



# The structure of jets in Gamma Ray Bursts G. Ghirlanda

## **Outline**: 1) Collimation in GRBs 2) Structure of GRB jets

- 4) Perspectives

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3) Hunting for structure signatures





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## **GRB** recipe

## Ingredient #2: relativistic & collimated

**Relativistic:** 

- $\tau_{\gamma\gamma,R} \sim \Gamma^{2\beta-2} \tau_{\gamma\gamma,NR}$ 1) Compactness
- 2) Proper motion and/or size expansion



*Mooley*+2018; *GG*+2019

## Collimated:

1) Energy budget

 $E_{jet} \sim \frac{E_{\gamma,iso}\theta^2}{\eta}$ 

2) Afterglow jet break



Berger 2014

### How they look like depends on **Intrinsic properties** and **orientation**

# Viweing angle effects



### Strong depression of the observed luminosity (more prompt than afterglow) for slightly off axis observers

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Top hat or Uniform jet: *E*,  $\Gamma \mid \theta_{iet}$ 

# **GRB** diversity



Uniform jet Top hat jet

### GRB diversity —> Intrinsic







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Structured jet <u>Universal</u> jet





Strctured jet + cocoon Uniform jet + cocoon Structured cocoon

2017 Aug  $\rightarrow$  Many ...

. . .

(*Lipunov et al. 2001*) *Rossi + 2002; Zhang+ 2002* 

Structured jet = Universality (only orientation matters)



## Why structured jet is appealing



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Lu et al . 2018

## Viewing angle effects: structured jet



## Why do we care about jet structure?

- Expected
- Determines observable properties

## Jet structure definition

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·It conveys information on otherwise unobservable phenomena (Jetlaunching mechanism, jet-star material interaction, central engine ...)

> $\frac{dE'}{d\Omega}(\theta,t)$ (Jet internal energy)

 $\Gamma(\theta, t)$ 

## We care about jet structure!



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## A Picassian view of a GRB

![](_page_9_Figure_1.jpeg)

• Nature of the central engine Energy extraction mechanism

• Jet Head formation Forward/reverse shock • Cocoon - jet confinement effect

 Jet-cocoon breakout - free expansion • First light (shock breakout emission)

Freezing of angular structure

Prompt emission

Afterglow emission

Non-relativistic transition —> Jet structure erase

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Initial conditions

Angular structure

![](_page_9_Picture_13.jpeg)

## Adolescence

### Gottlieb et al. 2022

![](_page_10_Figure_2.jpeg)

#### **PROMPT EMISSION**

![](_page_11_Figure_2.jpeg)

## consistent with rather than constraining jet structure

Salafia et al. 2015, Salafia & Ghirlanda 2022

#### LONG GRBs

![](_page_12_Figure_2.jpeg)

**Pescalli et al. 2015**, 2016; Salafia 2015; GG&Salvaterra 2022

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 $\theta^{-\alpha}$  with  $\alpha < -4$ Or Gaussian

![](_page_12_Picture_6.jpeg)

![](_page_13_Figure_1.jpeg)

SHORT GRBs

Salafia & Ghirlanda 2022

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 $\theta^{-\alpha}$  with  $\alpha < -3$ 

![](_page_13_Picture_6.jpeg)

![](_page_14_Figure_1.jpeg)

![](_page_15_Figure_1.jpeg)

Beniamini et al. 2022

![](_page_16_Figure_1.jpeg)

Duque et al 2022

Steep decay / plateau: S. Ascenzi talk follows

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![](_page_16_Picture_6.jpeg)

1.Relativistic Jets in GRBs

2.Jet ⇔ progenitor vestige interaction ⇒ Jet structure

3.Structure ⇔ Unobservable GRB prop.

- Initial conditions
- •CE duration/enegy
- Vestige properties

4. $E(\theta)$ ;  $\Gamma(\theta) \propto \theta^{-\alpha}$ : • $\alpha > 3$  luminosity function (Pescalli et al. 2015) •  $\alpha > 3 \, \text{GW} / \text{GRB170817}$ (e.g. Ghirlanda et al. 2019) • GRMHD simulations (Gottlieb et al. 2022)

![](_page_17_Figure_8.jpeg)

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## Conclusions

HUNTING FOR JET STRUCTURE		
Observable	Constraining power	Diff
Prompt emission		
(spectrum, spectral energy	Low	Ea
correlations, etc)		
Early Afterglow (photometry)	High	M
Late Afterglow imaging	High	Η
olarization (prompt/afterglow)	Low	Η
Populations	Medium	Ea

COMBINATION OF SEVERAL OBSERVABLES IN FEW GRBS AND/OR POPULATION STUDIES

Salafia & Ghirlanda, 2022, Galaxies, 10(5), 93 https://www.mdpi.com/2075-4434/10/5/93

![](_page_17_Picture_15.jpeg)

![](_page_18_Picture_1.jpeg)

## BACKUP SLIDES

![](_page_19_Figure_1.jpeg)

#### Salafia, Ghirlanda, Ascenzi, Ghisellini 2019

![](_page_19_Picture_3.jpeg)