

PRISMA DAYS 2022
TORINO, NOVEMBER 26

FROM THE OBSERVATIONS OF
METEORS
TO THE ORIGIN OF
METEOROIDS



MÁRIA HAJDUKOVÁ

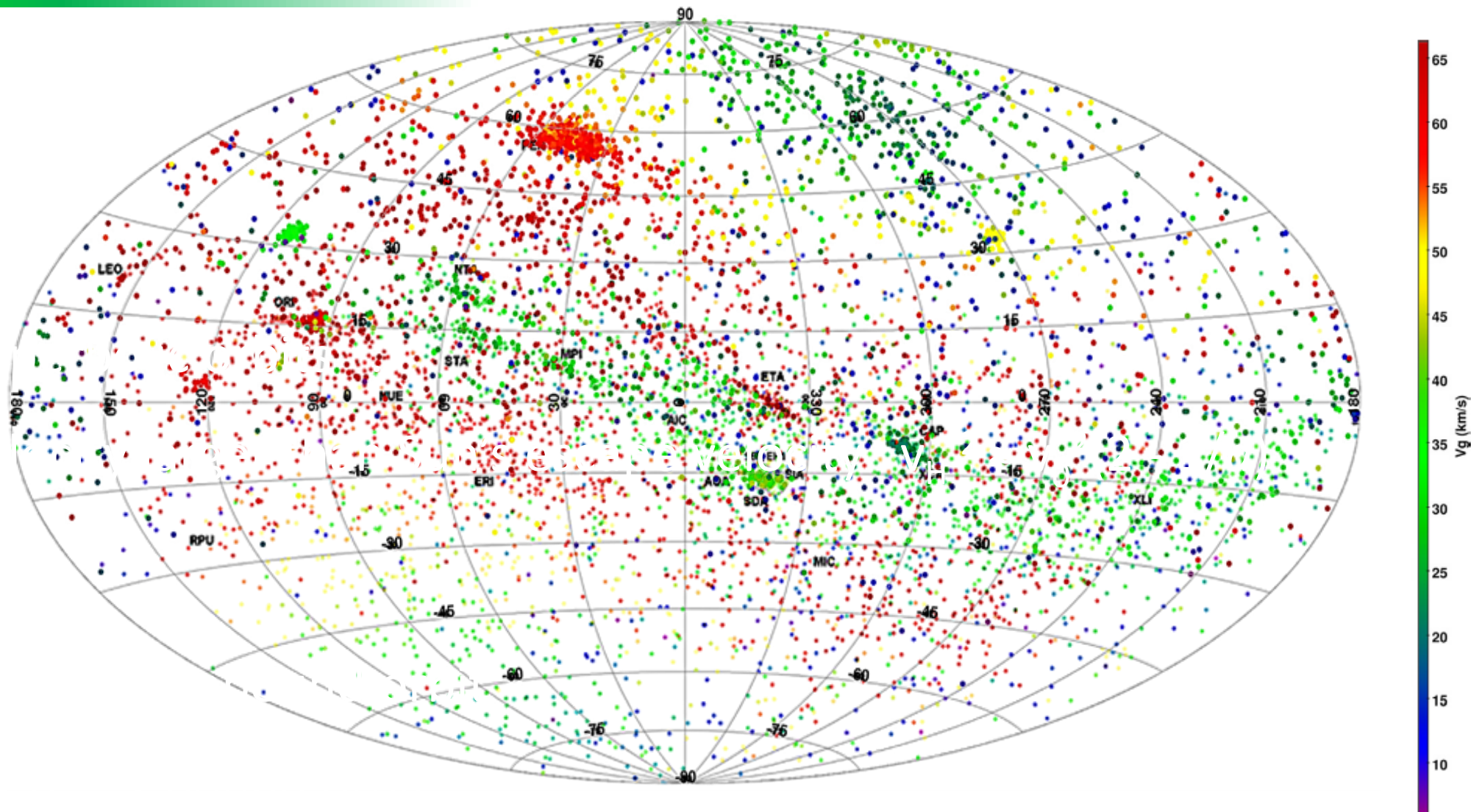
ASTRONOMICAL INSTITUTE
SLOVAK ACADEMY OF SCIENCES
BRATISLAVA, SLOVAKIA



v_{inf} and α, δ

METEOR RADIANTS

V_G



Tóth et al., AMOS all-sky video system, Toth, J. et al., PSS, 118, 102, 2015; Toth, J. et al, EPSC, Berlin, 2018

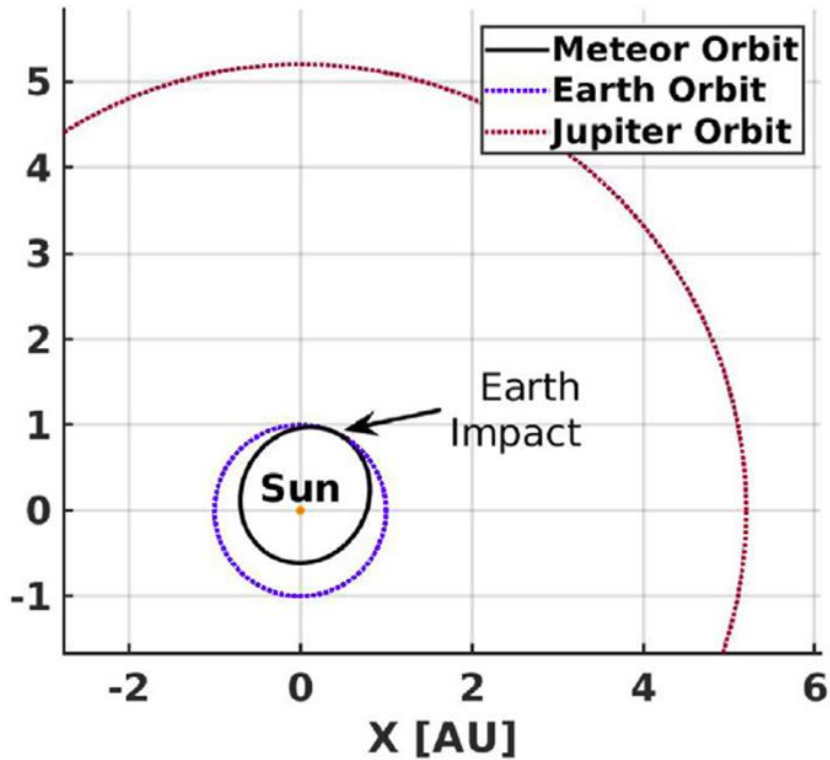


$a, e, q, \omega, \Omega, i, T$

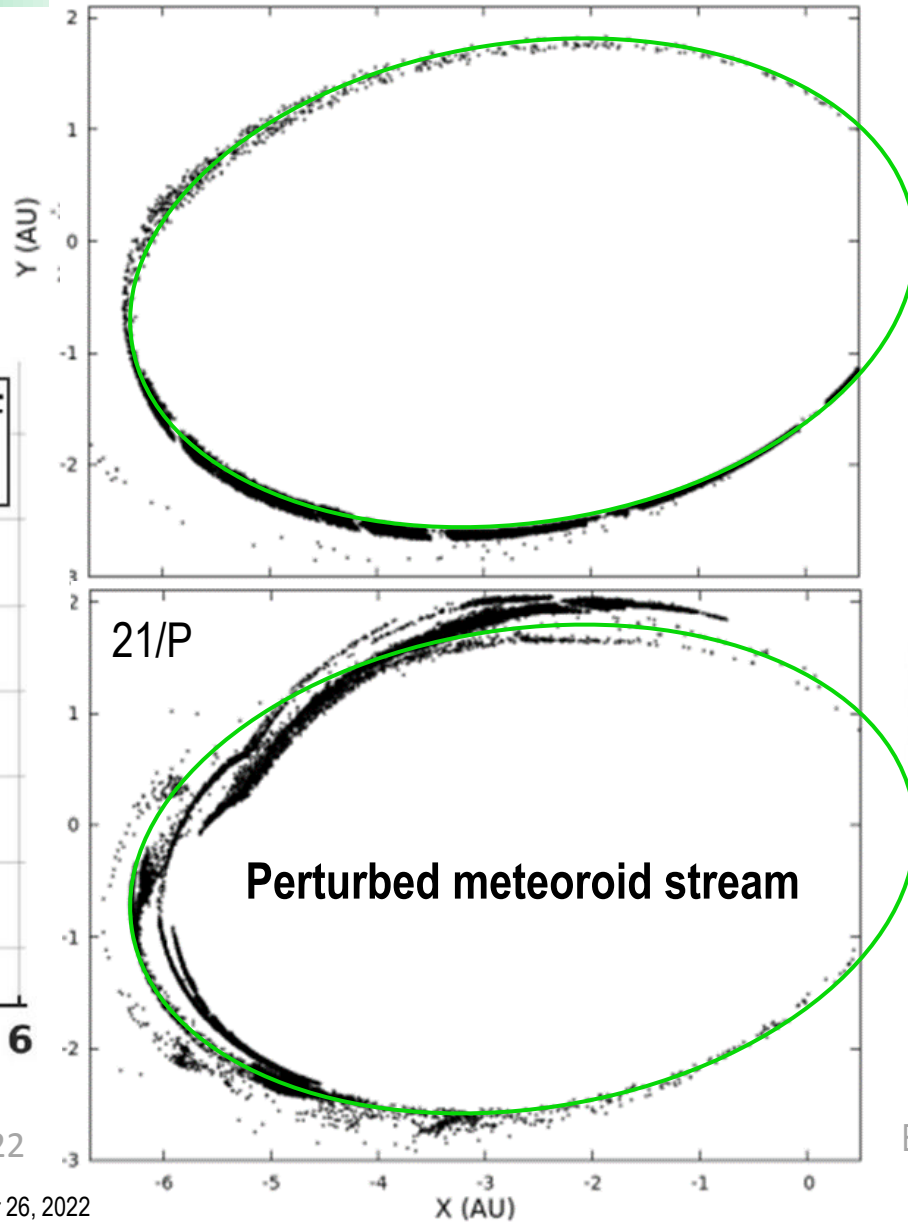
METEOROID ORBITS

ELLIPTIC ORBIT

Local meteoroids, meteoroid streams

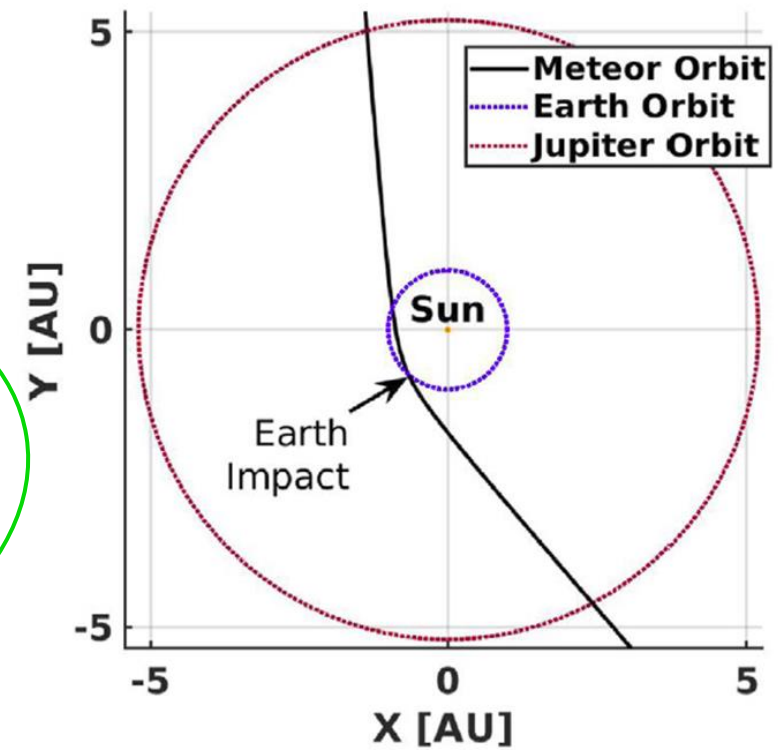


Blanchard et al., Icarus 386, 115144, 2022



HYPERBOLIC ORBIT

Interstellar meteoroids

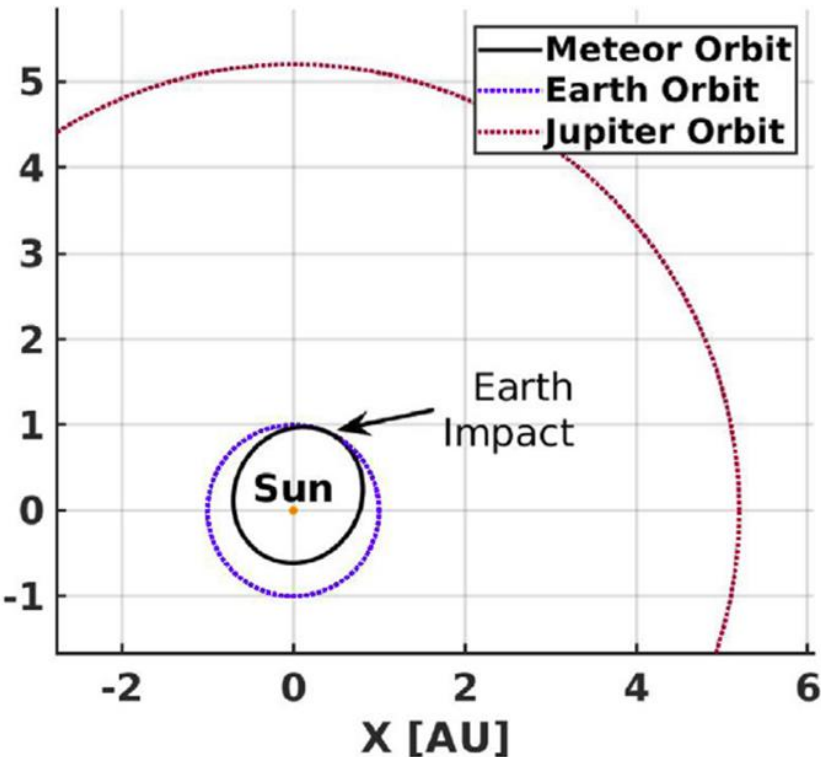


Egal, A., PSS, 185, 104895, 2020

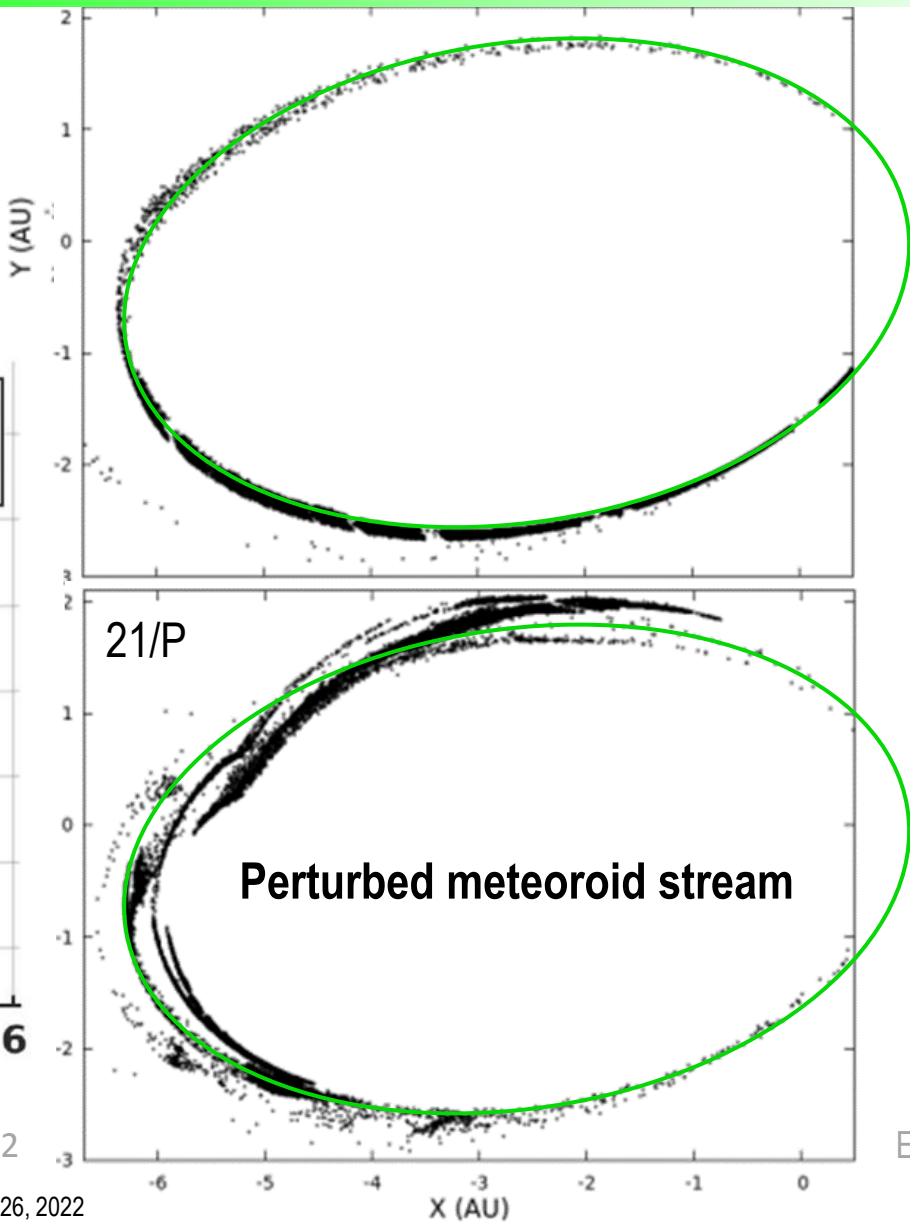
THE ORIGIN OF METEOROIDS

ELLIPTIC ORBIT

Local meteoroids, meteoroid streams



Blanchard et al., Icarus 386, 115144, 2022



Egal, A., PSS, 185, 104895, 2020

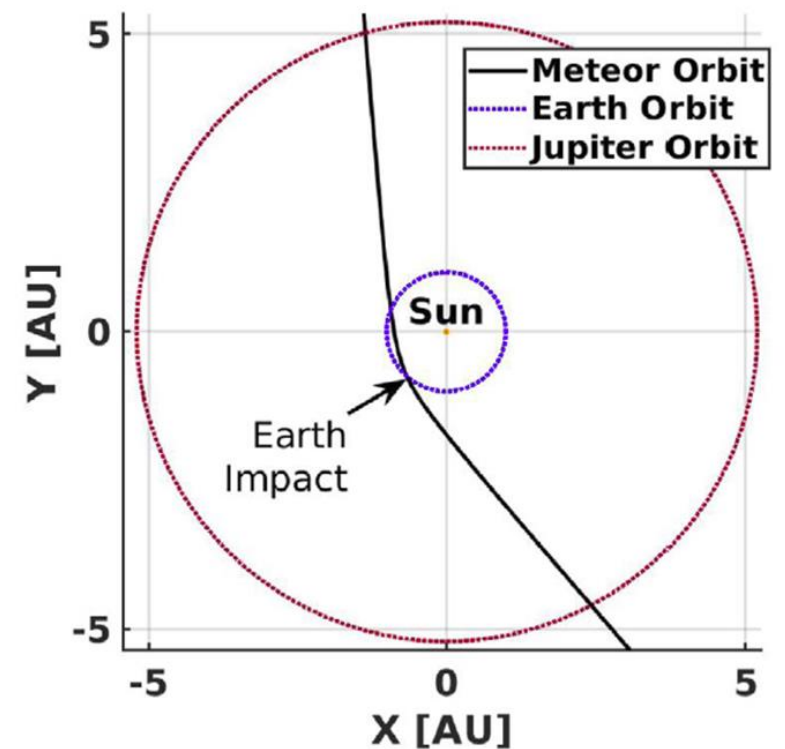
- in the solar system:
 $1/a > 0$ (bound orbit)
- parent bodies –
comets, asteroids
- D-criteria for orbital
similarity
Southworth, Hawkins, 1963
Drummond, 1980
Jopek, 1993
Valsecchi, Jopek, Froschle, 1999
Kholshchikov et al, 2016
- Modeling

THE ORIGIN OF METEOROIDS

- outside the solar system: $1/a < 0$ (unbound orbit)
- LIC, exoplanetary systems
- velocities larger than Sun's escape velocity
- expected excess velocities
 $v_H \sim 49 \text{ km/s}$ (for $v_a = 25 \text{ km/s}$)

HYPERBOLIC ORBIT

Interstellar meteoroids



Blanchard et al., Icarus 386, 115144, 2022

PROBLEMS

ELLIPTIC ORBIT



too many minor showers reported,
which do not exist (about 400 in the
last 15 years)

HYPERBOLIC ORBIT



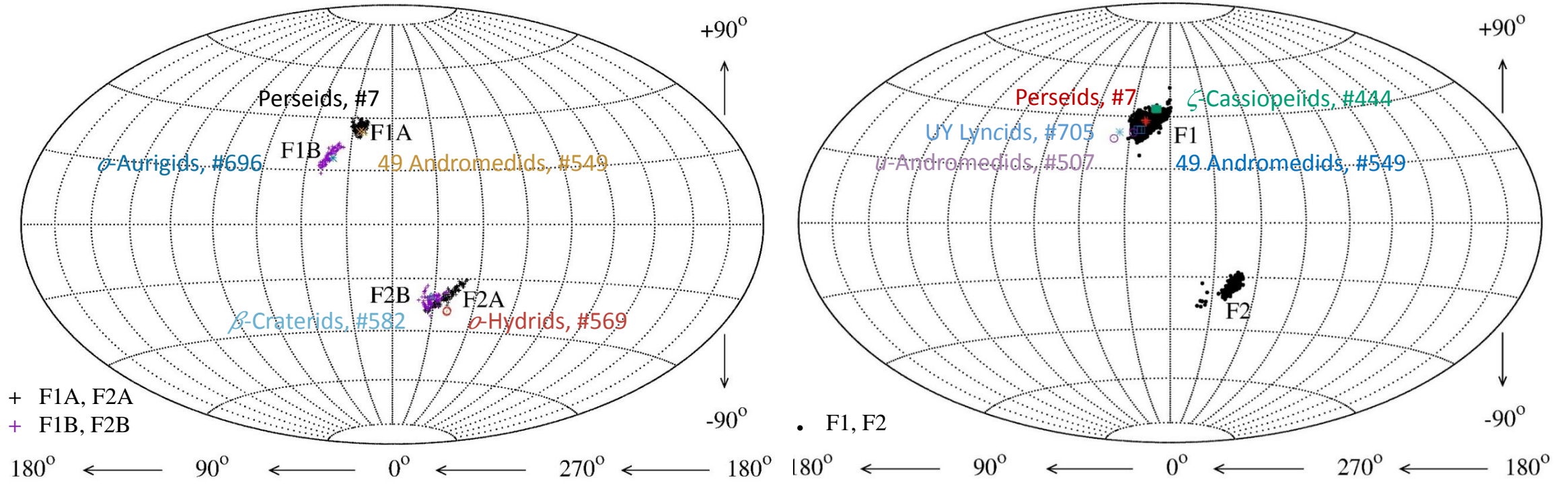
too many hyperbolic meteors observed,
which are not interstellar (about 10%)

PROBLEMS

MINOR METEOR SHOWERS

Comet 109P/Swift-Tuttle

Models based on the nominal and cloned comet orbit of 109P



predicted showers

mean radiants of the IAU MDC showers

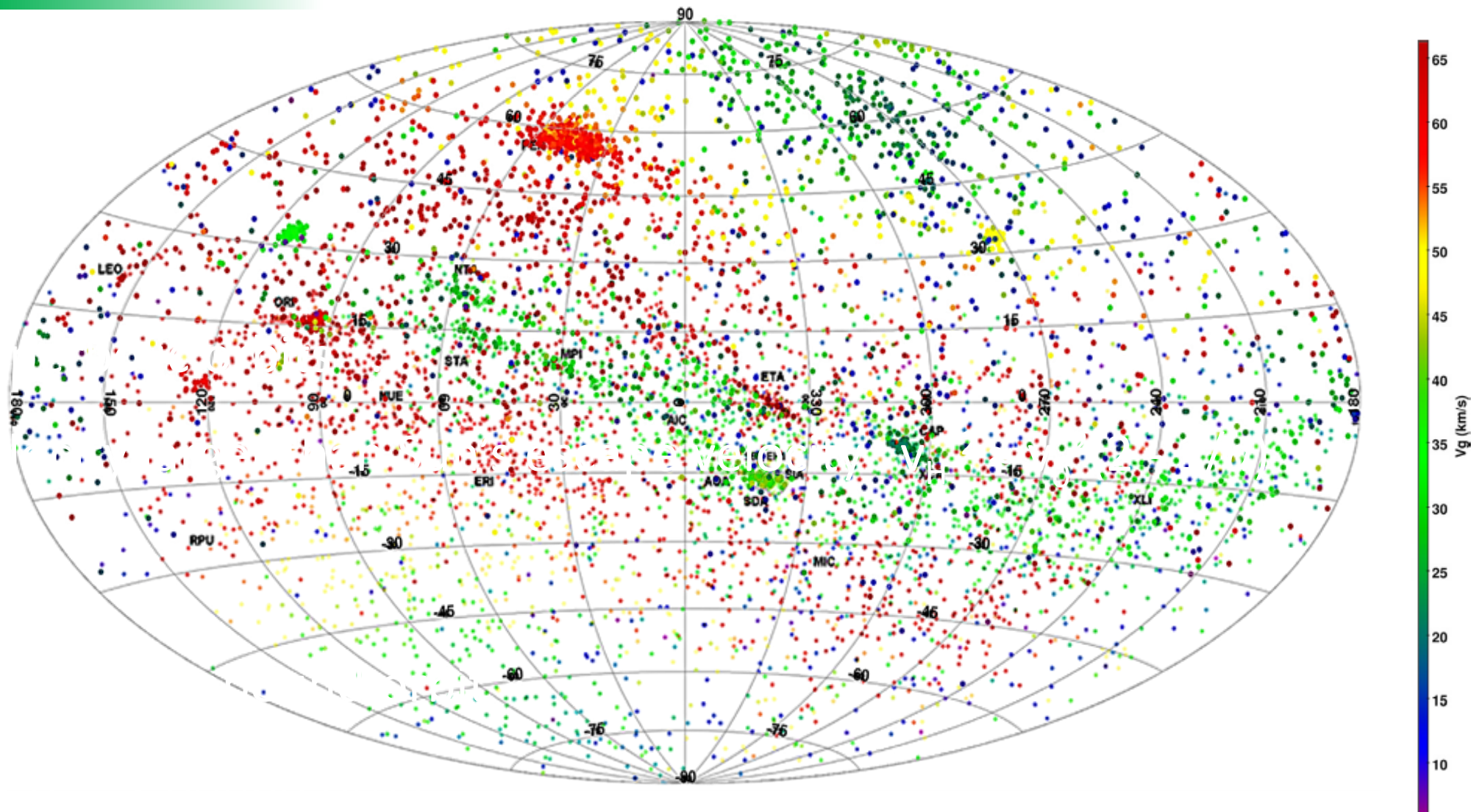
Neslušan, L., Hajduková, M.,
Icarus, 382, 115015, 2022
Hajduková, M., Neslušan, L.,
Icarus, 387, 115175, 2022



v_{inf} and α, δ

METEOR RADIANTS

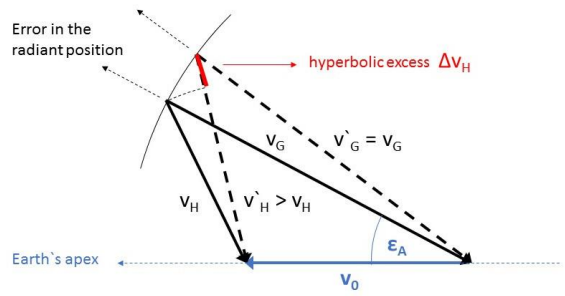
V_G



Tóth et al., AMOS all-sky video system, Toth, J. et al., PSS, 118, 102, 2015; Toth, J. et al, EPSC, Berlin, 2018

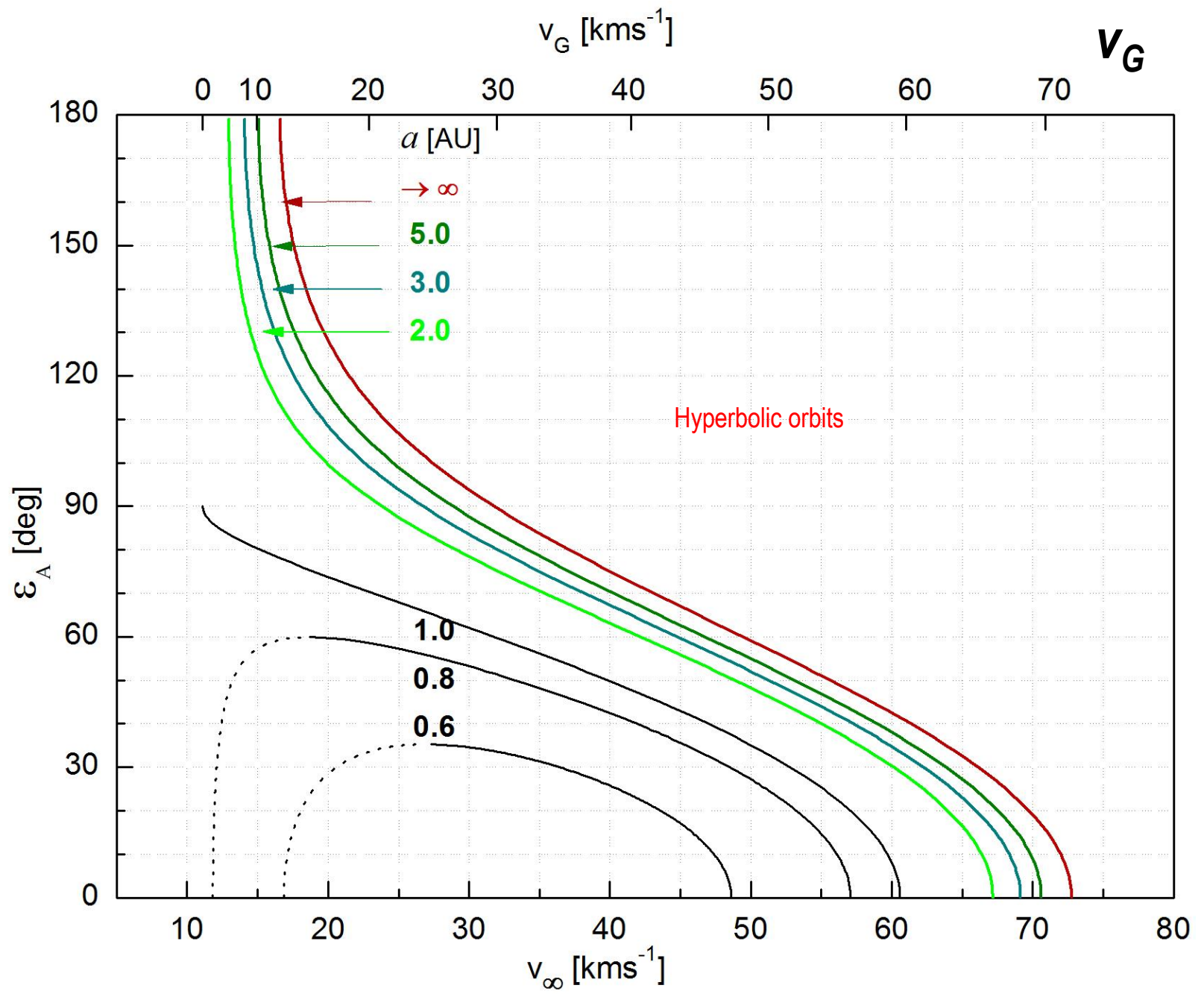
v_{inf} and α, δ

ϵ_A

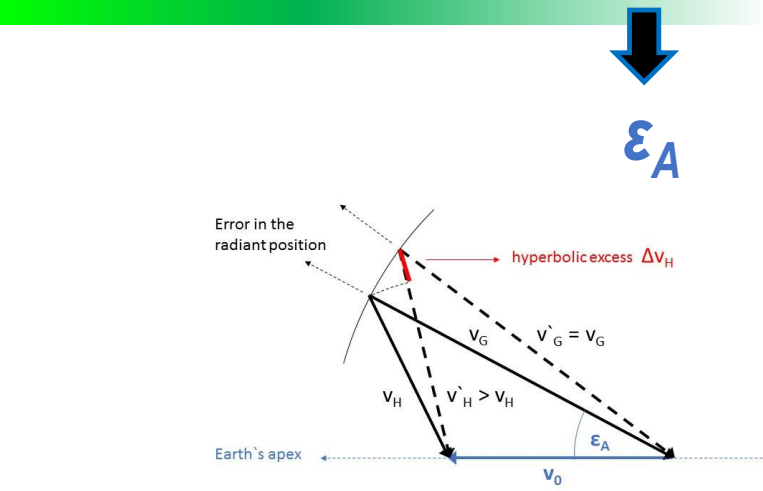


Kresák, L. and Kresáková, M., BAC, 27, 106, 1976

Hajduková, M., Sterken, V., Wiegert, P., Meteoroids CAP, 235, 2019



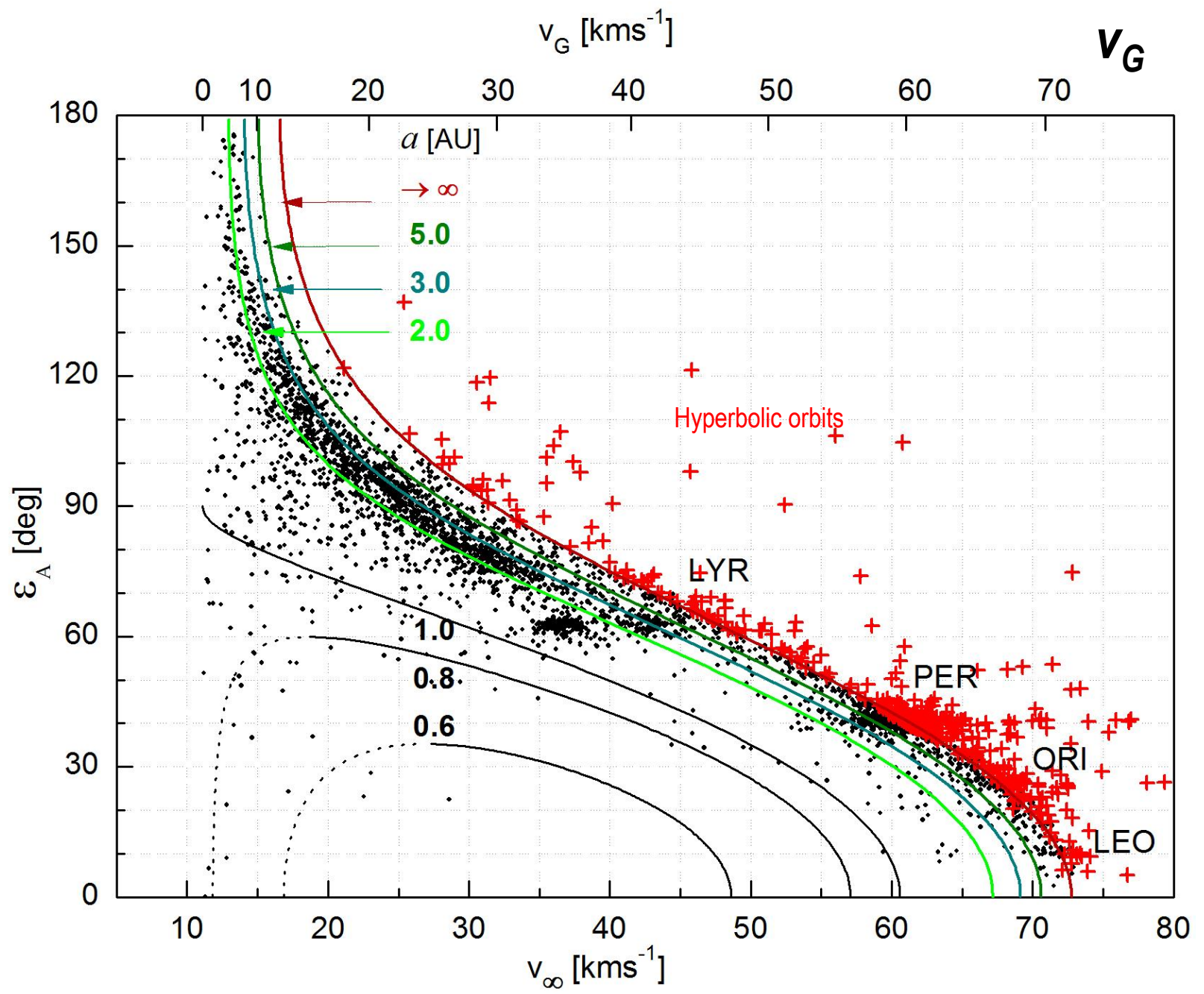
v_{inf} and α, δ



Kresák, L. and Kresáková, M., BAC, 27, 106, 1976

Hajduková, M., Sterken, V., Wiegert, P., Meteoroids CAP, 235, 2019

Photographic meteors from Lindblad, B. A., et al., EM&P, 93, 249, 2005



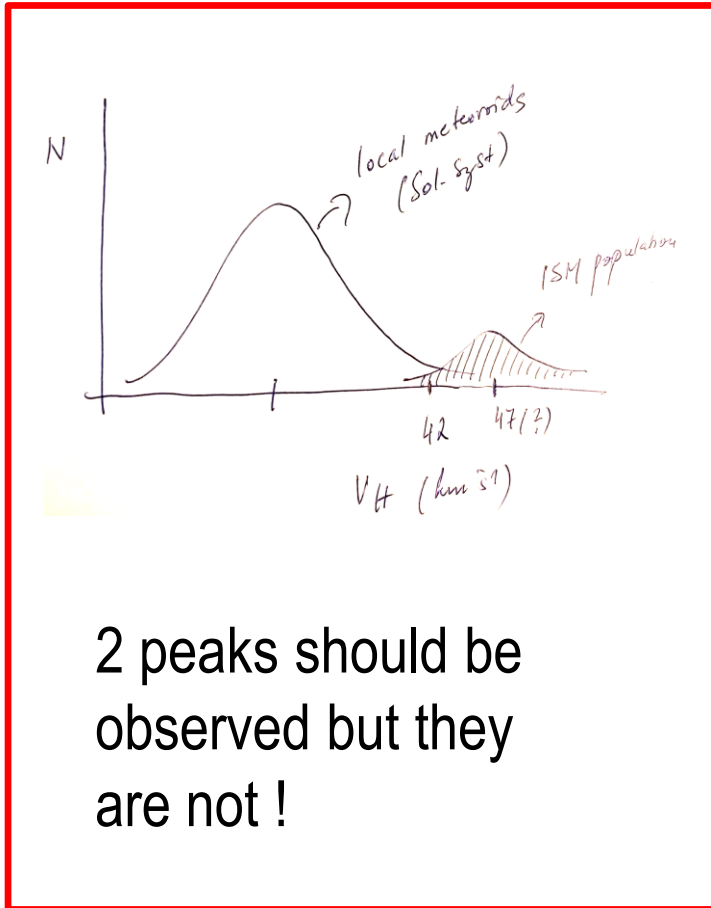
v_{inf} and α, δ

$a, e, q, \omega, \Omega, i, T$

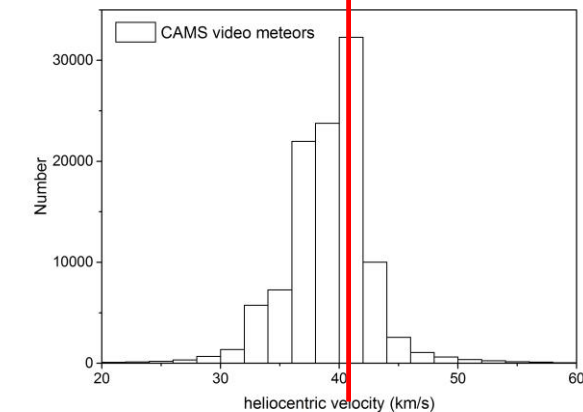
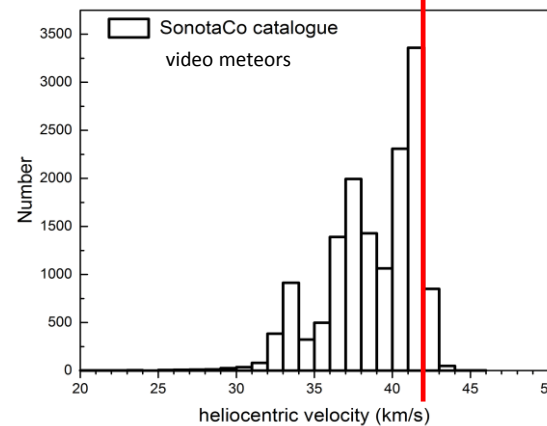
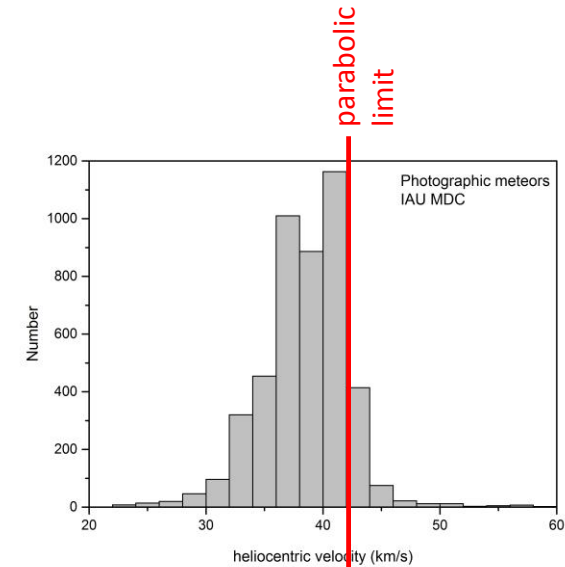
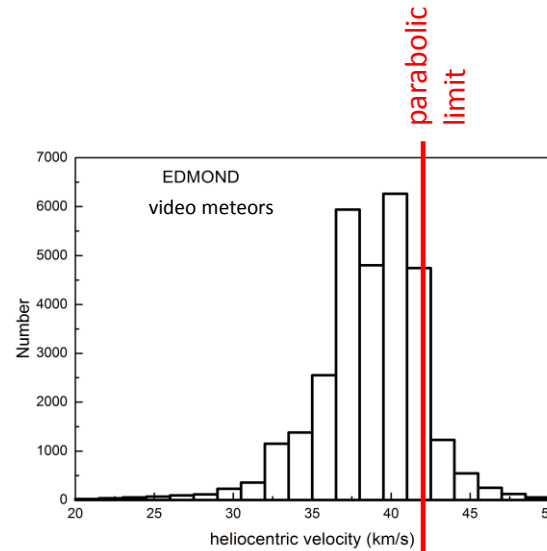
WHY IS IT A CONSEQUENCE OF ERRORS?

1. Distributions of the heliocentric velocity

$$v_H^2 \approx v_0^2 (2 - 1/a)$$



2 peaks should be observed but they are not !

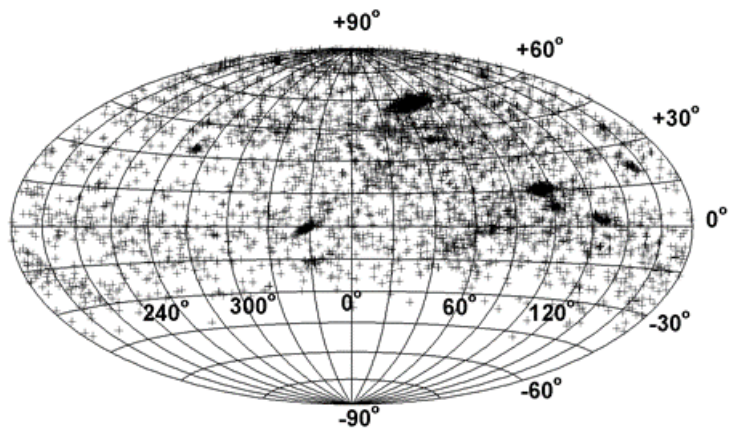




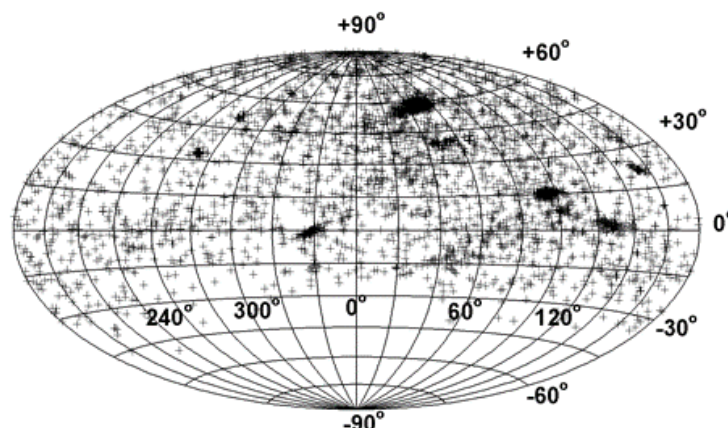
$$N_{e>1}/N = f(v_H)$$

WHY IS IT A CONSEQUENCE OF ERRORS?

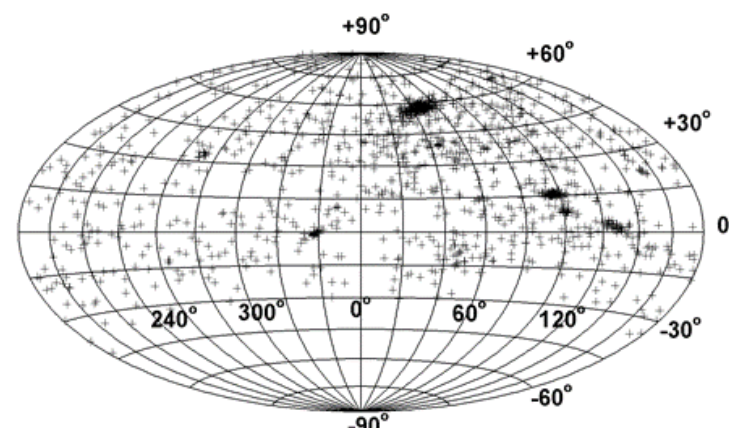
2. Hyperbolic orbits of shower meteors



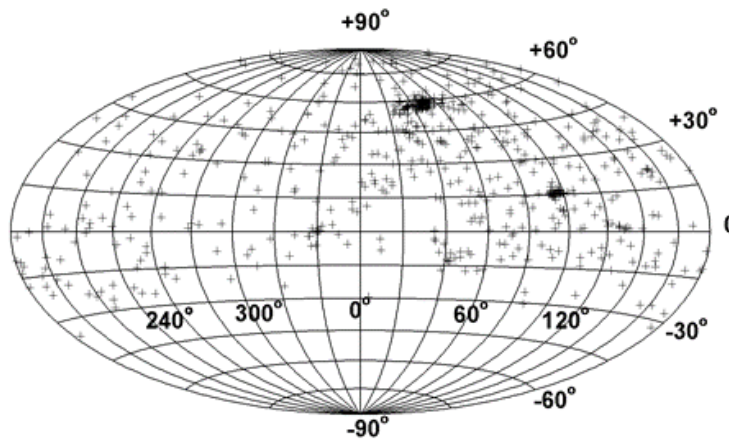
$0.2 < 1/a < 0.1$



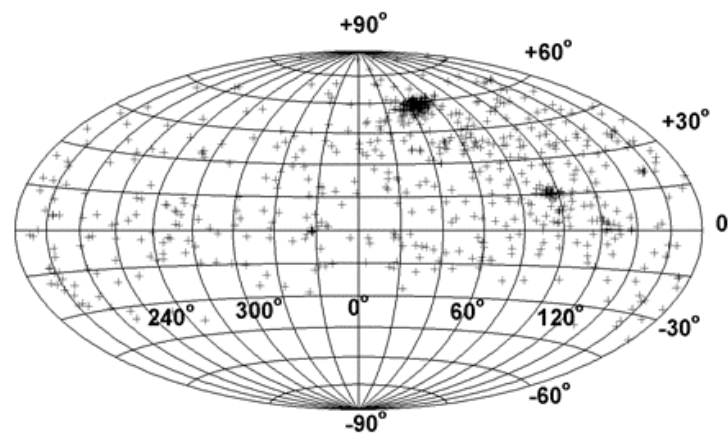
$0.0 < 1/a < 0.1$



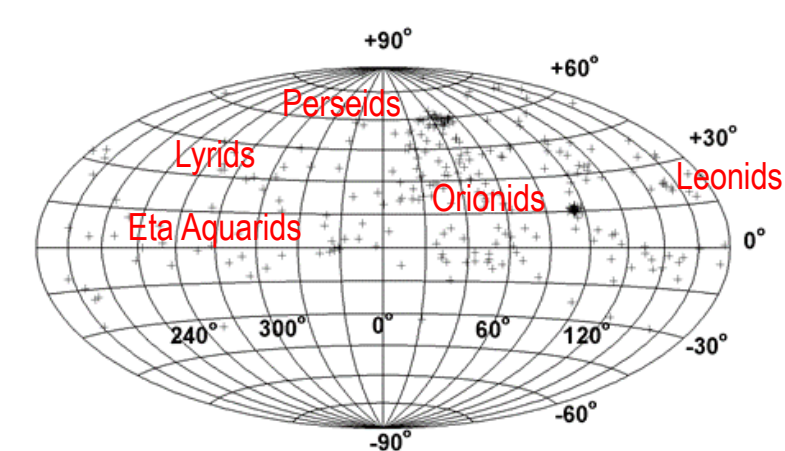
$-0.1 < 1/a < 0.0$



$-0.2 < 1/a < -0.1$



$-0.5 < 1/a < -0.2$



$1/a < -0.5$

Perseids
Lyrids
Eta Aquarids
Orionids
Leonids



v_{inf}
radiant

± 1 km/s
 $\pm 1^\circ$

WHY IS IT A CONSEQUENCE OF ERRORS?

3. Modeling the influence of errors

Nominal orbit

$q = 0.945$ au

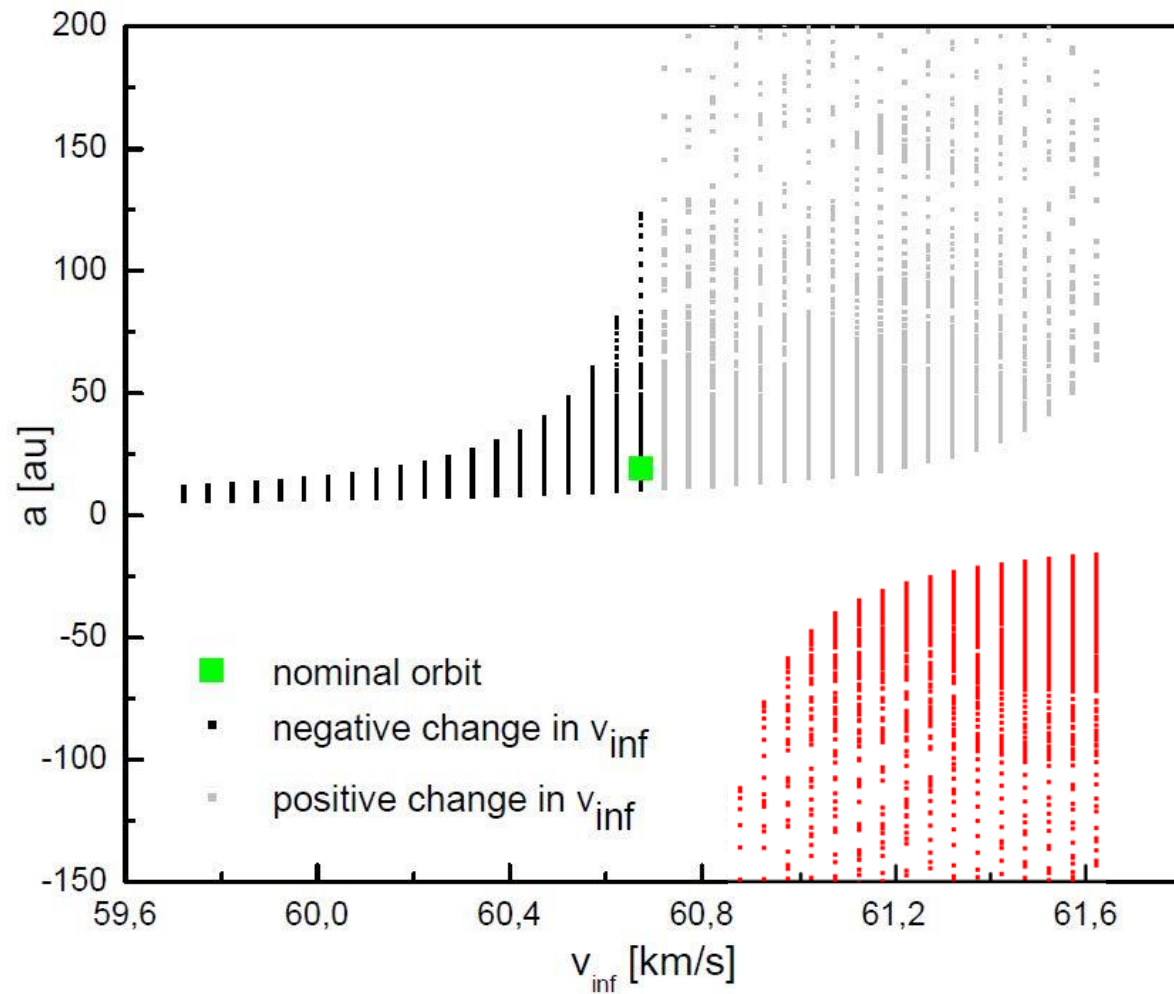
$a = 19$ au

$e = 0.95$

$Q = 37$ au

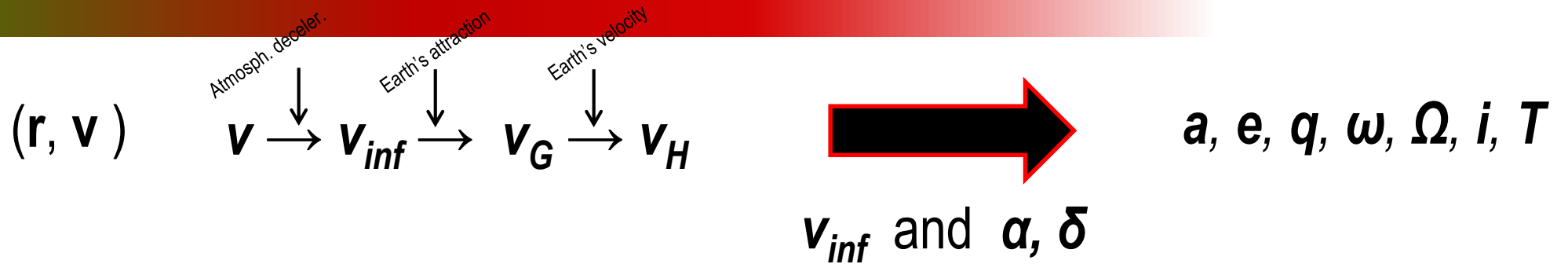
$i = 113^\circ$

Hajduková, M., Kornoš, L.,
PSS, 190, 104965, 2020



1/3 of orbits changed to hyperbolic

PROBLEMS



underestimation of measurement errors

small number of members \rightarrow mean parameters dispersed \rightarrow the match of a given shower with a predicted counterpart is not clear

problems with the \rightarrow duplicity of showers
 \rightarrow identity of showers

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THANK YOU
FOR YOUR
ATTENTION



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