# Deep neural networks for single-line event direction reconstruction in ANTARES

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## **ANTARES** in a nutshell

First large undersea neutrino telescope [1]:

- 12 vertical lines forming a 3D array of photo-sensors
- 25 floors per line
- 3 Optical Modules (OMs) per floor

It detects Cherenkov light induced by secondary particles from neutrino interactions

## Problems

The event reconstruction and background discrimination is challenging, specially for low-energy single-line events.

• Azimuth angle reconstruction is missing with classical approaches (BBfit) [2].

• The low-energy range is interesting for dark matter searches and other studies.

## Neural Network approach

We combined Deep Convolutional Networks (DCNs) [4] with Mixture Density Networks (MDNs) [5].





 MDNs allow to predict not only the direction angles but also their uncertainty.

ELU+1: σ output

The network was trained with Monte Carlo simulations [6] of track-like events: (anti)muonneutrino interactions:

- Training set (60%)
- Validation set (20%), to avoid over-fitting through Early Stopping
- Test set (20%), to evaluate the generality of the results

	Input (25,161,3)	Convolution 2D + ReLU Filters 16, Kernel 2x10	Max Pooling 2D Kernel 2x2	Convolution 2D + Re Filters 32, Kernel 2x		
/	Flatten + DropOut (0.2)	Max Pooling 2D Kernel 2x2	Convolution 2D + ReLU Filters 128, Kernel 2x10	Max Pooling 2D Kernel 2x2		
	Dense + ReLU + Batch Normalization. Size 128	Dense + ReLU + Batch Normalization. Size 128	Dense + ReLU + Batch Normalization. Size 32	Dense + ReLU + Bat Normalization. Size 3		

## From data to images

Two-dimensional images are created (floor x time) using the hits information. Images are then centered based on the reference hit.



The RGB colors of the pixels are obtained by weighting the angle of the OM with the voltage amplitude of the hit.

## Results

### DCN outperforms classical reconstruction methods

Absolute

error



### Our predictions are unbiased against the energy of the neutrino







Moreover, our uncertainty parameter behaves as expected for a standard deviation value.

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Mean

Bfit		DC	CN	<b>DCN</b> (50% lowest $\sigma$ )				
	Median	Mean	Median	Mean	Median			
	8.5°	7.4°	4.4°	3.7°	2.5°			
	-	41.4°	31.5°	29.3°	23.6°			
	-	- 28.1°		18.3°	13.2°			

DCN

DCN

### We checked the robustness of our results using 5-fold cross-validation

or		Randomized						Date-sorted						
Mean Absolute er		Fold 1	Fold 2	Fold 3	Fold 4	Fold 5		Fold 1	Fold 2	Fold 3	Fold 4	Fold 5		
	Zenith	7.4°	7.4°	7.4°	7.4°	7.4°	Zenith	8.1°	7.6°	7.4°	7.5°	6.9°		
	Azimuth	41.5°	41.3°	41.6°	41.4°	41.3°	Azimuth	44.4°	42.2°	41.4°	41.7°	39.7°		
	Total	28.2°	28.1°	28.4°	28.1°	28.1°	Total	30.0°	28.7°	28.4°	29.1°	26.3°		
	Total	28.2°	28.1°	28.4°	28.1°	28.1°	Total	30.0°	28.7°	28.4°	29.1°	26.3		

### Network predictions derived from background noise

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	0.02					AN <sup>-</sup>	TARE	ES pr	elin	ninary		-		
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	0.00	Ó	1	0	2	0	3	0		40	5	60		
Zenith absolute error (°)														





## Conclusion

**References:** 



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Predicted uncertainty (°)

As expected the network is not able to predict the direction using background noise and the predicted uncertainty is high.

### Reconstructing other types of simulated events

Good agreement between real data and simulations, especially when cutoffs based on quality parameters were applied.

Accurate predictions

despite these events

were not explicitly

trained.

for their direction

and uncertainty,

### • We clearly outperform classical reconstruction methods for single-line events. • An azimuth prediction is made, which was previously missing. • The method is robust and ready to be applied in physics analyses.

[1] ANTARES collaboration, Nucl. Instrum. Meth. A 656 (2011) 11 [arXiv:1104.1607] [2] ANTARES collaboration, Astropart. Phys. 34 (2011) 652 [arXiv:1105.4116]. [3] J. García-Méndez et al, JINST 16 (2021) C09018. [4] J. Gu et al., Pattern Recognition 77 (2018) 354. [5] C.M. Bishop, Neural Computing Research Group NCRG/94/004, (1994). [6] ANTARES collaboration, JCAP 01 (2021) 064 [arXiv:2010.06621].