

Fifth Workshop on Millimetre Astronomy in Italy

Bologna, June 12-14 2023

UMLAUT:

**A Data-Driven Machine Learning
Approach with Automated
Parameter Tuning and Outlier
Detection**

Ivano Baronchelli

INAF - IRA

ARC – Italian node



Identification of Single Spectral Lines through Supervised Machine Learning in a Large HST Survey (WISP): A Pilot Study for Euclid and WFIRST

I. Baronchelli¹, C. M. Scarlata², G. Rodighiero¹, L. Rodríguez-Muñoz¹, M. Bonato^{3,4}, M. Bagley⁵, A. Henry⁶,
M. Rafelski^{6,7}, M. Malkan⁸, J. Colbert⁹, Y. S. Dai (戴昱)¹⁰, H. Dickinson^{2,11}, C. Mancini¹, V. Mehta²,
L. Morselli¹, and H. I. Teplitz⁹

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Abstract

Future surveys focusing on understanding the nature of dark energy (e.g., Euclid and WFIRST) will cover large fractions of the extragalactic sky in near-IR slitless spectroscopy. These surveys will detect a large number of galaxies that will have only one emission line in the covered spectral range. In order to maximize the scientific return of these missions, it is imperative that single emission lines are correctly identified. Using a supervised machine-learning approach, we classified a sample of single emission lines extracted from the WFC3 IR Spectroscopic Parallel survey, one of the closest existing analogs to future slitless surveys. Our automatic software integrates a spectral energy distribution (SED)-fitting strategy with additional independent sources of information. We calibrated it and tested it on a “gold” sample of securely identified objects with multiple lines detected. The algorithm correctly classifies real emission lines with an accuracy of 82.6%, whereas the accuracy of the SED-fitting technique alone is low ($\sim 50\%$) due to the limited amount of photometric data available (≤ 6 bands). While not specifically designed for the Euclid and WFIRST surveys, the algorithm represents an important precursor of similar algorithms to be used in these future missions.



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THE ASTROPHYSICAL JOURNAL SUPPLEMENT SERIES, 257:67 (17pp), 2021 December

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<https://doi.org/10.3847/1538-4365/ac250c>



Identification of Single Spectral Lines in Large Spectroscopic Surveys Using UMLAUT: an Unsupervised Machine-learning Algorithm Based on Unbiased Topology

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Abstract

The identification of an emission line is unambiguous when multiple spectral features are clearly visible in the same spectrum. However, in many cases, only one line is detected, making it difficult to correctly determine the redshift. We developed a freely available unsupervised machine-learning algorithm based on unbiased topology (UMLAUT) that can be used in a very wide variety of contexts, including the identification of single emission lines. To this purpose, the algorithm combines different sources of information, such as the apparent magnitude, size and color of the emitting source, and the equivalent width and wavelength of the detected line. In each specific case, the algorithm automatically identifies the most relevant ones (i.e., those able to minimize the dispersion associated with the output parameter). The outputs can be easily integrated into different algorithms, allowing us to combine supervised and unsupervised techniques and increasing the overall accuracy. We tested our software on WISP (WFC3 IR Spectroscopic Parallel) survey data. WISP represents one of the closest existing analogs to the near-IR spectroscopic surveys that are going to be performed by the future Euclid and Roman missions. These missions will investigate the large-scale structure of the universe by surveying a large portion of the extragalactic sky in near-IR slitless spectroscopy, detecting a relevant fraction of single emission lines. In our tests, UMLAUT correctly identifies real lines in 83.2% of the cases. The accuracy is slightly higher (84.4%) when combining our unsupervised approach with a supervised approach we previously developed.



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⁸ Université de Genève

⁹ Space Telescope

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Received 2020 December 30;

The identification of an emission line in the same spectrum. However, in many cases, the redshift is unknown. We developed a freely available (UMLAUT) that can be used in a variety of cases. To this purpose, the algorithm size and color of the emitting source. In the case, the algorithm automatically identifies lines associated with the output parameters. We combine supervised and unsupervised machine learning. WISP (WFC3 IR Spectroscopic Pilot Survey) near-IR spectroscopic surveys that will investigate the large-scale structure in near-IR slitless spectroscopy. The algorithm correctly identifies real lines in 83% of the cases. This unsupervised approach with a supervised



ORIGINAL RESEARCH
published: 11 July 2019
doi: 10.3389/fpsyg.2019.01528



Inter- vs. Intra-Speaker Variation in Mixed Heritage Syntax: A Statistical Analysis

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¹ Dipartimento di Studi Europei, Americani e Interculturali, Sapienza Università di Roma, Rome, Italy, ² Dipartimento di Fisica e Astronomia, Università di Padova, Padua, Italy, ³ Istituto Culturale Möcheno, Palù del Fersina, Italy

Based on the novel data pertaining to five syntactic phenomena (the position of the finite verb in embedded clauses, in sentences with a modal verb, negative concord, the position of focused light/heavy objects in main clauses with a complex tense and scrambling) in the heritage language Möcheno collected via original fieldwork, we show that there are two populations – one exhibiting intra-speaker variation between German and Italian word orders, and one lacking it; and these two populations are the result of diatopic variation and, to a lesser extent, of diastatic variation. The results achieved using quantitative statistical analysis are partially convergent with those arrived at via the traditional theoretical syntax for Möcheno, but our analysis has allowed us to shed new light on a series of phenomena that have been neglected or poorly understood thus far. More specifically, and for the first time, we discovered that there is a micro-variation resulting from diastatic (age) variation within the Roveda variety, which represents the

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Software

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UMLAUT - Little snail

Ivano Baronchelli

UMLAUT 1.0

Unsupervised Machine Learning Algorithm based on Unbiased Topology (IDL VERSION)

DOI: 323649071.svg

<https://zenodo.org/badge/323649071.svg>

REFERENCE: Baronchelli I. et al. (2021):

<https://iopscience.iop.org/article/10.3847/1538-4365/ac250c/meta>

<https://arxiv.org/abs/2111.01807>

UMLAUT is a variant of the KNN (K-closest neighbor) algorithm.

- Given a set of reference data points (training set), for which the value of N+1 parameters is known,
- Given one analysis data point with N parameters known and the (N+1)-th parameter unknown,

--> UMLAUT estimates the value of the (N+1)-th parameter for the analysis data point. To this purpose, UMLAUT finds the closest data-points of the training set, in a ranked N-dimensional space "associated" with the physical parameter space. After finding the closest data points, the unknown parameter is obtained as the combination (ex. from the average) of the values assumed by the closest reference data points along the (N+1)-th dimension.

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sky in near-IR slitless spectroscopy
correctly identifies real lines in 83.
unsupervised approach with a supe

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Python

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UMLAUT

Unsupervised Machine Learning Algorithm based on Unbiased Topology

IDL

Public

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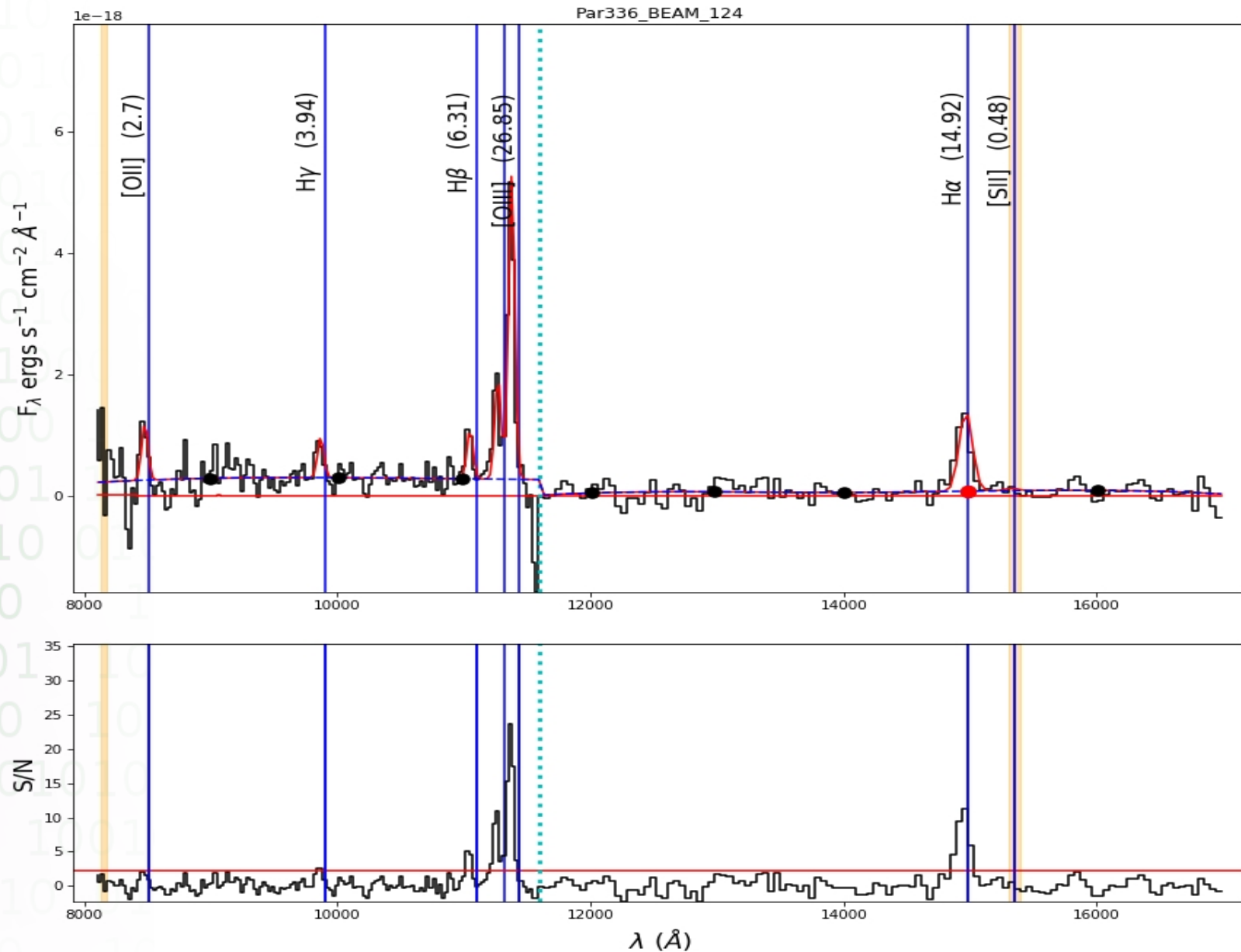
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Identifying emission lines is quite easy... ... when multiple lines are detected

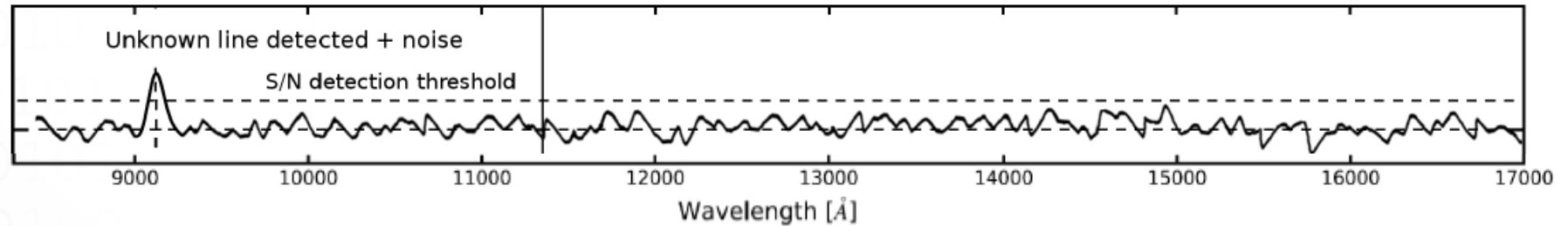


But when only one line is detected...

HST IR grisms (J and H band)

G102

G141



Unknown single line detected

Anything else is beneath the S/N detection threshold

But when only one line is detected...

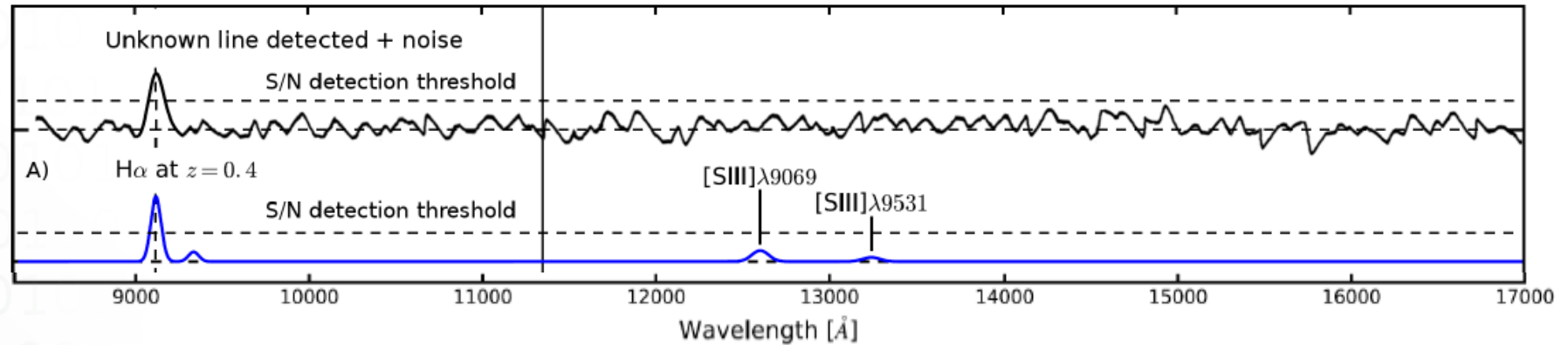
HST IR grisms (J and H band)

G102

G141

?

H α



Unknown single line detected

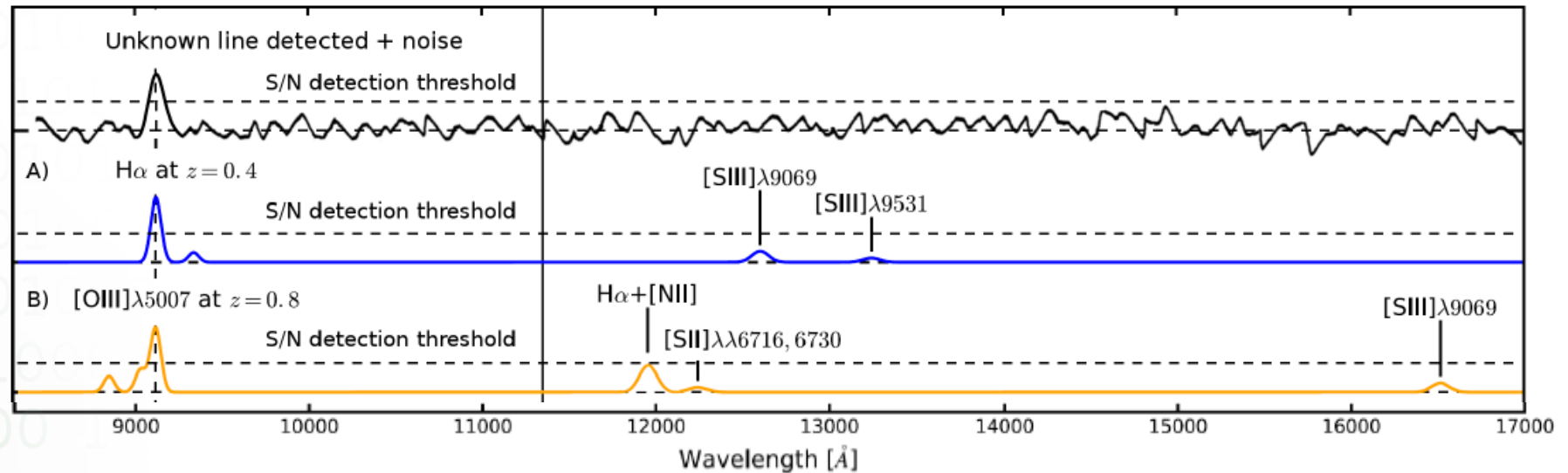
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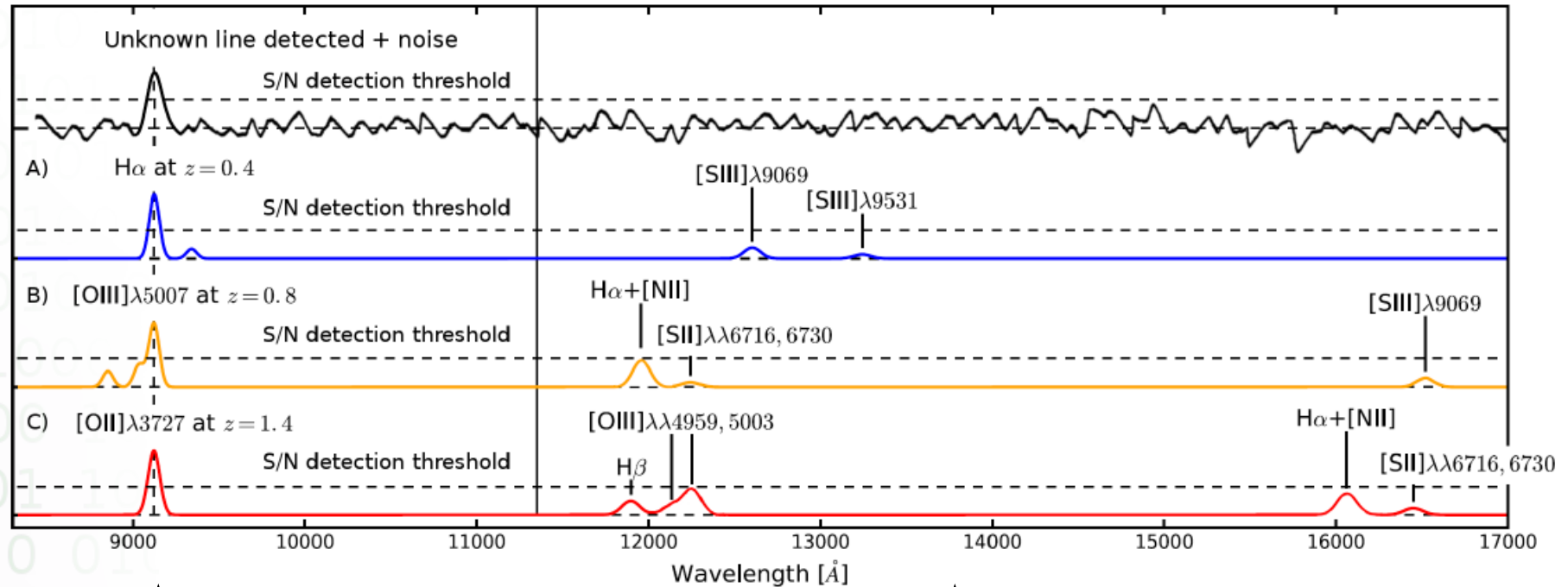
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HST IR grisms (J and H band)

G102

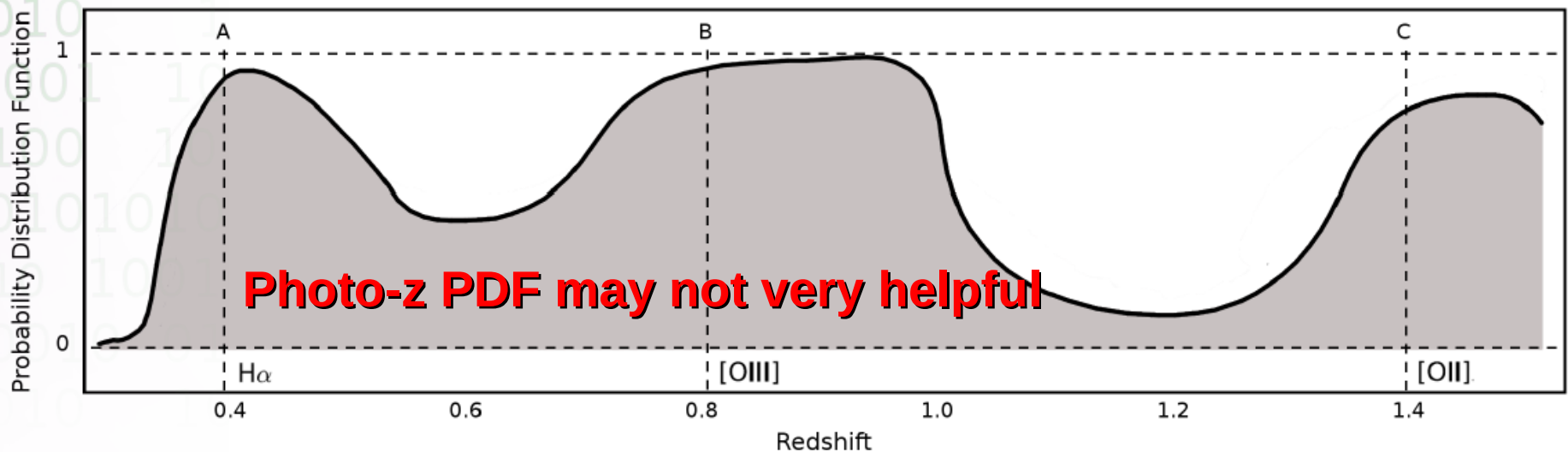
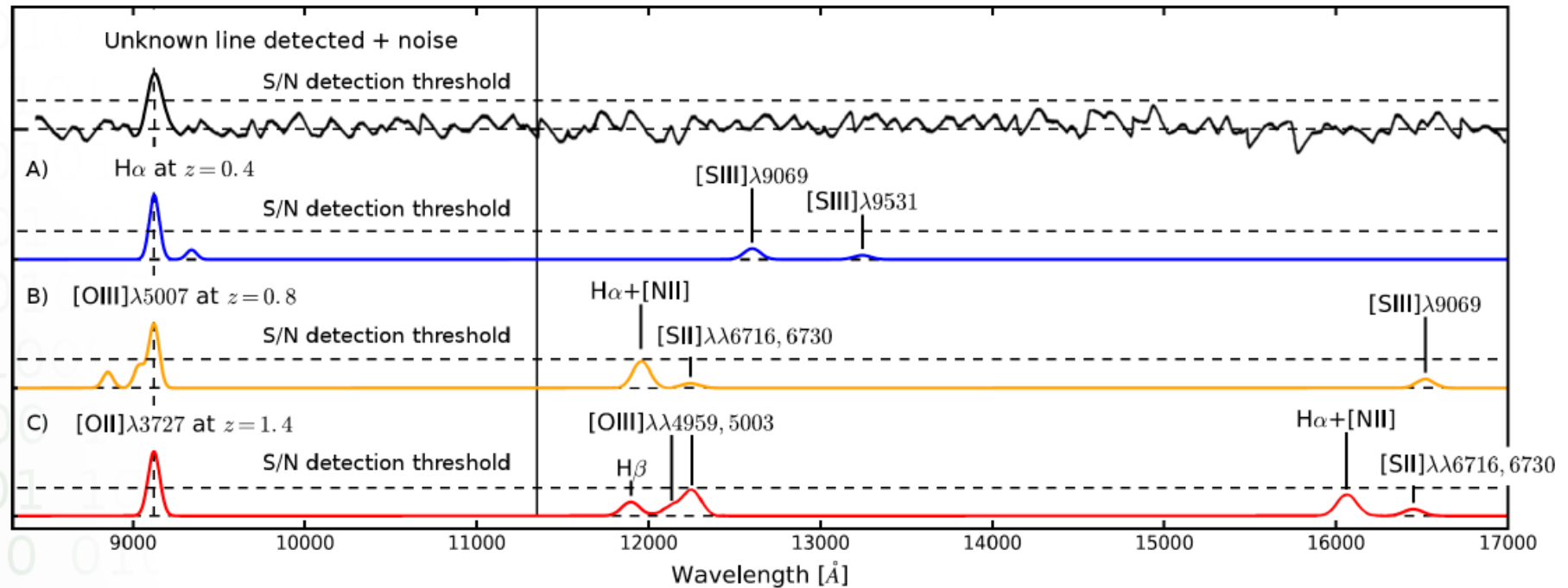
G141

?

$H\alpha$

$[OIII]$

$[OII]$



WISP

PI: M. A. Malkan (University of California - Los Angeles)

The WFC3 Infrared Spectroscopic Parallel survey

H. Atek et al. (2010)

HST pure parallel large program (~ 2000 orbits, >1500 arcmin²)

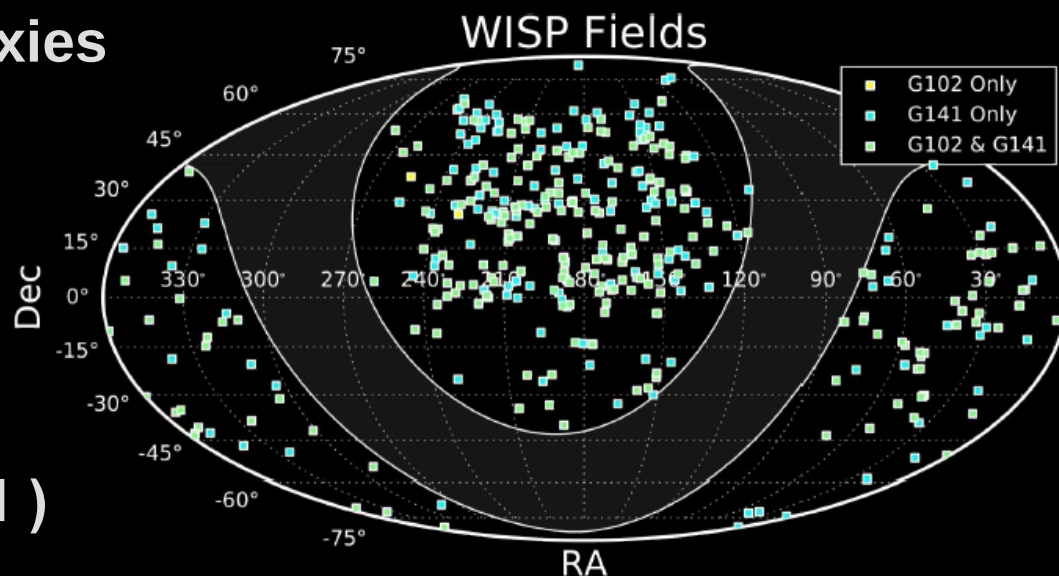
- slitless spectroscopy $\left\{ \begin{array}{l} \text{G102: } 0.8 - 1.1 \mu\text{m (R}\sim 210\text{)} \\ \text{G141: } 1.07 - 1.7 \mu\text{m (R}\sim 130\text{)} \end{array} \right.$
- direct imaging $\left\{ \begin{array}{l} \text{J band (F110W),} \\ \text{H band (F140W/F160W)} \\ \text{Ancillary UVIS + IRAC + ground based optical} \end{array} \right.$

- One of the most important proxies

For the future surveys of

Euclid and Roman

- Now collecting data with JWST (PASSAGE survey, P.I. M. Malkan, proposal ID: GO-1571 - Cycle 1)



An artificial intelligence is not an artificial GOD

**A learner can help exploiting hidden information,
but it does not *create* information from nothing.**

**Additional information that can help
identifying a spectral line:**

- Apparent Magnitude (F110W);

- Apparent Size (F110W);

- Color index (J-H);

- Line equivalent width;

... etc ...



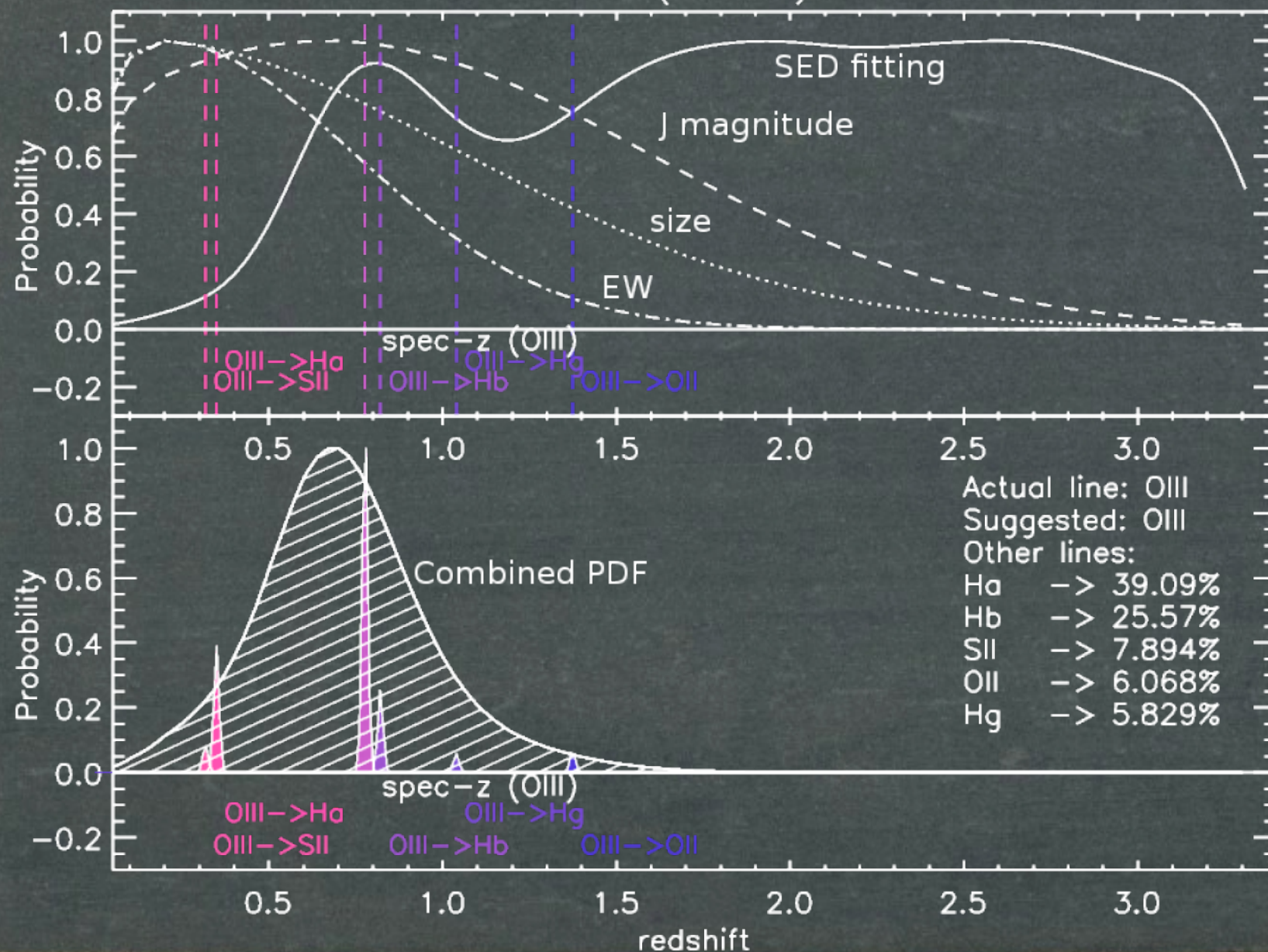
ALGORITHM 1 (Naïve Bayes approach)

PAR: 369

OBJ: 42

MAG (F110W): 21.83

z: 0.777



UMLAUT basics

Unbiased Machine Learning Algorithm Using Topology

Empirical principle:

Objects that are close to each other in N dimensions tend to be close also in an independent $N+1$ dimension, when $N \gg$

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Toucans that are similar to each other in 2 dimensions, tend to be similar also in 3 dimensions...

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Unbiased Machine Learning Algorithm Using Topology

Empirical principle:

Objects that are close to each other in N dimensions tend to be close also in an independent $N+1$ dimension, when $N \gg$



Toucans that are similar to each other in 2 dimensions, tend to be similar also in 3 dimensions...

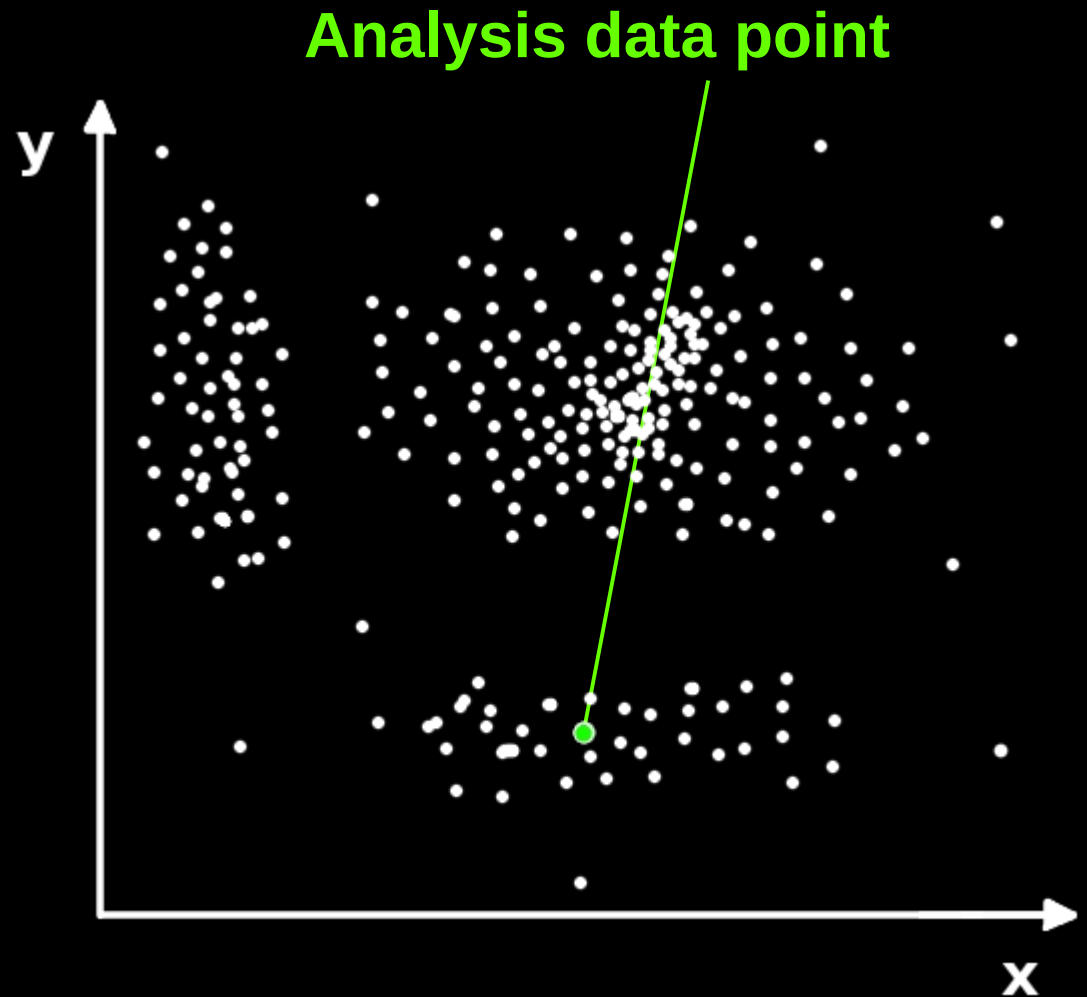
...unless one of the pictures is depicting the picture of a toucan.

How it works in practice:

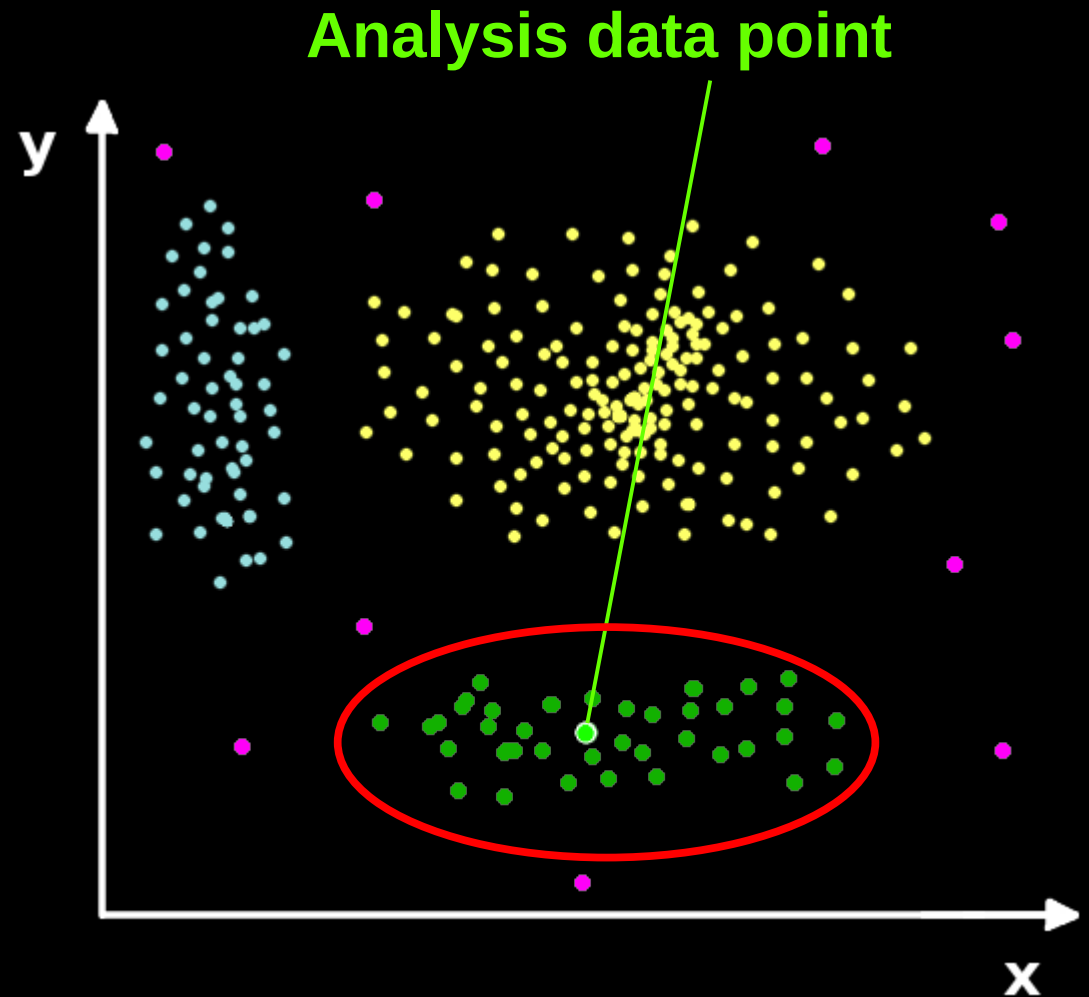
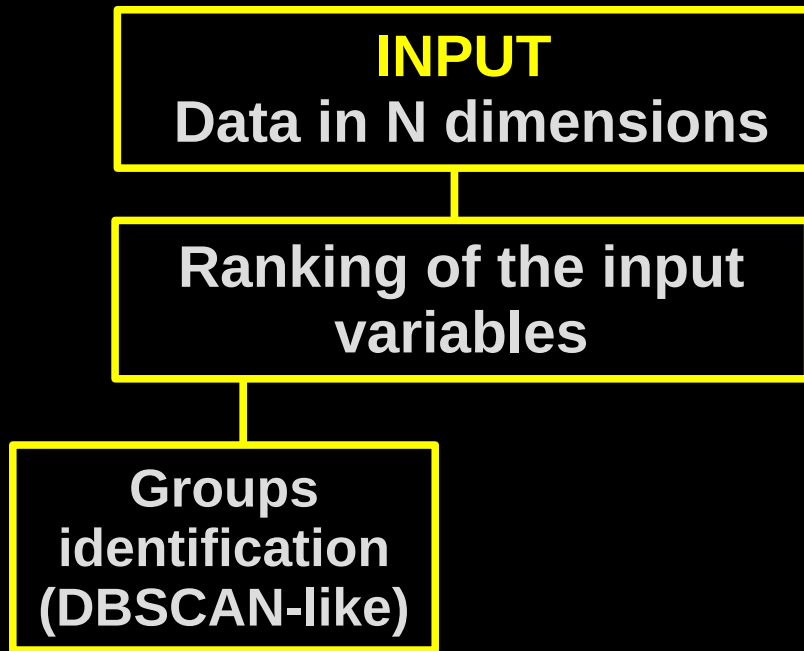
INPUT

Data in N dimensions

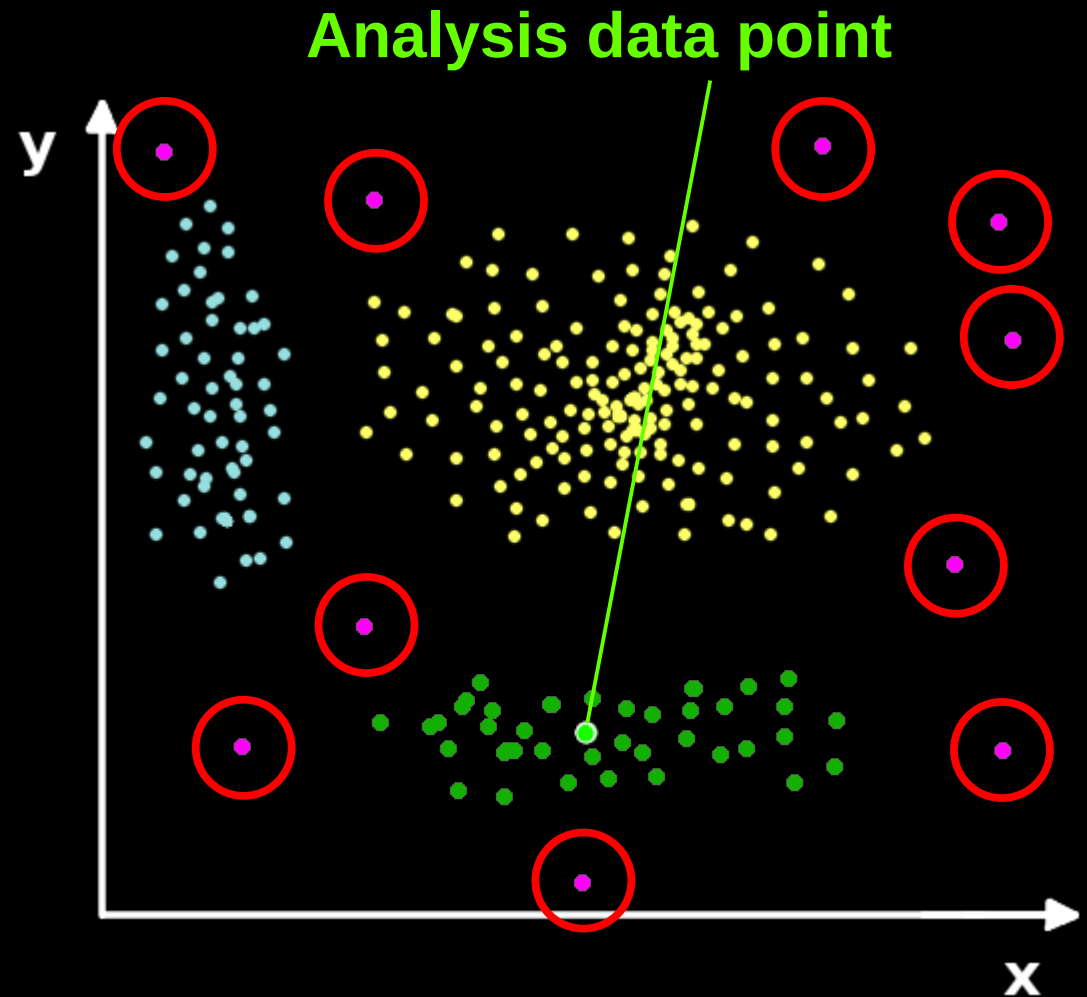
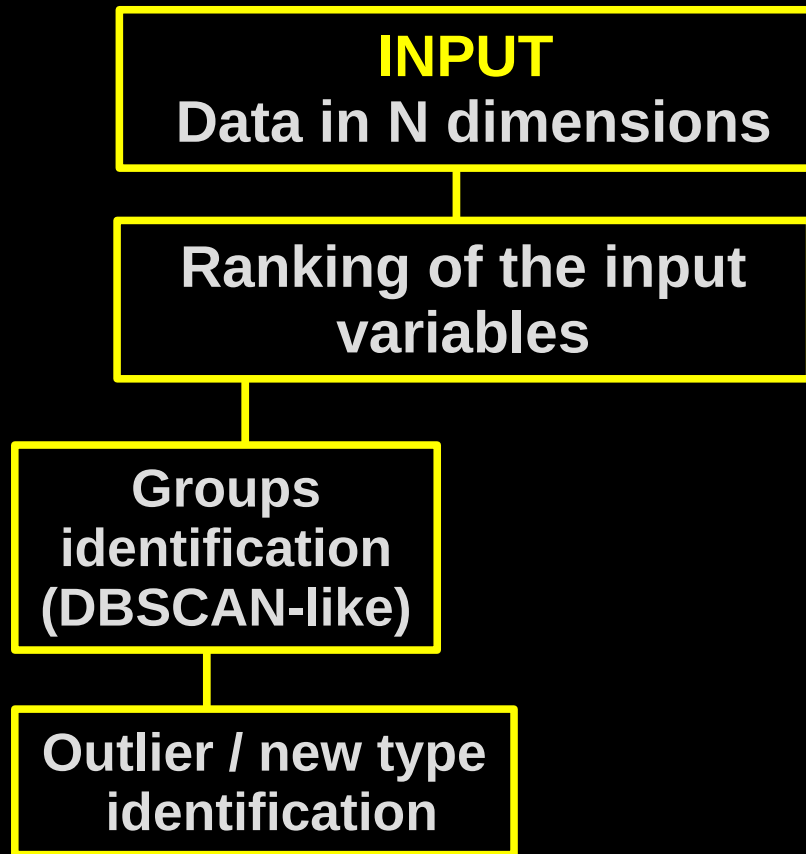
Ranking of the input
variables



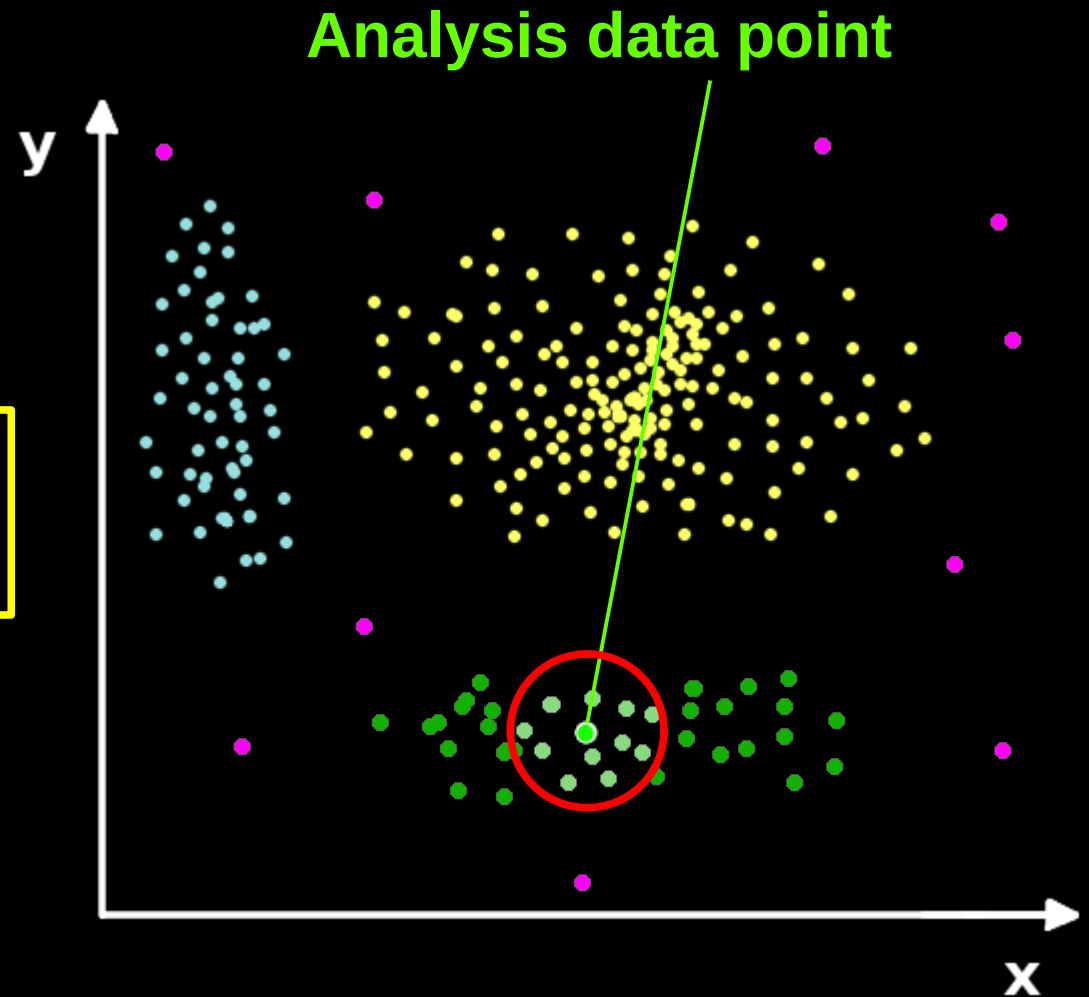
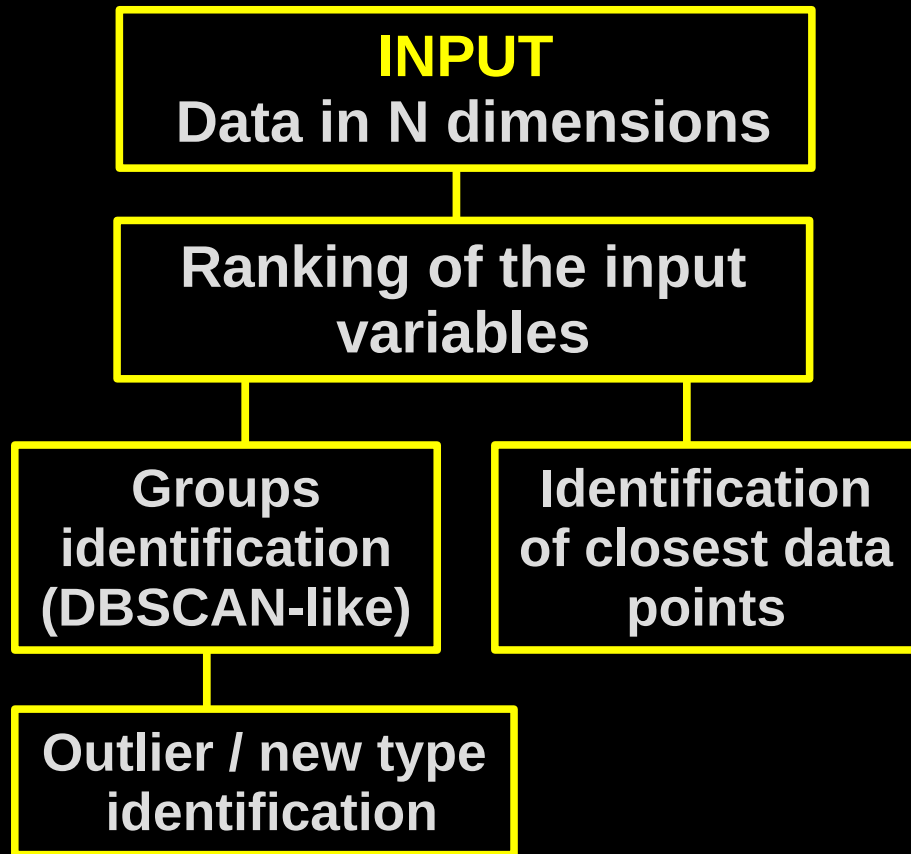
How it works in practice:



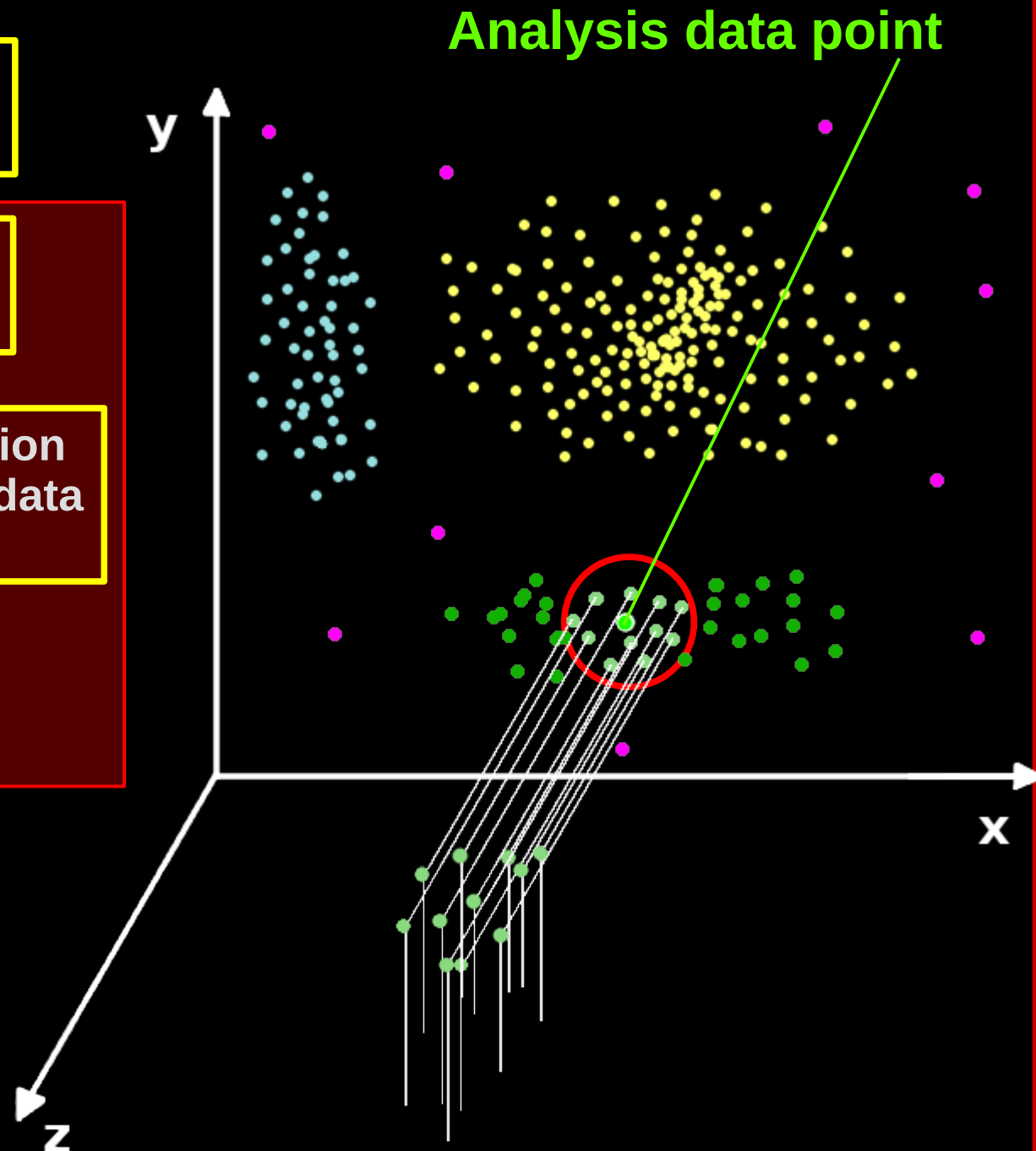
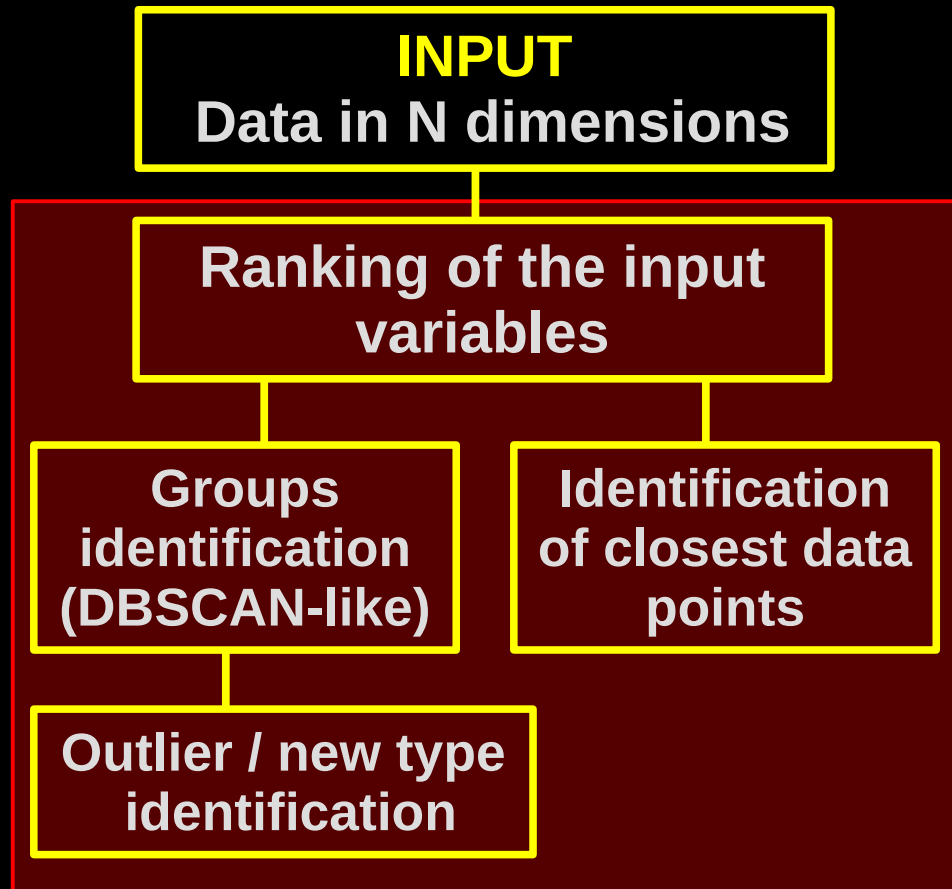
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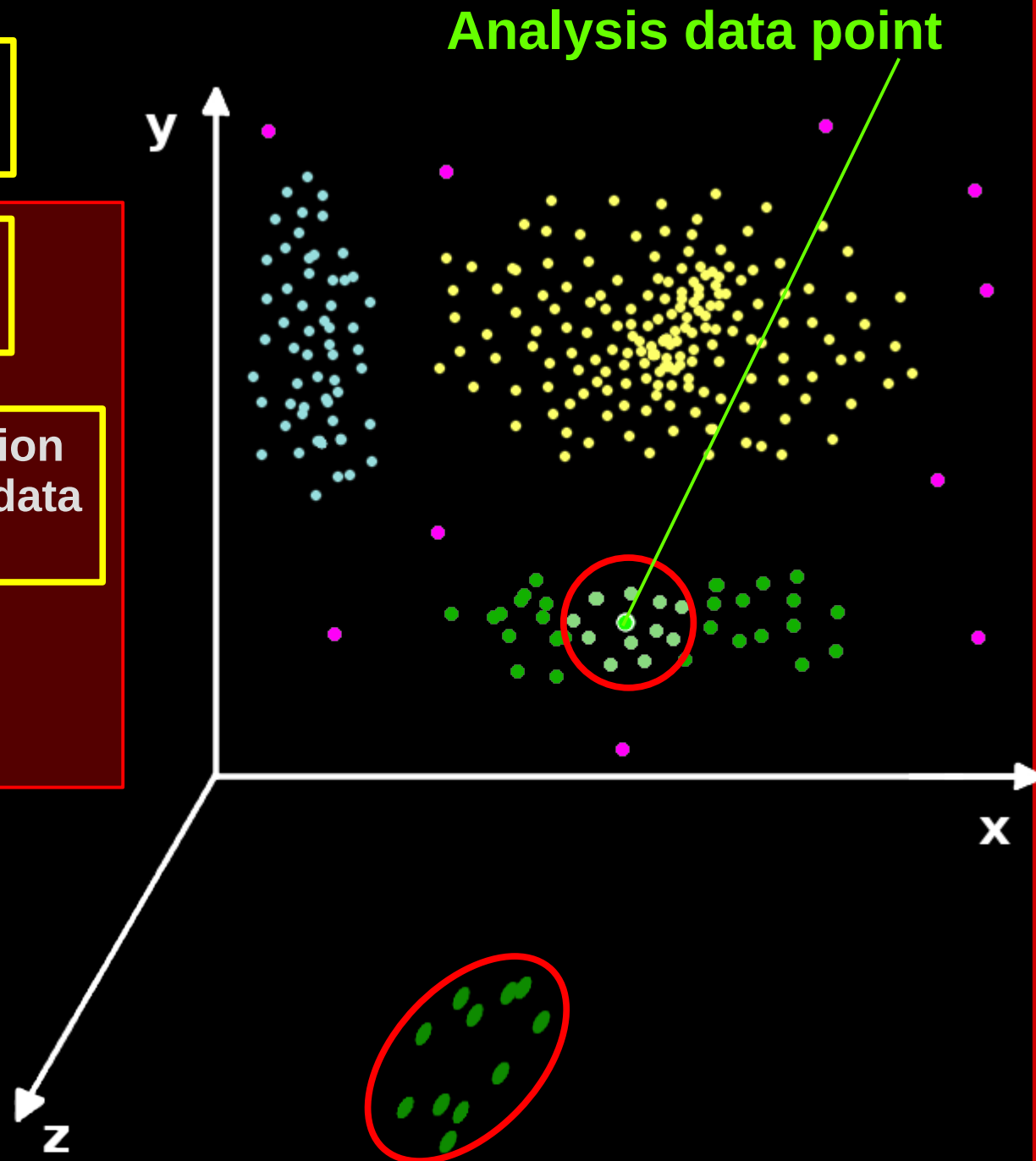
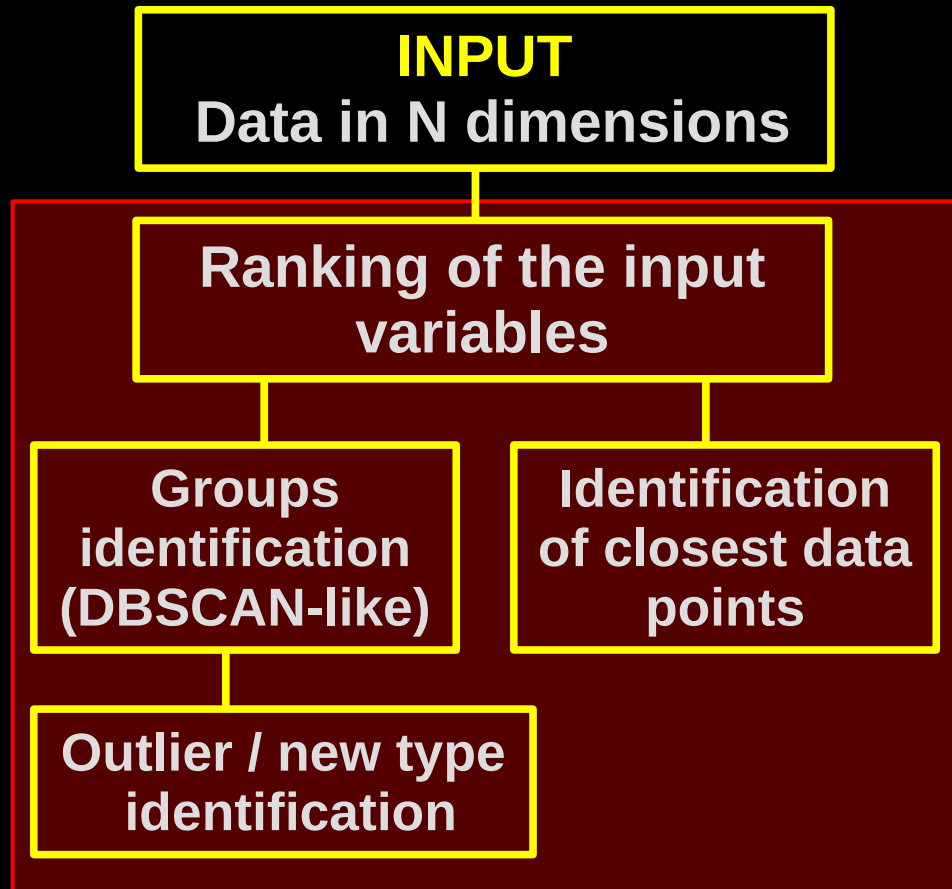
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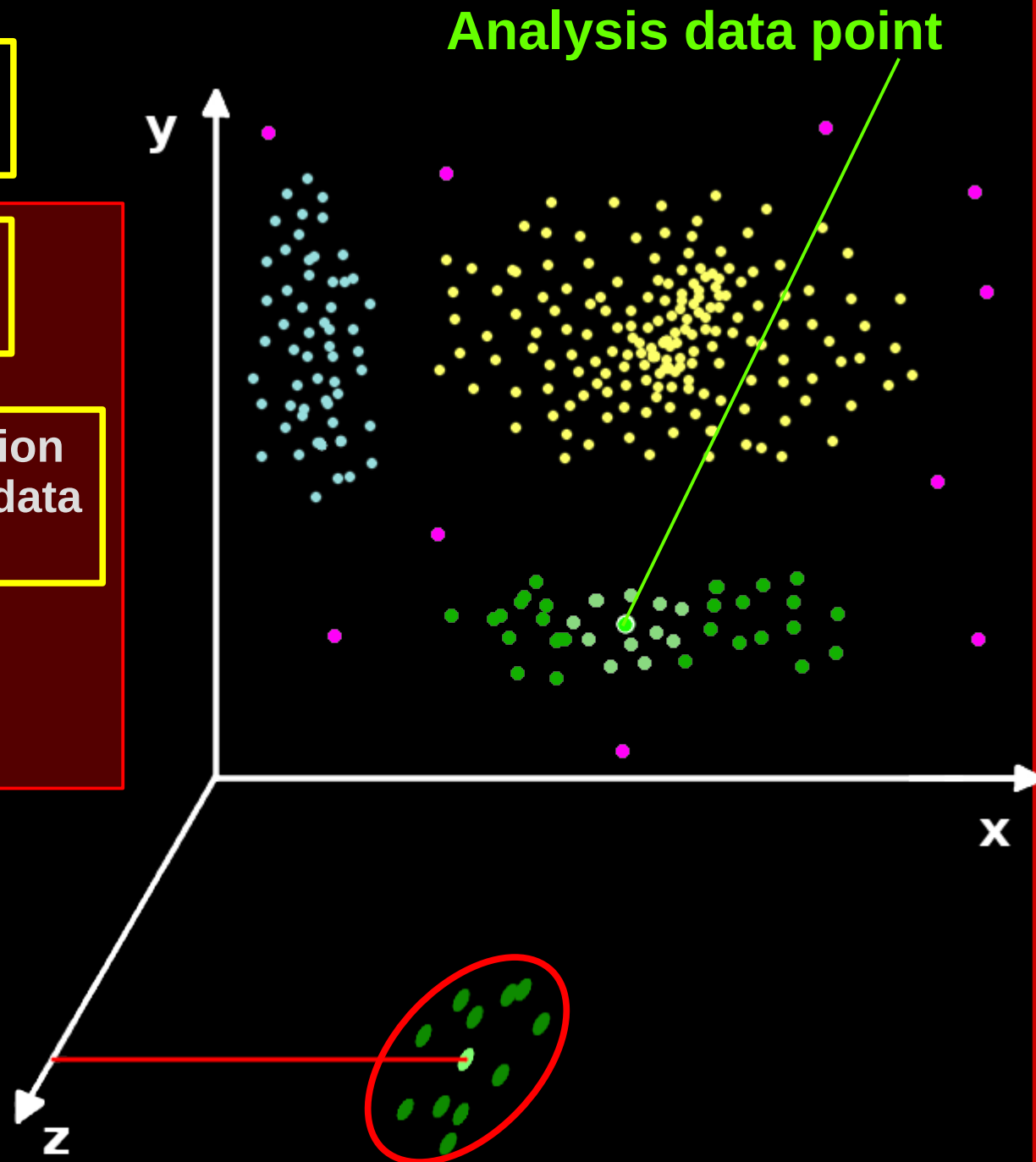
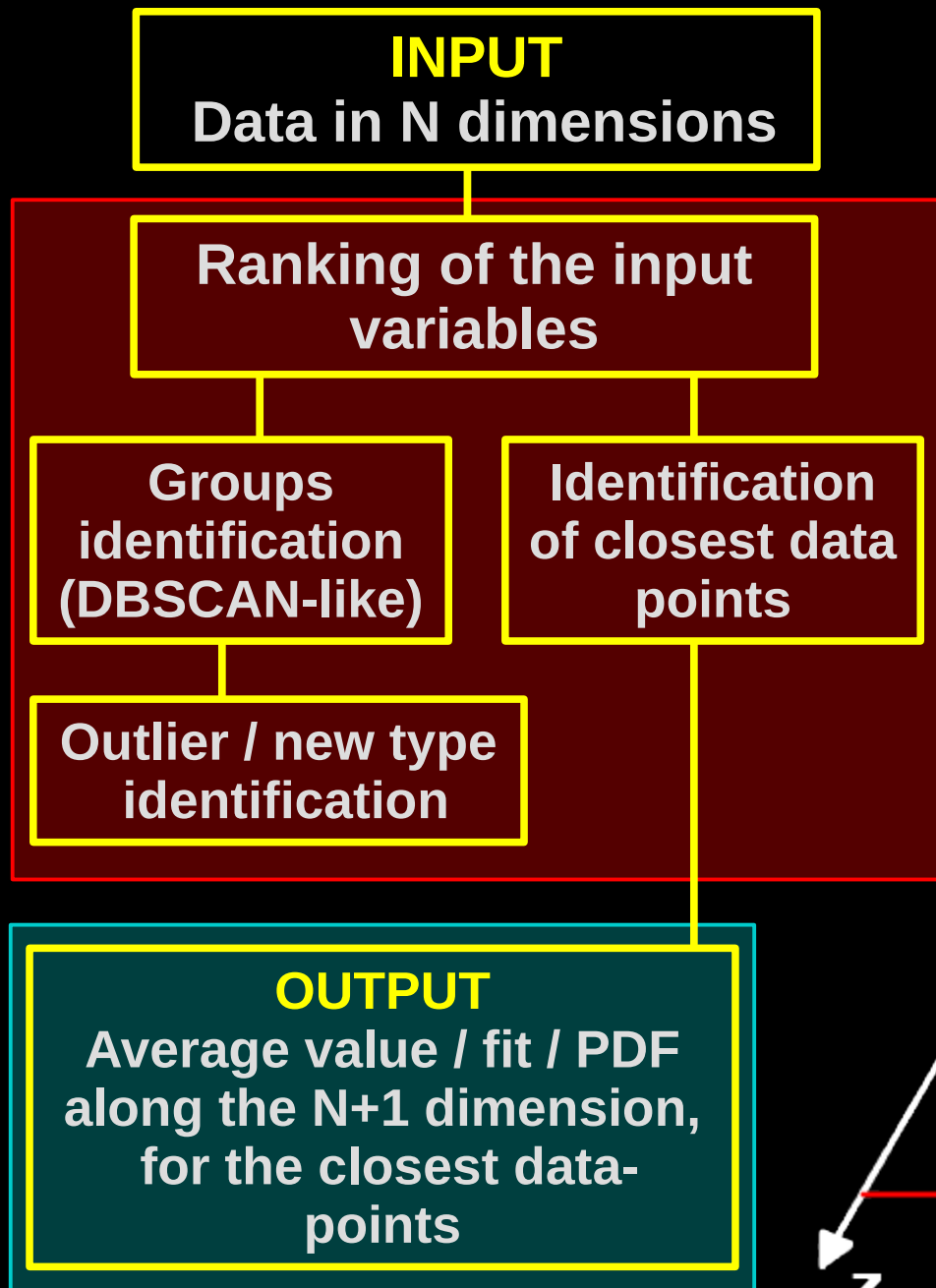
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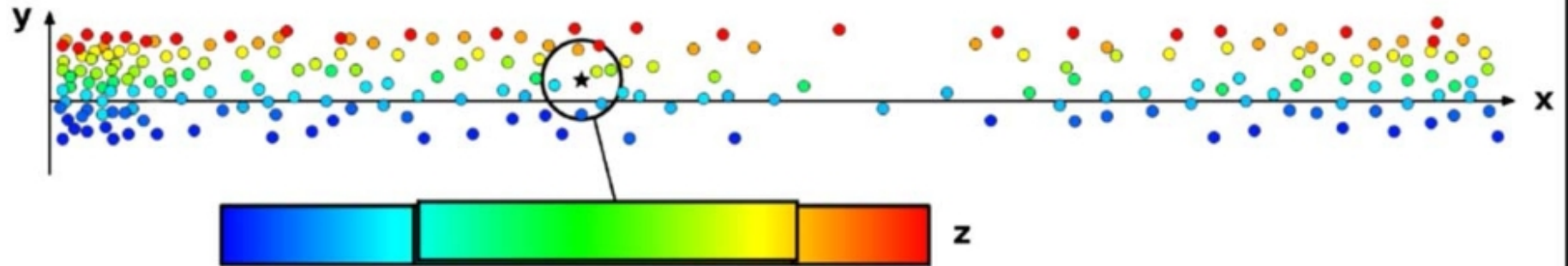
How it works in practice:



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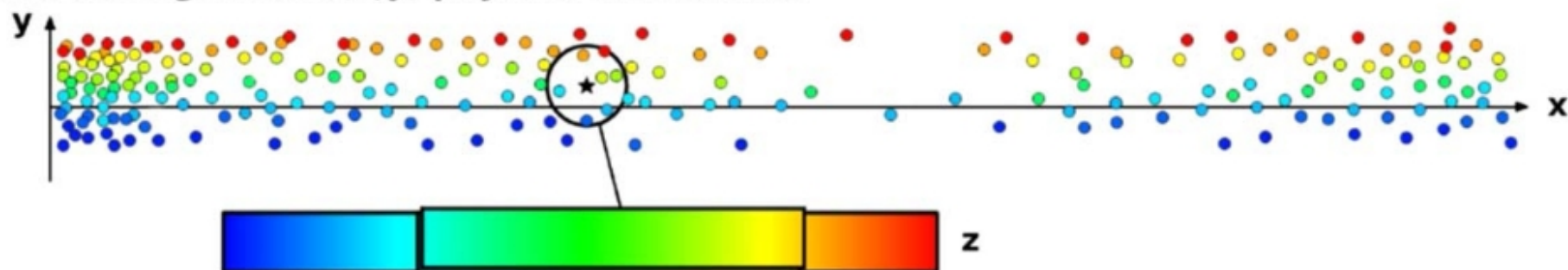


A) Not homogeneous (x,y) physical dimensions



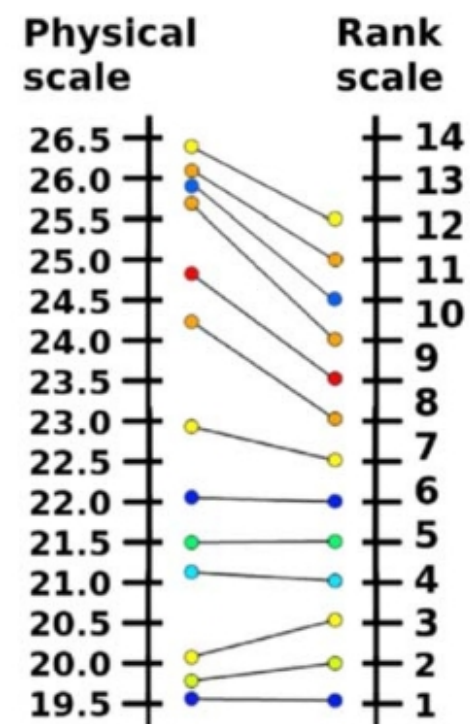
Dispersion along z

A) Not homogeneous (x,y) physical dimensions

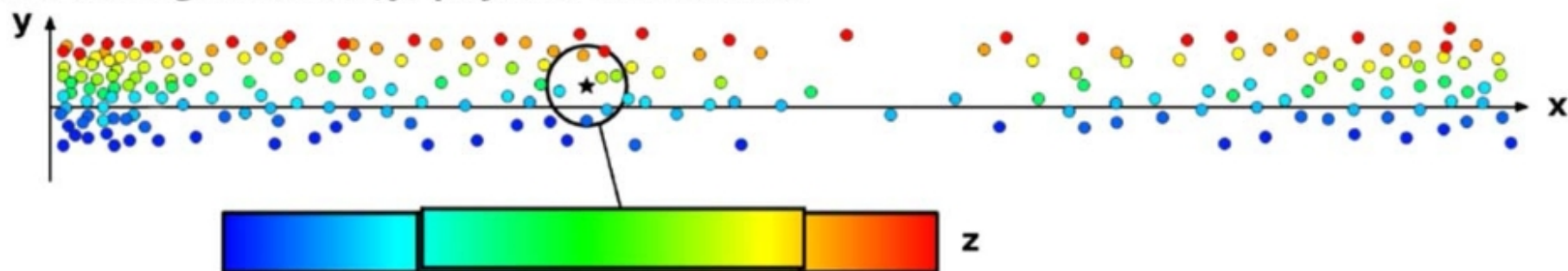


Dispersion along z

C)

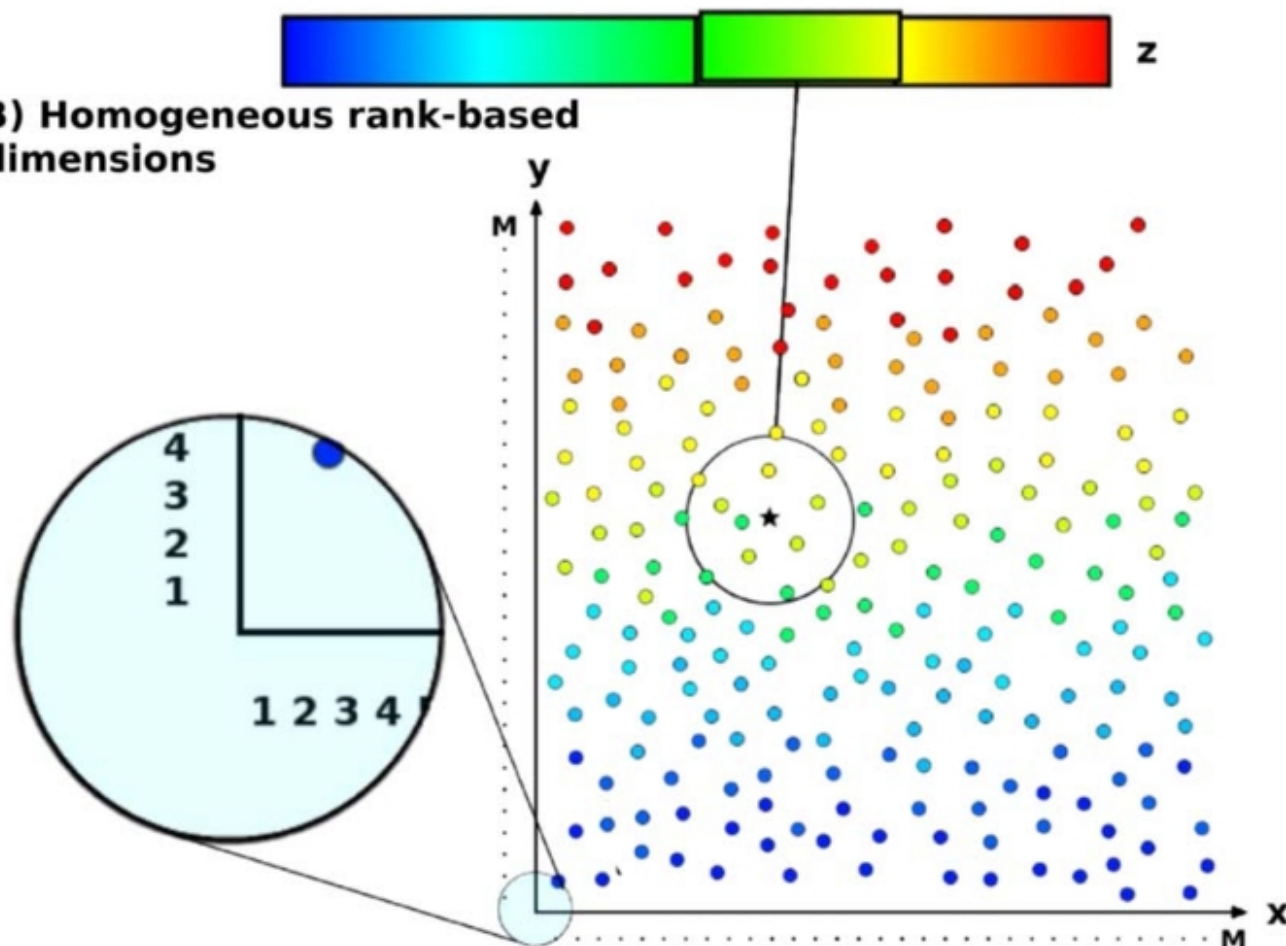


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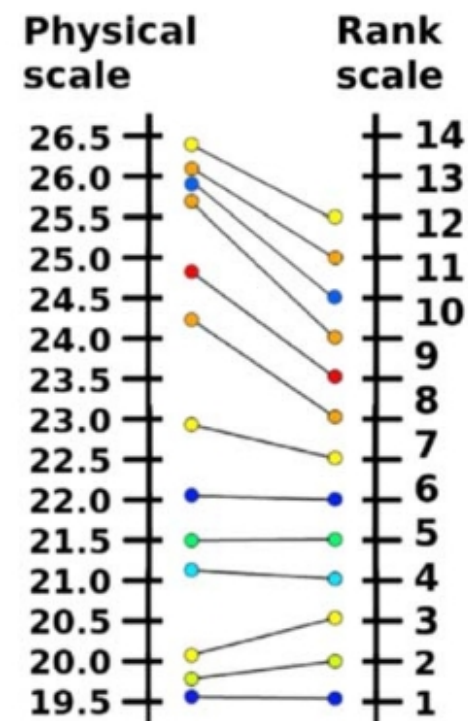


Dispersion along z

B) Homogeneous rank-based dimensions

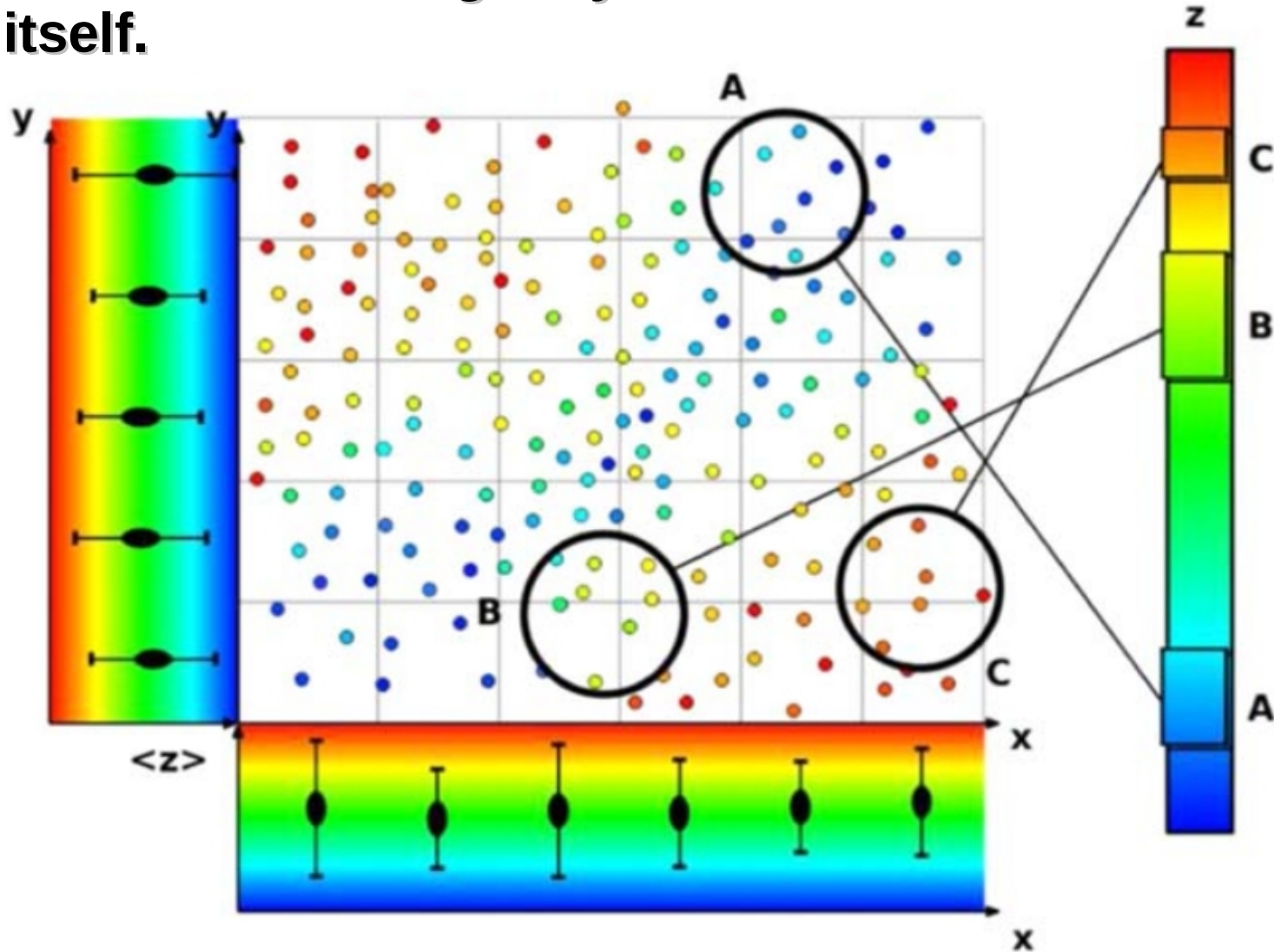


C)



Among the UMLAUT Advantages:

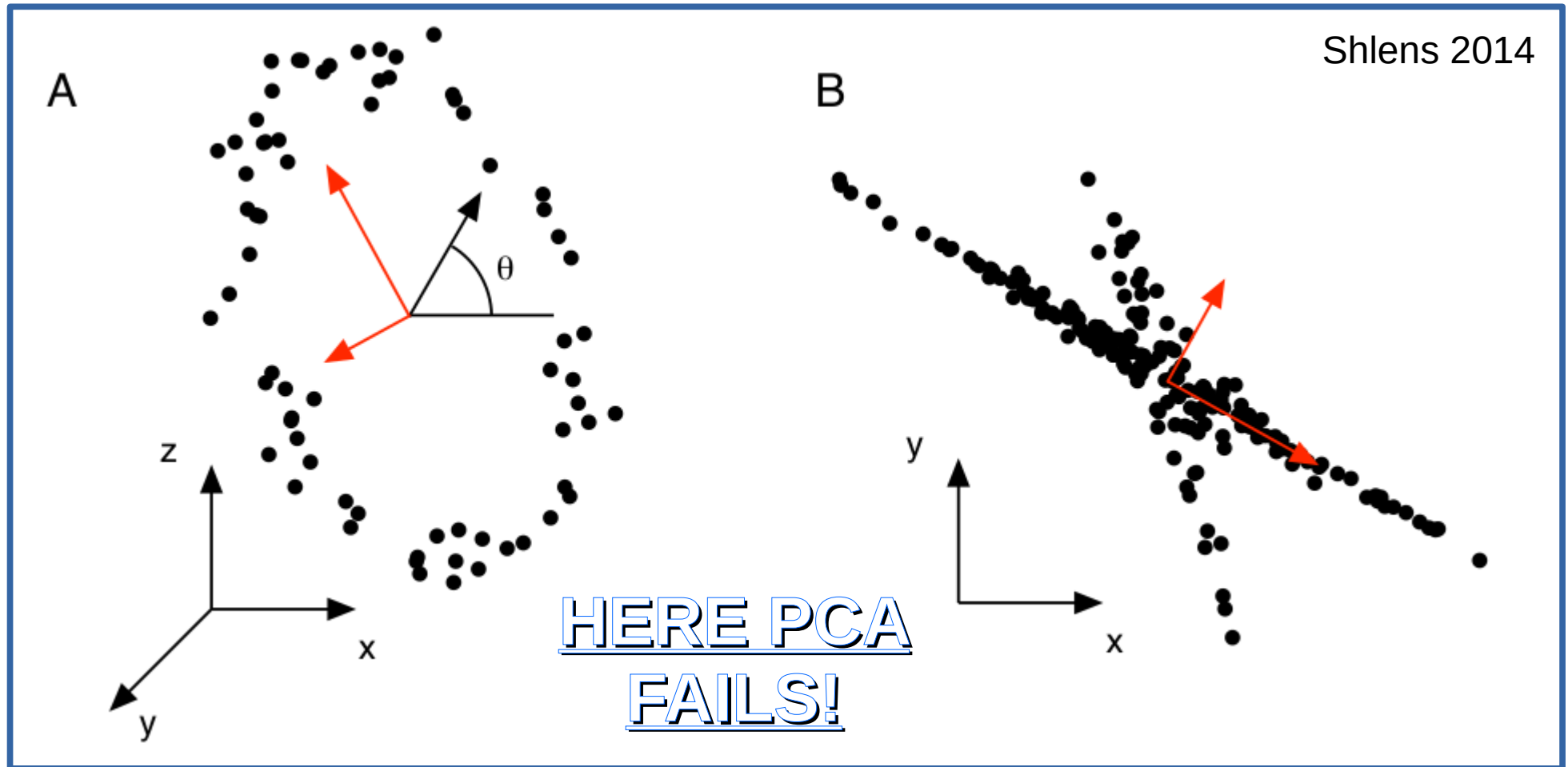
Single dimensions could retain useful information about the output without showing any correlations at all with the output itself.



→ UMLAUT takes into account all the input dimensions at the same time!

Among the UMLAUT Advantages:

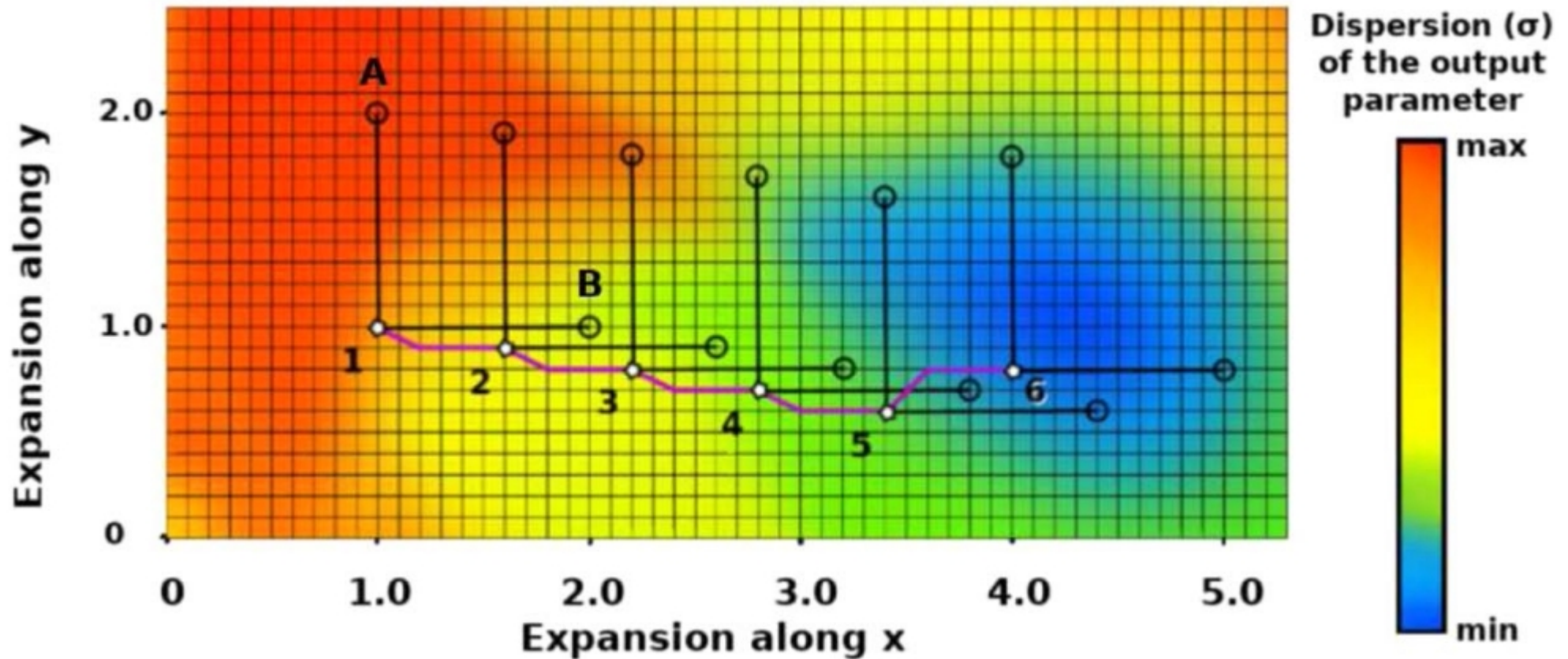
Dimensionality reduction strategies (such as PCA) often assume orthogonal native axes and/or Gaussian distributions



→ **UMLAUT does not require/assume orthogonal native axes or assumptions on the data distribution.**

Dimensionality reduction strategy

N-dimensional space of expansion of the N input parameters
(N=2 in this example)



Strategy (simplified):

- 1) Measure the dispersion on the output parameter in positions 1, A and B
- 2) Move into the direction of decreasing gradient (position 2)
- 3) Iterate until a local minimum is found

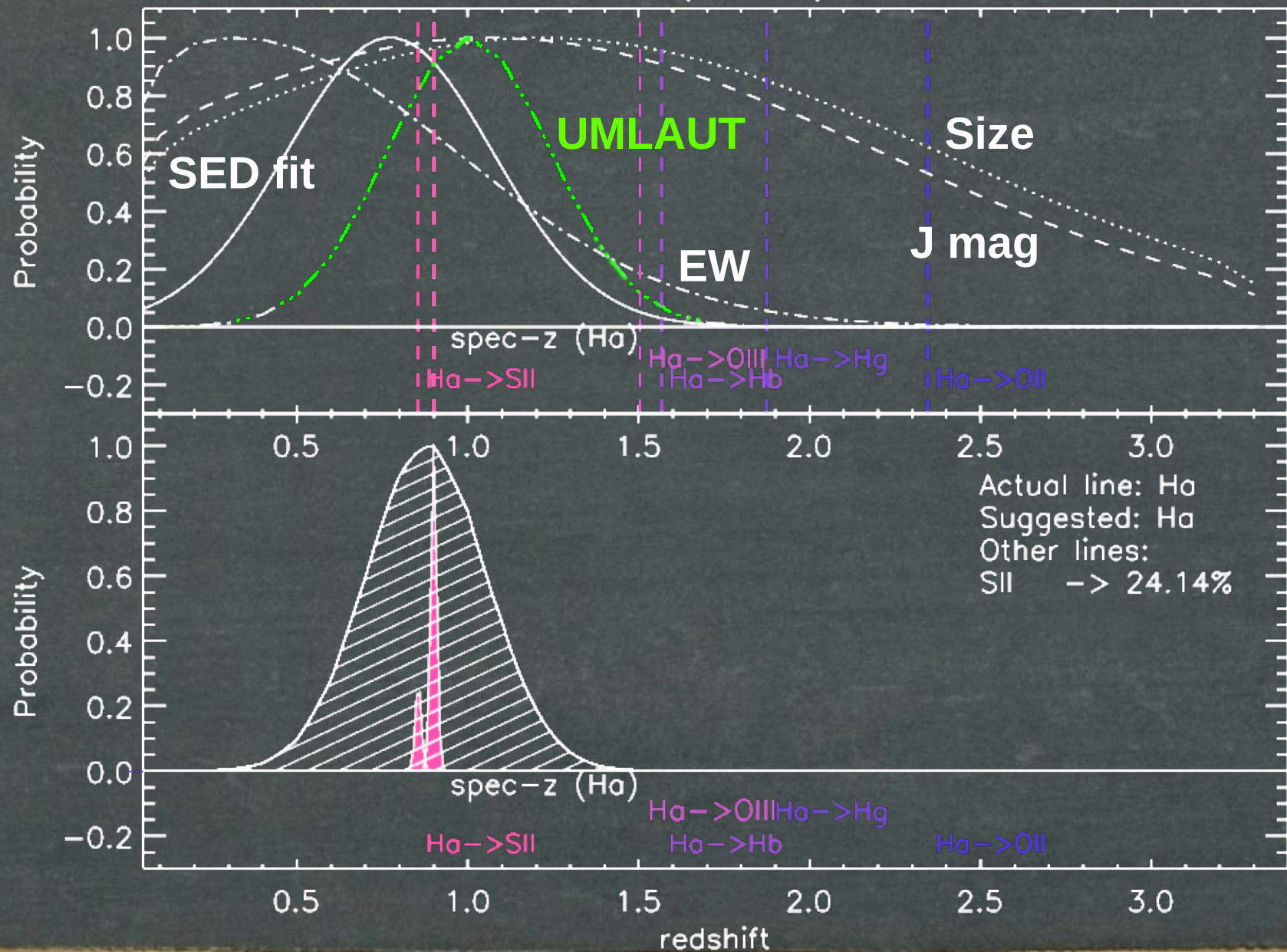
Naïve Bayes + UMLAUT

PAR: 454

OBJ: 54

MAG (F110W): 22.5570

z: 0.90060



Naïve Bayes alone

	Completeness	Contamination	Purity
Ha	88.3 %	15.2 %	84.8 %
[OIII]	77.0 %	17.9 %	82.1 %
[OII]	50.0 %	73.3 %	26.7 %

UMLAUT alone

	Completeness	Contamination	Purity
Ha	89.6 %	16.0 %	84.0 %
[OIII]	77.8 %	18.0 %	82.0 %
[OII]	12.5 %	33.3 %	66.7 %

Naïve Bayes + UMLAUT

	Completeness	Contamination	Purity
Ha	89.3 %	12.8 %	87.2 %
[OIII]	79.6 %	16.2 %	83.8 %
[OII]	66.7 %	70.9 %	29.1 %

ALMA Band 3 number counts

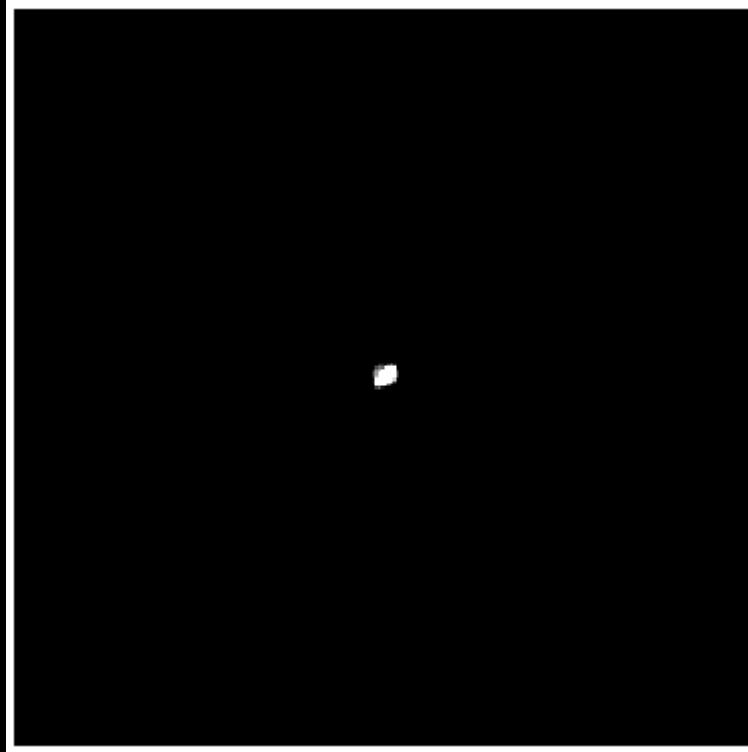


Image of a band 3 calibrator

- Are there any other serendipitous detections around?



Compute number counts !

ALMA Band 3 number counts

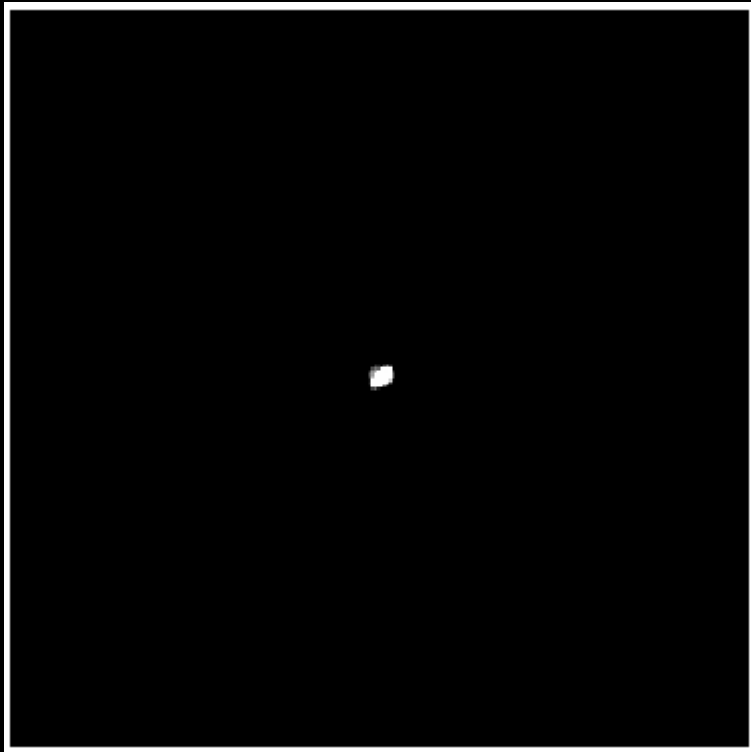
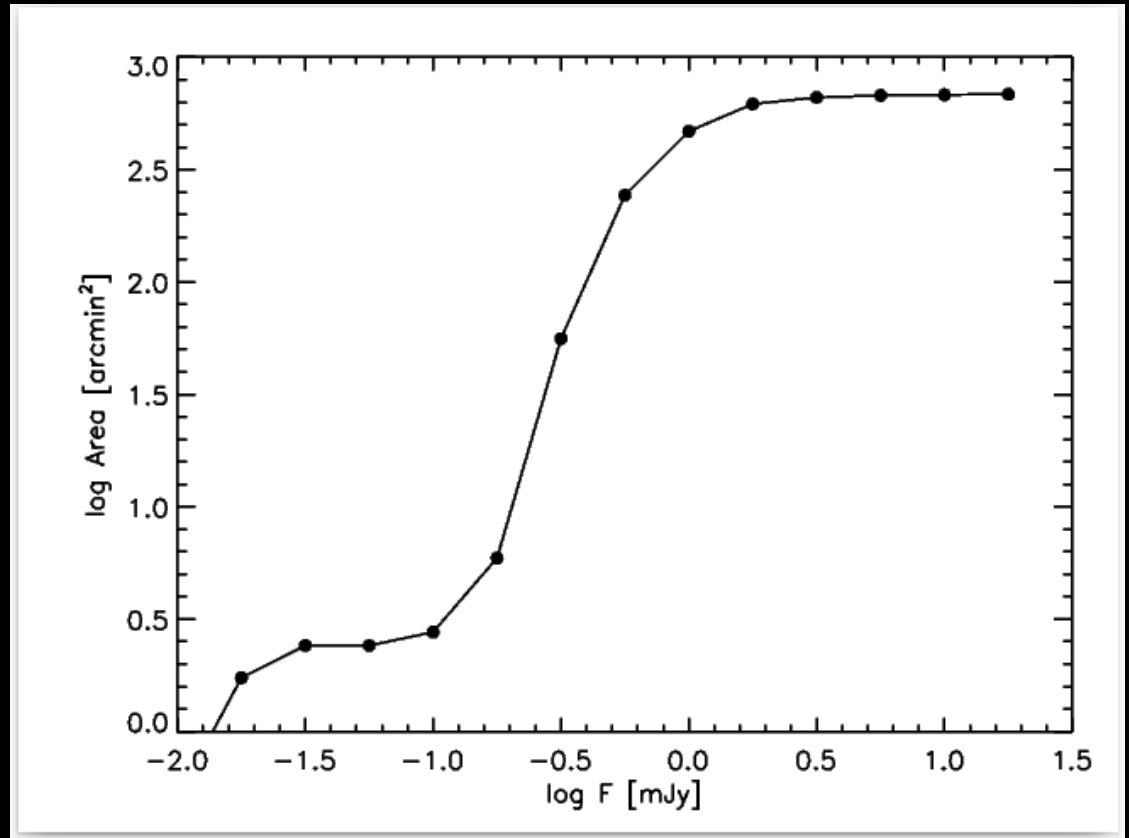


Image of a band 3 calibrator



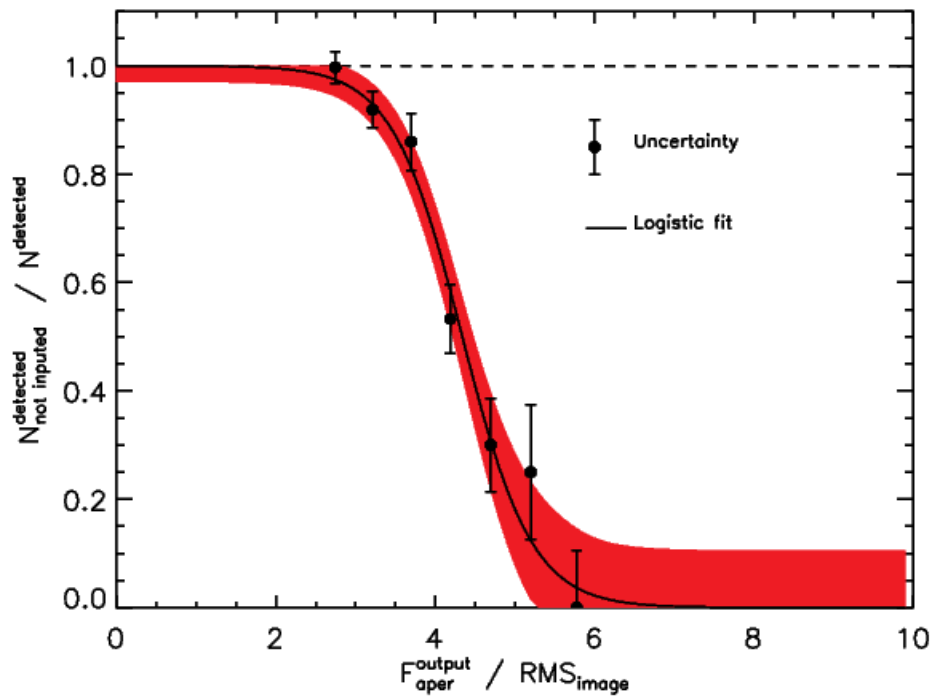
~ 500 valid calibrator images
~ 700 arcmin² (at 1mJy)

- Are there any other serendipitous detections around?

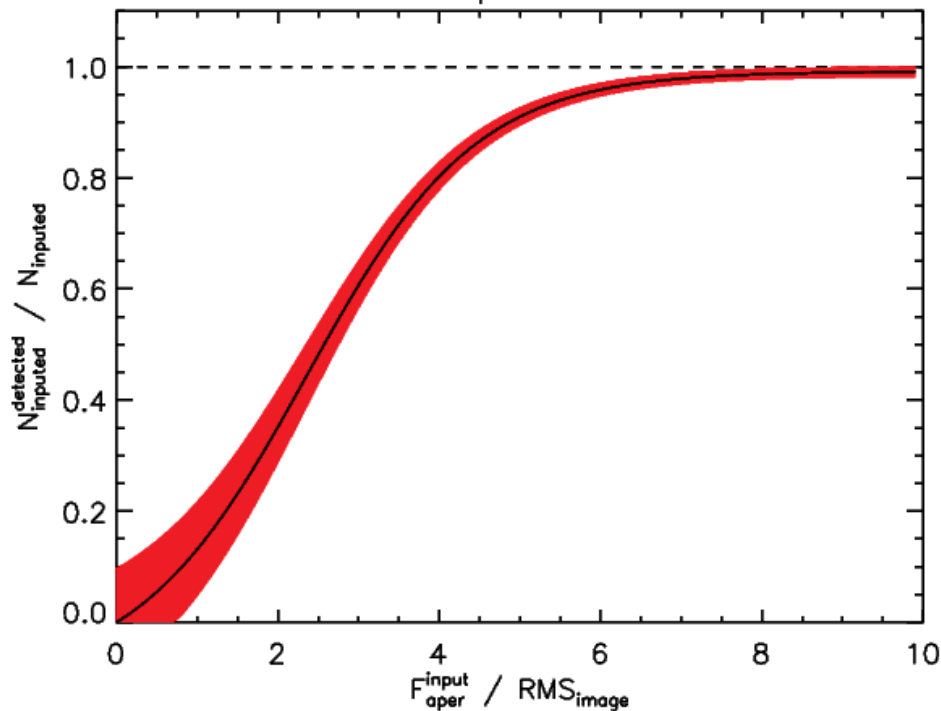


Compute number counts !

Contamination



Completeness



$$N_i(S_i) = \frac{1 - f_{\text{sp}}(S_i)}{C(S_i) \cdot A(S_i)},$$

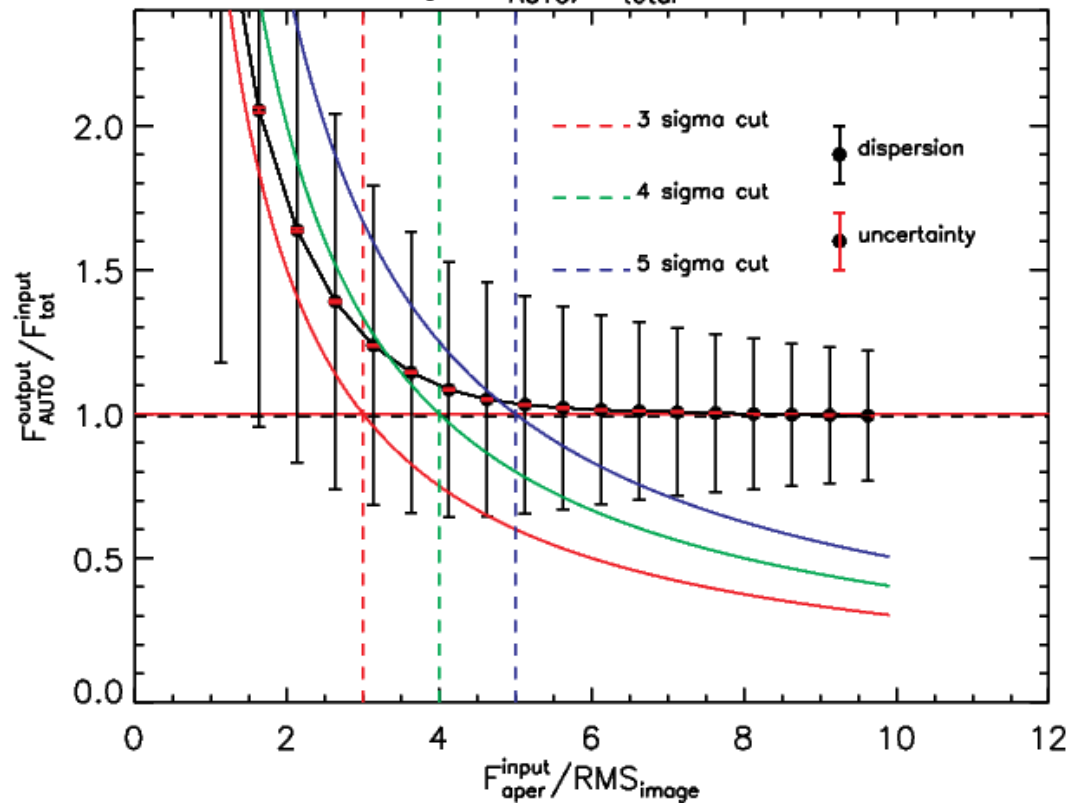
Contribution of each single detection

S_i is the flux density of the detection;

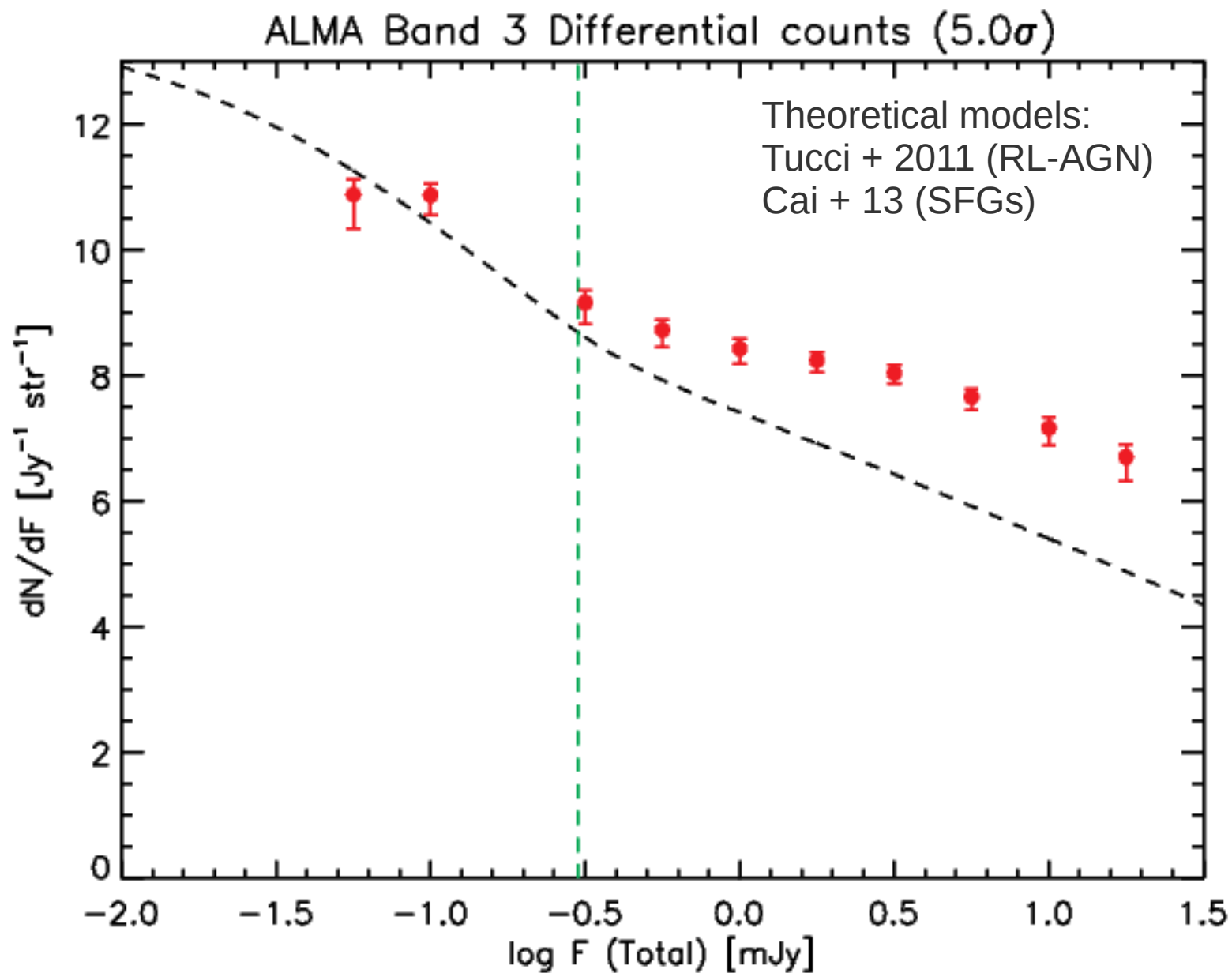
$f_{\text{sp}}(S_i)$ is the fraction of spurious sources at S_i ,

$A(S_i)$ is the effective area at S_i ,

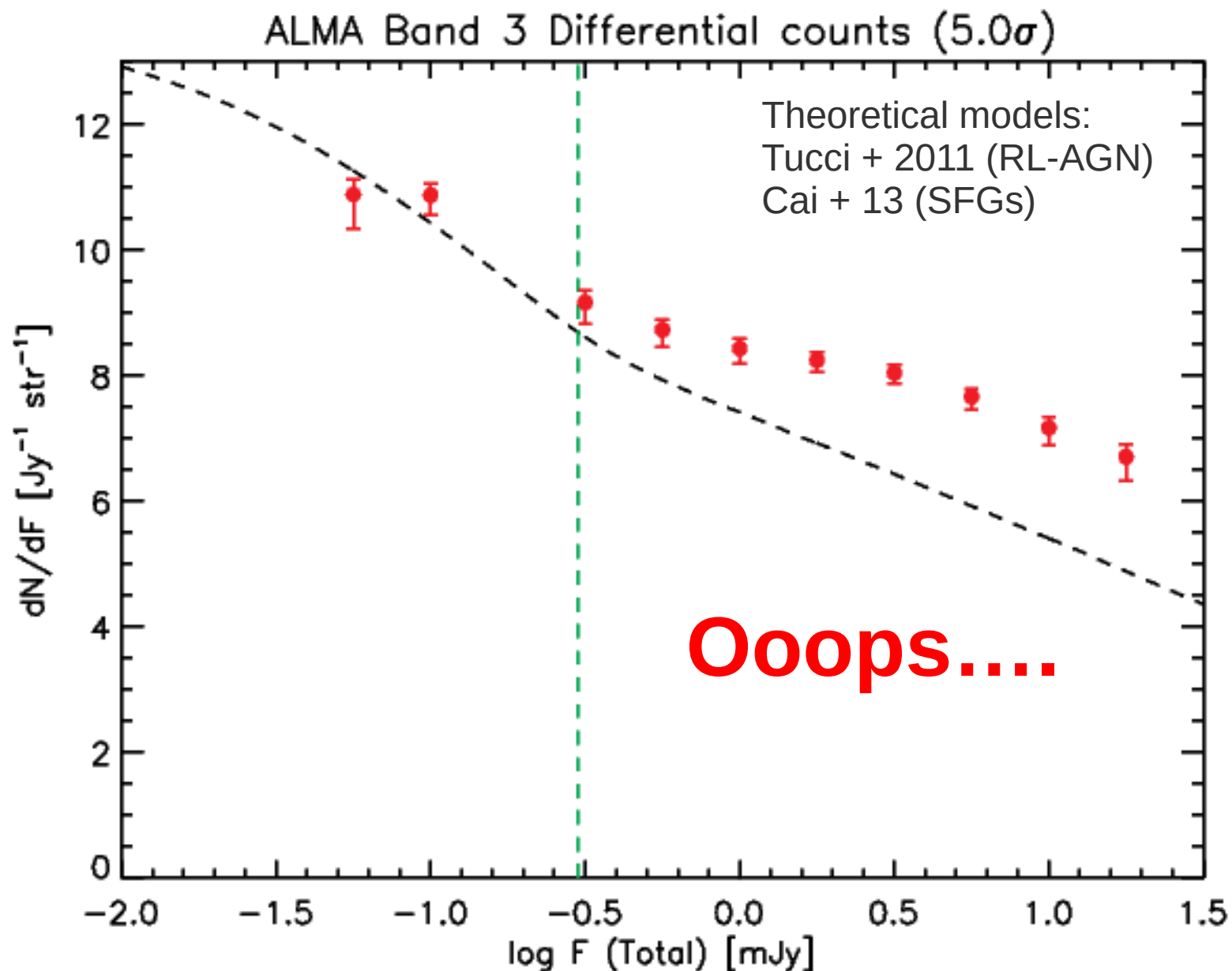
$C(S_i)$ is the completeness at S_i ,

Flux boosting $F_{\text{AUTO}} / F_{\text{total}}^{\text{input}}$ 

ALMA Band 3 number counts



ALMA Band 3 number counts



ALMA Band 3 number counts

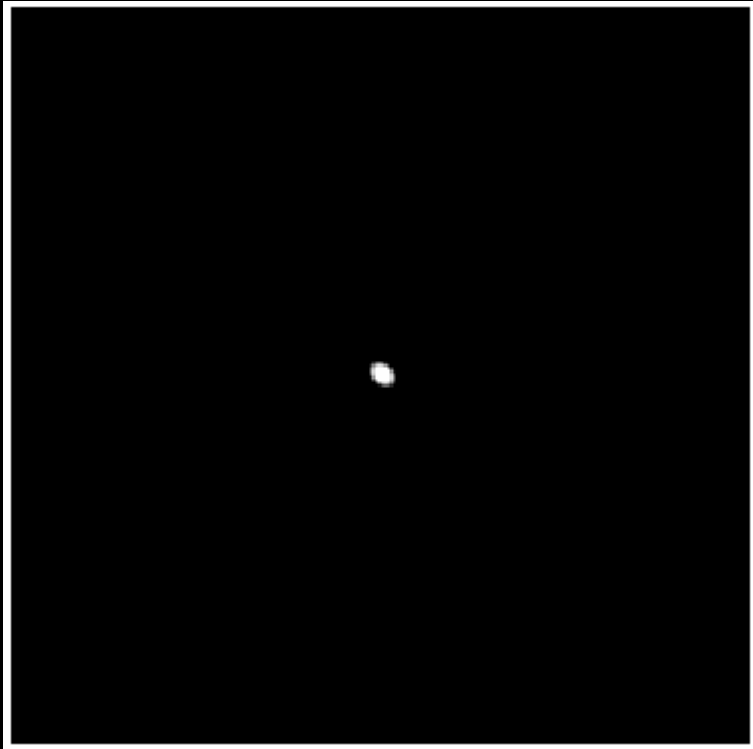


Image of a band 3 calibrator

- Are there any other serendipitous detections around?



Compute number counts !

ALMA Band 3 number counts

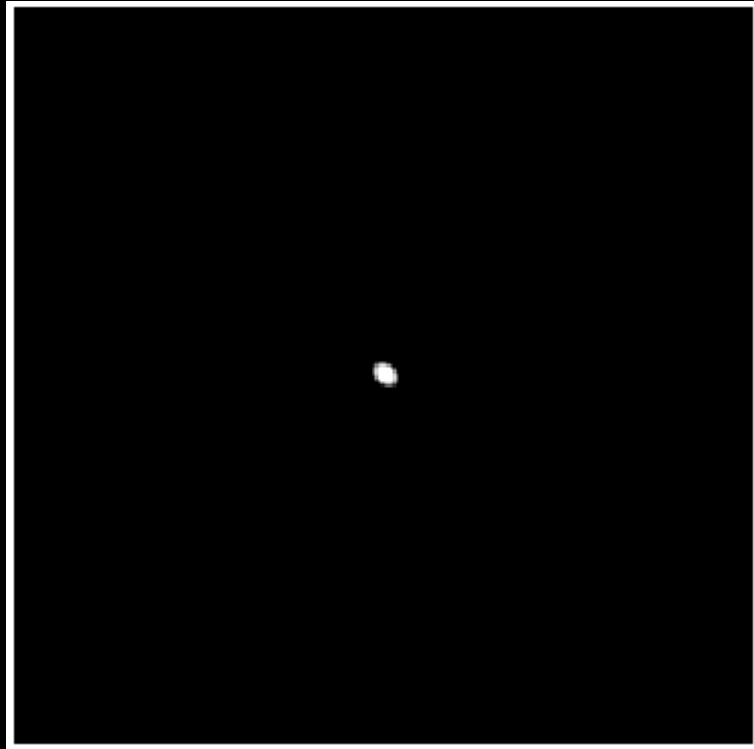
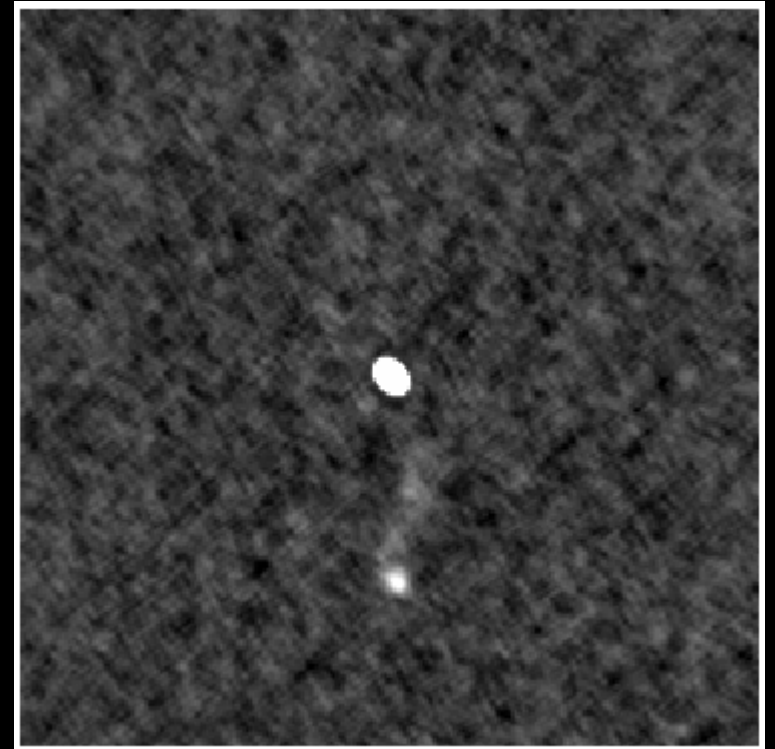


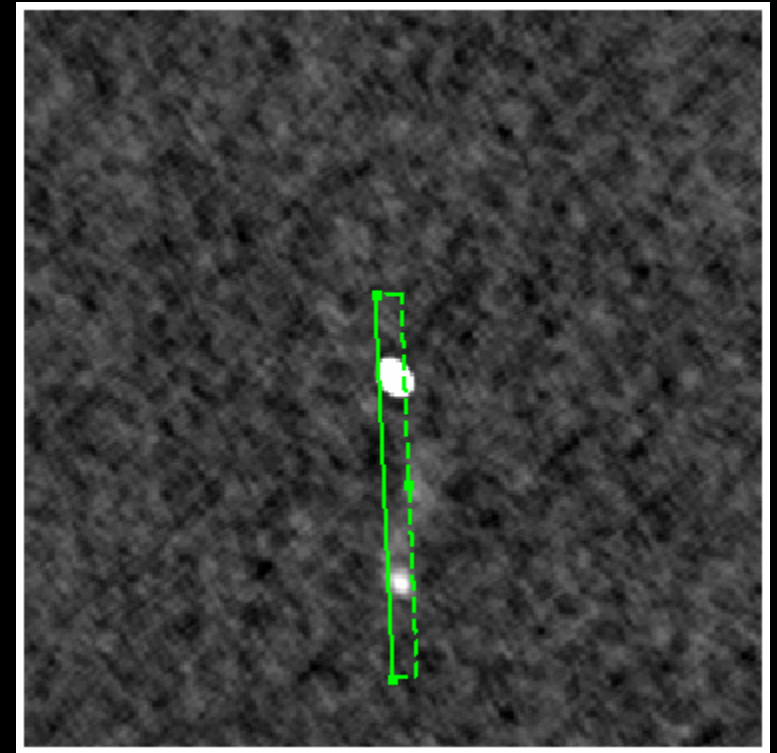
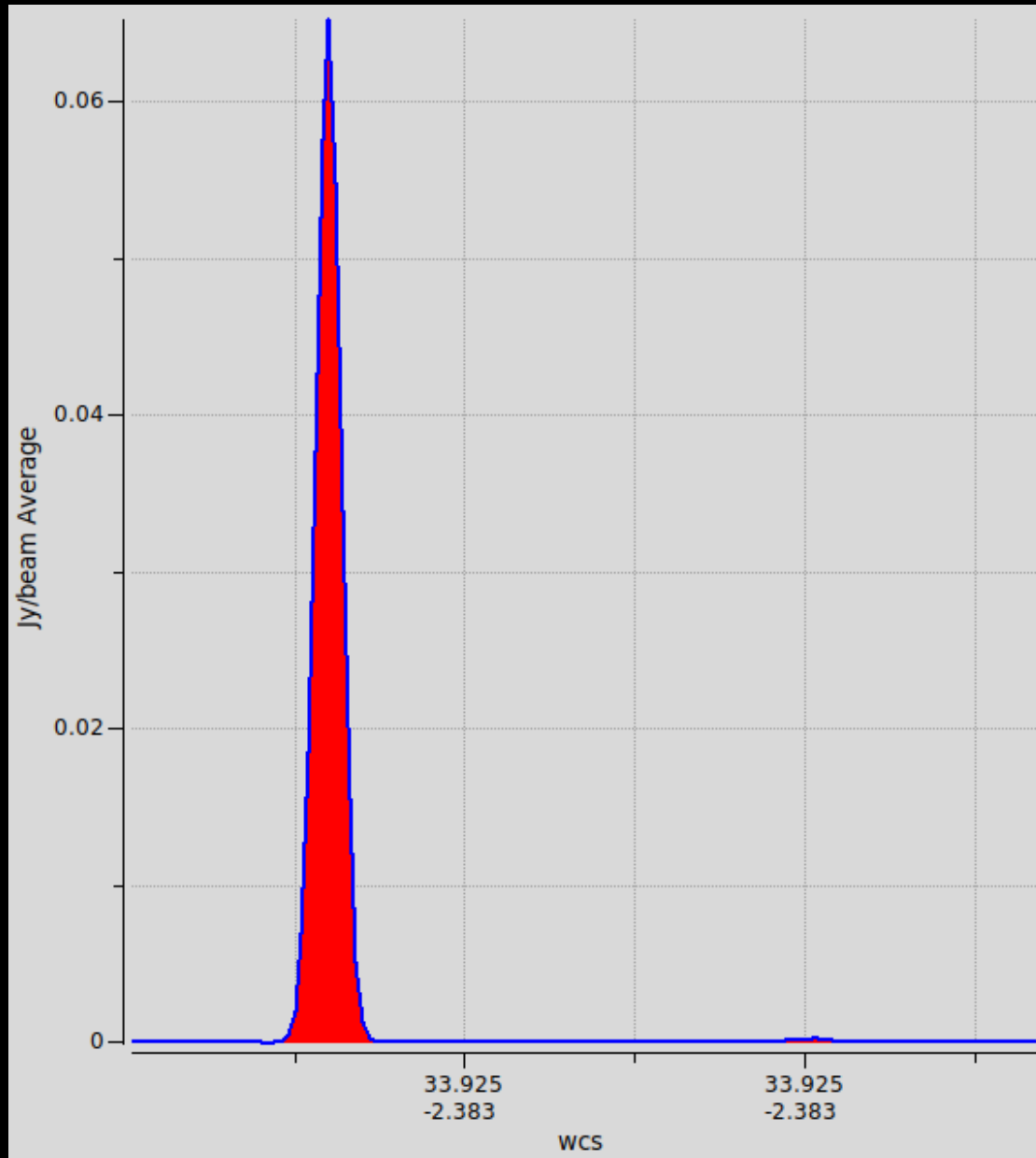
Image of a band 3 calibrator



Same calibrator (after changing scale)

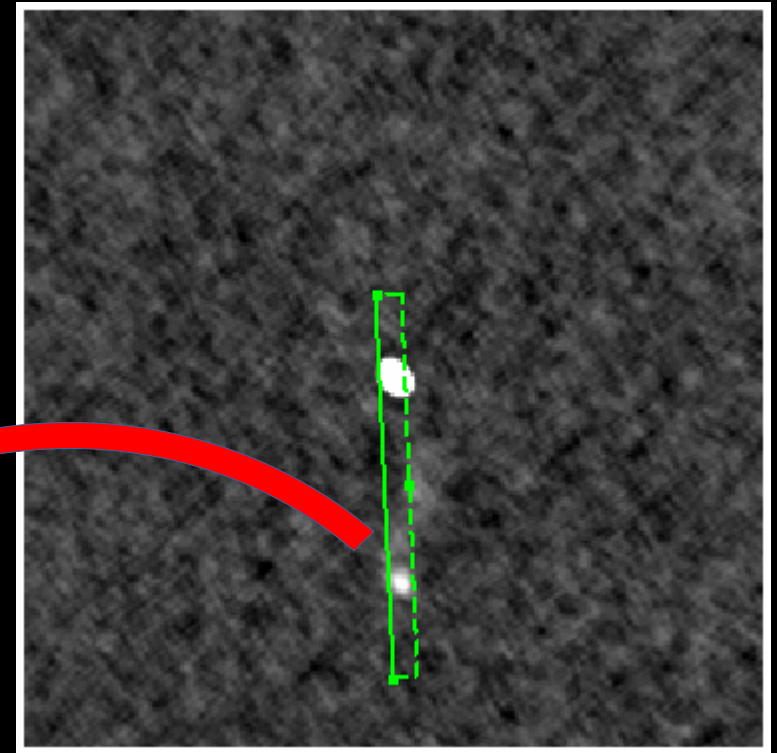
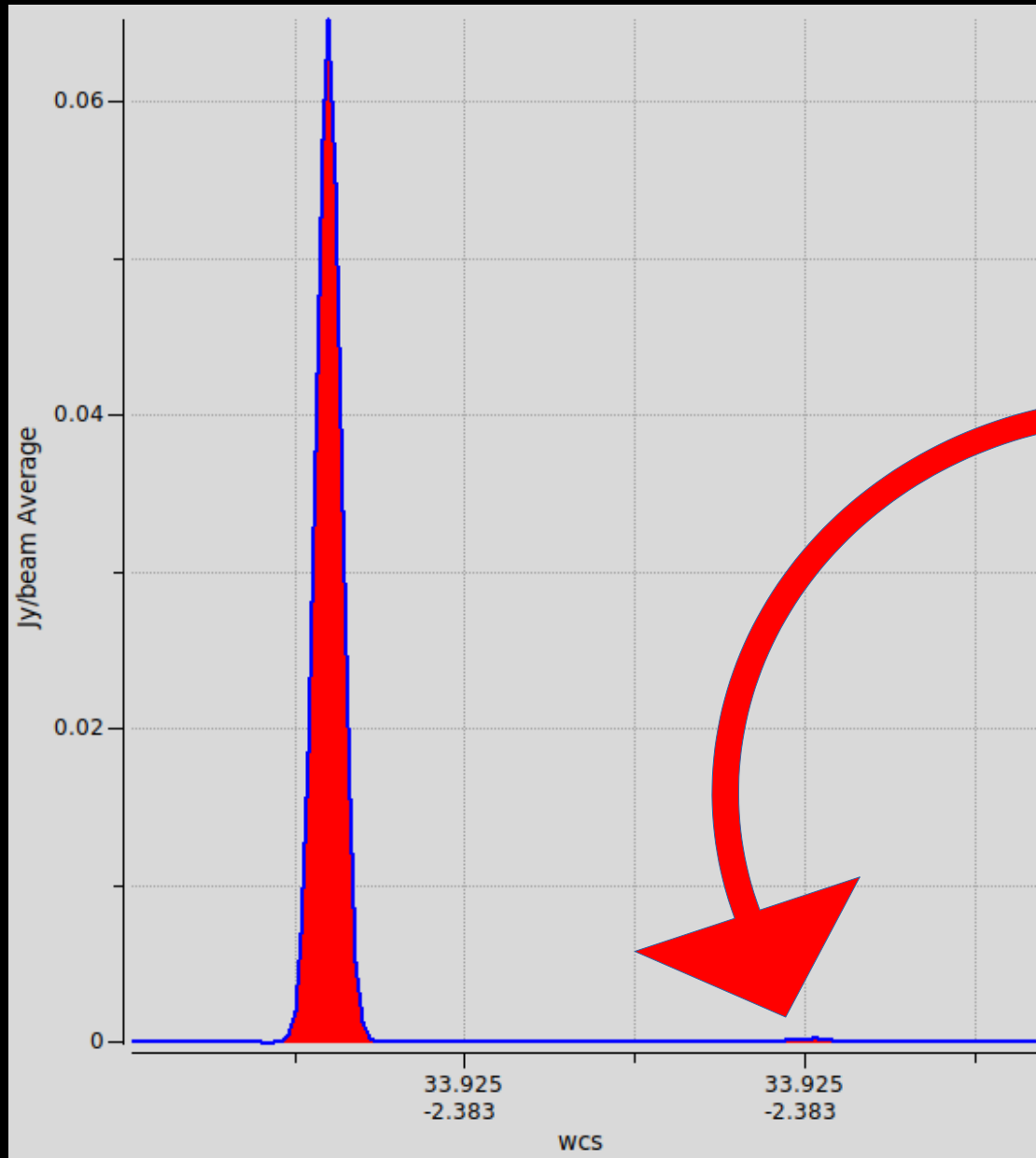
Extensions of the central calibrator are counted as (multiple) serendipitous sources

ALMA Band 3 number counts



Same calibrator (after changing scale)

ALMA Band 3 number counts



Same calibrator (after changing scale)

ALMA Band 3 number counts

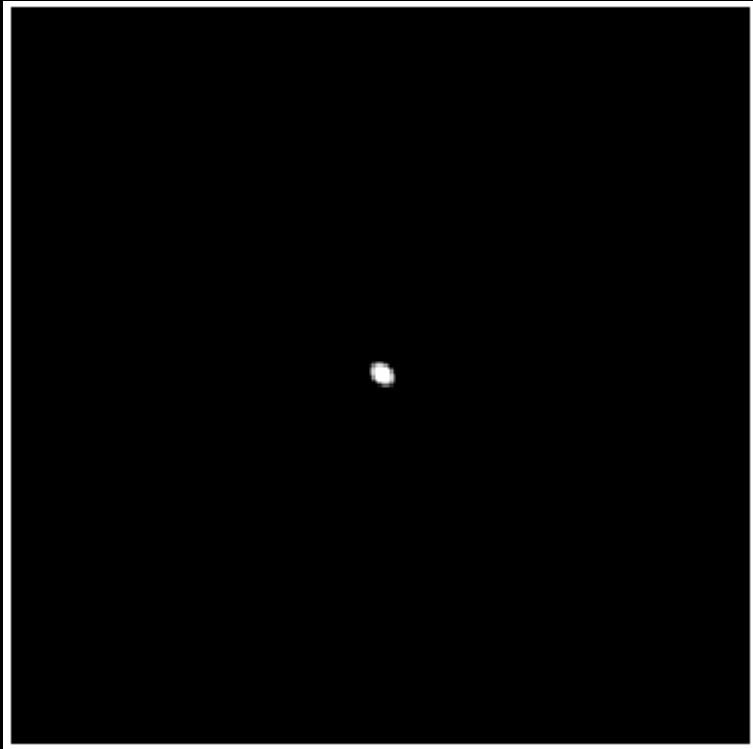
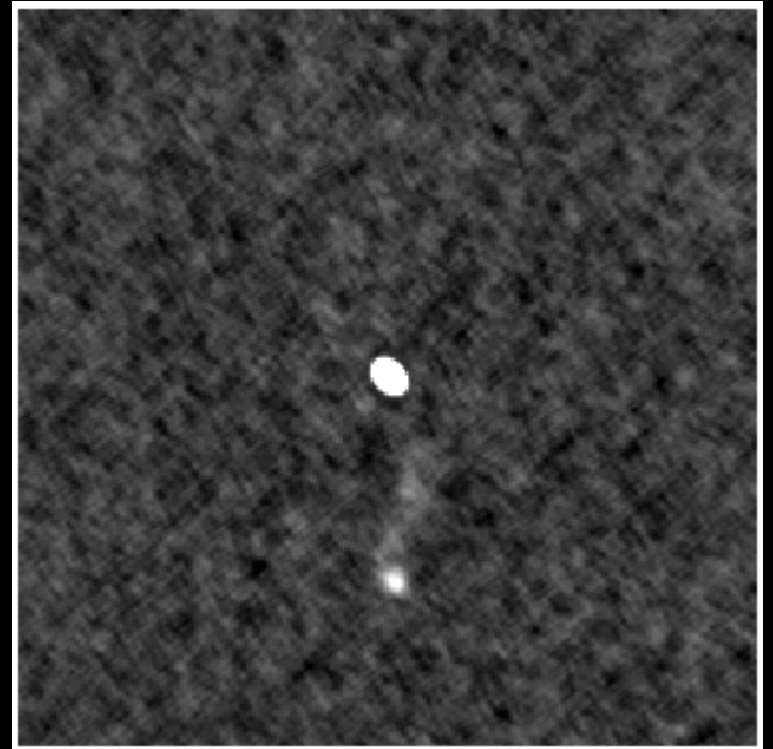


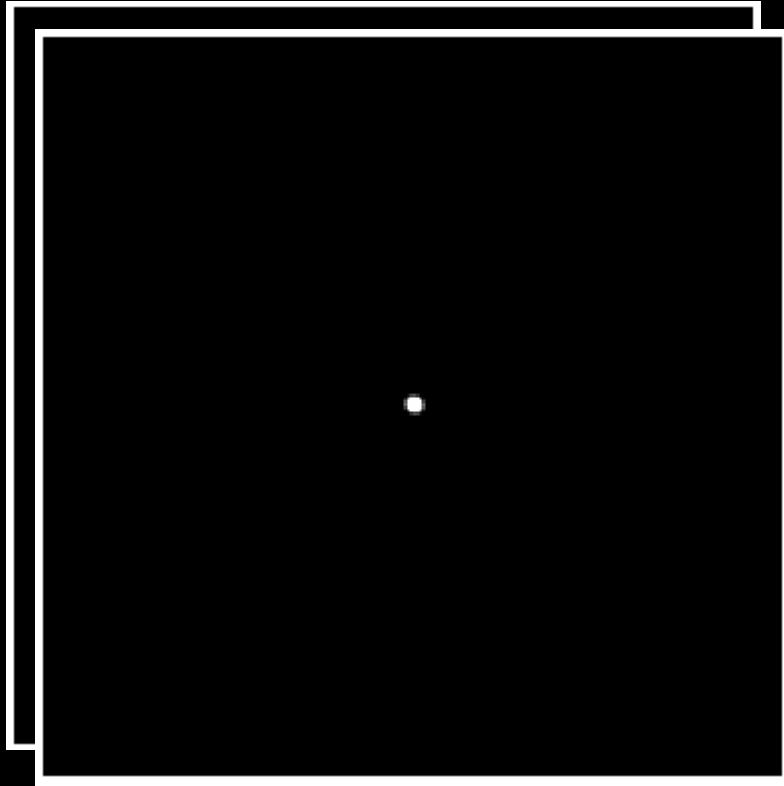
Image of a band 3 calibrator



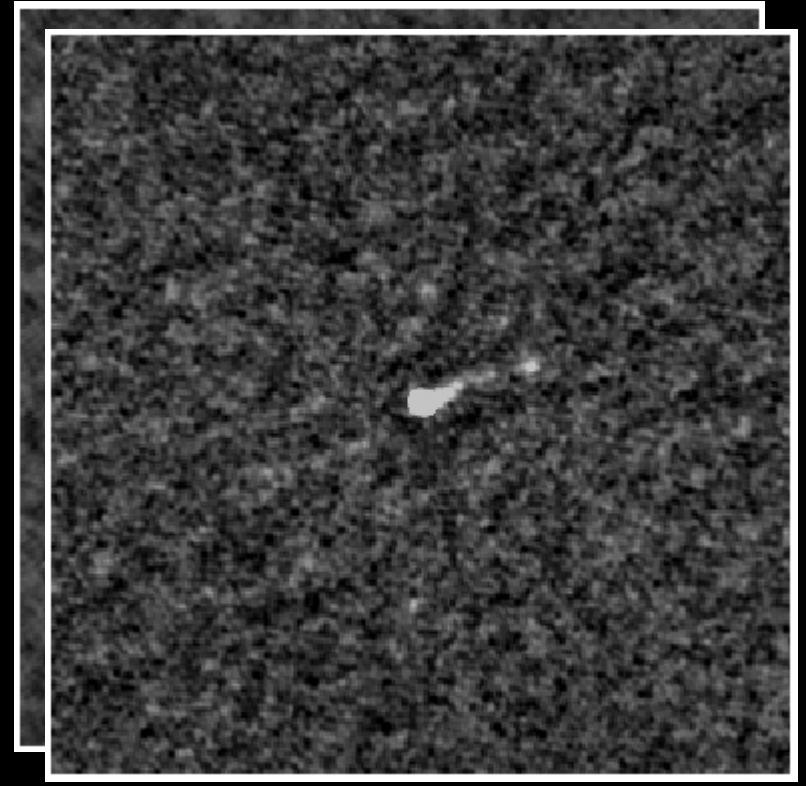
Same calibrator (after changing scale)

Extensions of the central calibrator are counted as (multiple) serendipitous sources

ALMA Band 3 number counts



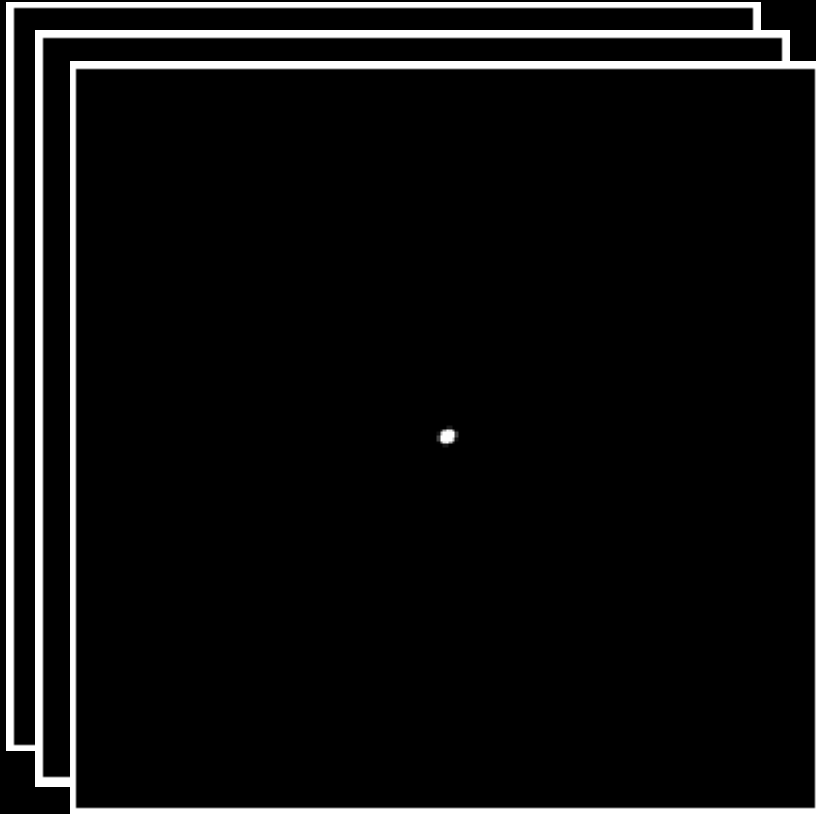
Another calibrator



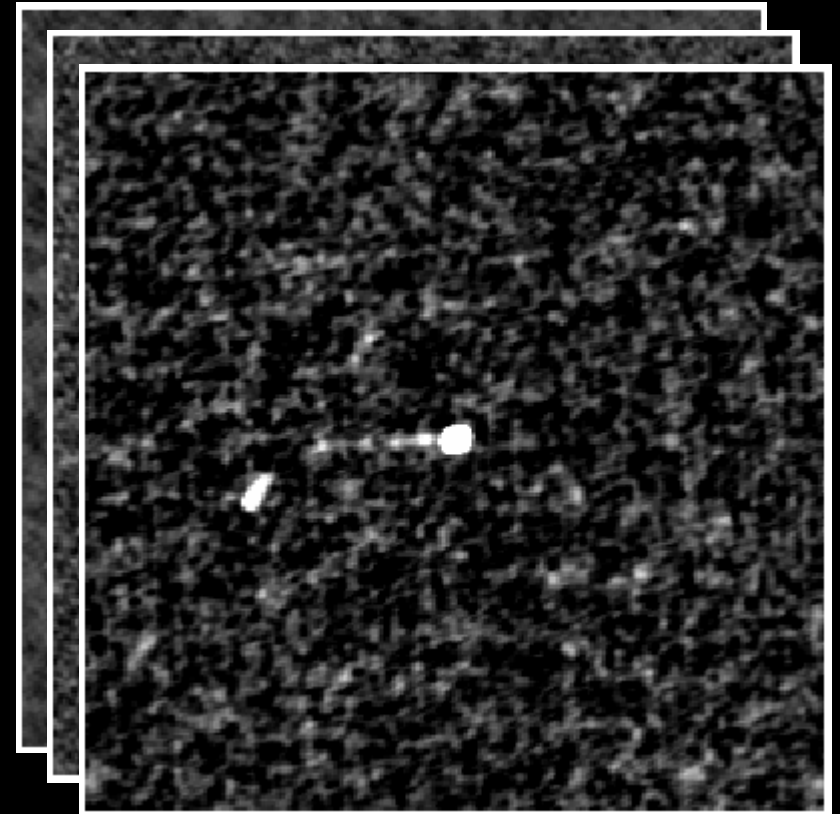
Same calibrator (after changing scale)

Extensions of the central calibrator are counted as (multiple) serendipitous sources

ALMA Band 3 number counts



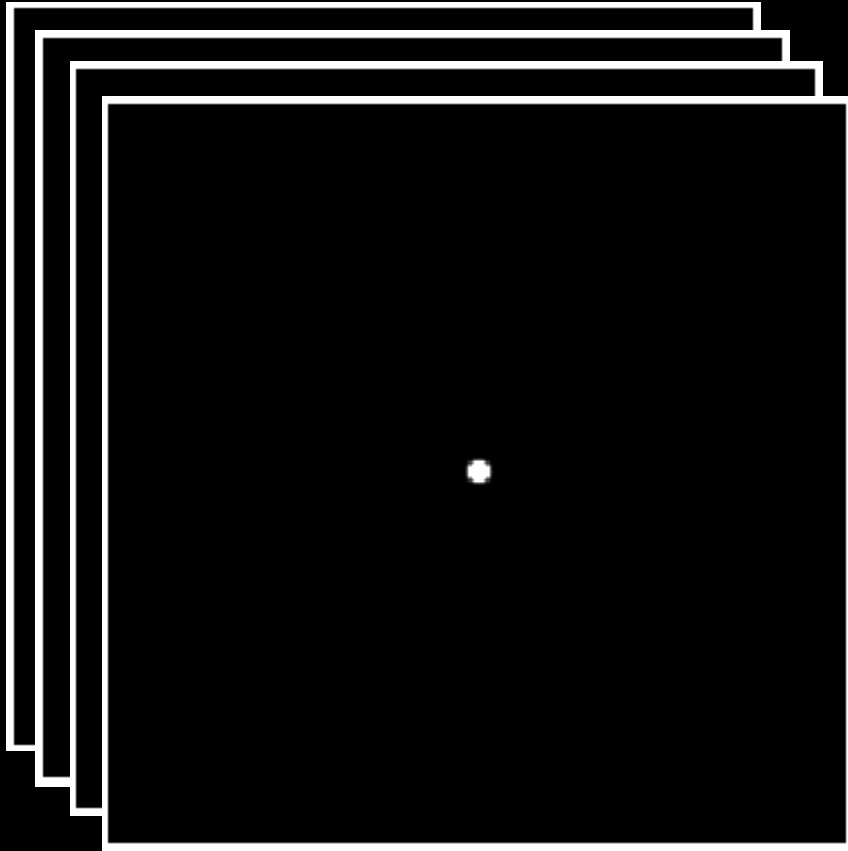
Another calibrator



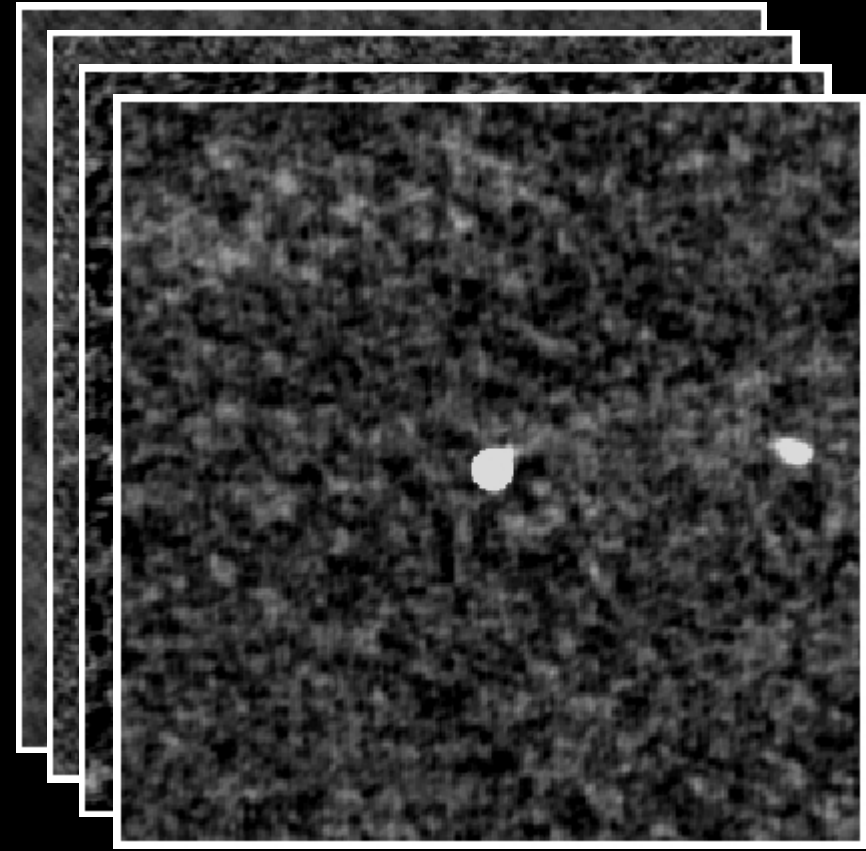
**Same calibrator (after
changing scale)**

**Extensions of the central calibrator are counted as (multiple)
serendipitous sources**

ALMA Band 3 number counts



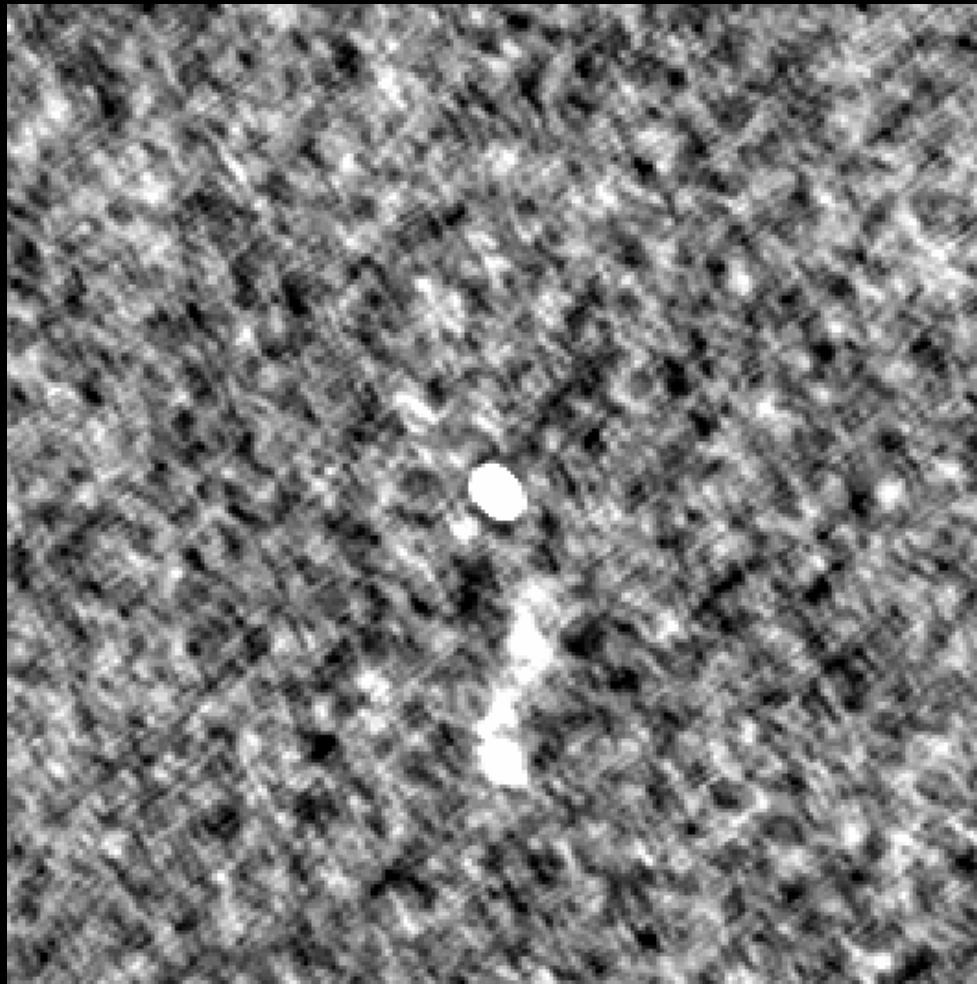
Another calibrator



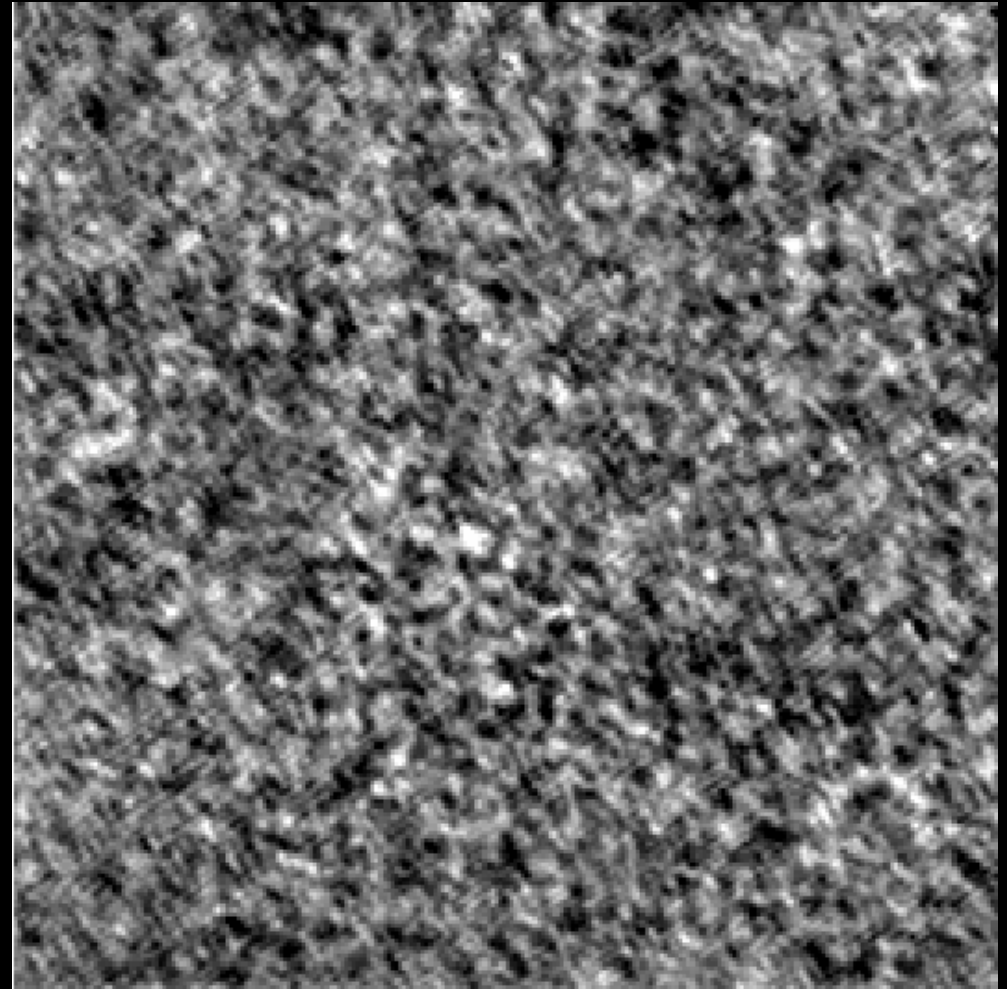
Same calibrator (after changing scale)

Extensions of the central calibrator are counted as (multiple) serendipitous sources → Down to a few sigma, it becomes more and more difficult to identify these contaminants by eye

ALMA Band 3 number counts - simulations

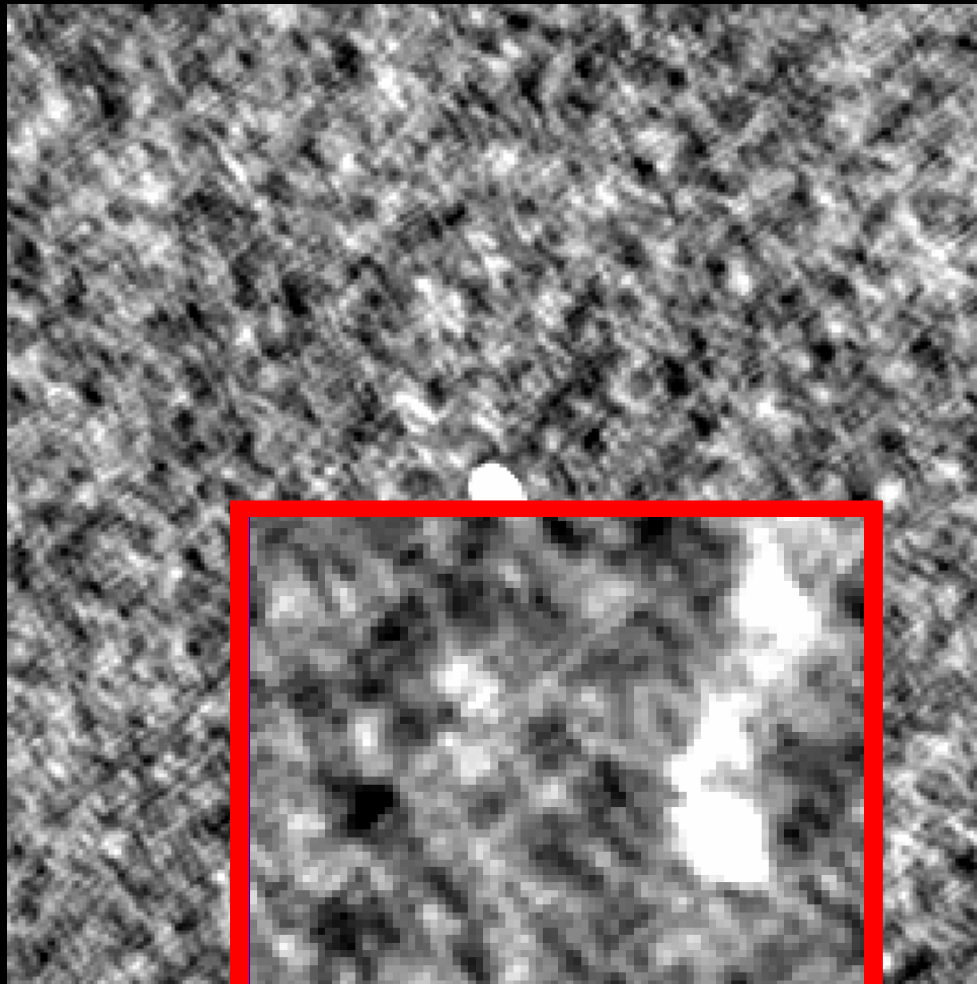


Original image

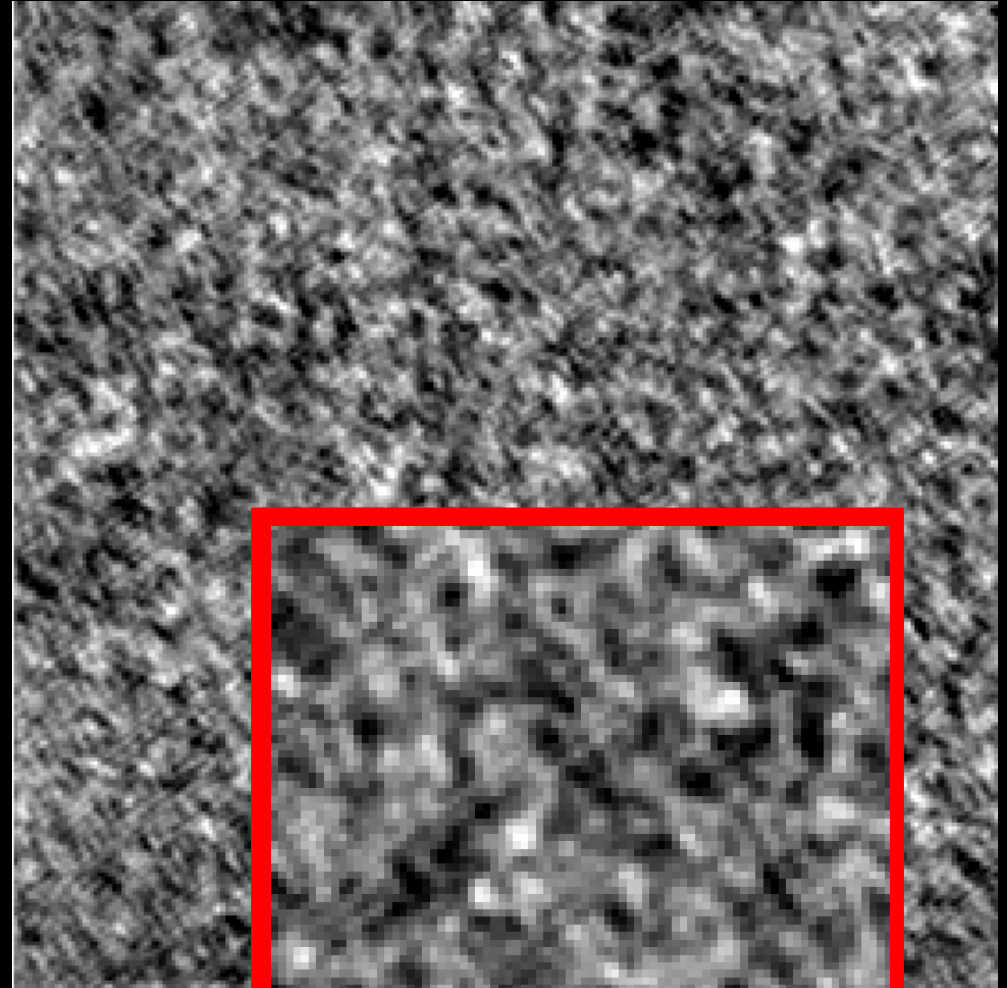


Simulated image

ALMA Band 3 number counts - simulations

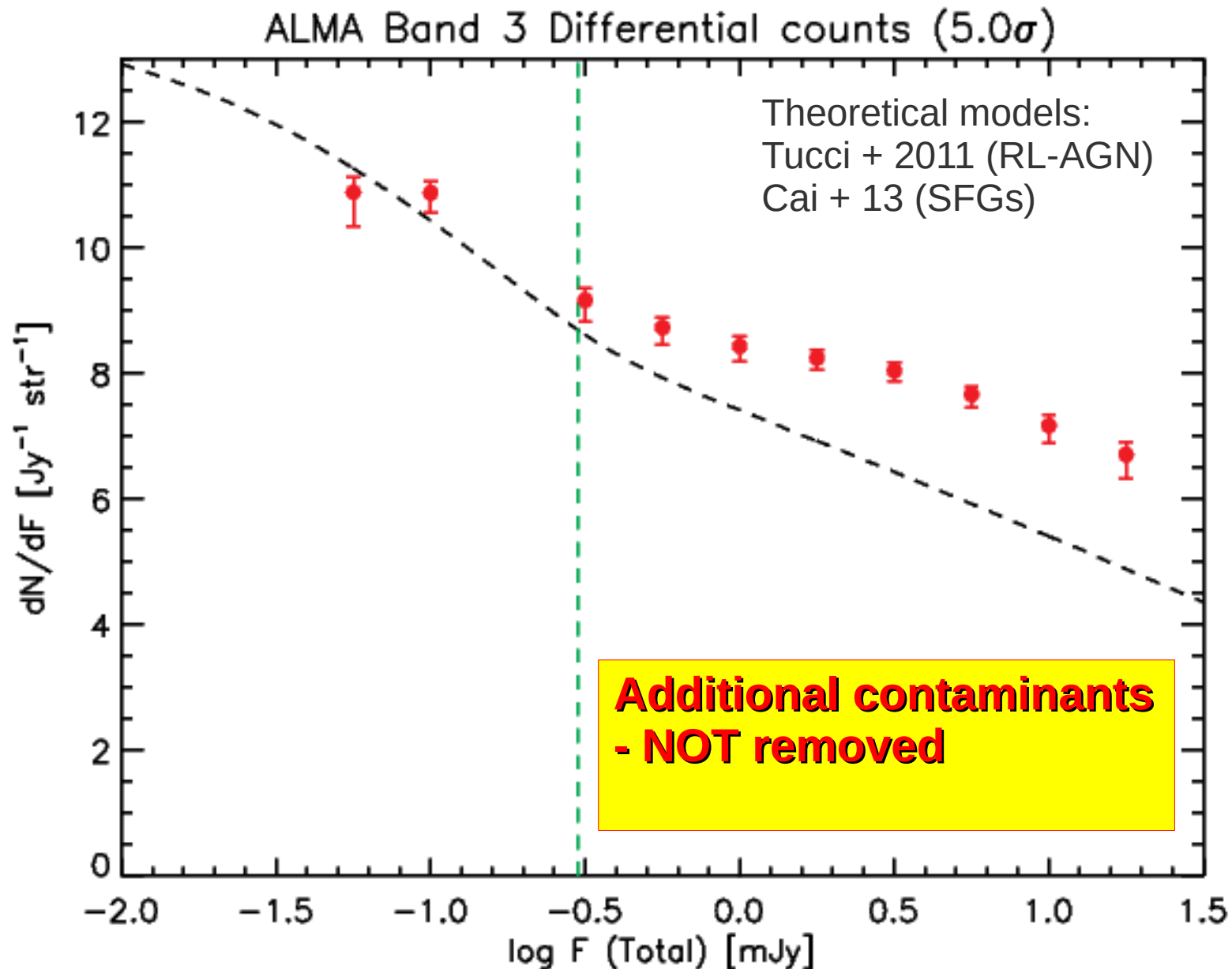


Original image

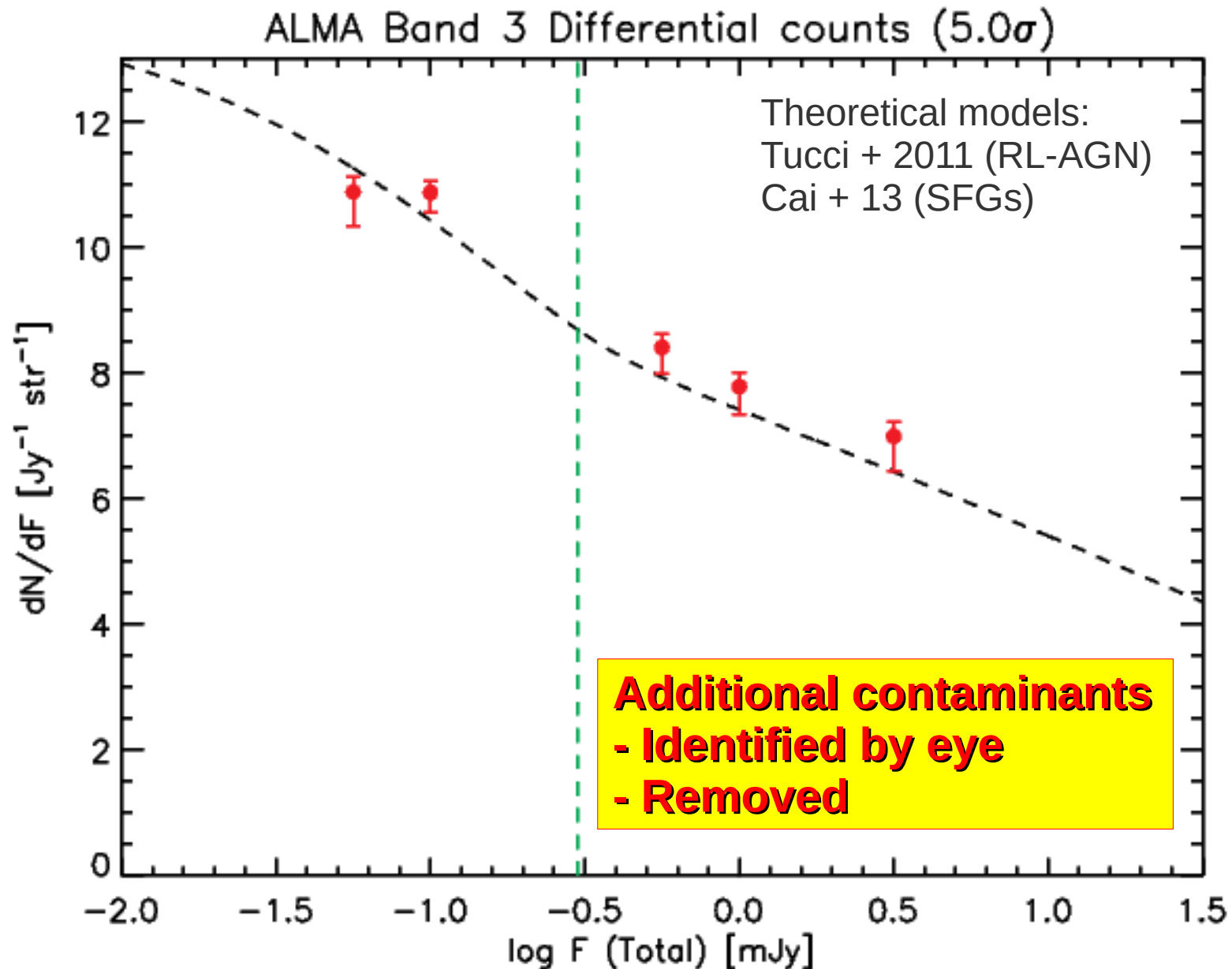


Simulated image

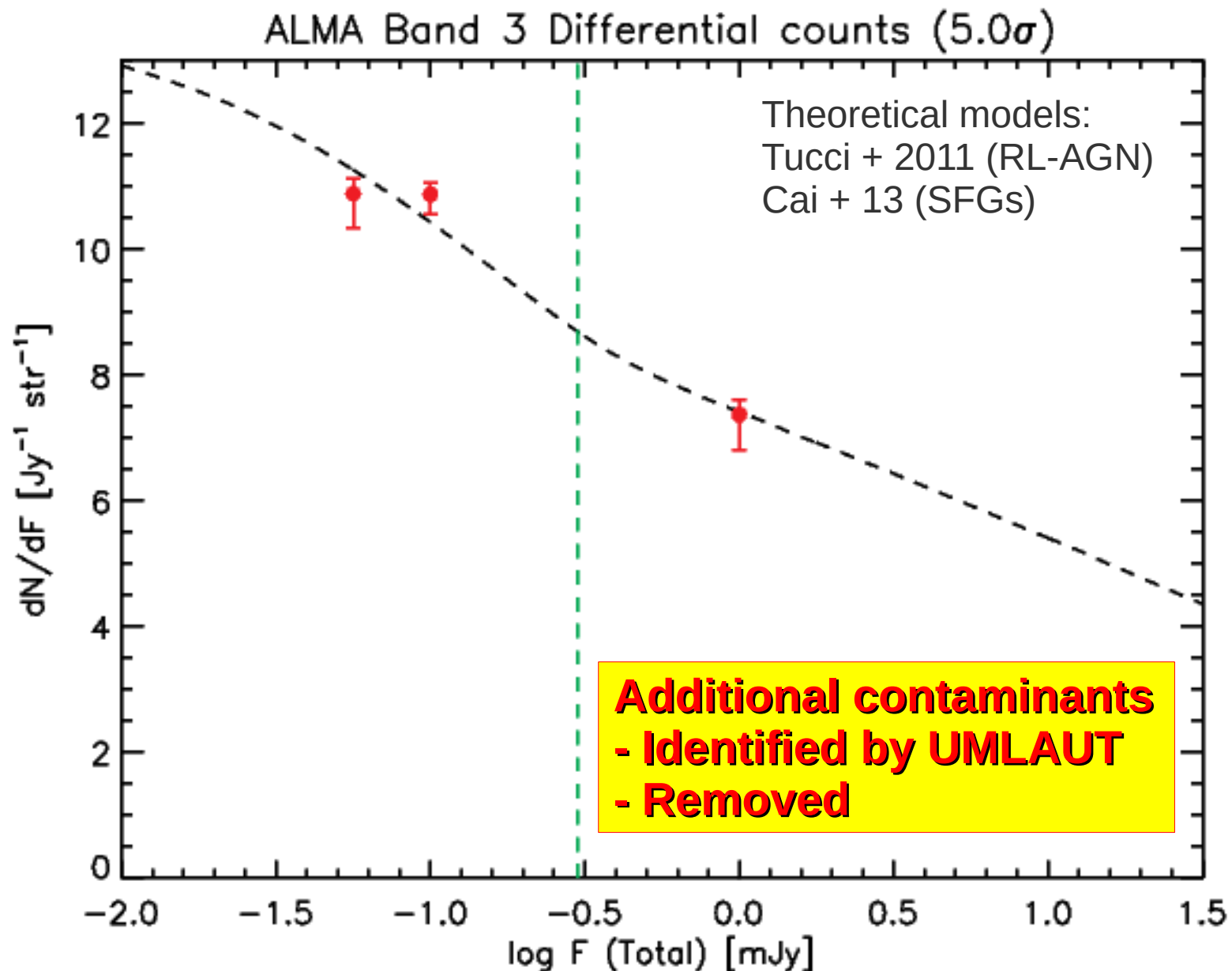
ALMA Band 3 number counts (5σ)



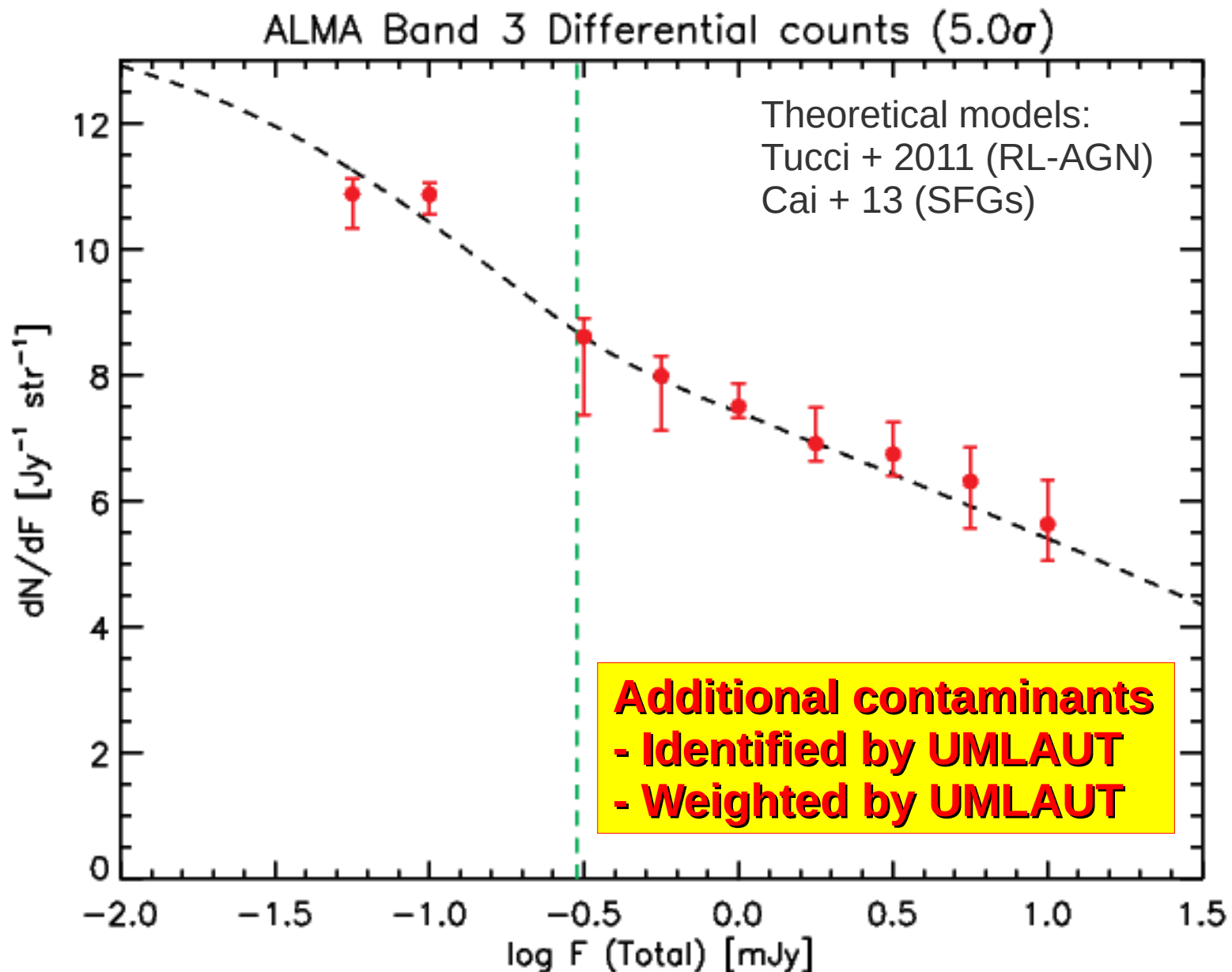
ALMA Band 3 number counts (5σ)



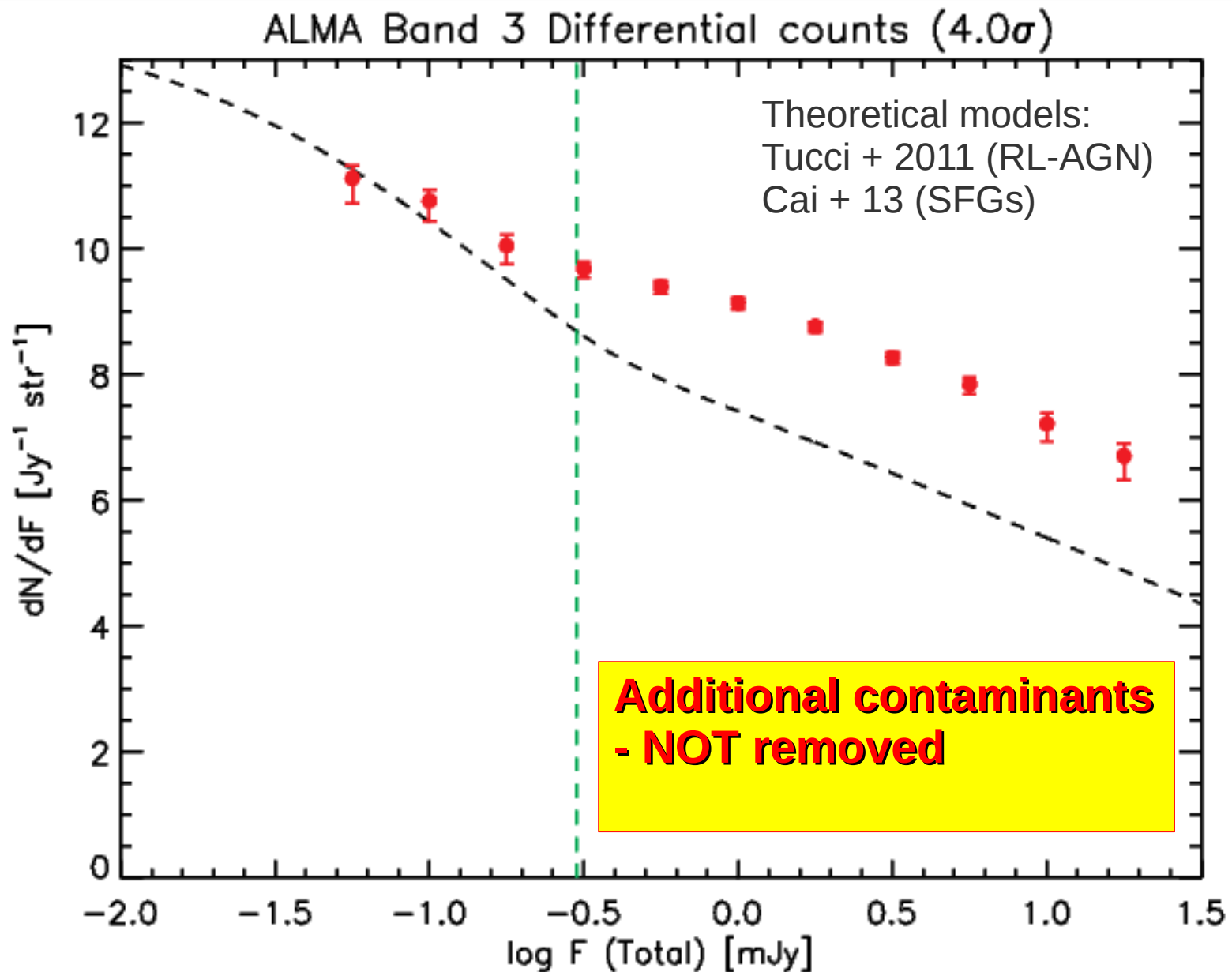
ALMA Band 3 number counts (5σ)



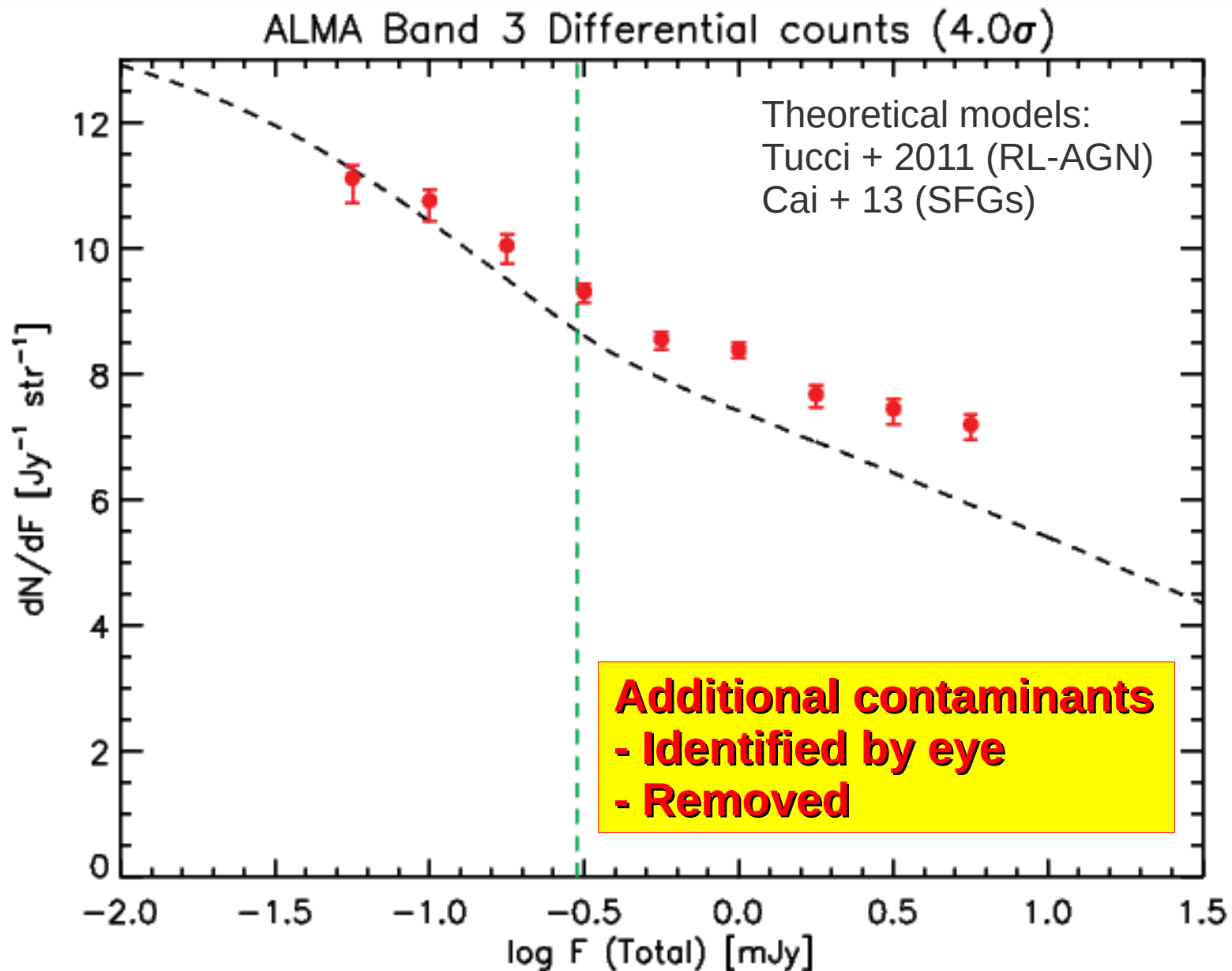
ALMA Band 3 number counts (5σ)



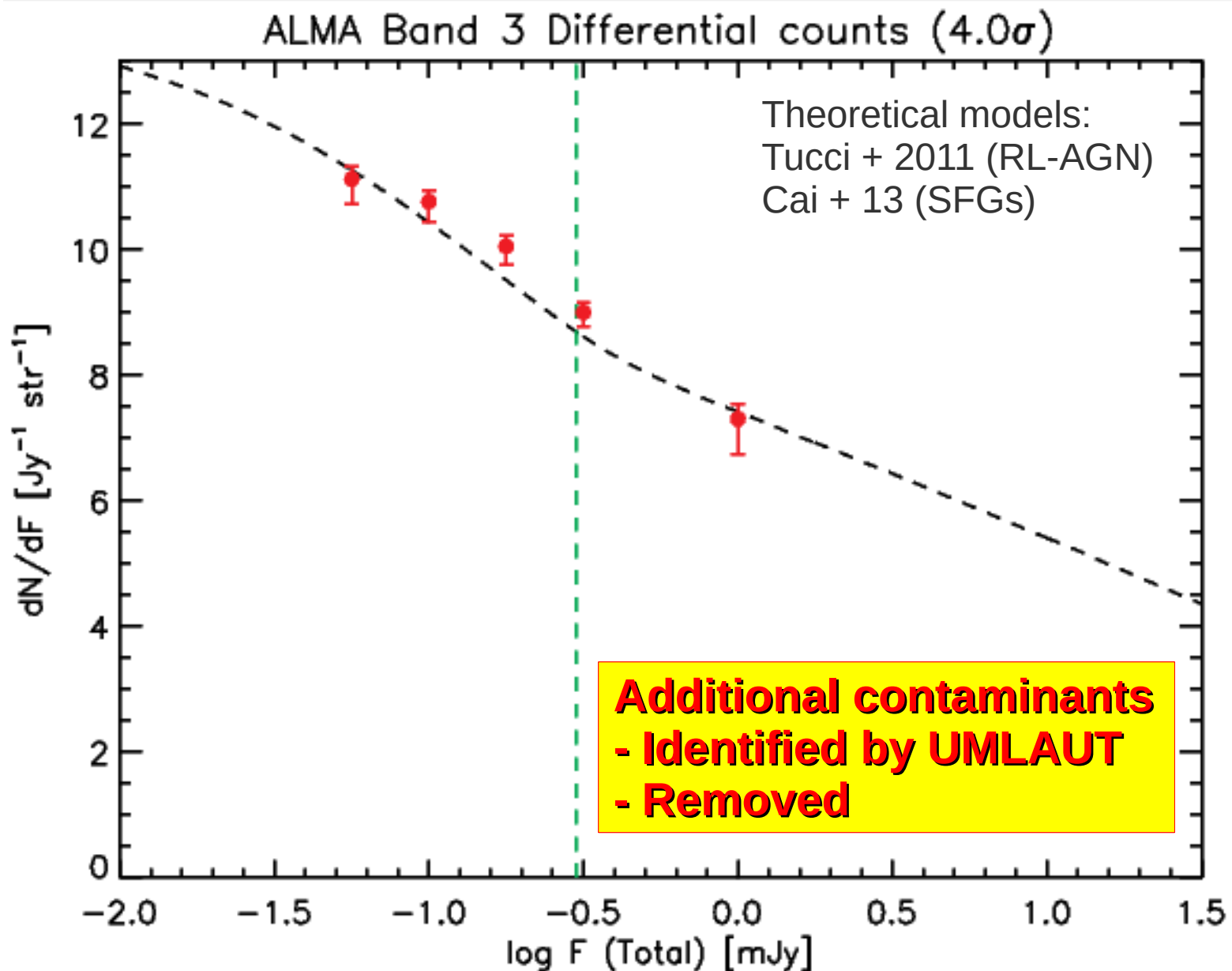
ALMA Band 3 number counts (4σ)



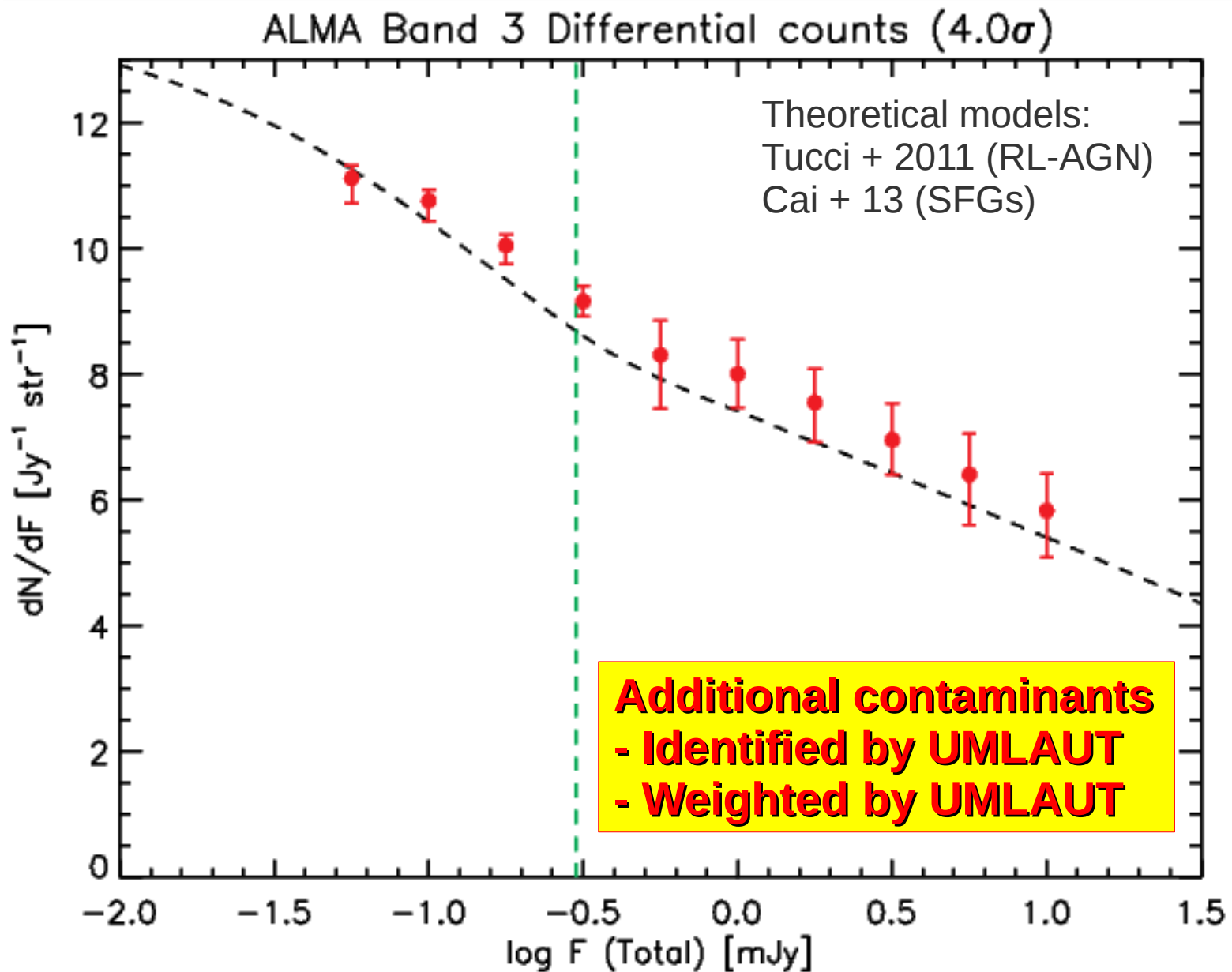
ALMA Band 3 number counts (4σ)



ALMA Band 3 number counts (4σ)



ALMA Band 3 number counts (4σ)



UMLAUT

Easily adaptable to a wide variety of problems



Inter- vs. Intra-Speaker Variation in Mixed Heritage Syntax: A Statistical Analysis

Federica Cognola^{1*}, **Ivano Baronchelli^{2*}** and Evelina Molinari³

¹ Dipartimento di Studi Europei, Americani e Interculturali, Sapienza Università di Roma, Rome, Italy, ² Dipartimento di Fisica e Astronomia, Università di Padova, Padua, Italy, ³ Istituto Culturale Mòcheno, Palù del Fersina, Italy

OPEN ACCESS

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Spain

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Kleanthes K. Grohmann,

Based on the novel data pertaining to five syntactic phenomena (the position of the finite verb in embedded clauses, in sentences with a modal verb, negative concord, the position of focused light/heavy objects in main clauses with a complex tense and scrambling) in the heritage language Mòcheno collected via original fieldwork, we show that there are two populations – one exhibiting intra-speaker variation between German and Italian word orders, and one lacking it; and these two populations are the result of diatopic variation and, to a lesser extent, of diastratic variation. The results achieved using quantitative statistical analysis are partially convergent with those arrived at via the traditional theoretical syntax for Mòcheno, but our analysis has allowed us to shed new light on a series of phenomena that have been neglected or poorly understood thus far. More specifically, and for the first time, we discovered that there is a micro-variation resulting from diastratic (age) variation within the Roveda variety, which represents the

Views

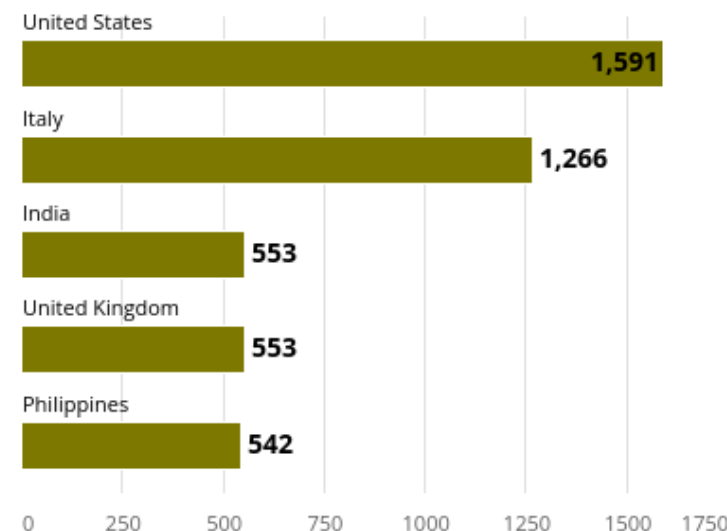
8,866 ⓘ

Frontiers

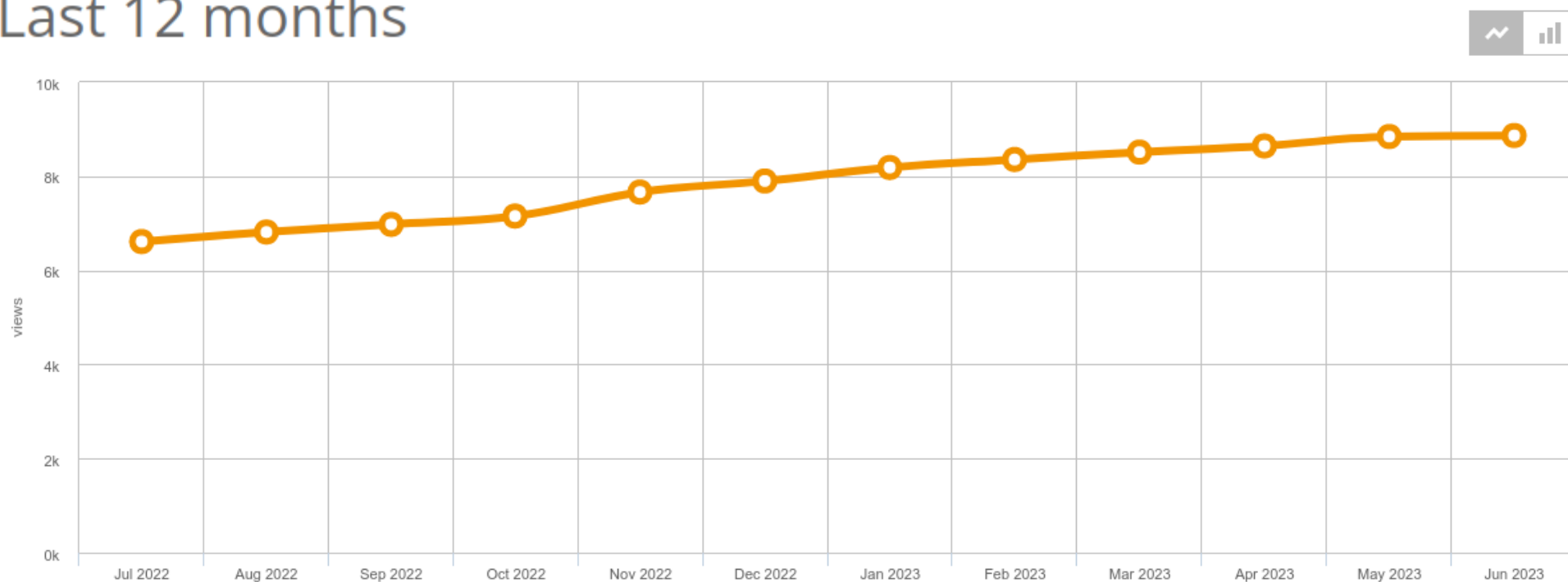


This article has more views than **93%** of all **Frontiers** articles.

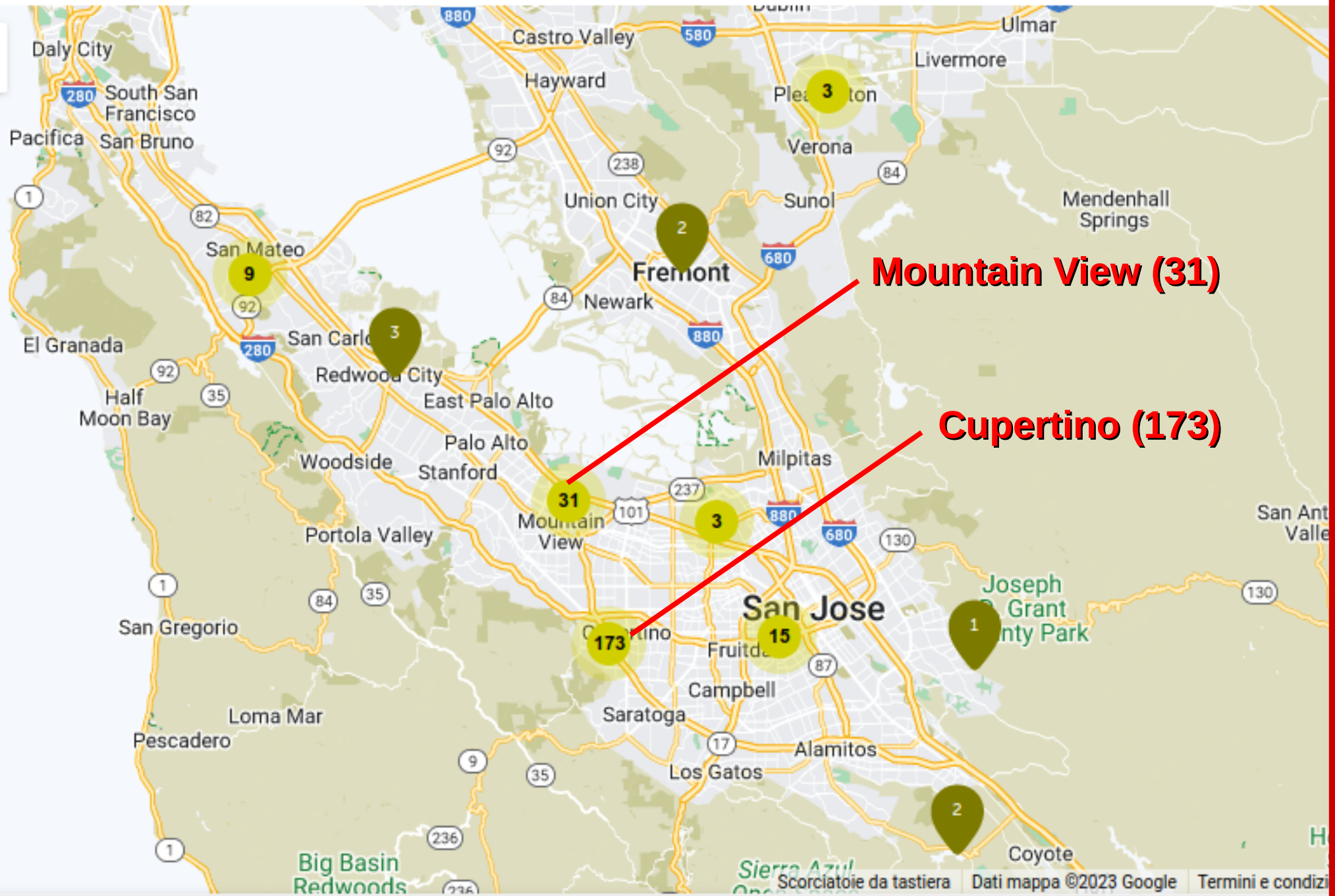
Top countries

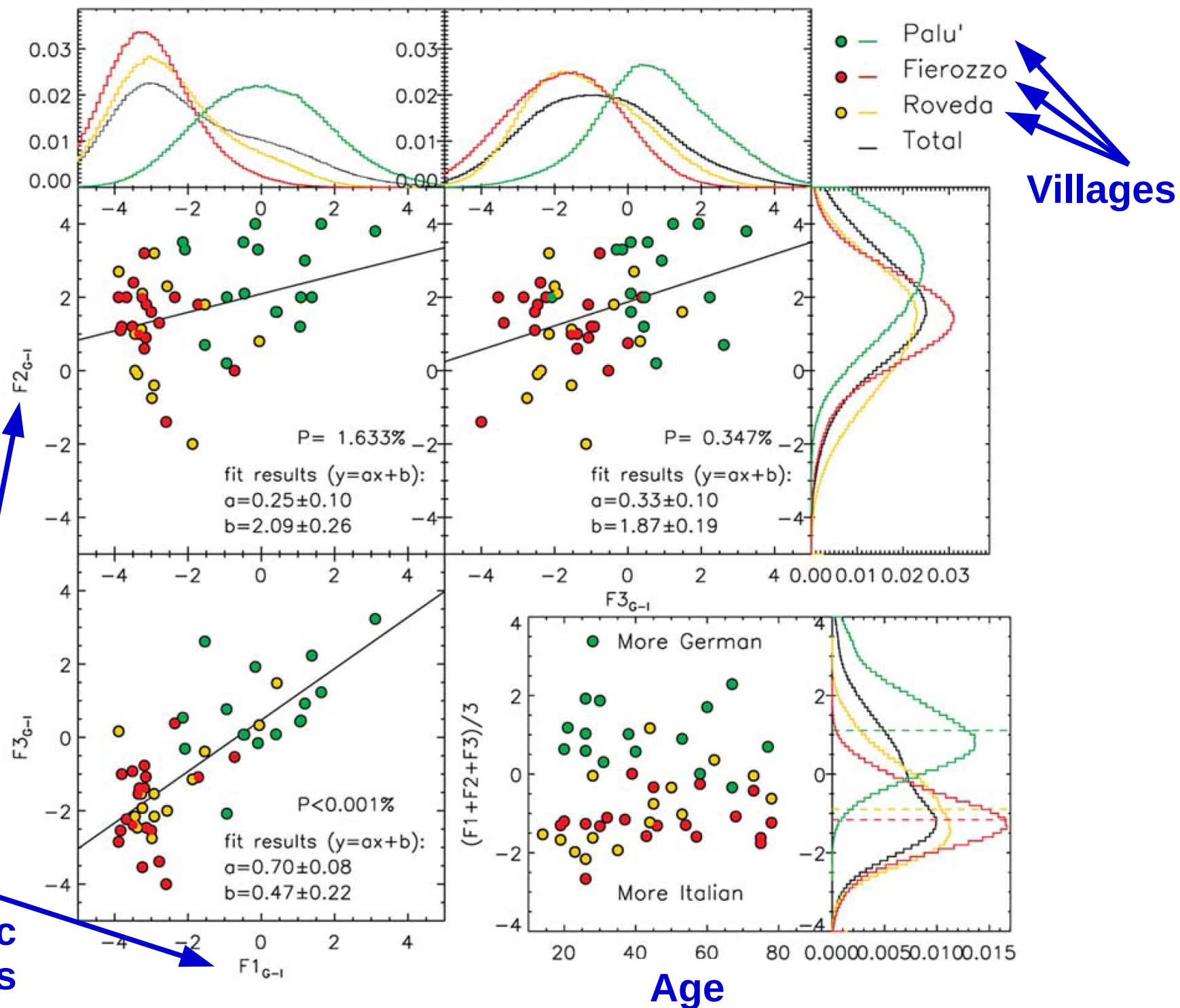


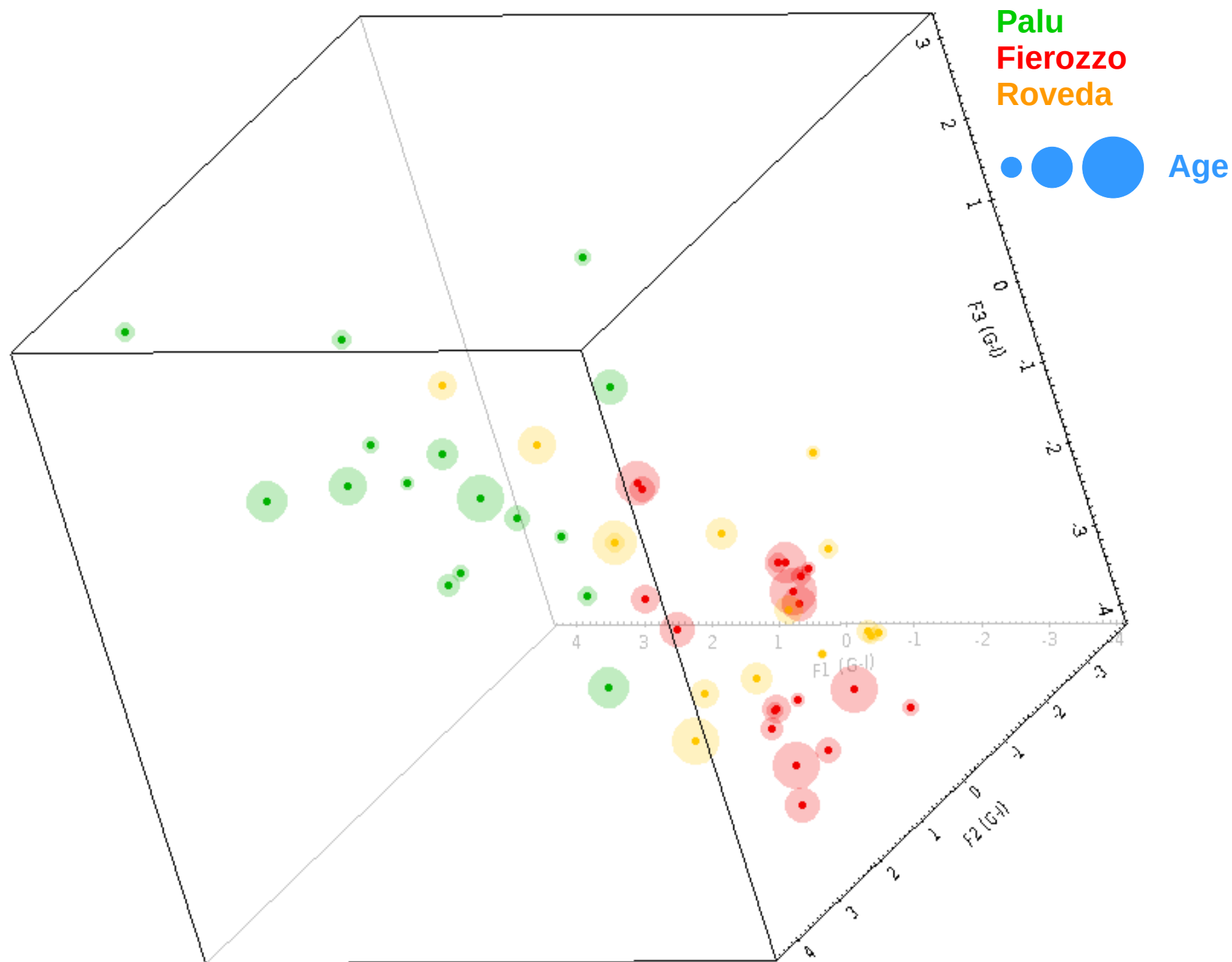
Last 12 months



Demographics



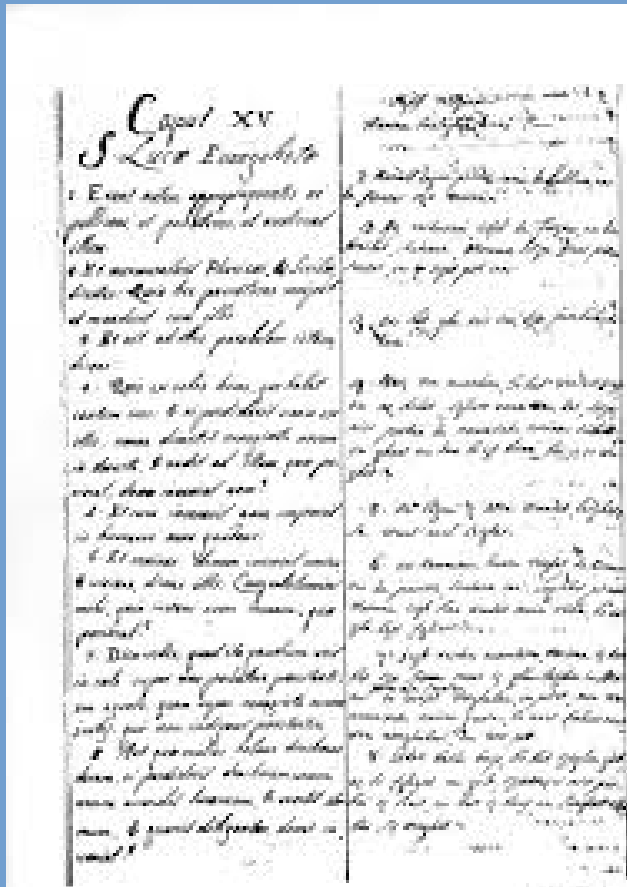




1811 – Napoleon wants a census of all the dialects spoken in the Italian kingdom

→ First important official document written in Mocheno language

Who wrote it?



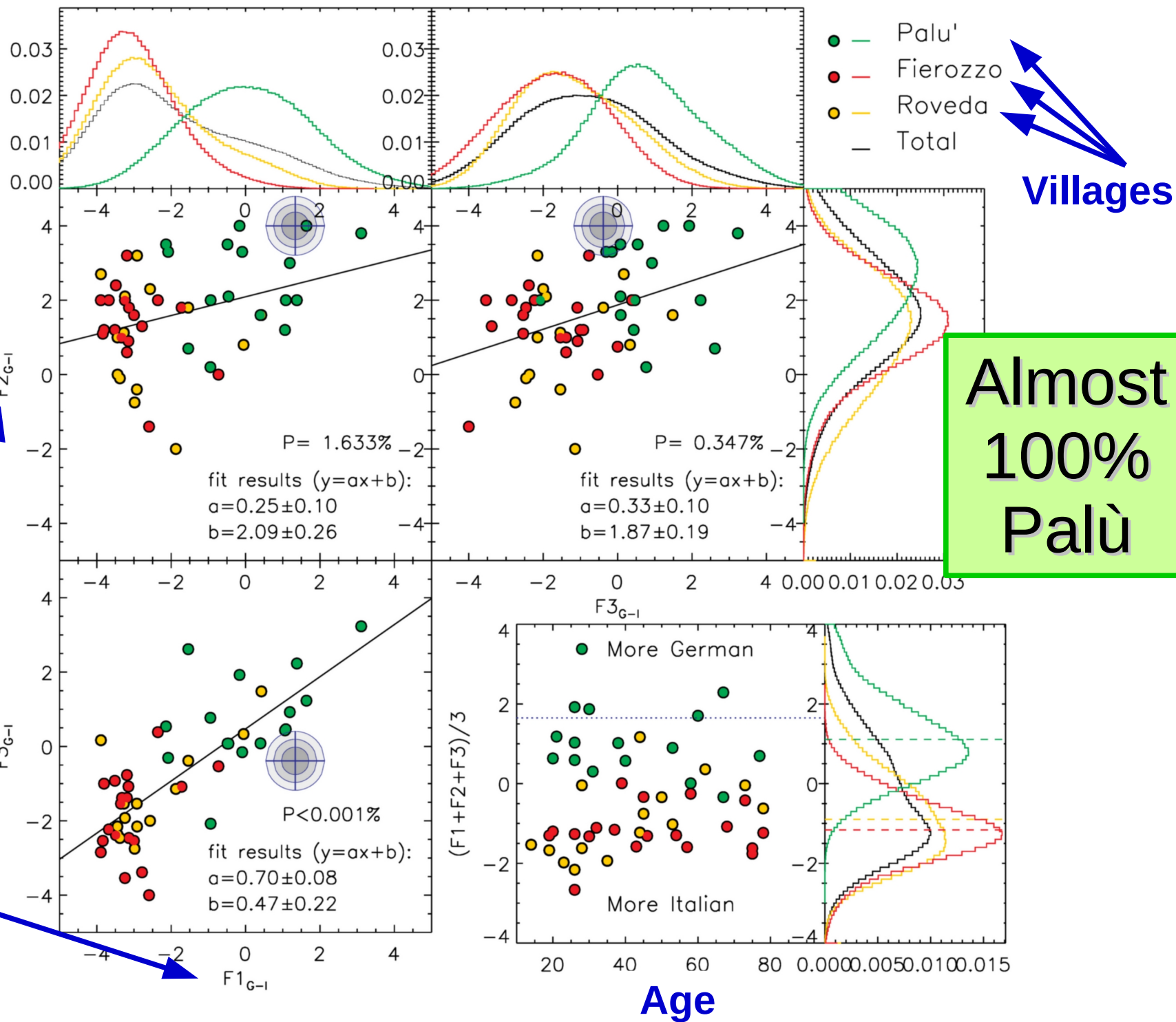
1811 – Napoleon wants a census of all the dialects spoken in the Italian kingdom

→ **First important official document written in Mocheno language**

Who wrote it?

Linguistic markers





Ongoing development:

- Code optimization (same algorithm - faster run);
- Translate UMLAUT in different languages (Python)
- Parallelization of the code (NVIDIA-CUDA);
- Improvement of the current dimensionality reduction strategy
 - example1: using a PCA-like strategy
 - example2: by replacing each of the input dimensions with random ranks and comparing the output variance.

Conclusions:

- New facilities producing unprecedented amount of data are going to be available in the next few years: every aspect of the treatment of such data must be automated;

UMLAUT is:

- a ML algorithm based on the analogy paradigm;
- extremely (and easily) adaptable to very different contexts;
- freely available online

3 applications were presented:

- Identification of single spectral lines
- Recognition of ancient linguistic variants
- Identification of contaminants in sub-mm number counts