Cold Gas Dynamics in High-z Galaxies



Federico Lelli Arcetri Observatory



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Rotation Curves

 \rightarrow Mass models (baryons & DM)

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 - \rightarrow Mass models (baryons & DM)
- Non-circular motions
- \rightarrow Inflows, outflows, bars, spirals

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- \rightarrow Star formation & feedback

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- Gas turbulence
- \rightarrow Star formation & feedback
- Environmental phenomena

 \rightarrow Interactions, mergers, gas stripping

Gas Dynamics across Cosmic Time



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Gas Dynamics across Cosmic Time



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Cold Gas Tracers at high-*z* with ALMA





De Ugarte Postigo+(2012, A&A)

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Cold Gas Dynamics in High-z Galaxies

Cold Gas Tracers at high-*z* with ALMA





High-resolution study of two disk galaxies:

• z**~1.5:** CO(2-1), CO(3-2)

• z≃2.2: CO(3-2), CO(4-3) Lelli+(2023, A&A)

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Two main-sequence galaxies at cosmic noon



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Star-forming main-sequence galaxy at $z \simeq 1.5$

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Star-forming main-sequence galaxy at $z \simeq 1.5$

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Star-forming main-sequence galaxy at z \simeq 2.2

Fit the datacube with ^{3D}Barolo (Di Teodoro+15): $\simeq 254$ km/s $\sigma_{\rm CO} < 15$ km/s $V_{rot}/\sigma_{CO} > 17$ (as local disks)

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Mass Models from HST+ALMA Synergy

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Mass Models from HST+ALMA Synergy

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Disk-halo degeneracy: Long-known issue at z=0

Cold Gas Dynamics in High-z Galaxies

Cold gas tracers at high z with ALMA

[CII] line at 158 µm:

- Main coolant of the ISM: strong & trace cold gas
- Ion. potential of 11.3 eV, similar to HI (13.6 eV)
- Multiphase tracer: atomic, molecular and ionized gas

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- 15 galaxies at $z \approx 4-5$ with known [CII] flux > 2 Jy km/s
- Follow-up observations at $0.05" 0.20" \rightarrow 0.5-1.0$ kpc
- Originally identified as SMGs, QSOs, LBGs, even a DLA
 - \rightarrow high mass (M_{*} $\simeq 10^{10}$ -10¹¹ M_{\odot}) & high SFR (~10²-10³ M_{\odot}/yr)
- Data analysis in progress (Lelli+ in prep.)

Cold Gas Dynamics in High-z Galaxies

TRICEPS: Ubiquitous regular rotation at $z \simeq 4-5$

See also: De Breuck+14; Jones+17, 21; Tadaki+19, 20; LeFevre+2019; Neeleman+20; Rizzo+20, 21; Lelli+21; Fraternali+21

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Rotation Curve: $V_{obs}^2 = -R\nabla\Phi_{tot}$

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Gas Distribution: $V_{gas}^2 = -R\nabla \Phi_{gas}$ Rotation Curve: $V_{obs}^2 = -R\nabla \Phi_{tot}$

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Mass Models with Baryons Only

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Mass Models with NFW Halo

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Cold Gas Dynamics in High-z Galaxies

Comparison with Massive Spirals at z=0

Rotation Curve Shapes

Radial Acceleration Relation

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Future Prospects:

Ultra-deep [CII] observations: → More extended rotation curves → Proposal currently under review

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Rest-frame UV-to-NIR imaging: → Actual stellar mass distribution → Proposal approved!!!

Conclusions:

1. Dynamically cold gas disk at high $z (V_{rot} / \sigma_v > 10)$

 \rightarrow in contrast to common view of turbulent high-*z* disks

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Conclusions:

1. Dynamically cold gas disk at high $z (V_{rot}/\sigma_V > 10)$

 \rightarrow in contrast to common view of turbulent high-*z* disks

2. Strong disk-halo degeneracy at high z

 \rightarrow DM fractions and halo properties are totally unconstrained

3. Mass models often require the presence of bulges \rightarrow fast bulge formation (early mergers? monolithic collapse?)

Thank you!

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Dynamically-cold gas disks may be common!

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ALESS 73.1 at $z \simeq 4.75$: weak noncircular motions

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Rotation curves from different emission lines

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Cold Gas Dynamics in High-z Galaxies

Beam smearing: Lessons from galaxies at z=0

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Beam Smearing: Lessons from HI radio data

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Ionized Gas Dynamics: Confusing Situation!

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ASTRONOMY Speedy

galaxy evolution Mature features are detected in an early galaxy Science "perspective" article by Julie Wardlow commenting on Lelli+(2021, Science).

→ Speedy galaxy evolution now confirmed by many JWST results

Galaxy evolution

Evidence from an early galaxy, ALESS 073.1, indicates the presence of a rotating stellar disk and a central bulge, two features that were thought to be hallmarks of mature, older galaxies. This suggests that the processes that generate these features are present much earlier than previously thought.

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