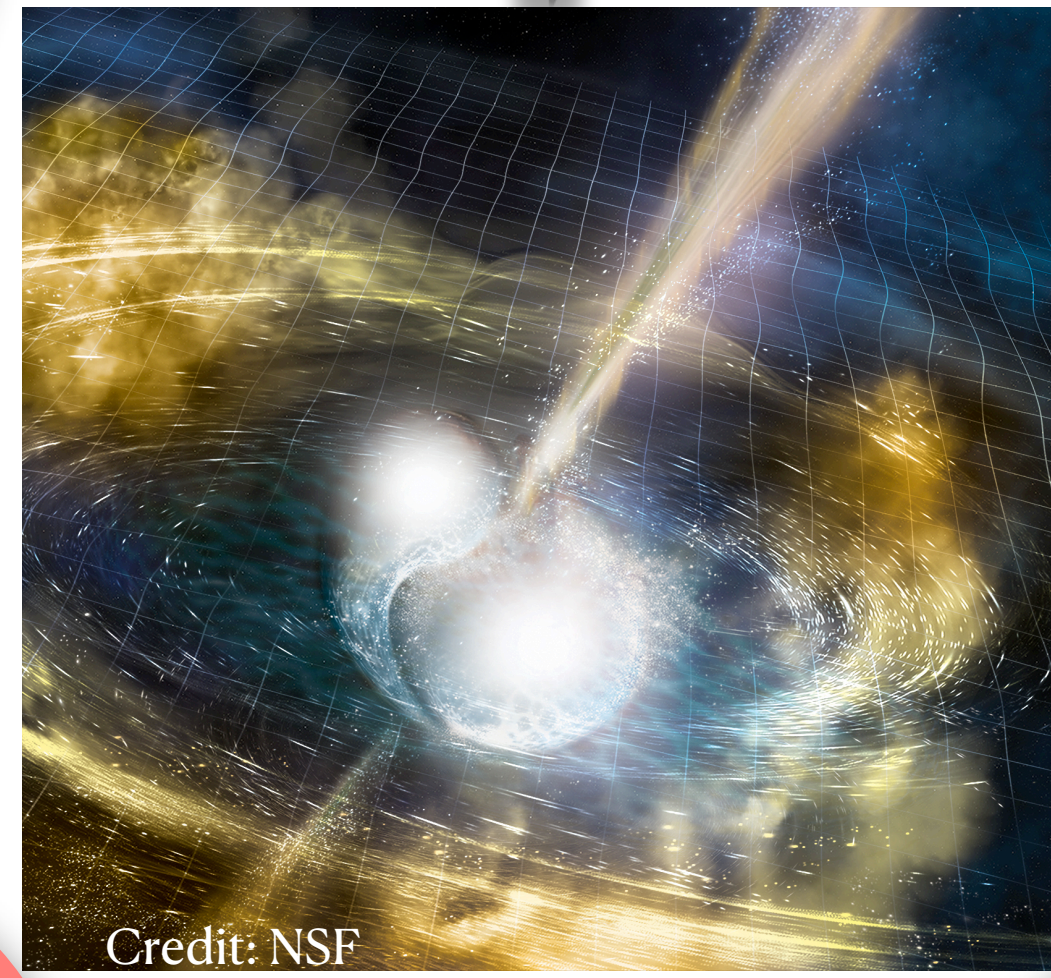


The trans-disciplinary character of GRBs

Stellar & Galactic
astronomy

Early Universe



Gravitational
waves

Relativistic
Astrophysics

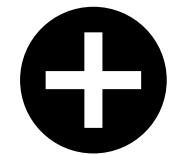
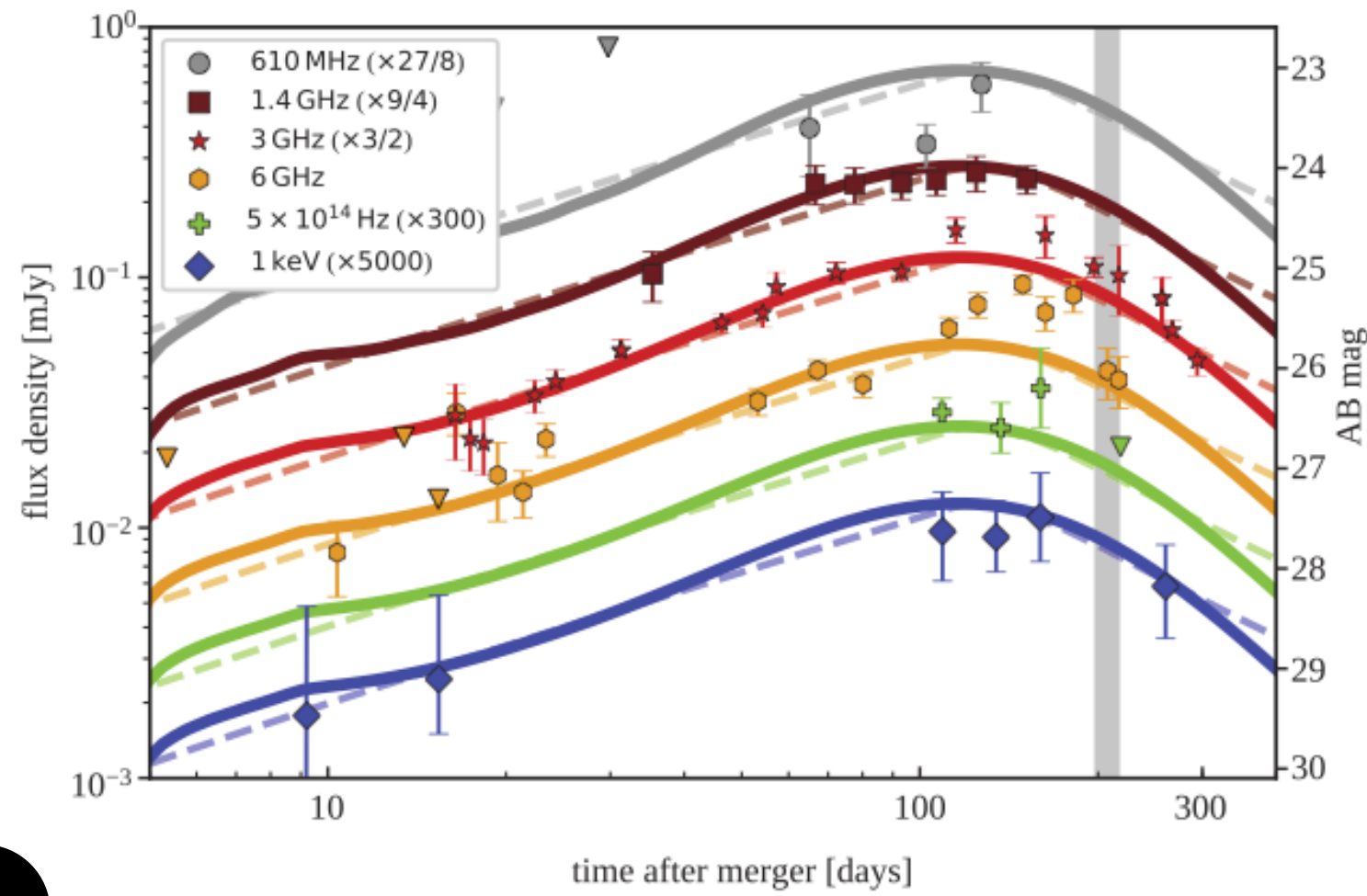
UHECR & ν

GRBs are bursts of knowledge. When nothing happens ... a GRB pops up ...

From GRB 170817 ... to structured jets

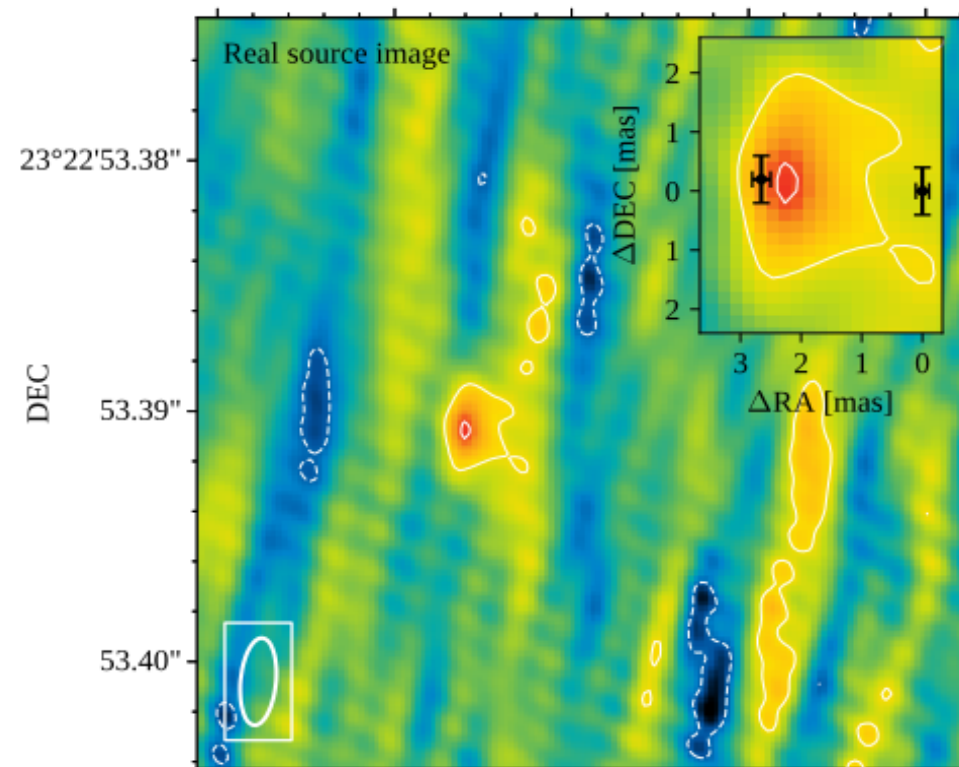
(for MM see Paolo's):

D'Avanzo, P.; Campana, S.; Salafia, O. S.; Ghirlanda, G., et al. 2018, A&A

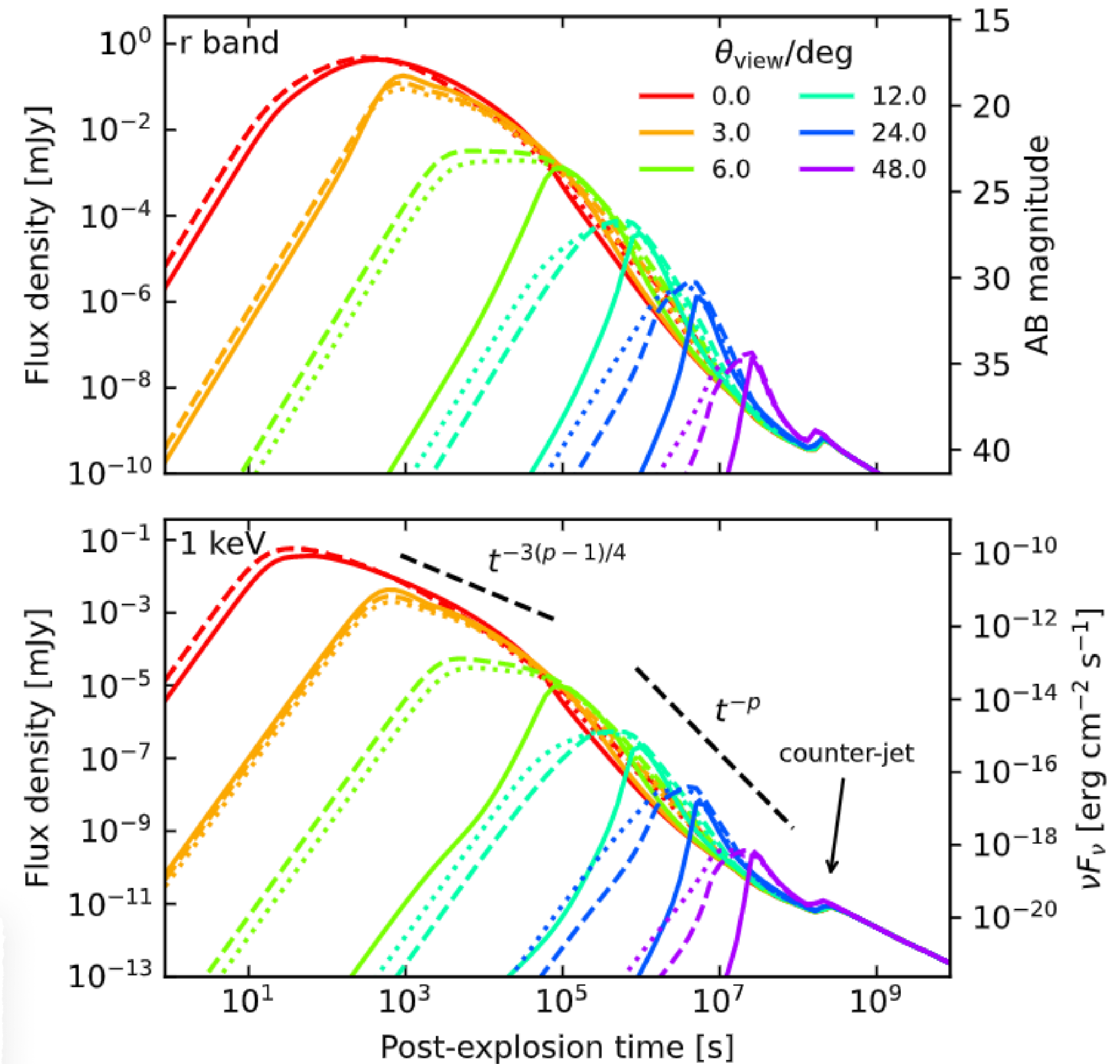


Structured jet in GRBs

Ghirlanda G., Salafia O. S., Paragi Z., Giroletti M., et al. 2019, Sci



What is the structure origin?
 Which observational features?
 Successful vs failed jets
 Jets, breakouts, cocoon

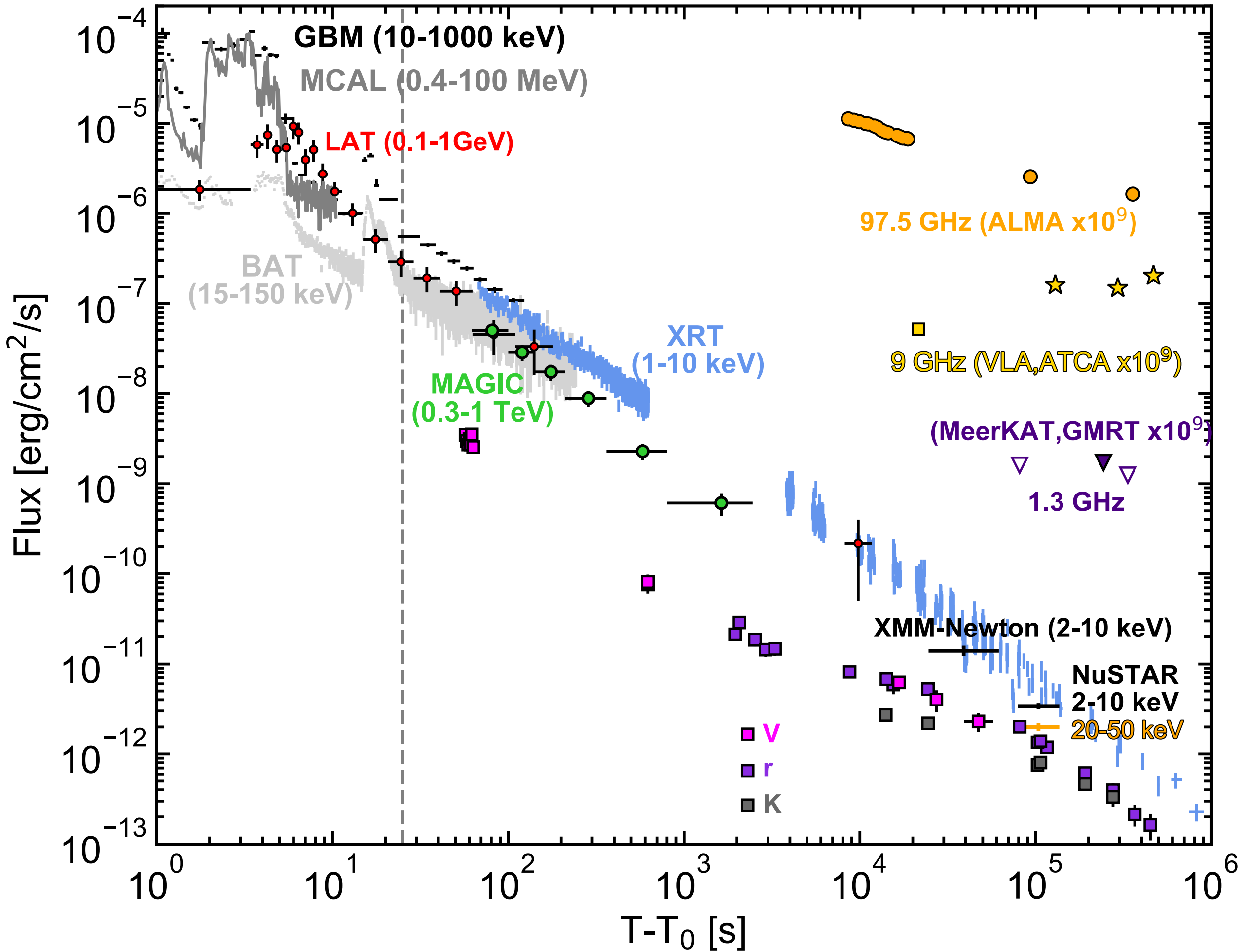


Review on structured jets in GRBs:
Salafia & Ghirlanda 2022

TeV emission of GRBs

MAGIC collab., 2019, Nat.

GRB 190114C



Name	z	T ₉₀ s	E _{γ,iso} erg	L _{X,11h} erg s ⁻¹	Energy TeV	Time s	IACT
180720B	0.653	49	^b 6 × 10 ⁵³	3 × 10 ⁴⁶	0.1–0.44	(3.6–4.3) × 10 ⁴	H.E.S.S.
190114C	0.424	25	^c 2.5 × 10 ⁵³	10 ⁴⁶	0.2–1	62–2454	MAGIC
190829A	0.078	^b 63	^a 2 × 10 ⁵⁰	4 × 10 ⁴⁴	0.18–3.3	(1.5–20) × 10 ⁴	H.E.S.S.
201216C	1.1	^b 30	^a 5 × 10 ⁵³	10 ⁴⁶	0.1	56–8 × 10 ³	MAGIC
201015A	0.42	10	10 ⁵⁰	2 × 10 ⁴⁴	0.14	33–1.4 × 10 ⁴	MAGIC

Reviews: *Nava L., 2021; Miceli & Nava 2022*

- What is the nature of the TeV emission?
- What the constraints on the GRB physics?
- Constraints on the EBL
- New physics?

Christmas gift: a long GRB with KN

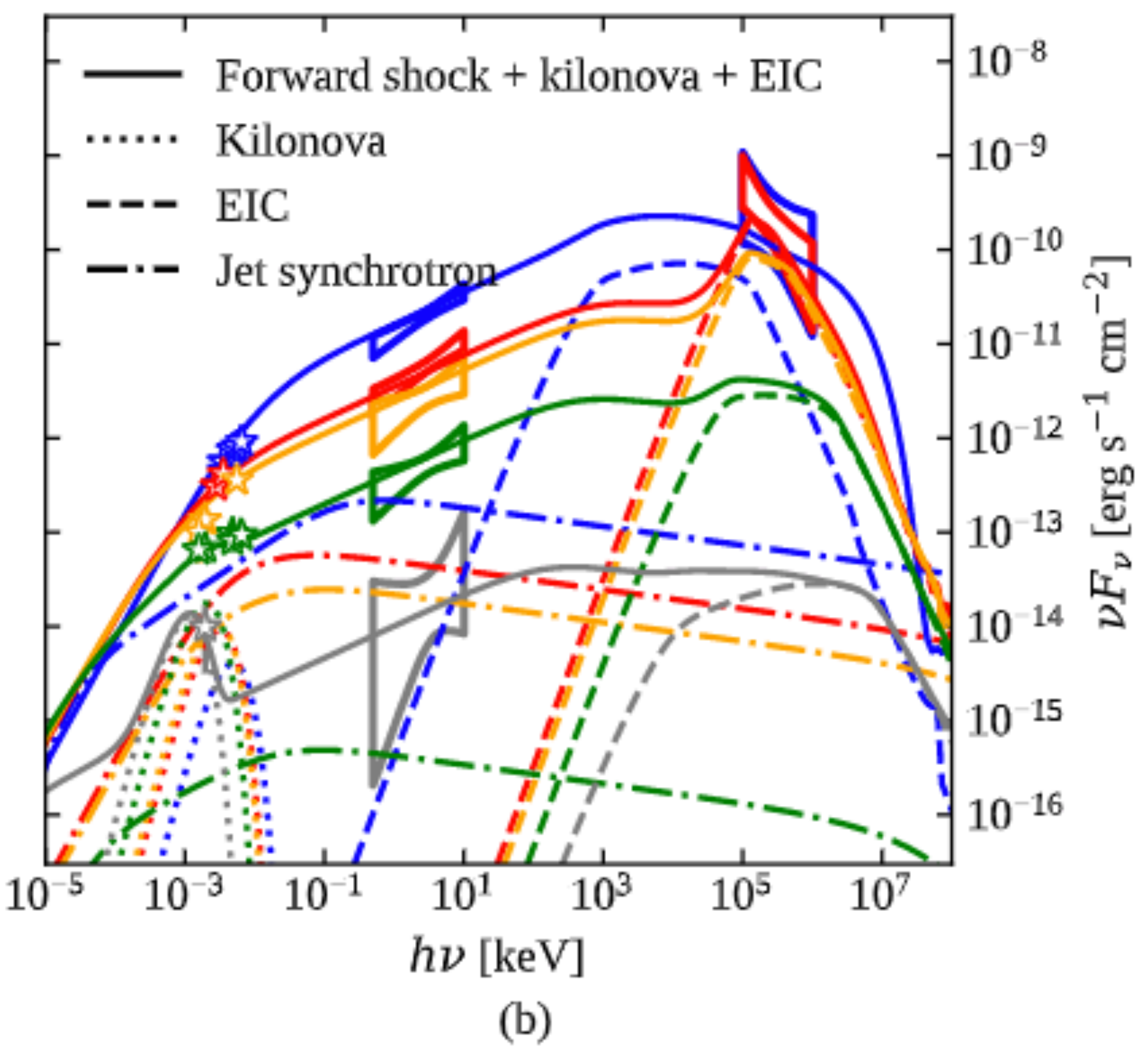
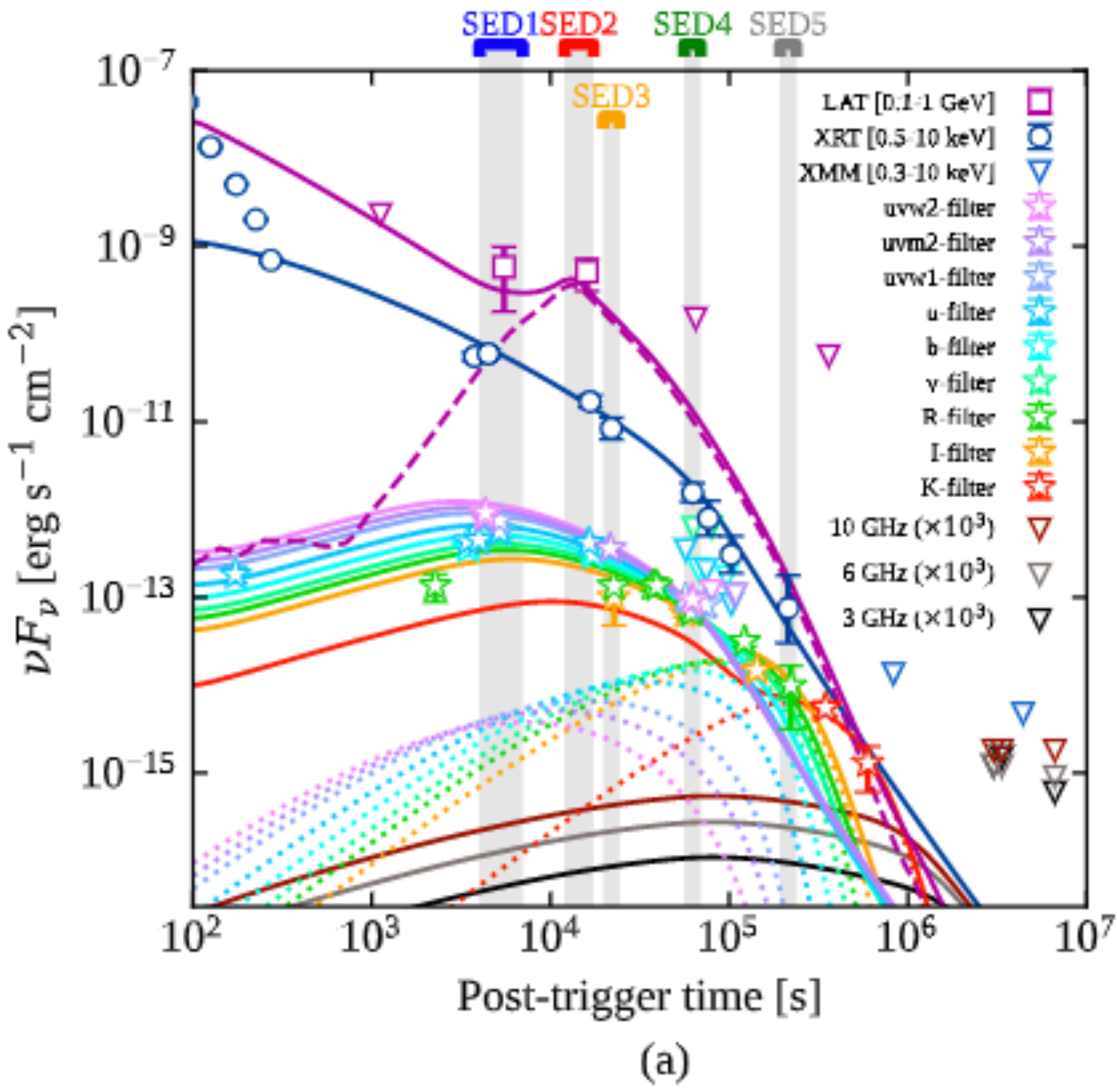
13 Papers + 30 GCN circulars



1. Kilonova signature (Rastinejad et al., 2022)
2. (another hint) for synchrotron prompt emission (Gompertz, **Ravasio**, et al. 2022)
3. Precursor periodicity (Xiao et al., 2022)
4. Late Fermi/LAT (GeV) detection (**Mei et al., 2022**)



A. Mei, B. Banerjee, G. Oganessian, O. S. Salafia, S. Giarratana, M. Branchesi, P. D'Avanzo, S. Campana, G. Ghirlanda, S. Ronchini, A. Shukla, P. Tiwari, 2022, Nat. submitted

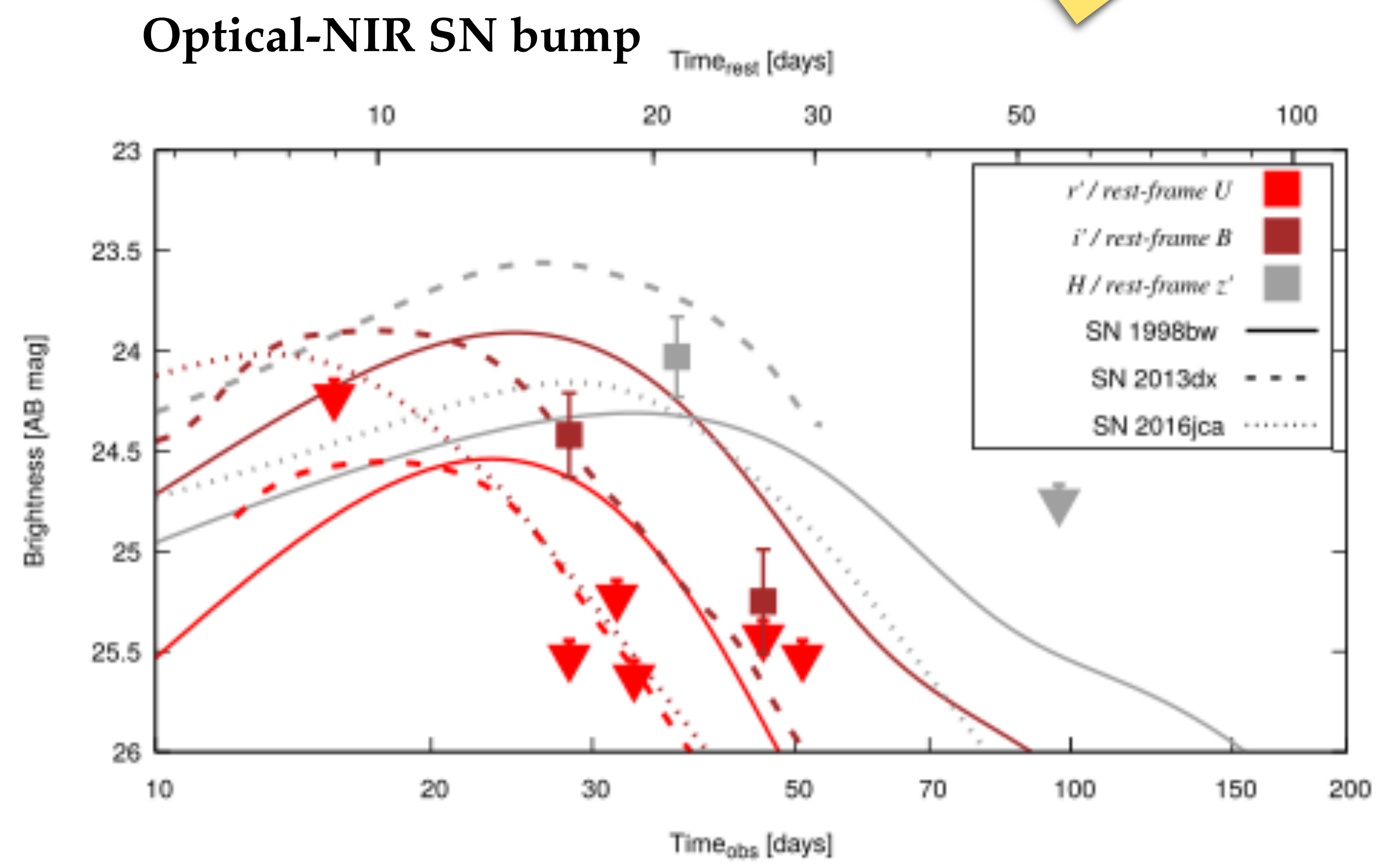
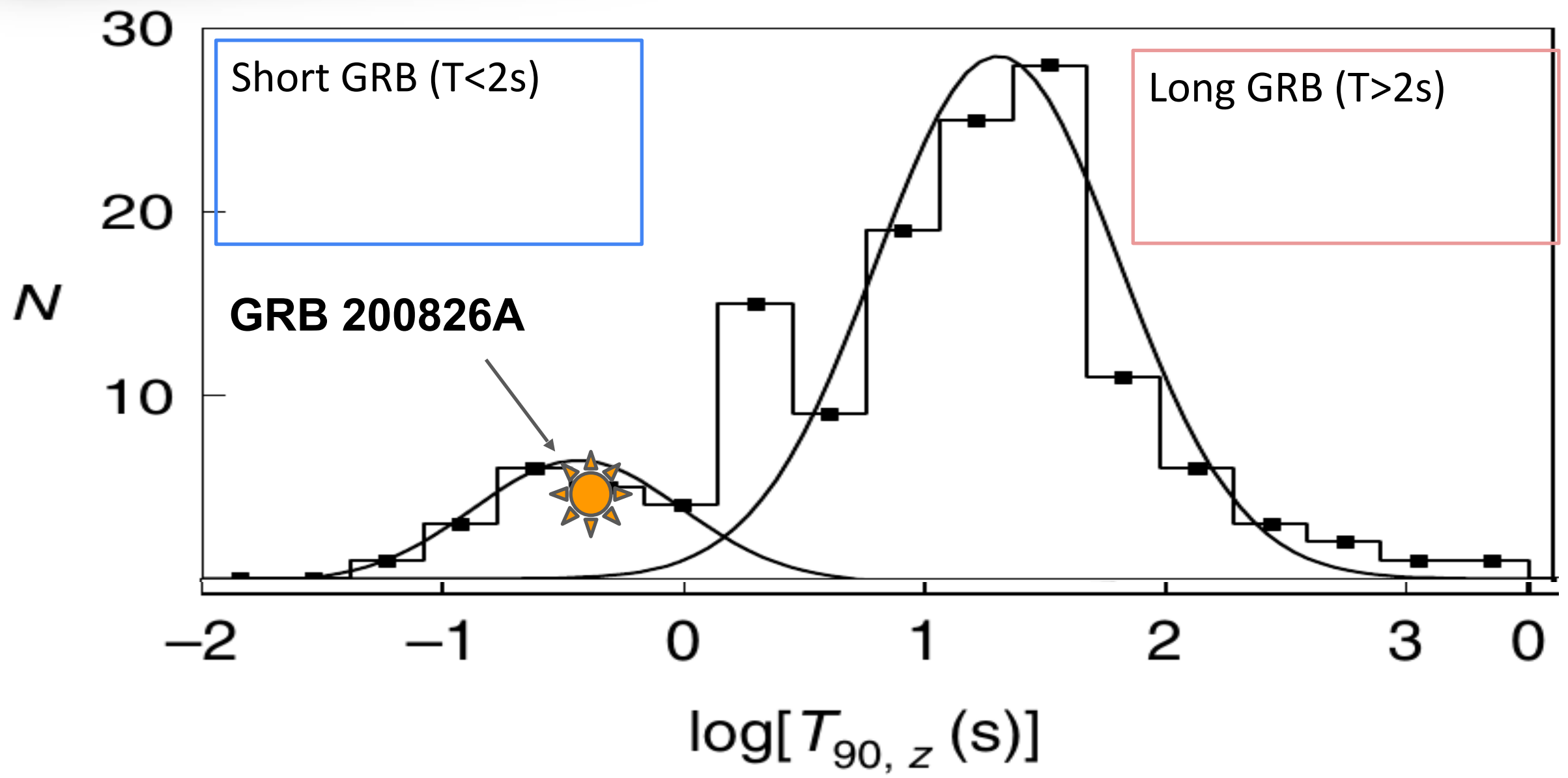


How the CE works?
Short and long
Extra components?

Summer gift: the shortest collapsar

Rossi, B. Rothberg, E. Palazzi, D. A. Kann, P. D'Avanzo, L. Amati, S. Klose, A. Perego, E. Pian, C. Guidorzi, A. S. Pozanenko, S. Savaglio, G. Stratta, G. Agapito, S. Covino, F. Cusano, V. D'Elia, M. De Pasquale, M. Della Valle, O. Kuhn, L. Izzo, E. Loffredo, N. Masetti, A. Melandri, P. Y. Minaev, A. Nicuesa Guelbenzu, D. Paris, S. Paiano, C. Plantet, F. Rossi, R. Salvaterra, 2022, ApJ, 932:1

GRB 200826A



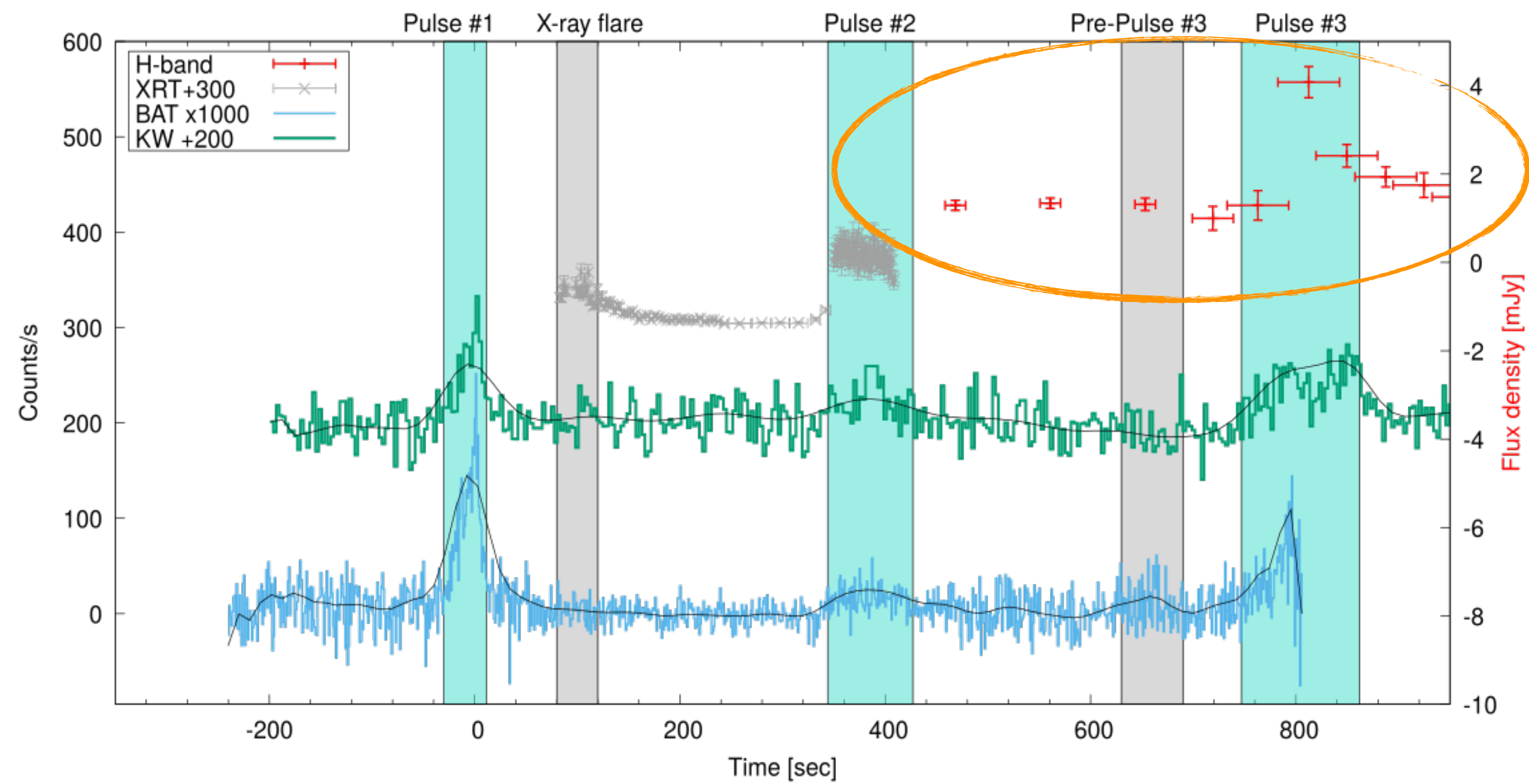
Revisiting short/long divide
Progenitor nature

Monster blasts from the past

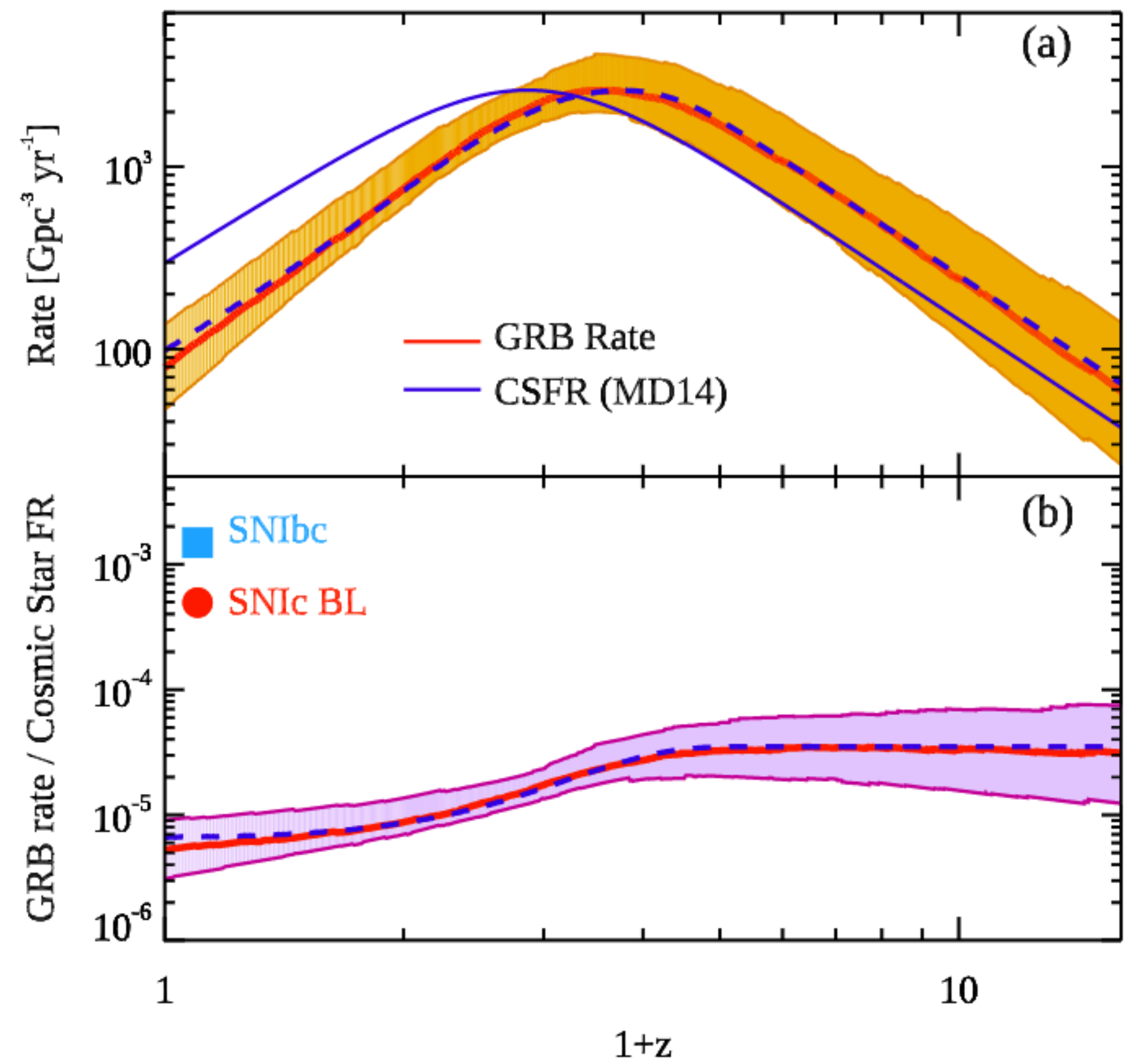
A. Rossi, D. D. Frederiks, D. A. Kann, M. De Pasquale, E. Pian, G. Lamb, P. D'Avanzo, L. Izzo, A. J. Levan, D. B. Malesani, A. Melandri⁶, A. Nicuesa Guelbenzu, S. Schulze, R. Strausbaugh, N. R. Tanvir, L. Amati, S. Campana, A. Cucchiara, G. Ghirlanda, M. Della Valle, S. Klose, R. Salvaterra, R. L. C. Starling, G. Stratta, A. E. Tsvetkova, S. D. Vergani, A. D'Ai, D. Burgarella, S. Covino, V. D'Elia, A. de Ugarte Postigo, H. Faushey, J. P. U. Fynbo, F. Frontera, C. Guidorzi, K. E. Heintz, N. Masetti, E. Maiorano, C. G. Mundell, S. R. Oates, M. J. Paegle, M. Pazzi, J. Palmerio, G. Pugliese^{34,35}, A. Rau³⁶, A. Saccardi¹⁹ et al.
 2022, A&A, 665, 125

GRB 210905

@ z=6.3



Ghirlanda, G. & Salvaterra, R., 2022, ApJ, 932, 10

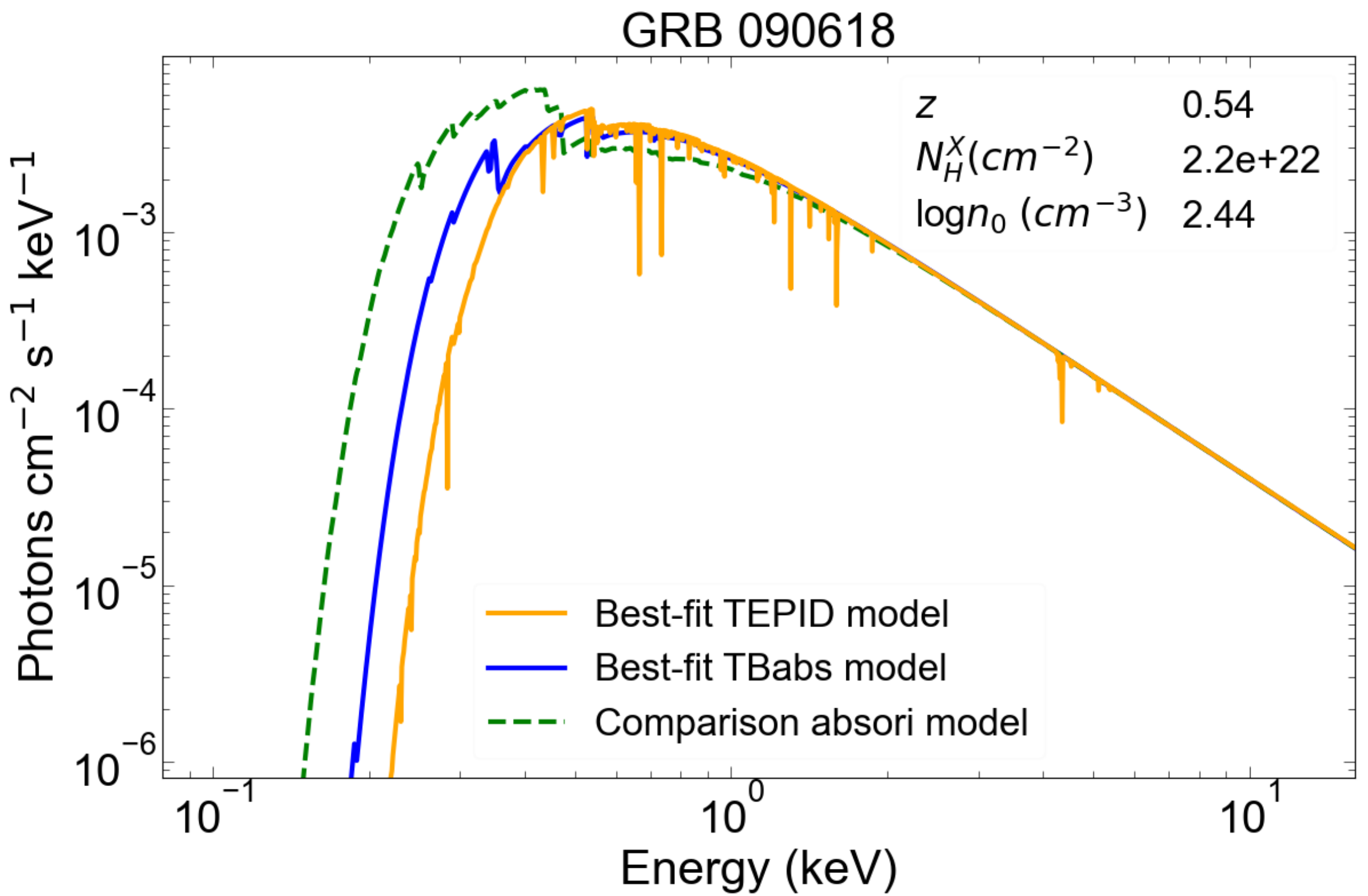


Population studies and simulations supporting mission concept design and science case studies [Theseus, HERMES, CTA ...]

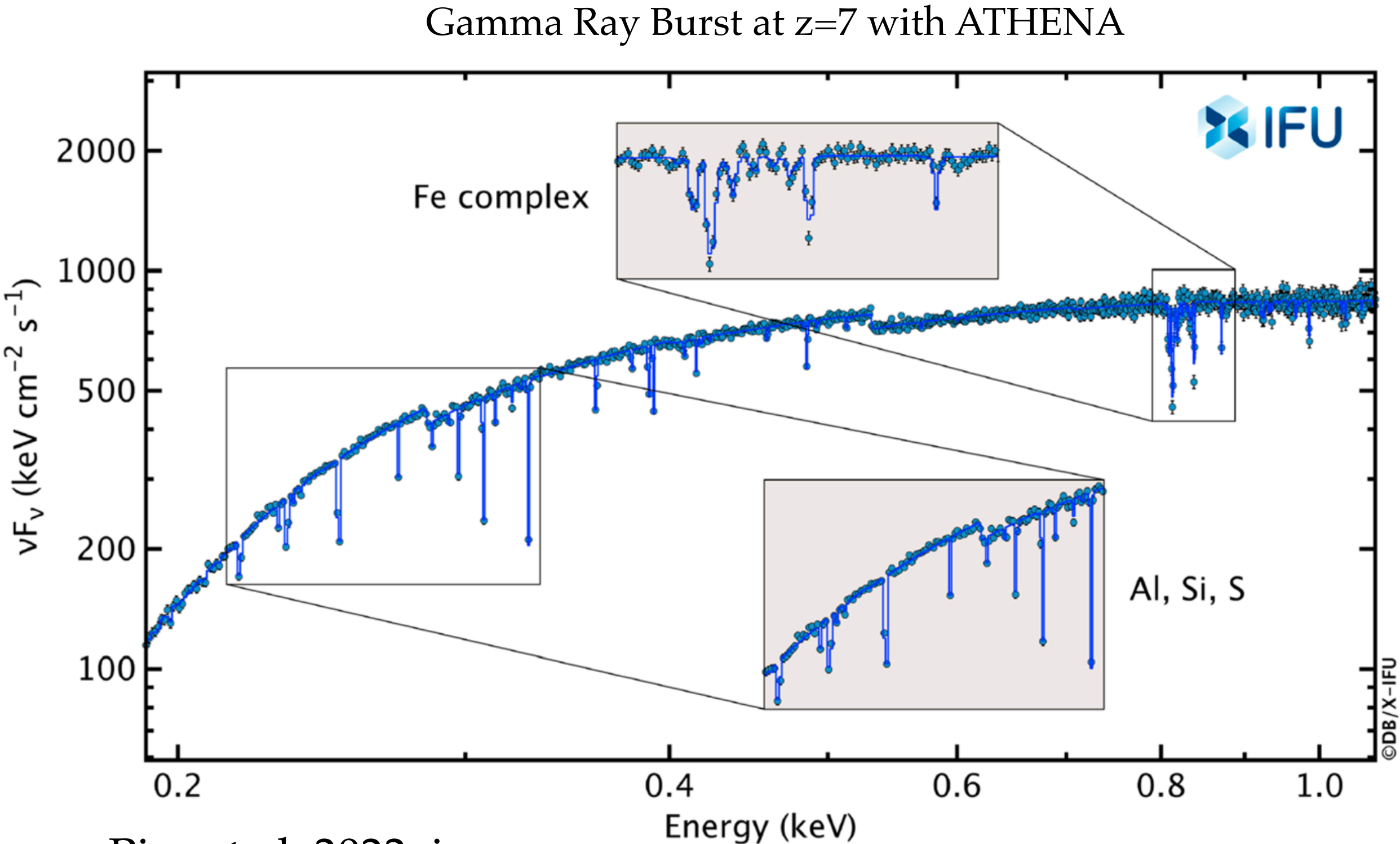
Environment matters

Highly variable and large GRB flux into the circumburst environs requires proper dependant modelling (TEPID: Luminari et al 2022)

Applied to a golden sample of XMM_Newton Obs (Thakur et al 2022, in prep.)



When did the first generation of stars explode to form the first seed black holes and disseminate the first metals in the Universe?

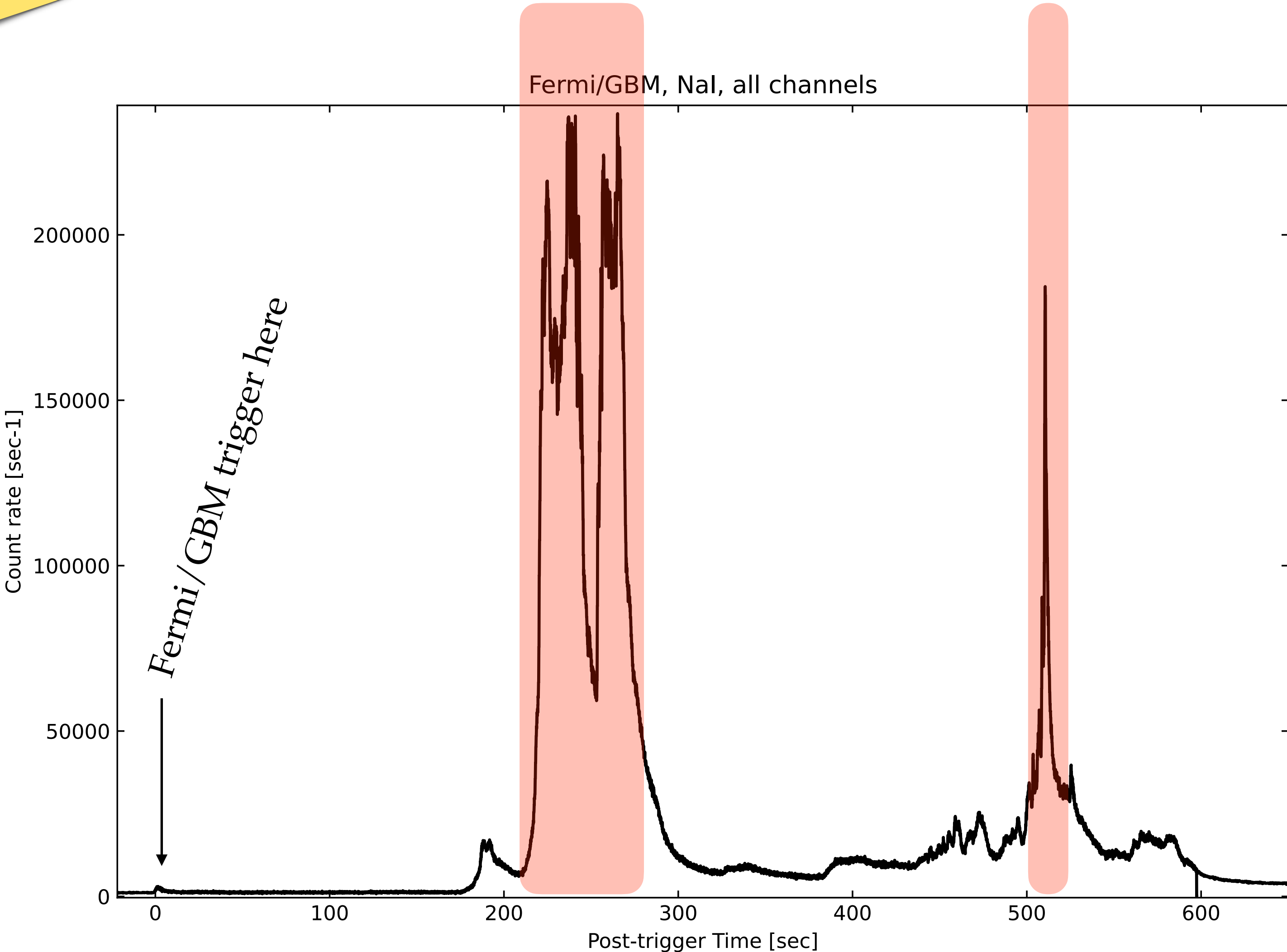


Piro et al. 2022; in prep.

A monster in the backyard

GRB 221009A

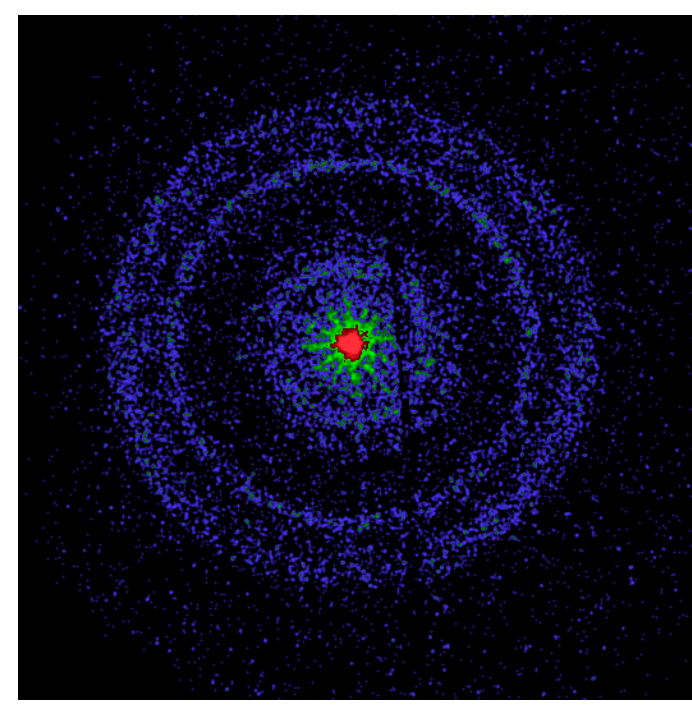
Data issues



Long, energetic, luminous
 $L > 6 \times 10^{52} \text{ erg s}^{-1}$ $E \gtrsim 3 \times 10^{54} \text{ erg}$

VHE detection (up to 18 TeV ...)

ALP interpretation (Galanti et al. 2022)

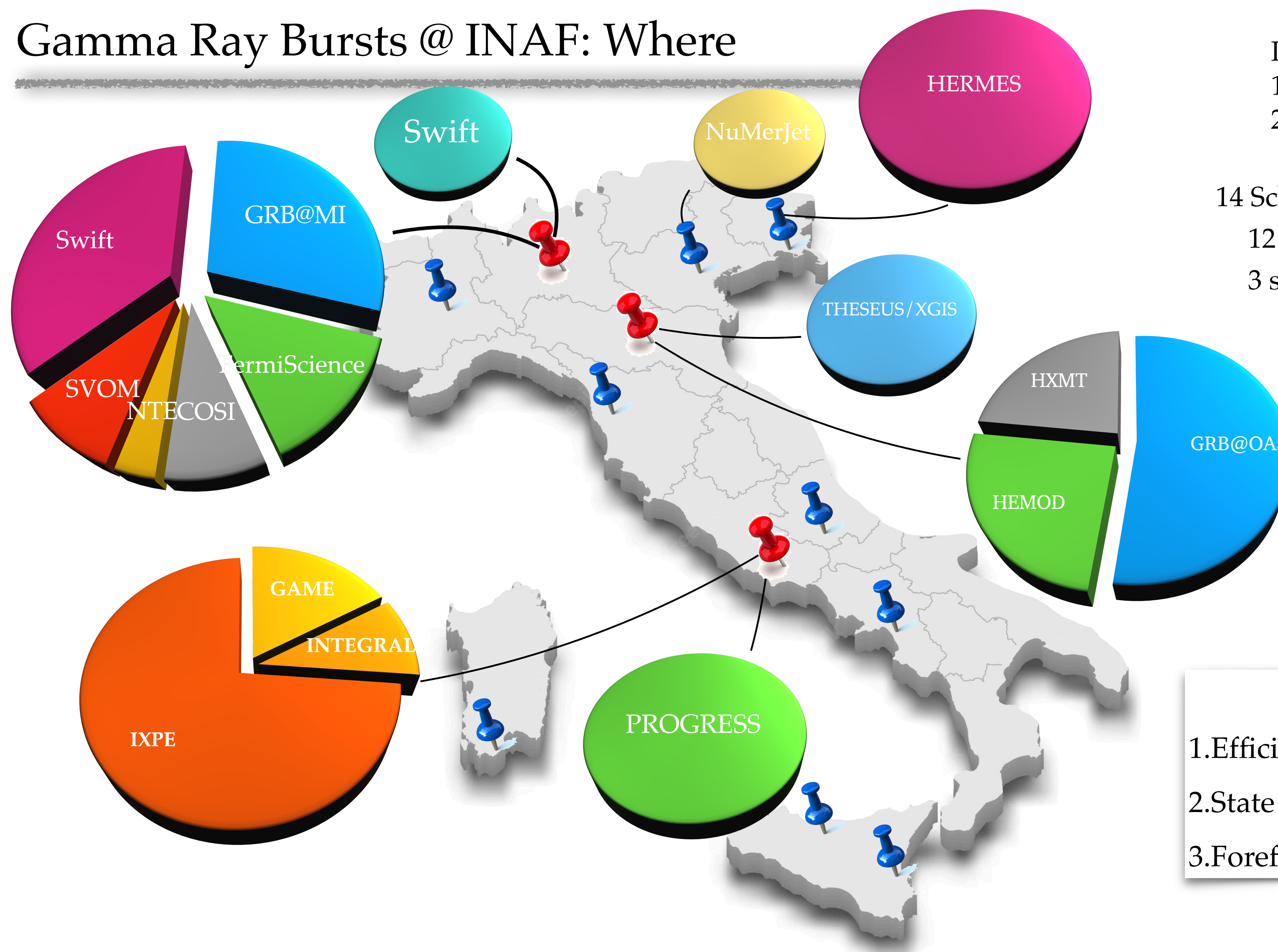


Dust rings
(Tiengo et al. 2022)



Much more to come

Gamma Ray Bursts @ INAF: Where



Preliminary remarks:

- 1) GRB only (MM → Paolo)
- 2) Schede RSN4 primaria

14 Schede in RSN4 primaria + 1 in RSN5

12 schede legate a satelliti / strumenti

3 schede su GRB (>50% non GW):

1. GRB@MI
2. GRB@OAS
3. GAME

GRBs @ INAF

1. Efficient top level Observations
2. State of the art modelling and theory
3. Forefront in tecno devel.

Gamma Ray Bursts @ INAF: What

Preliminary remarks:

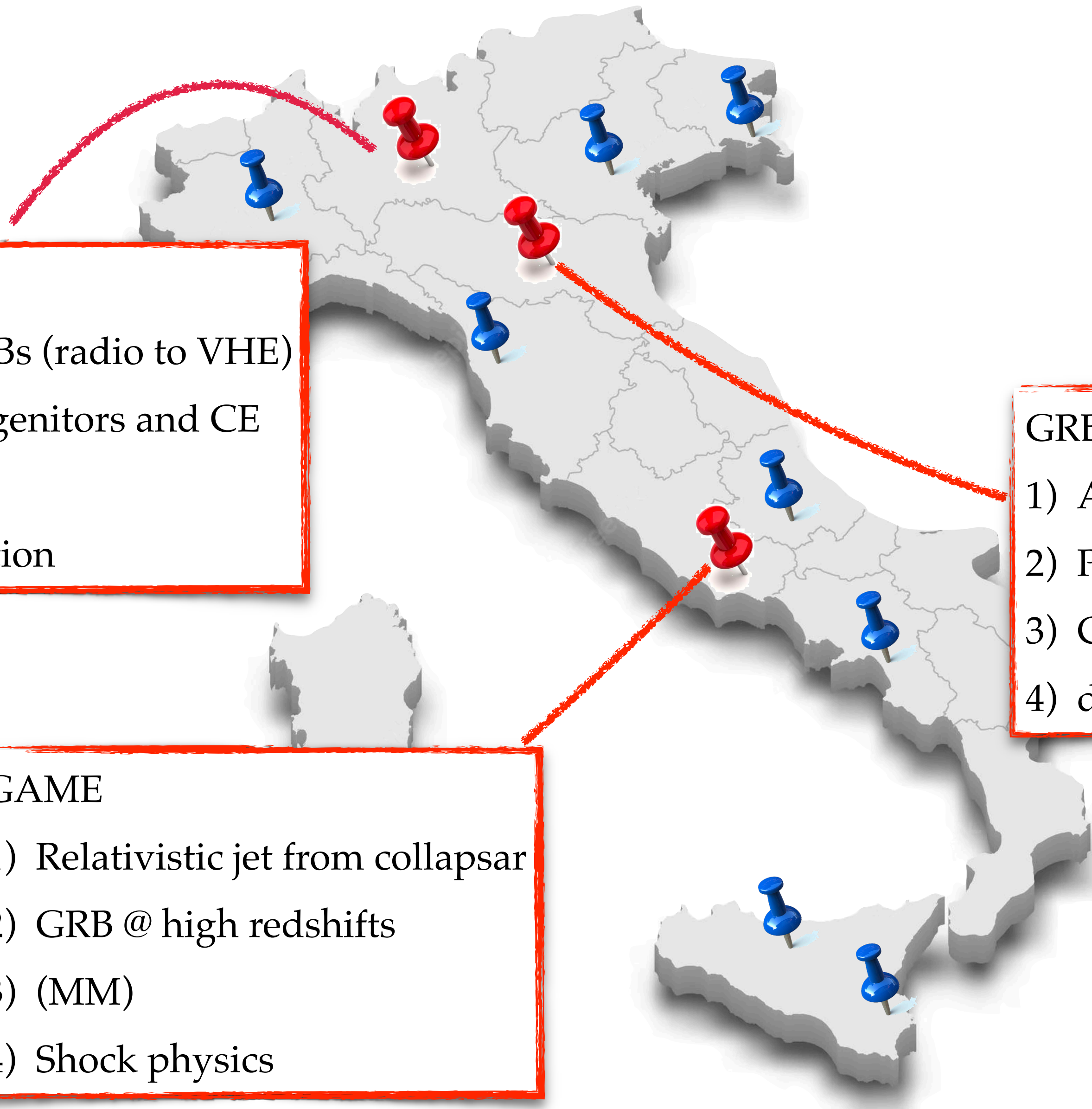
- 1) GRB only (MM → Paolo)
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12 schede legate a satelliti / strumenti

3 schede su GRB (>50% non GW):

- 1. GRB@MI**
- 2. GRB@OAS**
- 3. GAME**



GRB@MI

- 1) Physics of GRBs (radio to VHE)
- 2) Nature of progenitors and CE
- 3) (MM)
- 4) Cosmic evolution

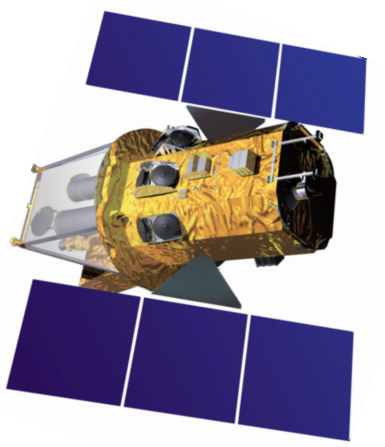
GRB@OAS

- 1) Afterglow observ / interpret.
- 2) Progenitor nature
- 3) Close / large environ.
- 4) detection / new instrument.

GAME

- 1) Relativistic jet from collapsar
- 2) GRB @ high redshifts
- 3) (MM)
- 4) Shock physics

National and international collaboration



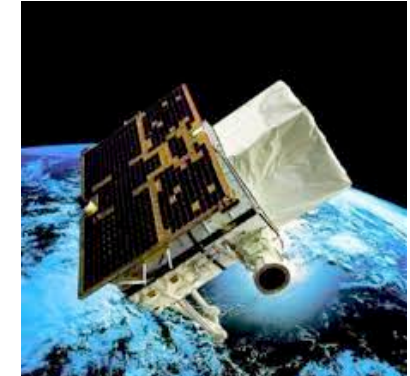
Italian *Swift* team

~ 20 people (INAF: OAB, IASF-Mi, OAR, IASF-Pa; ASI: SSDC)
 PI: Gianpiero Tagliaferri (OAB)

18 YEARS:
 ~1600 GRBs (90/yr)
 80% XRT detection
 30% UVOT detection
 50% optical/NIR detection
 7% radio detection
 26% with redshift



Integral



AGILE



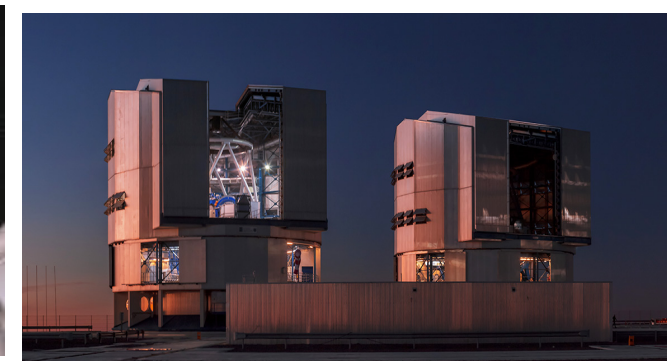
Fermi



Optical/NIR ground-based GRB follow-up

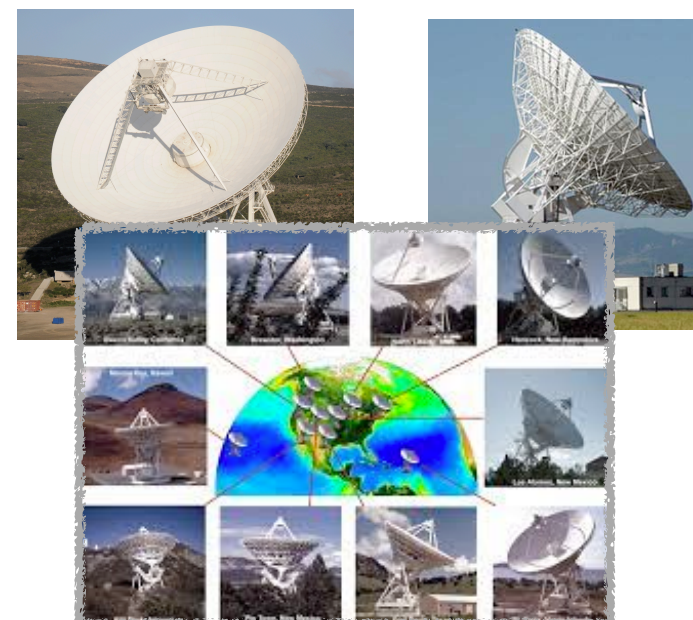
Coordinamento Italiano Burst Ottici: REM-TNG-LBT

~ 30 people, Italian collaboration INAF-led



STARGATE: ESO-VLT

~ 100 people, European collaboration
 significant participation of INAF researchers



Radio follow up (SRT, Medicina, EVN, EAVN, VLBA, e-Merlin, VLA)

~10-15 participants, international team - INAF led collaboration

Instruments & Facilities

Future

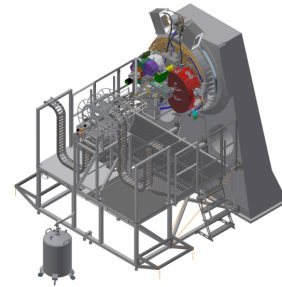
SKA



ELT



JWST



SoXS



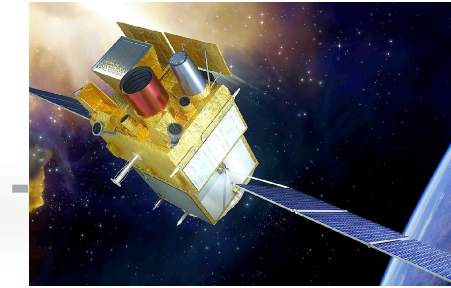
V. Rubin

Hermes

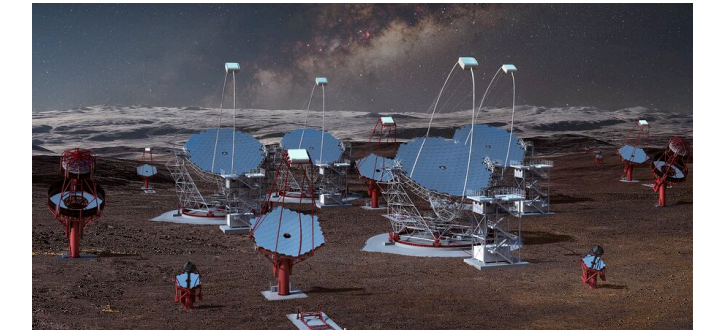


Hermes

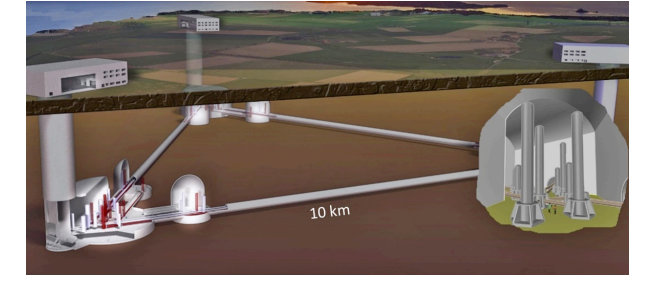
SVOM



CTA



Einstein Telescope



Medicina & Noto

SRT



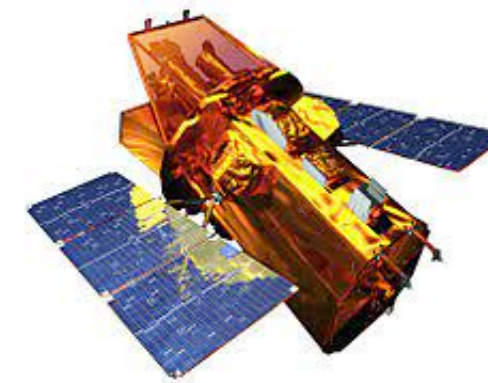
LBT



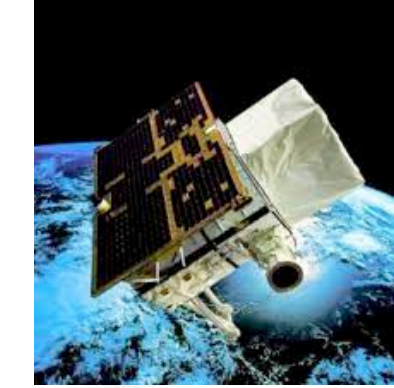
TNG



Niel Gehrels
Swift Obs.



Agile

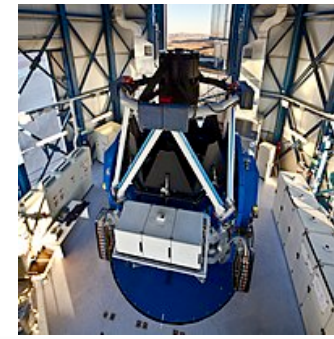


ASTRI

REM



VST



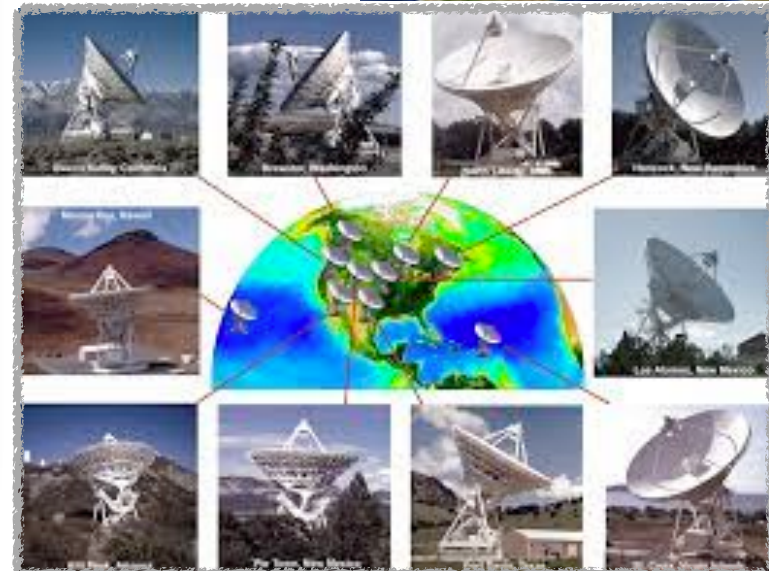
MAGIC



EVN,
g-VLBI,
eMerlin



VLBA



VLT



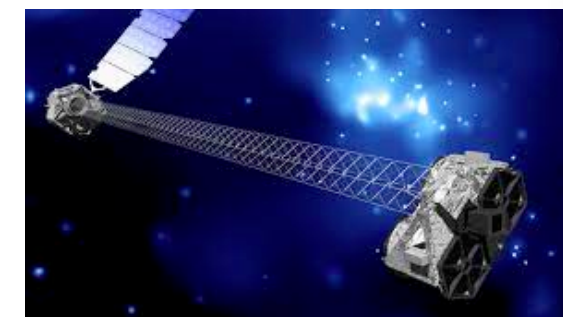
NTT



Chandra



Nustar



Integral



Fermi



HST



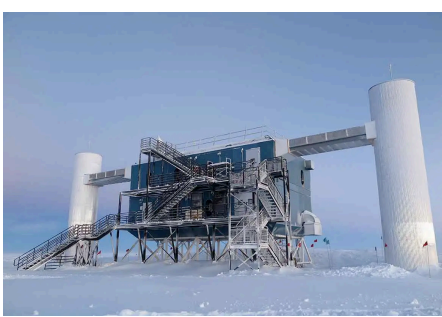
XMM-Newton



LIGO & Virgo



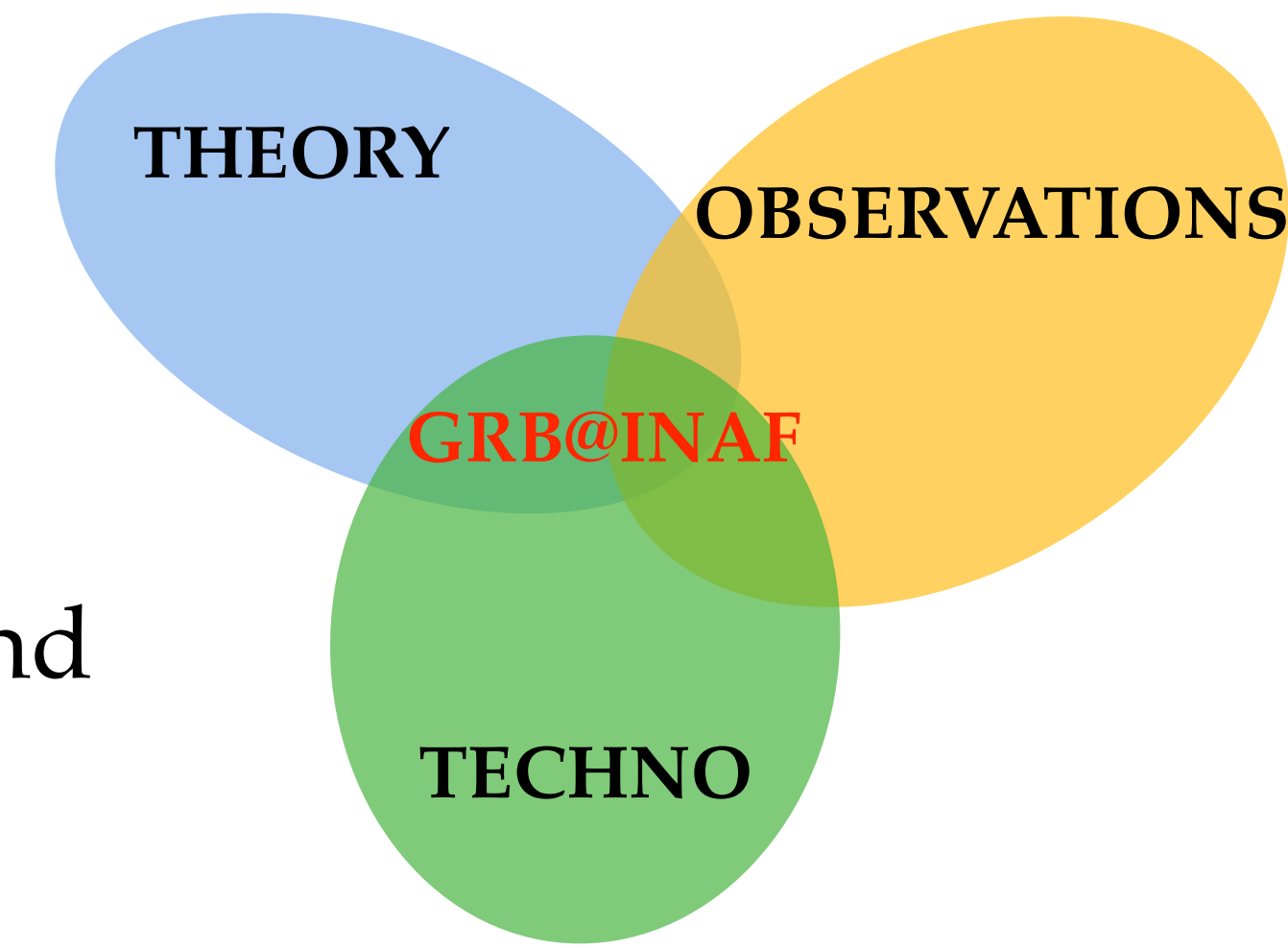
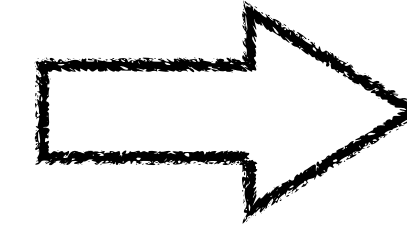
Ice-Cube



Strengths, Weaknesses and Needs



- The success of GRB science in INAF



- As a community we have evolved into large (national and international) collaborations.

- INAF is at the forefront in this research field

- Perspectives? What will come after (and better than) Swift and Fermi? How INAF wrt ESA, NASA, etc.?

- Programmatic views (earlier career support) & Alta formazione

- INAF follow up facilities: present (REM, LBT, TNG, Magic, SRT ...) and future (SOXS, CTA and new ideas - Hugo)

- Funding plans (ASI, INAF, MUR, EU)



Thank you