

IAPS

Bologna, 22-24/06/2022 – Forum della Ricerca Sperimentale e Tecnologica in INAF



Laboratori e Facilities

IAPS ISTITUTO DI ASTROFISICA
E PLANETOLOGIA SPAZIALI

Stefania Stefani

in rappresentanza di tutto il personale coinvolto nelle attività

Introduction

Facilities

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- Plasma Chamber SWIPS (Solar Wind and Ionospheric Plasma Simulation)
- Thermo-Vacuum chambers
- ROCT (Run Optical Chamber Test setup)
- Calibration of accelerometers , drop tower and MTGSE
- X-Ray
- Production of Polarized and Unpolarized Monochromatic X-rays

Laboratories

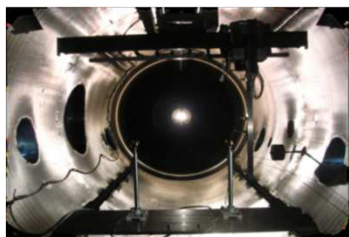
- Spectroscopy
 - ❖ Gas, dust, rocks, micro-meteorites
- Space materials Laboratory
- Space instrument for extra-terrestrial dust
- Ion-ENA Beam laboratory
- Cryogenic laboratory



Plasma Chamber SWIPS (Solar Wind and Ionospheric Plasma Simulation)

Referente: Piero Diego

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Test and Calibration



Tests on Ion Drift Meter and Retarding Potential Analyzer



Simulation of attitude effects on Electric Field Detector

Main features:

- Large volume experimental facility (about 9 m³) for **realistic ionospheric** (density 10¹¹-10¹² m⁻³, electrons T 0.2-0.4 eV, Vion @ 8000m/s) and **interplanetary** (density 10⁷-10⁸ m⁻³, electrons temperature 1-2 eV, Vion @ 350 km/s) environment simulation.
- Geomagnetic field nullified (residual < 0.01 G) or adjusted to simulate S/C orbits

Experimental activities concern:

- Instrument development test and calibration
- Space plasma compatibility for materials and technologies for space applications
- basic plasma physics experiments such as interaction of neutral, charged, and magnetized bodies with plasma for the evaluation of thrust, etching or any kind of plasma-bodies interaction

Current Projects and Collaborations

- o CNSA-ASI project Limadou2 : development of QM and FM of Electric Field Detector for CSES-02 satellite
- o ESA-INAF project CUBE (CME Catcher Carousel): development of Plasma Analyzer
- o IAPS-CIRA collaboration for the improvement of various space technologies
- o IAPS-NSSC collaboration for the development of plasma sources and diagnostic systems

Instrument development:

view of the electric field probe currently under development to be installed on CSES-02 satellite (launch foreseen for 2021 sun-synchronous orbit at 500 km altitude)



Mechanical sketch of retarding potential analyzer /ion capture meter developed to check

solar wind velocity and species, and also suitable for cubesat.



Thermo-vacuum chambers

Referenti: G. Piccioni, D. Biondi e A. Boccaccini

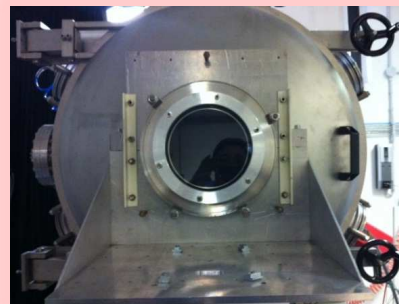
The two test facilities have been employed to test and to qualify electronic, mechanical and optical parts of spatial instrumentations.



Main features

- ✓ Volume 1600 l
- ✓ Vacuum limit 10^{-7} mbar
- ✓ Primary vacuum 10^{-2} mbar
- ✓ Temperature range $-60^{\circ}/+100^{\circ}\text{C}$

The thermo-vacuum chamber is placed inside a clean room of 10000



• Main features:

- ✓ Volume: 90 l
- ✓ Primary vacuum: 10^{-2} mbar
- ✓ Max vacuum : 10^{-6} mbar
- ✓ Temperature range: $-180/60^{\circ}\text{C}$
- ✓ Optical window: 165mm CaF2

Internal Collaboration:

- ✓ Fasy
- ✓ INAF-To
- ✓ IXPE
- ✓ MAJIS

External Collaboration:

- ✓ Bercella
- ✓ CGS
- ✓ IKR
- ✓ Kayser
- ✓ Leonardo
- ✓ OHB Italia
- ✓ Sab Aerospace



Actually the TV chamber is placed inside of a clean room ISO 6

The CRV was employed to test optical and mechanical parts of the VIRTIS spectrometer on board of the VENUS EXPRES, ROSETTA space mission of the SERENA/ELENA instrumentations on board of the BepiColombo

ROCT (Run Optical Chamber Test setup)

Referenti: G. Piccioni, S. Stefani, D. Biondi e A. Boccaccini

Project: MAJIS_JUICE

System:

Cryo-cooler 2nd stage:

Cooling Capacity:

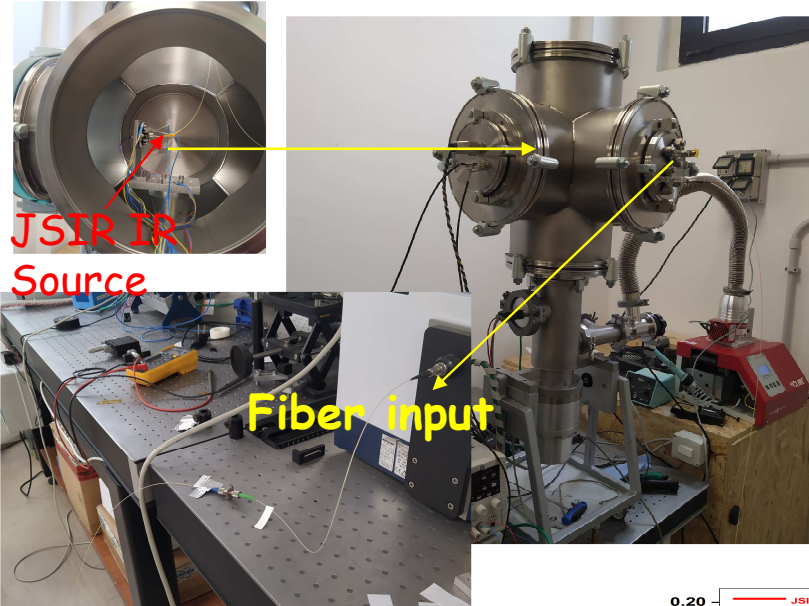
1st 50/40W at 43K

2nd 1.0 W at 4.2K

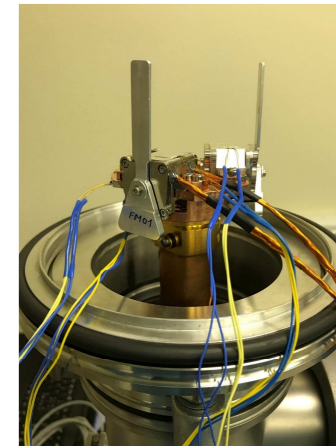
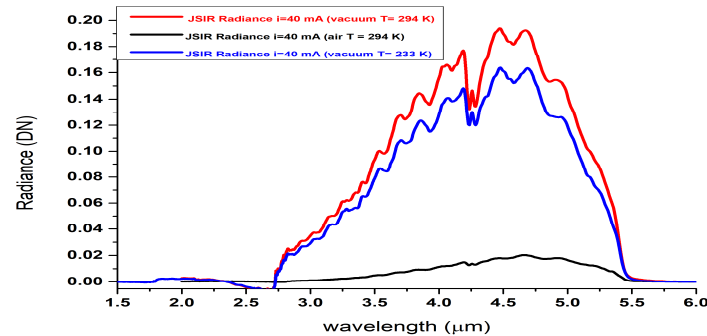
Vacuum chamber: 10^{-6} mbar

Principal activities:

- 1) Life time test on the FM of the Shutter which is a part of the MAJIS spectrometer on board of the JUICE space mission.



2) Optical test on the JSIR emitter source.
The signal coming from to the source, was acquires with a FT-IR spectrometer in the spectral range 2-6 μm



Operative Temperature 135 K
cycles performed without damage: 139,200

FACILITY di test, calibrazione e Mechanical and Thermal Ground Support Equipment (MTGSE)

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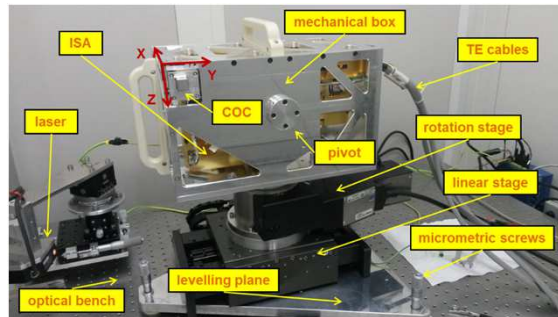
Referenti: C. Lefevre, M. Lucente, E. Fiorenza

Progetti: ISA/BepiColombo (ESA/ASI), HAA/JUICE (ESA), GReAT (ASI)

FACILITY DI CALIBRAZIONE ACCELEROMETRI

Sviluppata per ISA/BepiColombo e HAA/JUICE per la misura della direzione degli assi sensibili:

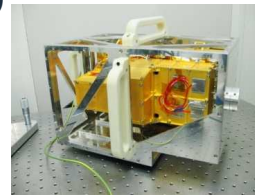
- ✓ Linear stage (0-50 mm) Aerotech ABL15005, con roll/pitch/yaw ± 0.5 arcsec, UUT fino a 35 kg
- ✓ Rotation stage Newport RV120HAT ($\pm 170^\circ$, $\sim 0.005^\circ$), UUT fino a 180 kg
- ✓ 3 laser a triangolazione Keyence (650 nm, 0.95 mW), per misure di piccoli spostamenti (range ± 18 mm, $\sim 0.5 \mu\text{m}$)



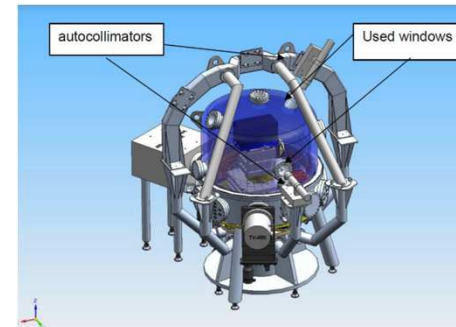
Torre di caduta

Struttura tubolare di 1.5 m di diametro e altezza 10 m, coibentata, acusticamente isolata, utilizzabile per prove di caduta.

Sviluppata per eseguire test di sgancio dell'accelerometro differenziale GReAT (General Relativity Accuracy Test in an Einstein Elevator), un esperimento per la verifica del Principio di Equivalenza di Einstein su pallone stratosferico (WEP).



MTGSE



Facility di test per effettuare analisi e caratterizzazioni di accelerometri in termo-vuoto controllato. Dotata di movimentazioni di precisione della Micos (tip-tilt).

- Sviluppata per accelerometri ISA e HAA
- Camera TermoVuoto (10-6/10-8 mbar, $-40/+60^\circ\text{C}$) con 2 cold plate.
- Volume interno ca. 600 l
- UUT fino a 30 kg
- Movimentazioni (tip-tilt, rotation stage)
- Autocollimatori
- Ambiente pulito (classe 100.000)
- Basamento e struttura esterna per scambiatori di calore, da realizzare.



X-Ray Facility

Referenti: Y. Evangelista, F. Ceraudo, M. Feroci

Projects: ADAM/PixDD, HERMES, eXTP

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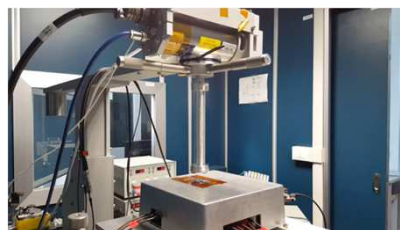
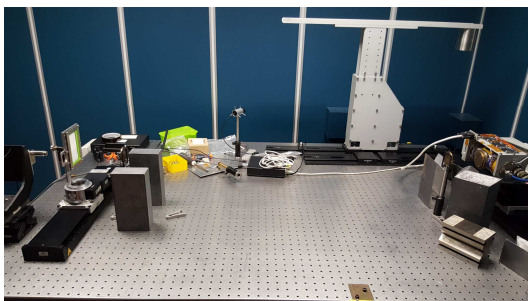
Evangelista et al. 2018, JINST 13, P09011, Campana, R. et al., 2011, NIMA, 633 11 22-30, Cirrincione, D. et al., 2019, NIMA, 936, 239

Lead-shielded chamber for X-ray studies

- Angular resolution of optical devices
- Point-by-point response of sensors

Equipment list:

- Optical bench
- Manual and motorized linear (horizontal and vertical) and rotational stages
- Motorized stages can be controlled remotely
- Single-anode X-ray generators (Fe, Ti, Mo, Rh...) by Oxford
 - Emission lines and Bremsstrahlung continuum
 - 50 W max power (50 kV, 1 mA)
 - Interlock safety system connected to chamber doors and temperature monitor of X-ray tube
- Range of collimators and diaphragms
 - Decrease spread of X-ray beam and increase spatial/angular resolution
- Independent air-cooling system
 - Allows moderate temperature control
- Lead-glass window
 - Enables optical inspection even during beam operation



Enhanced X-ray Timing Polarimetry (eXTP)

- Angular response of MCP collimators for the Large Area Detector (LAD)
- Achieved beam properties at device under test:
 - 0.7 mm-diameter
 - 2.4 arcmin-divergence
- Reference:
 - Ceraudo, F. et al., 2018, JINST 13 9 P09020 doi:10.1088/1748-0221/13/09/P09020

Pixelated Drift Detector (PixDD)

- Point-by-point response of multi-pixel Silicon Drift Detector (SDD)
- Achieved beam properties at device under test:
 - $< 50 \mu\text{m}$ -FWHM
- Reference
 - Evangelista, Y. et al., 2018, JINST 13 9 P09011, doi:10.1088/1748-0221/13/09/P09011

Cable feed-throughs

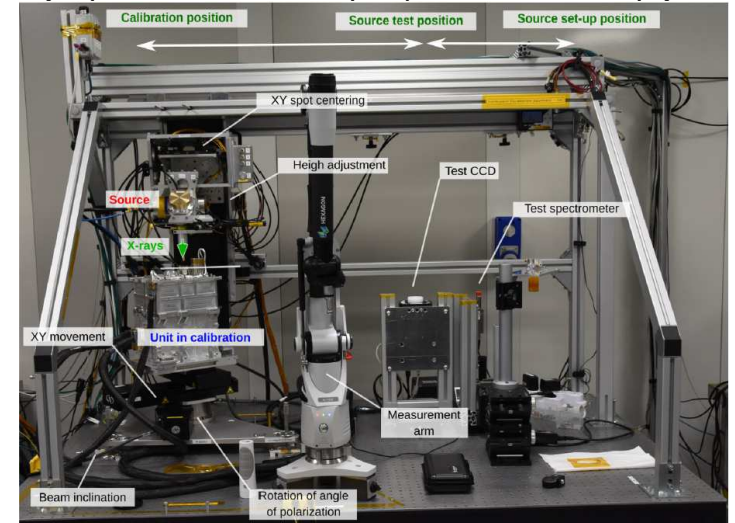


Facility for production of Polarized and Unpolarized Monochromatic X-rays

Referenti: P. Soffitta, F. Muleri, S. Fabiani and E. del Monte

Projects: IXPE (Imaging X-ray Polarimetry Explorer), CUSP (Cubesat Solar Flare Polarimeter), Hype-X (High Yield Polarimetry Experiments in X-rays)

- Bragg diffraction at 45° (Polarized Monochromatic)
- Fluorescence lines (Unpolarized Monochromatic)
- Metrology system for angle and position determination
- X-ray tubes 50 W 1 mA plus crystals and collimator for Polarized X-rays
- Head-on X-ray tube 2W 1 mA for producing unpolarized X-rays
- Two stations: one in a ISO-7 clean room one in lead-shielded room
- Equipped with a X-ray Andor CCD, Amptek-SDD
- Computer controlled linear and rotary stages



Successfully employed for the ground calibration of the IXPE four detector Units (DUs)

Operated in total 160 days 7/7 days/week 24/24 hours/day

Energy	Configuration	Source settings	Rate [cts/s]	Polarization	
				Degree [%]	Angle [deg]
2.04	Fluorescence of Zr target illuminated by Rh X-ray tube	19 kV, 0.95 mA	~177	1.12 ± 0.14	1.38 ± 3.62 ^a
2.29	Fluorescence of Mo target illuminated by Ag X-ray tube	17 kV, 0.77 mA	~199	0.85 ± 0.11	1.32 ± 3.63 ^a
2.70	Direct X-ray tube with Rh anode + 300 μm PVC filter	4.0 kV, 0.52 mA	~205	6.71 ± 0.10	-0.93 ± 0.41 ^a
2.98	Direct X-ray tube with Ag anode + 6 μm Ag filter	4.0 kV, 0.06 mA	~215	13.20 ± 0.09	-0.19 ± 0.19 ^a
3.69	Direct X-ray tube with Ca anode	5.4 kV, 0.003 mA	~213	Undetected, MDP(99%) = 0.22%	Undetected
5.89	⁵⁵ Fe nuclide	4 mCi	~180	Undetected, MDP(99%) = 0.19%	Undetected

keV	Crystal	Line	Bragg angle
1.65	ADP(101)	CONT	45.0
2.01	PET(002)	CONT	45.0
2.29	Rh(001)	Mo L _α	45.3
2.61	Graphite	CONT	45.0
3.7	Al(111)	Ca K _α	45.9
4.5	CaF ₂ (220)	Ti K _α	45.4
5.9	LiF(002)	⁵⁵ Fe	47.6
8.05	Ge(333)	Cu K _α	45.0
9.7	FLi(420)	Au L _α	45.1
17.4	Fli(800)	Mo K _α	44.8

Monochromatic Polarized Sources

Unpolarized Monochromatic Sources Muleri et al., 2022





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Laboratories

PASSxS (Planetary Atmosphere System Simulation for Spectroscopy)

Referenti: S. Stefani, G. Piccioni

Projects: VENUS-Express, JUNO-JIRAM- MAJIS-JUICE and ARIel

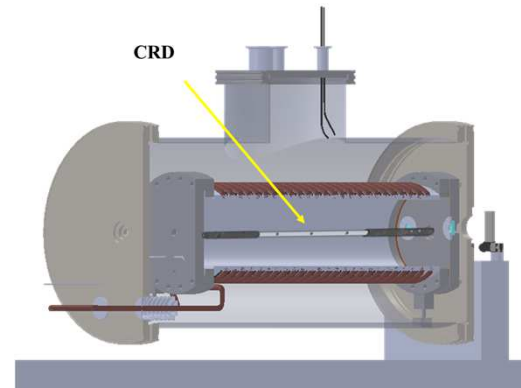
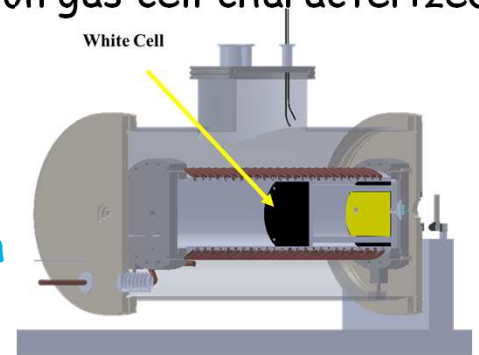
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Goal: experimental set up used to characterize the optical properties of gases at typical planetary conditions high/low pressure ($\text{mbar} < P < 70 \text{ bar}$) high/low temperature ($100 \text{ K} < T < 500 \text{ K}$). In particular CO_2 , H_2 and different mixing ($\text{H}_2 + \text{He}$, $\text{H}_2 + \text{H}_2$ and CH_4) typical of the Jupiter atmosphere. When coupled with a FT-IR spectrometer (as shown on the left), the chamber is equipped with a Multi Pass (MP) absorption gas cell characterized by an optical path of about 9.2 m.

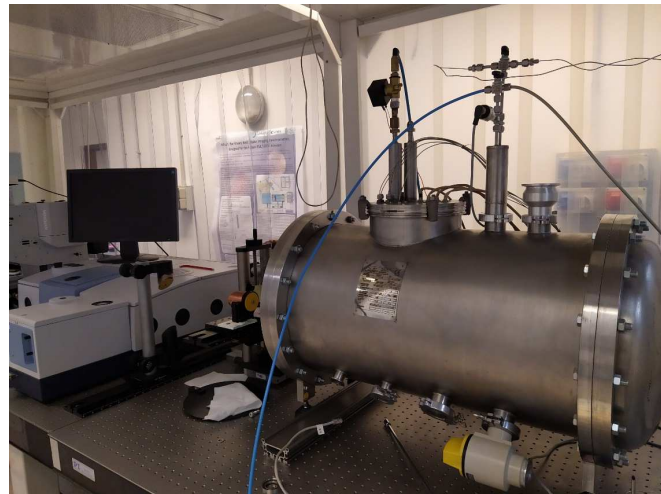
Resolution: $0.01 - 10 \text{ cm}^{-1}$

Spectral range : $[0.3 - 8] \mu\text{m}$

the chamber is equipped with a White Cell (MP) absorption gas cell characterized by an optical path of about 9.2 m.



Actually the set up is placed inside of a Soft-wall ISO 6 chamber



When coupled with a Cavity Ring Down (CRD) tube:
optical path of about 5 km

Thanks to the high reflectivity of the mirrors we can measure the total absorption of about 10^{-6}

Spectral range depends on the laser tunability



Reflectance and μ IR spectroscopy (PLAb)

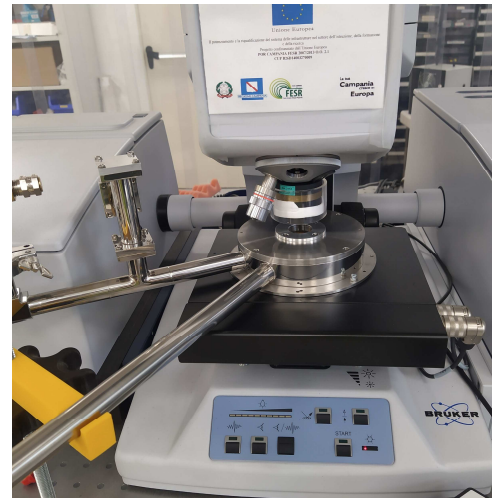
Referenti: S. Stefani and G. Piccioni

Projects: MAJIS-JUICE and others external collaborations

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Goniometer accessory of Vertex 80

- ✓ Bi-conical Reflectance (low angle 13°): variable incidence and emission angle: ($13-85^\circ$); minimum phase angle 26° ;
- ✓ Spatial Resolution: spot of **5 mm** in diameter;
- ✓ Motorized or manual controlled arm, fixed sample;
- ✓ **spectral range: variable 400-16000 nm**



μ IR-spectroscopy connected to the Vertex 80 interferometer

- working range $0.4-14 \mu\text{m}$
- used to acquire spectra in reflection and transmission mode on single feature with minimum dimensions of $50 \mu\text{m}$
- Used to acquire IR images

Integrated with a micro-cryostat

Temperature range $< 6 \text{ K}$ (LHe) to 475 K

Temperature stability $< 50 \text{ mK}$

Older dimension about $1'$

Will be employed to characterized optical filters, micro-meteorites, rocks @ different temperatures

Spectroscopy Laboratory (SLAb) and Sample Facility Preparation (SFP)

Referenti: C. Carli, E. Bruschini and A. Stephan

Main Project: SIMBIO-SYS (Bepicolombo) OI-BODIES (ASI-INAF 2018.HH.0) POSEIDON (H2020 Marie Curie)
Supported also by Dawn/ExoMars and different IAPS director funding for Laboratories.

VNIR FieldSpec Pro coupled with a Goniometer

- ✓ Bi-Directional Reflectance: variable incidence and emission angle (0-65°), minimum phase angle 28°;
- ✓ Spatial Resolution: spot of 6 mm in diameter;
- ✓ Motorized arm and X-Y plane, controlled by computer;
- ✓ spectral range: 350 - 2500 nm, resolution 2÷12nm

Sample preparation: Build in 2010, using an already prepared set-up at which we dedicated a specific laboratory to work systematically on solids as analogues for planetary surfaces.

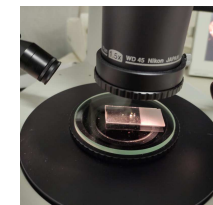
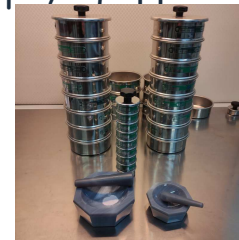
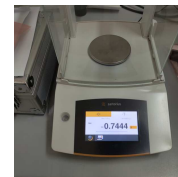
The facility is characterized by 2 machine rocks grinding, a saw for rocks cutting, an electronic sieving. Sieves are present for sizes from 2 mm up to 20 mm with a diameter of 10 cm and 2 series of 2.5 cm in diameter and from 200 up to 20 mm, plus ultrasound pool for cleaning.

Two different hand ground agata system. To produced samples from bulk rocks to cut sample and to powder at different grainsizes. We have also an high precision balance and a drying cab. Recently we had also a plate for fining polish. A preliminary analysis can be done using two different microscopes a stereoscope and a petrographic microscope, equipped with cameras.



This lab supported also the sample preparation for measurements at IPAG in Grenoble, PSL at DLR in Berlin, at IAS and at SOLEIL Paris. Support to:

- a) More then 20 publications,
- b) 5 Ph.D. students and 10 Master/Bachelor Thesis or
- c) SSHADE spectra library



Laboratorio Spettroscopia Superfici C-Lab

referenti: M.C. De Sanctis -1°ric, S. De Angelis -ric. TI, M. Ferrari -ric. TI, A. Frigeri -ric. TI, E. La Francesca -AdR, L. Rossi -BS, N. Costa -BS

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Missioni spaziali Sistema

Solare

- * ExoMars
- * Dawn
- * Prospect
- * JUICE



Spettroscopia VIS-NIR-MIR-Raman

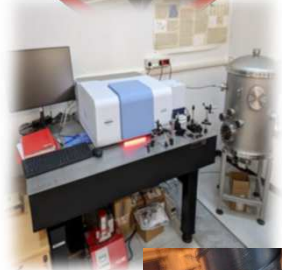
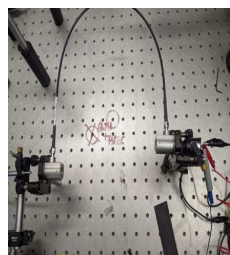
- * Analoghi superficiali planetarie (Marte, Cerere, Luna, asteroidi, ...)*
- Rocce/minerali/misture * Meteoriti *
- Target di riferimento (commerciali, FM spare)

Sviluppo camere UHV

- * Cella PT + Fspe4 / SPIM
- * Progetto CAPSULA + FTIR
- * Misure spettrali analoghi a condizioni variabili di P-T-atm

Caratterizzazione Setup, Fibre Ottiche + Sorgenti IR

- * Progetto TRIS



Attività sul campo

- * Misure su terreno con prototipi / strumenti portatili
- * Raccolta campioni per lab



Modelli di laboratorio / Spare di spettrometri spaziali

- * Ma_MISSION + Drill / ExoMars
- * SPIM / Dawn

Pubblicazioni (Peer Review)

- De Angelis S. et al., *PSS*, 101, 89-107, 2014
- De Angelis S. et al., *PSS*, 117, 329-344, 2015
- De Angelis S. et al., *Rev.Sci.Instr.*, 86, 093101, 2015
- BostN. et al., *PSS*, 108, 87-97, 2015
- Manzari P. et al., *Earth Sp.Science*, 3, 2016
- De Angelis S. et al., *Icarus*, 280, 315-327, 2016
- De Sanctis M.C. et al., *Astrobiology*, 17, 6-7, 612-620
- De Angelis S. et al., *PSS*, 144, 1-15, 2017
- Manzari P. et al., *MAPS*, 1-20, 2018
- Ferrari M. et al., *Icarus*, 321, 522-530, 2019
- De Angelis et al., *Rev.Sci.Instr.*, 89, 103107, 2019
- De Sanctis M.C. et al., *MNRAS*, 482, 2, 2407-2421, 2019
- Singh S.K. et al., *Nature Communications*, 12:2690, 2021
- Schroder S. et al., *Nature Communications*, 12:274, 2021
- De Angelis S. et al., *JGR:Planets*, 126, 2021
- Pisello A. et al., *Icarus*, 374, 114801
- De Sanctis M.C. et al., *PSJ*, 2022



Space materials Laboratory

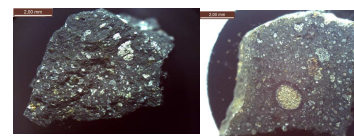
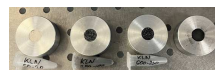
Referenti: E. Palomba, F. Dirri, A. Longobardo

Goals:

1. Meteorites and analogue mixtures characterization with FTIR spectrometer (spectral range:) for JAXA-Hayabusa2 (Ryugu), JAXA-DESTINY+ (3200 Phaethon asteroid) and ESA-HERA (Didymos-Dimorphos binary asteroid) Space Missions



2. Thermal setup at 500-700°C in vacuum chamber to simulate thermal metamorphism on asteroids of meteorites/analogue mixtures

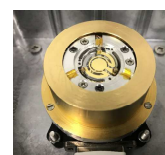


Goals:

1. Development of Microbalance device for Space Applications (dust and volatiles characterization in different planetary environments)
2. Development and testing of Microbalance device for laboratory use (Materials outgassing evaluation in dedicated vacuum chambers)



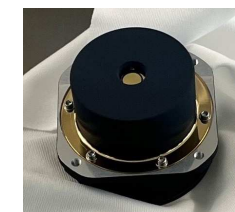
CAMLAB (Contamination Assessment Microbalance for Laboratory) developed for "Development of a European Quartz Crystal Microbalance" ESA-ITT Project (2017-2019)



CAM (Contamination Assessment Microbalance), developed for "Evaluation of an in-situ Molecular Contamination Sensor for space use" ESA-ITT Project (2014-2016)



VISTA (Volatiles In-Situ Thermogravimeter Analyser) developed for HERA-ESA Mission onboard Milani Cubesat (2020-2023)



Space instruments for extra-terrestrial dust:

- 'in situ' measurements
- collection and retrieval

Referenti: Vincenzo Della Corte¹, Alessandra Rotundi^{2,1}, Andrea Longobardo¹, Stefano Ferretti^{2,1}, Anna Musolino¹, Alice Maria

Extra-terrestrial dust particles dynamical properties «in situ» measurements: DISC

Project: DISC (Dust Impact Sensor and Counter), part of the DFP (Dust Field and Particles) suite, onboard ESA Fast space mission Comet Interceptor (Launch 2029).

Heritage: GIADA onboard ESA/Rosetta space probe.

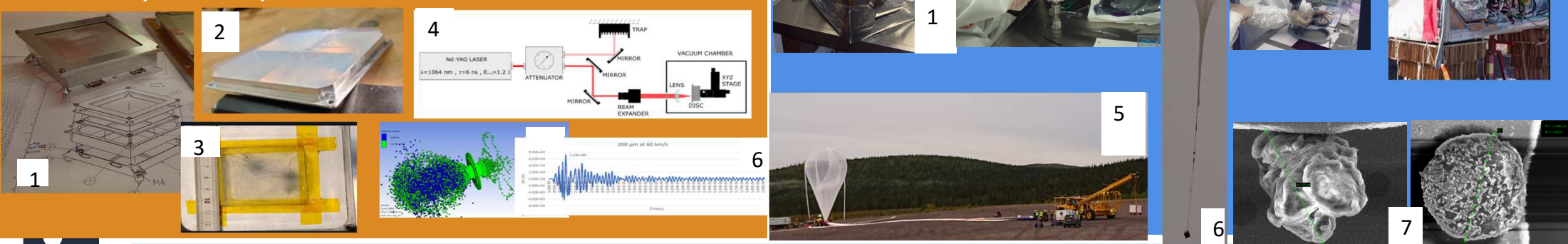
Activities: Instrument development and management.

- Bread Board sensor design and test (1).
- Dust shield for High Velocity impacts: design and test with hyper velocity facilities (2,3).
- Hyper Velocity Dust particle impacts simulation:
 - High power pulsed laser experimental setup (4)
 - Ansys/Autodyn simulation chain (5,6)

Extra-terrestrial dust particles collection in the Earth stratosphere & retrieval for laboratory analyses: DUSTER Project: DUSTER (Dust in the Upper Stratosphere Tracking Experiment and Retrieval), onboard stratospheric balloons.

Activities: Instrument development and management:

- Instrument design/manufacturing (1)
- Campaign preparation (2,3,4,5,6)
- Campaigns: 2008,2009 Svalbard (Norway), 2011, 2019, 2021 Kiruna (Sweden)
- Collected dust particles identification (7)
- Collected dust particles analyses and classification (7)

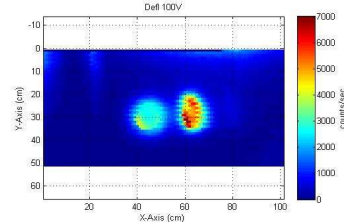
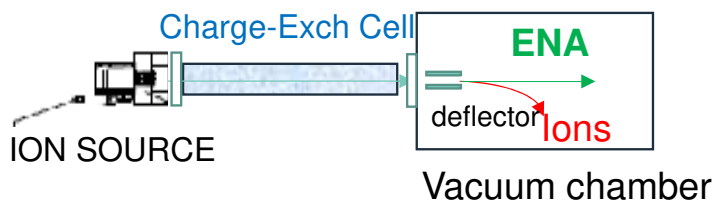


Ion-ENA beam Laboratory

Referenti: E. De Angelis, R. Rispoli, L. Colasanti, N. Vertolli, F. Nuccilli and M. Moroni

Main general tasks:

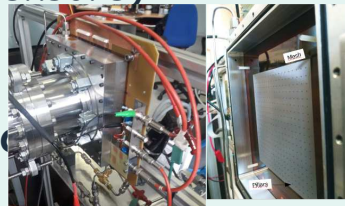
- Development, testing and calibration of space instruments for Ion and Energetic Neutral Atoms (ENA) detection devoted to Planetary Space Weather
- Laboratory simulation of interaction phenomena generated by Solar Wind with solar system bodies and environments.
- Interaction investigation of selectable ion/neutral beam with several samples (instrument devices, detectors, planetary analogues).



SWEATERS project (Space Weather ENA Sensors)

INAF-INFN with CERN and SWRI

- Development of innovative ENA sensor :
First ENA instrument based on Gas detector (MPGD-Micro Pattern Gas detector)



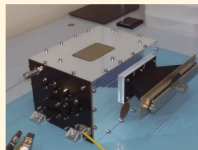
for Space Weather application: Ion/ENA beam on MPGD type detector

- Test @IAPS, set-up for detector characterization
- Test on CarbonFoil/Graphene entrance window
- Study, simulation and testing for parameters refining

[ASI contract: «SWEATERS project»]

SERENA/ELENA: instruments on board the BepiColombo/Mercury Planetary Orbiter (launched October 2018- at Mercury 2025) ESA-JAXA

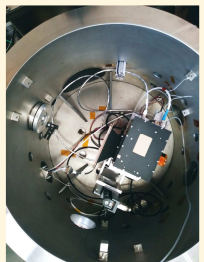
- ELENA Flight Spare activities
Test on ground during BC cruise phase:



ELENA FM
(on-board BC mission)

Satellite simulator (MIS) allows to command Serena\ELENA FS in lab @Mercury beam condition:

- Mercury Lab simulation
- Test with ion and ENA beam to verify different flight instrument configuration
- Development and testing of on-board procedure and SW updates



ELENA FS (vacuum chamber @IAPS)

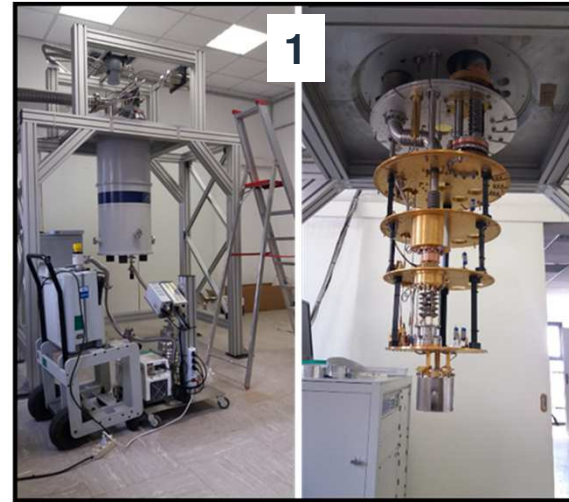
(ASI contract: «Serena/BC-Fase E»]

Cryogenic laboratory for X-ray Astrophysics:

more details at: <http://www.iaps.inaf.it/2020/09/29/cryogenic-laboratory-for-x-ray-astronomy-2>

Referenti: C. Macculi, L. Piro, M. D'Andrea

Skills	Facilities	Comments
Cryogenics	<p>1) Pulse Tube + Dilution unit He3/He4 (T_{base} < 10 mK, Cooling Power @50mK = 115 μW)</p> <p>2) Pulse Tube + ADR (T_{base} ~ 40 mK, Cooling energy @50mK = 17 mJ)</p> <p>3) Pulse Tube (T_{base} ~ 2.3 K, Cooling power @4K = 0.4 W)</p>	<ul style="list-style-type: none"> • Transition Edge Sensor work at ~ 0.1 K • Electronics works at ~ 4K • Equipped with PID-type active thermoregulation systems, and dedicated instruments for monitoring and control • Optical inputs up to the lower temperature stages available
EMI/magnetic shielding + filtering	<ul style="list-style-type: none"> • Faraday cage • Magnetic shielding (at cryogenic and warm temperatures) 	<ul style="list-style-type: none"> • The use of the SQUID calls for maximum noise current of $i_n(f) \sim 5 \text{ pA}/\sqrt{\text{Hz}}$ up to few MHz • TESs require a residual magnetic field < μT in the sample space.



DRIVER:

- High resolution spectroscopy in X-ray
- Cryogenic Anticoincidence particle Detector
- Cryogenic and Warm Readout Electronics

REFERENCE MISSION:

Athena. At present the facility is currently 100% committed until 2025.