

IGMF limits with gamma rays

Discussion session

State of the art

Current limits:

- GeV (reprocessed emission) -> Fermi/IACTs
- Multi TeV (primary spectrum) -> IACTs

Observational remark:

- FoV of IACT is small, observations should be planned
- Observation strategy for transients (search for delayed emission) has to be optimized





Sources

- Need: multi TeV primary photons
- Obvious candidate: extreme blazars
 - BUT: faint both at GeV (Fermi) and TeV (IACTs), few sources known
- Multimessenger observations of blazars
- Recently: **GRBs** to probe smallest B range
 - BUT: still the population is small

Blazar sequence



 10^{-11} 10^{-13} <u>ල</u> භ 10⁻¹⁵

 10^{-17}

 $10^{-19} -$

 10^{-21}

 10^{-4}

Comments

- Nice limits... but we do not have to forget our assumptions:
 - Intrinsic spectrum (including cutoff at TeV)
 - Variability
 - Extragalactic background light (uncertain at large z)
 - Iterative propagation model
- Future instruments will boost this science case!
 - CTA
 - LHAASO/SWGO

Sensitivity to Assumptions on Spectra

- Consider impact on predicted $f_{\rm c}$ and $f_{\rm c}$ upper limit of
 - EBL model
 - From Gilmore 2012 (arXiv:1104.0671, fiducial model) & Franceschini 2008 (arXiv:0805.1841)
 - Intrinsic spectrum
 - Assume spectral index Γ = 1.660, based on Fermi measurement
 - · Does not account for possible variability on longer timescales
 - Consider Γ = [1.460, 1.660, 1.860]
 - Assume intrinsic spectrum described by exponentially cut-off power law
 - No cutoff in VERITAS spectrum != no cutoff
 - Highest energy spectral point @ 4 TeV
 - E_c = [5, 10, 20] TeV





IGMF detection with gamma rays

Discussion session

Open questions

- Can we set upper limits from gamma-ray data
- How solid are current limits?
 - Dependance on the line of sight \rightarrow can this be optimized/tested/...
- Can we remove some of the assumptions?
- Do we have observables to test IGMF models?