



A scientific guide to Gaia Astrophysical Parameters

organized by Coordination Unit 8



"A scientific guide to Gaia Astrophysical Parameters" Day XX





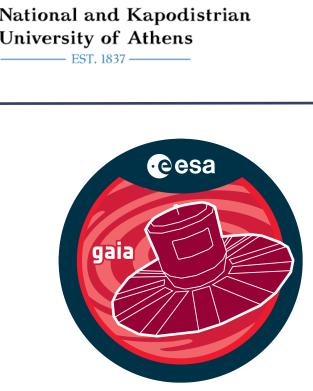




Royal Observatory *of* Belgium



University of Athen



Day 3

The Discrete Source Classifier in Gaia DR3

Sara Jamal, Coryn Bailer-Jones Max Planck Institute for Astronomy, Heidelberg

CU8 webinar, March 2023

- I. Extragalactic results in GDR3
- 2. The Discrete Source Classifier (DSC) in GDR3: design and performance
- 3. The extragalactic candidate tables in GDR3
- 4. Improving DSC in GDR4



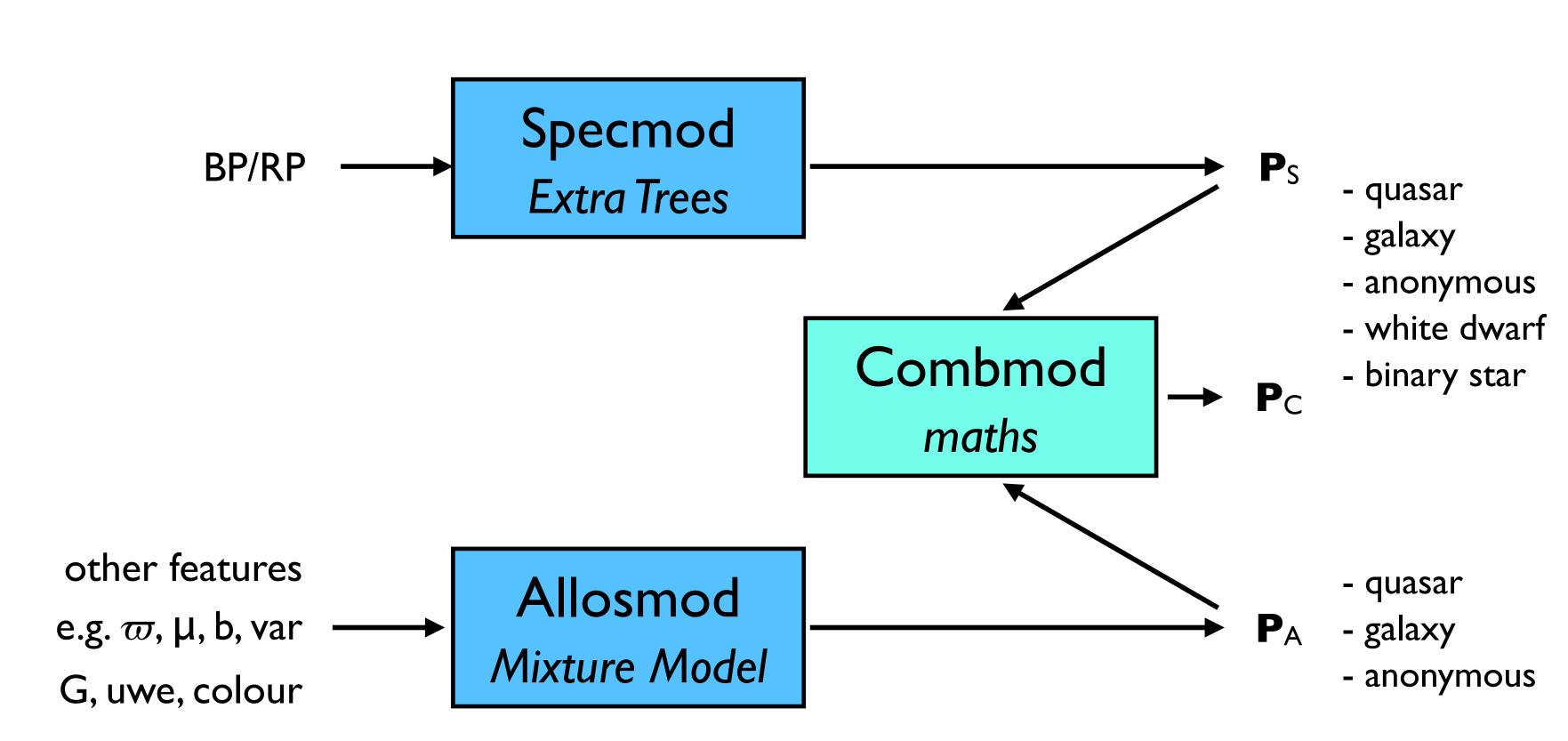
Extragalactic results in GDR3

Extragalactic results in GDR3

- Classification of Gaia objects
 - using BP/RP spectrum and astrometry (Discrete Source Classifier, DSC)
 - using photometric light curves (Vari)
- Input lists
 - fit 2D brightness profiles to extended objects (Extended Objects, EO)
 - identify objects from astrometry, similar to Gaia-CRF3
- Redshift estimates
 - quasars (QSOC) and galaxies (UGC)
- Link: Gaia Collaboration (Bailer-Jones et al.) 2023

The Discrete Source Classifier (DSC)

The Discrete Source Classifier (DSC)



Link: DSC overview Link: DSC details

A probabilistic classifier

if

 $P_{C}(quasar) > 0.5$

then

classlabel dsc = "quasar"

if

 $P_{s}(quasar) > 0.5 \& P_{A}(quasar > 0.5)$ then

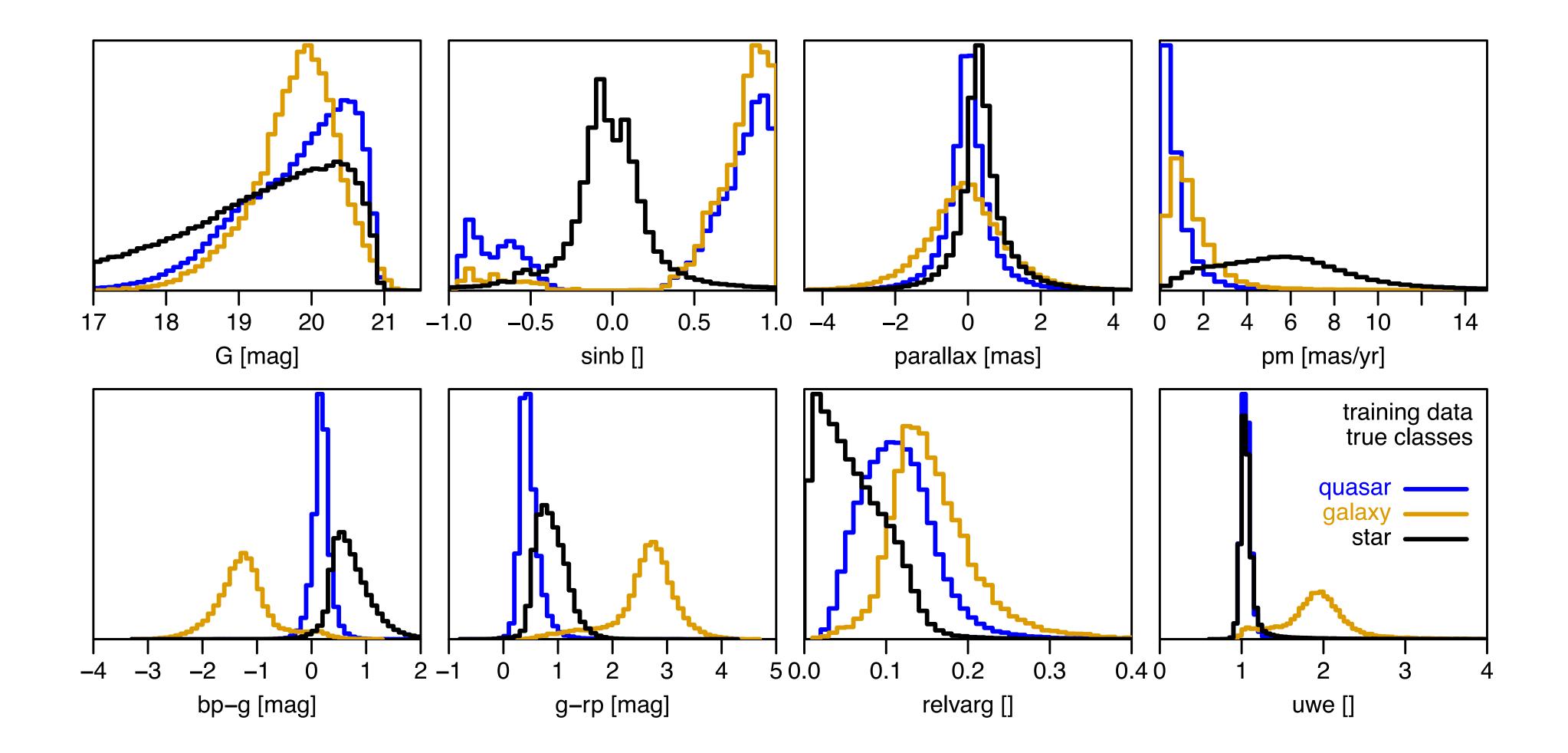
classlabel_dsc_joint = "quasar"

and similarly for "galaxy"

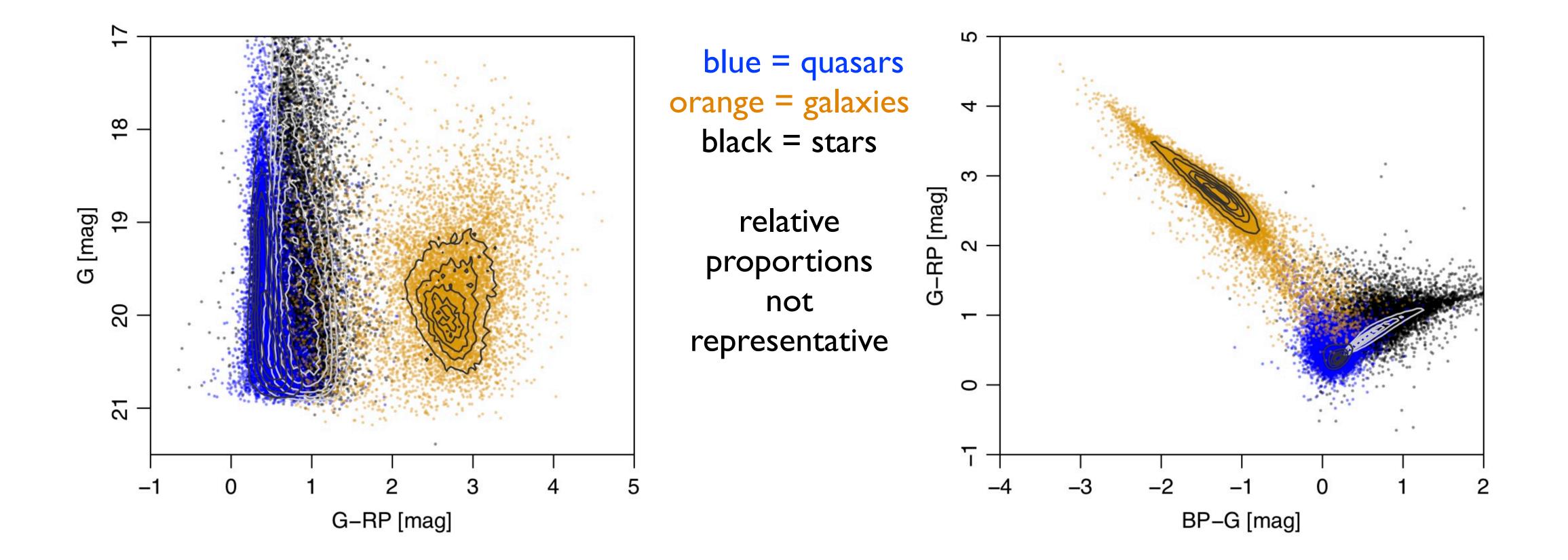




DSC: Allosmod training features

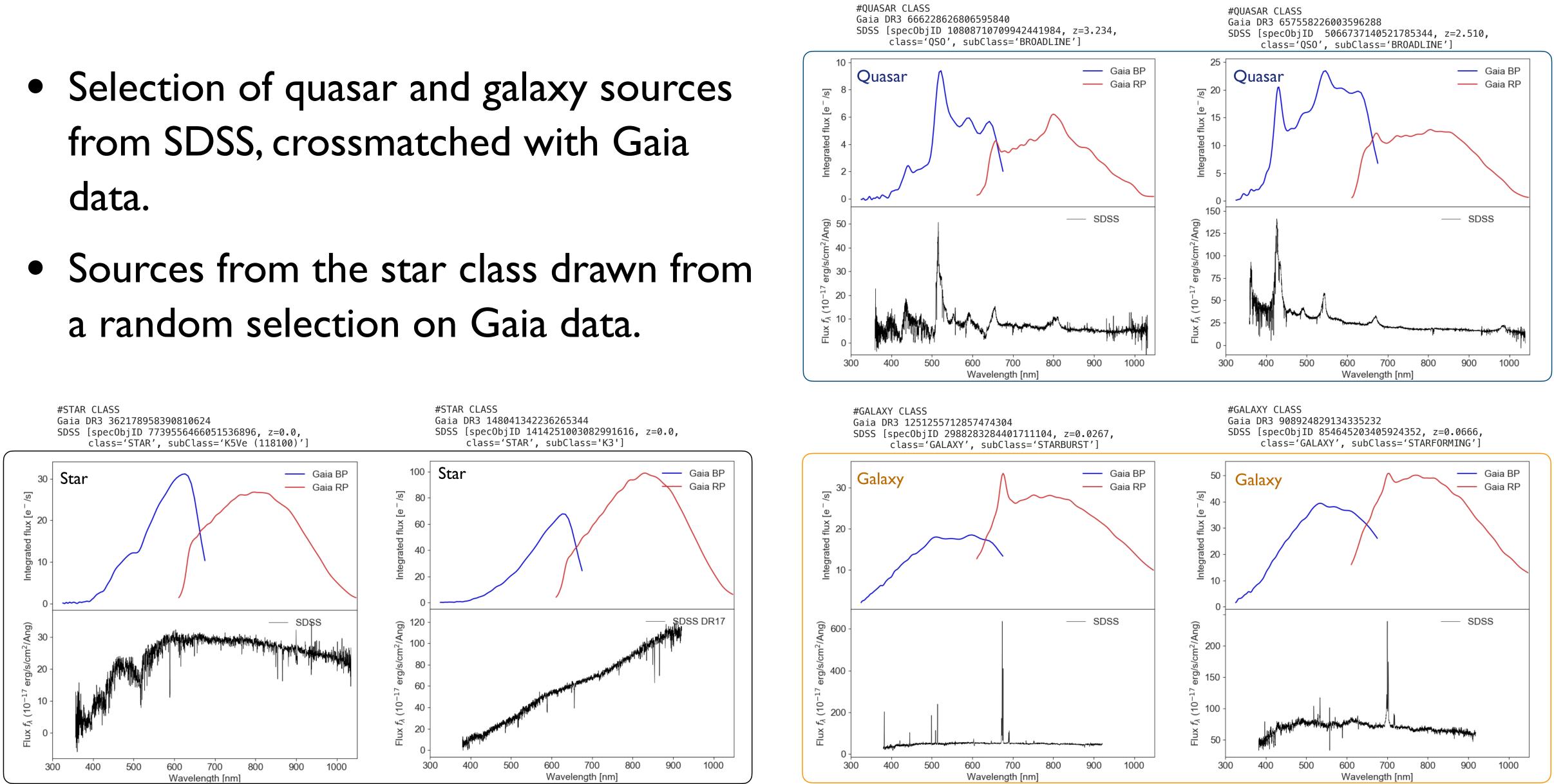


DSC: CMD and CCD of training data



DSC: XP spectra of training data

- from SDSS, crossmatched with Gaia data.
- a random selection on Gaia data.



DSC: Class prior

| | quasar | galaxy | star | white dwarf | physical binary star |
|---|----------|----------|----------|-------------|----------------------|
| x | 1/1000 | 1/5000 | 1 | 1/5000 | 1/100 |
| = | 0.000989 | 0.000198 | 0.988728 | 0.000198 | 0.009887 |

The prior must also be taken into account when estimating the purity on validation data because validation data usually does not have enough stellar contaminants. See <u>Bailer-Jones et al. (2019; MNRAS 490, 5615</u>) section 3.4

Important

| | Specmod | | Allosmod | | Combmod | | Spec&Allos | |
|--------|---------|--------|----------|--------|---------|--------|------------|--------|
| | compl. | purity | compl. | purity | compl. | purity | compl. | purity |
| quasar | 0.409 | 0.248 | | | | | | |
| galaxy | 0.831 | 0.402 | | | | | | |
| star | 0.998 | 0.989 | | | | | | |

classification by P > 0.5



| | Specmod | | Allosmod | | Combmod | | Spec&Allos | |
|--------|---------------|-------|----------|--------|---------|--------|------------|--------|
| | compl. purity | | compl. | purity | compl. | purity | compl. | purity |
| quasar | 0.409 | 0.248 | 0.838 | 0.408 | | | | |
| galaxy | 0.831 | 0.402 | 0.924 | 0.298 | | | | |
| star | 0.998 | 0.989 | 0.998 | 1.000 | | | | |
| | | | | | | | | |

classification by P > 0.5



classification by P > 0.5

| | Specmod | | Allosmod | | Combmod | | Spec&Allos | |
|--------|---------|--------|----------|--------|---------|--------|------------|--------|
| | compl. | purity | compl. | purity | compl. | purity | compl. | purity |
| quasar | 0.409 | 0.248 | 0.838 | 0.408 | 0.916 | 0.240 | | |
| galaxy | 0.831 | 0.402 | 0.924 | 0.298 | 0.936 | 0.219 | | |
| star | 0.998 | 0.989 | 0.998 | 1.000 | 0.996 | 0.990 | | |

classlabel_dsc



classification by P > 0.5

| | | | | | classlab | el_dsc | classlabel_ | _dsc_joint |
|--------|---------|--------|----------|--------|----------|--------|-------------|------------|
| | Specmod | | Allosmod | | Combmod | | Spec&Allos | |
| | compl. | purity | compl. | purity | compl. | purity | compl. | purity |
| quasar | 0.409 | 0.248 | 0.838 | 0.408 | 0.916 | 0.240 | 0.384 | 0.621 |
| galaxy | 0.831 | 0.402 | 0.924 | 0.298 | 0.936 | 0.219 | 0.826 | 0.638 |
| star | 0.998 | 0.989 | 0.998 | 1.000 | 0.996 | 0.990 | _ | _ |
| | | | | | | | | |

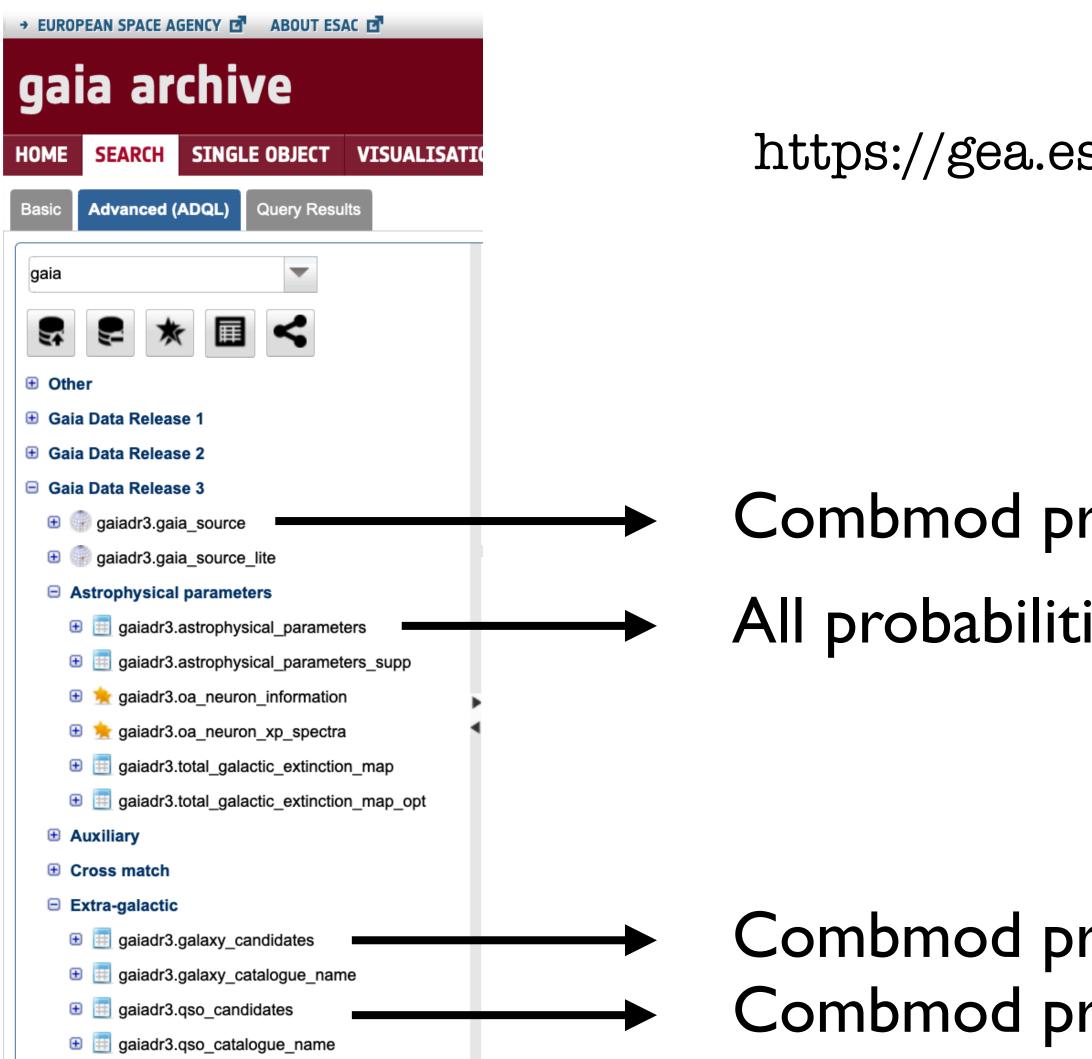


classification by P > 0.5

| | | | | | classlabel_dsc | | classlabel_ | _dsc_joint |
|--------------------------|---------------|-------|----------|--------|----------------|--------|-------------|------------|
| | Specmod | | Allosmod | | Combmod | | Spec&Allos | |
| | compl. purity | | compl. | purity | compl. | purity | compl. | purity |
| quasar | 0.409 | 0.248 | 0.838 | 0.408 | 0.916 | 0.240 | 0.384 | 0.621 |
| galaxy | 0.831 | 0.402 | 0.924 | 0.298 | 0.936 | 0.219 | 0.826 | 0.638 |
| star | 0.998 | 0.989 | 0.998 | 1.000 | 0.996 | 0.990 | — | _ |
| | | | | | | | | |
| quasar, $ \sin b > 0.2$ | 0.409 | 0.442 | 0.881 | 0.603 | 0.935 | 0.412 | 0.393 | 0.786 |
| galaxy, $ \sin b > 0.2$ | 0.830 | 0.648 | 0.928 | 0.461 | 0.938 | 0.409 | 0.827 | 0.817 |
| | | | | | | | | |



Where are the DSC results in the Gaia archive?



https://gea.esac.esa.int/archive/

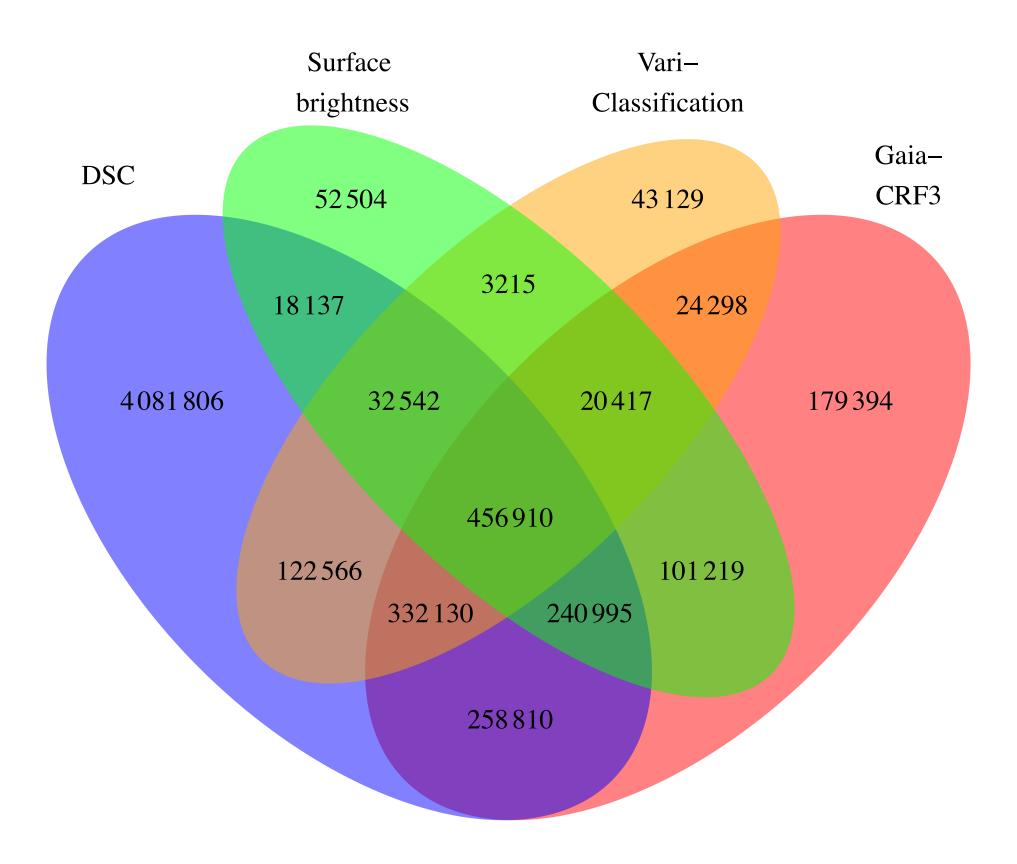
Combmod probabilities; all sources All probabilities and class labels; all sources

Combmod probabilities and class labels; galaxy candidates Combmod probabilities and class labels; quasar candidates

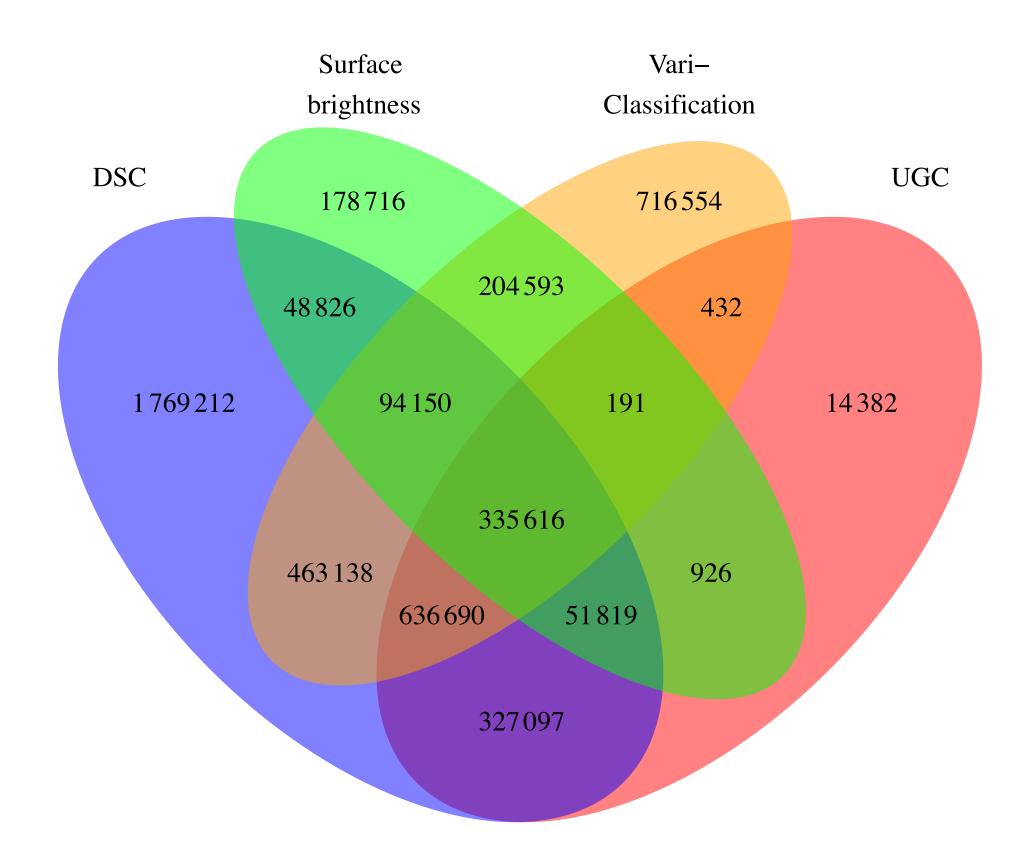
Extragalactic results in Gaia DR3

Contributions to the extragalactic candidate tables

Quasar candidates 6.6 million, 52% pure



Galaxy candidates 4.8 million, 69% pure



Comments on extragalactic candidate tables

- Heterogeneous sample
 - no common definition of "quasar" or "galaxy" across the modules
- Purity and completeness varies among subsets contributed by the modules
 - input lists and Vari driven by purity; DSC driven by completeness
 - higher purity subset achieved with further selections (at cost of completeness)
- If relative contamination is constant, absolute contamination follows source density
- Classification uses only Gaia data (higher purity expected with additional data)

How to select purer subsets in the extragalactic tables

Table 11. ADQL query to select the purer quasar sub-sample.

SELECT * FROM gaiadr3.qso_candidates WHERE (gaia_crf_source='true' OR host_galaxy_flag<6 OR</pre>

Table 12. ADQL query to select the purer galaxy sub-sample.

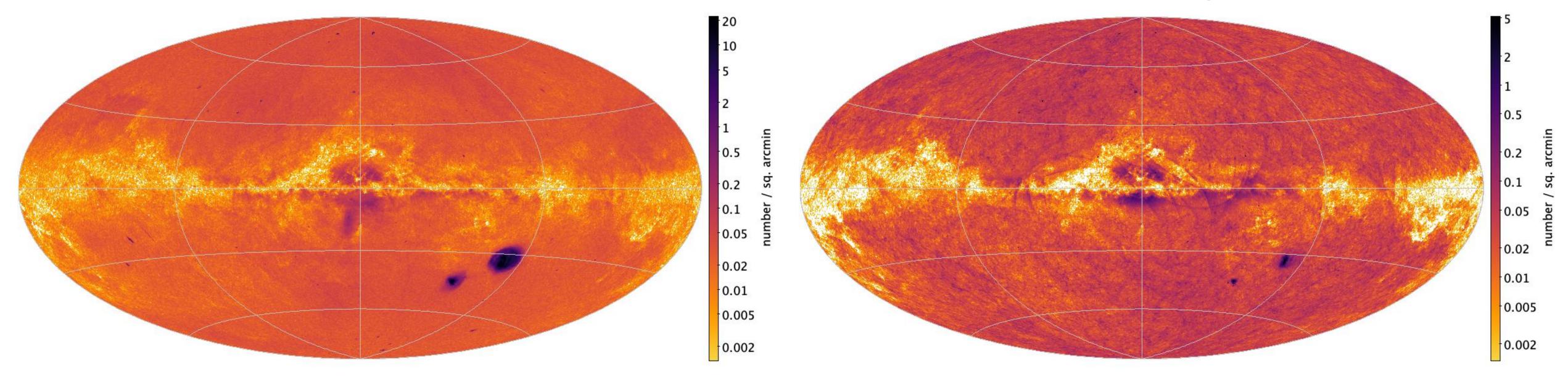
WHERE (radius_sersic IS NOT NULL OR

```
classlabel_dsc_joint='quasar' OR
vari_best_class_name='AGN')
```

```
SELECT * FROM gaiadr3.galaxy_candidates
      classlabel_dsc_joint='galaxy' OR
      vari_best_class_name='GALAXY')
```

Sky distribution of extragalactic candidates Full set

Quasar candidates 6.6 million, 52% pure

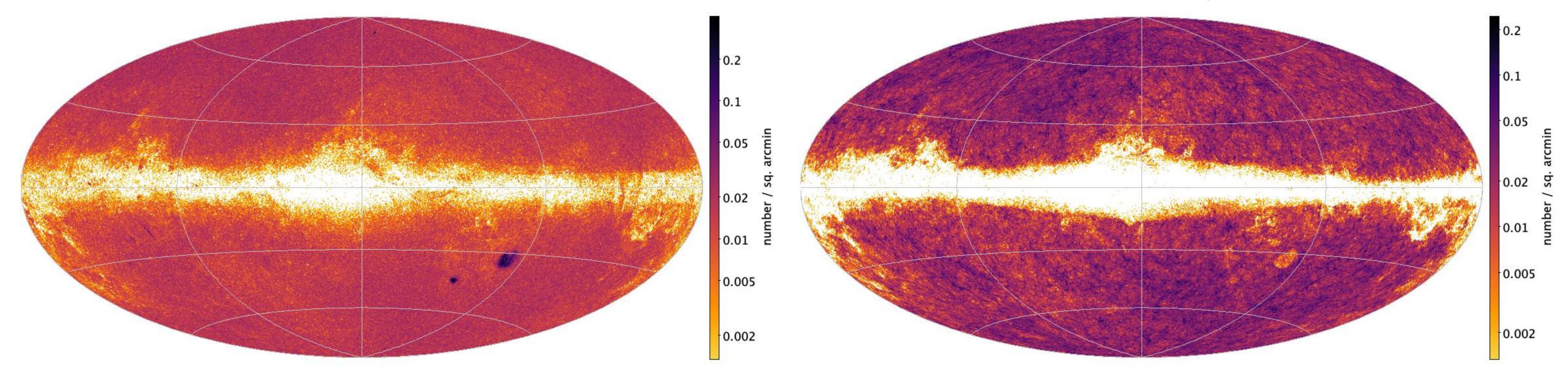


Hammer-Aitoff, Galactic coordinates, HEALpixel level 7 (0.21 sq. deg.)

Galaxy candidates 4.8 million, 69% pure

Sky distribution of extragalactic candidates Purer subset

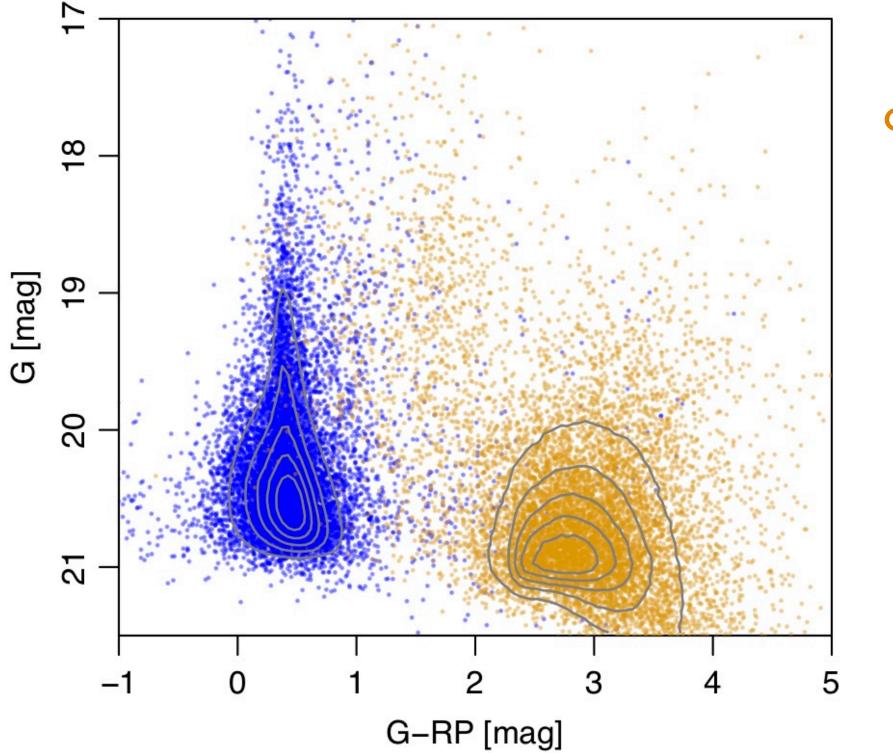
Quasar candidates 1.9 million, 95% pure

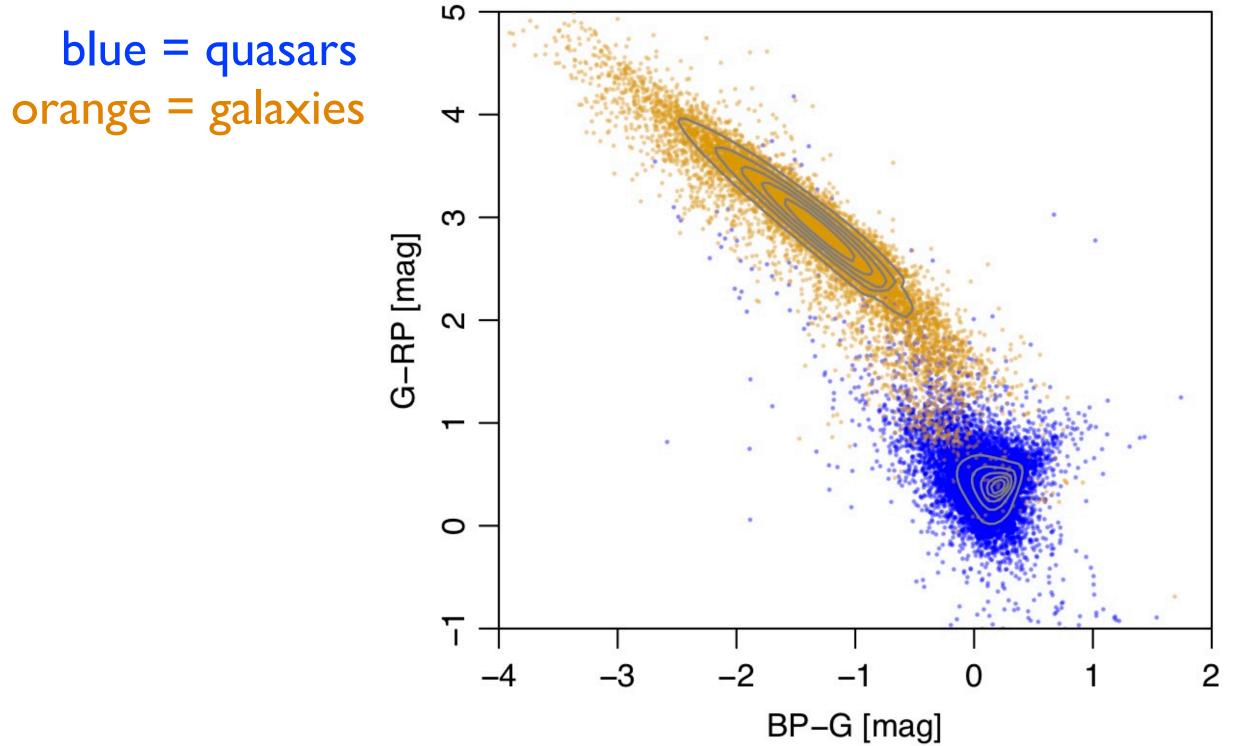


Galaxy candidates 2.9 million, 94% pure

Hammer-Aitoff, Galactic coordinates, HEALpixel level 7 (0.21 sq. deg.)

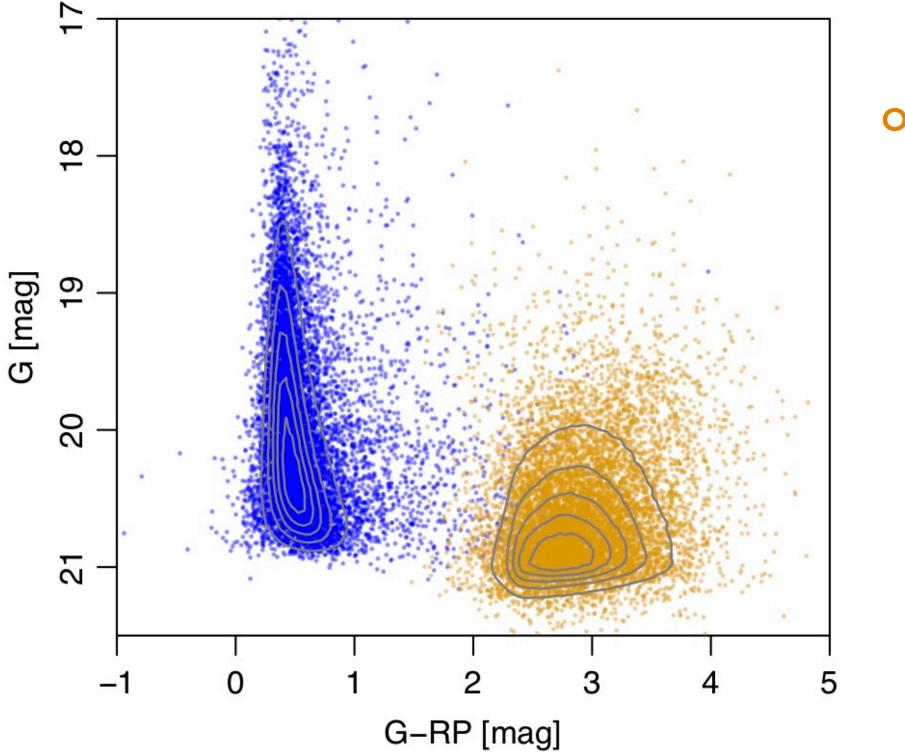
Colour, magnitude distributions of extragalactic candidates Full set

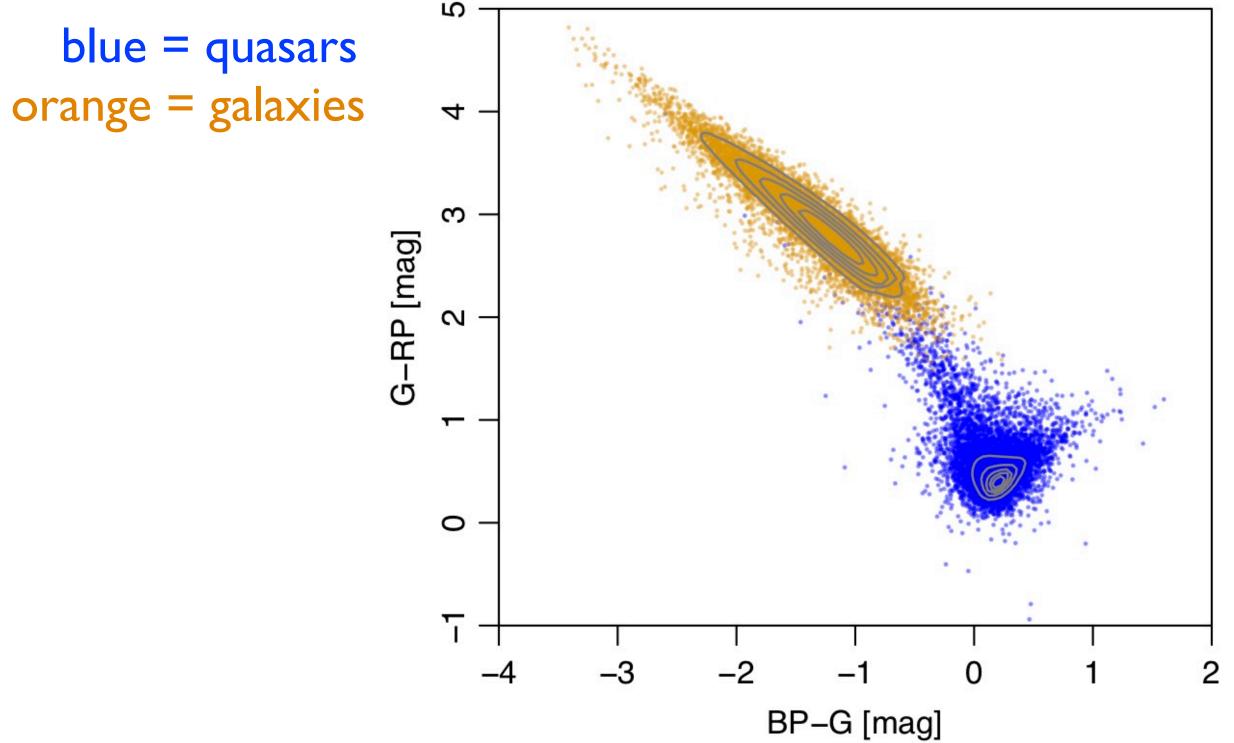




10 000 random sources of each class, linear density contours of full set

Colour, magnitude distributions of extragalactic candidates Purer subset





10 000 random sources of each class, linear density contours of full set

Planned improvements to DSC in GDR4

- Calibration of the posterior class probabilities.
- Algorithms, hyperparameters optimisation.
- Variable priors as function of brightness and galactic latitude [Hughes, Bailer-Jones and Jamal, A&A, 668, A99 (2022)]
- Rejection of contaminants in the training set.
- Evaluation of the classifier performances as function of brightness and galactic latitude.

- DSC classifies sources into three classes (star, quasar, galaxy) using BP/RP spectra, astrometry, colours, and variability information
- DSC extragalactic classes are determined by the SDSS training set
- DSC provides prior-adjusted probabilities
- DSC purity estimates take into account the rareness of extragalactic sources
- The GDR3 extragalactic candidate tables come from several modules, not just DSC
 - Can select a purer subset (~95% pure) using a simple ADQL query with:
 - 1.9 million quasar candidates, 2.9 million galaxy candidates

Summary