



Additional Cosmology For Euclid Italia 2023

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Outline

- General Remarks on Additional Cosmology
- From Small to Large Scales
 - Strong Lensing
 - Clusters
 - CMBX
 - Theory
- Pre-Launch Considerations
 - **KPs**
 - Coordination
 - Common Interfaces







Additional Cosmology: Structure & Goals

- Definition: Cosmology Science for Euclid, Complementary and Coordinated with Primary Probes
- Additional Cosmology WGs
 - Strong Lensing: Lens Dark Matter, Background Sources, Cosmography, ...
 - Clusters: Dark Matter Profile, Hot Gas Properties, Arc Statistics, Weak Lensing, ...
 - **CMBX**: Integrated Sachs Wolfe, Lensing, Thermal and Kinetic Sunyaev Zel'dovich, ...
 - Theory: Dark Energy & Modified Gravity, Dark matter, Initial conditions, ...
- Observables, Models, Codes
- Coordination with Primary Probes, Cosmological Simulations











Additional Cosmology: Consortium Interface & Organization

- Additional Cosmology absorbs Information from SWGs through Regular Monthly Teleconferences
- Information flows from and to
 - Editorial Board, monitoring Consortium Papers, Organizing Internal Referees, Coordinate Publication
 - Inter-Task Force Science Coordination
 - Science Coordination Group
- Examples:
 - Choices for Internal Paper Referees with SWG Leads,
 - Coordination of CLOE Development within IST Likelihood and Non-Linear Coordinated with Model Building and Modules in SWGs,
 - SWG Lead Rotation Plan, ...
- Information from and to SGS, OUs is spontaneous and functional to needs, the Launch offers the opportunity to make these links stronger, structural.











Strong Lensing: Organization

SWG coordinators: R. Gavazzi (LAM, Marseille), M. Meneghetti (INAF-OAS, Bologna), J.P. Kneib (EPFL, Lausanne, rotating off)

OU-SHE-SL coordinator: F. Courbin (EPFL, Lausanne)

Active members: ~25 people (I, CH, F, UK, US, NL, S, B, D)

Two active KPs (pre-launch) with dedicated working sub-groups for each ongoing paper.

Telecons and meetings:

- One general telecon per month (usually on the last thursday of each month, details of the redmine)
- Telecons dedicated to each KP (once per month)
- In person meetings: resumed last summer (Leiden), second meeting in December (Marseille). Next meeting likely in May (TBD)

Communications:

- Wiki: <u>https://euclid.roe.ac.uk/projects/slswg/wiki/Wiki</u> (being re-styled)
- Mailing list: ecsls@googlegroups.com
- **Slack:** ecslsworkspace.slack.com









Strong Lensing: KP Plan

Two Key Projects coordinated by the SLWG (and ongoing paper leading authors, **purple=italian leadership**):

• KP-SL-1: Galaxy-scale Strong Lenses (G. Despali, G. Vernardos) (GAEV-SWG, OU-SHE)

- Realistic image simulations of strong lensing galaxies (B. Metcalf)
- Algorithms for finding strong lensing galaxies (B. Clément, L. Leuzzi)
- Lens modeling techniques (J. Nightingale)
- Information content on the galaxy evolution model (A. Sonnenfeld)

• **KP-SL-2: Cluster-scale Strong Lenses (L. Moustakas, G. Mahler)** (*WL-SWG, CL-SWG, GAEV-SWG, OU-SHE*)

- Realistic image simulations of strong lensing clusters and groups (P. Bergamini)
- Algorithms to find strong lensing features in galaxy clusters and groups (G. Angora)
- Lens modeling techniques (G. Mahler)
- Strong lensing by cluster substructures (M. Meneghetti)











Strong Lensing: Consortium Interfaces & Coordination

- Report to ECB last June 2022
- Very strong connection with OU-SHE-SL and SDC CH to build the SL processing functions (e.g. sharing person power; actively participating in the coding activity)
- Good interface with other SWGs. Sharing tools (e.g.: image euclidization tool HST2Euclid with LU-WG; image simulations of double AGNs with GE-WG; lens mapping tools with CL-WG) and expertises.









Strong Lensing: Status of Activities

- **Completion of the Pre-launch KPs:** 1 paper completed and to be submitted to ECEB in the next days (led by L. Leuzzi); overleaf existing for other 2 papers; other papers planned to be completed by June
- Largest effort dedicated to **developing, testing, and validating the SL lens finding and modeling pipeline:** ongoing work, but in advanced stage
- Definition of Q1 and DR1 projects
- **Readiness for Q1:** Aim at quick release of automated lens finding (trained with sims with their own limitations); involvement of citizen science to validate lens detections (ZooUniverse?)
- **Planning follow-up observations:** e.g. spectroscopy on 8m class telescopes (to be discussed and coordinated with other WGs); T. Collet led a joint Euclid/Rubin 4MOST proposal (4SLSLS, now absorbed into TiDES) to follow-up Euclid and LSST lens candidates: southern sky, 36000 fiber hours granted









Clusters: Organization

- ~150 Members: ~30% both Italy and France, 13% Germany, 7% both UK and Spain, 5% USA, plus few more people in Canada, Japan, Netherlands, Finland, Norway, Austria, Portugal...
- 11 Working Packages, 7 of them with an Italian co-lead
- 6 pre-launch KPs (see below)
- Communications:
 - General bi-weekly zoom calls, alternating a general SWG telecon and one reserved to the KP/WP leads
 - Wiki: <u>https://euclid-redmine.roe.ac.uk/projects/cgswg/wiki</u>
 - Strong connection/overlap with OULE3-Clusters (IT-lead)
- Meetings:
 - Twice per year (spring and fall), plus possible splinters at Consortia
 - next one, in Bologna, March 14th-16th, 2023









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Clusters: KP Plan: 6 KPs, 28 'papers', 4 already completed

- KP-CG-PL-1: Validation of the Cluster Cosmology Pipeline (managed by CG-WP3 IT-lead)
 - Paper 1.1: The HSC AMICO cluster catalog
 - Paper 1.2: Modelling of Euclid Cosmological likelihood-systematics and statistics (IT-lead)
 - Paper 1.3: Constraints from the AMICO HSC Cluster catalogs with the Euclid cluster cosmology pipeline (IT-lead)
- KP-CG-PL-2: Establishing the Cluster Selection Function (managed by CG-WP9 IT-lead Jointly with OULE3)
 - Paper 2.1: Comparison of Euclid Cluster selection function to other surveys
 - Paper 2.2: Preliminary evaluation of the Euclid cluster selection functions
 - Paper 2.3: Data-driven evaluation of the Euclid cluster selection function
 - Paper 2.4: Definition of the cluster sample for cosmology
- KP-CG-PL-3: Precision Simulations for Cluster Cosmology (managed by CG-WP9 IT-lead)
 - Paper 3.1: Calibration of the cluster halo mass function in vanilla LCDM models (IT-lead): PUBLISHED
 - Paper 3.2: Characterization of the covariances for cluster halo mass function and halo bias (IT-lead): 2 PUBLISHED
 - Paper 3.3: Baryonic effects on halo mass function and bias (IT-lead)
 - Paper 3.4: Halo mass function and bias in non-standard models (IT-lead)
 - Paper 3.5: 3D and 2D profiles and concentrations in LCDM and non-standard models
 - Paper 3.6: Convergence and shear maps of simulated clusters: calibration of WL masses (IT-lead): SUBMITTED
 - Paper 3.7: Impact of hydrodynamical modelling on the WL signal of massive halos
- KP-CG-PL-4: Euclid galaxy clusters legacy science (managed by CG-WP7 and CG-WP11, both with IT-lead)
 - Paper 4.1: Astrophysical properties of selected/unselected clusters: X-ray simulations
 - Paper 4.2: Optical counterparts of the X-ray selected clusters
 - Paper 4.3: The X-ray counterpart of the optical clusters
 - Paper 4.4: Impact of cluster physics on the cluster selection
 - Paper 4.5: Tools and forecasts for ICL detection
 - Paper 4.6-4.9: Definition and optimization of algorithms for the identification of clusters at z>1.5 and protoclusters (more papers, IT-lead)
- KP-CG-PL-5: Cluster mass-observable relation)managed by CG-WP5)
 - Paper 5.1: Covariance of ICM and optical/NIR observables in simulations (IT-lead)
 - Paper 5.2: Constraining WL mass-richness covariance from ICM observables
 - Paper 5.3: Mass calibration from galaxy kinematics (IT-lead)
- KP-CMBX-6: Meta-catalogues for Euclid Cluster science (managed by CG-WP8)
 - Paper 6.1: Velocity dispersions meta-catalogues
 - Paper 6.2: Consistency of masses in optical, X-ray, and SZ meta-catalogues





Clusters: Consortium Interfaces & Coordination

• The main SWG strength is coming from the strong collaboration/synergy/overlap with OULE3 Galaxy Clusters, where Italy is also having a strong leading role: a large fraction of the OULE3 processing functions necessary for cluster science are developed by Italian researchers who are at the same time very active members of SWG ⇒ most of them are leading the pre-launch KP papers planned inside OULE3 (KP-LE3-CL-1,2,3).

• Good interfaces with the rest of the Consortium:

SWGs: mainly with Strong Lensing, Galaxy clustering, Weak lensing, Cosmological simulations and Galaxy evolution;

IST: -Likelihood and -Nonlinear









Clusters: Status of Activities

- **Completion of the Pre-launch KPs:** 4 papers completed, **all with Italian leads** (Fumagalli, Tiago, again Fumagalli, Giocoli); overleaf existing for other 4-5 papers; preliminary results for the majority of the rest; few cases of delays due to missing person-power
- The largest effort is to apply the whole Euclid cluster cosmological pipeline (cluster detection, abundance and clustering measurements, mass-richness calibration, likelihood, final inference of the cosmological parameters, ...) to the best existing public data: **Hyper Suprime Cam (HSC)**
- Definition of Q1 and DR1 projects
- **Detailed plans for specific external data:** eRosita, SPT, other surveys.... (necessity of new MoUs? Public data only? Observational proposals?)









Theory: Organization

- 117 Members (Jan. 20): approx. 26% from Italy, 18% France, 10% UK, 9% both Portugal and Switzerland, 6% Germany and Spain, plus few more people distributed in Netherlands, Finland, Japan, Norway, Sweden, US...
- 10 Work Packages, 7 of them with an Italian co-lead
- 3 pre-launch KPs (see below)
- Communications:
 - Monthly calls, general one alternate with WP-leads one, plus biweekly calls for each WP.
 - Wiki: <u>https://gitlab.euclid-sgs.uk/SWG/SWG-TWG</u>
 - Mailing list: swg-cosmo-theory-group@euclid-projects.org
- Meetings:
 - One per year, plus possible Splinter at Consortium
 - Last one (first hybrid after the pandemy), in Bologna, October 10-11, 2022, next one TBD









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KP-TH-1: Forecasts for beyond-standard models in cosmology and fundamental physics

- **KP-TH-1 P1** Euclid preparation: xx. Forecast constraints on dark energy and modified gravity (IT-WP co-led)
- SP-TH-[WP1/2]-P1 Euclid: Constraints on f(R) cosmologies from the spectroscopic and photometric primary probes (w/ ECEB)
- SP-TH-[WP1/2]-P2 Euclid: Constraining linearly scale-independent modifications of gravity with the spectroscopic and photometric primary probes (IT-WP co-led) (w/ ECEB)
- SP-TH-[WP1/2]-P3 Euclid: Constraints on model independent parametrizations of gravity (IT-WP led)
- KP-TH-1 P2 Euclid preparation: xx. Forecast constraints on neutrinos in LCDM and beyond (IT co-led)
- SP-TH-[WP3]-P1 Euclid: Particle Dark Matter
- SP-TH-[WP3]-P2 Euclid: Generalized Dark Matter
- SP-TH-[WP3]-P3 Euclid: Primordial Black Holes
- SP-TH-[WP3]-P4 Euclid: Validation of MontePython
- KP-TH-1 P4 Euclid preparation: xx. Forecast constraints on initial conditions and implications for inflation (IT-WP led)
- SP-TH-[WP4]-P1 Euclid: The search for primordial features (IT-WP led)
- SP-TH-[WP4]-P2 Euclid forecasts for early Universe: Bayesian inference of cosmic initial conditions
- **KP-TH-1 P6** Euclid preparation: xx. Impact of non-linear clustering on beyond-LCDM constraints with Euclid

• KP-TH-2: Relativistic effects (IT co-led)

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- **KP-TH-2 P1** The impact of wide-angle, lensing, and the 'finger of the observer' effects on two-point statistics for the Euclid spectroscopic sample (IT-WP led) (published in MNRAS)
- KP-TH-2 P2 Euclid preparation XIX. Impact of magnification on photometric galaxy clustering (published in A & A)
- **KP-TH-2 P3** Euclid preparation: xx. Impact of lensing magnification on the spectroscopic sample of the *Euclid* mission
- SP-TH-[WP9]-P1 Relativistic effects beyond the ΛCDM model (IT-WP led)
- SP-TH-[WP9]-P2 Relativistic effects: the dipole of the correlation function (IT-WP led)
- KP-JC-6: Simulations and non-linearities beyond LCDM (IT co-led)
 - KP-JC-6 P1 Simulations and non-linearities beyond LCDM
 - KP-JC-6 P2
 Scientific results results from non-standard cosmological simulations
 - KP-JC-6 P3 Cosmological constraints on non-standard cosmologies from simulated Euclid probes (IT-WP Led)
 - **KP-JC-6 SP1** Exact title TBD (MCMC constraints on extended cosmologies from Euclid photometric survey)





Theory: Consortium Interfaces & Coordination

• Interfaces and coordination:

SWGs: mainly with Galaxy clustering, Weak lensing, Cosmological simulations (see WP 12 and joint pre-launch project) and CMBX;

IST: -Likelihood and -Nonlinear









Theory: Status of Activities

- KP Progress:
 - A full day was dedicated to the status of KPPs and SPPs in the SWG meeting last October
 - Progress monitored by different KP and WP leads in WP dedicated and general telecons, also supervised by SWG leads

• Contribution to CLOE:

- Interface with IST:L and IST:NL settled
- New WP11 (Likelihood interface) co-lead appointed fully dedicated to CLOE implementations in the last fall
- An heritage of models investigated with Fisher within pre-launch KPPs to be implemented in CLOE with manageable adjustments due to the evolving specifications/theoretical modelling
- An internal release of CLOE v2.0 is needed to proceed
- A clear schedule and wise choice of contents of IST:L KPPs is needed to better motivate work in the direction of CLOE implementation of cosmology beyond LCDM within TWG
- Joint meetings and/or hands on meetings with IST:L would be highly desirable (was Heidelberg 2016 the only one organized?)
- DR1
 - The discussion of the science which could be extracted from DR1 has started in the last Theory meeting and is proceeding.









CMBX: Organization

- Members: 70% equally splitted between France, Italy, 15% in UK, 5% Germany, 3% Spain, 3% Switzerland , 3% the Netherlands, 1% US, Gender Balance at 20%
- Communications:
 - General bi-weekly zoom calls: 705 044 4609
 - WP dedicated calls: Likelihood, Simulations, ...
 - Wiki: <u>euclid.roe.ac.uk/projects/cxswg/wiki</u>,
 - Slack Channel: euclid-cmbx.slack.com
- Meetings:
 - Twice per year, Italy (fall), France (Spring)
 - Splinters at Consortia,
 - o next one, in Marseille, March 23rd, 24th, 2023









CMBX: KP Plan

KP-CMBX-1: Forecasts for CMBX Paper 1.1: Forecasts validated on simulations [Series to KP 2.2] 0 Paper 1.2: Fisher based forecasts for cross-correlated Euclid and CMB data [Published] Ο Paper 1.3: Extended forecasts for CMBX [Summer] **IT-WP Lead** Ο KP-CMBX-2: Simulations for CMB-LSS cross-correlations Paper 2.1: Lightcone comparison project [Published] Ο Paper 2.2: CMBX Mock Simulations [Spring] IT-WP Lead Ο KP-CMBX-3: ISW-GC Cross-Correlation Paper 3.1: Estimators of the ISW-GC cross-correlation [Spring] IT-WP Lead Ο KP-CMBX-4: tSZ/CMBL-WL Cross-Correlation Paper 4.1: Estimators for tSZ-WL cross-correlation [Series to KP 4.2, Link to CWG] Ο Paper 4.2: Estimators for CMBL-WL cross-correlations [Fall] **IT-WP Lead** Ο KP-CMBX-5: tSZ/CMBL-GC Cross-Correlation Paper 5.1: Estimators for tSZ-GC cross-correlations [Series to KP 5.2, Link to CWG] Ο Paper 5.2: Estimators for CMBL-GC cross-correlations [Fall] **IT-WP Lead** Ο **KP-CMBX-6:** Estimators for the reconstruction of kSZ signal Paper 6.1: Methods for reconstructing the kSZ effect of galaxy clusters from Euclid-CMB cross-correlations [Series to KPs 4.1, 5.1]









CMBX: External Datasets







0.0 FDS dust emist

0.10 mK RJ



Activities and National WPs

Activities

- Aosta: CMB-N-Body & Ray Tracing, Simulations and Analysis
- Trieste: Analysis of CMB-N-Body Simulations
- Milano: Production of CMB-N-Body & Ray Tracing Simulations, Analysis
- Padova: Implementation of Estimators, Analysis, Theory
- Ferrara: Implementation of Estimators & Likelihood
- Bologna: Production of CMB-N-Body Simulations, Analysis, Estimators & Likelihood, Theory
- Rome: Productions of Estimators & Likelihood, Simulations, Analysis



National WPs:

- Trieste [ASI-INAF, Baccigalupi]: Simulations
- Milano [INFN, Guzzo]: Simulations, Estimators & Likelihood
- Padova [INFN, Stanco]: Estimators & Likelihood
- Ferrara [ASI-INAF-INFN, Lattanzi, Natoli], Estimators and Likelihood
- Bologna [ASI-INAF, INFN, Finelli, Sirri]: Estimators & Likelihood, Simulations, Theory
- Support for Computing Facilities (INFN)









CMBX: Status of Activities

- KP Schedule:
 - **Drafting 1.3, 2.2, 3.1,**
 - Starting 4.2, 5.2
- Increasing Integration:
 - e2e Simulation Pipeline Finalizing,
 - Estimators are Operational (ISW-G), Implemented (CMBL-Main Probes)
 - Likelihood Modules under Implementation
- Robust Interfaces through Group Members:
 - Simulation WG,
 - o IST-L,
 - IST-NL,
 - Clusters,
 - Theory,
- ..

- Near Term Goals:
 - Delivering KPs,
 - Completing Sets of Estimators & Validation on e2e Simulations,
 - Interface with SGS Objects,

• Interface with CMB Probes:

- Science Oriented, Performance Assessments in Papers, see KP1.2
- Using Public Objects, see KP1.2,
- Cross-Membership with CMB-S4, LiteBIRD, LSPE, SO,
- Consortium Level
- Meetings:
 - Orsay, March 23rd, 24th
 - Splinter at Consortium Meeting
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Pre-Launch Considerations

- Completion of Pre-Launch KP Plan
 - Capitalizing Significant Leadership of the Italian Community
 - Charting Inter-WG KPs
 - Discussing Post-Launch KPs,
 - Transition to Analysis Tools, Support to OUs, SGS,
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- Coordination
 - Success of the Consortium through the Italian Community
 - Areas of Improvement: OUs, SGS, Possibility for Discussing Stable Links
 - Plan for Hand-On Joint Meetings?
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- Common Interfaces & Needs: Numerical Simulations
 - Simulations are Supported through CINECA
 - Study Plan for post-Launch Support for Calculus
 - Catalogue of non-Standard Cosmological Simulations: Marco Baldi's <u>non-Standard Cosmologies Simulation</u> <u>Chart</u>,
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