



GRB 080928 afterglow polarization analysis and GRBs spectropolarimetry

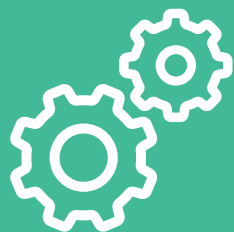


GRBV – V Congresso Nazionale GRB - TRIESTE

*R. Brivio, S. Covino, P. D'Avanzo, K. Wiersema, J.R. Maund,
M.G. Bernardini, S. Campana, A. Melandri*



GRB
POLARIMETRY



DATA ANALYSIS



RESULTS

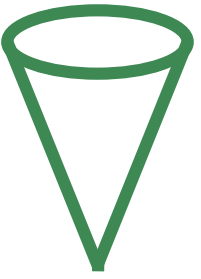


MODELING



CONCLUSIONS

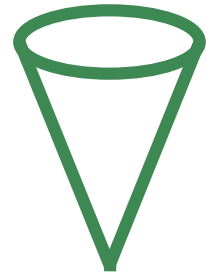
GRBs POLARIMETRIC ANALYSIS



POLARISATION OF ELECTROMAGNETIC RADIATION

$$\begin{bmatrix} I \\ Q \\ U \\ V \end{bmatrix} \quad P_l = \frac{\sqrt{Q^2 + U^2}}{I}$$
$$\theta = \frac{1}{2} \text{atan} \left(\frac{U}{Q} \right)$$

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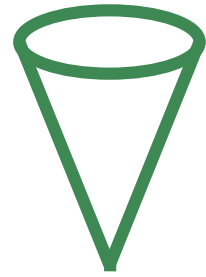
POLARISATION OF
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GAMMA-RAY BURSTS

PARADIGM?

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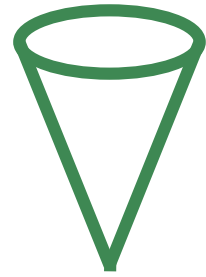
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Emitting region features,
jet geometry, magnetic fields, ...

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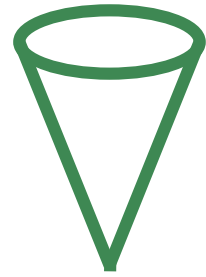
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POLARISATION ANALYSIS!

GRBs POLARIMETRIC ANALYSIS



POLARISATION OF
ELECTROMAGNETIC RADIATION

Asymmetries in the
emitting region

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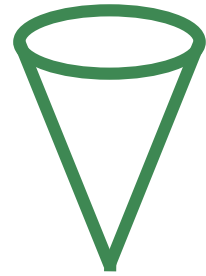
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GRBs POLARIMETRIC ANALYSIS



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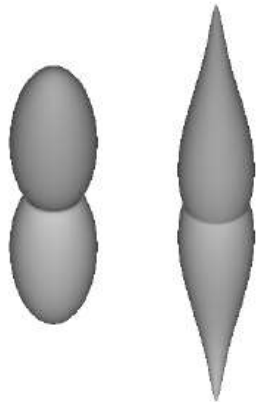
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Asymmetries in the emitting region

INTRINSIC?

GAUSSIAN/
STRUCTURED JET



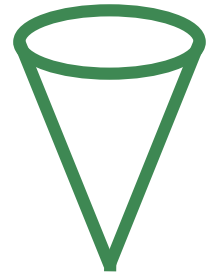
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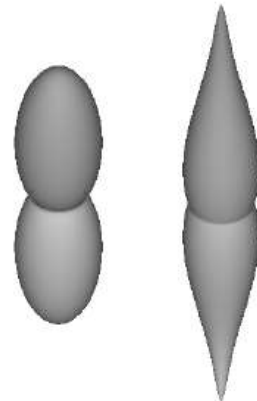
Asymmetries in the
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INTRINSIC?

ANGLE-DEPENDENT?

GAUSSIAN/
STRUCTURED JET

HOMOGENEOUS JET



GAMMA-RAY BURSTS

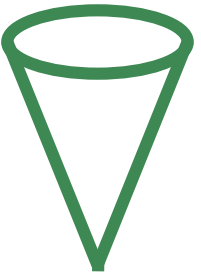
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POLARISATION ANALYSIS!

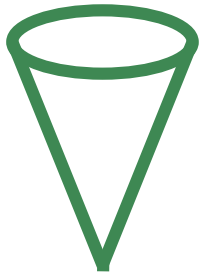
From Rossi E. et al (2004)

GRB AGs SPECTRO-POLARIMETRIC OBSERVATIONS



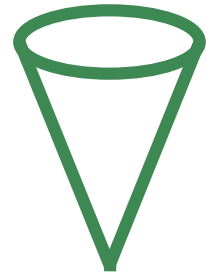
GRB	Telescope	Instrument	$t - t_0$ [h]	P [%]	ϑ [°]	Ref
020813	Keck	LRISp	4.7 – 7.9	1.8 – 2.4	148 – 162	Barth et al. (2003)
021004	VLT	FORS1	18.83	0.8 – 1.7	100 – 147	Wang et al. (2004)
	VLT	FORS1	“	1.9	118	Lazzati et al. (2004)
030329			24 – 72	0.3 – 1.5		Greiner et al. (2004)
	VLT	FORS1	86.4	0.5 – 0.9	73 – 83	Covino et al. (2003)
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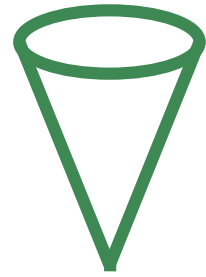


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No evidence of
 λ -dependence

Mild ISM
polarisation

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080928	VLT	FORS1				Brivio et al. (accepted in A&A)

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GRB 080928 - INSTRUMENTATION

INSTRUMENTATION: ESO-VLT FORS1

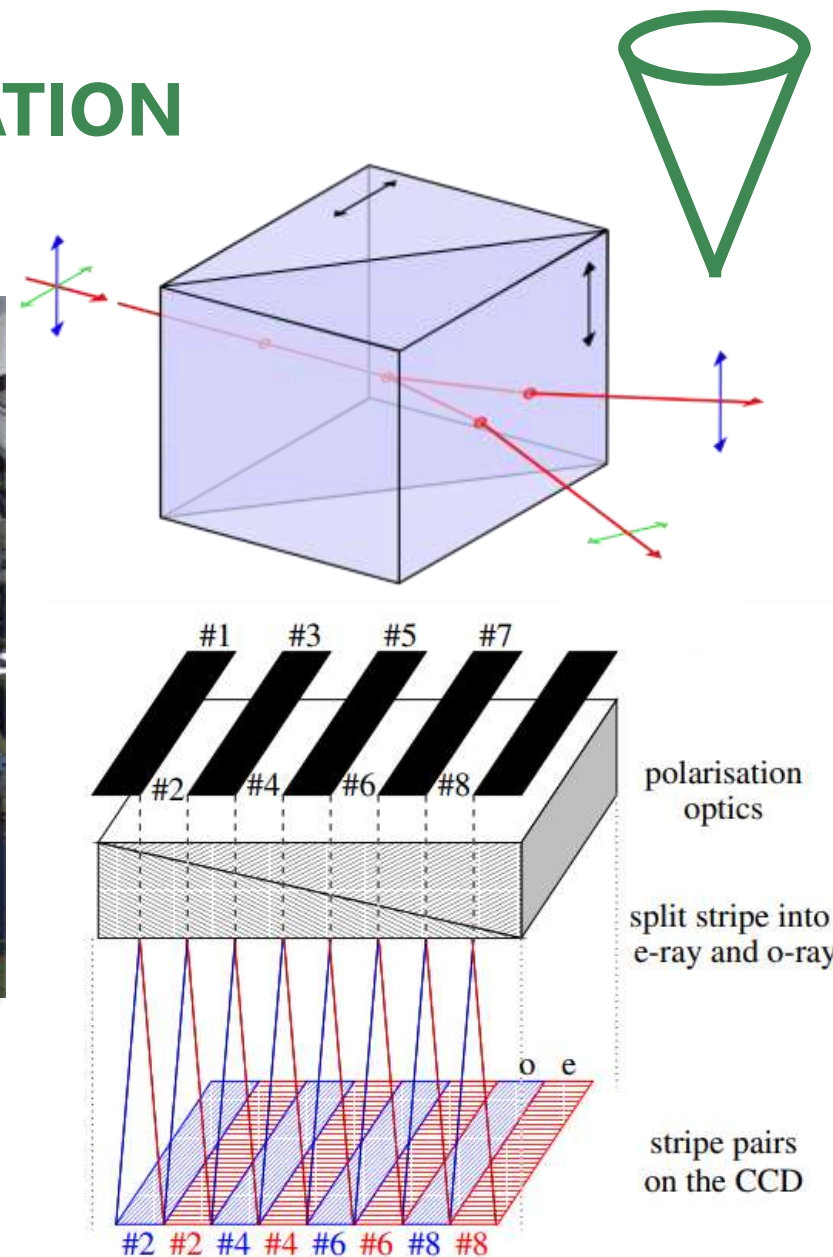
LOCATION: CASSEGRAIN focus of UT1 (ANTU)

OBSERVING MODES: Imaging - spectroscopy
Imaging/Spectro-polarimetry

WAVELENGTH: OPTICAL, 330-1100 nm

SPATIAL RESOLUTION: 0.25"/pixel (SR) / 0.125"/pixel (HR)

SPECTRAL RESOLUTION: 260 to 1600 (low to medium)



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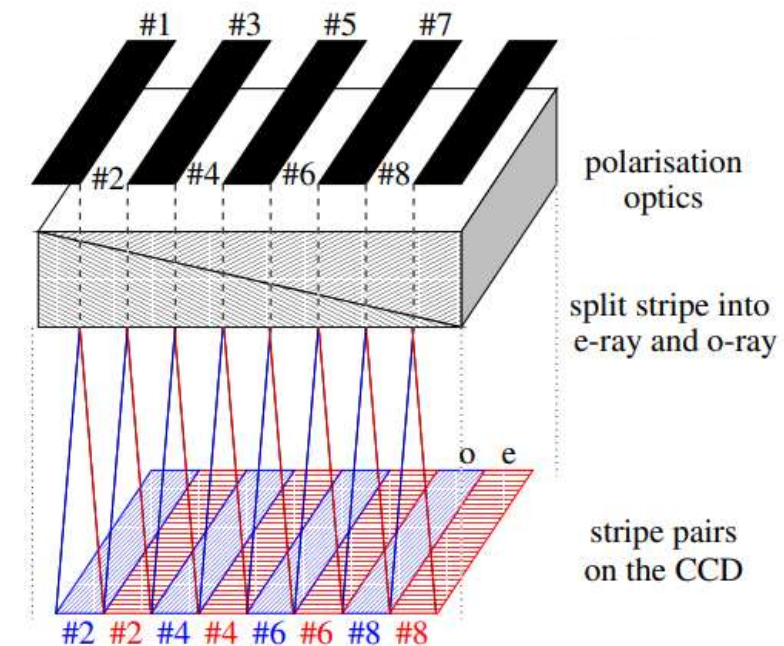
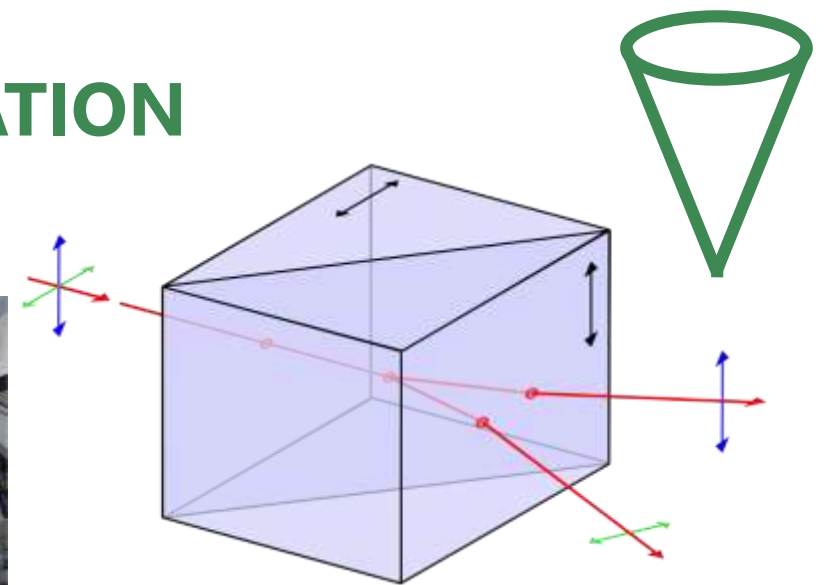
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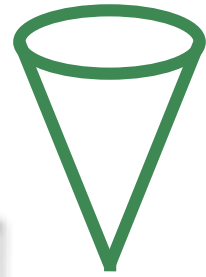
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$$\frac{Q}{I} = \frac{1}{2} \left[\left(\frac{f_o - f_e}{f_o + f_e} \right)_{0^\circ} - \left(\frac{f_o - f_e}{f_o + f_e} \right)_{45^\circ} \right]$$

$$\frac{U}{I} = \frac{1}{2} \left[\left(\frac{f_o - f_e}{f_o + f_e} \right)_{22.5^\circ} - \left(\frac{f_o - f_e}{f_o + f_e} \right)_{67.5^\circ} \right]$$



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OBSERVATIONS

IPOL	14.01 – 14.81 h	V_HIGH filter
PMOS	14.95 – 16.60 h	300V grism
IPOL	39.39 – 42.01 h	V_HIGH filter

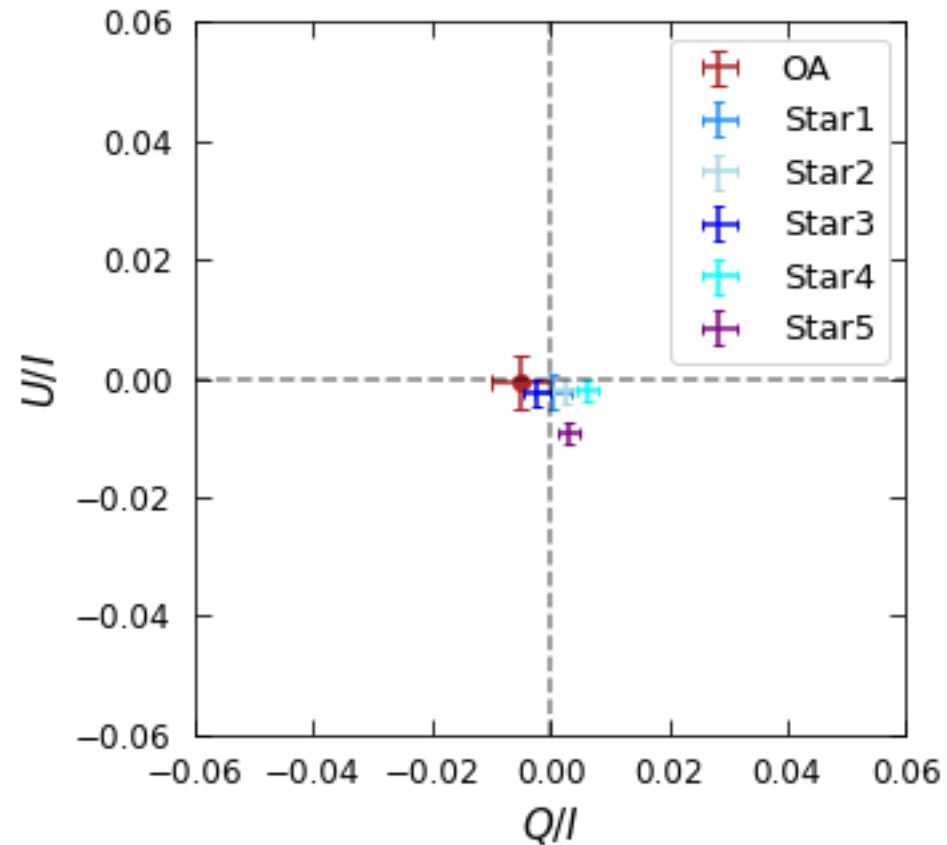
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IMAGING POLARIMETRY RESULTS

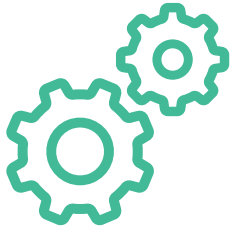


1st night – $t - t_0 \sim 0.62$ d

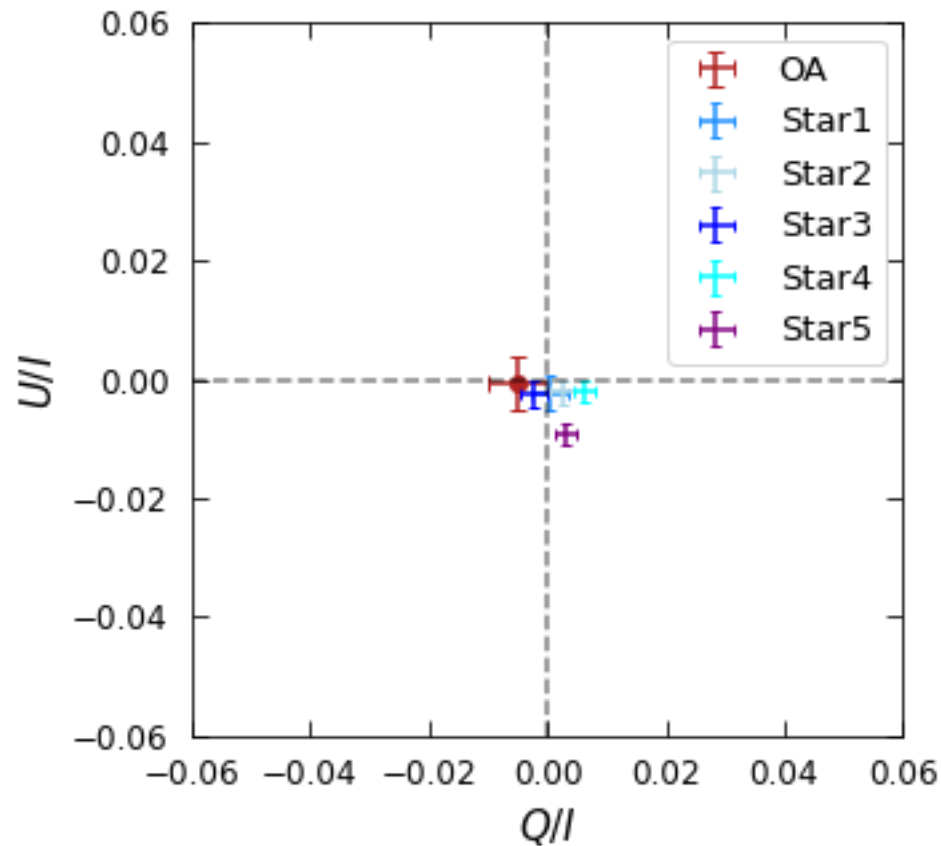


$P_{OA} < 1.99\%$

IMAGING POLARIMETRY RESULTS

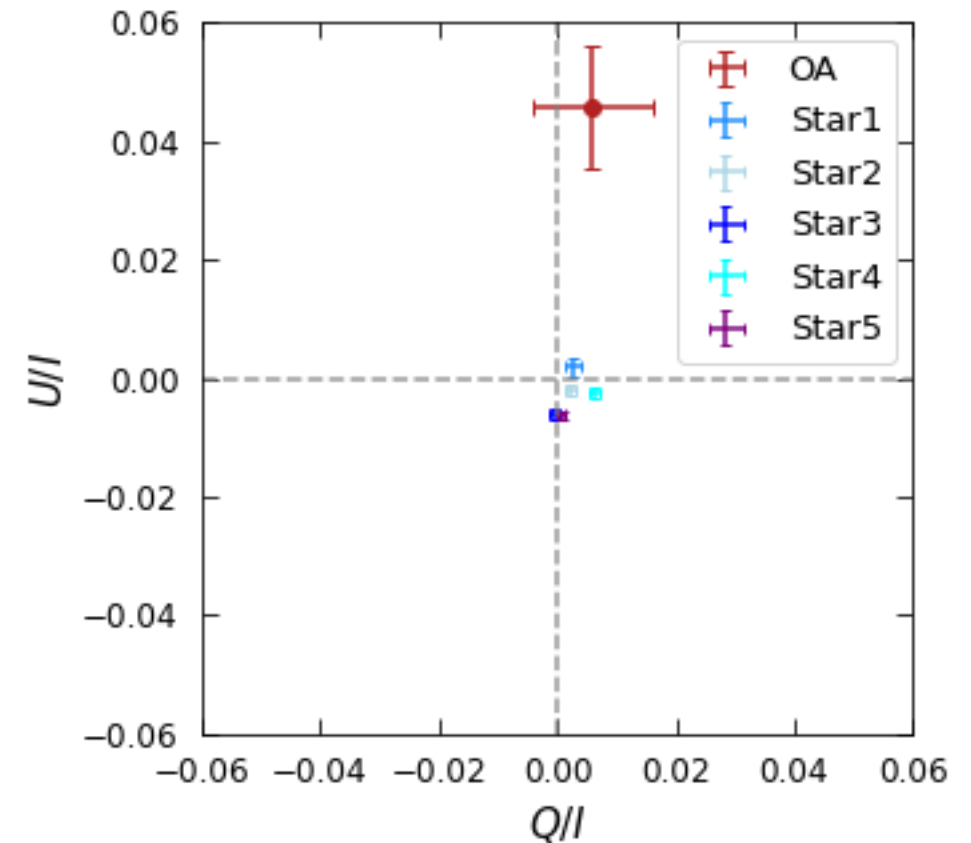


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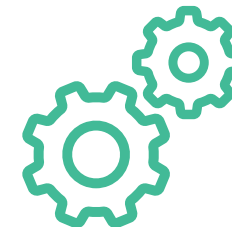
$$P_{OA} < 1.99 \%$$

2nd night – $t - t_0 \sim 1.70$ d



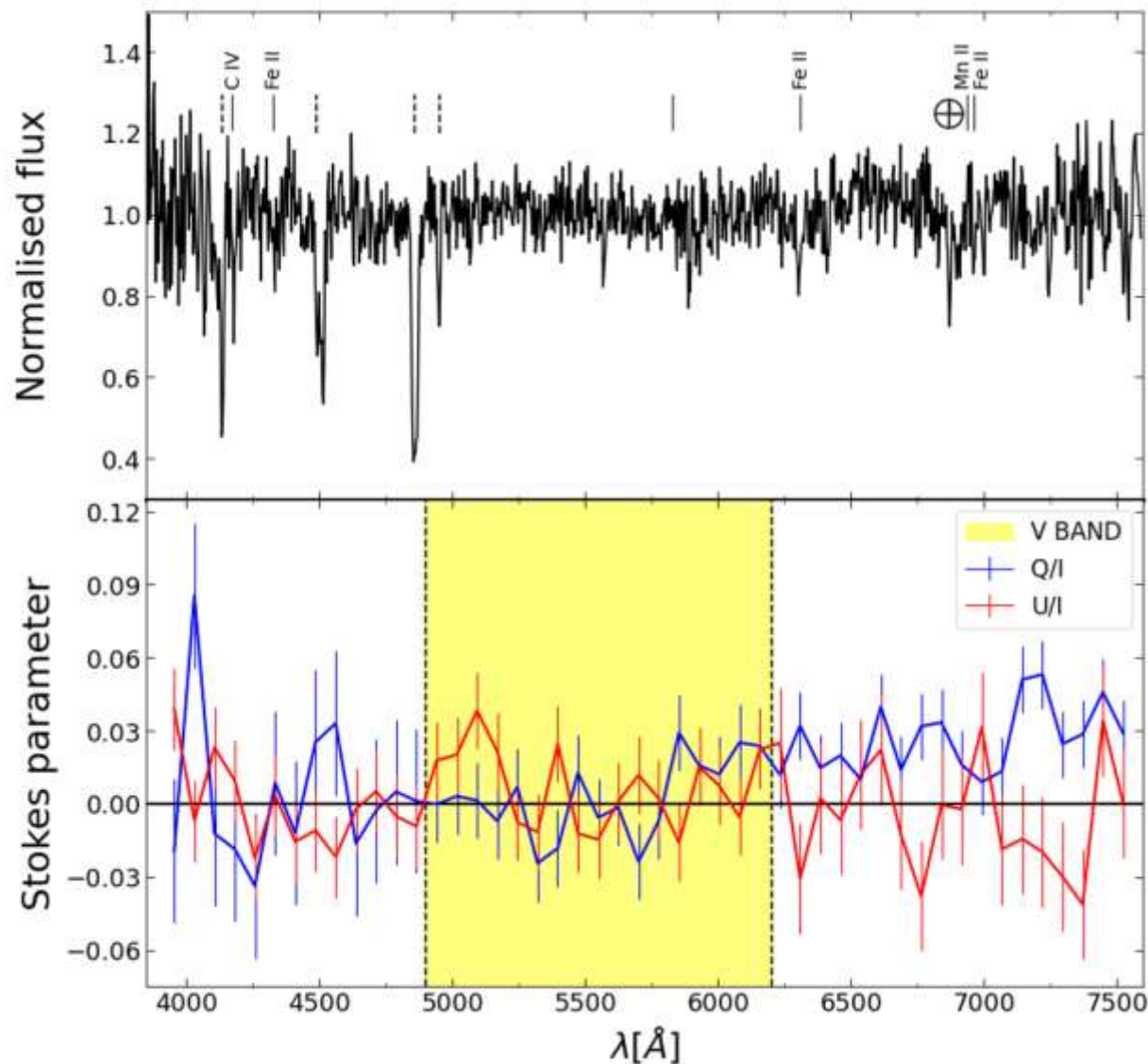
$$P_{OA} = 4.49^{+1.16}_{-0.96} \%$$

SPECTRO-POLARIMETRY RESULTS

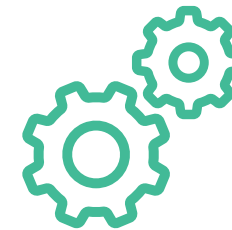


1st night – $t - t_0 \sim 0.66$ d

$P_{0A} < 1.57\%$

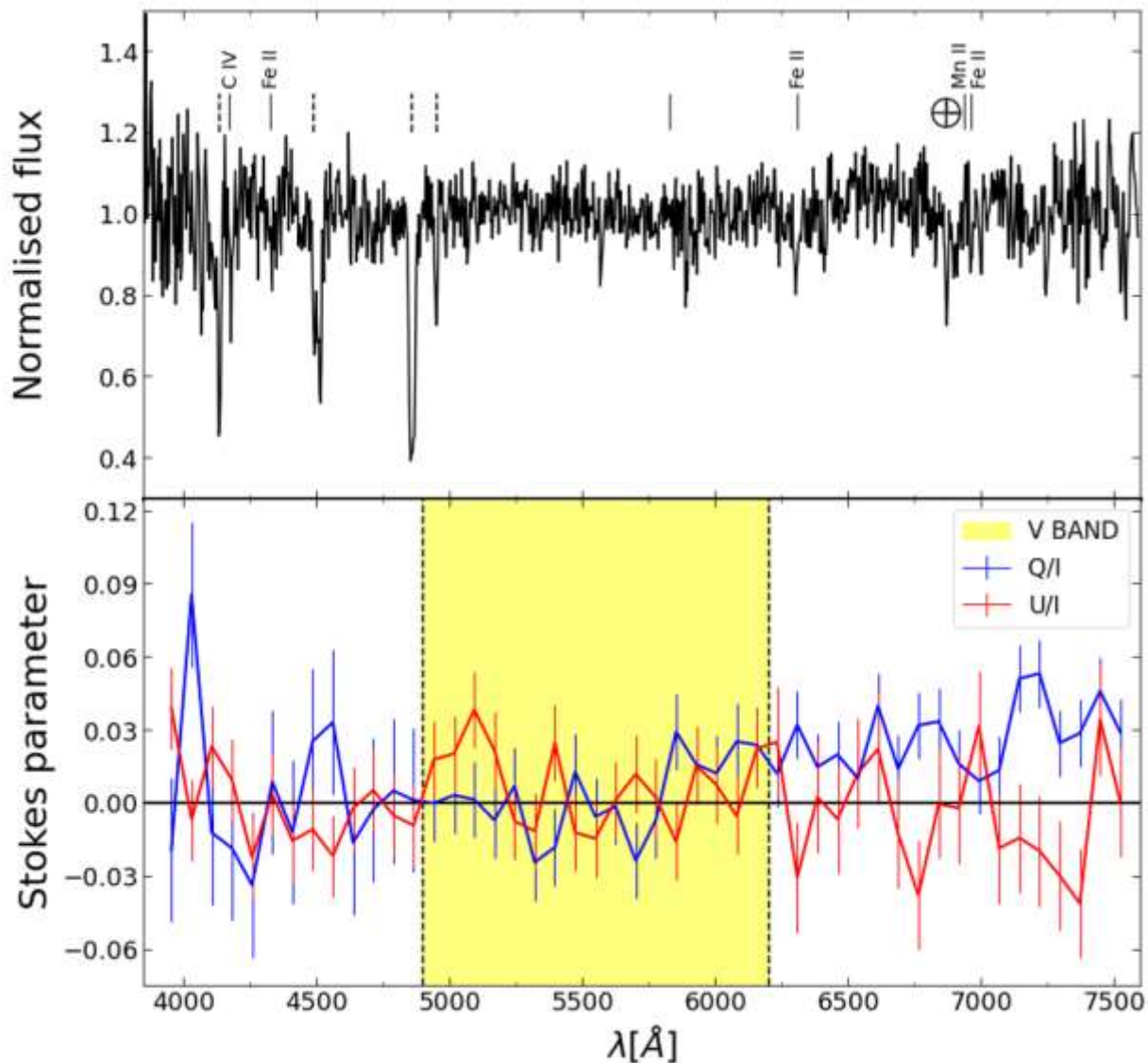


SPECTRO-POLARIMETRY RESULTS



1st night – $t - t_0 \sim 0.66$ d

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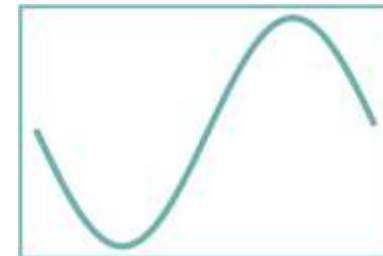
$$P_{max}(\%) \leq 9.0 \times E(B - V)$$



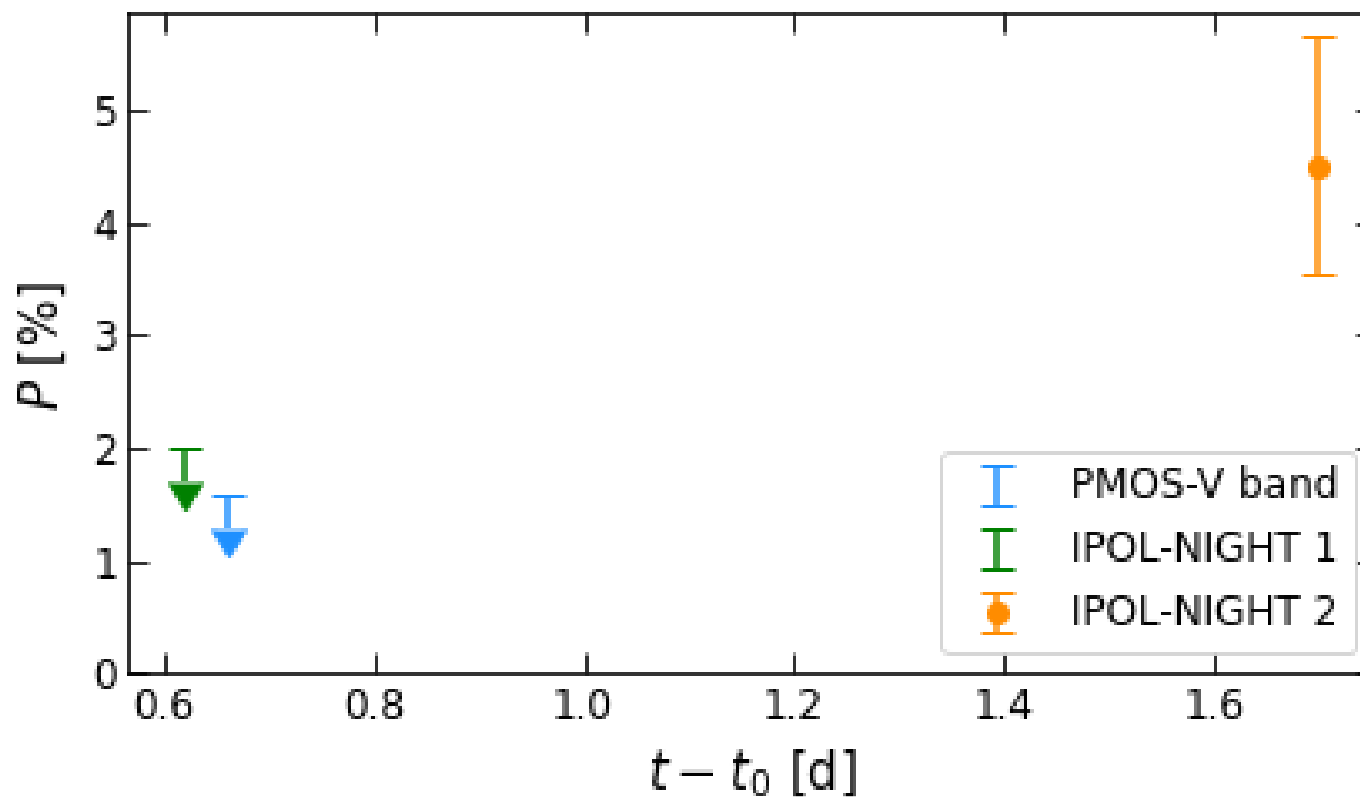
MW-LIKE
 $A_V^{host} = 0.37$
Rossi A. et al (2011)

$$P_{HG} \leq 1.1\%$$

POLARISATION CURVE



$t - t_0$ [d]	MODE	Q/I	U/I	P [%]	θ [°]
0.62	IPOL	-0.0049 ± 0.0050	-0.0008 ± 0.0045	< 1.99	—
0.66	PMOS	0.0031 ± 0.0034	0.0037 ± 0.0038	< 1.57	—
1.70	IPOL	0.0059 ± 0.0101	0.0457 ± 0.0104	$4.49^{+1.16}_{-0.96}$	41.3 ± 6.3



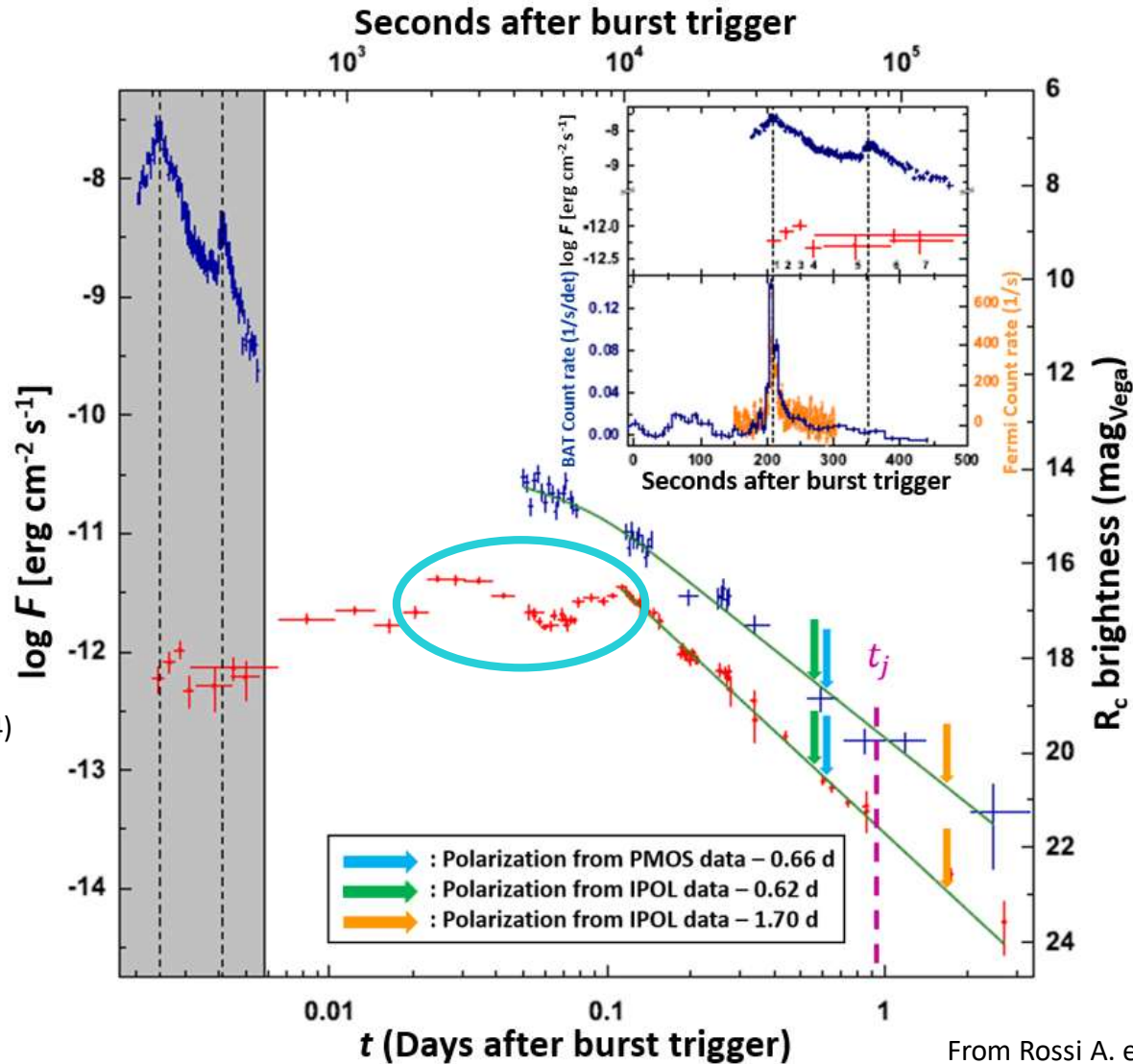
COMPARISON WITH MODELS



LONG-LIVED
ENGINE?

$$t_j = 0.93 \text{ d}$$

Leventis et al (2014)



From Rossi A. et al (2011)

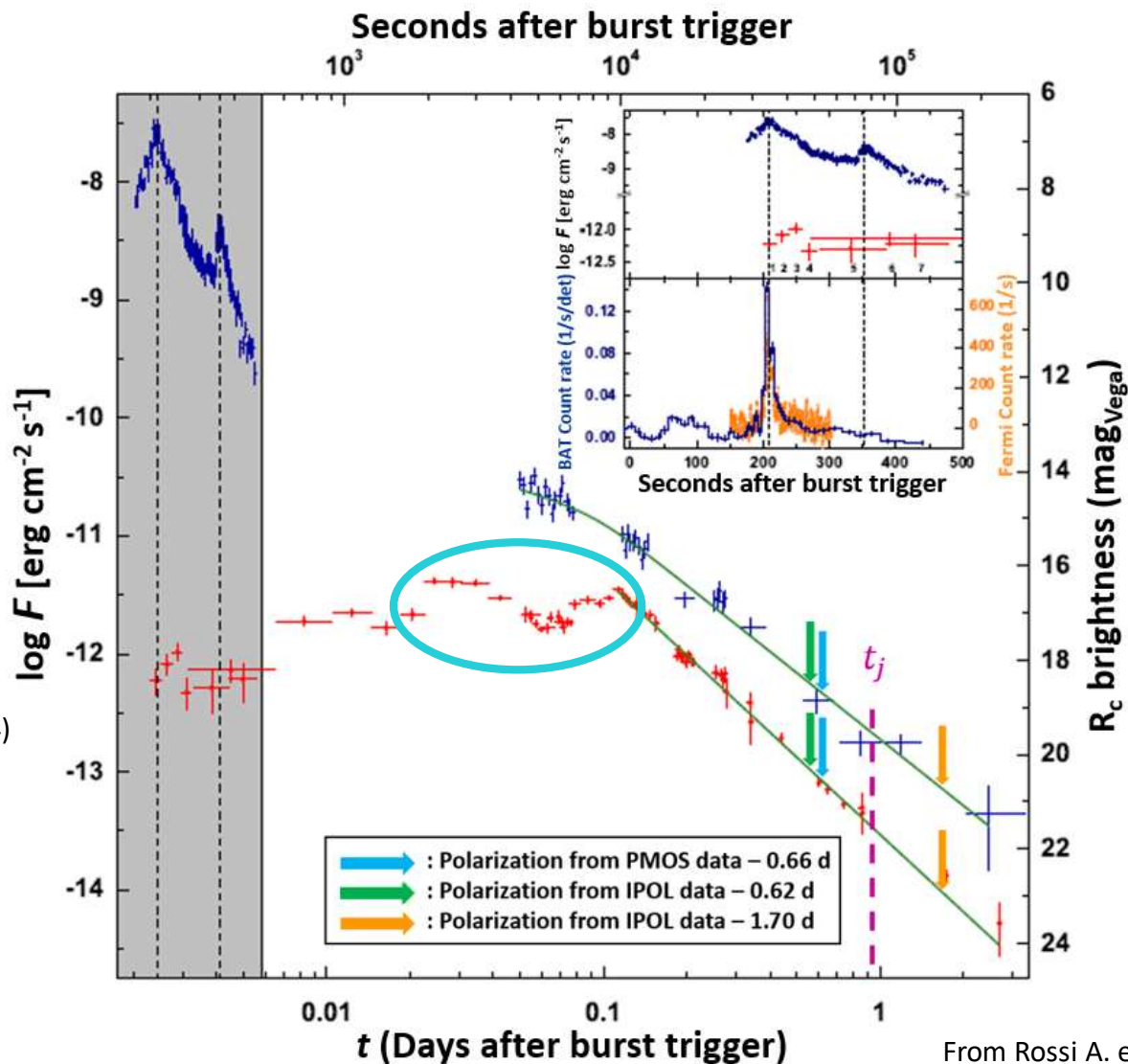
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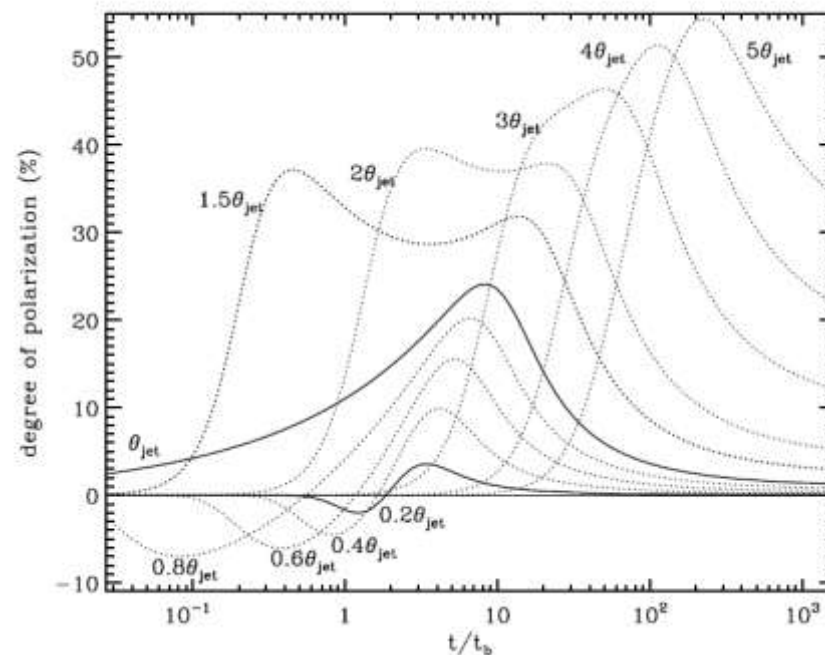
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Leventis et al (2014)



From Rossi A. et al (2011)

EXPECTED POLARISATION CURVES
FOR HOMOGENEOUS JET



From Rossi E. et al (2004)

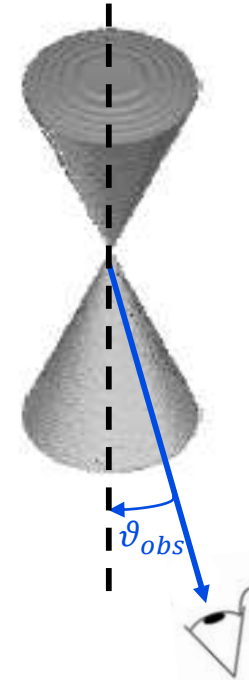
CONCLUSIONS

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Structured and Gaussian jets have been excluded after comparison with models;



CONCLUSIONS

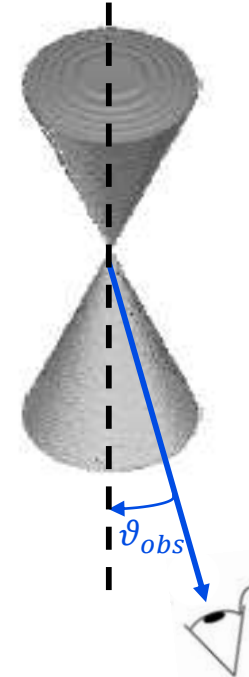
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- Polarisation degree results are consistent with $0.6 < \vartheta_{obs}/\vartheta_{jet} < 0.8$. Polarisation detection coincides with the second rise in the curve;



CONCLUSIONS



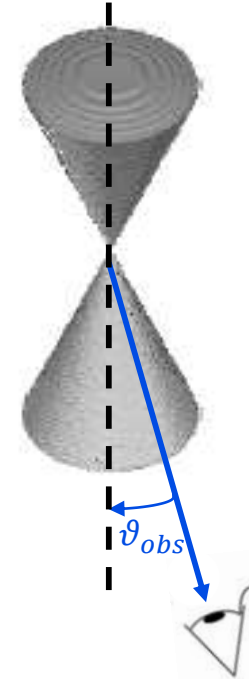
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- The fifth GRB spectro-polarimetric analysis overall, one of the relatively few GRB polarisation detections;



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- The fifth GRB spectro-polarimetric analysis overall, one of the relatively few GRB polarisation detections;
- The same method can be applied to more GRBs to increase the sample and look for similarities and differences.

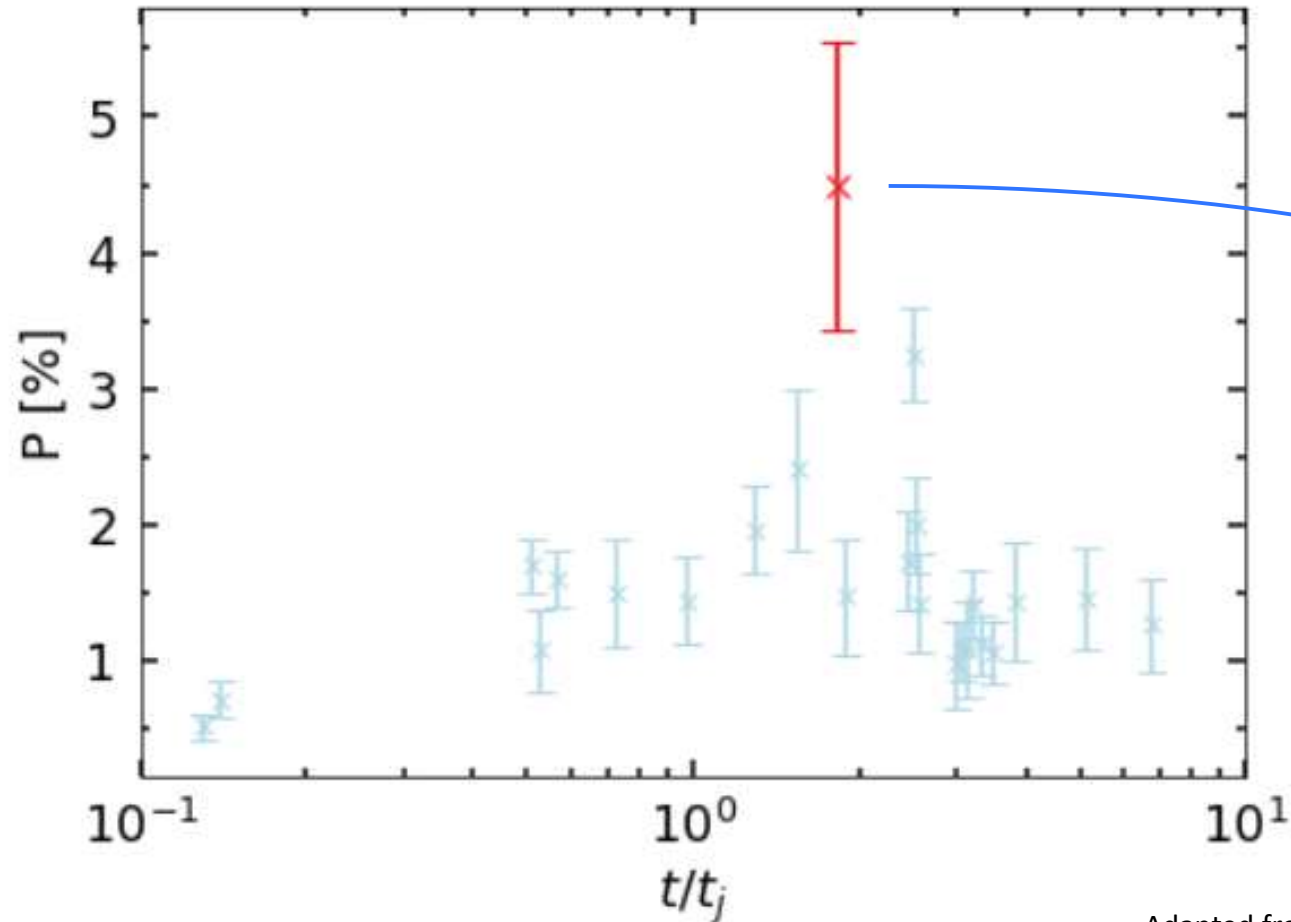




**Thank you for
the attention!**

BACK-UP SLIDES

POLARISATION DETECTION vs. JET BREAK TIME



Largest P of the sample.
1σ errors..

Adapted from Stringer&Lazzati (2020)