







The ASTRI Mini-Array

S. Scuderi – IASF Milano for the ASTRI Project

OAS Very High Energy meeting: towards ASTRI and CTA, 08-09 June 2022









Layout of the presentation

- The ASTRI Mini-Array project
- Architecture of ASTRI Mini-Array
- Teide infrastructure
- Status of production:
 - > Mechanical structure & mirrors
 - > Telescope's auxiliaries
 - Cherenkov cameras

 - > Software
- Operation concept
- Mini-Array implementation





The ASTRI Mini-Array Project

ASTRI (Astrofisica con Specchi a Tecnologia Replicante Italiana) was born as "Progetto Bandiera" funded by MIUR with the initial aim to design and realize an innovative end-toend prototype of the 4 meters class telescopes in the framework of the CTA observatory

The ASTRI Mini-Array is the second step of project whose purpose is to construct, deploy and operate an array of 9 Cherenkov telescopes of the 4 meters class at the **Observatorio del Teide in Tenerife (Spain) in collaboration with IAC.**

More than 150 hundred researchers belonging to • INAF institutes (IASF-MI, IASF-PA, OAS, OACT, OAB, OAPD, OAR) Italian Universities (Uni-PG, Uni-PD, Uni-CT, Uni-GE, PoliMi)

- Spain).

Italian and foreign industrial companies are and will be involved in the ASTRI Mini-Array project with important industrial return.



• International institutions (University of Sao Paulo – Brazil, North-West University – South Africa, IAC

The ASTRI Mini-Array Project

- The ASTRI mini-array can be considered a new pathfinder of the arrays of Cherenkov telescopes
- Hosting agreement foresees 4 + 4 years of operations for the ASTRI Mini-Array starting from beginning of operations
- During the first 3/4 years of operations the array will be run as an experiment
- The ASTRI Science team will develop a strategy to concentrate the observational time on a limited number of programs with clearly identified objectives (Vercellone's talk)
- After this initial period the project will gradually move towards an observatory model in which a fraction of the time will be assigned to scientific proposals through a Time Allocation Committee procedure



Mini but not small...



PoS(ICRC2021)884 (2021)

Wide-field stereoscopic observations in the 1 – 300 TeV energy band

- Restricted number of targets/deep exposures (\geq 200 h)
- Galactic sources: wide FoV \rightarrow multi-target fields
- Extragalactic sources: survey of a few promising targets at $> \sim 10$ TeV scale
- Fundamental physics: studies on LIV, EBL, Axion-Like Particles, ...
- Stellar Hambury-Brown intensity interferometry in the visible band
- Direct measurements of cosmic rays



Largest Imaging Atmospheric Cherenkov Telescopes facility until CTAO will start to operate

Expected performance

• Sensitivity: better than that of current IACTs (E > a few TeV):

- Extend the spectra of already detected sources and/or measure cut-offs • Energy/Angular resolution: ~ 10% / ~ 3' (E > a few TeV)
 - Characterize the morphology of extended sources at the highest VHE

• Wide FoV (≥ 10°), with almost homogeneous off-axis acceptance

- Optimal for multi-target fields, surveys, and extended sources
- Enhanced chance for serendipity discoveries









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Operation modes

- moderate moonlight conditions. Calibration activities are included in the normal operation mode.
- vary from 1 hour to 1 day.
- those observations, will be scheduled well in advance.
- (e.g. network, data center, etc). This is the only daytime operation mode.



• Normal (science) observation mode: this is used to observe the targets as defined by the Science Operation Plan. Usually science observations require dark time, although it is also possible operate also during

• ToO Mode: the science operation plan will identify some astrophysical targets (either a specific object or a class of objects) that, giving raises to transient phenomena, will require a response from the night operator and a change in the night schedule. This means that no dedicated automatic software procedure to react to these transient phenomena is foreseen. Depending on the type of transient object the reaction time will

• Coordinated Mode: Synergies with the current VHE arrays (MAGIC, LST) in the northern hemisphere are foreseen in the science operation plan. This means that simultaneous observations will be possible. Usually,

• Maintenance mode: this mode deals with all activities necessary for the maintenance of the telescopes, the on-line control software, the monitoring, characterization and calibration devices, and the infrastructures

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Science Operations

- No real time analysis of the data is foreseen but only a data quality check. Data analysis policy adopted will then be next day processing.
- No array trigger (stereo trigger) will be implemented at the site. Any search for Cherenkov events detected in coincidence by more than one telescope will be performed via software off-line at the Rome Data Centre.
- No subarray operation is foreseen.
- Night science operations will be controlled remotely from La Laguna @ IAC \rightarrow no people required at the site during the night.
- The local control room at the Themis Observatory will be used during commissioning and science verification phase, during maintenance activities or in case of other special activities.
- Other Array Operation Centers (control room) from Italy.



Duty cycle for Cherenkov observations

Moonless Night Hours Fraction of clear nights (cloud coverage < Fractional loss due to bad weather Fractional loss due to "Calima"

Average Annual Observation Time





	1565 h
<20%)	0.79
	0.04
	0.07
	1104 h

Setting 15 NSB as limit \rightarrow AAOT \sim 1800 h

Duty cycle: 21%

The ASTRI Mini-Array locations

The ASTRI Mini-Array in Tenerife

- Telescope Array & auxiliaries (Observatorio del Teide OT) Data Centre in Rome
- Local Control Room @ THEMIS building (OT)
- On site Data Centre @ IAC Teide Residencia (OT)
- Array operation center @IACTEC in La Laguna





The ASTRI Mini-Array in Italy

Remote Array operation centers

The ASTRI Mini-Array architecture: PBS



Infrastructure: composed by all those parts needed to make the observational site suitable to host the telescopes of the ASTRI Mini-Array. Safety & Security: an independent system for the protection of people and site assets Telescopes: include mainly the hardware used to collect and image Cherenkov light from air showers and the auxiliary assemblies needed to support this function.

ICT: includes all computing/storage hardware, the overall networking infrastructure (including cabling and switches) and all system services (operating system, networking services, name services, etc.) necessary on site and off site to control and monitor the array and to archive and analyse the scientific and engineering data.

Software: The Mini-Array software will provide to the user a set of tools from the preparation of an observing proposal to the execution of the observations, the analysis of the acquired data online and the retrieval of all the data products from the archive. **Monitoring, Characterization and Calibration**: the set of devices that allows the environmental monitoring the atmospheric characterization and the array calibration.

Logistics Support: includes all the hardware & software necessary for the preventive and corrective maintenance of the ASTRI Mini-Array.



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The ASTRI Mini-Array architecture: context diagram **Mini-Array**



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Time



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The Teide Infrastructure

- Civil Work (including foundations for telescope and auxiliaries, roads, trenches)
- Power supply network (including transformer station, UPS and emergency power generator)
- Telecommunication network
- Control room @ Themis
- **Onsite Data Centre**
- Service cabinets
- Facilities in La Laguna











Infrastructure: foundations







Infrastructure: Telescope's area







Service cabinet







Infrastructure: Power network









Infrastructure: Telecommunication network







Infrastructure: Care of environment





- Planting new specimen



Status of the infrastructure

- All telescope foundations completed
- Excavations and tubes for power and data network completed
- Control room @ Themis completed
- Data Centre completed by half June
- Transformation Centre to be completed by the end of June with UPS and Power diesel generator
- Electric cables & optical fibres being laid





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Status of production: ASTRI-1 integration







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ASTRI-1: disassembling, packing and transportation







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Telescope's auxiliaries



Pointing Monitoring Cameras (Uni-PG)

 \rightarrow First three delivered at Italy integration site \rightarrow PMC of ASTRI-1 shipped mounted on the telescope



CCD camera placed on the M2 support structure used to monitor pointing and tracking performances of the telescope.



Optical Camera (IASF-MI, OAPD, OACT, OA Brera) \rightarrow Ready to ship



CCD camera placed on the telescope focal plane to align the panel of M1









Cherenkov Cameras

Design & simulation



Camera thermal model: evolution of SiPM temperature (T_{env}=25 °C)



Focal plane support structure

Ongoing contract for the construction of 11 Cherenkov Cameras

Prototyping







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Production









Delivered 346 out of 450





ICT – On site Data centre



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Virtual Telescope Control System: the system hosting the virtual machines that will be used for the telescopes control.

Camera Servers: are the physical servers, one for each telescope, for the Cherenkov camera and stellar intensity interferometry data acquisition.

Computing System: is the set of physical servers dedicated to the on-line analysis of scientific data for quality check and of monitoring data for the alarm management.

Storage System: is the collection point of the raw scientific data, of the monitoring and of the alarm data. It also the location from where all these data are accessible for remote transfer and for all on-site uses.

Network System: is the set of devices responsible for internal and external network connections.

Tender still to be issued \rightarrow documentation ready



mini-ICT

m-ICT

- Reduced version of the onsite ICT
- Allow to run single telescopes
- Preliminary installation @ Themis completed (F. Gianotti & coll.)
- Final installation @ data centre \rightarrow week of 13th of June







Control room ICT architecture







ICT – Offsite data centre





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A distributed archive solution will be adopted for the ASTRI Mini-Array archive with three main nodes

INAF – OAR: central storage and computing nodes

- INFN LNF: for access to Data Grid infrastructure 2.
- ASI SCDC: to interface high level ASTRI Mini-3. Array science archive with multiwavelength SCDC facilities and science tool services







Software: architecture





- **Supervisory Control And Data Acquisition (SCADA) System** The software system devoted to control all the operations carried out at the Mini-Array site, including the startup of the Mini-Array system. SCADA is a central control system which interfaces and communicate with all equipment and dedicated software installed On-Site.
- **Archive System** The software service that provides storage and organization for all data, data products, and metadata generated for and by the Mini-Array, and defined by the Mini-Array Data Models.
- Data Processing System The software system used to calibrate and reduce \bullet the data acquired. This software is also used to check the quality of the final data products.
- **Science Support System** The software system which provides the main point of access for the exchange of science-related data and information with the ASTRI Science Users, and which supports the whole sciencerelated workflow, from the Observing Project submission to the access to the archived high-level Mini-Array science data products and the corresponding Science Tools to support data analysis.
- **Simulations System** The software system that runs Monte Carlo simulations to provide simulated data for the development of reconstruction algorithms and for the characterization of real observations.
- *Local Control Software* Firmware and low-level software dedicated to the low-level hardware control operations.















Software: data & information flow



Interaction with actors



The ASTRI Mini-Array software is envisioned to handle an observing cycle, i.e. the end-to-end control and data flow system. The observing cycle can be divided into the following main phases:

- **Observation preparation** 1.
- **Observation execution** 2.
- 3. Data Processing
- Dissemination 4.

Software development mainly by INAF

Support by external companies/research institutions on specific aspects as for example:

- Central Control System by Universidad Tecnica Federico Santa Maria (SCADA)
- Operator Human Machine Interface by University of Geneve (SCADA)







Implementation phases: schedule











Implementation phase 0 - Summer 2022





- ASTRI-1
- Transformer Station
- Control Room
- Data Center (m-ICT)
- Meteo station 1
- Power & Telecommunication Network



Implementation: populating the array









ASTRI Mini-Array



View from Vacuum Tower Telescope



