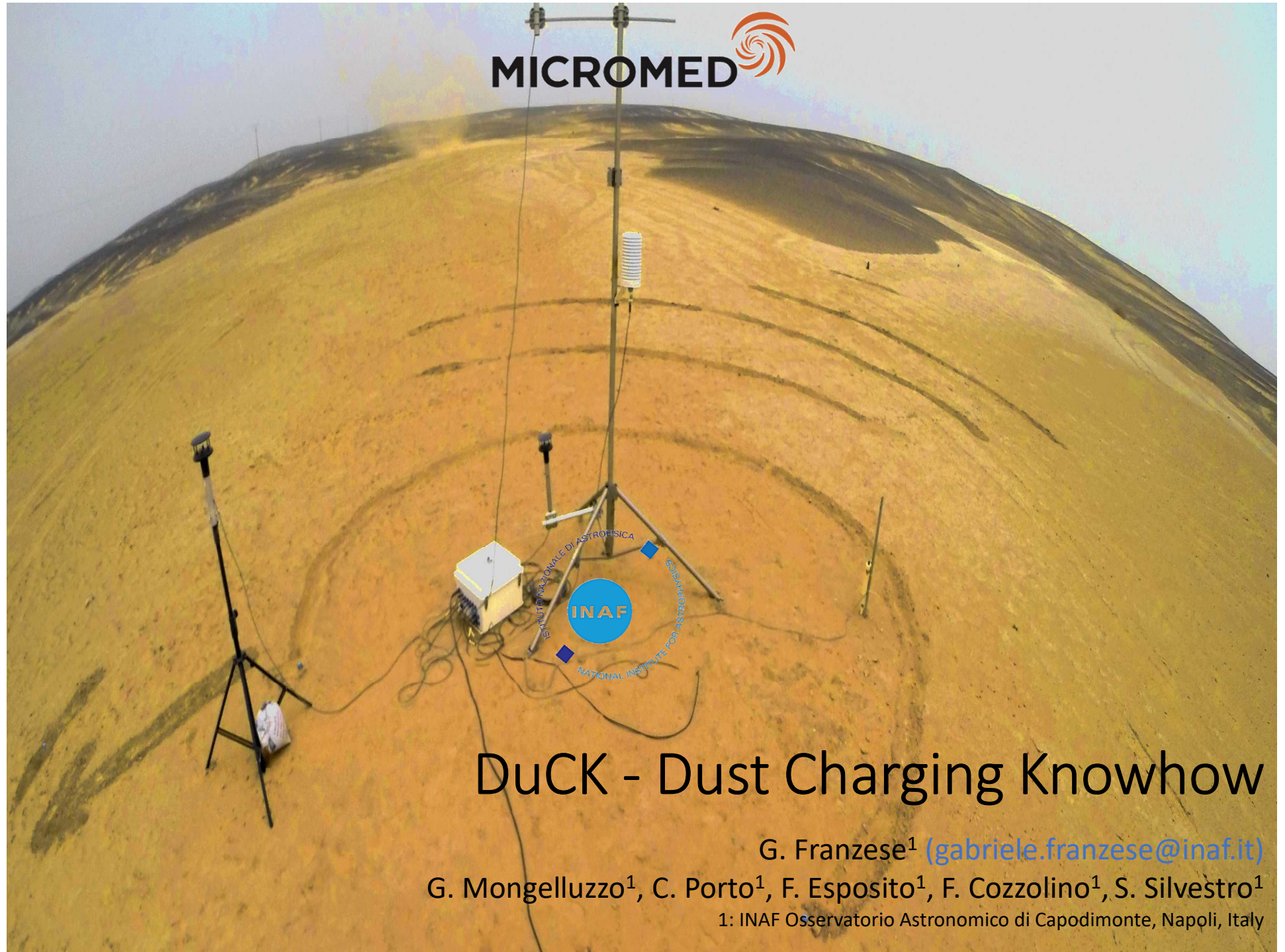


MICROMED 



DuCK - Dust Charging Knowhow

G. Franzese¹ (gabriele.franzese@inaf.it)

G. Mongelluzzo¹, C. Porto¹, F. Esposito¹, F. Cozzolino¹, S. Silvestro¹

¹: INAF Osservatorio Astronomico di Capodimonte, Napoli, Italy

Mineral Dust

from suspension of minerals constituting the soil
grain size $\sim [10^{-2} \mu\text{m} : 65 \mu\text{m}]$

effects on the atmosphere:

it scatters and absorbs visible
and infrared radiation

it cools down the surface
and the low atmosphere,
it heats the upper layers
favouring a stable atmosphere

it inhibits the turbulence
it reduces the vertical exchange of
momentum, decreasing the near surface
wind

it influences the composition
and vertical structure
of the clouds

depending on the relative position of
the dust layer and the cloud coverages

it can:

enhance the
cloud
evaporation

enhance the
cloud
lifetime

it alters the optical proprieties of clouds
the formation of droplet and
the precipitation formation

contribution to atmospheric dust:

$\sim 50\%$

Dust Storms

$\sim 50\%$

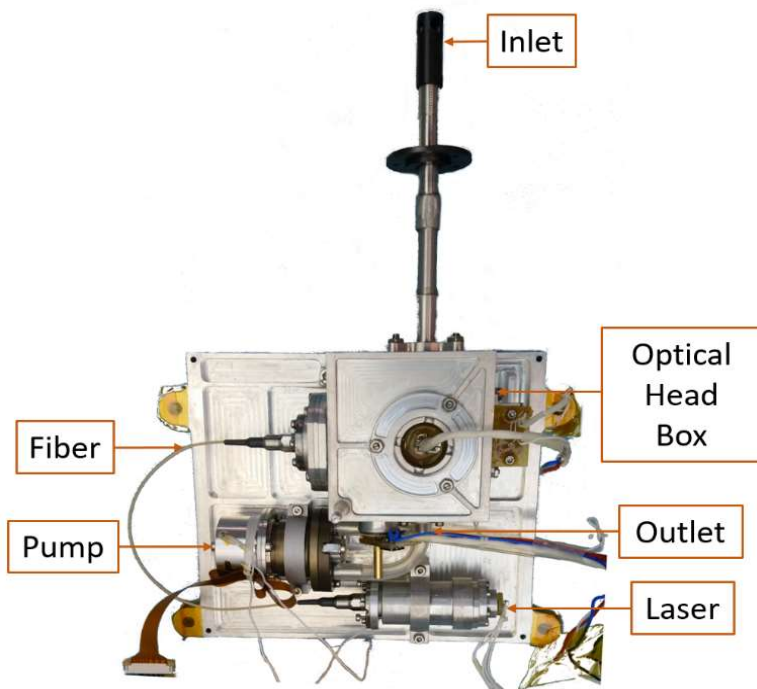
Dust Devils

Dust Complex

Dust Complex:

suite of 4 instruments specifically aim to characterize martian dust lifting phenomena and electrification.

- impact sensor: detection of the saltating sand and charge measurement;
- electric probe for atmospheric E-field;
- electromagnetic activity sensor (EMA) for possible electromagnetic discharges;
- optical particle counter (MicroMED) dust concentration and size distribution.



MicroMED

Primarily scientific purposes:

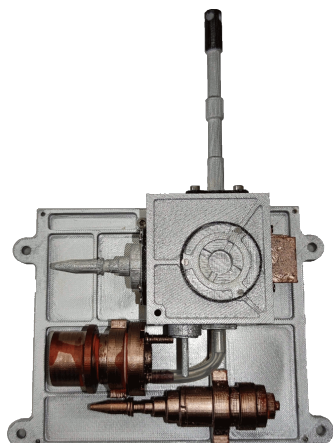
First in situ analysis of the primary lifted dust.

In particular it will measure the lifted dust:

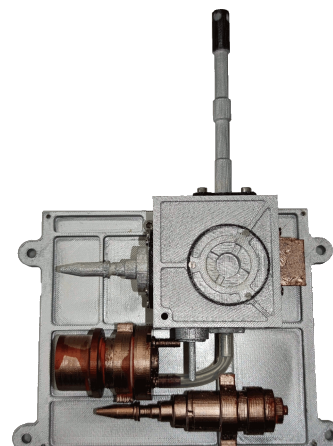
- concentration
- size distribution

MicroMED

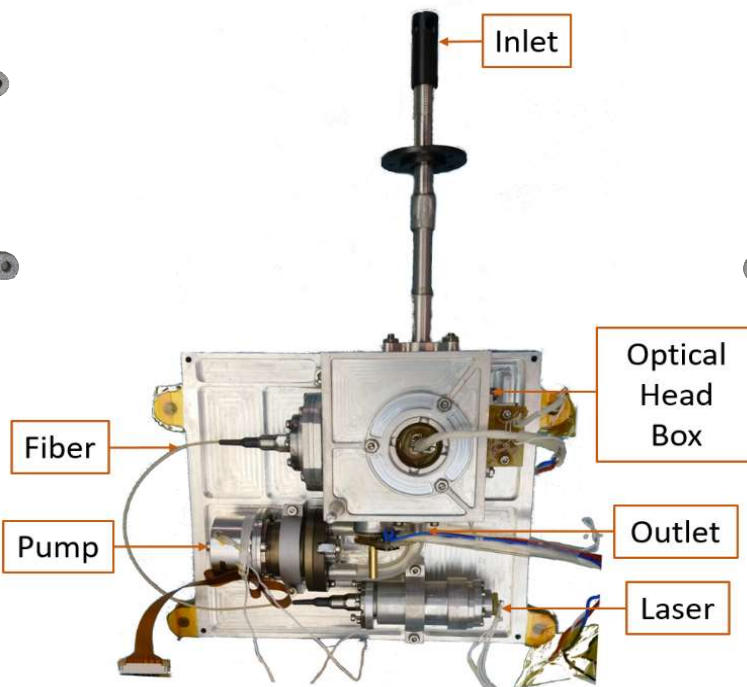
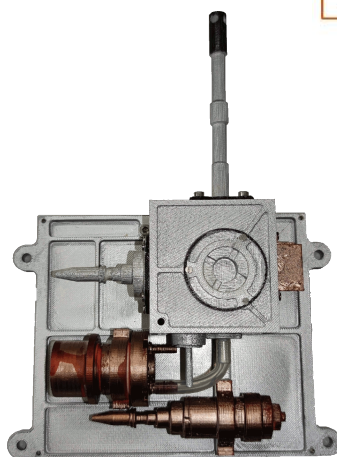
Optical System



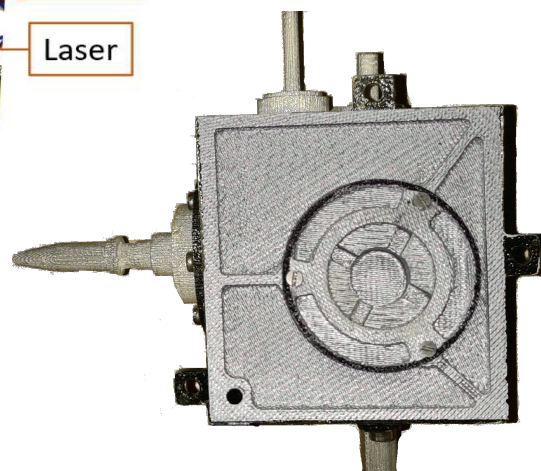
Optical Head



Pump System

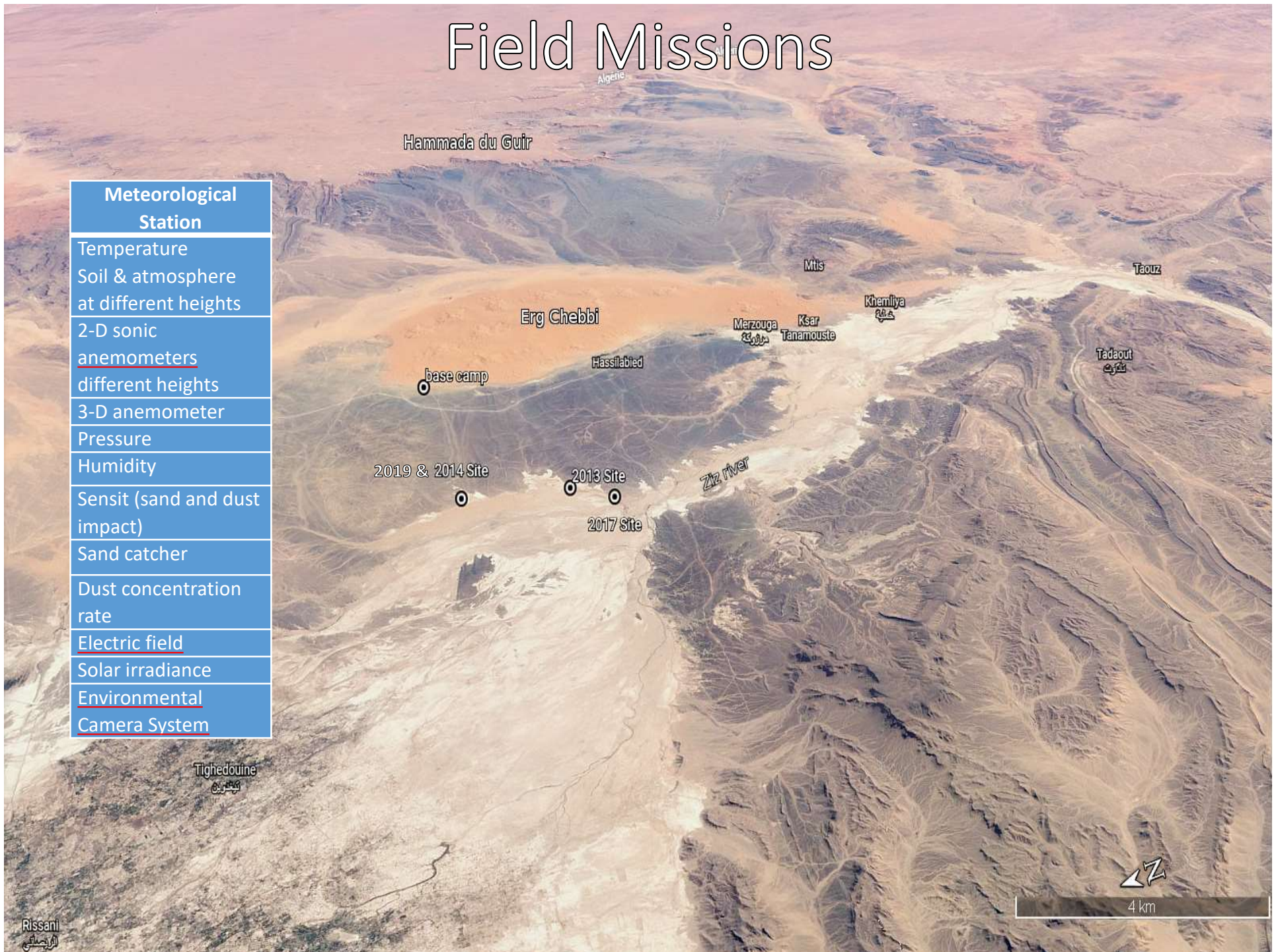


Inside Optical Head



Field Missions

Meteorological Station
Temperature
Soil & atmosphere at different heights
2-D sonic anemometers
different heights
3-D anemometer
Pressure
Humidity
Sensit (sand and dust impact)
Sand catcher
Dust concentration rate
Electric field
Solar irradiance
Environmental Camera System



Collaboration

Key experiences

INAF-OACN
team



Lanzhou University
team

Martian Science

Terrestrial Science

Field measurements

Field & laboratory
measurements

Dust storms
& Devils

Dust storms

Induced E-Field

Induced E-Field

Will to learn

Numerical Simulation of
Saltating process

proper lifting model for terrestrial environment

lifting model for martian environment



Objectives

Current Knowledge

Presence of an atmospheric electric circuit is not confirmed on Mars (lack of dedicated instruments)

Laboratories and simplified study cases suggest how

E-field could lower the lifting threshold however

current saltation models don't consider the E-field contribution

Fail to apply a single electrification model due to the complexity of the process main dependencies on size and composition

Most common configuration is: dipole like

dust cloud negatively charged over a positive saltating sand bed

Our contribution

Modeling in steps of increasing difficulty

I

unidirectional wind (dust storm)

single mode grain population (only sand)

a priori charge-to-mass ratio (empirically derived)

II

vortex wind (dust devils)

bi mode grain population (sand + dust)

parametric description charge-to-mass ratio (tuned on field data)

III

adaptation

from terrestrial → to martian environment

different ionic species
possible electron avalanches and discharges.

Thanks to all

