

## **MRI introduction**

- HIRES initiative
- Hires blue book (2013-2016)
- Hires Phase-A (2016-2018)
- Hires pre-Phase B (2019-2021)

## **Overall concept**

The design of an instrument operating from the blue to the K band requires different detector technologies to be encompassed in the design not to mention the different required temperature regimes, e.g. cryogenic environment in the K band. A modular fibre-fed cross dispersed echelle spectrograph is considered as a promising concept to be deeply studied

#### **General description**

The possible overall concept is summarized in figure 1. The light from the telescope is split, via dichroics, into N wavelength channels. Each wavelength channel includes several telescope optical interfaces that feed, through separate groups of fibres, a dedicated spectrograph module. Each telescope-interface and fibre-bundle corresponds to an observing mode. Based on a preliminary analysis we consider four spectrograph modules as shown in Figure 1 but both the number of spectrograph and splitting in wavelength may change following the dedicated trade-off study during Phase-A.

### **Modules location**

The split in wavelengths over the modules is influenced, among all other parameters by the optical transparency of the different types of fibres available on the market; therefore, the different modules can be positioned at different distances from the telescope focal plane.

## **Observing modes**

All spectrometer modules have a fixed configuration, i.e. no moving parts inside the spectrometers. They include a series of parallel entrance slits, each generated from a separate set of fibres that, in turn, determines the observing mode. The definition of the baseline observing modes will be carried out during phase-A study.

# **Blue book Requirement**

Requirement	HR Single Obj. Mode	MR Multiplexed Mode	HR AO-assisted IFU					
Spectral	>= 100'000	10'000-20'000	>= 100'000					
Resolution								
Multiplexing	1	1-10 (few arcmin FoV)	IFU					
Spectral		0.37-2.5						
coverage (µm)								
Minimum Blue		370 nm						
wavelength								
Allowed	No substantial							
wavelength gaps								
Wavelength	Espresso template for the							
calibration	VISible, TBD for IR							
Stability	10cm/s for VIS (goal	<u>n.a</u> .	10cm/s					
	2cm/2), TBD for IR							
throughput	Espresso template for the	Espresso template for the	Espresso template for the					
	VISible (12%), ECHO VISible (12%), ECHO VISible (12%), E							
	template for IR (8mag s/n	template for IR (8mag s/n	template for IR (8mag s/n					
	10'000 per res element in	10'000 per res element in	10'000 per res element in					
	100min or better)	100min or better)	100min or better)					
Polarimetry	Preferrable	<u>n.a</u> .	<u>n.a</u> .					

#### **Blue book Functional scheme**



#### **Blue Book Architecture**



#### **Blue Book WBS**



### **Blue Book Conclusion**

- Hires can be modulated to provide reasonable fraction of the required science with almost ANY early 1<sup>st</sup> light of ANY ELT.
  - + Hires is feasible
  - + Time adaptability (different module timeline)
  - + Location adaptability
  - + Any Telescope pupil
  - + No AO dependence
  - + No Mass issue

## **Phase A Top Level req**

R-TLR-A.1 Spectral	The Instrument shall provide a spectral resolution of at least 100000	1
Resolution		
R-TLR-A.2 Spectral	The spectrometers shall provide a target sampling of at least 2 pixels per	1
sampling	resolution element. Goal of 3 pixel sampling is desirable.	
R-TLR-A.3 Wavelength	The instrument shall provide simultaneous and as complete as possible spectral	1
coverage	coverage of the 500-1800 nm wavelengths range.	
R-TLR-A.4 Wavelength	The instrument shall be able to achieve radial velocity accuracy better than 1m/s	1
calibration accuracy	between 500 and 1800nm	
and Instrument		
Stability		
R-TLR-A.5 Spectral	For a source of suitable (TBD) brightness it shall be possible to achieve $S/N$ >	1
fidelity on spectra	1000 (TBC) per resolution element in the 1D extracted spectrum from a single	
of bright sources	exposure using daytime calibrations	
R-TLR-A.6 Sky	For single science targets, a sky spectrum shall be recorded simultaneously with	1
Subtraction	the spectrum of the science target.	
R-TLR-A.7 Sky	The instrument shall have seeing limited capability (no AO support is foreseen	1
aperture	from the telescope)	
R-TLR-A.8 Lifetime	The instrument shall have a lifetime of 10 yr (goal 20 yr)	1
R-TLR-A.9 Wavelength	The Instrument shall extend the spectral coverage (defined in R-TLR-A.3) at	2
coverage +	shorter wavelengths down to <b>400 nm (goal 370 nm)</b>	
R-TLR-A.10 Integral	The instrument shall provide an Integral field unit with variable scale up to 2	3
field unit	(TBC) spaxel sampling of the diffraction limit core of the PSF in the J band	
R-TLR-A.11 Wavelength	The instrument shall be able to achieve radial velocity accuracy better than 0.02	4
calibration accuracy	m/s between 400 (goal 370) nm and 670 nm	
and Instrument		
stability +		

 ${}^{[\underline{1}]}$  Will be defined with simulations during Phase B

<sup>[2]</sup> Will be confirmed during Phase B

<sup>[3]</sup> Will be defined during Phase B

#### **Phase A Architecture**



### **Overall deployment**



## <mark>Key parameters</mark>

- + 2 arcmin patrol field of view
- + Wavelength precision > 1m/s
- + Transmission > 7%
- + R=100 000
- + Sampling 4pix Vis, 2.5 pix IR
- + Spectral coverage 400 to 1800 (goal 350 to 2400 nm)
- + Contrast 1000 at 31/D

#### pre phase B Architecture





- Scheme of the current baseline of the system
- Roles in the System Team
- SE Meetings
  - + Tiger team at first (SE meeting slot with custom participants)
- Interactions, modeling
  - + Cameo (with html interface to setup)
  - + Excel (to exchange inputs ad update the model)
  - + Ticketing (Jira/Confluence)
- Next events and Objectives

## TRS, Requirements

[R-HIR- 20] D/A/I/T	The instrument shall have two observing modes: 1. Seeing-limited mode - 2. IFU mode
[R-HIR- 22a] D/A/I/T	The baseline design of HIRES covers the wavelength range specified in [R-HIR-40]. As a goal, the wavelength range will be extended to U-band and/or K-band.
[R-HIR- 255]	Exposure time
D/A/-/T	The instrument shall support individual exposure times between 1 seconds and 900 seconds (goal: 3600 seconds). The on-chip exposure time shall be known to 0.1 seconds, or 1% of the exposure time, whichever is smaller
[R-HIR- 28]	Patrol field
D/A/-/T	The Instrument Field-of-View (area on sky) in seeing-limited mode shall allow picking- up two targets with a separation of up to 2 arcmin (goal: 5 arcmin) and recording their spectra simultaneously.
[R-HIR- 29]	Minimum separation
D/A/-/T	The minimum separation between two objects, which can be fed into the two fiber channels, shall be 15 arcsec (goal: 10 arcsec)
[R-HIR- 30]	Secondary guiding
D/A/-/T	HIRES shall stabilize the de-rotation and centering. The instrument shall allow tracking the object(s) with the pick-off arm(s). The tracking error shall be lower than 50mas rms in the seeing limited mode and lower than 2mas rms in the IFU mode over a timescale of 1 hour.
[R-HIR- 32]	Wavelength precision (TLR-A.12 & TLR-A.20)
D/A/-/T	The Instrument shall achieve a precision in the wavelength calibration of better than 1 m/s RMS (goal: 0.1 m/s) averaged over the whole wavelength range specified in R-HIR-40.
[R-HIR- 33]	Stability of Wavelength Calibration Accuracy (TLR-A.12, TLR-A.17 & TLR-A.20)
D/A/-/T	The instrument shall achieve a short term (24 hours) stability of wavelength calibration accuracy of better than 1 m/s RMS over the wavelength range specified in R-HIR-40. As a goal, the instrument shall achieve a long-term stability of better than 0.02 m/s RMS over a 10 year period between 400 (goal: 350nm) and 670 nm.
[R-HIR- 34]	Transmission (TLR-A.8 & TLR-A.21)
D/A/-/T	The instrument shall have an average transmission >7% with a global minimum >4% (goal: >5%) at wavelengths longer than 400nm (goal: at wavelengths longer than 350nm). The transmission includes the detector quantum efficiency but excludes slit losses due to seeing.
[R-HIR- 35] D/A/-/T	Limiting magnitude (TLR-A.13, TLR-A.21 & TLR-A.22) The instrument shall provide a SNR=10 per spectral resolution element on the extracted spectrum at R=100,000 for a magnitude AB=20 (goal AB=21) in an exposure time Texp=1hr under median seeing conditions (as specified in AD5) at the reference wavelengths of V=550nm, I=850nm and J=1250nm (one for each spectrometer). Note, see [R-HIR-258] for a conversion to the photon flux at the instrument.
[R-HIR- 36]	Spectral Resolution (TLR-A.1)
D/A/-/T	The instrument shall provide an average spectral resolution of R=100,000 with a variation within [-10%,+30 %] across the wavelength range.
[R-HIR- 40]	Spectral Wavelength Coverage (TLR-A.3, TLR-A.11 & TLR-A.19)
D/A/-/T	The instrument in seeing-limited mode shall provide a simultaneous spectral coverage from 400 nm to 1800 nm (goal: 350 nm to 2400 nm). Only one gap, less than 100 nm, centered between 1360 nm and 1410 nm shall be tolerated.
[R-HIR- 259] D/A/-/T	Seeing-limited image quality The instrument shall deliver under the best seeing conditions (seeing<0.3arcsec) an image quality of a point source better than 85% (goal: 95%) encircled energy within 0.4 arcsec radius over the full wavelength range specified in [R-HIR-40].
[R-HIR- 41]	Aperture and Field-of-View (TLR-A.6, TLR-A.7 & TLR-A.10)
D/A/-/T	The instrument in seeing-limited mode shall have at least two sub-FoV, which allow the simultaneous observation of two science targets or one science target and a sky background. The full spectrograph FoV (i.e. the full slit) projected on the sky shall cover more than 1 square-arcsec (goal: 2 square-arcsec).
[R-HIR- 42]	Spectral Wavelength Coverage (TLR-A.15 & TLR-A.19)
D/A/-/T	The instrument in IFU mode shall provide a simultaneous spectral coverage from 980 nm to 1800 nm (goal: 600 nm to 2400 nm). Only one gap, less than 100 nm, centered between 1360 nm and 1410 nm shall be tolerated.

#### **Parameters of the current design**

#### Seeing limited observing modes

IFU Observing Modes (no R&D WPs)

Parameter	value	Options and comments	] [P	Parameter	value	Options and comments			
Resolving power	100 000	Could be increased to 130,000 for RIZ & (U)BV	R	Resolving power	100 000	Same as seeing limited			
Spectral sampling	2.7 (YJH); >4 (U-Z)	pixels	S	pectral sampling	2.7	Pixels			
Simultaneous Iambda coverage	400-1800 nm	Optional extension to U (350-400 nm) and/or K (1900-2400 nm).	Si	imultaneous ambda coverage	950-1800 nm	Only the YJH spectrometer; optional extension to K-band (1900-2400 nm)			
Apertures	2 x D=0.75"	Obj. + sky; ~5% of slit for simultaneous calib.							
		Each spectrometer aperture is uniformly illuminated regardless of input PSF. Each spectral res. element is sampled by many (~200) pixels; detector noise dominates for faint objects		# of spaxels	61	Organized in a hexagonal matrix			
Baseline observing mode	Optimized for spectral fidelity on bright objects			paxel scale	Up to five scales, e.g., 5, 10, 20, 50, 100 mas/spaxel	Scales always available for observation. Option with single scale (mono-mode IFU) proposed by SCAO-IFU WG.			
Optional observing mode	Optimized for faint objects	IFU-like illumination of spectrometer slit, to take advantage of telescope PSF. Entrance apertures mapped on 0.125" spaxels. <u>Important: SCAO correction is</u> <u>not available for this observing mode.</u>	A	AO correction	SCAO	SCAO loop can be closed on the scientific target (on-axis) or on a nearby object, maximum off-axis distance is 3"			

# **Map of Hires**

- Not a functional diagram or a product tree but a visual map to understand the overall system
- Specific diagrams will be generated by MBSE software





## **System Engineering**

#### **Team: Architects**

• Refer

	Ø 0 1_WP_Requirements Management.docx	1 Apri con Microsoft Word						
	Project: HIRES Phase: B WP ID: 1	TBD						
Defen to	WP Title: Requirements Management							
Refer LO	WP Objective:							
the sheets	<ul> <li>Support the System Engineer in the definition of the system requirements flow down</li> </ul>							
for tasks	<ul> <li>Manage and control the flow down of system requirements to subsystem requirements</li> </ul>							
	<ul> <li>Supply the model manager with requirements input</li> </ul>							
and	<ul> <li>Compile the subsystem requirements document w architects</li> </ul>	<ul> <li>Compile the subsystem requirements document with inputs from other architects</li> </ul>						
objectives	WP Manager: Marco Xompero Institute	e: OAA						
5	Deputy: - Deputy i	institute: -						
	Other contributors (with affiliation):							
	Alessio Zanutta (INAF-OABr), Matteo Genoni (INAF-OABr), OABr), Giorgio Pariani (INAF-OABr), Igor Coretti (INAF-OATs OAA), Marco Riva (INAF-OABr)	Matteo Aliverti (INAF- s), Ernesto Oliva (INAF-						
	Duration: T0 + 24 m							
	WP Required Resources							
	Total WP: 0,4 (already included in SE work package)							
	0,2 FTE/yr (already included in SE work package)							
	WP Costs							
	Included in SE WP							
	Required Inputs							
	<ul> <li>Technical Specifications</li> </ul>							
	<ul> <li>– ESO Applicable Documents</li> </ul>							
	Interfaces							
	SE, SWSE (software-SE), s-SE (subsystem-Engineers), other Architects							
	Expected Outputs/Deliverables							
	- Subsystem requirements datapack							
- Requirement tracing matrixes								

Difference btw sSEs and Architects.

Arch.s Present themeseves



SubSys. Eng.s	Manuel Abreu (FE SE)					
[your presence at the SE meetinas is	Andrea Tozzi (FL SE)					
very important]	Michael Weber (UBV SE)					
	Bruno Chazelas (RIZ SE)					
	Mike Macintosh (YJH SE)					
	Enrico Pinna (SCAO-IFU SE)					
	Philip Huke (CU SE)					
	Roberto Cirami (SW SE)					
Architects	Livia Origlia (Op.&Cal. Arch.)					
[you are always welcome to ioin the	Tino Oliva (Opt. Arch.)					
SE meeting, but you	Marco Xompero (Req. Arch.)					
invited when a	Alexandre Cabral (Interf. Arch.)					
in the agenda	Matteo Aliverti (Mec. Arch., Term. Arch.)					
requires your expertise]	Igor Coretti (Ele. Arch.)					
	Giorgio Pariani (AIV Arch.)					
	Marco Riva (Model Arch.)					

## Tiger Teams

- For initial Tradeoffs and important activities
- Themes:
  - + Monomode IFU, AO
  - + Impact of K band (cryo stuff in IFU?)
  - + Consequences on the Calibration Unit
  - + AoB
- Timeslot SE biweekly (odd weeks) next 21 gen 2022 @ 14:00 CET



Customization for Astronomical Instrument Model Managing



④ 100% ≎ 🔍 🔍 🖼

## **Requirement Managing**

#### Example, not HIRES

#	riangle Name	Text	Property	Value	Margin
1	TRS #1 ABA_version				
2	🗆 🛅 Derived Requirements				
3	🗆 🛅 Derived Requirements				
4	🚔 TRS-1.1.1 sub-sub-system R1	the subsubsystem mass must not exceed 3 kg			
5	🚔 TRS-1.1.2 sub-sub-system R2	the subsystem mass must not exceed 5 kg			
6	🖀 TRS-1.1 sub-system R1	the subsystem total mass must not exceed 7 kg			
7	🖀 TRS-1.2 sub-system R2	the system mass must not exceed 6 kg			
8	🚔 TRS-1 TRS #1	Total mass must not exceed 16 kg	🔽 totalMass : Real	30	-14
9	TRS #2 AZA_version				
10	🗆 🛅 Derived Requirements				
11	6 Derived from TRS #2.1	a solar panel shall be implemented to compensate excessive power loads			
12	10 Derived from TRS #2.1	the component 1 must allow sleep mode			
13	□ 🚔 TRS-2 TRS #2	the power consumption must not exceed 30 W and the plug must be EU standard	✓ totalPower : Real	10	20
14	🖻 🖀 TRS-2.1 TRS #2.1	the power consumption must not exceed 30 W			
15	🖀 TRS-2.1.1 TRS #2.1.1	the power consumption of component 2 must not exceed 10 W			
16	🚔 TRS-2.1.2 TRS #2.1.2	the power consumption must not exceed 19 W			
17	🖻 🚔 TRS-2.2 TRS #2.2	the plug must be a Schuko			
18	TRS-2.2.1 other	other			



# **CAMEO export-inport**

#### Via excel sheets stored in owncloud

•••	🖳 Cameo Systems Modeler 2021x - HIRES_mode	l.mdzip [/Volumes/GoogleDrive/Drive c	ondivisi/HIRES/MODEL/]								
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🔁 Containment 🐉 Diagrams 🖥 Structure	😹 HIRES product tree 📓 TRS #2 📓 TRS #1 📰 TRS Co	ompliance Matrix 🗙 🌆 Requirements Verific	atio Requirement Deri	ivation M   📾 Model	4 4						
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TRS #1 ABA_version	Excel Import Status: 🔄 New 🗌 Updated 🔲 Obsolete 🗋 Unchanged										
Relations	# Name	Text	Document Reference	Achieved Value	Compliance Ph						
Requirement Derivation Map	1 🗆 🛅 TRS #1 ABA_version										
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go to sleep after 10 min of inactivity	3 🖃 🛅 Derived Requirements			Home	Inserisci Disegno Layout di pagina F	- Formule Dati Revisione	Visualizza 🔉 Dimmi				🖻 Condividi 🛛 🖵 Commenti
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HR HIRES	7 🊔 sub-sub-system R2	the subsystem mass must <u>not</u> exceed 5 kg			А	В	c	D	E	F	G
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HIR-CUN CU	9 🗆 🛅 TRS #2 AZA_version			2	Name	Text	Document Reference	Achieved Value	Compliance	Phase- A verification	on Phase- B verification
HIR-FEN FE HIR-FLI FL HIR-FLI FL HIR-FLI FL HIR-RIZ RIZ	the power consumption must j exceed 30 W and the plug mu be EU standard	the power consumption must <u>not</u> <u>exceed 30</u> W and the plug must be EU standard	link to confluence page that contains the analysis: https://docs.nomagi c.com/display/MD20 21x/Web+Publisher +2.0+report	TRS #1	l stem R1	Total mass must not exceed 16 kg the subsystem total mass					
				sub-su	ıb-system R1	must not exceed 7 kg the subsubsystem mass must not exceed 3 kg					Analysis
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Documentation of Diagram TRS Compliance Matrix	12 🚔 TRS #2.1.1	the power consumption of component 2 must <u>not exceed 10</u>		sub-sy 7	rstem R2	the system mass must not exceed 6 kg					
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	14 🖂 🚔 TRS #2.2	the plug must be a Schuko				standard	isplay/MD2021x/Web+Publi				
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# **HTML interface**

- As reference for:
  - + Requirements
  - + Structure, PT, numbering, BoM
  - + Interfaces
  - + Functions, Actions
  - + Budgets



Horizon Europe Call for instrument

development funding

<mark>(aka OPTICON22)</mark>



## **Next Flow of Events & Objectives**

- test of the communication lines and interactions (SE team with subsystems)
  - + Collaborative tools implementation, usage directions
- Consolidation of the Architecture (from Tech Specs and Science Cases to:)
  - + Critical items identification
  - + subsystem feasibility assessments
  - + PT, numbering
  - + Interfaces definition between subsystems
  - + subsystem Req.s flow-down
- System Design (definition of the System from the Architecture)
  - + Build the Drawings basing on the Architectural scheme