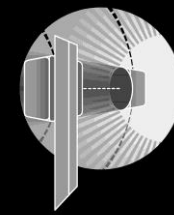


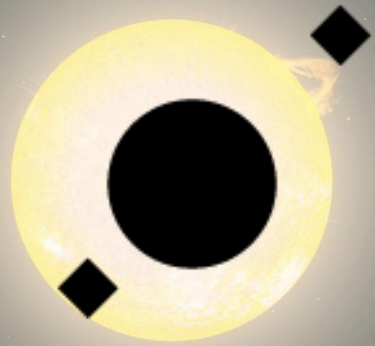
SOLI INVICTO



proba-3



esa



**INAF**

ISTITUTO NAZIONALE  
DI ASTROFISICA

# PROBA-3/ASPIICS Coronagraph



*Silvano Fineschi*  
*INAF Osservatorio Astrofisico di Torino*

**Audizione RSN-3 Sessione 8**  
**28 May, 2021**

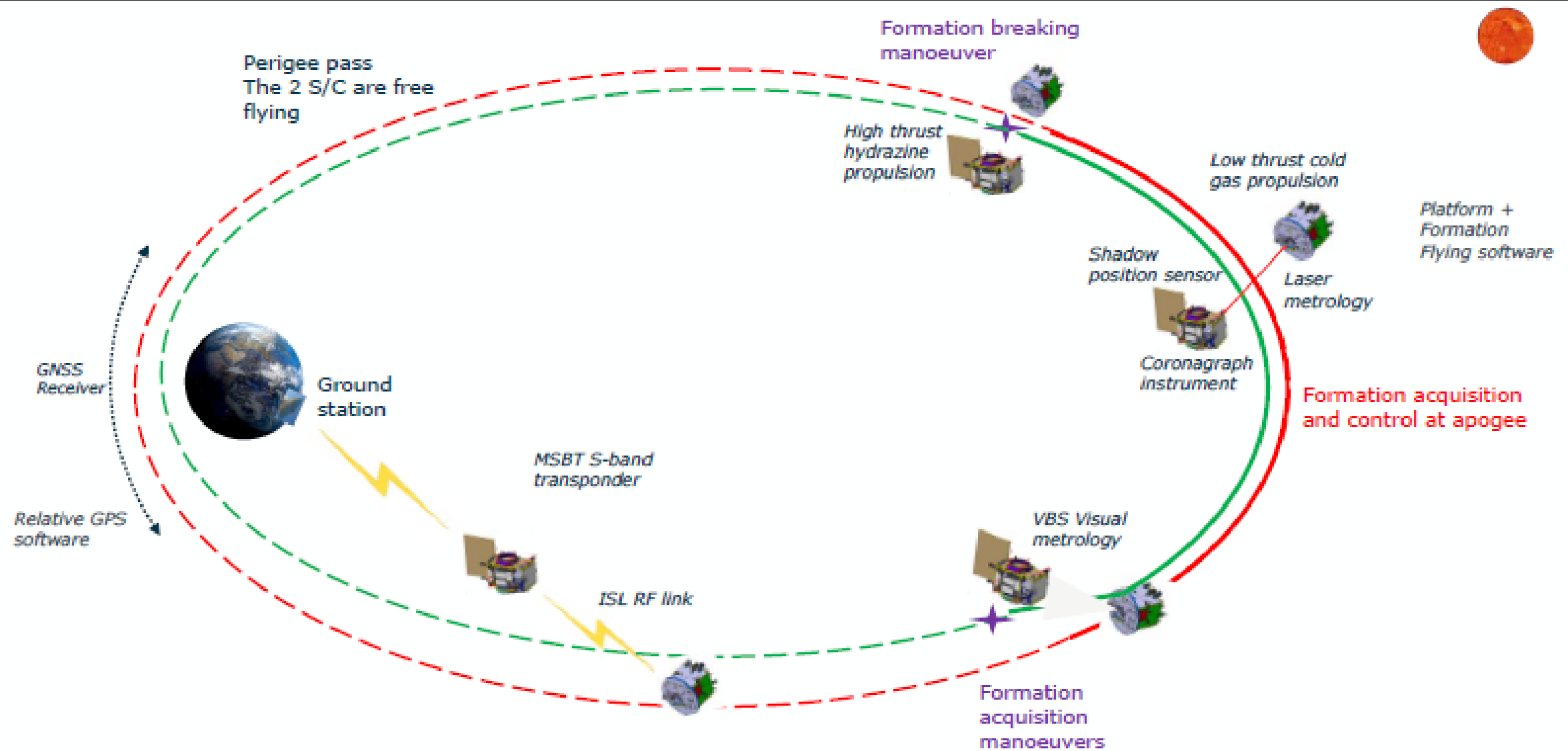
# Heliophysics in a nutshell

- Heliophysics is the science that unites all of the linked phenomena in the region of the cosmos influenced by the Sun:
  - \* solar physics,
  - \* space physics
  - \* heliospheric physics,
  - \* physics of planetary magnetospheres, etc.
- Objects under study:
  - \* Sun and its corona,
  - \* solar wind,
  - \* solar cosmic rays,
  - \* plasma environment of planets.
- Measurements performed:
  - \* remote-sensing measurements,
  - \* in situ measurements.



(movie courtesy B. Nicula & the JHelioviewer team at ROB)

# ESA PROBA-3: 1<sup>st</sup> Formation-flying Mission Technology Demonstrator



***A coronagraph with the external occulter on one spacecraft and the optical instrument on the other spacecraft at ~150 meters from the first one.***

***Relative transverse positioning accuracy:  $\pm 1.5$  mm***

# PROBA-3 Development Timeline

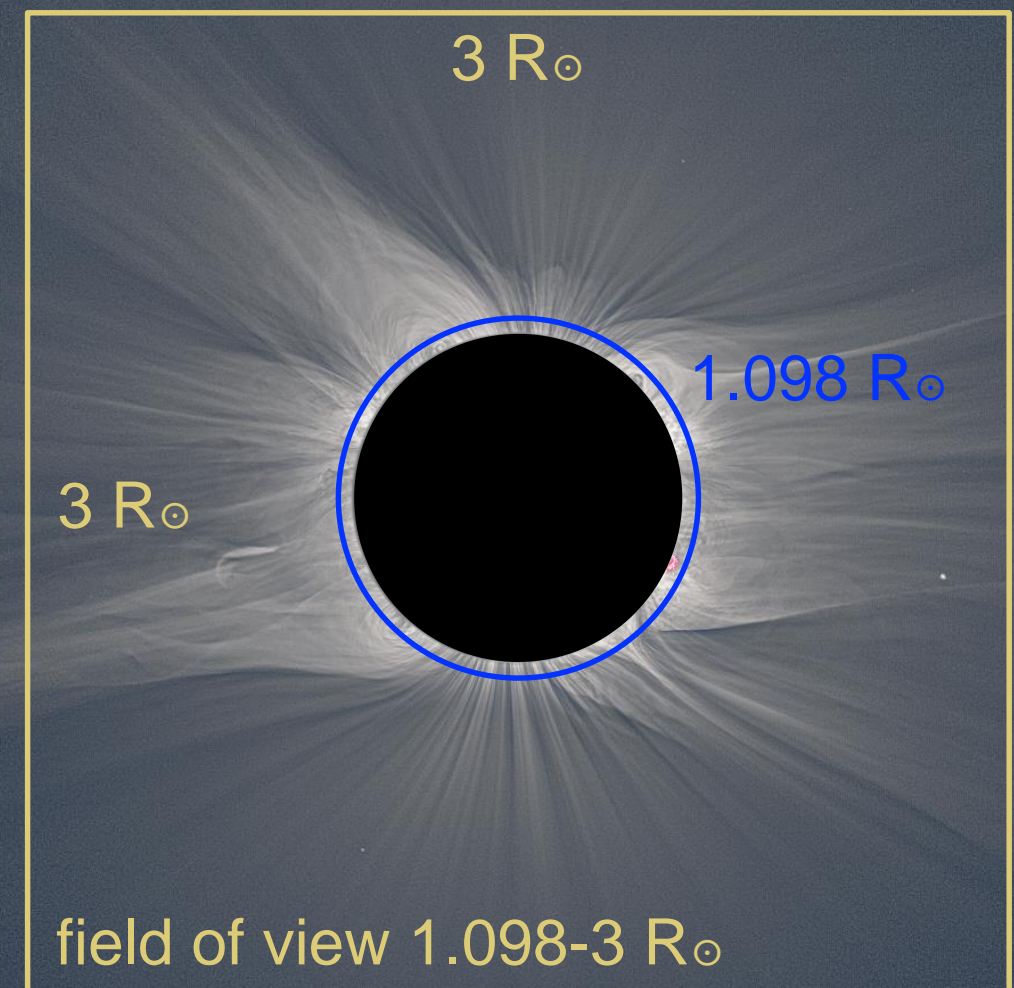
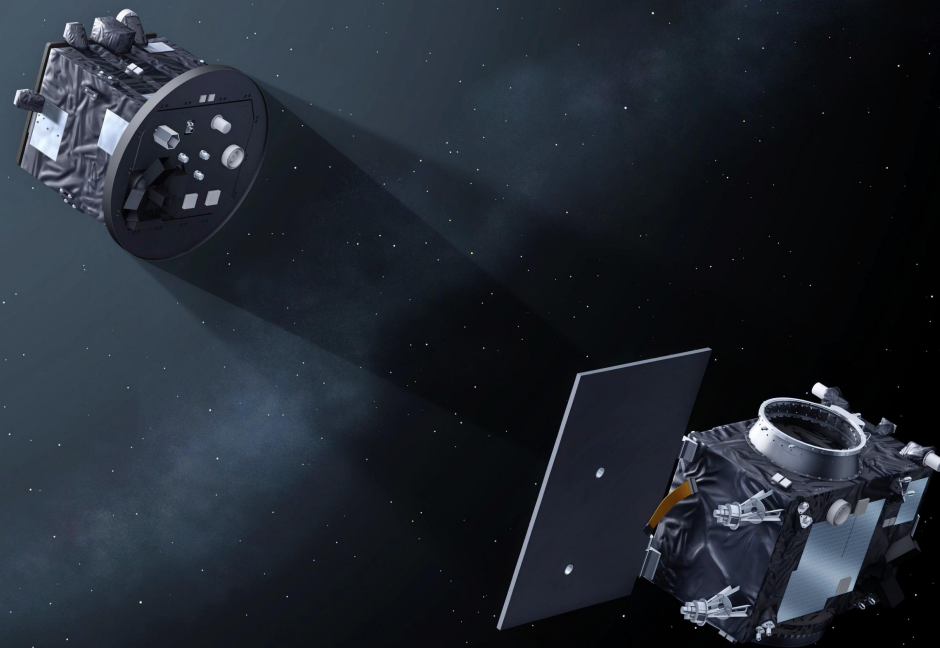
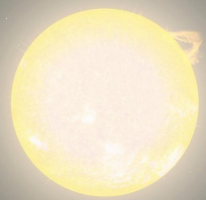
- 2005 **Technological** Mission concept study
- 2009 Phase A studies - “Startiger”
- 2012 Phase B
- 2014 Phases C/D/E1 (change of industrial organisation)
- 2018 System CDR
- 2017 Mission of opportunity adoption by the ESA **Science** Programme.
- 2021 ASPIICS delivery to PROBA-3 S/C
- 2023 June Launch
- 2023 + 6 months End of commissioning
- 2025 End of nominal mission

➤ **INAF involved since Phase A (“Startiger”) 2009**



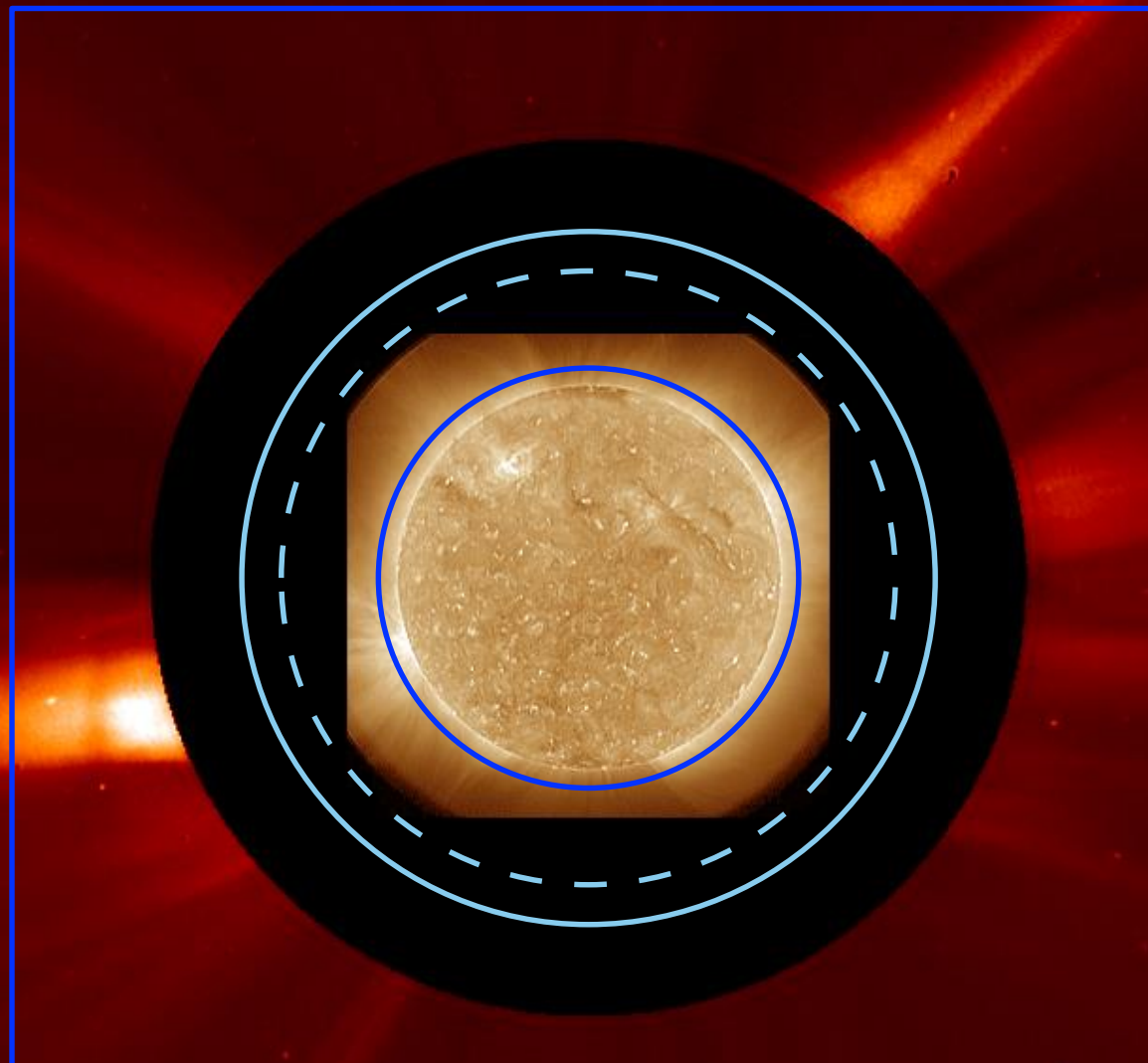
# PROBA-3/ASPIICS: the ultimate coronagraph!

- The formation flying will be maintained over 6 hours in every 20-hour orbit: *around a factor 100 improvement* in the duration of uninterrupted observations in comparison with a total eclipse.
- PROBA-3 will observe the corona two orbits per week on average: *around a factor 50 improvement* in the occurrence rate in comparison with a total eclipse.
- 6 spectral channels:
  - white light (5350-5650 Å),
  - 3 polarized white light,
  - Fe XIV passband at 5304 Å.
  - He I D3 passband at 5877 Å.
- 2048x2048 pixels (*2.8 arc sec per pixel*)
- *60 s nominal synoptic cadence*
- 2 s using a quarter of the field of view.



The PROBA-3/ASPIICS coronagraph will examine the structure and dynamics of the corona in the crucial region that is difficult to observe.

**ASPIICS (Association of Spacecraft for Polarimetric and Imaging Investigation of the Corona of the Sun)** will fill the gap between the typical fields of view of EUV imagers and externally occulted coronagraphs.



SDO/AIA: below  $1.27 R_{\odot}$   
above  $2.2 R_{\odot}$

SOHO/LASCO C2:

PROBA-3/ASPIICS:  $1.098 - 3.0 R_{\odot}$

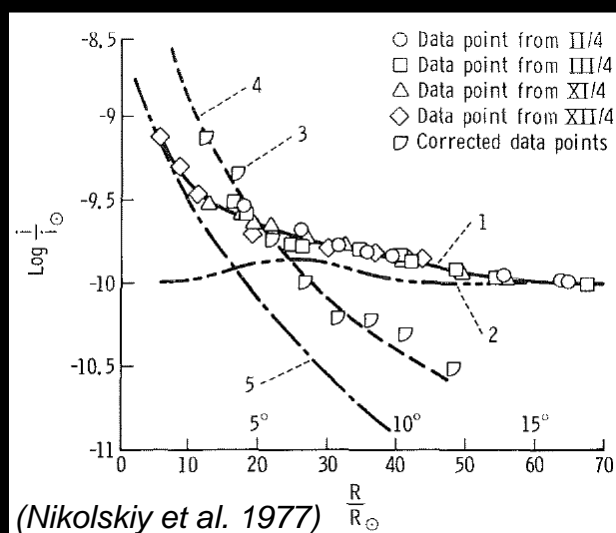
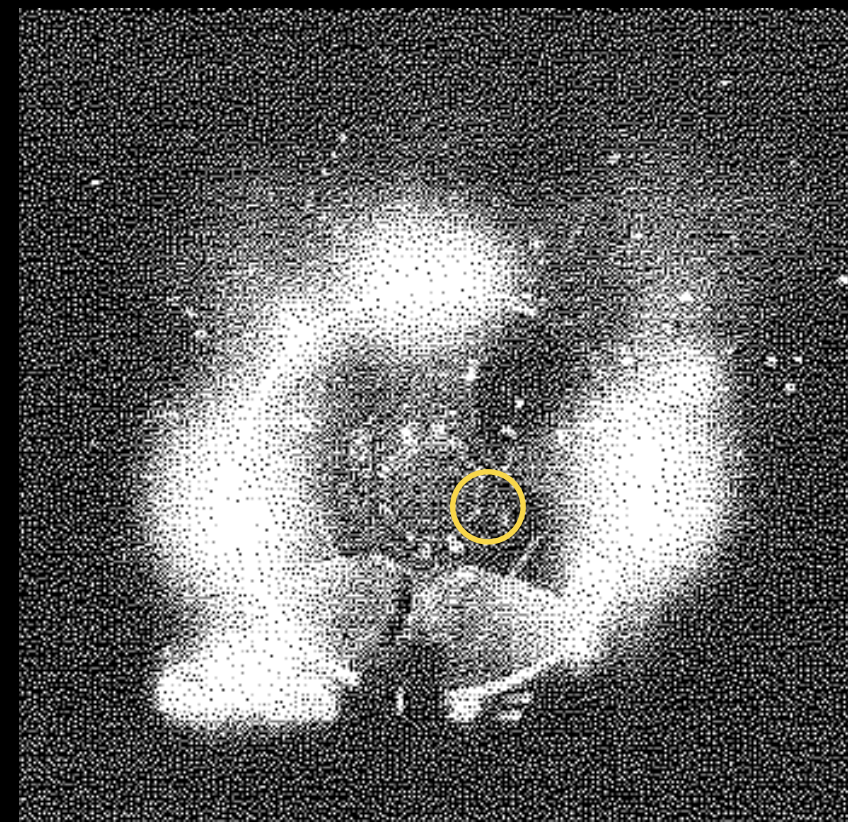
Solar Orbiter/Metis: above  $1.8 R_{\odot}$  (at the first perihelion, solid)

Solar Orbiter/Metis: above  $1.6 R_{\odot}$  (at the closest perihelia, dashed)



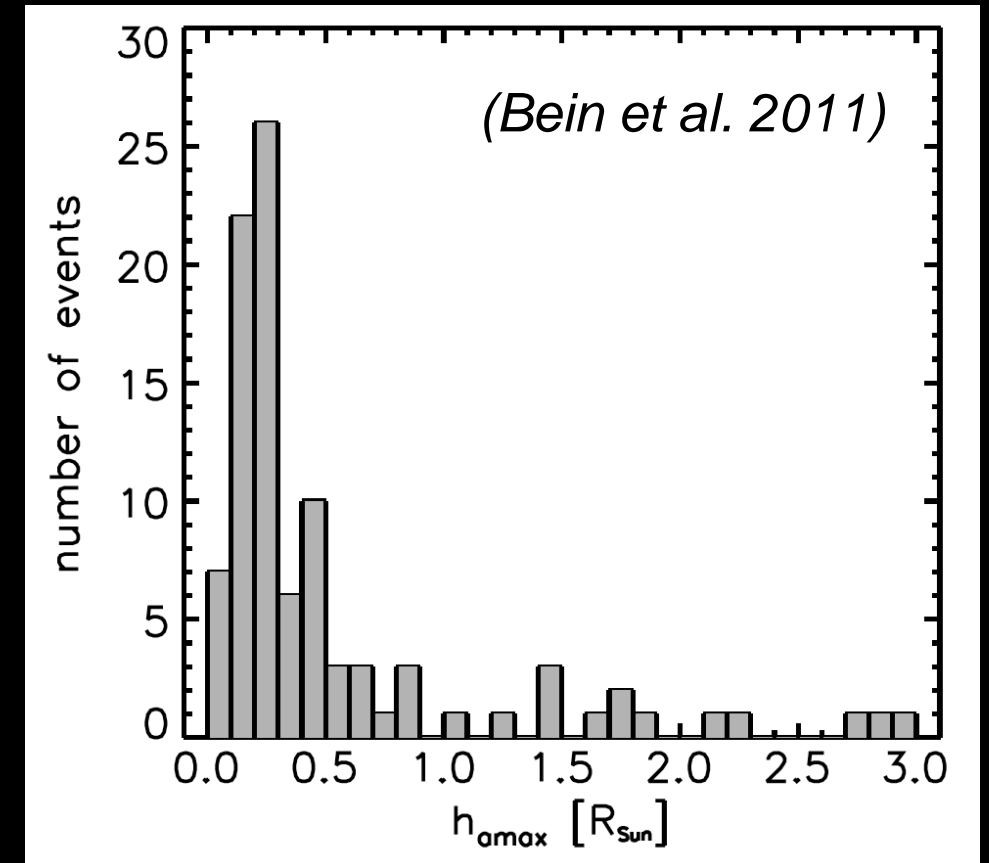
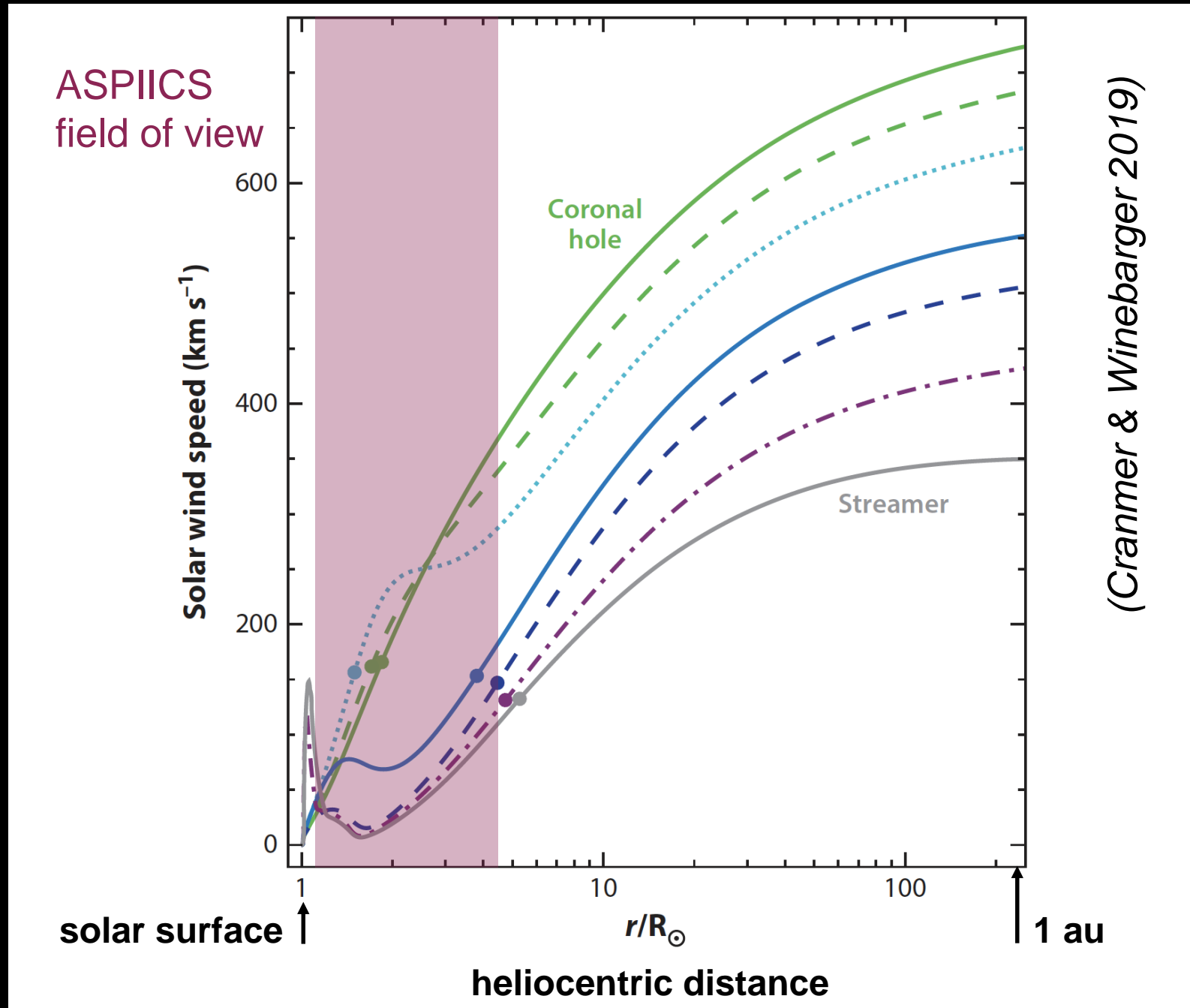


# Observing solar corona using formation flying: Apollo-Soyuz Test Project (1975)



First formation flying  
coronagraphic experiment:  
images taken by the crew of  
Soyuz with Apollo occulting  
the Sun.

# Why are observations of the inner solar corona so important?

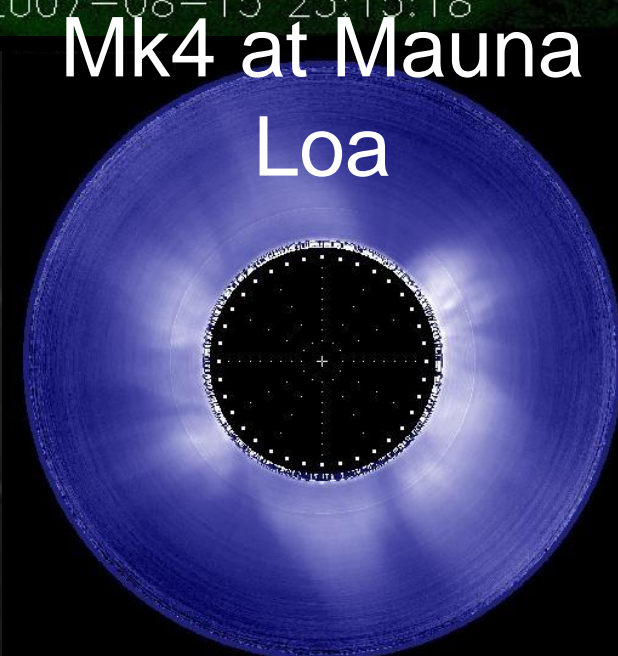
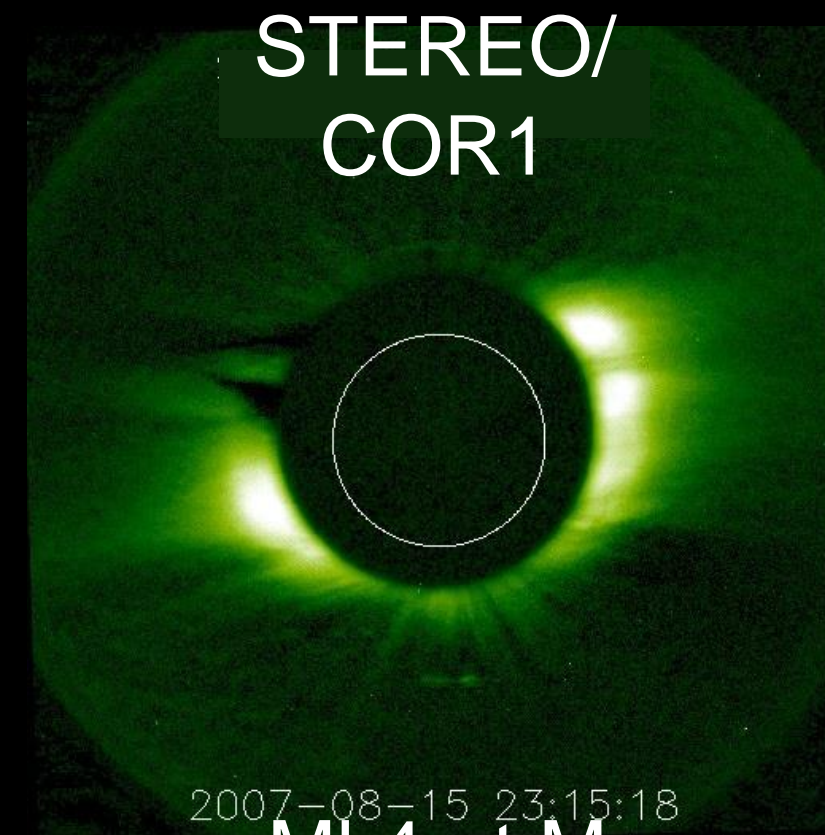
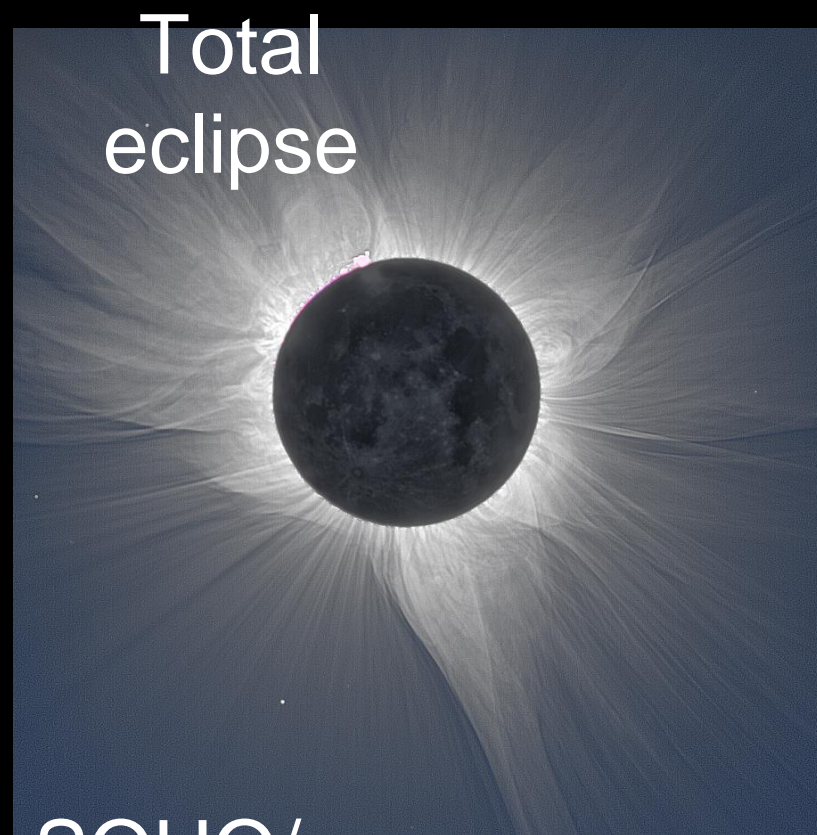


- Simulations show important acceleration of the solar wind (to supersonic velocities) around  $2\text{--}3 R_{\odot}$  from the center of the Sun.
- The peak acceleration of coronal mass ejections (CMEs) also occurs in this region.



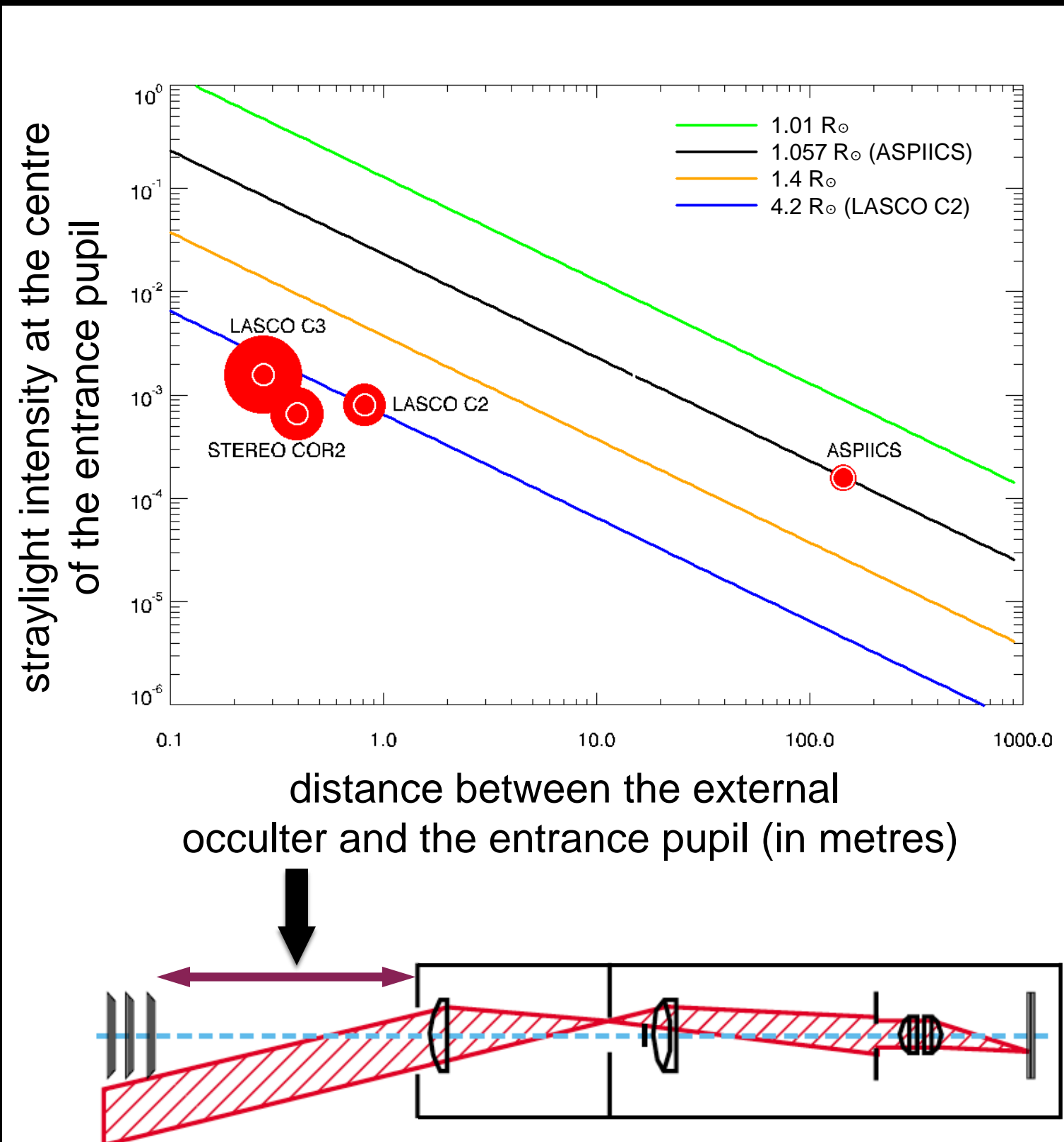
# Why wasn't the inner corona observed in much detail before?

- A short answer: **straylight**
- Straylight in coronagraphs is **very difficult to suppress** when the corona below  $2 R_{\odot}$  is observed.
- High straylight means that signal-to-noise ratio, the contrast of small-scale features, and the effective spatial resolution become low.



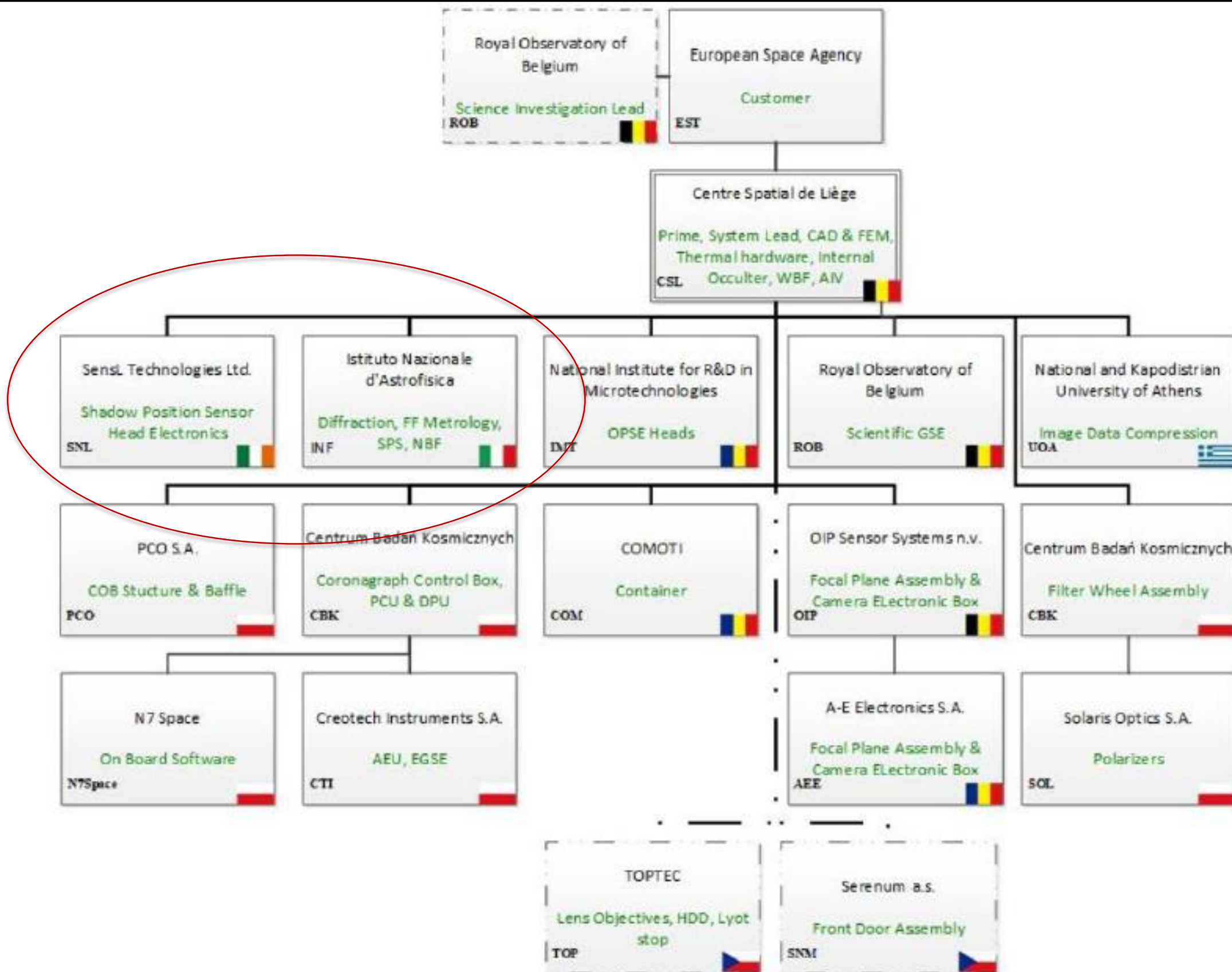
# Why will PROBA-3/ASPIICS do the job?

- The straylight critically depends on the **distance between the external occulter and the entrance pupil**.
- In ASPIICS, this distance is around two orders of magnitude larger than that in any other coronagraph built so far.
- This increase of distance allows in the same time:
  - to **reduce the position of the inner edge** of the field of view from  $2.2 R_{\odot}$  (LASCO C2) to  $1.098 R_{\odot}$  (ASPIICS),
  - to have the **straylight 5 times lower** than that in other coronagraphs.





# PROBA-3/ASPIICS Industrial Organization



# ASPIICS Overview



**Container**



**COMOTI**  
ROMANIAN RESEARCH &  
DEVELOPMENT INSTITUTE FOR  
GAS TURBINES

**COB Structure**



**PCO**

POLSKI HOLDING OBRONNY

**FPA/CEB/Radiator**



**OIP  
Sensor Systems**



**Optics**



**toptec**

**FDA**



**SERENUM**



**OPSE**



**SPS**

INAF

**sensL**

sense light



**FWA**



**SOLARIS**  
OPTICS

**CCB**



**CBK TECH**  
INSTRUMENTS S.A.



**SciEGSE**



**Royal Observatory  
of Belgium**



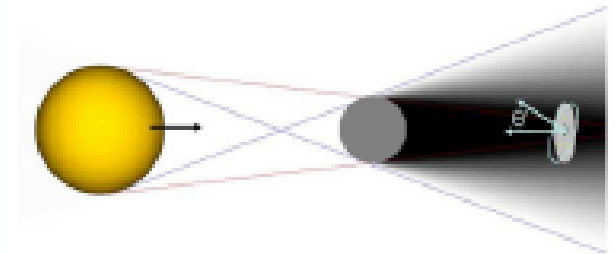
**UoA**





# Italy's Contribution to PROBA-3/ASPIICS: Formation-flying Metrology System

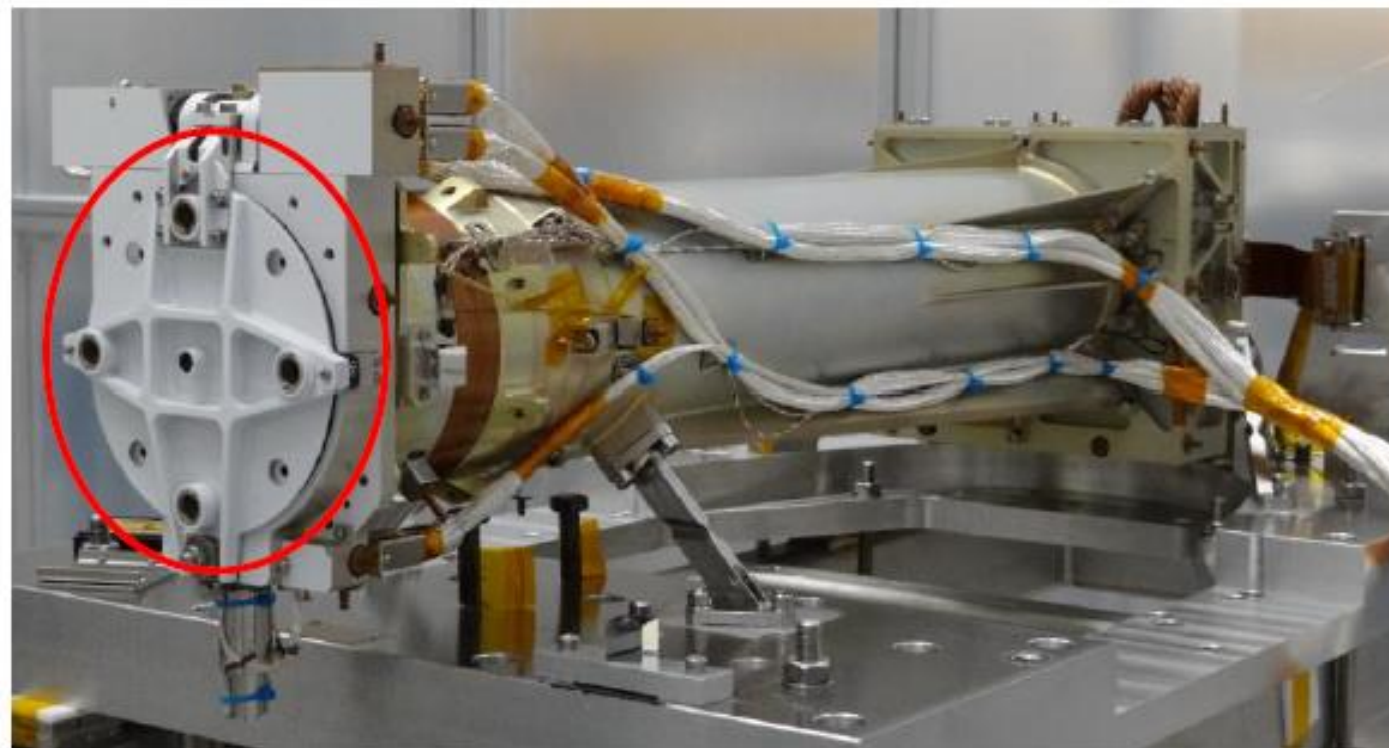
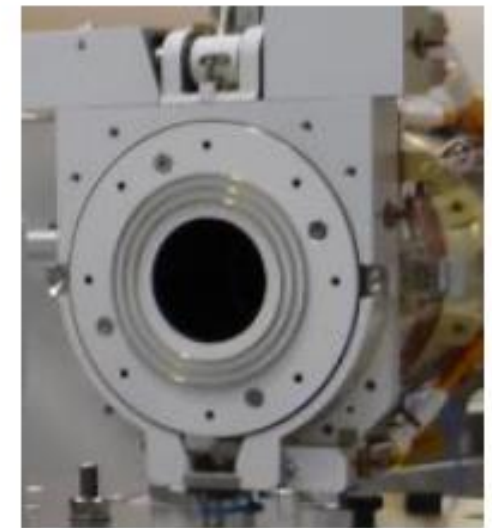
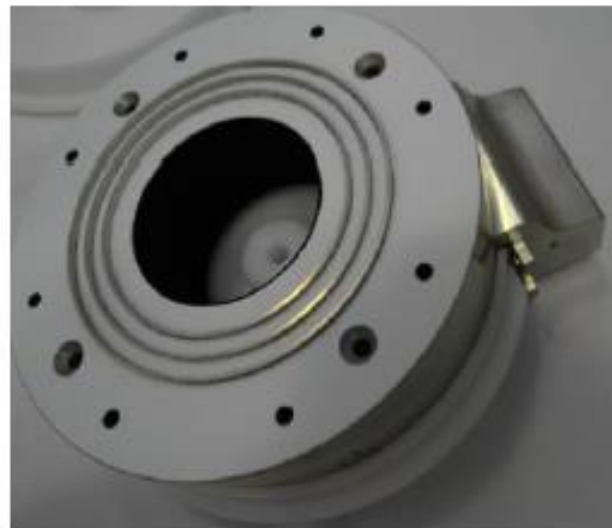
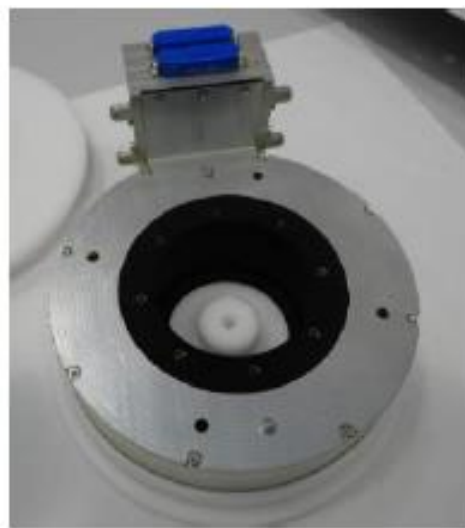
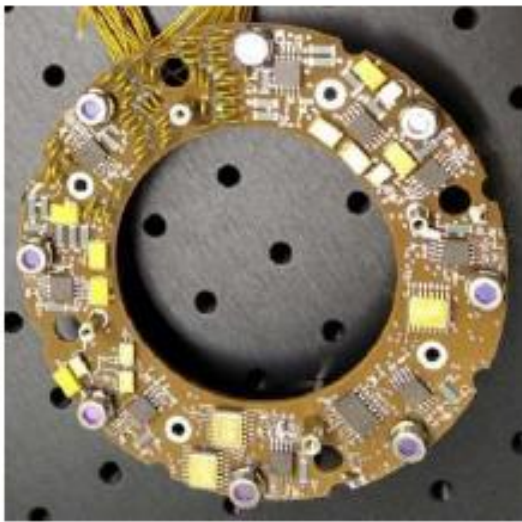
- Shadow Position Sensor (SPS)
  - 8 photodiodes surrounding the coronagraph pupil
  - On-board algorithm for position estimation with respect to shadow



- SPS: 8 SiPM (Silicon PhotoMultiplier) assembled, together with the control electronics, within a mechanical flange installed in front of the ASPIICS telescope
- SPS will monitor the symmetry of the shadow generated by the occulter on the ASPIICS's pupil plane
- A specialized algorithm will measure with a high accuracy ( $<0,5\text{mm}$ ) any displacement of the shadow with respect to the telescope reference system

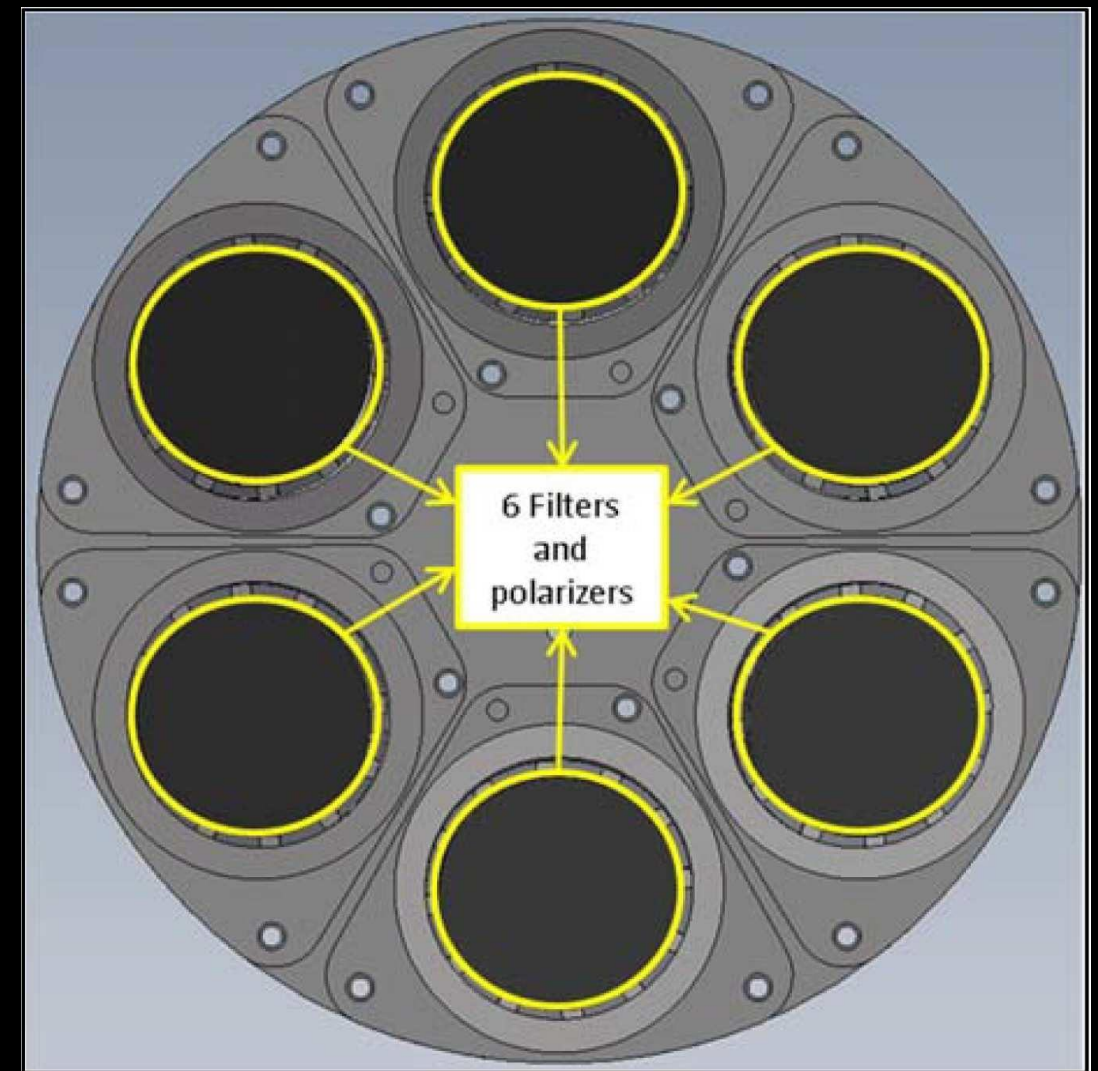
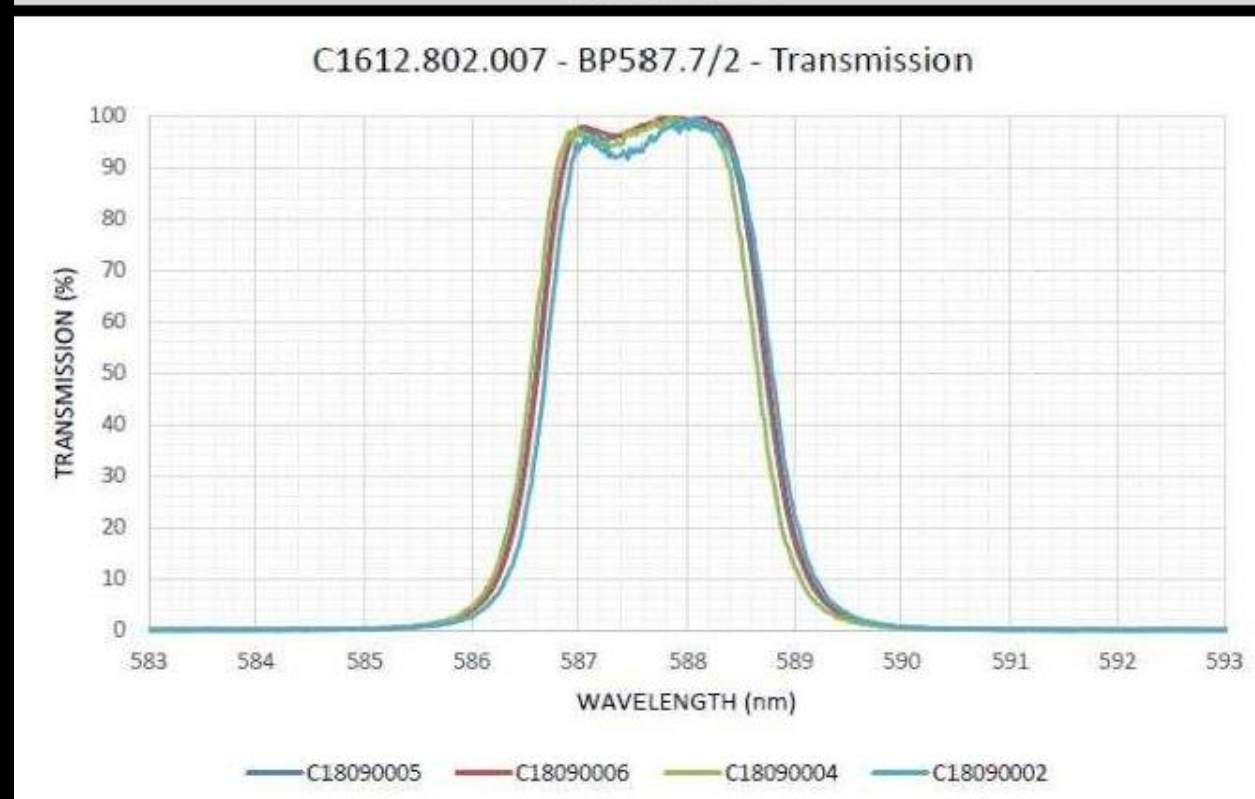
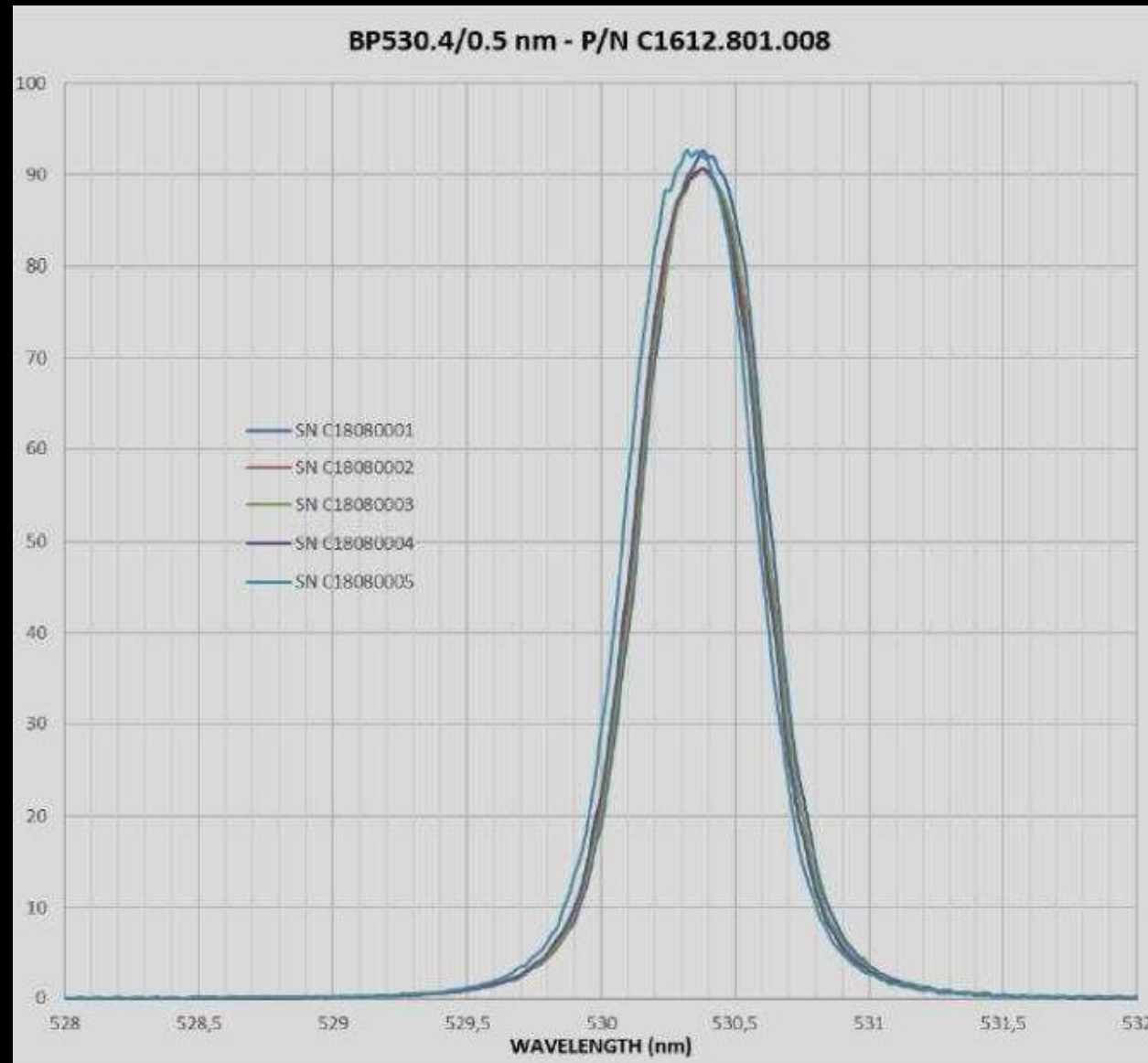
# INAF Shadow Positioning Sensors and formation-flying for ASPIICS/PROBA-3

SPS FM successfully tested (FFT) and baked-out





# INAF-Filters for ASPIICS



**BP filter 530.4/0.6 nm**

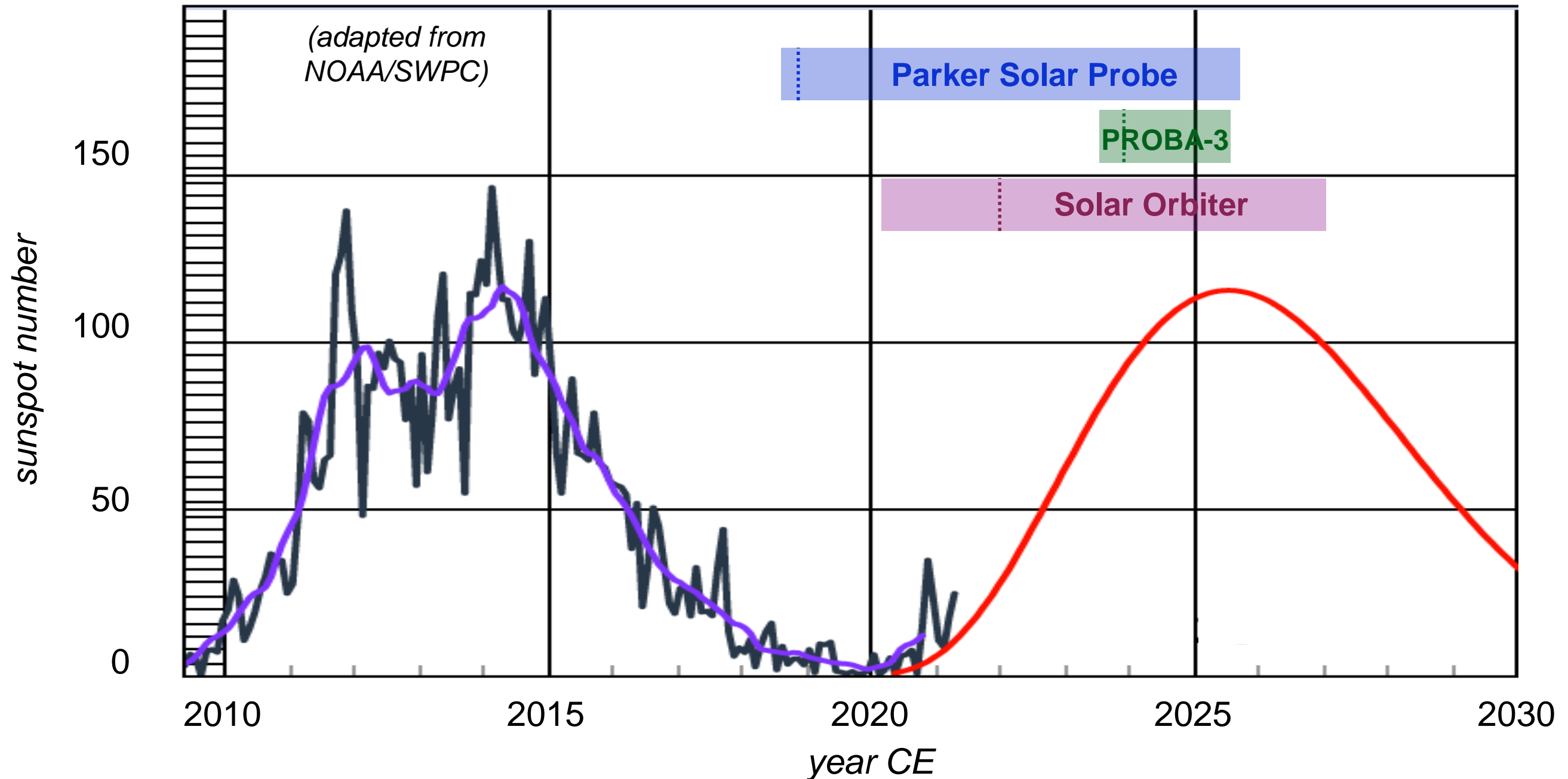
**BP filter 587.7/2 nm**

# ASPIICS Calibration in the INAF Optical Payload Systems (OPSys) Facility





# PROBA-3 in synergy with other missions with INAF participation







# Criticalità

- **Finanziamento ESA solo fino al commissioning:  
2023 Giugno + 6 mesi**
- **Finanziamento per i due anni di operazioni  
scientifiche (2023-2025) – da acquisire**
- **Possibile inclusione delle operazioni di  
ASPIICS/PROBA-3 nel rinnovo del contratto ASI  
(2022-2025) per le operazioni di Metis-Solar Orbiter?**

Thank you  
for your attention!

