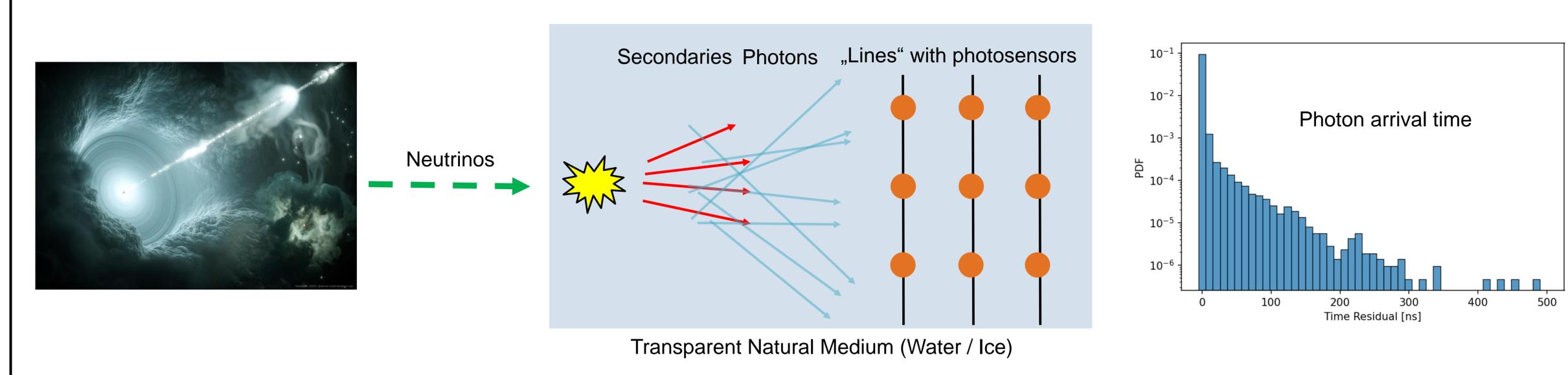
Machine-learning based detector optimization of the future P-ONE neutrino telescope

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Neutrino Telescopes



Science:

Multi-messenger astronomy

P-ONE

 Astrophysical accelerators Galactic & Extra-Galactic

• Multi-disciplinary science

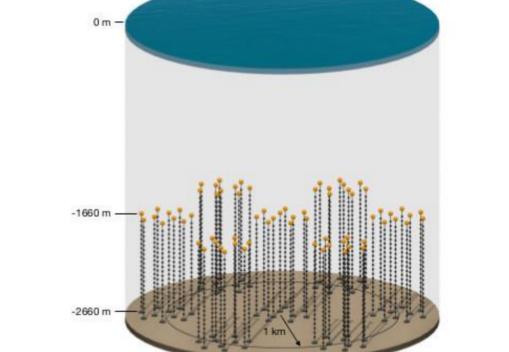
- Transient phenomena
- Cosmic ray transport
- BSM-Physics

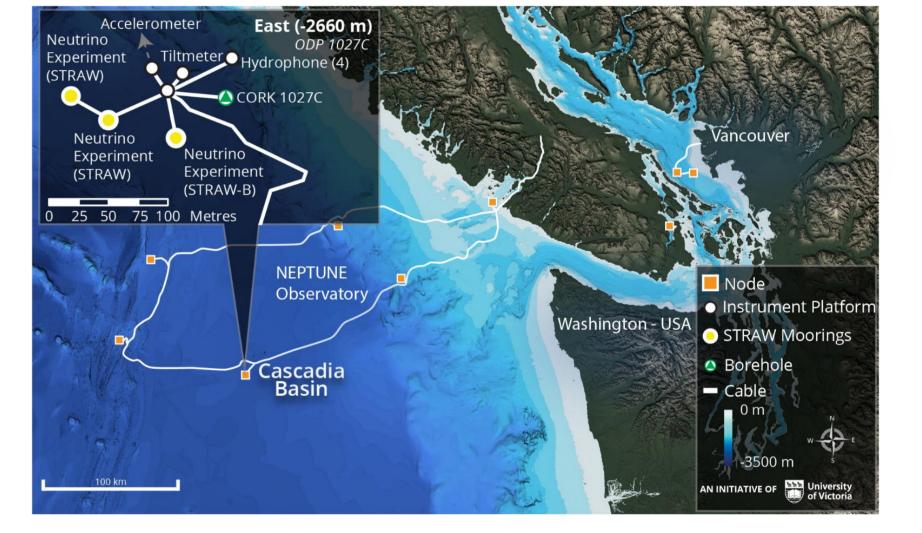


- Next-generation neutrino telescope in the Pacific Ocean ullet
- Vision for complete detector: around 70 mooring lines, multi-km³ instrumented volume
- P-ONE-1 (1st detector line) with 20 modules and total length of 1000 m ullet



- Sensor spacing (vertical / horizontal)
- Cluster structure
- Trigger





Integrated into deep-sea infrastructure: **Ocean Networks Canada**

P-ONE-1

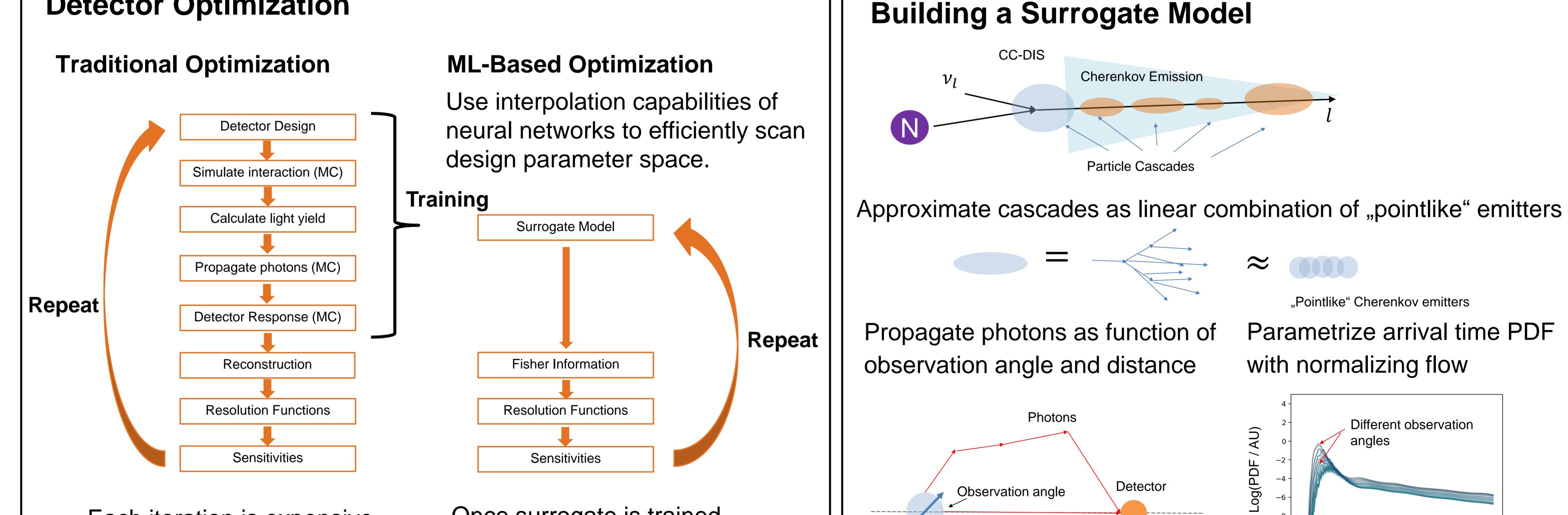
Design Optimization

Vision

Emitter

[1] Agostini, M., Böhmer, M., Bosma, J. et al. The Pacific Ocean Neutrino Experiment. Nat Astron 4, 913–915 (2020). https://doi.org/10.1038/s41550-020-1182-4

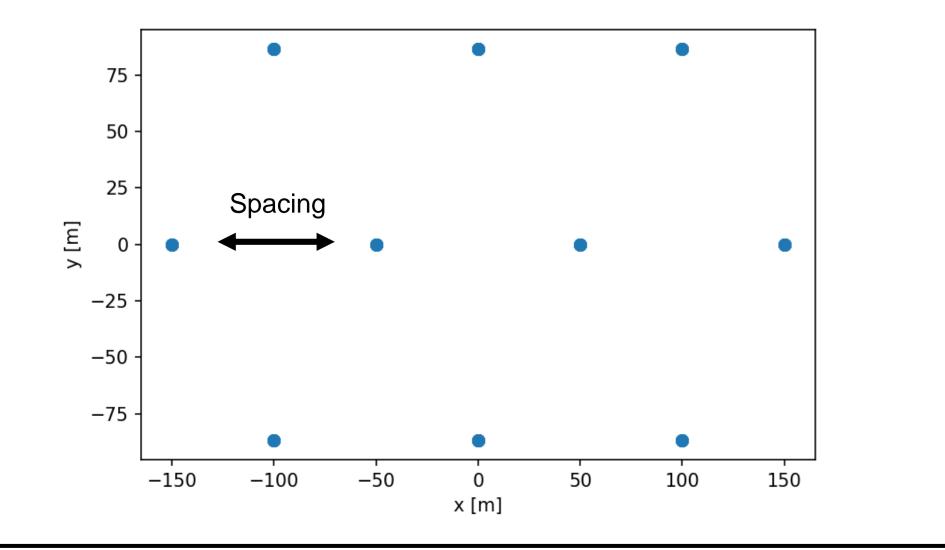
Detector Optimization



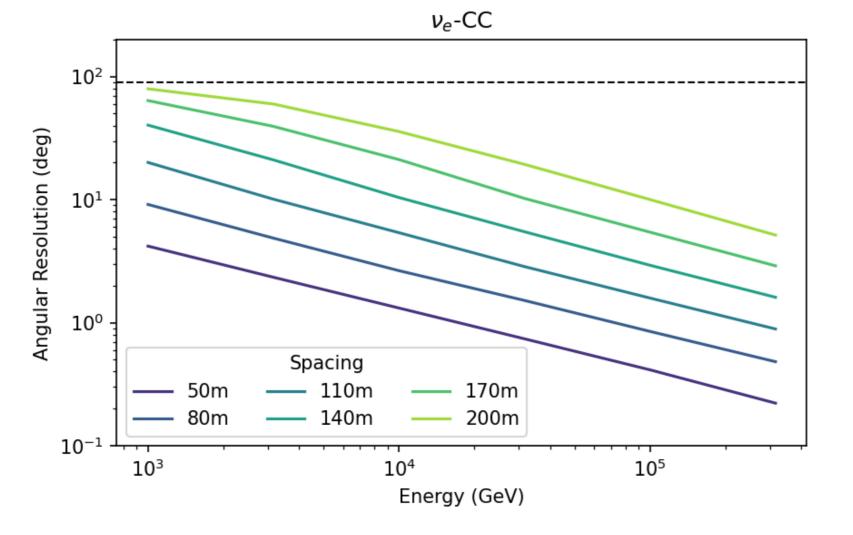
Each iteration is expensive (1000's of CPU & GPU hours) Once surrogate is trained, iteration is cheap

Example Optimization Results for Toy Detector

Toy detector with 10 lines and 20 modules per line, 50m vert spacing



Statistical limit of electron neutrino angular resolution



Future Plans

Extend optimization workflow to all detection lacksquarechannels

Photon arrival time (ns)

- Integrate resolution functions into analyses framework
- Define figures of merit (physics sensitivities)
- Optimize detector design with external constraints (site limitations, budget, ...)