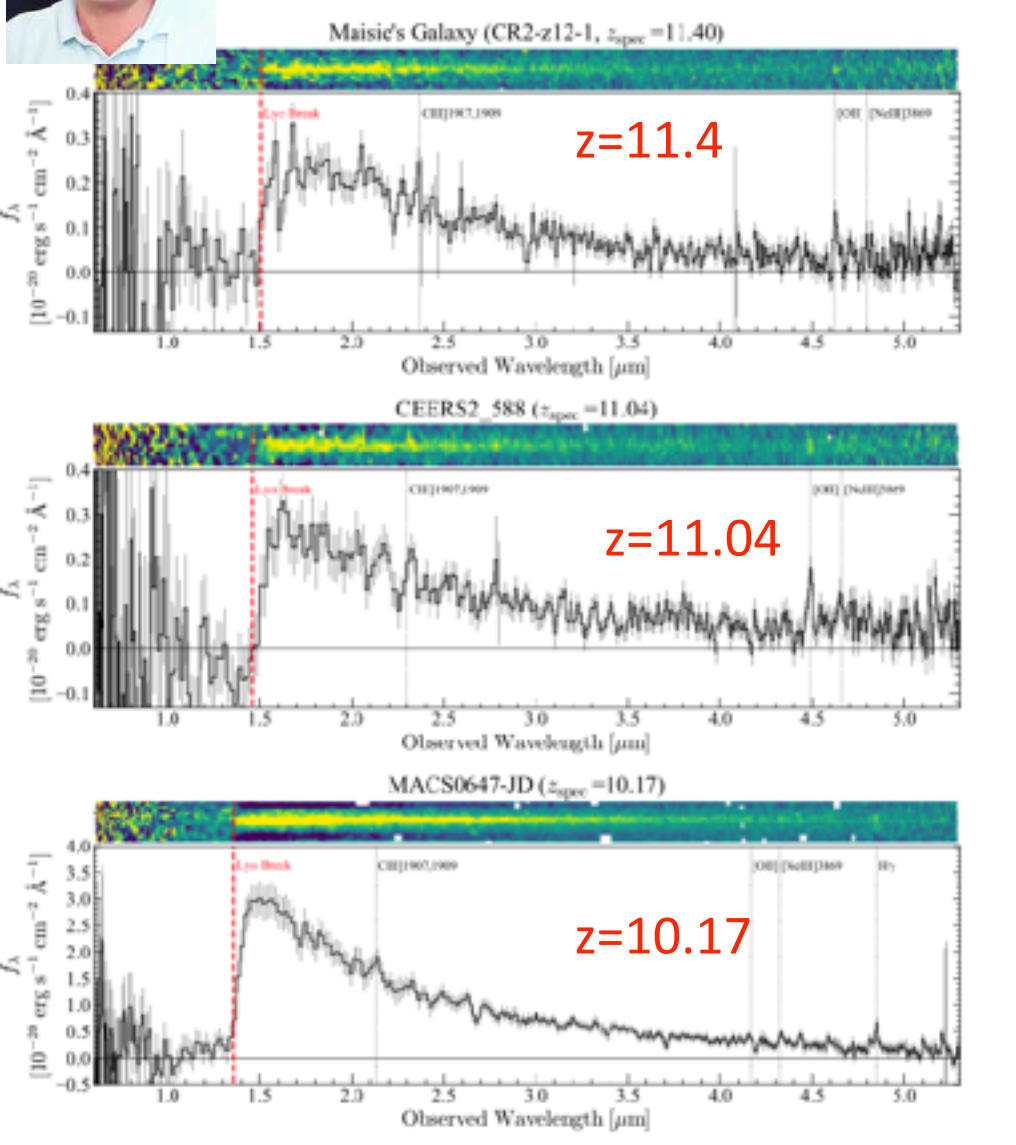
Yung+23, 2304.04348



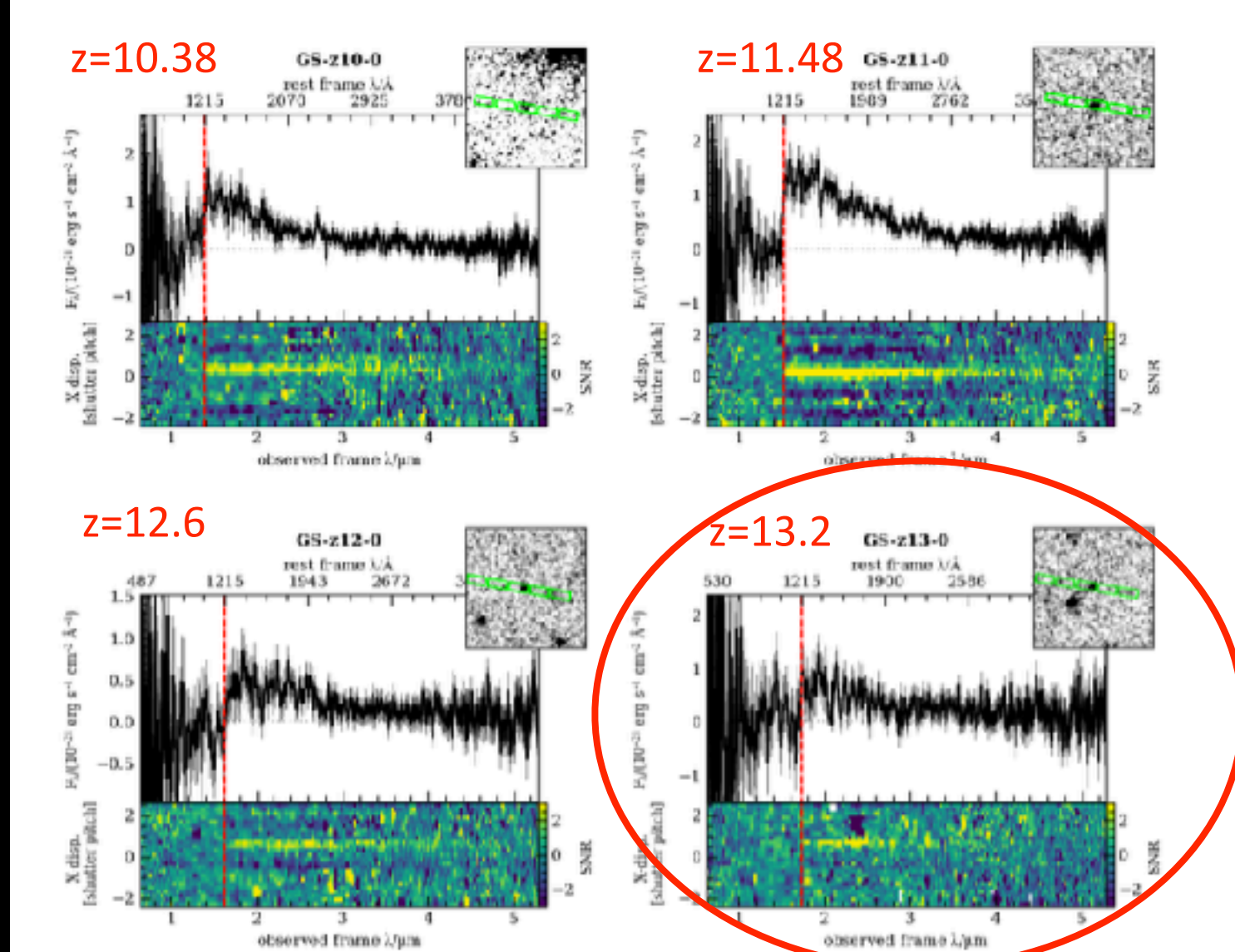
Finkelstein+23  
also Mason+23



# Surprise#2: We can get spectra up to $z \sim 13$ !



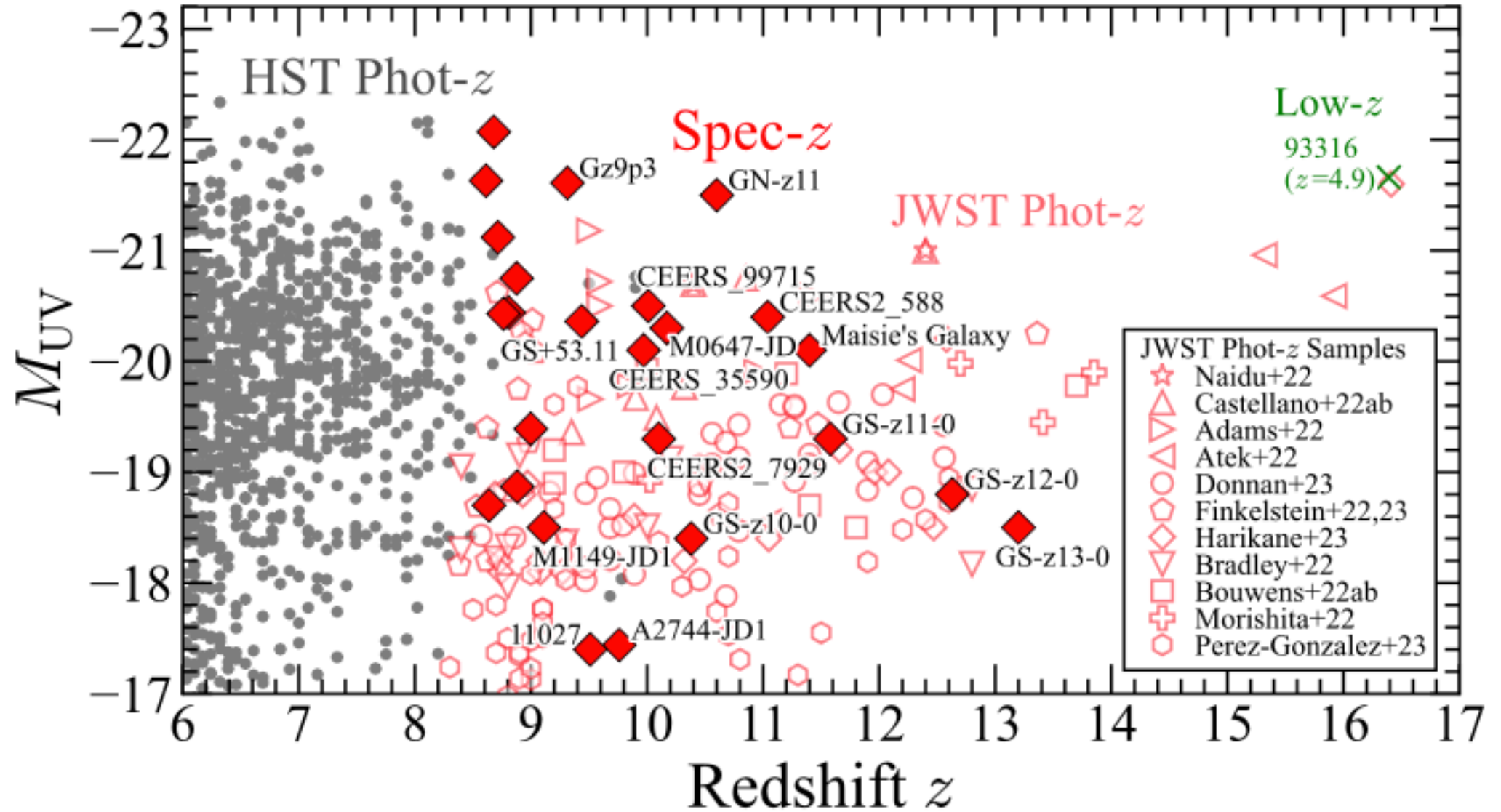
Harikane+23, 2304.06658



Curtis-Lake+23, Nature

Age of the Universe = 0.32 Gyr

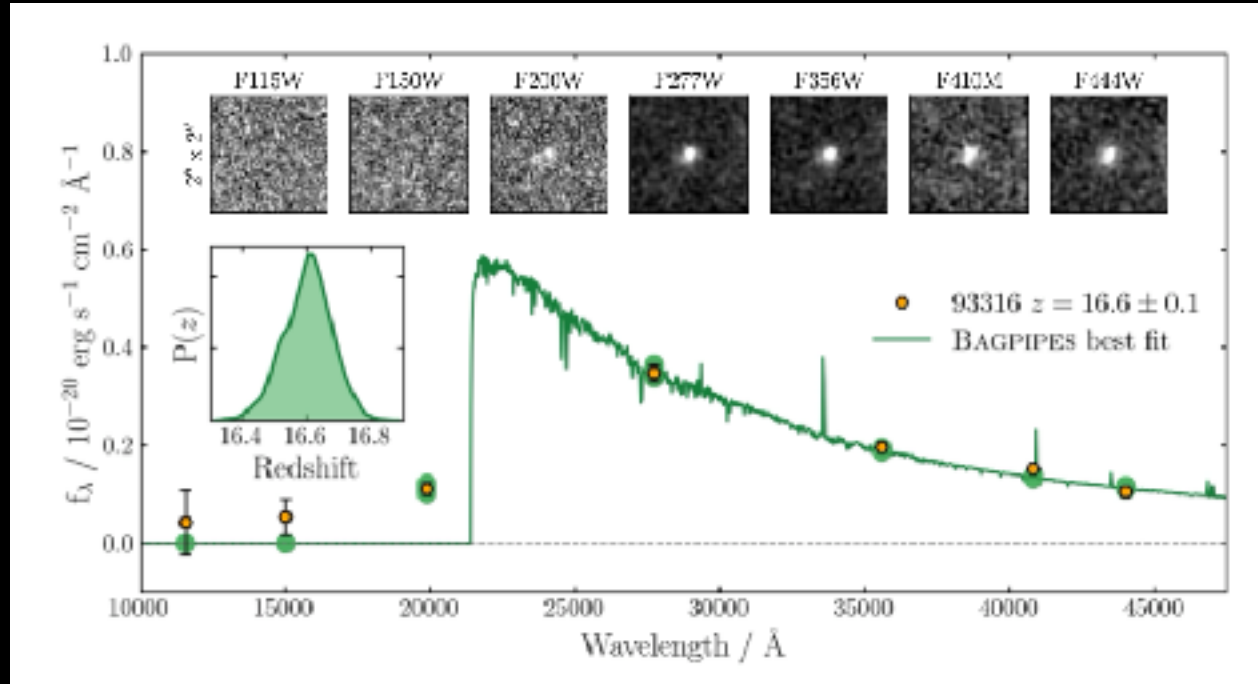




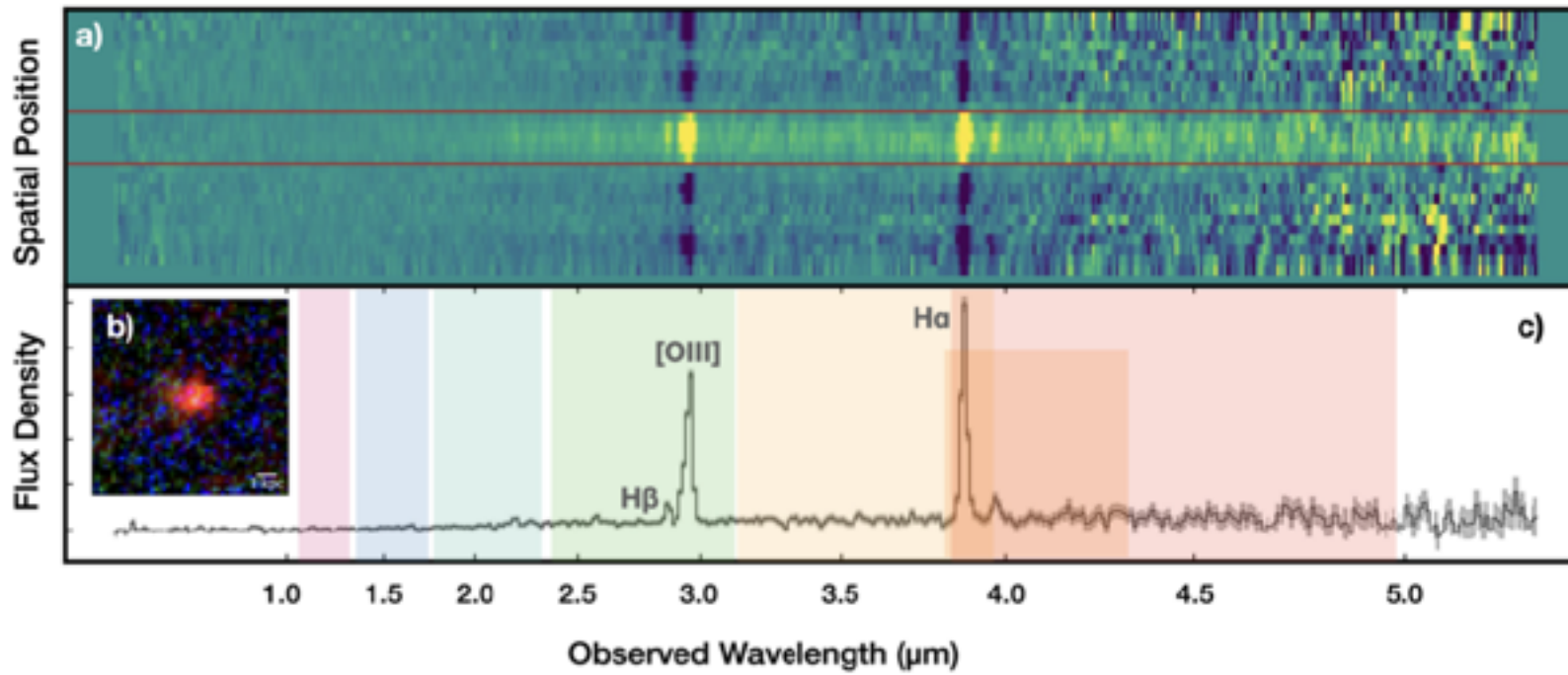


# Surprise#2b: You can't win them all...

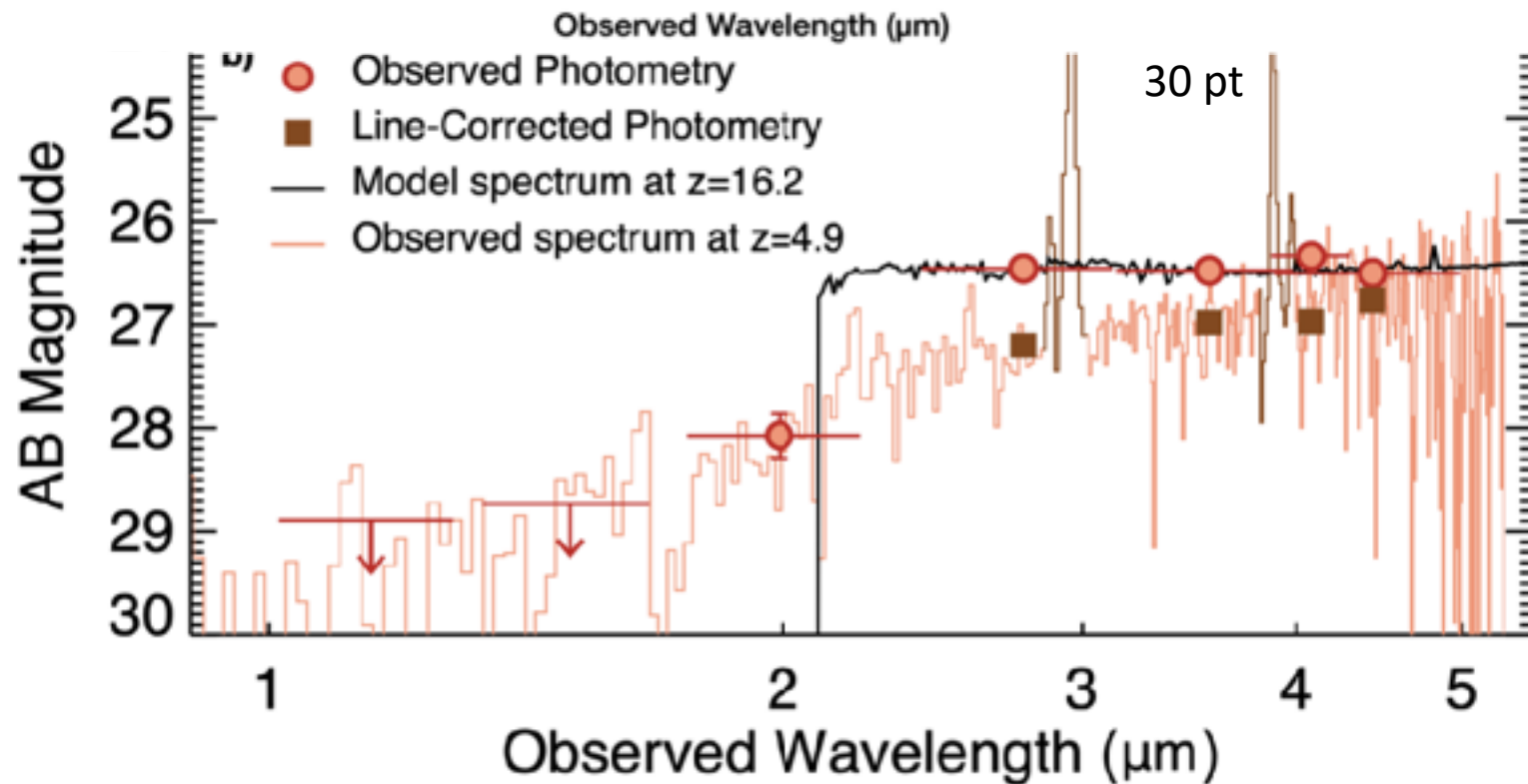
The best  $z \sim 16$  candidate so far...







Is actually at  $z \sim 5.5$ ...



Arrabal-Haro+,  
2303.15431

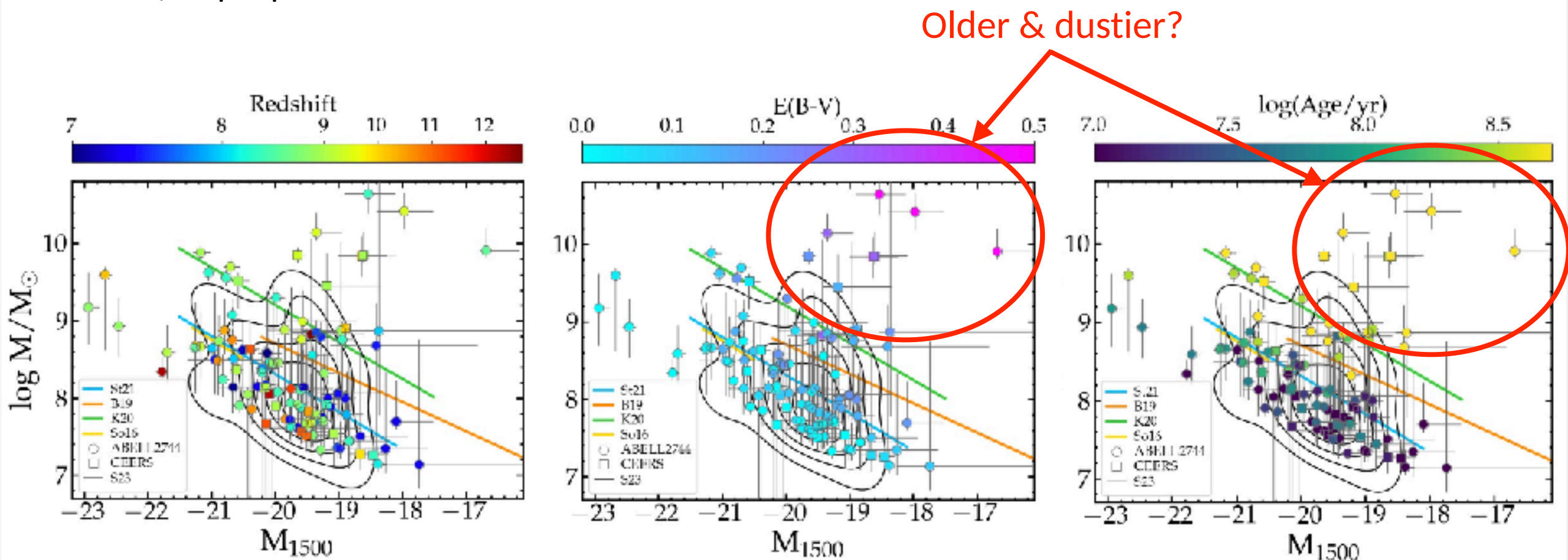


# Surprise#3: Galaxies at $z > 8$ come in a variety of flavours.. ...indicative of different evolutionary status

Santini, AF+22

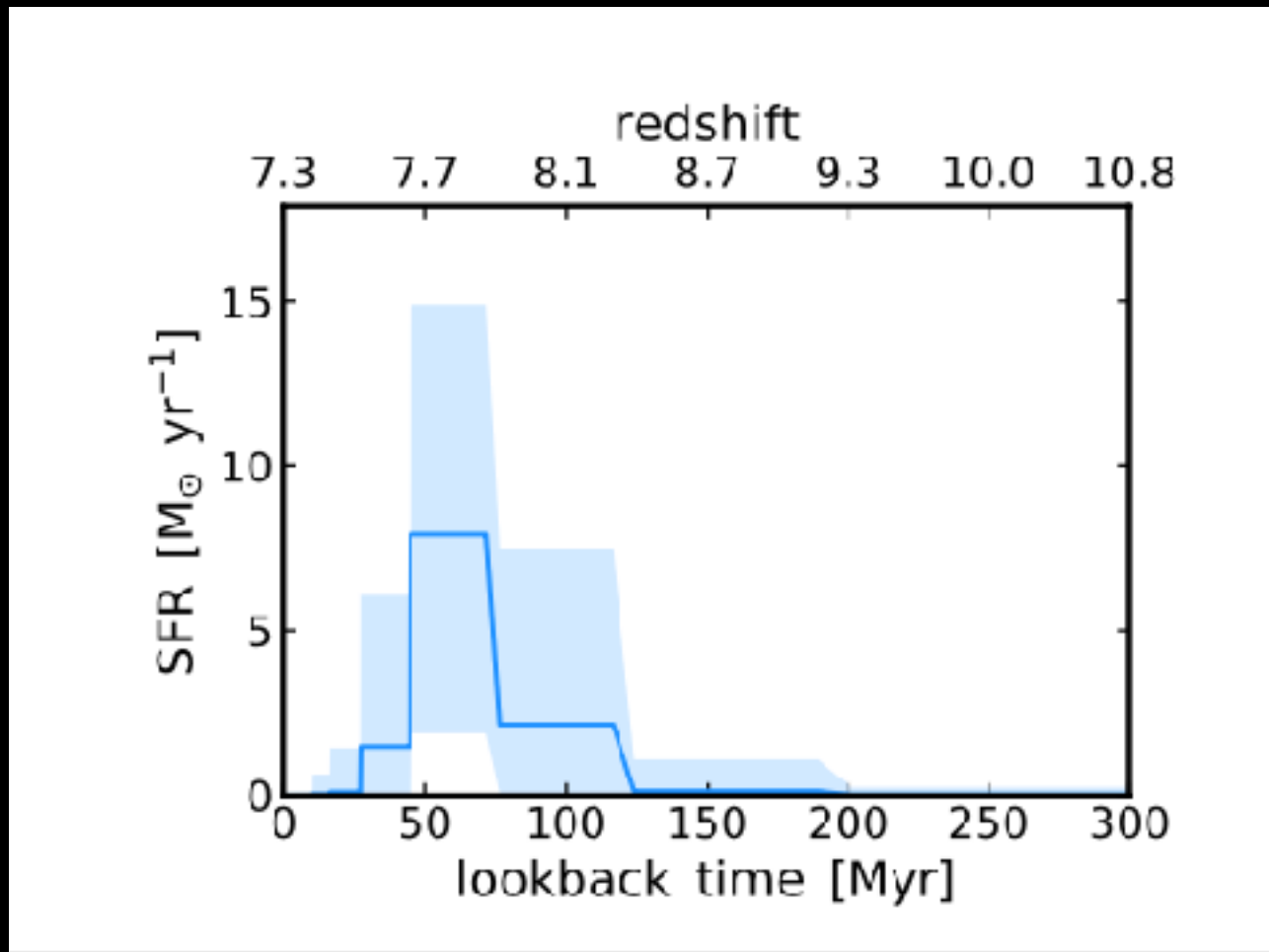
Merlin+, in prep.

## THE MASS – UV LUMINOSITY RELATION



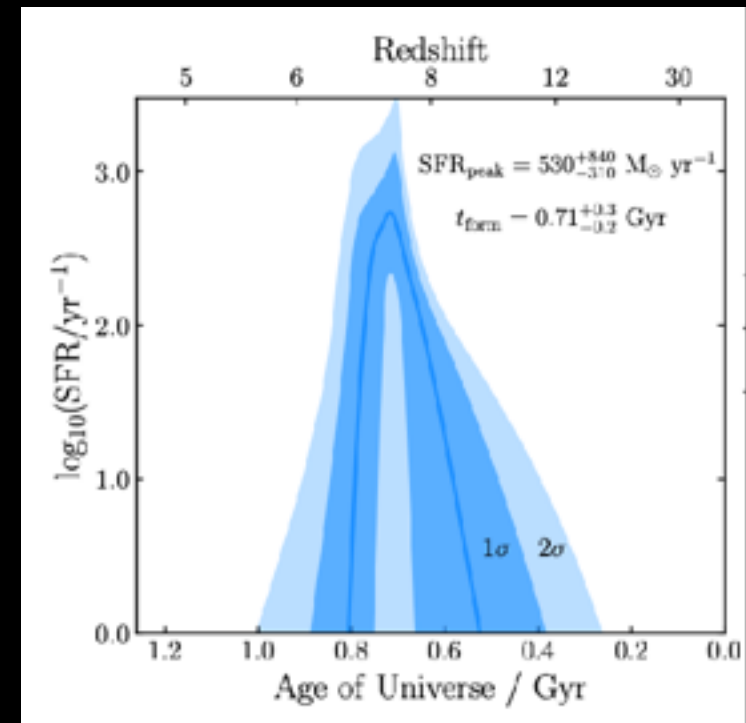
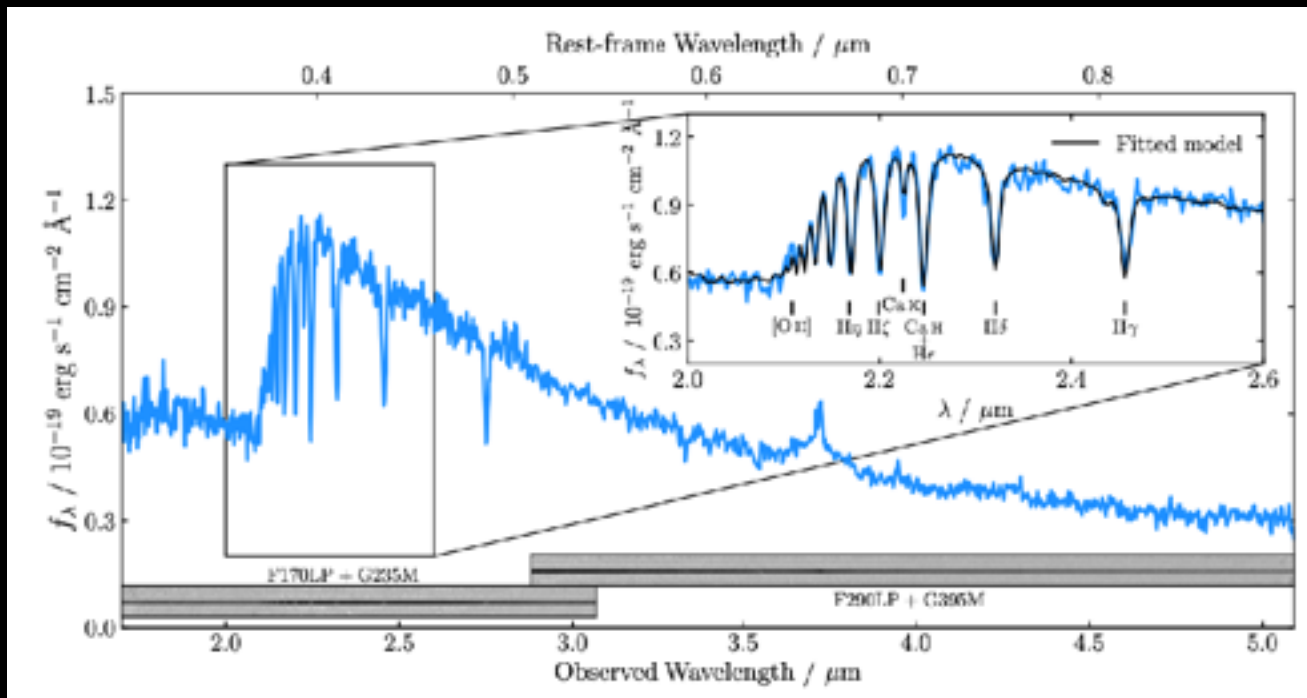
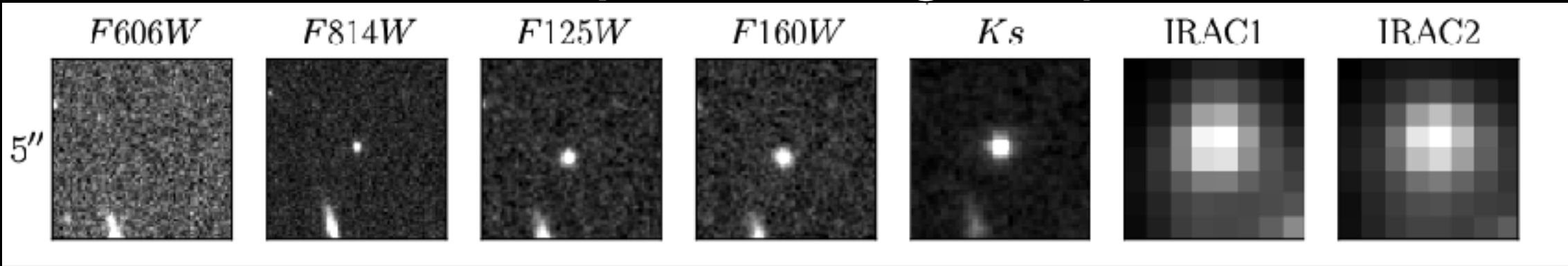


# A recently quenched galaxy at $z \sim 7.3$





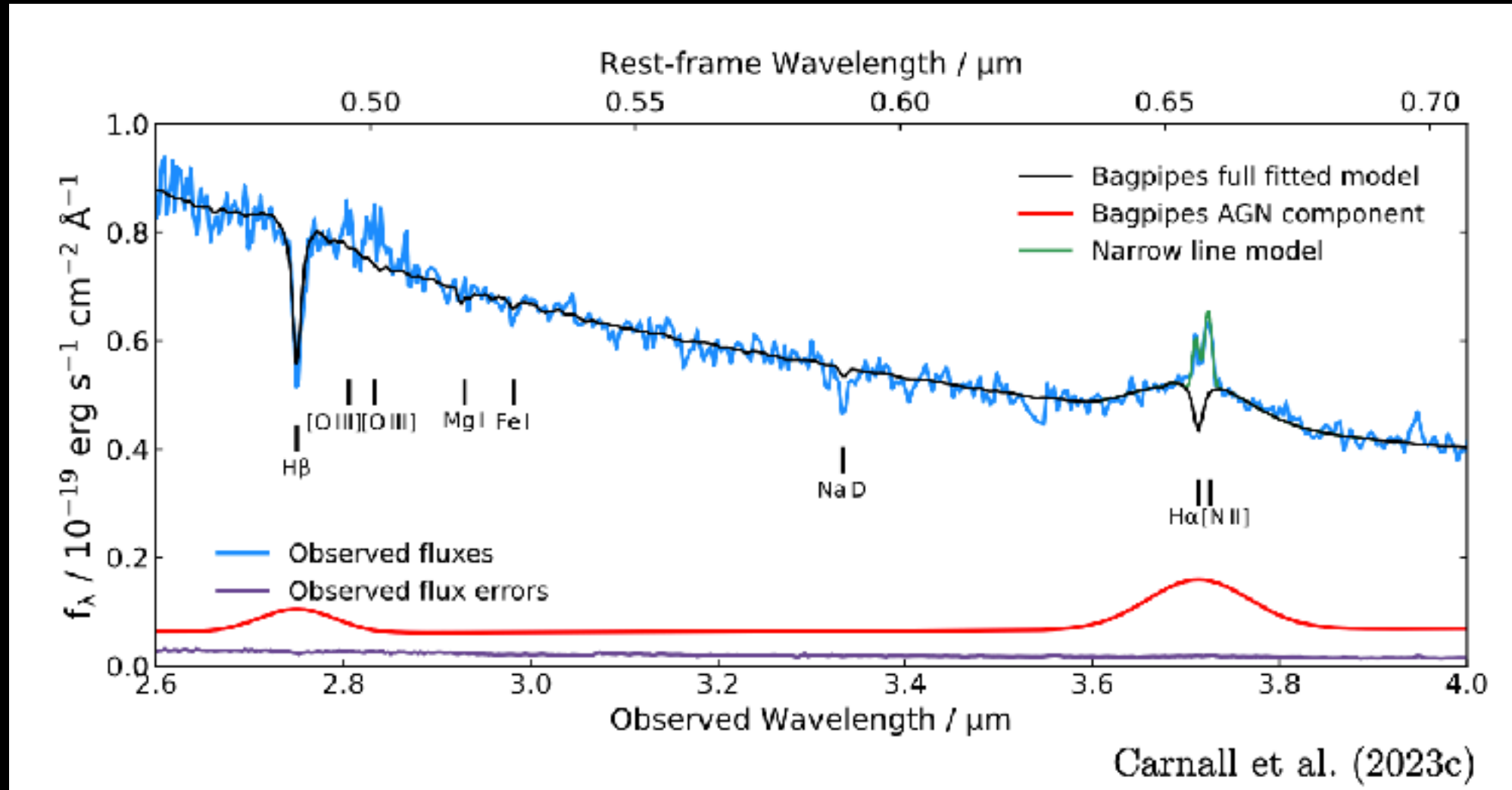
# A massive quiescent galaxy at $z=4.7$



# Carnall+2301.11413

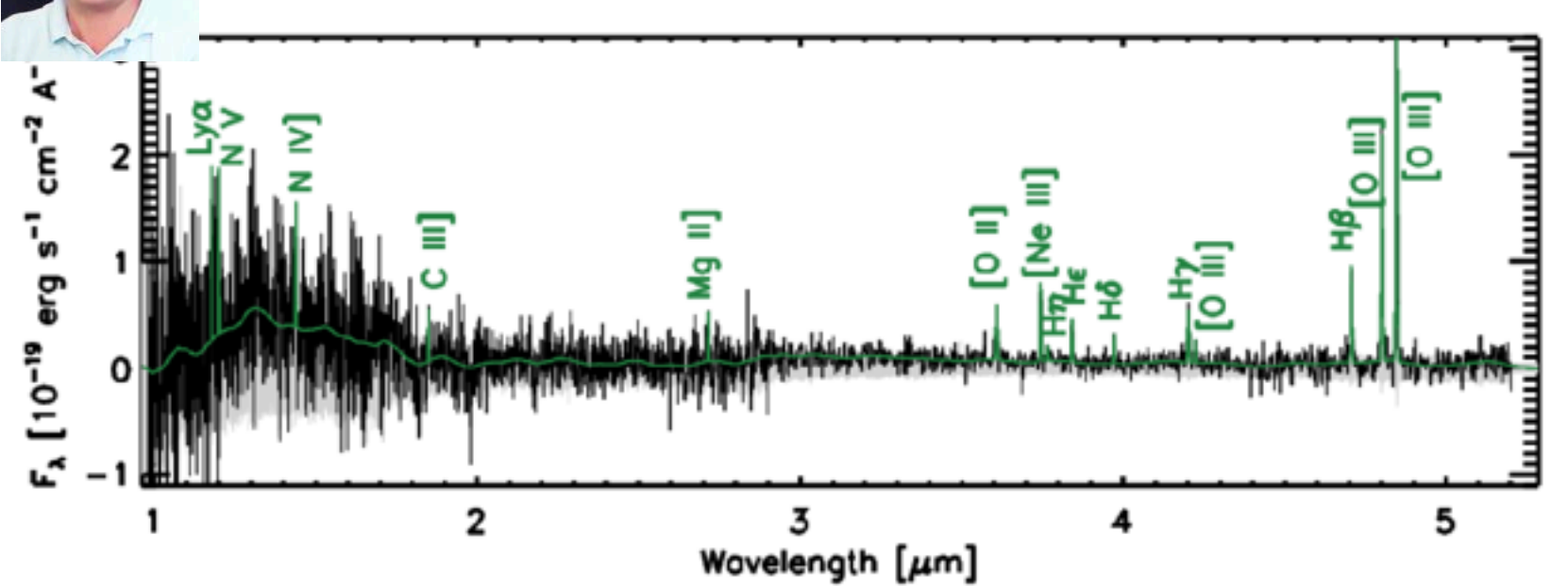
Age of the Universe=1.3Gyr

# A massive quiescent galaxy **with AGN** at $z=4.7$



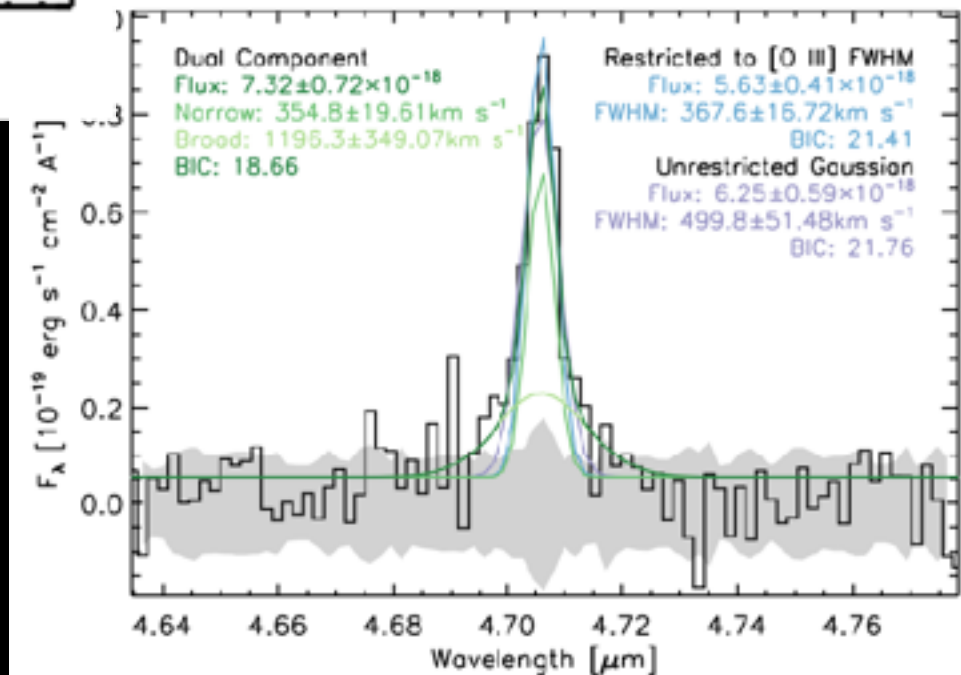


## Surprise#5: A lot of AGNs....



An AGN at  $z \sim 8.67$

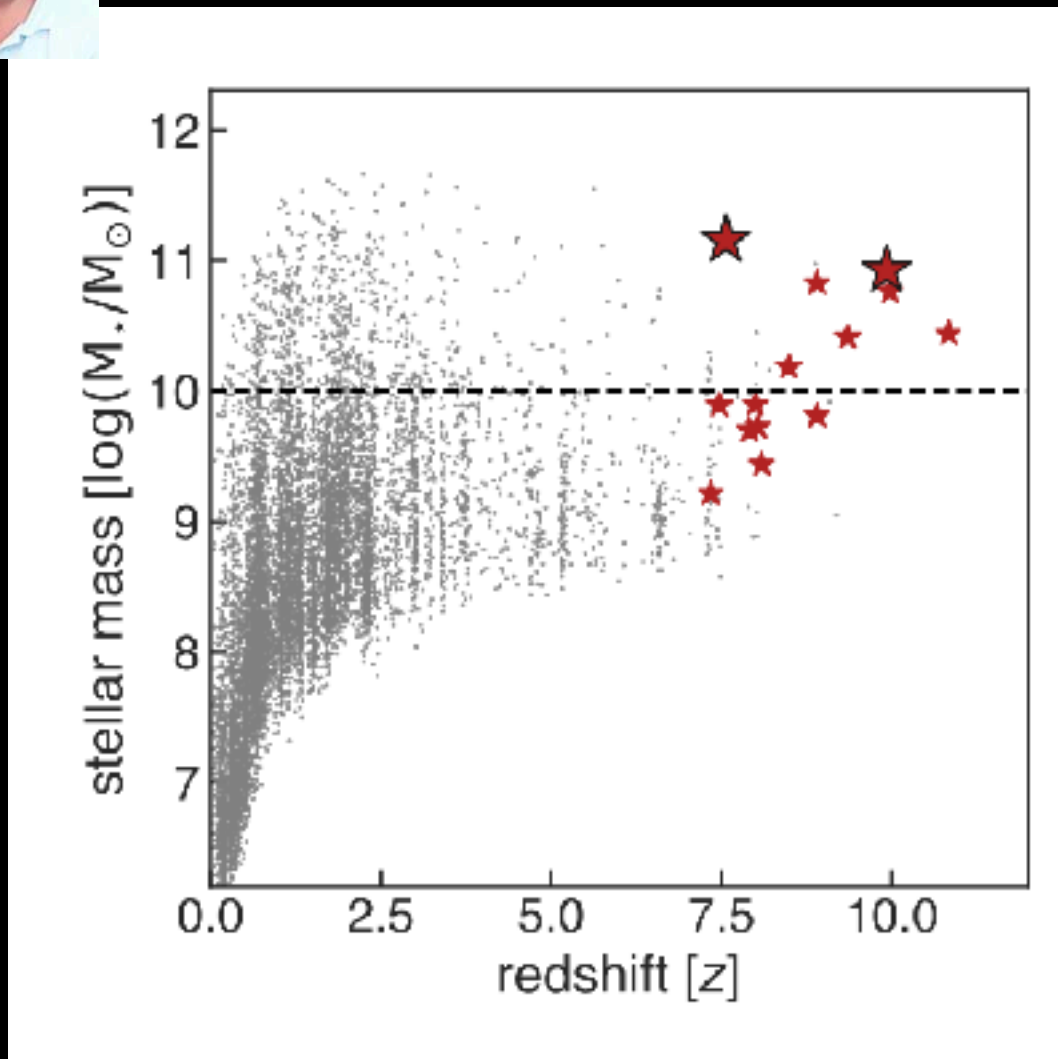
Larson+23, 2303.08918



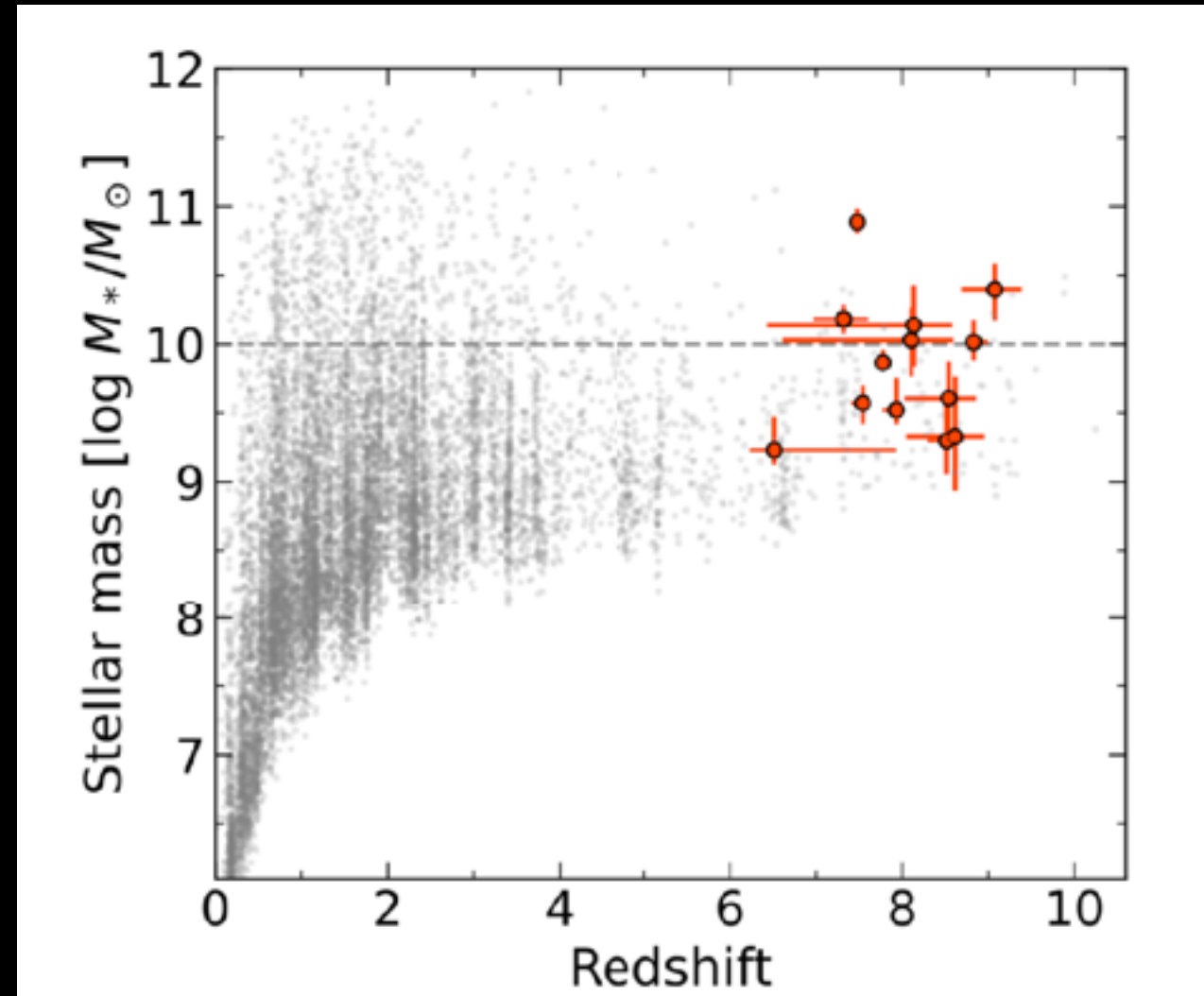




# Surprise#6: Massive Galaxies in the 1Gyr...

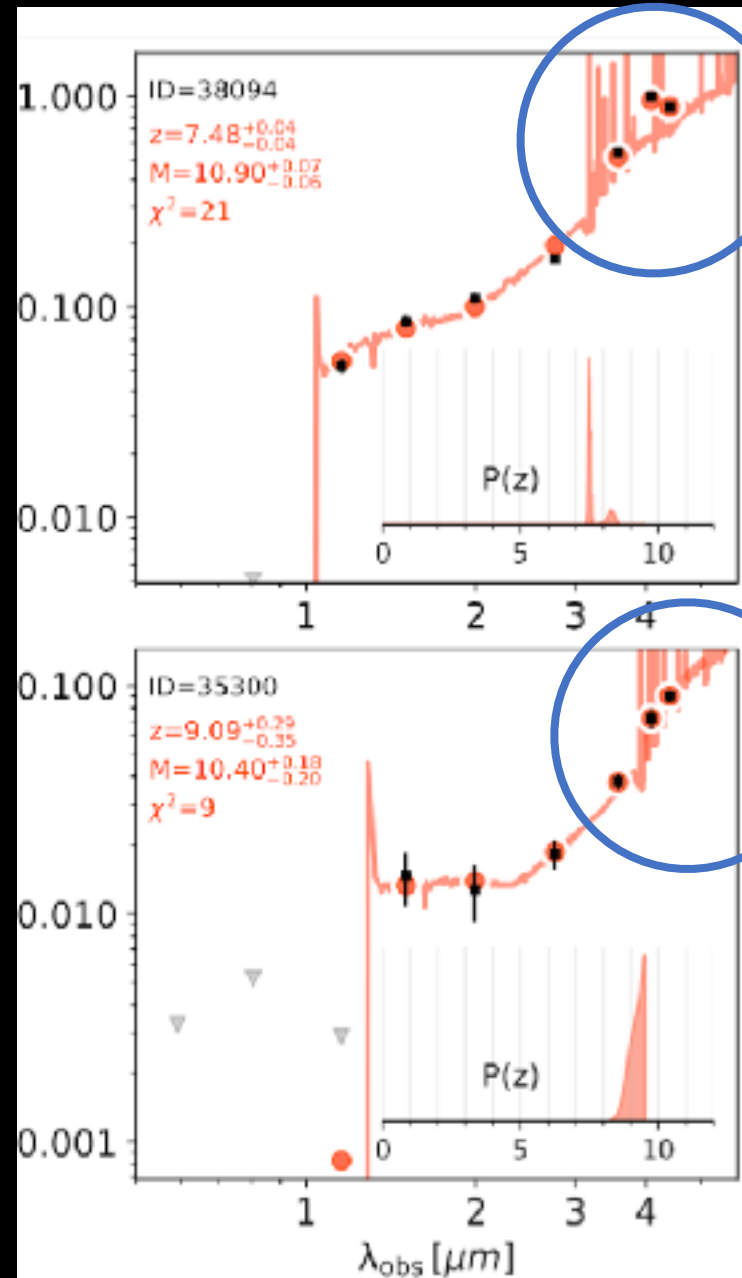


Labbe+23,2207.12446 Version 1



Labbe+23 Version 2 (Nature)

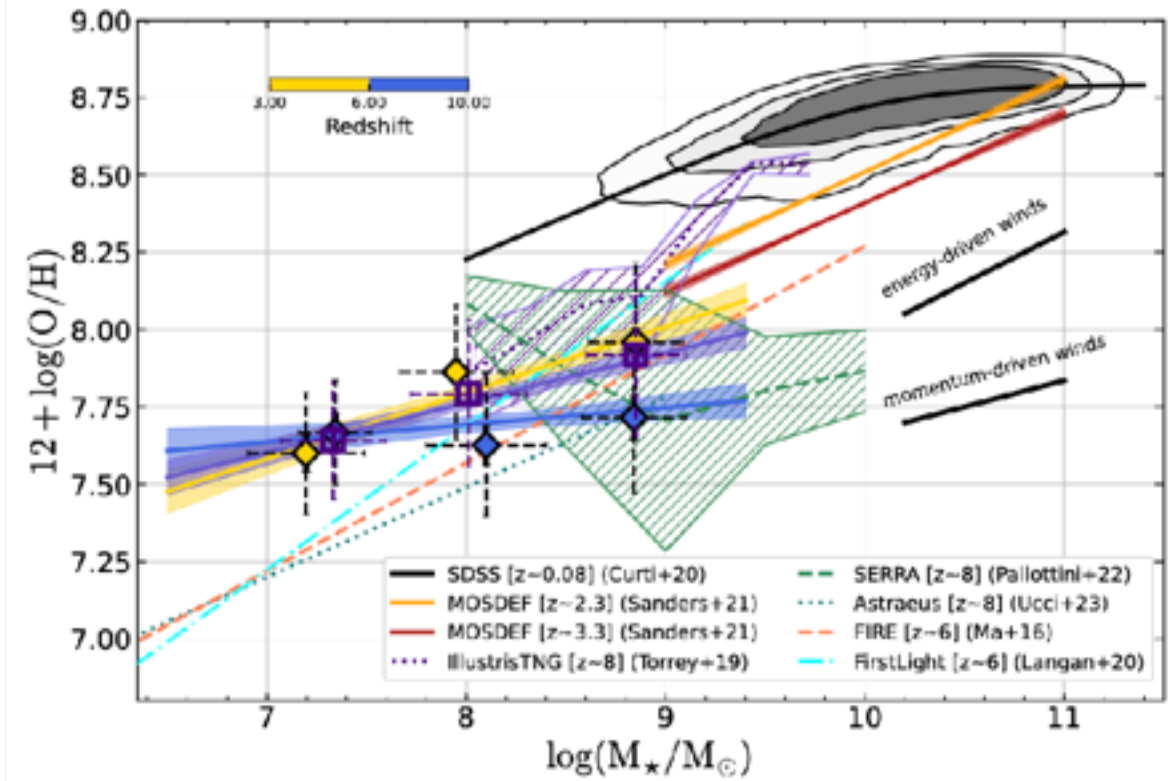
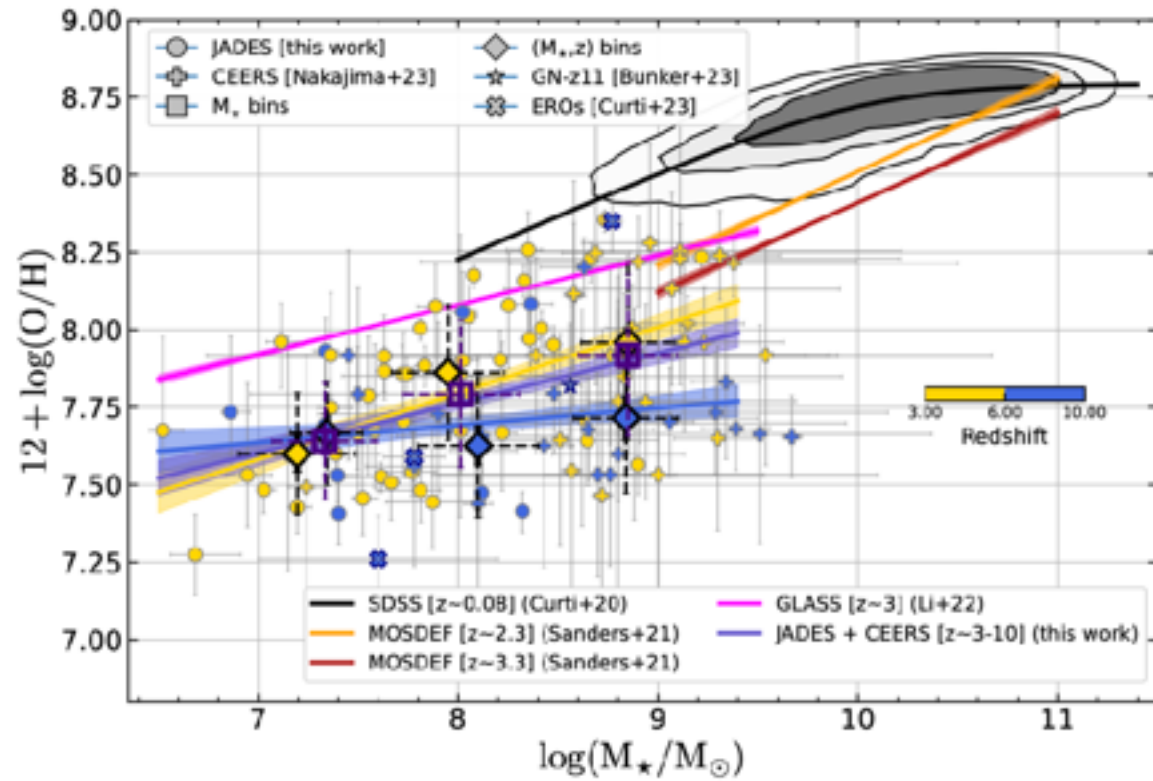
# Surprise! Six Massive Galaxies At $z \sim 8$ in the 1 Gyr...



The SED suggests the presence of a hidden AGN which biases upward both the stellar mass and the redshift



# Surprise#7: Metallicity evolution is lower than expected



Curti, Maiolino+23, 2304.08516

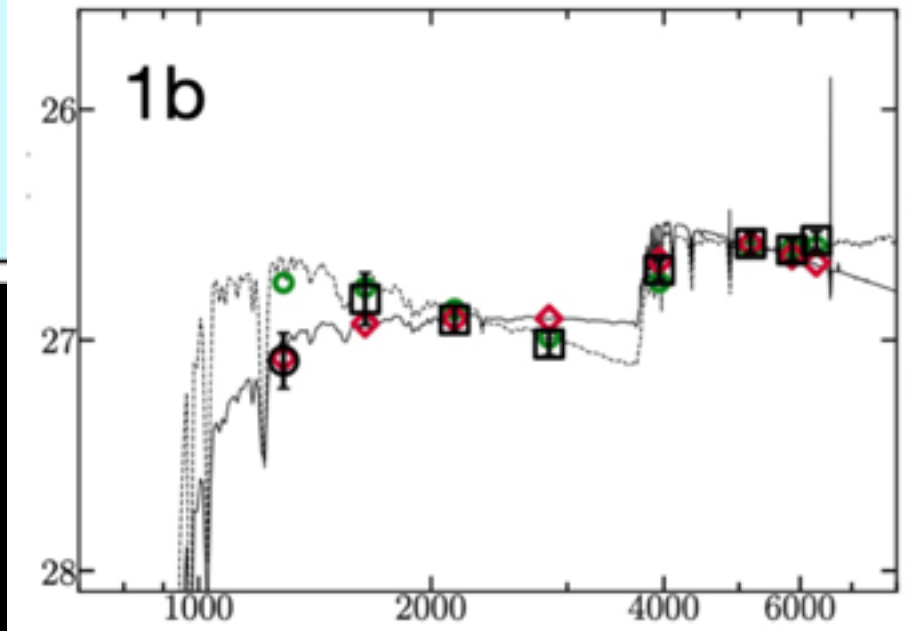
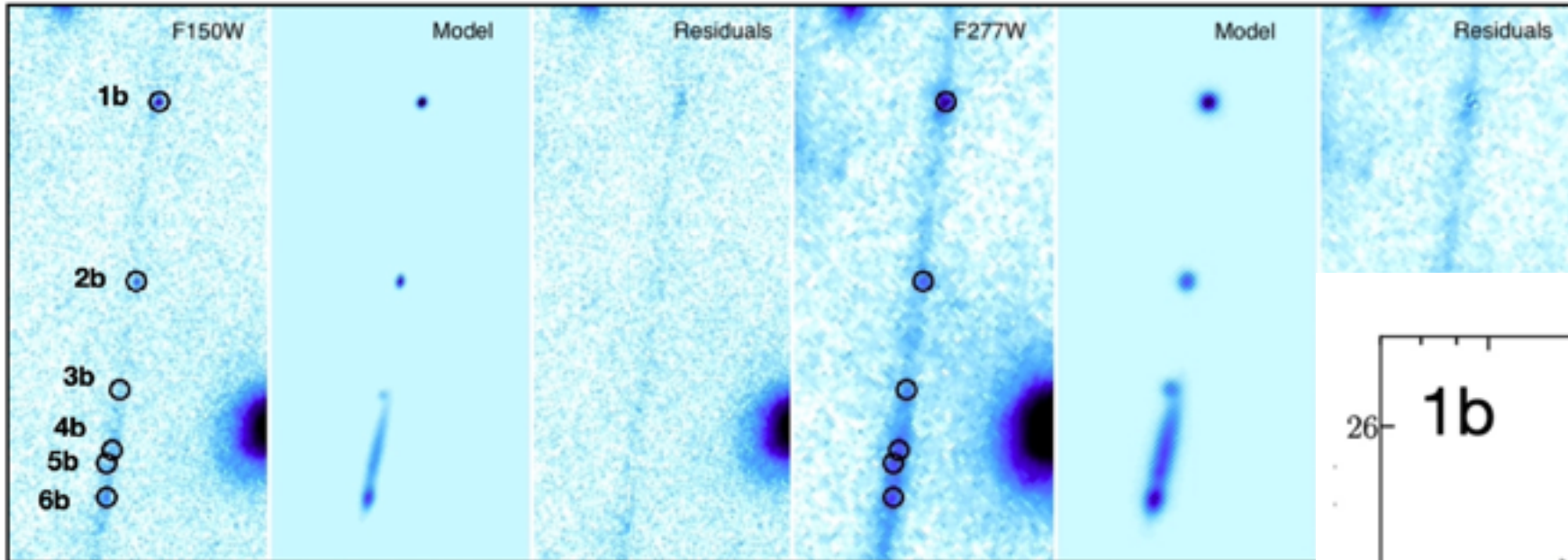




# Surprise#8: We detect proto-globular clusters at $z \sim 6$ !

JWST PROBES STAR CLUSTERS AT  $z \simeq 6$

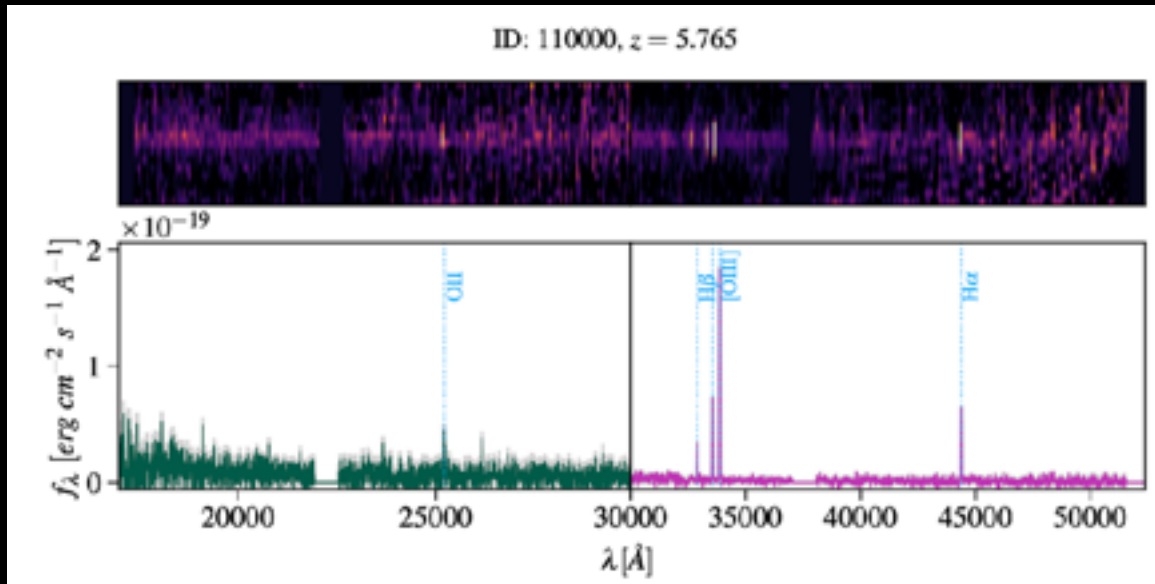
7



Vanzella, Castellano+23

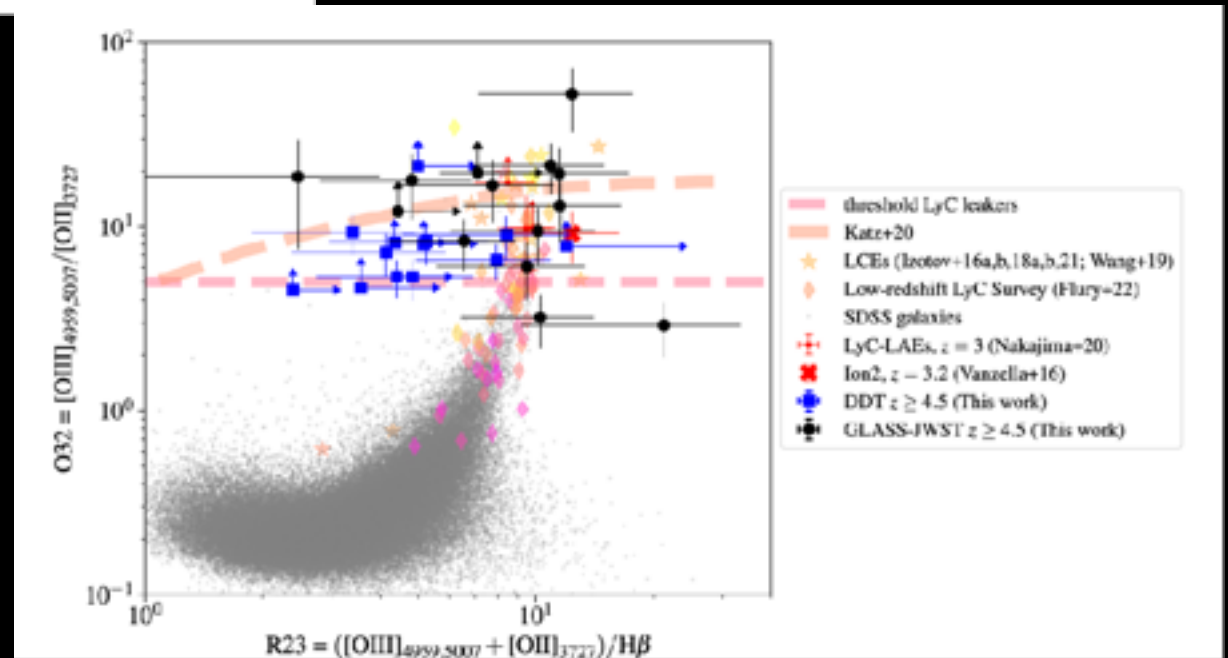


# Surprise#9: Closing in on the sources of cosmic reionization

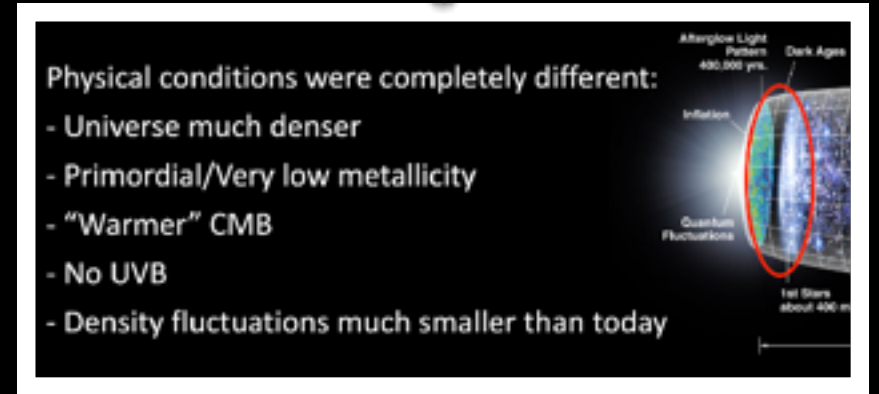
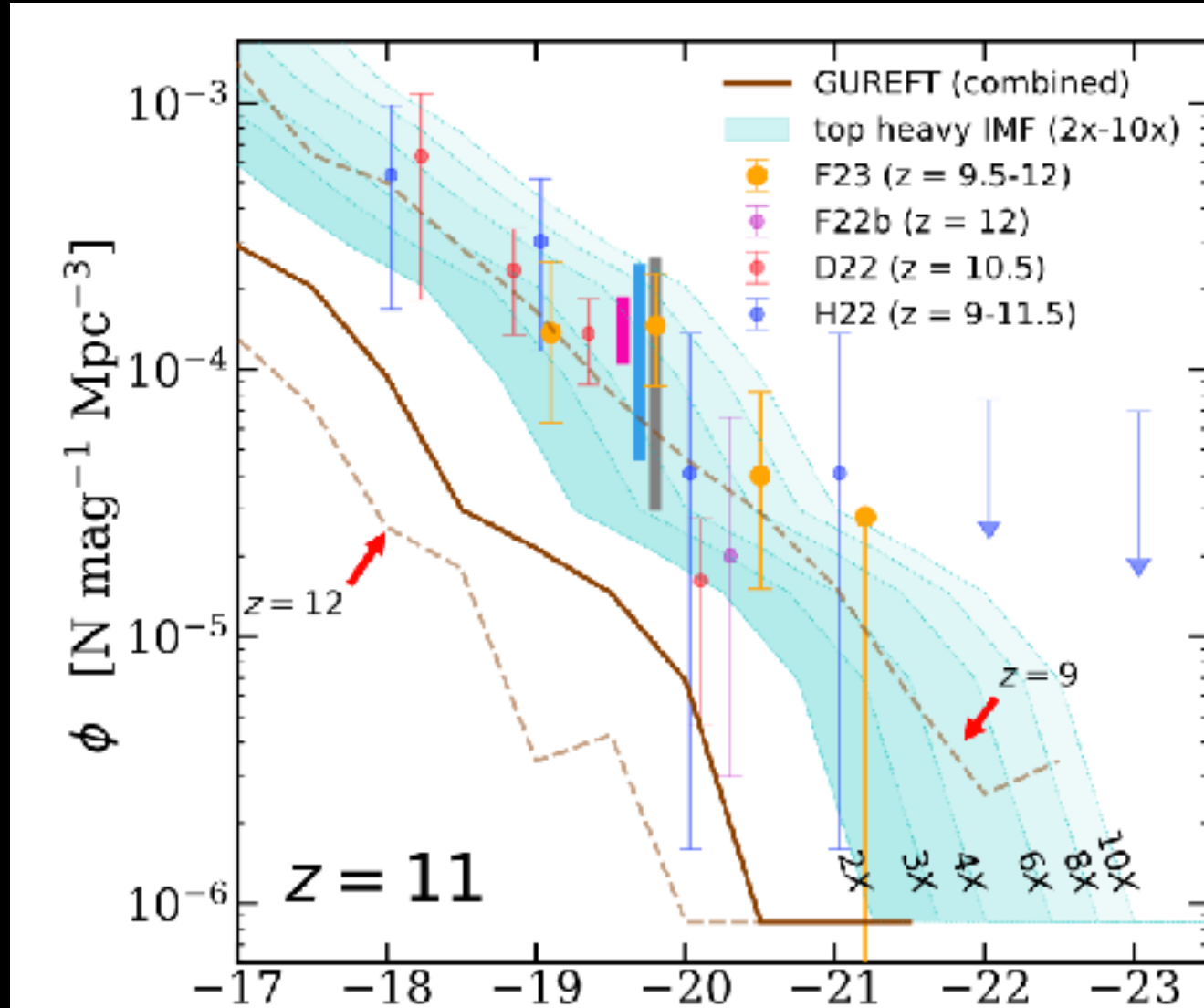


Spectroscopic evidence that low-mass galaxies are LyC leakers and contribute to reionisation

Mascia, Pentericci+23



# These early surprises are challenging our theoretical models.



- Top Heavy IMF?
- Very low Z / PopIII?
- Stronger feedback from SN/AGN?



# Is there space for “new physics”?

## High-redshift Galaxies from Early JWST Observations: Constraints on Dark Energy Models

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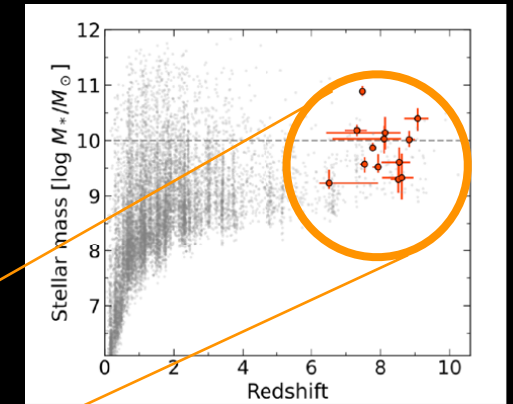
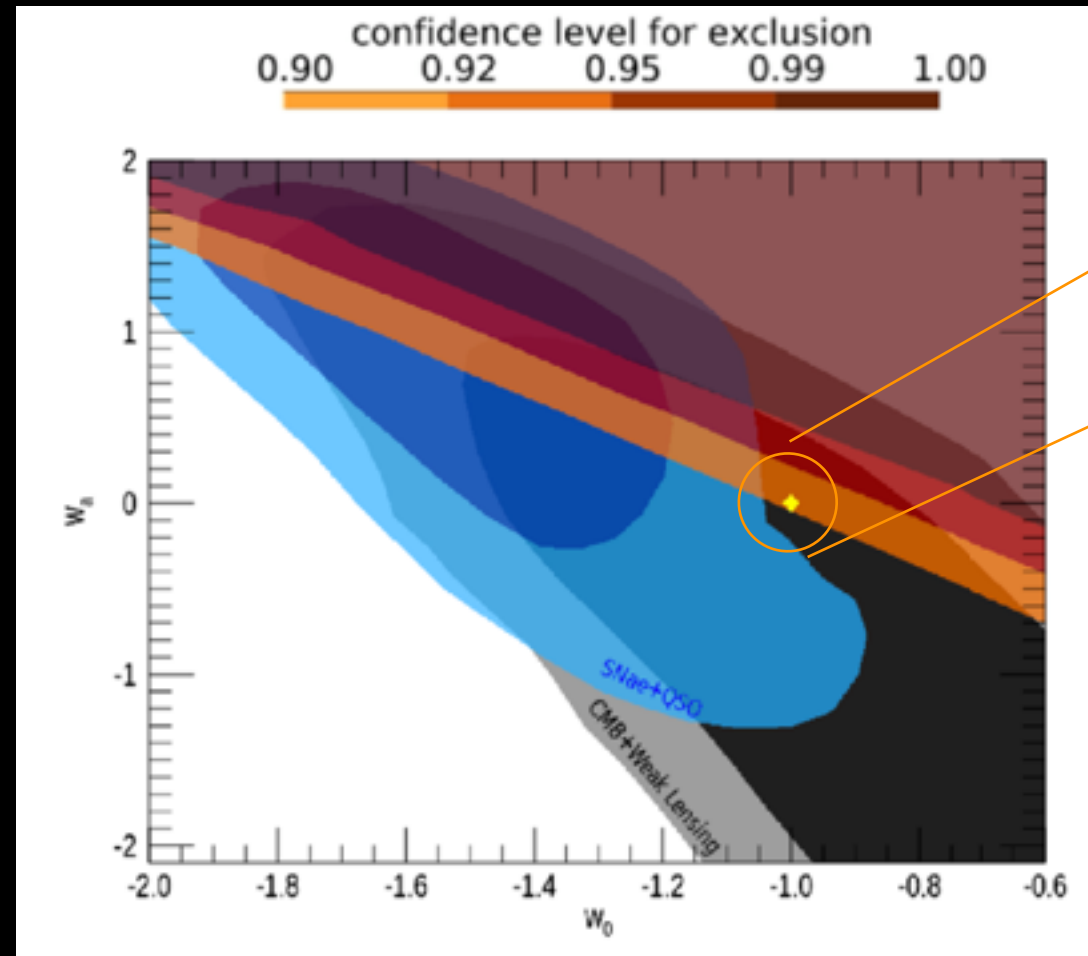
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We need more  
statistics and  
bulletproof results.

*The jury is still  
out...*



Labbe+23