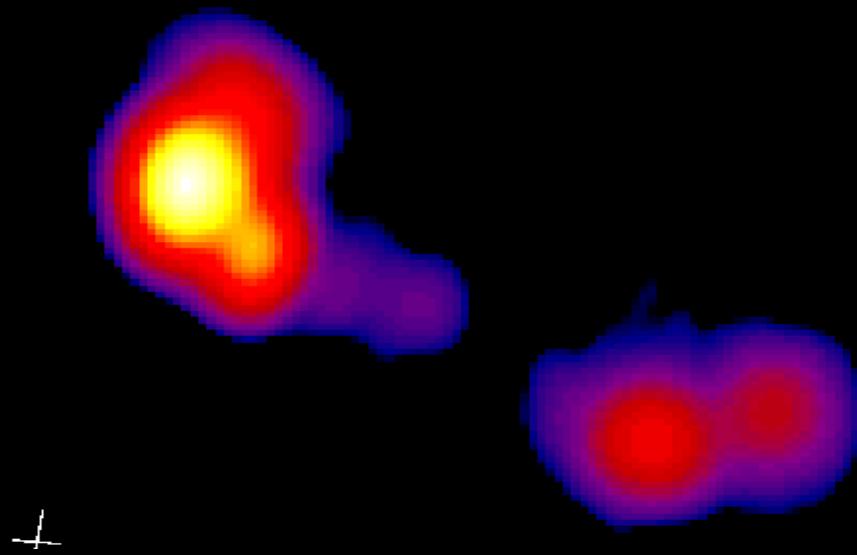
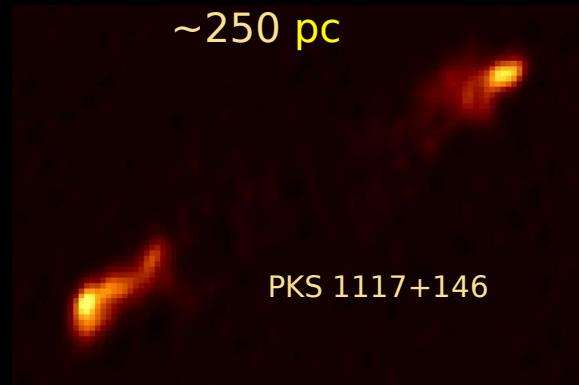


Physics and Evolution of Young Radio Galaxies (YRG)

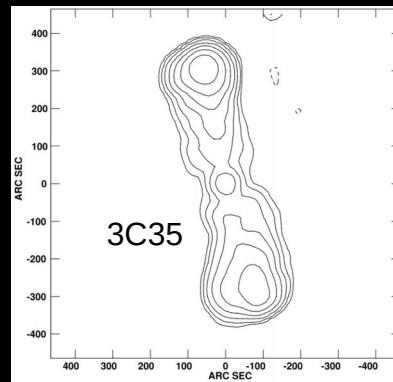
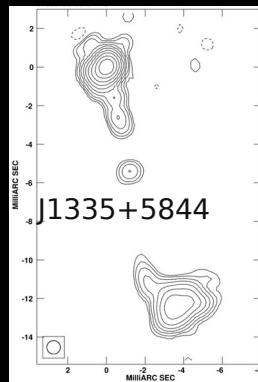
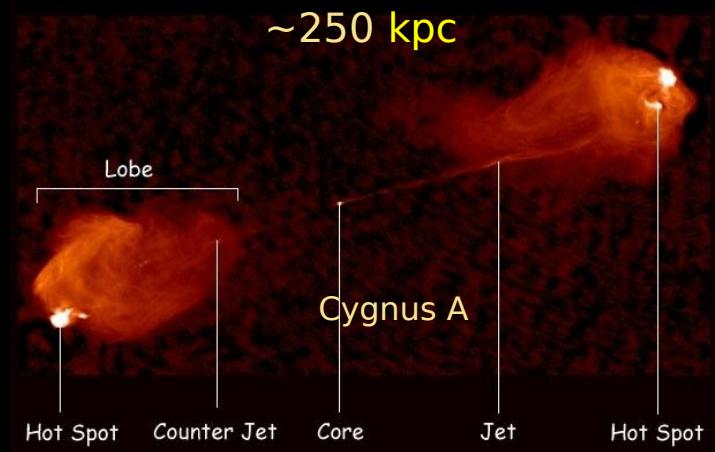


Carlo Stanghellini - audizioni INAF - 10 maggio 2022

YRG



Classical Double



Nomenclatura

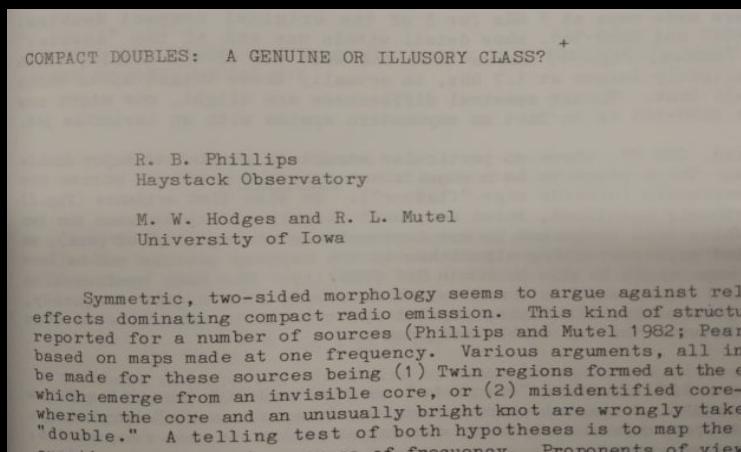
Le YRG consistono di classi di oggetti classificate con acronimi diversi in base ai criteri di selezione

CSS, GPS, HFP, CSO, MSO, etc...

ma sono intrinsecamente raggruppabili sotto un'unica linea di ricerca

Linea di ricerca tradizionale per INAF

- La prima menzione di questa classe di radiosorgenti (symmetric compact double) è di Phillips and Mutel (1982,1983)
- I primi studi (all'interno di IRA, allora IRA-CNR) risalgono agli stessi anni



EVN Observations of Compact Steep Spectrum Radio Sources at 18-CM

Show affiliations

Fanti, C. ; Fanti, R. ; Parma, P. ; Schilizzi, R. T.

The authors consider compact ($\lesssim 2 - 4$ arcsec) radio sources in the 3CR catalogue with normal straight spectrum and low frequency ($v_c \lesssim 200$ MHz) turnover. They have undertaken an observing program with the European VLB Network at 18 cm. Maps of 8 radio sources are presented here.

Publication: VLBI and Compact Radio Sources. Symposium no. 110 held in Bologna, Italy, June 27-July 1, 1983. Edited by R. Fanti, K. Kellermann, and G. Setti. p. 57, 1984

Pub Date: 1984

Bibcode: 1984IAUS..110...57F

Forte interesse dagli anni '90

“Workshop on Compact Steep Spectrum and GHz-Peaked Spectrum radio sources”

1st Dwingeloo 1990

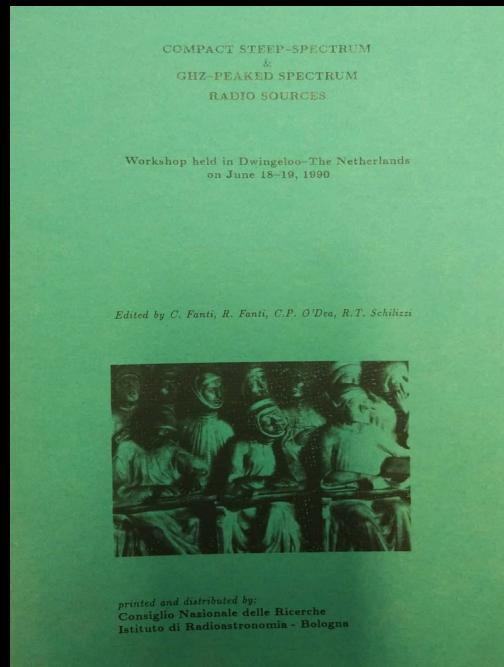
2nd Leiden 1996

3rd Kerastari 2002

4th Riccione 2008

5th Rimini 2015

6th Torun 2021



Forte interesse dagli anni '90

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Leadership INAF

Tutti i lavori significativi su questa linea di ricerca citano articoli di ricercatori INAF.

Nella review di O'Dea and Saikia (2021) sono citati 100 articoli (su ~700) **con ricercatori INAF come primo autore.**

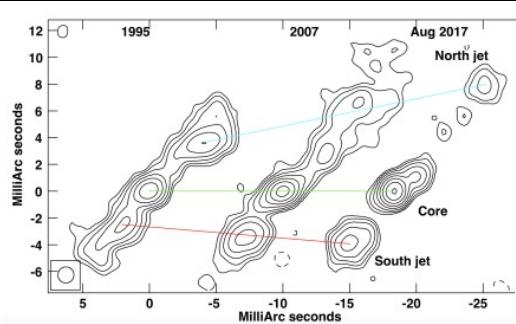
Dopo 40 anni di ricerca su questo tema, perché ha senso continuare ad investirci tempo e risorse?

Il problema di fondo

La natura delle YRG

- *cresceranno fino a diventare doppie classiche? (youth scenario)*
- *scompariranno prima di crescere ed espandersi oltre la galassia ? (short lived)*
- *rimarranno confinate a tempo indeterminato? (frustration scenario)*

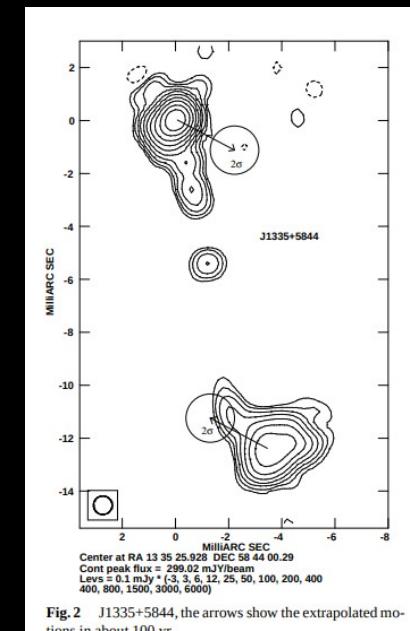
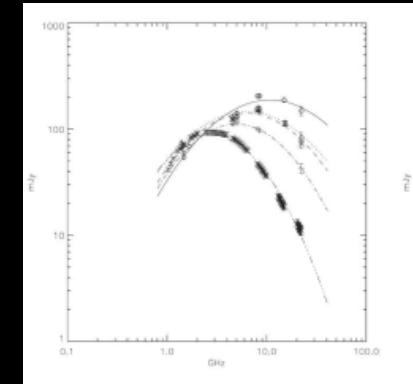
young/expanding



Principe et al. 2020

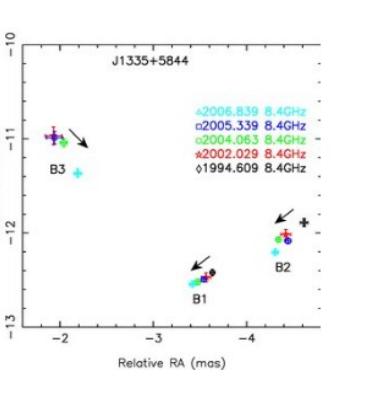
Stato dell'arte (I)

Orienti e Dallacasa 2021



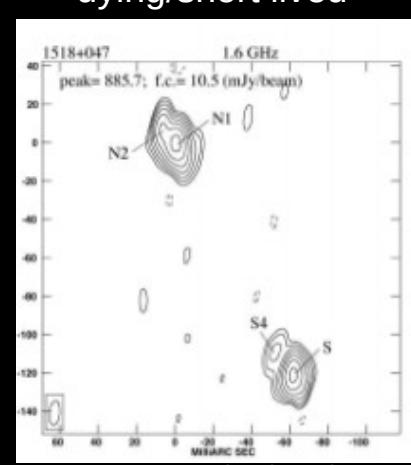
Stanghellini et al. 2009

confined/frustrated



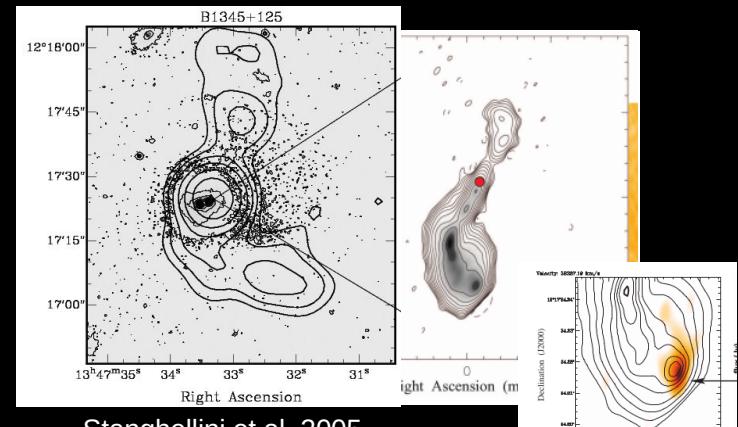
An et al. 2012

dying/short lived



Orienti et al. 2009

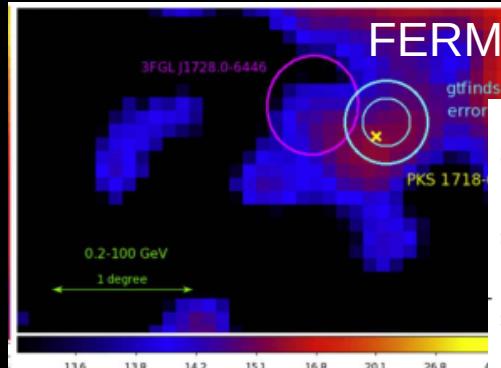
restarted



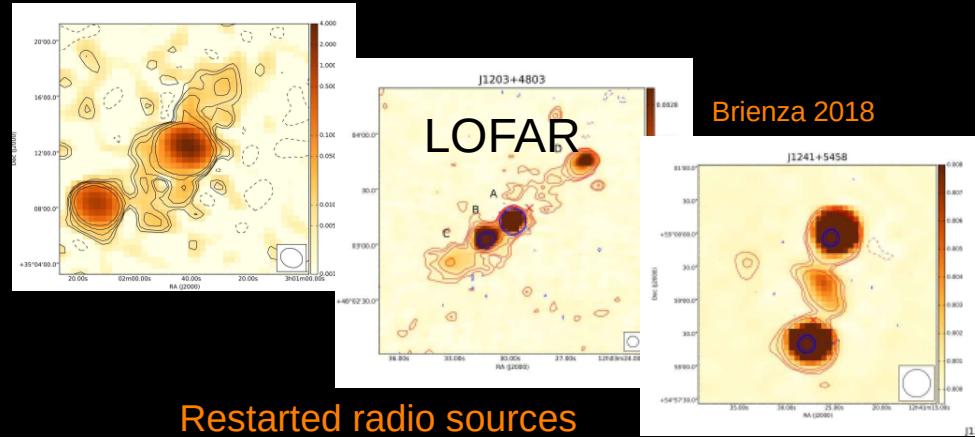
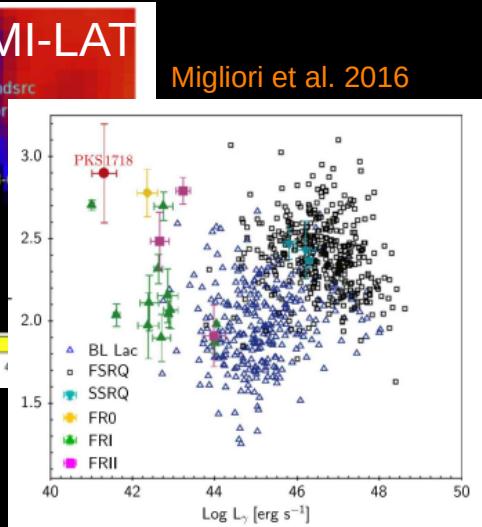
Stanghellini et al. 2005

Interaction/outflow

Morganti et al. 2013



First detection in γ -ray



Stato dell'arte (II)

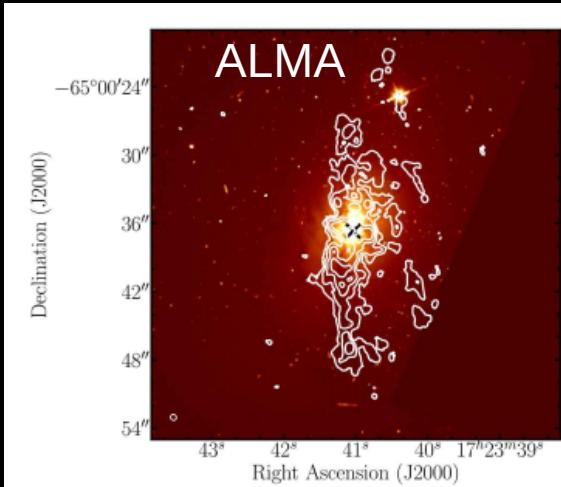


Fig. 2. Hubble Space Telescope WFPC2 image of PKS B1718-649 with

Molecular gas fuelling the AGN

Maccagni et al. 2018

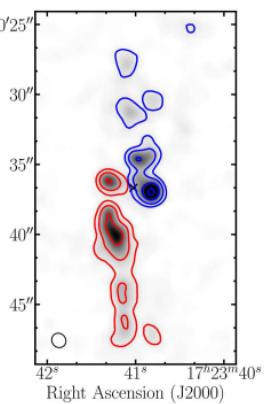


Fig. 2.

Fig. 2.

YRG at High-Energies

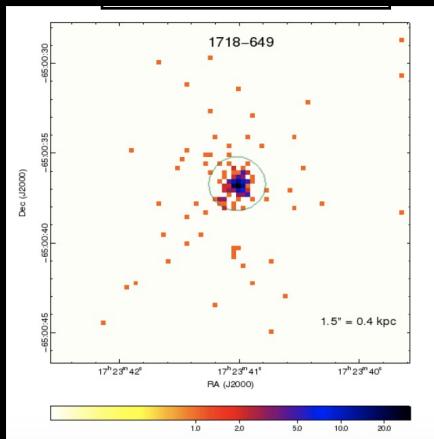
X-rays: sample studies with snapshots + dedicated studies with deep, broadband observations

- nature of the high-energy emission (accretion? young radio structures? Siemiginowska+'16,Migliori+'14)
- probe of the environment (obscuration, Ostorero+'10'17, NuStar observations in Bronzini et al. in prep.)
- young sources in hard X-ray selected giant RGs (Bruni+'19,'20)

Gamma-ray studies with Fermi:

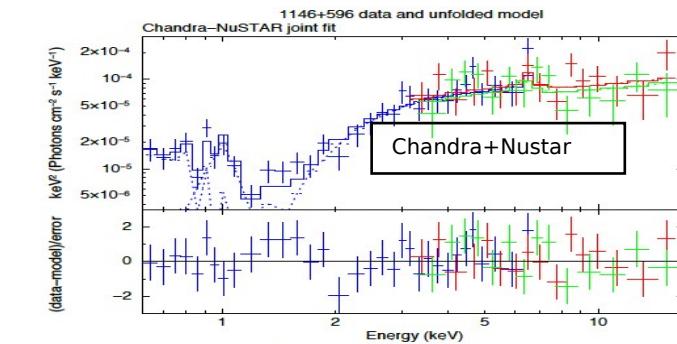
- individual studies (Principe+'20) + CSO in the 4LAC
- first stacking analysis for a sizeable YRS sample with 11 years of Fermi data (Principe+'21)

XMM/Chandra
snapshots



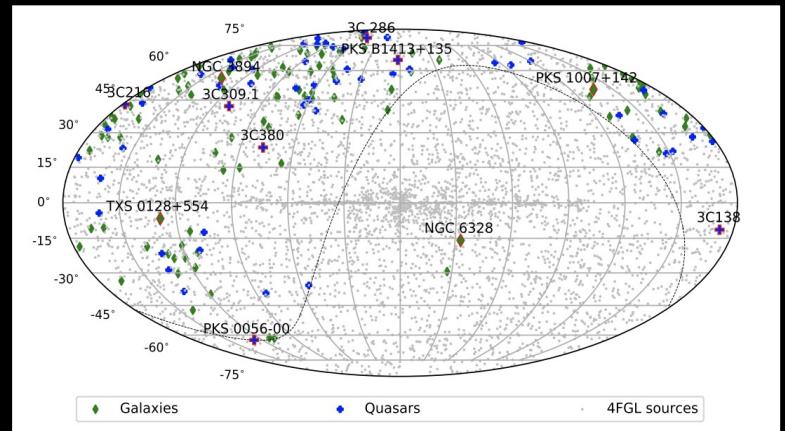
Siemiginowska+16
Sobolewska+19

CSO 1146+596 broadband X-ray spectrum



Bronzini+in prep

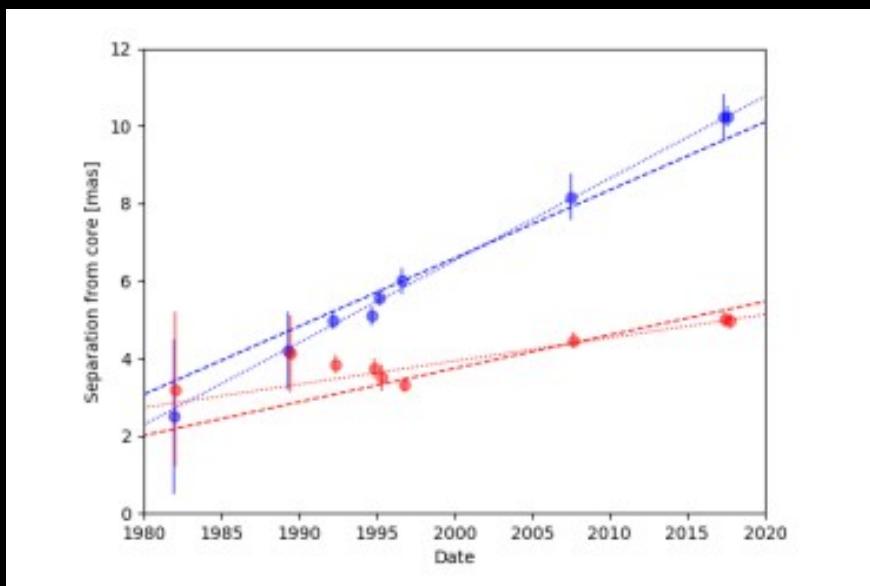
Sky map of young radio sources & 4FGL sources



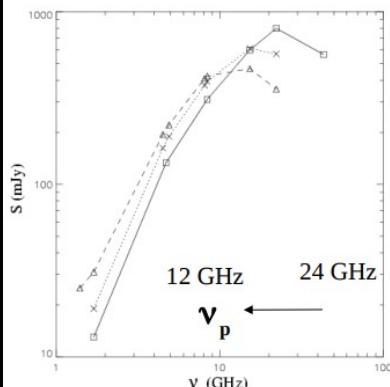
Principe+'21

Direzione futura della ricerca (I)

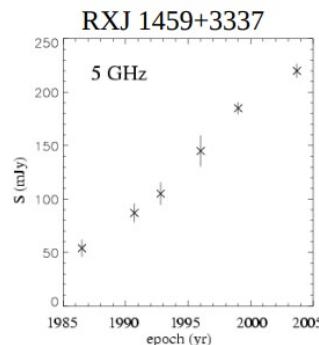
Misurare moti propri o espansione adiabatica può richiedere decenni



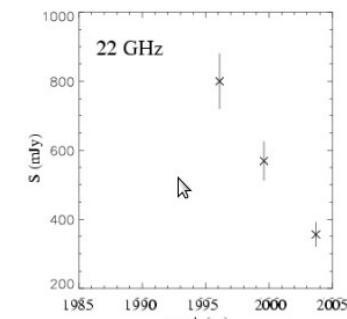
Variability due to
adiabatic expansion of
very young objects



Edge+96, Orienti+08



Optically-thick



Optically-thin



Il gravitational lensing può essere di aiuto - lensing magnification ----> tempi scala ridotti

Direzione futura della ricerca (II)

Natura ed evoluzione delle YRG:

- contributo relativo di sorgenti in espansione, frustrate, short lived
 - distribuzione statistica di dimensioni e moti propri di un campione ampio
 - meccanismo responsabile del turnover (FFS vs SSA) – osservazioni profonde a bassa frequenza
- radiosorgenti di bassa luminosità - short lived? - contributo delle FR0

AGN feedback:

- presenza di outflows
- interazione con l'ambiente
- SFR e contenuto molecolare

Simulazioni numeriche

- propagazione della radiosorgente compatta e confronto con i dati osservativi

Direzione futura della ricerca (III)

I nuovi strumenti

ASKAP – Norris et al 2011; Allison et al 2016

LOFAR – Snellen et al. 2009; Shimwell et al 2019

SKA – Falcke et al. 2004; Kapinska et al. 2015; Afonso et al. 2015; Norris et al. 2011

ngVLA – Patil et al. 2018

Nuove survey radio profonde sono destinate a fornire campioni estesi di YRG a bassa luminosita` e alto z.

Il team INAF

Competenze specifiche: radio, alte energie, simulazioni numeriche

- 16 ricercatori INAF di 5 strutture (IRA,OAS, IAPS, OAC, OATo) – 1.3 fte + 0.6 extra
- 9 ricercatori associati INAF (1.0 fte)
 - inclusi 6 TD

Ranieri Baldi, Marisa Brienza, Gabriele Bruni, Alessandro Capetti, Pietro Cassaro, Annalisa Celotti, Daniele Dallacasa, Filippo D'Ammando, Gabriele Giovannini, Marcello Giroletti, Paola Grandi, Rocco Lico, Elisabetta Liuzzo, Filippo Maccagni, Francesca Panessa, Karl Heinz Mack, Giulia Migliori, Matteo Murgia, Monica Orienti, Luisa Ostorero, Giacomo Principe, Paola Rossi, Cristiana Spingola, Carlo Stanghellini, Eleonora Torresi

Progetti in corso (I)

- ES101A/B (Old radio activity in two young radio galaxies) - **EVN+eMerlin** at 18 cm - 12+8 hours, **PI Stanghellini** - observed on 1/6 Mar 2022
- BS309 (Diffuse emission at mas scale in a young radio galaxy: relic or shocked winds?) - **VLBA** at 90/50 cm – **PI Stanghellini** - 8 hours to be scheduled in first semester 2022
- EM157 (What drives an outflow? The radio emission of MaNGA 1-166919 at mas scales) - **EVN+eMerlin** at 18 cm - 5 hours, **PI Migliori** - observed on 3 Nov 2021
- EG113A/B (Unveiling the pc-scale structure of the gamma-ray young radio galaxy PKS1007+142) - **EVN/eMerlin** at 18 cm - 2x10 hours, **PI Giroletti** - observed on 2/10 Mar 2022
- EAVN21B-231 - (Structural evolution in two compact radio sources) - East Asian VLBI Network (**EAVN**) at 1.3 cm – **PI Giroletti** - 20 hours observed in November 2021 and being correlated
- BO064 (VLBA observations of low-power jets in compact symmetric objects)- **VLBA** at 6 cm - 37 hours - **PI Orienti** - observed in 2021
- EG111 **eMerlin+EVN** at 6 cm – 24 hours – **PI Giovannini** - scheduled on 14-15-16/06/2022

Progetti in corso (II)

Non solo VLBI

- 19A-082 (Observations of FR0 as young radio sources candidates) **JVLA** at 6 and 18 cm- 9 observing blocks da agosto ad ottobre 2019 – **PI Baldi**
- 35-22 (Probing restarting activity in giant radio galaxies: young cores and duty cycle) – **Effelsberg** – 20 hours – **PI Bruni** – proposal approved

Non solo Radio

- ID: 23700588 (Extended X-ray emission around a young AGN radio jet) – **CHANDRA** 60 ksec PI: Sobolewska (CfA); **Co-I INAF: Migliori, Ostorero**
- ID 24700051 (First CHANDRA observations of 6 highly magnified lensed sources in search for binary and offset agn in the early Universe) **CHANDRA – PI Spingola - submitted**

Criticità (I)

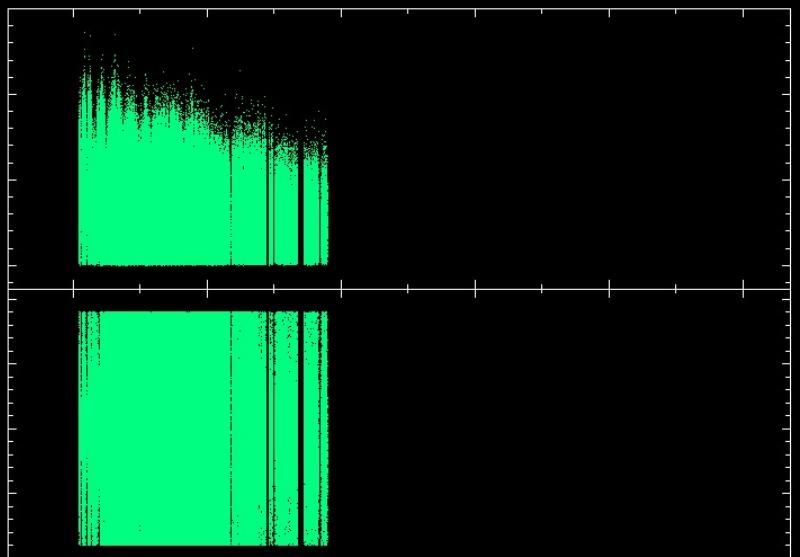
Rischio di perdere la leadership acquisita in seguito a:

- Perdita di competenze importanti (pensionamenti) e conseguente necessità di adeguato turnover
- Mancanza di fondi per progetti specifici su YRG

Soluzione: YRA for YRG

- 6 ricercatori TD su questa linea di ricerca (tesi, dottorato, post doc)
- 4 di questi a FTE 0 perché pagati ora su altri progetti
- 2 di questi sono post doc all'estero

DIFMAP

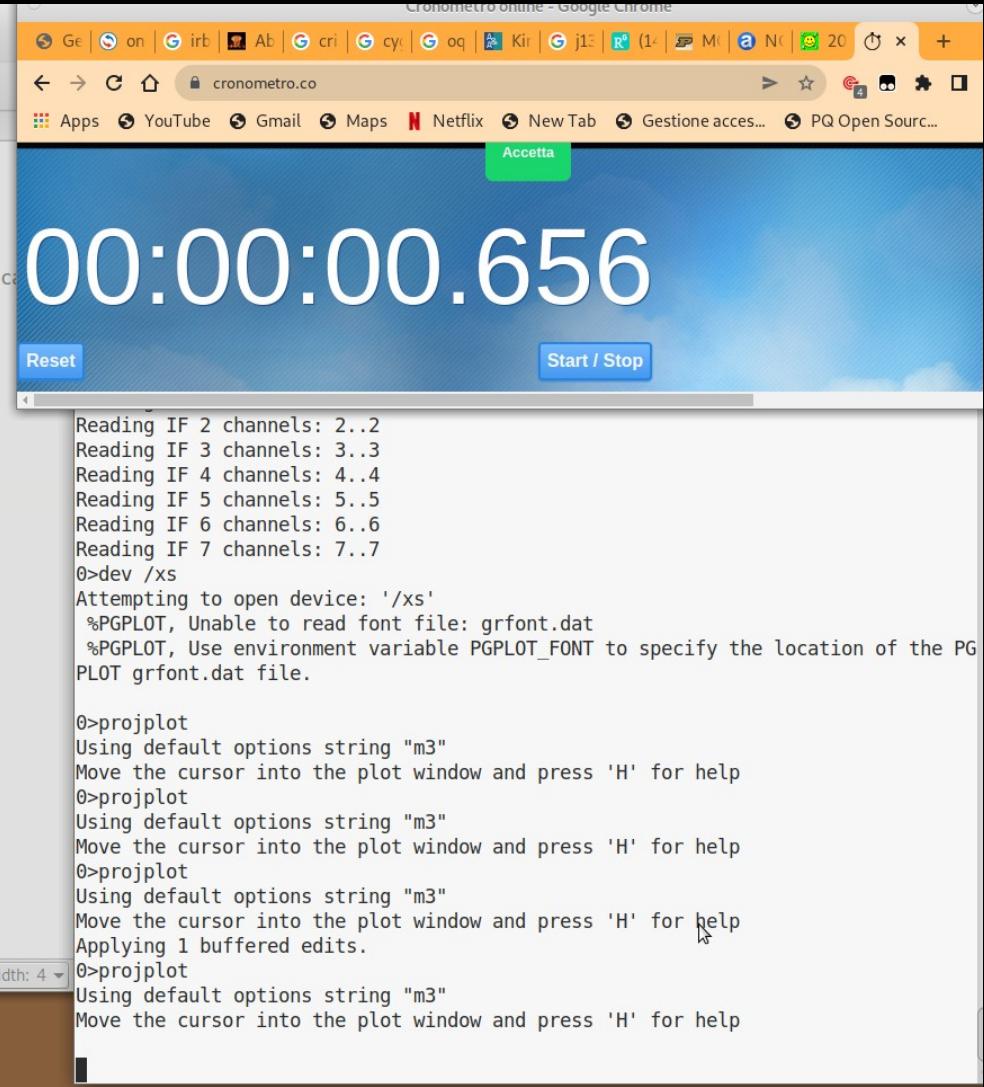


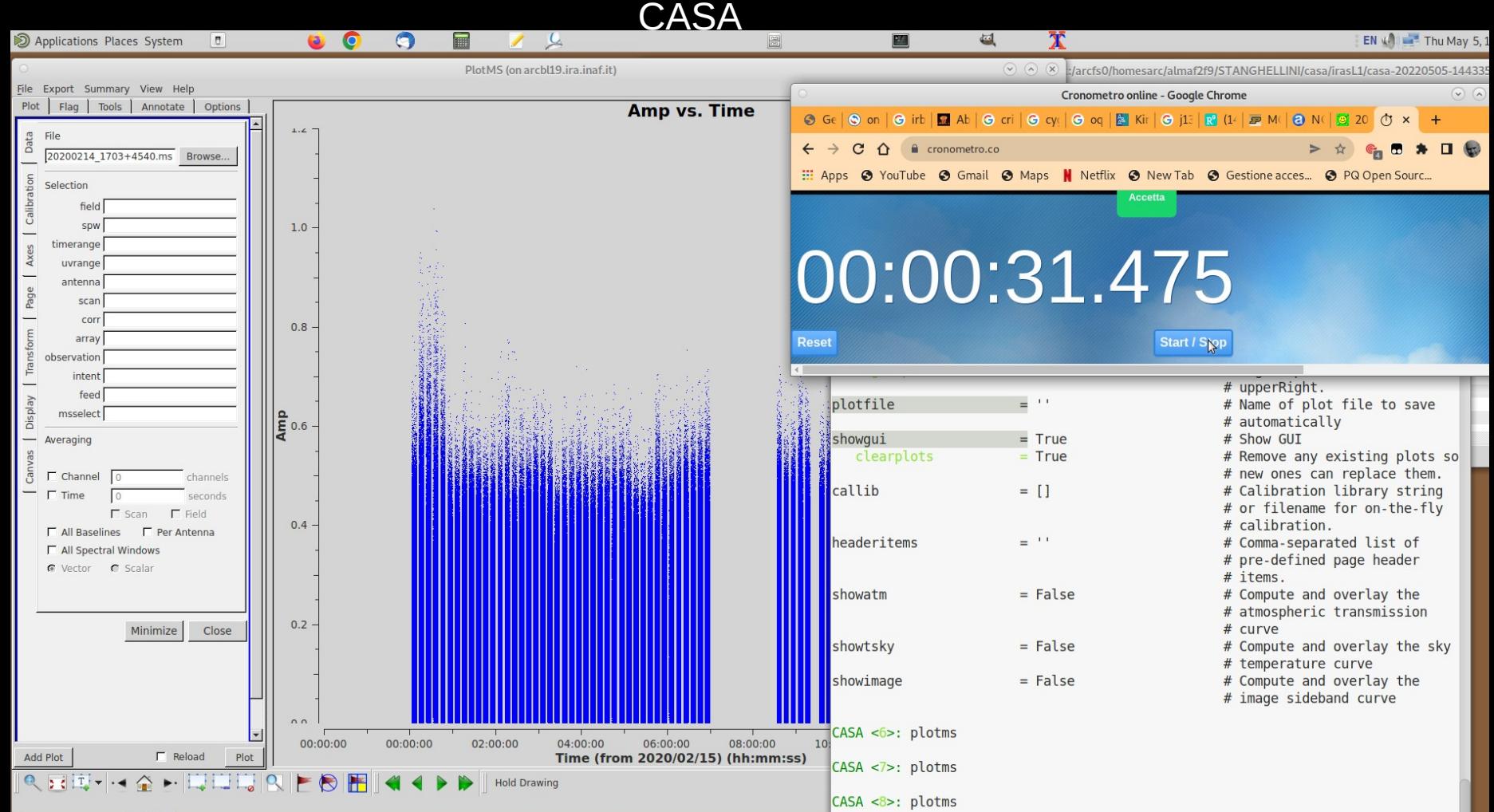
lavoro,
ta

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apr26.pdf

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Per affrontare le criticit

- GTO / GO Grant (normal) - coord Carlo Stanghellini - *Discovering the origin of parsec scale radio emission in young radio galaxies by VLBI observations* - 19 keu per hardware e missioni
- GTO / GO Grant (normal) collegato anche ad altre schede - coord Monica Orienti - *VLBA observations of low-power jets in compact symmetric objects* - 49 keu per Post Doc position, hardware e missioni