

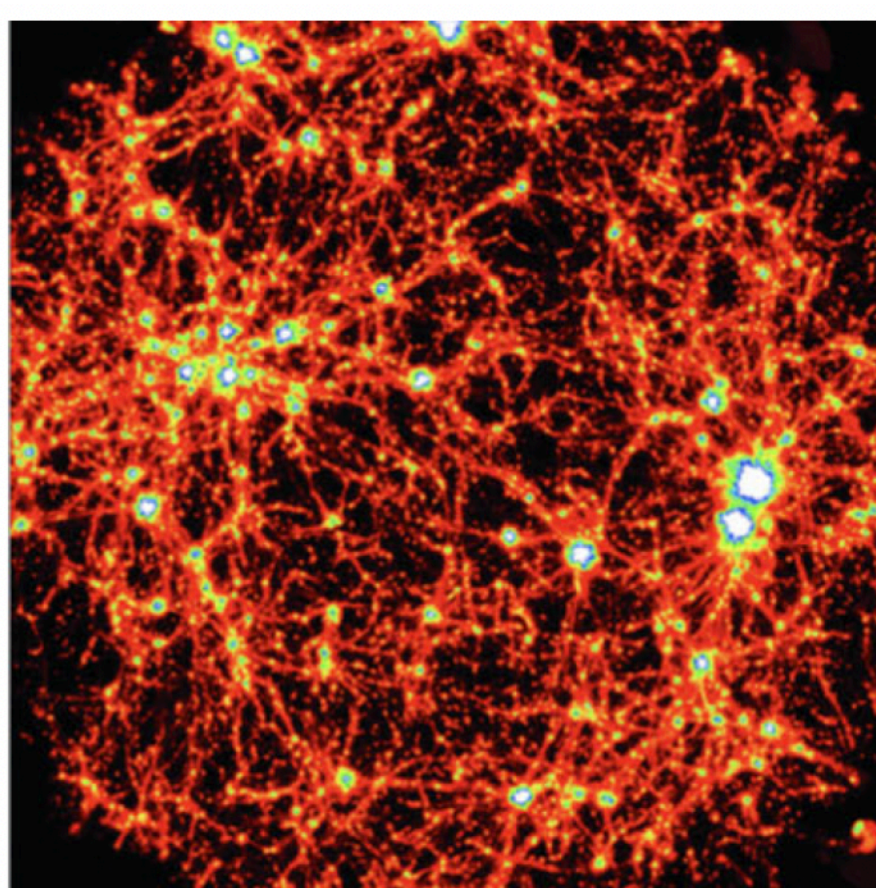
Discussion (2)



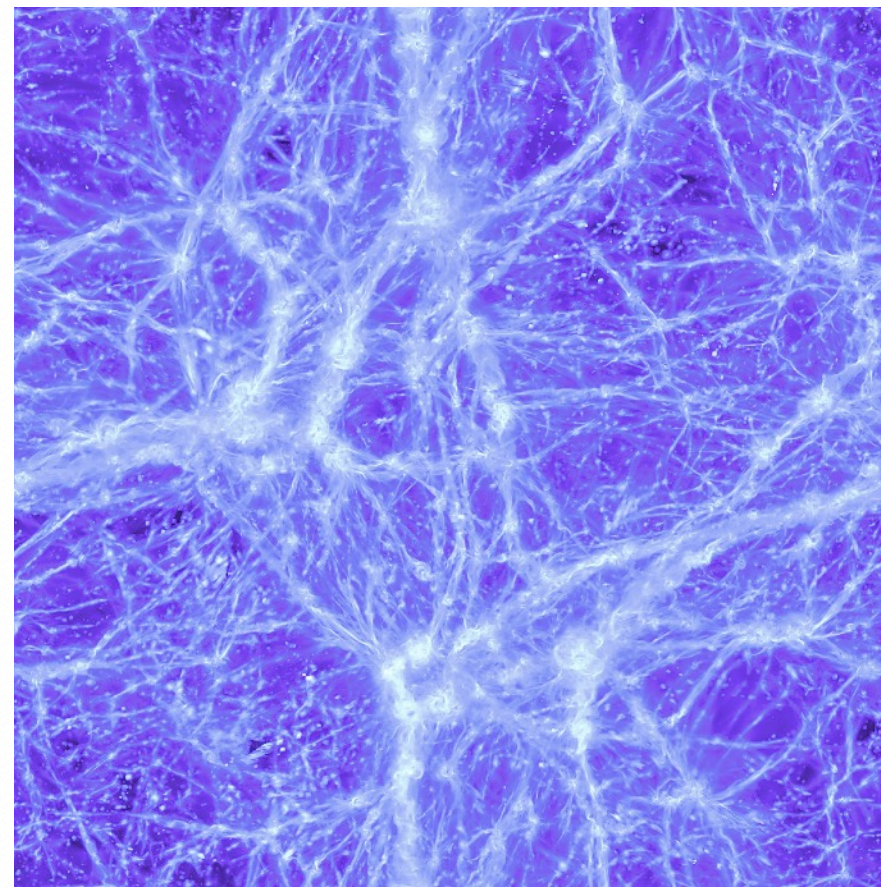
What can we trust and what we still do not trust in numerical predictions of cosmic magnetism?

MHD cosmological simulations

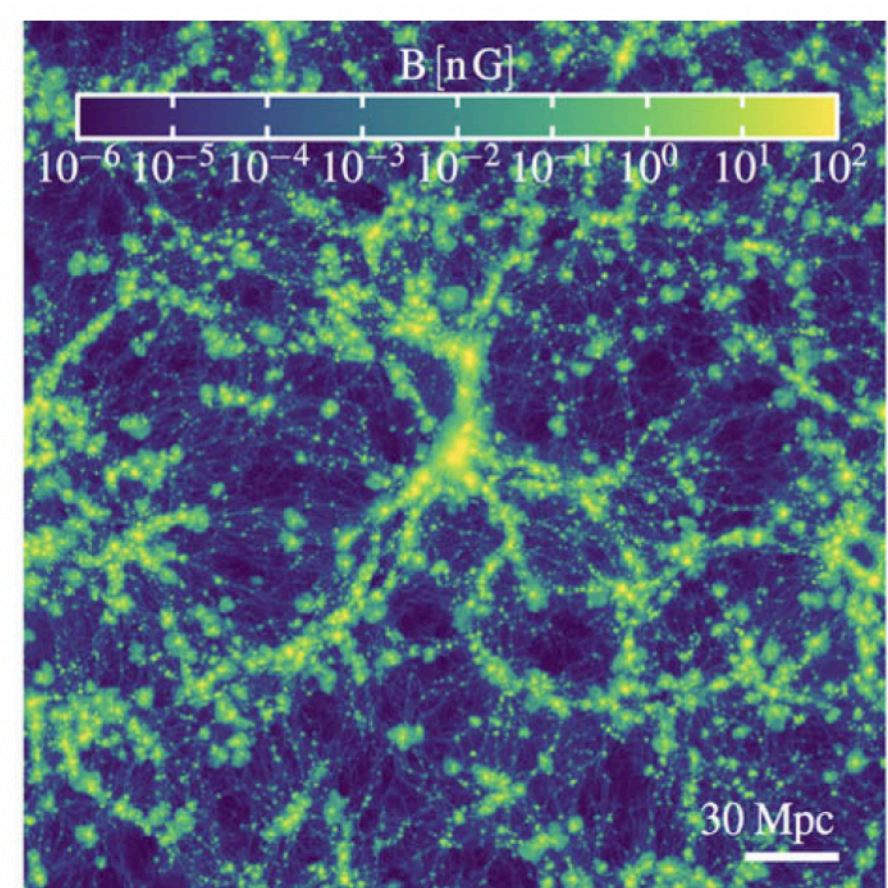
Broad agreement between MHD simulations of simple initial B fields evolving in large scale structures



Donnert+08 (Gadget - SPH)



Vazza+14 (ENZO-Eulerian)



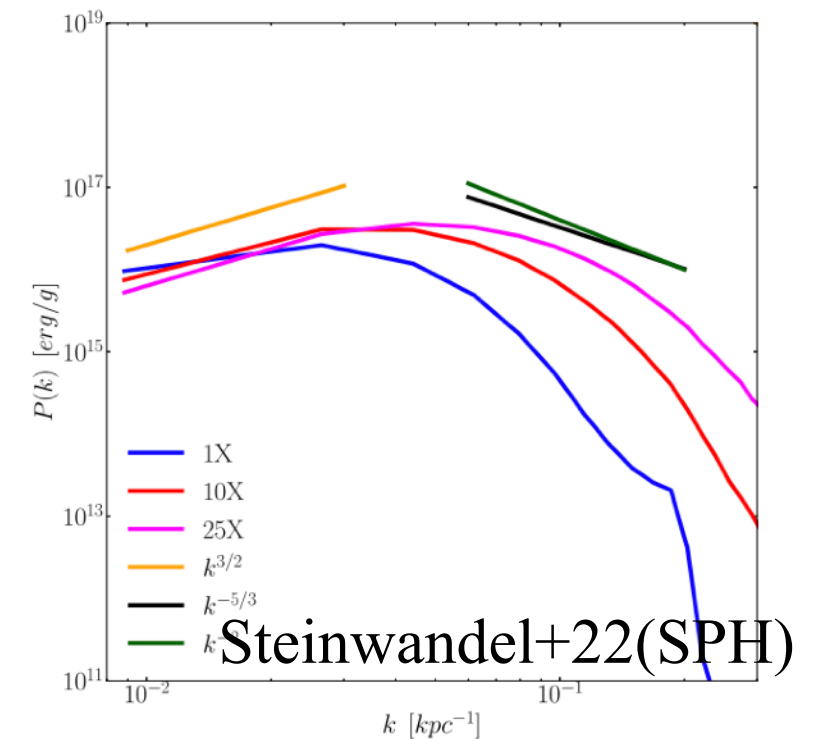
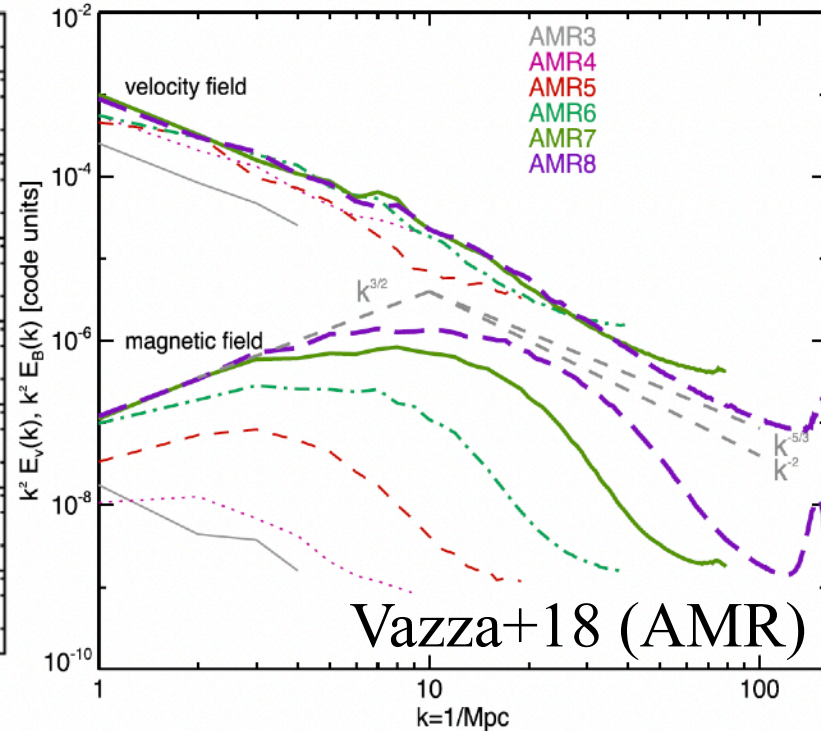
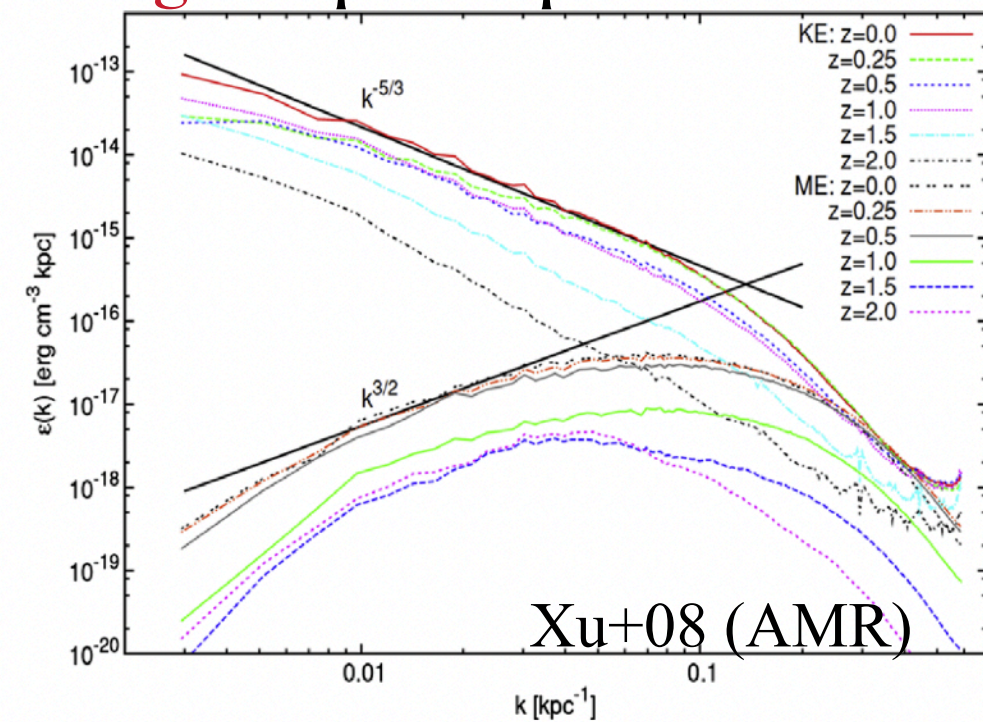
Marinacci+15 (AREPO-Moving Mesh)

But, what is the level of agreement by cosmological simulations on these facts:

- **galaxies and clusters of galaxies** do not retain memory of primordial seed fields
- **voids (also filaments?)** are the best place to look for primordial magnetic fields

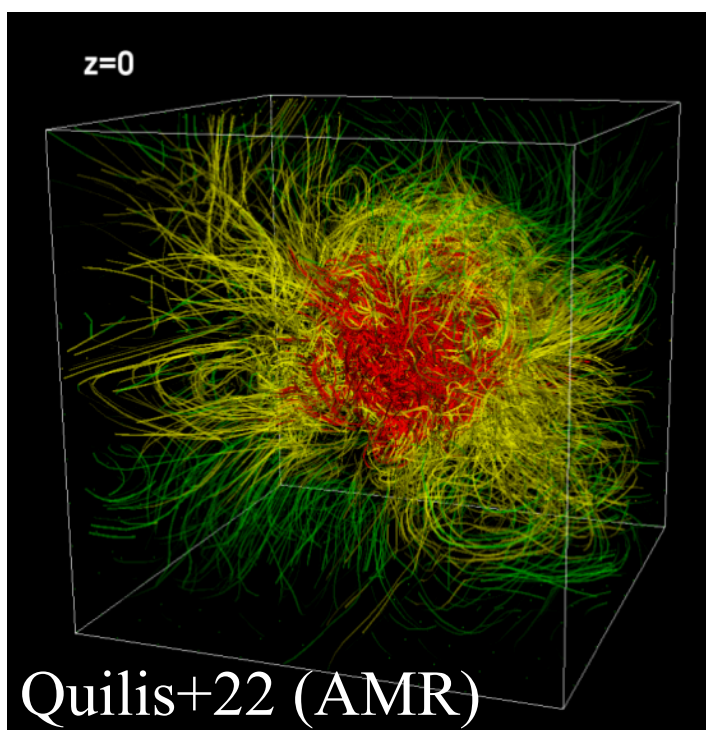
Intra Cluster Medium: dynamo

magnetic power spectra within clusters:



Small scale dynamo in the ICM:

- ▶ seen by all (?) ideal MHD simulations with high res. (≤ 10 kpc)
- ▶ a few % of conversion kinetic \rightarrow magnetic energy (Beresnyak+, Ryu+)
- ▶ ~compatible with radio observations of clusters
- ▶ **memory of seed fields is erased.**

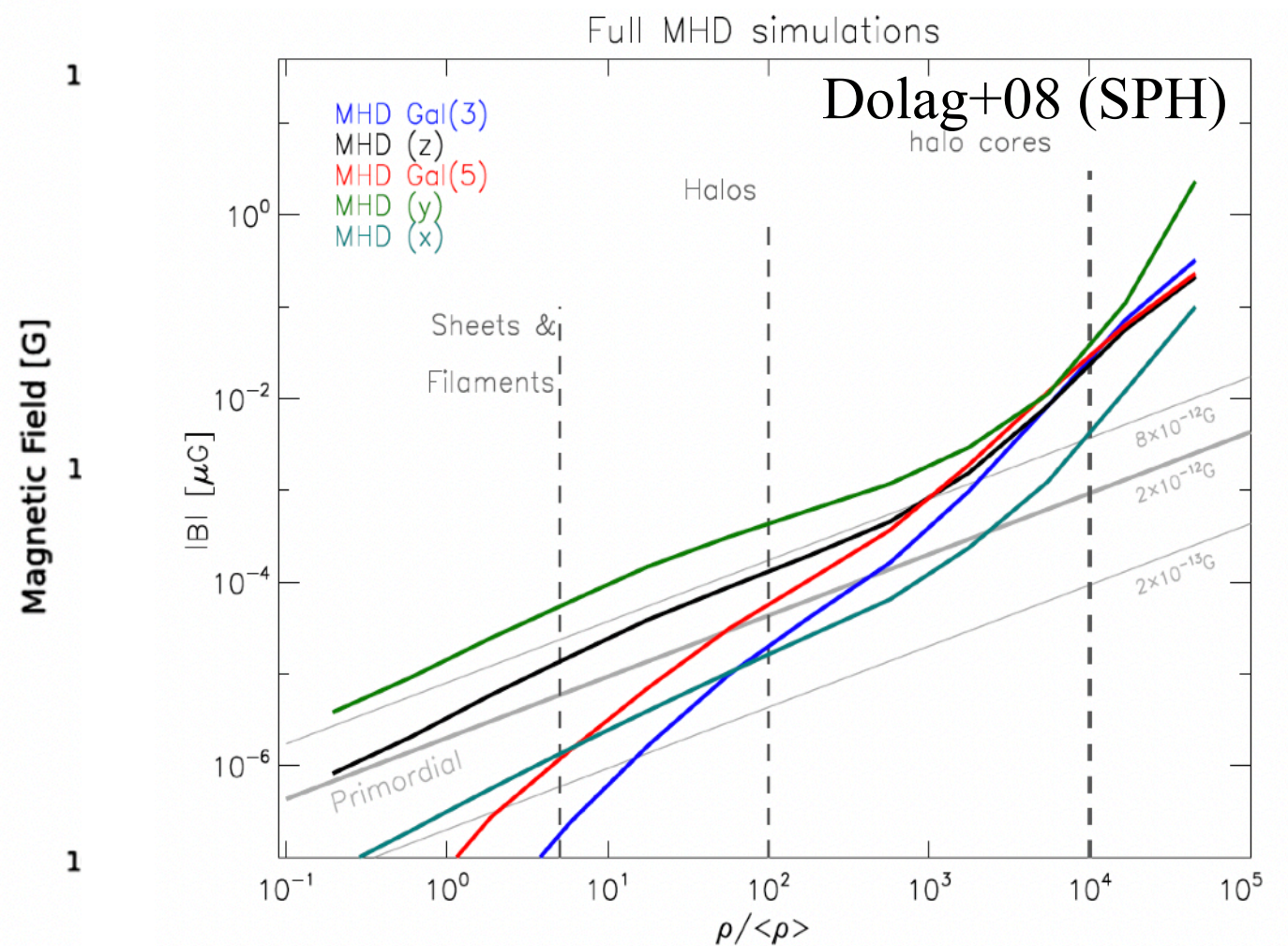
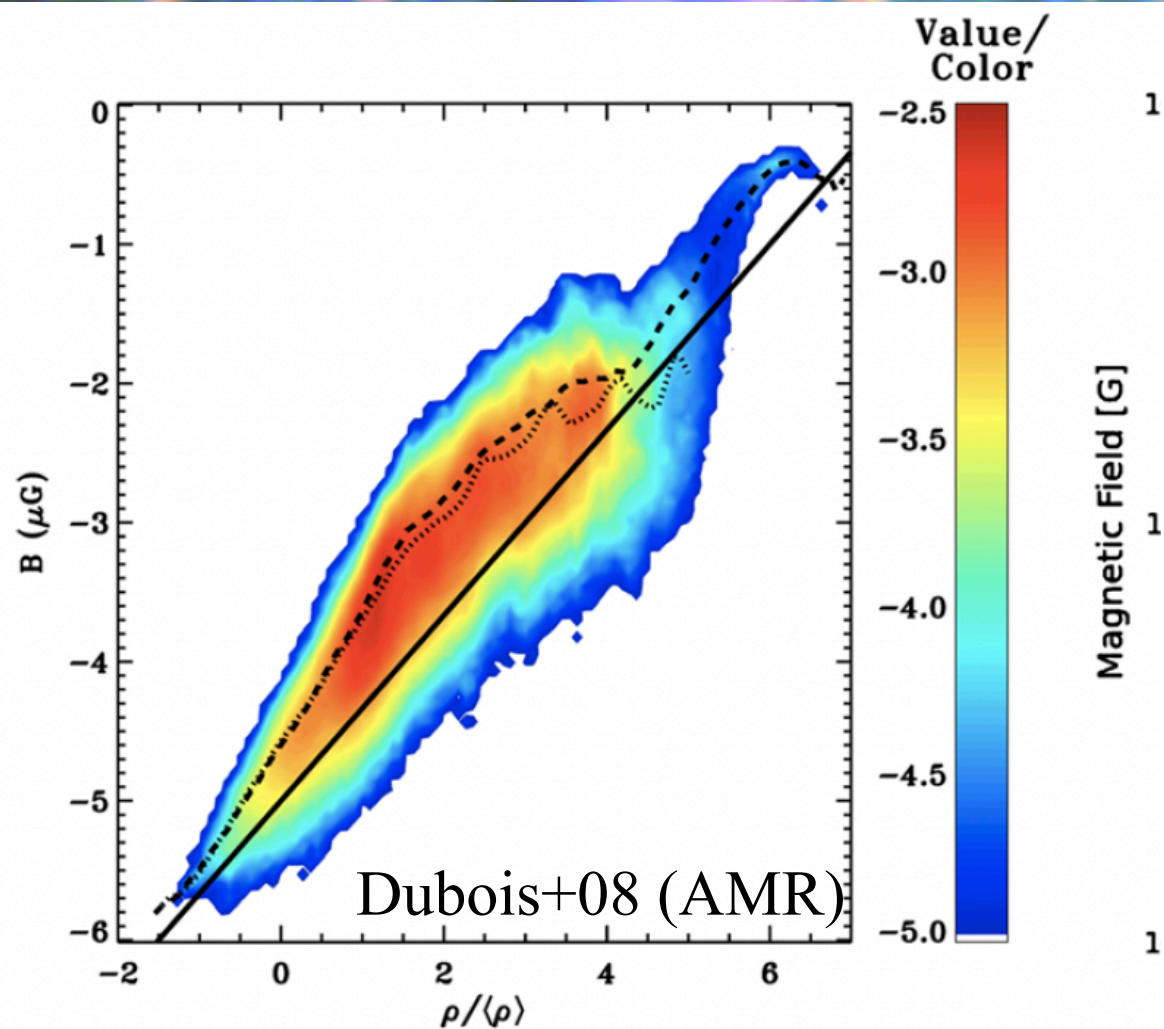


Do we agree?

Is there are critical radius for this (R_{100})?

Different story with helical fields?

Beyond clusters: no dynamo (?)



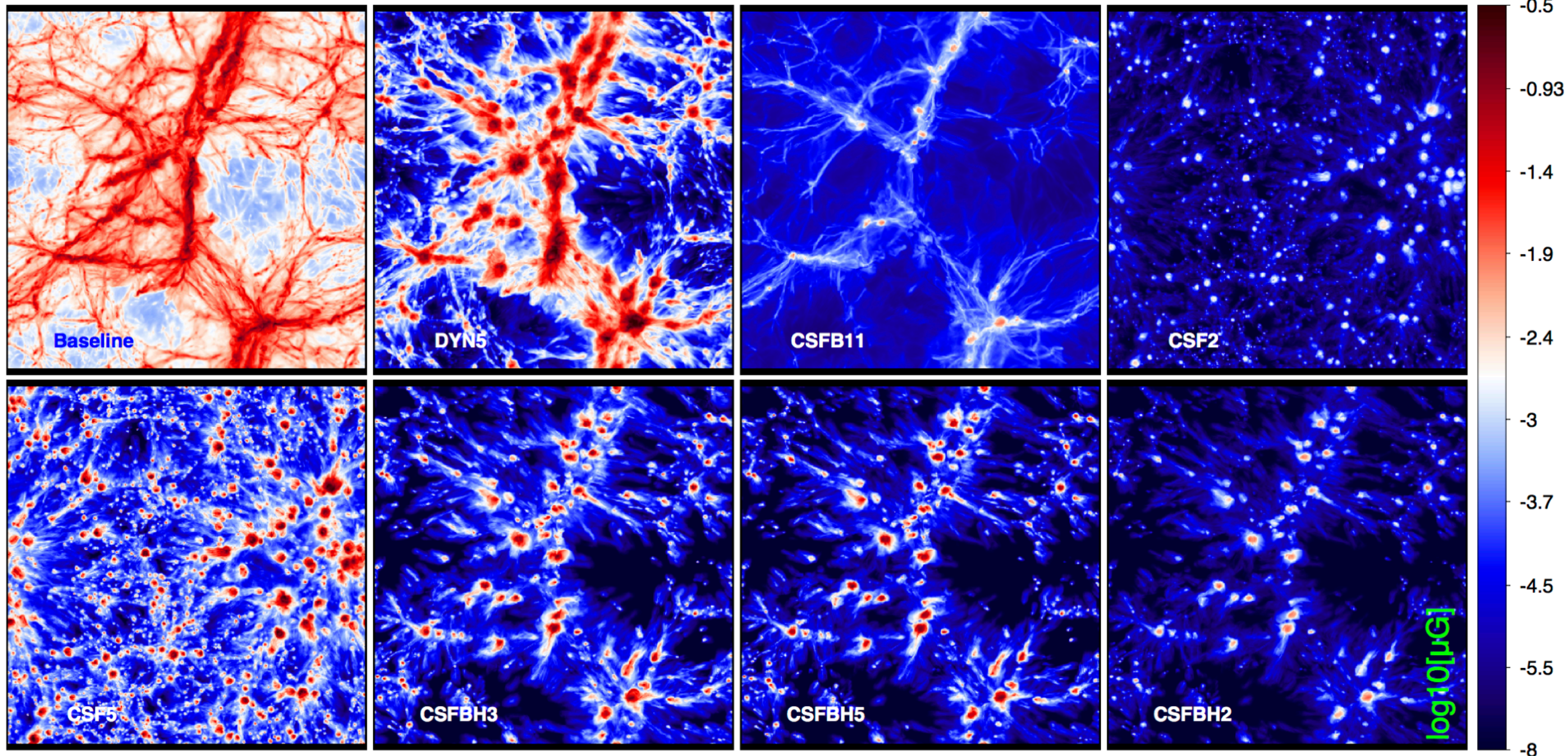
Magnetic field distribution in the cosmic web:

- ▶ mostly follow the $B \sim B_{seed} (\rho/\rho_0)^\alpha$ scaling ($\alpha \sim 2/3$)
- ▶ **memory of seed fields is conserved**

How much do “astrophysical B-fields” contaminate primordial B-fields?

Astroph. vs Primordial seeding

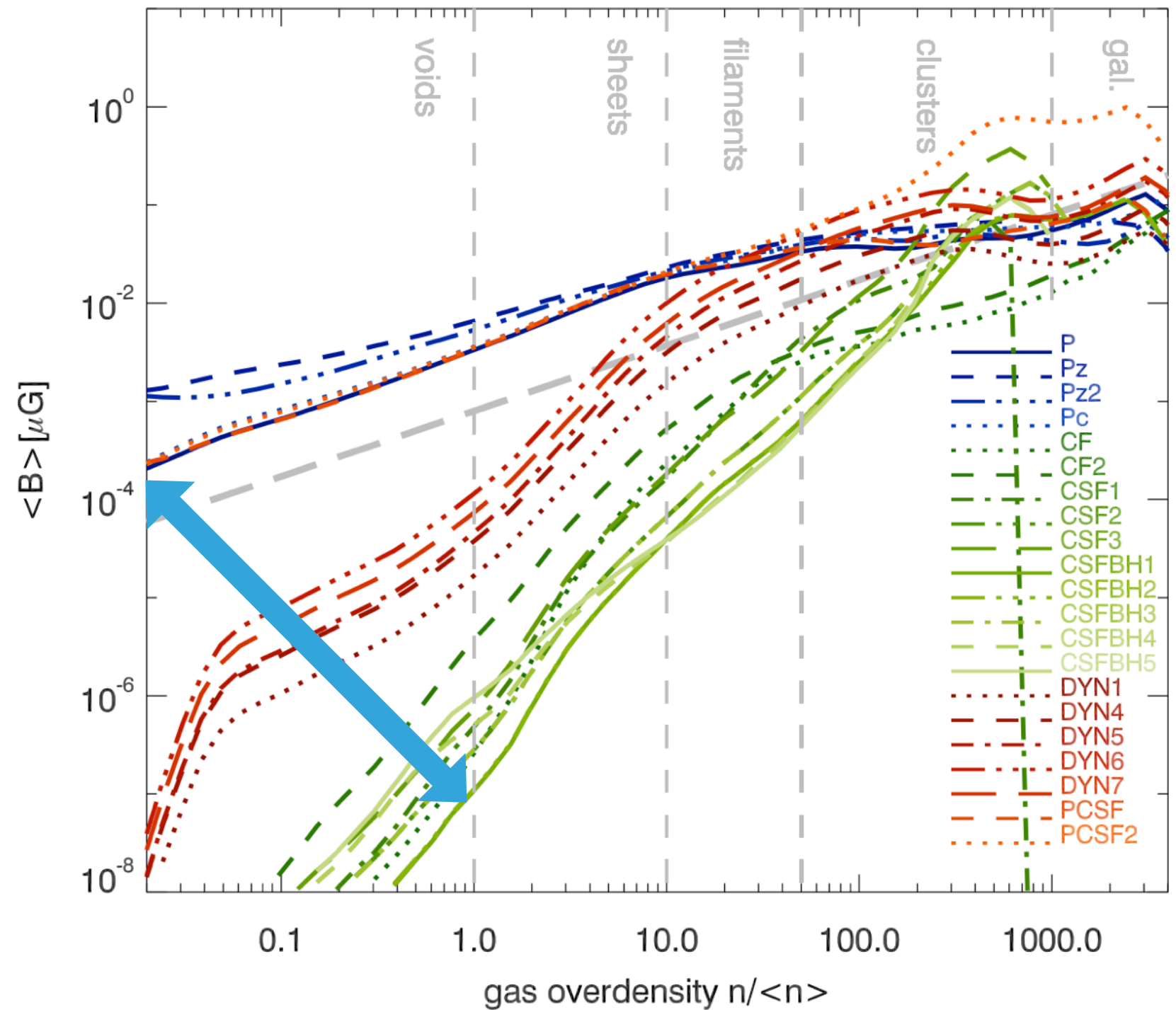
Vazza+17,21



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Astroph. vs Primordial seeding

Vazza+17,21



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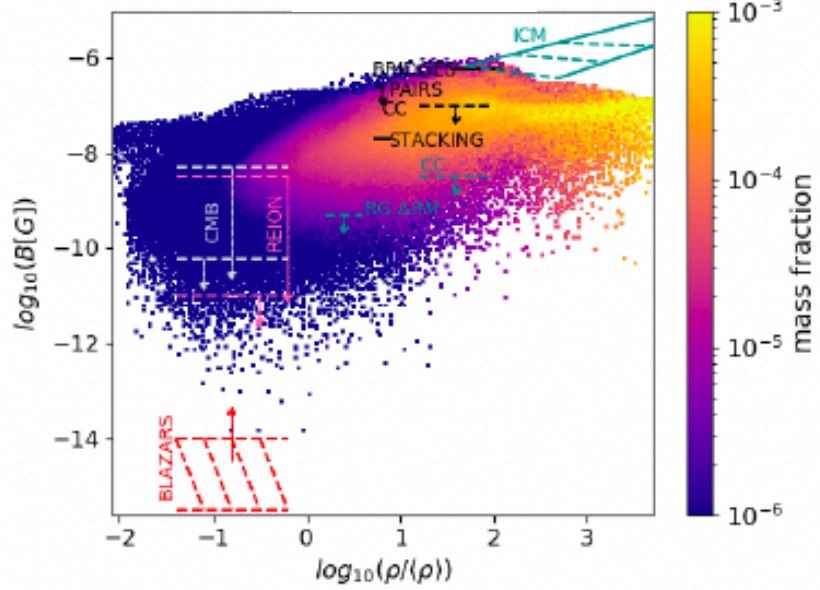
Compared with obs. limits, simple uniform primordial fields or purely astrophysical scenarios fail (somewhere).

$P_B(k) \propto k^{1.0-2.0}$ primordial B-fields give the best match to low-z observables

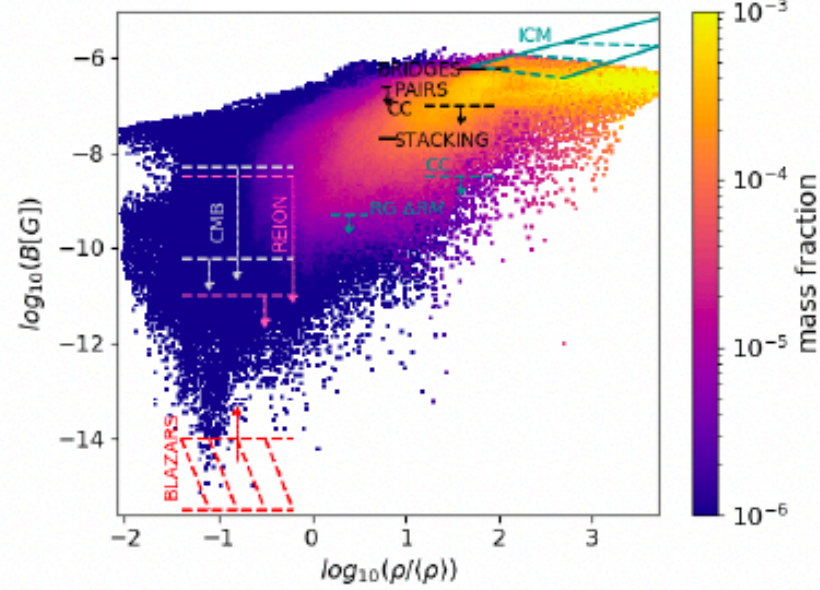
This relies on a simple \sim MHD evolution of B in voids!

$P_B(k) \propto k^{\alpha_B}, \alpha_B = 1.0$

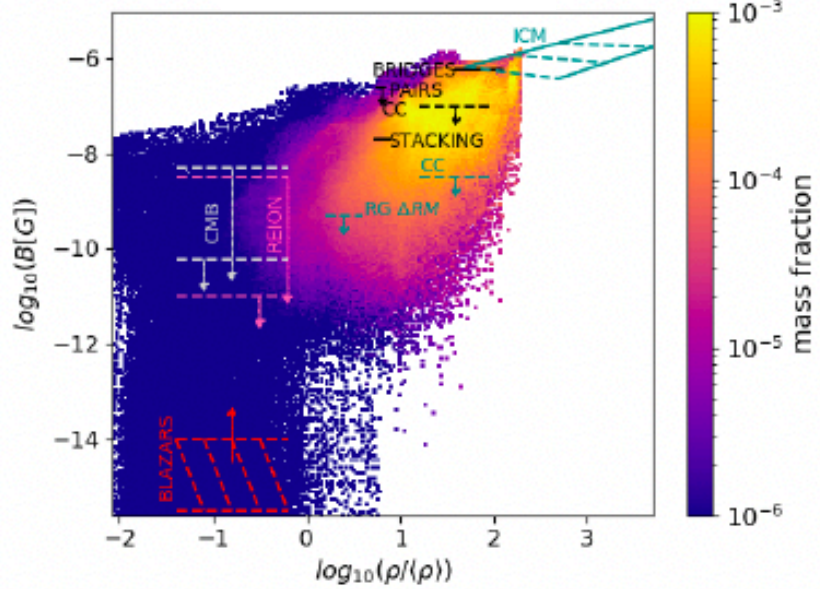
$B_0 = 1\text{nG}$



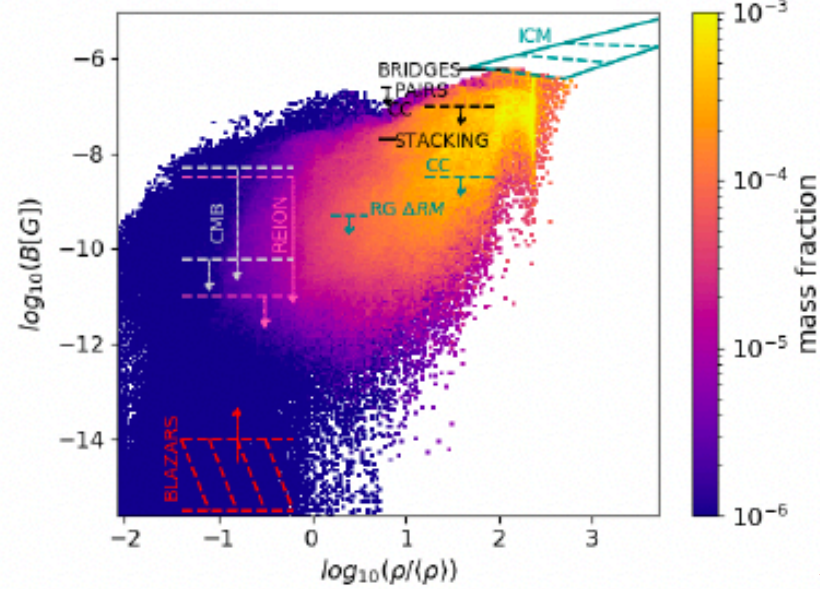
$B_0 = 10^{-7}\text{nG} + \text{dynamo}$



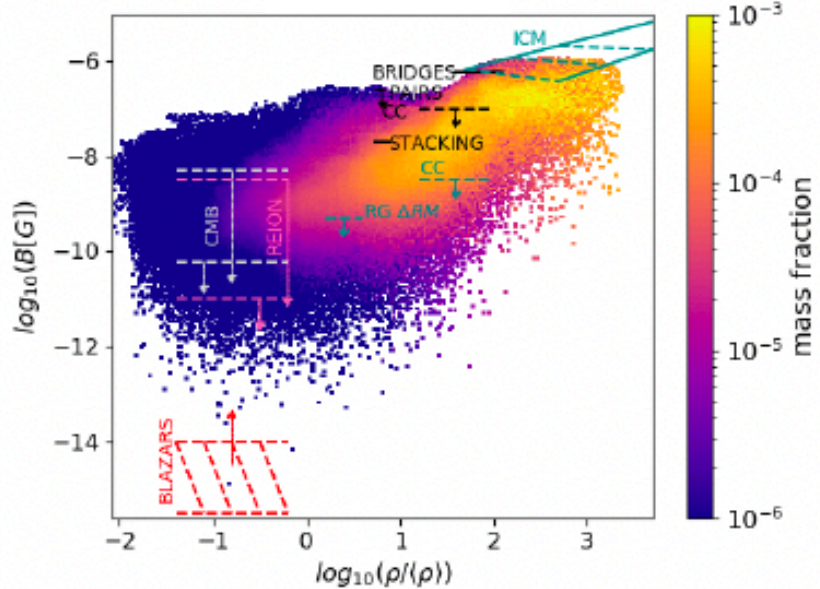
only star seeding



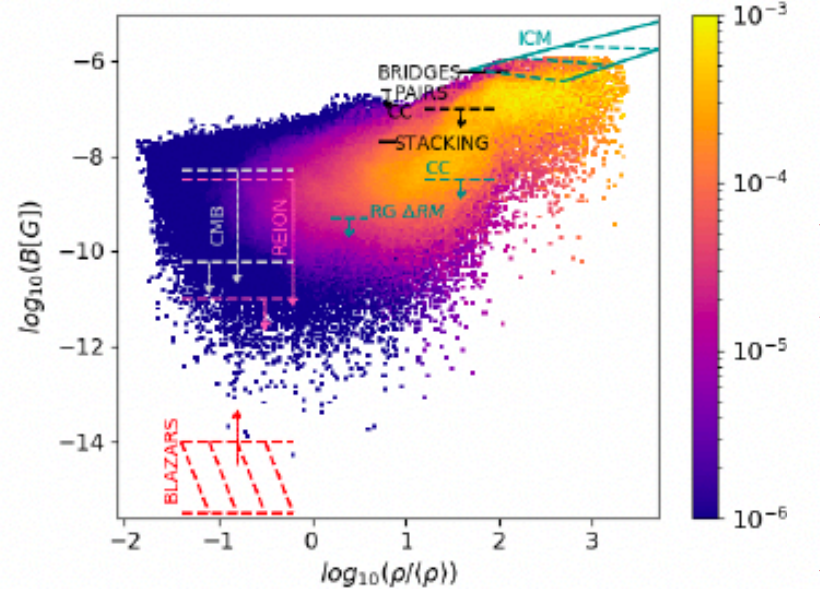
star and AGN seeding



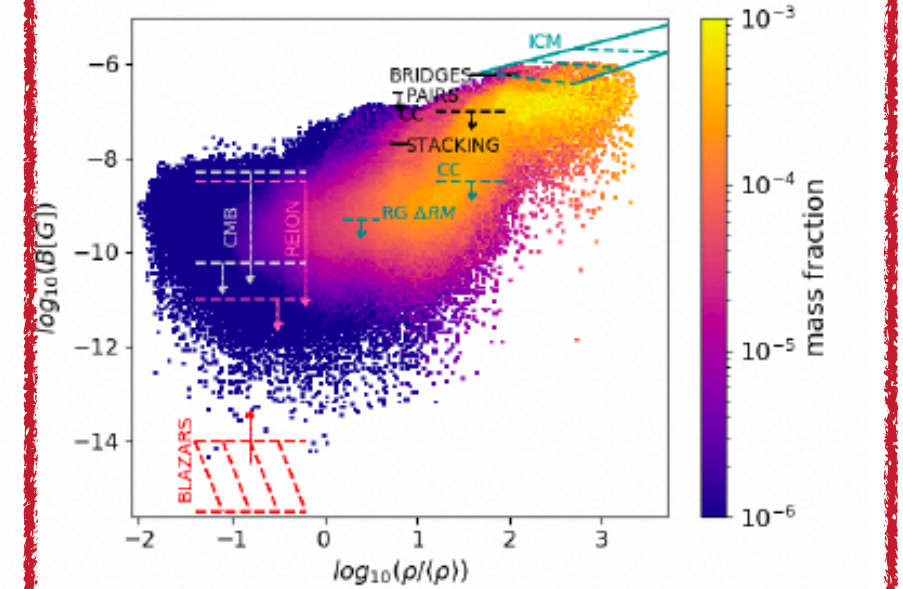
$B_0 = 0.1\text{nG}$



$P_B(k) \propto k^{\alpha_B}, \alpha_B = 0.0$

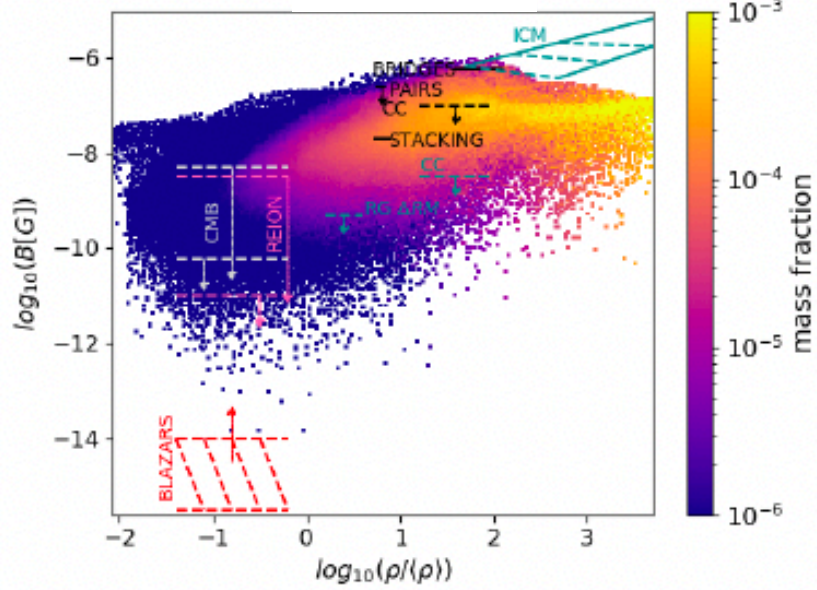


$P_B(k) \propto k^{\alpha_B}, \alpha_B = 1.0$

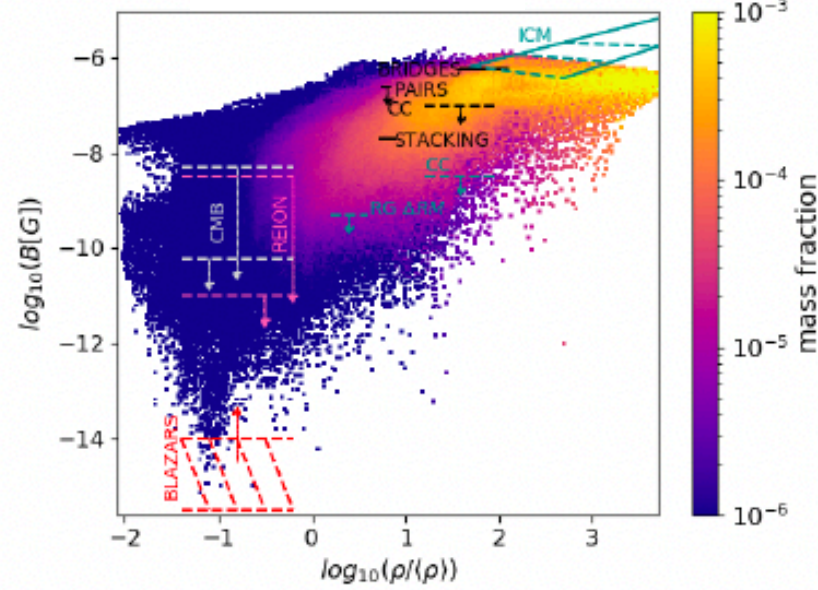


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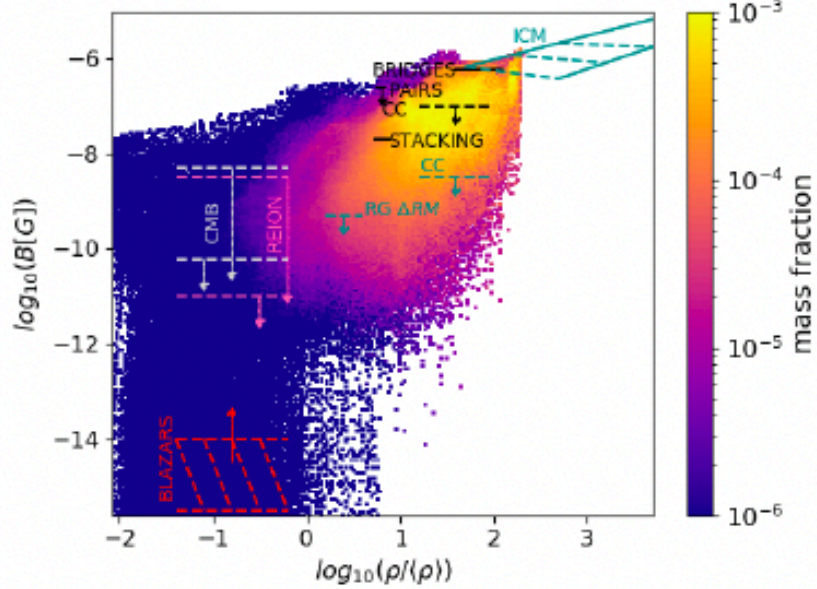
$B_0 = 1\text{nG}$



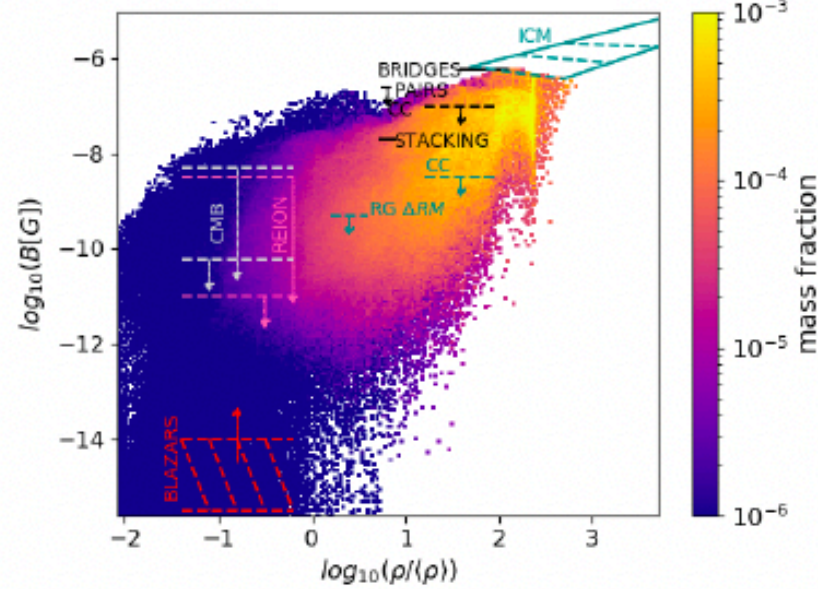
$B_0 = 10^{-7}\text{nG} + \text{dynamo}$



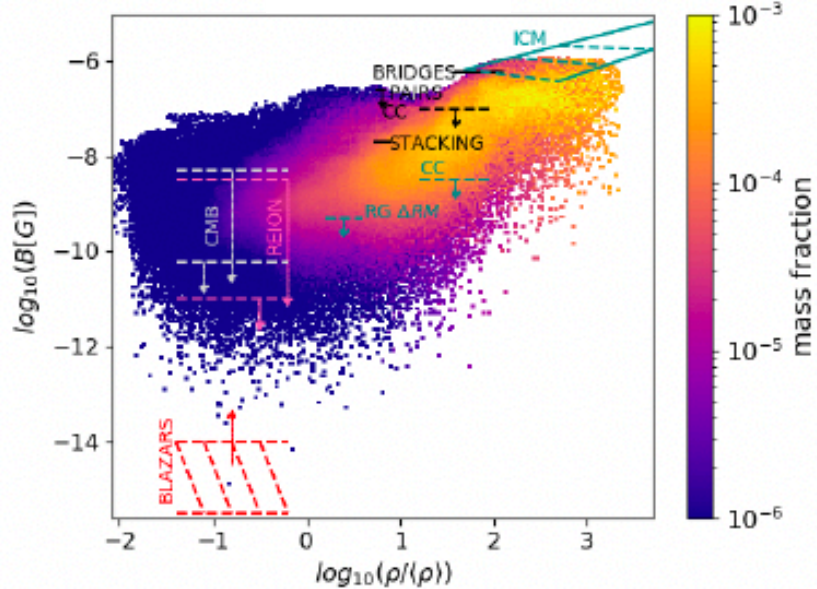
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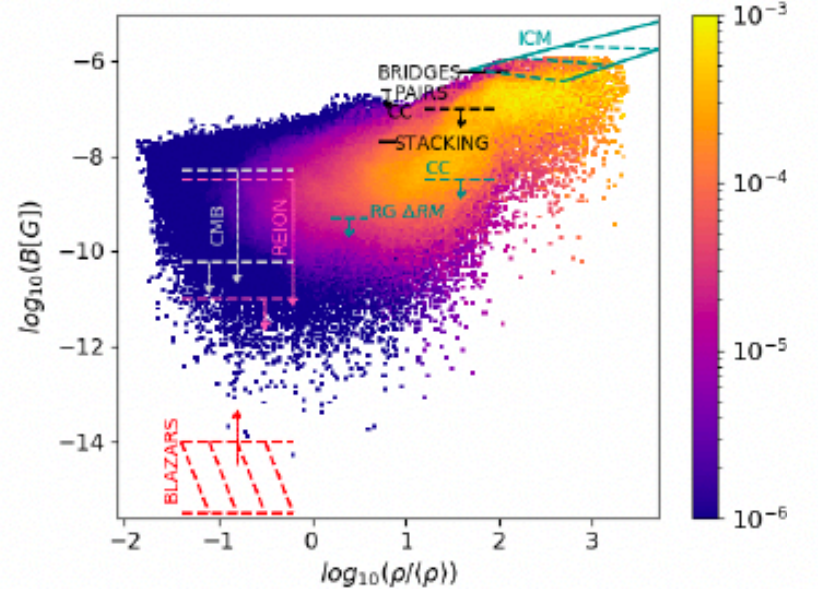
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